

# Mantle anisotropy beneath the Earth's mid-ocean ridges

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## Abstract

Observations of seismic anisotropy at oceanic spreading centres offer insights into mid-ocean ridge processes and the formation of new plates. Here, remote observations of seismic anisotropy beneath mid-ocean ridges are made using measurements of source-side shear wave splitting. Over 100 high-quality measurements are made using earthquakes that occur near mid-ocean ridges and transform faults, but are observed at teleseismic distances. In general, for off-axis ridge events, the polarisation of fast shear waves,  $\phi''$ , is approximately parallel to the spreading direction. Nearer the ridge ( $\lesssim 50$  km),  $\phi''$  becomes more scattered and is often ridge-parallel. Delay times,  $\delta t$ , tend to increase from  $<1$  s near the ridge axis to  $\sim 3$  s further away. Slow-spreading regions (Gakkel and Southwest Indian Ridges) show smaller amounts of splitting than faster spreading centres. At transform zones, the pattern is more complex. Coverage beneath the East Pacific Rise is especially good, and we observe a systematic increase in delay times in S wave splitting measurements compared to previous SKS splitting observations made at ocean-bottom seismometers. One compatible explanation is the presence of horizontally-aligned, connected layers of melt at depth; this is also compatible with other observations of the 'LAB' discontinuity and surface-wave derived measurements of radial anisotropy.

*Key words:* mid-ocean ridges, seismic anisotropy, LAB, mantle dynamics, LPO

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<sup>1</sup> **1. Introduction**

<sup>2</sup> Although it is well known that mid-ocean ridges (MORs) mark sites where oceanic lithosphere is created, there is still considerable uncertainty about mantle processes near ridges and how melt is extracted to form new crust. It has been long understood that viscous shearing leads to the lattice-preferred orientation (LPO) of mantle minerals at spreading centres (e.g., Hess, 1964; Blackman et al., 1996; Tommasi et al., 1999). Additionally, upwelling and decompression lead to melt generation, and shearing and strain partitioning can cause melt segregation (Phipps Morgan, 1987; Holtzman and Kendall, 2010). Both effects can impart a significant anisotropic signature on seismic waves, measurements of which can be therefore used to probe the dynamics of the Earth's upper mantle (UM) beneath ridges.

<sup>11</sup> Measurements of two orthogonally polarised and independent shear waves (i.e., shear wave splitting) are the most unambiguous observation of anisotropy, and are now routinely made in continental regions, or on oceanic islands (for reviews, see for instance Savage, 1999; Long and Silver, 2009). With UM anisotropy, the orientations of fast shear waves, as derived from splitting measurements, are usually interpreted in terms of LPO in peridotites, where olivine a-axes align roughly parallel to mantle flow directions (e.g., Mainprice, 2007). The delay time between the fast and slow shear-waves is proportional to the magnitude of the anisotropy and the extent of the anisotropic region.

<sup>19</sup> Whilst subduction zones and orogens are well sampled, MORs have not been routinely investigated because of significant logistical problems with placing seismometers on the seafloor. Experiments using ocean-bottom seismometers (OBSs) (Blackman et al., 1993, 1995b; Wolfe and Solomon, 1998; Hung and Forsyth, 1999; Barclay and Toomey, 2003; Harmon et al., 2004) have provided vital insights into MOR processes, though there are still very few observations of shear wave splitting at MORs. Using teleseismic phases (e.g., SKS), these few studies generally reveal fast shear wave polarisations parallel to the direction of plate spreading, with increasing values in delay times moving away from the ridge axis (Wolfe and Solomon, 1998; Hung and Forsyth, 1999; Harmon et al., 2004). These observations are consistent with interpretations of olivine LPO as originally proposed by Hess (1964) (based on

29 observations of P-wave anisotropy) and as modelled by Blackman et al. (1996). In contrast,  
30 shallow earthquakes measured within the axial valley show a fast shear-wave orientations  
31 in the crust that are parallel to the ridge axis, which are attributed to aligned cracks and  
32 layered intrusions of volcanic material (Barclay and Toomey, 2003). Blackman et al. (1996,  
33 1995a, 1993) explained the early arrival of P-waves across the southern Mid-Atlantic Ridge  
34 in terms of the vertical alignment on olivine a-axes in a mantle wedge beneath the ridge axis.  
35 Subsequent modelling has suggested that the vertical alignment of melt in films, pockets or  
36 bands would also be very effective in generating shear-wave splitting in near-vertically ar-  
37 riving teleseismic phases (e.g., Kendall, 1994; Blackman and Kendall, 1997; Holtzman and  
38 Kendall, 2010), and would also predict ridge-parallel fast shear-wave polarisations.

39 Previous studies of anisotropy beneath MORs in a global context have been undertaken  
40 using surface waves to infer azimuthal anisotropy (see e.g., Becker et al., 2007). Debayle  
41 et al. (2005), for instance, show that beneath MORs, fast orientations are generally similar  
42 to the spreading direction, however the behaviour beneath transform zones is more complex  
43 and such surface wave studies are limited in their horizontal resolution. It is also the case  
44 that even for the simpler case of global inversions for radial anisotropy in the UM, *a priori*  
45 corrections for the crust have a strong effect on the results of such inversions (Ferreira et al.,  
46 2010). Hence whilst this should be less of a problem in the region of MORs, where the crust  
47 is simple, caution in directly interpreting such results is still advisable. In a more localised  
48 study Gaherty (2001) and Delorey et al. (2007) mapped vertical and lateral variations in  
49 anisotropy beneath the Reykjanes Ridge. Using sources on the Gibbs fracture zone and  
50 receivers on Iceland, differences in Love and Rayleigh wave arrival times revealed faster  
51 vertically-polarised Rayleigh waves than horizontally-polarised Loves waves near the ridge  
52 axis and at depths less than 100 km. This observation is consistent with either the vertical  
53 alignment of olivine a-axes or a melt-induced anisotropy, but Holtzman and Kendall (2010)  
54 argue that the latter is more likely.

55 In this study we evaluate MOR anisotropy using measurements of shear wave splitting  
56 which occur beneath the earthquake, rather than the receiver, using direct S waves—a  
57 technique often termed ‘source-side splitting’ (e.g., Schoenecker et al., 1997; Nowacki et al.,

58 2010; Foley and Long, 2011). Using seismic stations with well-characterised anisotropy in  
59 the UM beneath the receivers, we can remove the effect of the splitting on the receiver side  
60 and measure only that which occurs beneath the source. We then attempt to interpret these  
61 observations in the context of previous observations and proposed mechanisms for anisotropy  
62 beneath a MOR.

63 **2. Methods and data**

64 *2.1. Shear wave splitting*

65 We aim to measure the seismic anisotropy beneath MORs around the world using the  
66 primary observable it produces, shear wave splitting. We use the ‘minimum eigenvalue’  
67 technique of Teanby et al. (2004) (which is an extension of that of Silver and Chan (1991)),  
68 which removes splitting by effectively maximising the linearity of the horizontal particle  
69 motion for a given pair of splitting parameters: the fast direction,  $\phi$ , and the delay between  
70 the fast and slow waves,  $\delta t$ . Where measurements are available for an event at more than one  
71 station within an azimuthal range of  $15^\circ$ , we use the method of Wolfe and Silver (1998) to  
72 stack the small eigenvalue ( $\lambda_2$ ) surfaces, with a backazimuth-independent implementation.  
73 This significantly reduces the errors when for some stations the measurement is very near  
74 null, as the initial polarisation is close to the fast direction beneath the event.

75 In this study, we make the common assumption that the lower mantle above D'' is not  
76 significantly anisotropic: despite some evidence of its presence in the uppermost lower mantle  
77 (Wookey and Kendall, 2004), several studies support this assumption (e.g., Meade et al.,  
78 1995; Montagner and Kennett, 1996; Panning and Romanowicz, 2006; Kustowski et al.,  
79 2008). Hence we can infer that any splitting is caused by anisotropy in the UM beneath the  
80 source and receiver. If we have prior knowledge of splitting in the UM beneath the receiver,  
81 we may correct for this and analyse the S phase, retrieving the splitting caused by anisotropy  
82 beneath the source. We interpret the fast direction of the receiver-corrected signal simply  
83 by considering the fast orientation at the source,  $\phi'' = \text{azimuth} + \text{backazimuth} - \phi$ . This  
84 simple geometric relationship is true for rays which are vertically incident at the surface,

85 but is only less accurate by a few degrees than a fully slowness-dependent expression, for  
86 the range of slownesses in this study. This error is generally less than the uncertainty in the  
87 method.

88 *2.2. SKS UM splitting corrections*

89 Seismic anisotropy in the continental UM (where our stations are located) appears to be  
90 ubiquitous, and is typically measured using phases such as SKS, PKS and SKKS; SKS is  
91 the most commonly used. It converts from a compressional to an S wave upon exiting the  
92 outer core, so begins its ascent through the mantle with no splitting present. It is polarised  
93 radially, hence it is also polarised parallel to the backazimuth at the receiver. SKS also  
94 propagates steeply through D'', which is known to be anisotropic in various places in the  
95 lower mantle (see reviews by Kendall and Silver, 1998, 2000; Lay et al., 1998; Nowacki et al.,  
96 2011). However, we assume that any contribution to splitting in the phase along this section  
97 is minor, as it has spent relatively little time in D''. Studies on a global scale support this  
98 approximation (Niu and Perez, 2004; Restivo and Helffrich, 2006), though any strong effects  
99 should be visible and display backazimuthal variation in splitting parameters (Hall et al.,  
100 2004).

101 Because we wish to remove UM anisotropy from the S phase, we choose seismic stations  
102 which have many SKS splitting measurements along a variety of backazimuths. If dipping  
103 or multiple layers of anisotropy exist beneath the station, then we expect the results to show  
104 a 90° or 180° periodicity to the measurements of  $\phi$  and  $\delta t$  (Silver and Savage, 1994). We do  
105 not use stations which exhibit such measurements, as complicated UM anisotropy beneath  
106 the receiver is difficult to infer uniquely, and therefore we cannot confidently remove its  
107 effects on direct S phases, as they will be arbitrarily polarised compared to the backazimuth,  
108 depending on the source mechanism and anisotropic fabric they have encountered near the  
109 source. Stations which exhibit backazimuthal variation in SKS splitting may also do so  
110 because of laterally heterogeneous anisotropy beneath them. We also avoid using such  
111 stations for similar reasons. Our approach is slightly different from some authors, who opt  
112 to use stations which appear to show no anisotropy beneath them (Foley and Long, 2011),

113 however these are rare and UM anisotropy appears to be the norm, rather than isotropy.

114 Some studies using surface waves (e.g., Gaherty, 2004) or combining long-period waves  
115 with SKS splitting measurements (Yuan and Romanowicz, 2010) show evidence for multiple  
116 layers of anisotropy beneath North America, including beneath stations which do not exhibit  
117 backazimuthal variation in SKS splitting. It is possible that the same might also be true  
118 beneath Ethiopia. It therefore may be that any complexity of anisotropy (e.g., multiple  
119 layers) is not imaged in backazimuthal variations in splitting in SKS alone at our stations,  
120 despite the apparent requirement for it in surface waves, thus breaking the assumption of  
121 simple sub-station anisotropy. However, because our study is concerned only with shear wave  
122 splitting, so long as the splitting experienced by SKS and S is similar enough, this should  
123 not impact on the use of SKS splitting measurements as corrections here. The similarity  
124 between splitting in SKS and S for a given anisotropy beneath the seismic stations used here  
125 is therefore the critical assumption we make in this study.

126 In order to be confident of our measurements, we wish to make several for each MOR  
127 event, and so we choose from sets of stations in North America and Ethiopia, where extensive  
128 SKS splitting studies have been conducted (Ayele et al., 2004; Barruol et al., 1997; Evans  
129 et al., 2006; Fouch et al., 2000; Kendall et al., 2005; Liu, 2009; Niu and Perez, 2004; J.O.S.  
130 Hammond, pers. comm., 2010). As explained, we reject stations with apparently complicated  
131 sub-station anisotropy. SKS measurements for two example stations used in this study are  
132 shown in Supplementary Figure 1. The stations used in this study and the SKS splitting  
133 parameters used as UM corrections are shown in Supplementary Figure 2.

134 We use these SKS-derived corrections and analyse the direct S phase from events beneath  
135 MORs, applying the correction during the analysis. We note that even though reciprocity  
136 must apply along the ray path (see, for example, Kendall et al., 1992), the splitting operators  
137 are not commutative (Wolfe and Silver, 1998), so it is essential to make the corrections in  
138 the correct order (see Wookey and Kendall, 2008; Wookey et al., 2005). As a further check  
139 that the correction is valid, after the measurement we check that the source polarisation of S  
140 matches that predicted by the event's focal mechanism (see section 2.4). This helps mitigate  
141 against the possibility that the S phase we analyse is contaminated by depth phases (sS and

<sup>142</sup> pS), as these will generally alter the apparent source polarisation of the combined phase to  
<sup>143</sup> be different to that expected from the CMT solution. A difference in the measured source  
<sup>144</sup> polarisation may also occur due to the application of an incorrect receiver correction in the  
<sup>145</sup> analysis (see below), which also leads us to reject measurements where the two are not in  
<sup>146</sup> agreement within 15°.

<sup>147</sup> *2.3. Testing the use of receiver corrections*

<sup>148</sup> Whilst we make every effort to ensure that we use seismic stations which have very well-  
<sup>149</sup> characterised anisotropy beneath them, some error will be present in the measurement. Part  
<sup>150</sup> of the difference will result because of the different slownesses between the S waves we study  
<sup>151</sup> and the SKS phases used to make the splitting measurements we use as station corrections,  
<sup>152</sup> but the difference is usually negligible in  $\phi$  and very small in  $\delta t$  (see discussion in Nowacki  
<sup>153</sup> et al., 2011). The majority of the error therefore likely comes from the assumption that  
<sup>154</sup> the anisotropy is simple beneath the station, and that the SKS splitting measurements are  
<sup>155</sup> accurate.

<sup>156</sup> We conduct synthetic tests to determine how large the uncertainty in the measured  
<sup>157</sup> source splitting parameters are when an ‘incorrect’ receiver correction is used. We apply a  
<sup>158</sup> known initial amount of splitting (the ‘source-side’ splitting,  $\phi_s^{\text{true}}, \delta t_s^{\text{true}}$ ) to a synthetic wave  
<sup>159</sup> of dominant frequency 0.1 Hz, then a known receiver-side splitting,  $\phi_r^{\text{true}}, \delta t_r^{\text{true}}$ . We then  
<sup>160</sup> analyse the splitting in the wave with a range of receiver corrections ( $\phi_r^{\text{trial}}, \delta t_r^{\text{trial}}$ ) to obtain  
<sup>161</sup> the ‘observed’ splitting parameters at the source ( $\phi_s^{\text{trial}}, \delta t_s^{\text{trial}}$ ) and compare the known and  
<sup>162</sup> measured source-side splitting. The procedure can be repeated for any combination of true  
<sup>163</sup> source and receiver splitting operators, and all receiver ‘corrections’.

<sup>164</sup> Supplementary Figure 3 shows the difference between the true and measured splitting  
<sup>165</sup> parameters where  $\phi_s^{\text{true}} = 20^\circ$ ,  $\delta t_s^{\text{true}} = 1.0\text{ s}$ , and  $\phi_r^{\text{true}} = 0^\circ$ ,  $\delta t_r^{\text{true}} = 1.0\text{ s}$ . The difference in  
<sup>166</sup> fast orientation,  $\Delta(\phi'') = \text{abs}(\phi_s^{\text{trial}} - \phi_s^{\text{true}})$ , is within about 15° whilst the trial receiver  
<sup>167</sup> correction is within about 40° and 0.4 s of the true receiver splitting parameters. In these  
<sup>168</sup> limits, the difference in source delay time,  $\Delta(\delta t) = \text{abs}(\delta t_s^{\text{trial}} - \delta t_s^{\text{true}})$ , is up to 0.6 s. Consis-  
<sup>169</sup> tent with previous tests using real data (Russo and Mocanu, 2009), we find that errors in  $\delta t_r$

170 appear to cause the largest uncertainty in the ‘observed’ source-side splitting parameters.  
171 Supplementary Figure 4 shows the case when  $\phi_s^{\text{true}} = 45^\circ$ ,  $\delta t_s^{\text{true}} = 1.0 \text{ s}$ .

172 We also show in Supplementary Figures 3 and 4 the difference between the known and  
173 measured source polarisation for a range of  $\phi_r^{\text{trial}}$  and  $\delta t_r^{\text{trial}}$ . The initial polarisation is  $0^\circ$  in  
174 both cases. Again, the difference in the true and trial receiver delay times plays a large rôle,  
175 and when the ‘observed’ source-side splitting parameters are most inaccurate, the source  
176 polarisation is often incorrect by about  $10\text{--}20^\circ$ . Hence the use of the source polarisation as  
177 a diagnostic of the quality of the result is important and helpful.

178 Finally, manual inspection of the results indicates that in several instances the ‘observed’  
179 source splitting parameters would be classified as null events, especially where the delay times  
180 are large as shown in Supplementary Figures 3 and 4. This also highlights the strength of  
181 using manual inspection or an automated null-classifying scheme to maintain the integrity  
182 of measurements (Wuestefeld et al., 2010). When all of these diagnostics are included, and  
183 the receiver corrections are within an acceptable uncertainty range of within about  $20^\circ$  for  
184 the fast direction and 0.4 s for the delay time, we can be confident that the source-side shear  
185 wave splitting measurement is a true reflection of the splitting which has affected the wave  
186 in the source anisotropic region.

#### 187 *2.4. Event locations and focal mechanisms*

188 In order to make inferences about anisotropy beneath MORs, it is obviously important  
189 to accurately know the earthquake location. Because MOR events typically have large  
190 uncertainties on their locations in time and space, where possible (for events before 2008)  
191 we take these parameters from the ISC’s relocations using the EHB algorithm (Engdahl et al.,  
192 1998). The published horizontal uncertainty in the standard ISC locations is approximately  
193 20 km; for the EHB locations in this study, the average uncertainty is 7 km.

194 The location of an event—whether beneath a ridge segment or transform zone—may  
195 affect the type of anisotropy we expect, hence each event was assigned to one of these  
196 categories based on its location relative to the bathymetry (Smith and Sandwell, 1997), and  
197 in part its focal mechanism. These were taken from the Global CMT catalogue. Where

198 there was ambiguity from bathymetry, the event was classified as being located on a ridge  
199 if the focal mechanism was mainly dip-slip, and as on a transform if mainly strike-slip.

200 *2.5. Dataset*

201 We consider events of  $M > 5.0$ , depth  $\leq 35$  km, in the epicentral distance range  $55^\circ \leq$   
202  $\Delta \leq 82^\circ$ , which are located on the East Pacific Rise (EPR), Mid-Atlantic Ridge (MAR),  
203 Gakkel Ridge, and the Southwest and Southeast Indian Ridges (IRs) (Figure 1). At distances  
204 less than  $\sim 55^\circ$ , the difference in incidence angle between SKS and S becomes large enough  
205 that the vertical-incidence approximation may no longer be appropriate, and increases the  
206 possibility that an SKS-correction for UM anisotropy is inaccurate; beyond  $\sim 82^\circ$ , the S  
207 phase interferes with ScS, or there may be a triplication due to the presence of the D'' layer,  
208 contaminating the S signal in the splitting analysis. We of course also wish to avoid D''-  
209 traversing rays due to the anisotropy present there. Events deeper than 35 km are unlikely  
210 to occur near MORs, and such depths may indicate a poor event location. The seismograms  
211 were band-pass filtered between 0.001 and 0.3 Hz.

212 After selection, over 2000 events matched the criteria between 1979 and 2009, according  
213 to the USGS National Earthquake Information Center (NEIC) and International Seismolog-  
214 ical Centre (ISC) catalogues. Due mainly to signal-to-noise requirements,  $\sim 400$  events were  
215 retained for analysis, leaving  $\sim 820$  event-station pairs.

216 During analysis, we apply a strict set of criteria to select the optimum splitting results.  
217 Only non-null results which meet the following are retained: (i) acceptable signal-to-noise  
218 ratio on both horizontal components; (ii) clear elliptical particle motion before analysis; (iii)  
219 clear linearisation of particle motion when corrected; (iv) measured source polarisation is  
220 within  $15^\circ$  of the CMT-predicted source polarisation; (v) clear minimum on the  $\lambda_2$  surface.  
221 A quality of 1 (excellent) to 4 (very poor) is assigned manually to each measurement. Null  
222 measurements are retained, provided the signal-to-noise ratio is adequate and particle motion  
223 is clearly linear before analysis, but after correction for receiver anisotropy.

224 Following analysis, 350 measurements of splitting of ‘fair’ (3) quality or better beneath  
225 67 events comprise the dataset. Of these, 122 are of quality ‘good’ (2) or better. There are

226 189 null measurements. The events have magnitude range  $4.4 \leq M_b \leq 6.7$ , and depth range  
227 0–33 km.

### 228 3. Results

#### 229 3.1. East Pacific Rise

230 The EPR is the best-sampled MOR segment in this work. Our results agree excellently  
231 with SKS splitting results from ocean-bottom seismometers (OBSs) deployed as part of the  
232 MELT and GLIMPSE projects (Wolfe and Solomon, 1998; Harmon et al., 2004) (orange  
233 bars, Figure 2). Here and in the OBS experiments,  $\phi''$  or  $\phi_{\text{SKS}}$  is approximately parallel to  
234 the spreading direction, with  $\delta t$  varying from 1–3 s, depending on distance from the ridge  
235 axis. Figure 3 shows the variation of splitting parameters with distance for results on the  
236 EPR which are classified as ‘ridge’ events, alongside the MELT and GLIMPSE data.

237 Away from the straightest segments of the EPR, where frequent fracture zones offset  
238 the ridge axis, the pattern of observed splitting is different. There is no clear spreading  
239 direction-parallel trend to  $\phi''$ , and the change in  $\delta t$  is also complicated. At about  $-5^\circ$   
240 latitude, for example,  $\phi''$  seems to change over a short distance by  $\sim 70^\circ$  from spreading  
241 direction-parallel to transform zone-perpendicular. Similarly, the pattern of  $\phi''$  and  $\delta t$  east  
242 of the Pacific–Nazca–Antarctic triple-junction is also complex, with a variation of  $\phi''$  from  
243 parallel to perpendicular to the Challenger Fracture Zone (at about  $-35^\circ$  latitude).

#### 244 3.2. Mid-Atlantic Ridge

245 Events which produced ‘good’ source-side splitting measurements were limited to lati-  
246 tudes between  $-40^\circ$  and  $15^\circ$ . Very few events of sufficient magnitude are reported in the  
247 catalogues along the Reykjanes Ridge, and no ‘good’ measurements could be made north  
248 of the equator. We note that few measured earthquakes occur along clear, linear ridge seg-  
249 ments along the MAR, and most seismicity for which we have results is located instead on  
250 the transform zones. Nonetheless, the few events clearly beneath ridges (e.g., the stack at  
251  $-30^\circ$  latitude) do seem to show spreading direction-parallel  $\phi''$ . This agrees approximately

252 with SKS splitting measurements made at ASCN (Butt Crater, Ascension Island; Wolfe and  
253 Silver, 1998) and SHEL (Horse Pasture, St. Helena; Behn et al., 2004).

254 Along transform zones, about half the results show  $\phi''$  close to the spreading direction,  
255 whilst many show large ( $\sim 2.5\text{--}3$  s)  $\delta t$  and  $\phi''$  roughly perpendicular to the strike of the  
256 transform. The dependence of splitting parameters upon distance along the transform zone,  
257 away from the nearest ridge segment, is shown in Supplementary Figure 5. The pattern  
258 shows considerable variation near the ends of the transform zones, close to the ridge axes,  
259 perhaps related to the complex tectonic environment and resultant shearing and melt pro-  
260 duction. However, there is a decrease in the maximum delay time as distance from the ridge  
261 axis increases, possibly indicating a reduced contribution from a mechanism of anisotropy  
262 arising due to melt or other sub-axial process.

263 Two events on the MAR gave results at stations in both North America and Ethiopia. In  
264 this case, we may examine the azimuthal dependence of the splitting. Figure 5 shows equal-  
265 area lower-hemisphere stereoplots of the splitting parameters, which are notably different  
266 along the two different azimuths. Splitting measured at North American stations for both  
267 events has smaller  $\delta t$  (stacked splitting parameters:  $\phi'' = (25 \pm 4)^\circ$ ,  $\delta t = (1.9 \pm 0.1)$  s), whereas  
268  $\delta t$  is larger when measured along the other azimuth at Ethiopian stations ( $\phi'' = (76 \pm 2)^\circ$ ,  
269  $\delta t = (2.6 \pm 0.1)$  s). At this limited range of slownesses, there is not much variation in the  
270 angle away from the vertical for the rays, so the differences primarily arise due to azimuth.  
271 The Fresnel zones of the two rays of period 20 s stop overlapping significantly when deeper  
272 than  $\sim 200$  km, so if heterogeneity were the cause, then the majority of the anisotropy would  
273 need to be present below this.

### 274 3.3. Gakkel Ridge

275 Ten results from events on the Gakkel Ridge were of ‘good’ or better quality, with an  
276 equal number of null results. The splitting parameters are shown in Figure 6A. It is notable  
277 that most results show a small amount of splitting ( $\langle \delta t \rangle = 1.1$  s), and there is a higher  
278 proportion of null results than in other regions. The splitting that is present is often ridge-  
279 parallel. The spreading rate predicted by NUVEL-1A (DeMets et al., 1994) increases from

280 ~6 to 18 mm a<sup>-1</sup> from right to left in Figure 6A, however there is no clear corresponding  
281 trend in the amount of splitting. There is also no obvious systematic variation of parameters  
282 for the cluster of events furthest north (rightmost in Figure 6A, circled) with azimuth. A  
283 lower amount of splitting beneath such extremely slow-spreading ridges might be related to  
284 reduced melt production caused by slow exhumation of material and a consequently small  
285 amount of adiabatic decompression melting. If this is the case, the dominant contribution  
286 to seismic anisotropy at teleseismic distances would then be from LPO, yet the axis-parallel  
287 fast orientations we observe are hard to explain via mineral alignment.

### 288 3.4. Southwest and Southeast Indian Ridges

289 Beneath events on the SWIR and SEIR, 43 individual results, allowing three stacked  
290 results, and seven null measurements were made. These are shown in Figure 6B. Again, the  
291 pattern is complicated, and few events lie on ridge segments: most are on transforms. The  
292 Southwest Indian Ridge shows some of the most oblique spreading of any MOR, so it may  
293 help distinguish between processes which lead to anisotropy which is ridge-perpendicular  
294 or spreading direction parallel. However, there are insufficient large earthquakes to make  
295 any strong inferences from source-side splitting. Interestingly, all measurements made from  
296 beneath the ridge segment at longitude 20° appear to be null. This might result from the  
297 absence of anisotropy in the region, but it may also occur if the source polarisation is parallel  
298 or perpendicular the local fast orientation of some anisotropy. With only one azimuth of  
299 measurements and no other events with different source polarisations, it is not possible to  
300 distinguish these scenarios.

## 301 4. Interpretation and discussion

302 In interpreting our results, it is difficult to draw firm conclusions about the behaviour of  
303 MORs in general because of poor sampling, arising from the lack of stations outside USA  
304 and Ethiopia with comprehensive studies published on the backazimuthal variation of SKS  
305 splitting parameters. For the purposes of studying relatively small-magnitude earthquakes  
306 at teleseismic distances such as is done here, networks of stations with such measurements

307 are necessary to allow stacking of data, especially when fast directions are near the source  
308 polarisation. This limitation also means that comparisons between fast and slow ridges are  
309 hard to make. However, our measurements do suggest that splitting near the ridge axis is  
310 greater beneath fast-spreading ridges (full-rate  $> 100 \text{ mm a}^{-1}$ ) than slow-spreading ones.

311 Our measurements of splitting show  $\phi''$  to be very similar to the fast orientations observed  
312 by regional SKS studies near MORs, but these are extremely limited in coverage because  
313 of the practical difficulties in operating such OBS sites. Surface wave studies examining  
314 azimuthal anisotropy globally (e.g., Debayle et al., 2005) can provide better coverage near  
315 MORs, but limited horizontal resolution means changes over relatively small distances (up  
316 to few tens of kilometres and less) cannot be imaged well. Such global measurements tend  
317 to show fast orientations approximately parallel to the spreading direction, but this can vary  
318 by up to  $45^\circ$  in some places, notably near large fracture zones.

319 *4.1. Doldrums FZ observations*

320 The multi-azimuth observations beneath the Doldrums Fracture Zone (FZ) in section  
321 3.2 are interesting (Figure 5). The observation of this azimuthal dependence in splitting  
322 parameters appears to be a robust feature: tests requiring  $\phi''$  to be the same for stations in  
323 Ethiopia and North America show the splitting experienced by the direct S wave beneath  
324 both sets of stations would have to be different by around  $45^\circ$  to that observed in SKS  
325 waves and for which we correct. Such a strong incidence dependence in splitting beneath  
326 the receiver would in all likelihood appear as complexity in SKS splitting observations, and  
327 here we deliberately avoid stations where this is the case. Whilst the limited number of data  
328 prevents detailed analysis, we can speculate on the likely causes of the observed pattern.

329 With two azimuths of observations, we can seek to define an hexagonal symmetry, ori-  
330 ented arbitrarily (called ‘tilted transverse isotropy’, or TTI). If we assume Thomsen’s (1986)  
331 anisotropic parameters  $\delta = \epsilon$  (the case of elliptical anisotropy), we can use the two azimuths  
332 of observations to find the plane of isotropy, or axis of symmetry, by simple trigonometry  
333 (Nowacki et al., 2011). This dips shallowly to the southwest, as shown by the dashed line  
334 in Figure 7. For TTI derived from aligned material, for instance, this would correspond to

335 penny-shaped inclusions having their short axis aligned about  $35^\circ$  from the vertical. This  
336 is in some sense similar to the orientations predicted by simulations (Weatherley and Katz,  
337 2010), which suggest melt should be focussed along northwest–southeast flow lines for a  
338 transform in this orientation. Intriguingly, it also would be consistent with the suggestion  
339 of van Wijk and Blackman (2005), who speculate that the transform fault itself would dip  
340 towards the ridge segment near the ends of the transform.

341 Another likely contributor to seismic anisotropy in the FZ would be the alignment of  
342 olivine in response to flow. Natural samples and deformation experiments show that the  
343 dominant way in which olivine develops an LPO is by slip along [100] (a-direction), on  
344  $\{0kl\}$  or (010) (b-planes), known as D- and A-type olivine respectively. We examine the  
345 possibility that the observed anisotropy at the Doldrums FZ is caused by olivine LPO by  
346 using the method described by Wookey and Kendall (2008) and Nowacki et al. (2010). We  
347 use the single-crystal elastic constants of olivine (Abramson et al., 1997) and mix them in  
348 all proportions with an isotropic average. We then rotate these constants to all possible  
349 orientations and compute the shear wave splitting accrued over a 200 km thick layer for the  
350 two raypaths observed, and plot the orientations producing splitting compatible with the  
351 observations. Figure 7 shows the compatible orientations and degree of alignment as the a-  
352 and b-axes of the aligned olivine on a lower-hemisphere equal-area projection.

353 Compatible orientations of the a-axes are northwest–southeast, with glide planes dipping  
354 north, northwest or west, all shallowly. This direction of shear is not parallel to the spreading  
355 direction or FZ strike, or to the absolute plate motion (APM) (Figure 4). It is consistent  
356 with some dynamics models of ridge transforms (e.g., van Wijk and Blackman, 2005; Sparks  
357 et al., 1993; Phipps Morgan and Forsyth, 1988), though such models also predict low strain  
358 rates in such regions and relate to shorter FZs. This interpretation is, however, inconsistent  
359 with the expected behaviour for a pure strike-slip fault, where strains are sufficiently high  
360 that olivine [100] directions should be parallel to the FZ strike. This incongruity might  
361 support the hypothesis that SPO due to melt or another material is the cause: if olivine [100]  
362 directions were parallel with APM or FZ strike, but some other anisotropy were overprinted,  
363 then we would not necessarily retrieve the olivine orientations with this method. Equally,

such conditions as are present beneath the FZ might lead to the dominance of a different slip system in olivine, in which case flow might still be parallel to the FZ, but not olivine [100] axes. Whilst the uncertainty in our measurements is not insignificant, even with more relaxed constraints on the orientations the picture is much the same (Supplementary Figure 6). We of course neglect other anisotropic phases in this approach, but would expect this to require a stronger texturing in the olivine itself to match observations.

Finally, heterogeneity between the two raypaths away from the source may also be important, as paths to Ethiopia spend considerable distance close to the FZ, whilst paths to North America travel away from this region. This would have to be true at depths of  $\sim$ 200 km, as the Fresnel zones of waves of this period overlap significantly until that depth. Hence the signal may reflect different anisotropies along these two paths, perhaps with the addition of a common anisotropic region immediately below the source. However, other studies (e.g., ) suggest that such strong anisotropy at depths is unlikely, and we prefer to interpret the observations in terms of a mixture of olivine alignment and SPO

#### 4.2. Spreading rate and strength of anisotropy

Dynamic models of MOR accretion have predicted the amount and orientation of shear wave splitting near spreading ridges on the basis of LPO of pure olivine (Blackman et al., 1996), LPO of olivine and enstatite (Blackman et al., 2002; Blackman and Kendall, 2002; Nippress et al., 2007), and combined olivine LPO with the effect of oriented melt pockets (Blackman and Kendall, 1997). In such models, the spreading rate controls the behaviour of upwelling beneath the ridge axis and the shape of the melt-rich region, and hence the orientation and amount of shear wave splitting observed at the surface. In Blackman and Kendall's (2002) simulations, slow spreading ridges (full-rate  $\sim$ 40 mm a $^{-1}$ ) show significant ridge-parallel splitting within  $\sim$ 20 km from the axis because of the requirement that buoyant flow beneath the ridge axis supplies the upwelling material in a small region. Fast-spreading (full-rate  $\sim$ 140 mm a $^{-1}$ ) ridges, by contrast, do not focus material so efficiently towards the centre and should not produce much observable difference in splitting times between the ridge axis and at distance ( $>$  50 km). Both cases show  $\sim$ 0.5–1 s of splitting away from

392 the axis. Hence their model predicts there should be little observable difference in splitting  
393 in SKS for fast ridges at the axis compared to at distance ( $>50$  km). This agrees with  
394 SKS measurements at the EPR (Wolfe and Solomon, 1998; Harmon et al., 2004), where  
395 full-spreading rates are  $>60$  mm  $a^{-1}$ , and on average  $\sim 150$  mm  $a^{-1}$ .

396 Figure 8 shows  $\delta t$  versus spreading rate for measurements made beneath ridge events,  
397 with filled circles indicating those within 50 km of the ridge axis. The above models would  
398 predict a negative trend in the near-axis data, with larger splitting times observed at the  
399 slowest spreading centres; however, there is no strong trend apparent in our results and the  
400 data suggest a weakly positive correlation if any. (Weighted least-squares linear regression  
401 for results  $<50$  km from axis gives  $R^2 = 0.42$ .)

402 Whilst most authors predict increased splitting at the slowest MORs, some observa-  
403 tions suggest that anisotropy away from the ridge increases with palaeo-spreading rate. P  
404 wave anisotropy in the shallow lithosphere beneath the northwest Atlantic (spreading full-  
405 rate  $\sim 20$  mm  $a^{-1}$ ) is significantly less at  $\sim 3\%$  (Gaherty et al., 2004) than that observed  
406 at present-day fast-spreading sites near the East Pacific Rise ( $\sim 6\%$ , rate  $\sim 100$  mm  $a^{-1}$ )  
407 (Dunn and Toomey, 1997) and old lithosphere in the western Pacific ( $\sim 6\%$ , palaeo-rate  
408  $\sim 60$  mm  $a^{-1}$ ) (Shearer and Orcutt, 1986). Such observations constrain the anisotropy in  
409 the uppermost mantle, hence probably reflect the effect of ‘frozen-in’ olivine LPO and pro-  
410 cesses contemporaneous with lithosphere creation. Gaherty et al. (2004) suggest a spreading  
411 rate dependence could be due to slower ridges accommodating more deformation by brittle  
412 failure in the crust, leading to reduced LPO in the uppermost mantle. However, our shear  
413 wave splitting measurements integrate anisotropy over the complete ray path in the upper  
414 mantle, and it is not clear that this effect could cause the change in  $\delta t$  we observe, given the  
415 thickness of the brittle crust.

416 It is also important to note that this discussion ignores any azimuthal dependence on  
417 splitting parameters. With stations only in North America and Ethiopia, there may be an  
418 azimuthal bias between ridges of different spreading rates, which could account for some of  
419 the variability we observe.

420    *4.3. Splitting in S and SKS at the EPR*

421    Figure 3 indicates that the amount of shear wave splitting in direct S increases away from  
422    the ridge more quickly than that in SKS phases. Assuming that the splitting is accrued over  
423    a horizontal layer of constant anisotropy, this difference cannot be accounted for simply by  
424    the difference in incidence angle between the phases (which would predict a difference in  
425     $\delta t$  of  $\lesssim 0.3$  s for the largest  $\delta t_{\text{SKS}}$ ). Accruing splitting operators through a model of LPO  
426    development at the EPR (Blackman and Kendall, 2002) along the raypaths of the S and  
427    SKS phases also indicates that the contrasting azimuths and incidences of the waves are  
428    insufficient to produce the observed discrepancy.

429    One possible explanation might be along-ridge variability in the strength and style of  
430    mantle anisotropy, leading to varying splitting dependent on the location along ridge seg-  
431    ments. This would imply a sampling bias, whereby events nearest the ridge are in weak  
432    anisotropy regions, whilst those furthest are in strong anisotropy regions. This correlation  
433    arising by chance or through some earthquake mechanism seems unlikely. More likely may  
434    be the influence of azimuth and large-scale heterogeneity. SKS waves travelling to the MELT  
435    and GLIMPSE OBSs are along a backazimuth of  $\sim 280^\circ$ , approximately ridge-perpendicular,  
436    whilst S waves in this study travel along an azimuth of  $\sim 20^\circ$ , closely parallel to the ridge.  
437    Hence the S waves may be more sensitive to ridge structure for events near the ridge, and  
438    less so at distance, magnifying the contrast in structure between the on- and off-axis mantle.

439    Another simple explanation may be that current LPO models do not include effects  
440    of short-wavelength segregation of material such as that observed in experiments which  
441    deform partially molten olivine aggregates (Holtzman et al., 2003a,b), and in numerical  
442    experiments incorporating porosity and strain rate-dependent viscosity (Katz et al., 2006).  
443    These observations predict lenses of melt at MORs will be aligned approximately with long  
444    axes of the strain ellipses expected for corner flow, forming bands which dip away from  
445    the ridge axis, becoming approximately horizontal beyond about 50 km from the centre.  
446    The alignment of seismically distinct material on scales shorter than the seismic wavelength  
447    would lead to a shape-preferred orientation (SPO). Such an SPO with inclusions of uniaxial  
448    symmetry would lead to transverse isotropy (TI), where the seismic velocities and amount

of splitting vary only away from the axis of rotational symmetry (e.g., Hudson, 1980). This would mean SKS waves travelling perpendicular to the axis (such as those measured by the MELT and GLIMPSE OBSs) would not be split due being polarised in the sagittal plane, whilst near the ridge fast orientations for S waves would be between ridge-parallel and ridge-perpendicular, depending on the dip of the fabric. Further from the ridge, where bands are horizontal, SKS would be unsplit in any azimuth; S waves travelling parallel to the ridge would have ridge-perpendicular fast orientations. This mechanism would increase delay times in S over SKS away from the axis, as we observe at the EPR. Aligned inclusions leading to SPO could also explain surface wave observations that  $V_{SH} > V_{SV}$  beneath the Pacific (e.g., Ekström and Dziewoński, 1998; Nettles and Dziewoński, 2008), and receiver function observations of a suboceanic seismic discontinuity at depths of 50 to 150 km (the so-called ‘LAB’ discontinuity; Rychert and Shearer, 2009; Kawakatsu et al., 2009; Rychert et al., 2010; Kumar and Kawakatsu, 2011).

Several studies show that the structure beneath the EPR is asymmetric (e.g., Conder, 2007; Harmon et al., 2004; Podolefsky et al., 2004; Wolfe and Solomon, 1998), hence this might play some part in the observed difference in splitting times between S and SKS phases. However, our data also sample both sides of the ridge, so presumably are also affected by the same asymmetry, yet still consistently show larger splitting in S than SKS. Blackman and Kendall (1997) calculate the splitting times for vertical-incidence shear waves on a suite of asymmetric model of MOR LPO development, as for the EPR, and for the best-fitting case predict splitting times for SKS of up to 2–3 s on the Pacific plate, and up to 1.5 s on the Nazca plate. This asymmetry is observed to a lesser extent in the data (Figure 3). However, again the effect of azimuth is not tested and the path of the S waves (ridge-parallel) may negate some of the effects on  $\delta t$  expected for an asymmetric spreading centre.

#### 4.4. Plate motion and mantle flow

Several authors interpret shear wave splitting measurements in terms of the alignment of olivine due to shearing of the asthenosphere by the relative motion of the lithosphere above it (e.g., Tommasi, 1998; Conrad et al., 2007). In this study, it is hard to discriminate

477 between this and spreading processes for the EPR, MAR and SWIR because the directions of  
478 spreading and APM are very similar (Figures 2, 4 and 6). APM is also small ( $<10 \text{ mm a}^{-1}$ )  
479 at the SWIR. Beneath our measurements on the SEIR, the APM is eastwards, but no fast  
480 orientations are parallel to this. At the eastern Gakkel ridge, plate motion is approximately  
481 parallel to the ridge, and is faster than the spreading rate ( $|APM| \approx 12 \text{ mm a}^{-1}$ ). It may be,  
482 therefore, that the ridge-parallel  $\phi''$  reflects the shearing of the North American and Eurasian  
483 plates over the asthenosphere. Another interpretation would be that because spreading rate  
484 varies along the ridge, mantle material again flows parallel to the axis, again leading to  
485 olivine LPO with [100] directions which are compatible with our observations. In both  
486 cases, however, the mechanisms which cause this at spreading centres themselves is still  
487 unclear. Conrad et al. (2007) point out that parallelism between APM direction and olivine  
488 a-axes—assuming LPO in olivine to be the cause of the observed anisotropy—may not be  
489 a good assumption beneath MORs, because of the complicated combination of radial and  
490 horizontal flow present there. One might also expect along-axis flow of material due to the  
491 fact that spreading rate varies along the ridge, which might also explain ridge-parallel fast  
492 orientations.

493 A more involved explanation of upper mantle shear wave splitting results can be invoked  
494 by considering more complex flow regimes. Behn et al. (2004), for instance, combine plate  
495 motion models with a model of mantle flow derived from seismic tomography, and compare  
496 SKS splitting observations with the fast orientations predicted by the flow models. In  
497 this case, the authors conclude that plate spreading directions adequately describe SKS  
498 fast orientations on oceanic plates within 500 km of the ridge. In this study, only one  
499 measurement is further than this from the spreading centre (westernmost stack in Figure  
500 2), and here the fossil spreading direction and APM are the same within a few degrees.  
501 This might serve to reinforce the fossil anisotropy in the lithosphere, but  $\delta t$  here is modest  
502 ( $(1.7 \pm 0.5) \text{ s}$ ).

503 Where inferred mantle flow beneath Africa might affect our measurements, on the SWIR,  
504 we see several fast orientations parallel to the spreading direction, however one stacked  
505 measurement (Figure 6) at  $\sim 32^\circ\text{E}$  is perpendicular to spreading. This direction does not

506 correlate with APM. Whilst more rigorous modelling of deeper flow could be attempted (e.g.,  
507 Forte et al., 2010), this is beyond the scope of this study, and may be more appropriate when  
508 further data are presented.

## 509 **5. Conclusions**

510 We present measurements made using the source-side shear wave splitting technique of  
511 upper mantle anisotropy beneath mid-ocean ridges around the world. We correct for the  
512 UM on the receiver side for seismic stations where the anisotropy beneath is very well char-  
513 acterised, and can resolve the source anisotropy, subject to a series of rigorous tests. With  
514 122 new observations, the presented dataset adds significantly to the current knowledge of  
515 anisotropy beneath MORs. There is no strong trend that corroborates the prediction of more  
516 splitting beneath slow-spreading ridges, and it may be true that more splitting is present at  
517 fast-spreading ones. For the EPR, comparisons with previous SKS splitting measurements  
518 show more splitting in S away from the ridge. We suggest that TI dipping away from the  
519 ridge axis, becoming horizontal at distance, is compatible with our observations, in addition  
520 to LPO development. This would be consistent with other observations of anisotropy and  
521 a seismic discontinuity beneath the oceans. We find anisotropy at MORs appears to be  
522 dominated by ridge processes, rather than plate motion over the asthenosphere. As further  
523 rigorous study of UM anisotropy using SKS phases becomes routine, more stations can be  
524 used to measure the seismic shear wave splitting beneath MORs and other remote parts of  
525 the Earth where earthquakes occur, and hence our understanding of mantle dynamics in  
526 these regions will be vastly improved.

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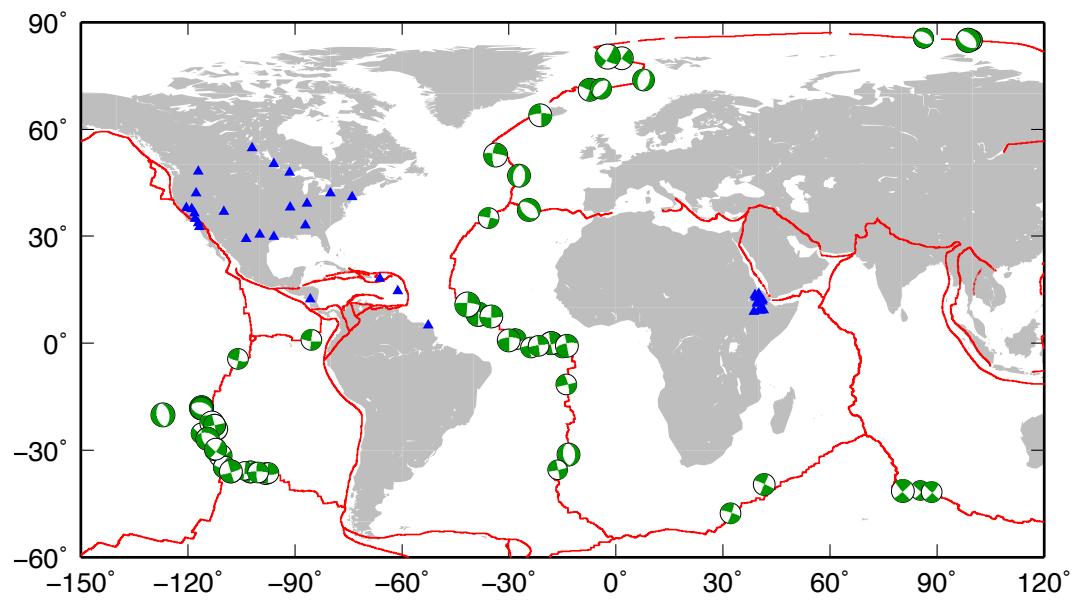


Figure 1: Location of events used in this study with plate boundaries of Bird (2003). Lower-hemisphere focal mechanisms are the best-fitting double-couple solutions as given by the Global CMT project. Blue triangles are seismic stations. Magnitude range ( $4.4 \leq M_b \leq 6.7$ ) shown by size of hemispheres.

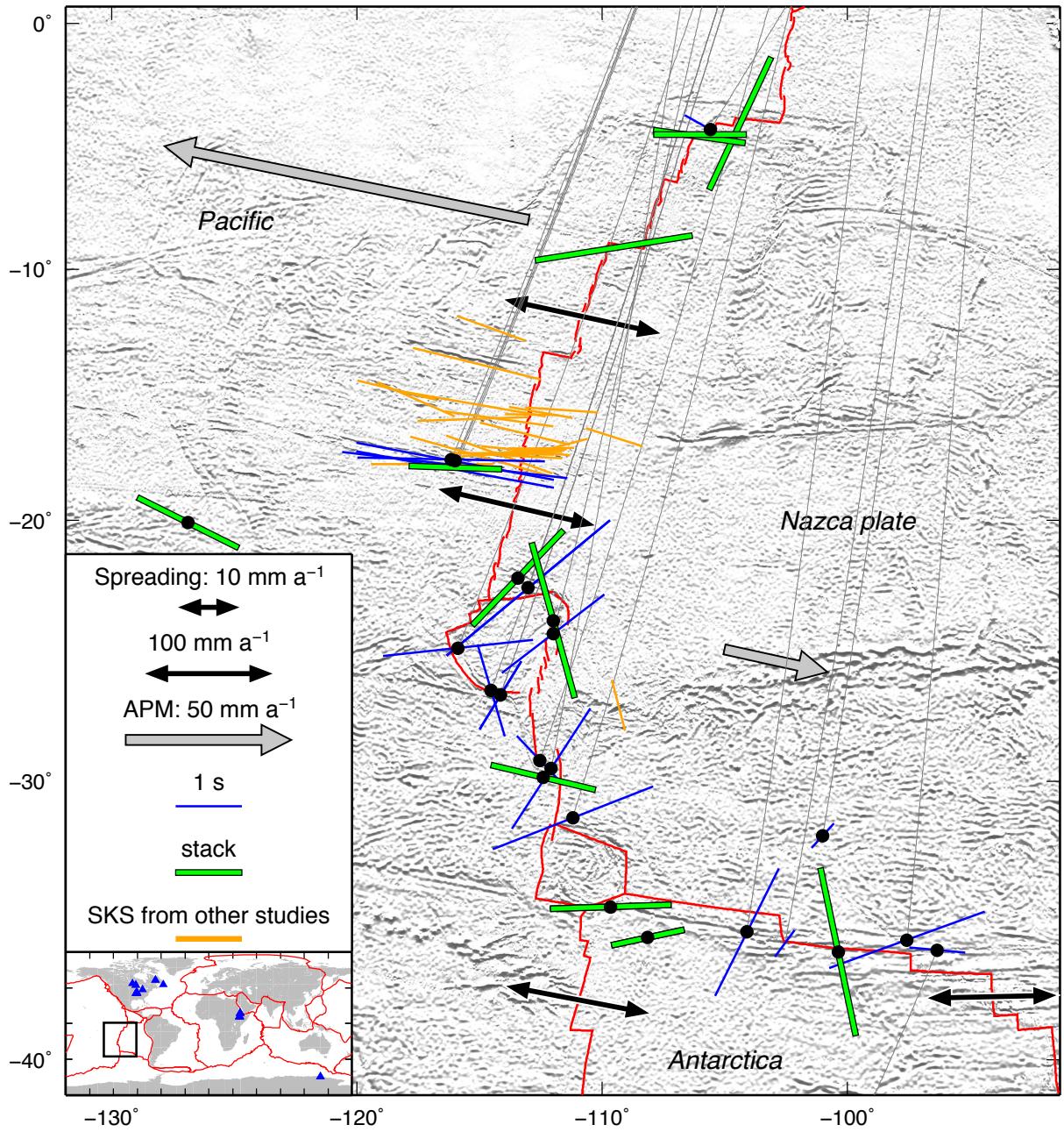


Figure 2: Source-side splitting beneath events on the EPR. Dots show earthquake locations, with bars indicating splitting parameters, where the orientation shows  $\phi''$  and the length  $\delta t$ , as in the legend. Blue bars are for single measurements; green for stacks. Orange bars show SKS splitting parameters from previous studies (Wolfe and Solomon, 1998; Harmon et al., 2004). Thin grey lines show raypaths to stations (blue triangles, inset map). Shading indicates bathymetry (Smith and Sandwell, 1997), and thin red lines are plate boundaries. Black double-headed arrows show base-10 logarithm of NUVEL-1A full spreading rates at selected locations along the ridge. Grey arrows show the absolute plate motion (APM) in the HS3 reference frame of the NUVEL-1A model (Gripp and Gordon, 2002). The legend indicates spreading and APM rates. Results include stacks from Nowacki et al. (2010).

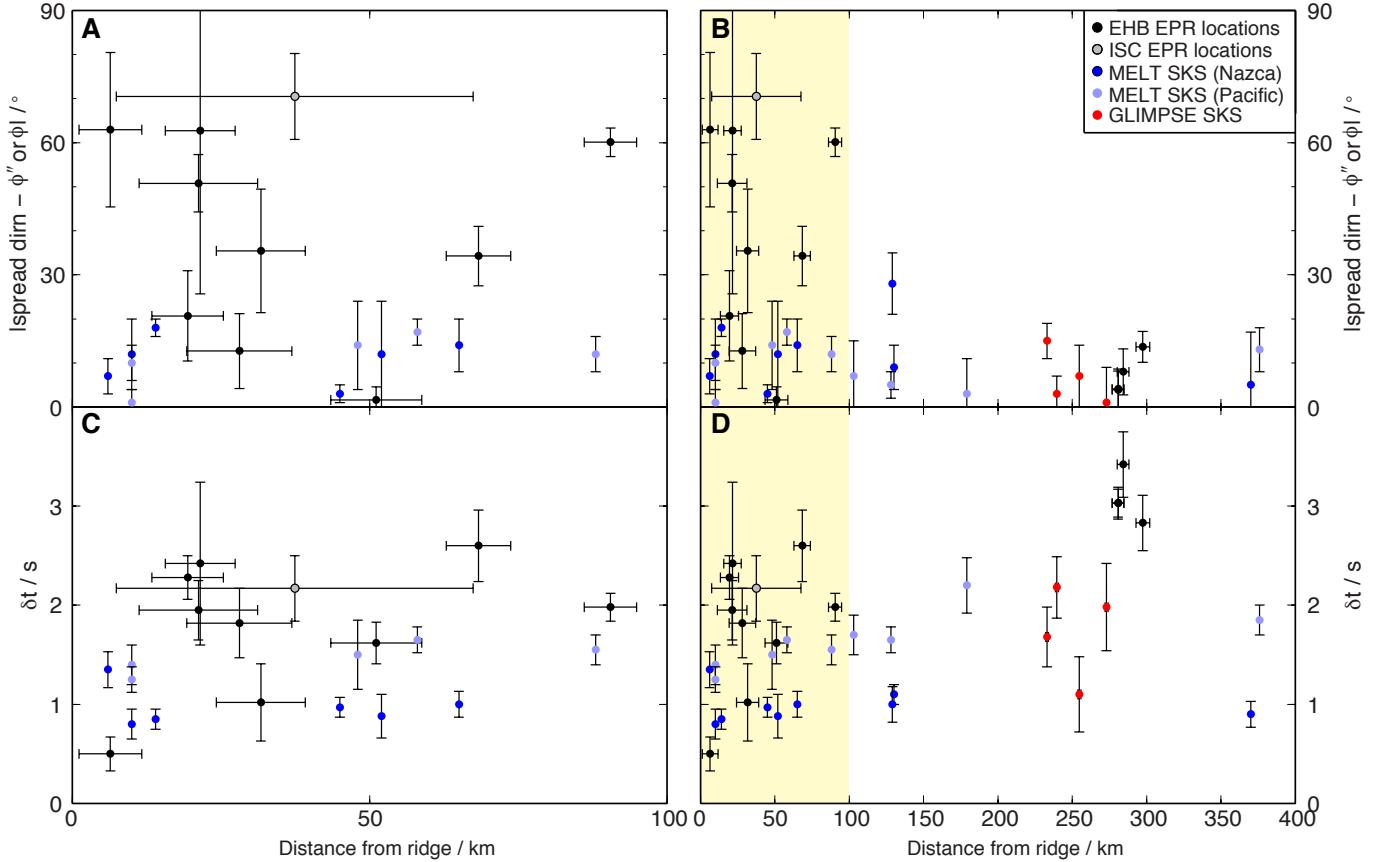


Figure 3: Variation of splitting parameters beneath ‘ridge’ events with distance away from the axial ridge at the EPR. Black and grey circles indicate respectively the EHB and standard ISC locations of the events in this study. Error bars show 95% confidence interval in splitting parameters and stated uncertainty in event locations. Coloured circles indicate MELT (blue) and GLIMPSE (red) SKS splitting parameters as shown in the legend, where MELT stations on the Pacific and Nazca plates are coloured lighter and darker respectively. All GLIMPSE stations are on the Pacific plate. Shaded part of panels on right shows regions shown by panels on left. A) and B) Modulus of difference in angle between  $\phi''$  or  $\phi_{SKS}$  and the plate spreading direction (DeMets et al., 1994). C) and D) Splitting times for S or SKS.

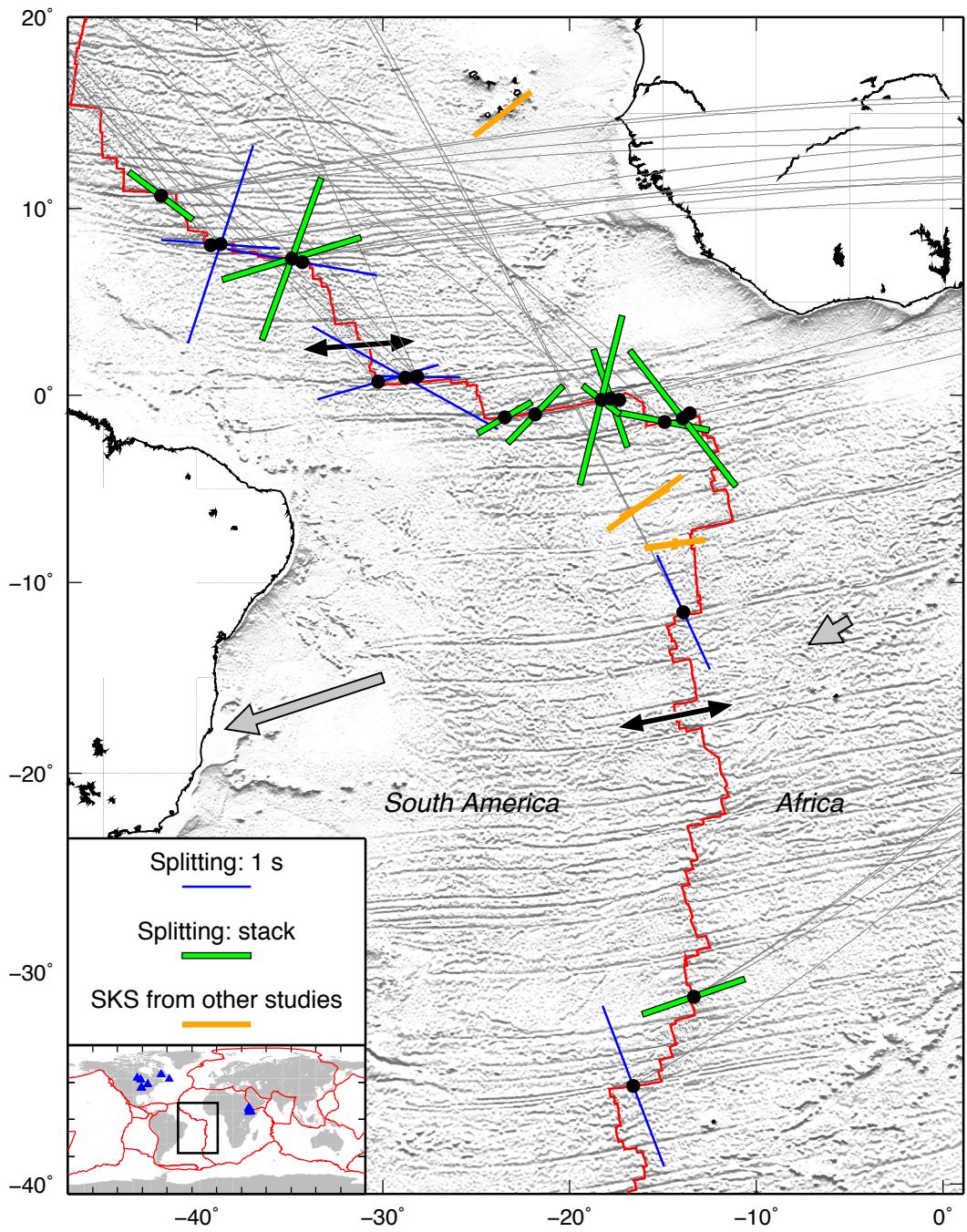


Figure 4: Splitting parameters beneath events on the MAR. Symbols as for Figure 2. Orange bars show SKS results of Wolfe and Silver (1998) and Behn et al. (2004). Note that some events are measured at stations in both North America and Ethiopia, in which case stacks of results for both directions are shown. Spreading rates and APM are of same scale as Figure 2.

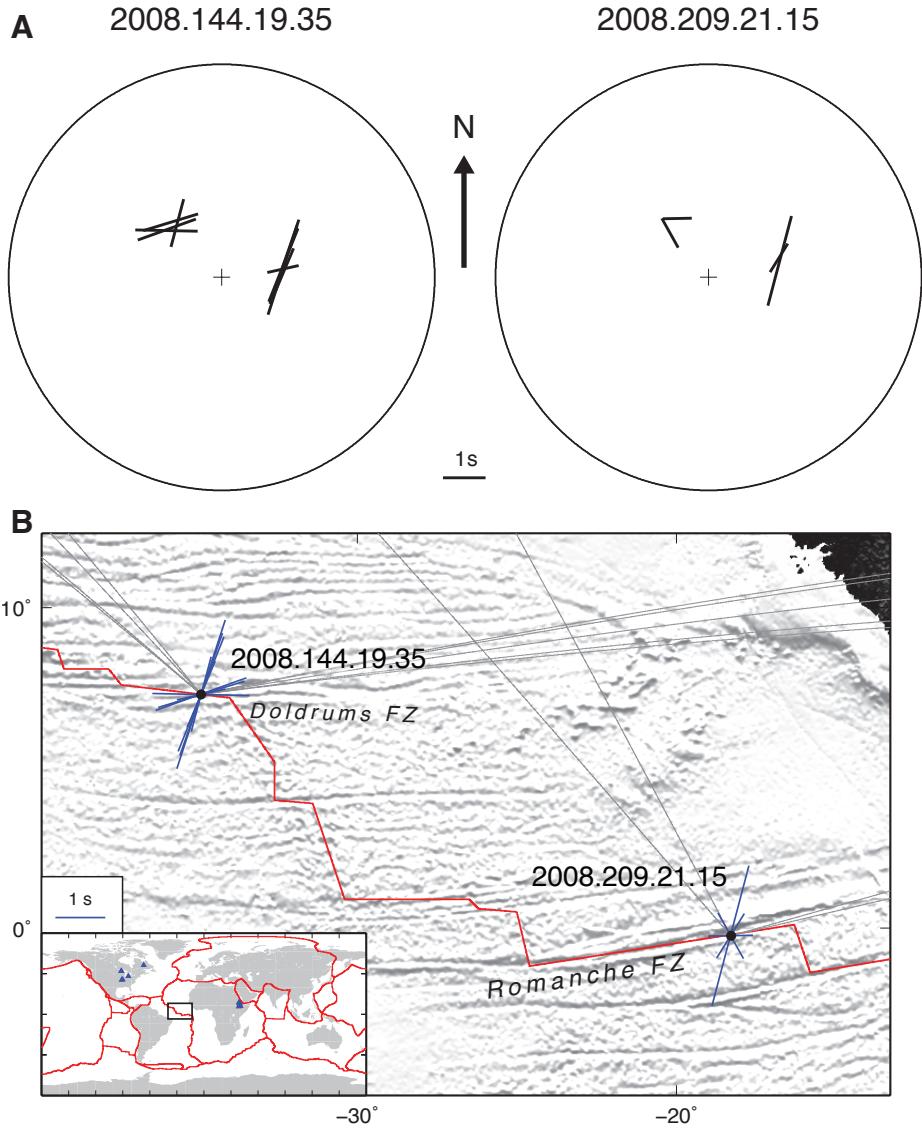


Figure 5: Lower-hemisphere diagrams for splitting parameters measured beneath two events on the MAR. A) Azimuthal and inclination-dependence of splitting parameters shown on equal-area lower-hemisphere projections. Average inclination of downgoing rays in top 150 km of IASP91 (Kennett and Engdahl, 1991) is shown by radial distance (with vertical at the centre). Azimuth corresponds to azimuth at the event. Bar orientation and length corresponds to  $\phi''$  and  $\delta t$  respectively, as per the scale, centre. The splitting times measured at Ethiopian stations (group on right of hemispheres) are larger for both events, and  $\phi''$  is also different. B) Location of events and individual splitting measurements shown at earthquake location. Bars correspond to splitting parameters as for previous figures, with delay time indicated by length as per the legend (left). Inset map show location of larger map by thick black box. The raypaths to Ethiopia run along the transform zones, whilst those to North America move away from the transforms.

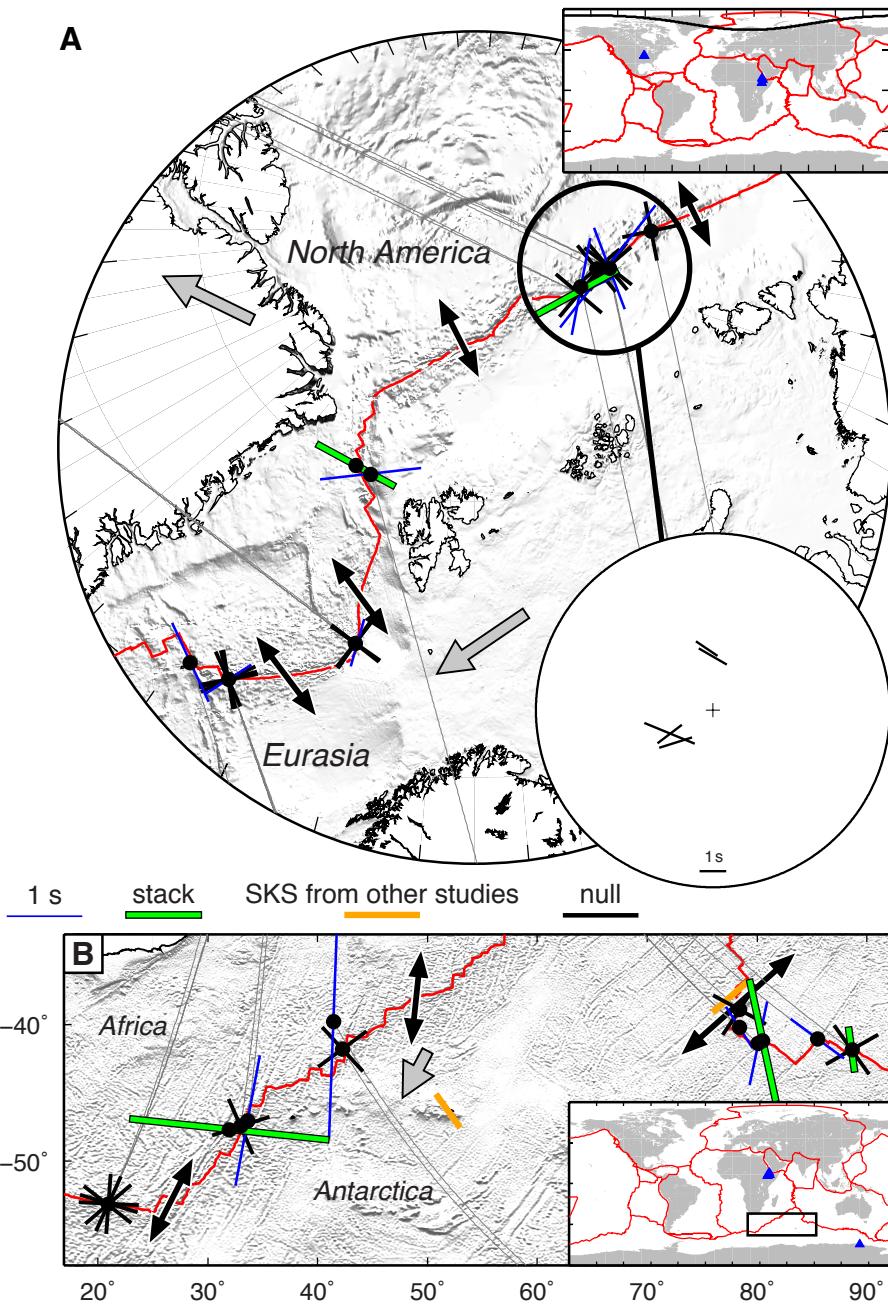


Figure 6: Splitting parameters beneath events on the Gakkel, Southeast and Southwest Indian Ridges. Symbols as for Figure 2, except null results are shown by black crosses with bars parallel to the null directions. A) Results for events on Gakkel Ridge. Thick black circle shows results included in lower hemisphere stereoplot, inset lower right. Bars above the centre show measurements made at North American stations; those to the lower left show measurements at Ethiopian stations. Scale indicated at bottom. B) Results for events on Southwest and Southeast Indian Ridges. SKS splitting at CRZF (Base Alfred Faure, Crozet Islands) and AIS (île Nouvelle-Amsterdam; Behn et al., 2004) is shown by the orange bar. APM is less than  $10 \text{ mm a}^{-1}$  for African plate, and parallel to spreading direction at  $\sim 65 \text{ mm a}^{-1}$  for Australian plate (northeast corner, not labelled).

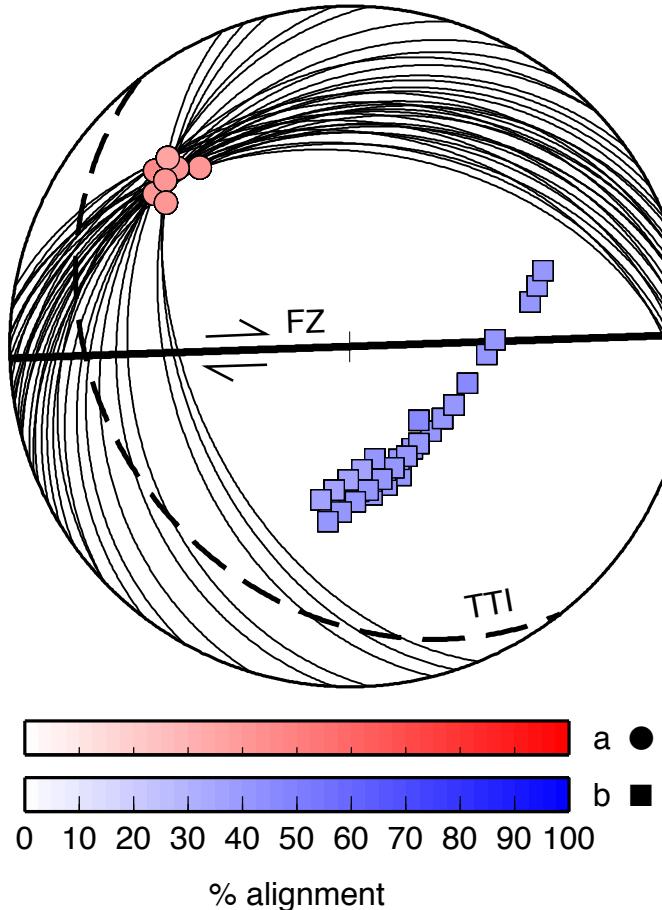


Figure 7: Orientations of olivine a- and b-axes compatible with observations from event 2008.144.19.35 on the Doldrums FZ, MAR. Lower-hemisphere equal area plot shows north upward and the vertical direction out of the page. Red circles (a-axes) and blue squares (b-axes) are shaded per the degree of alignment according to the scale below. Thick black solid line shows approximate strike of FZ and spreading direction, with strike-slip arrows indicating sense of shear. Thick dashed line is best-fitting plane of isotropy from fit of TTI to fast orientations. Thin solid lines are crystallographic slip planes (b-planes) for the case of ‘A-type’ olivine LPO.

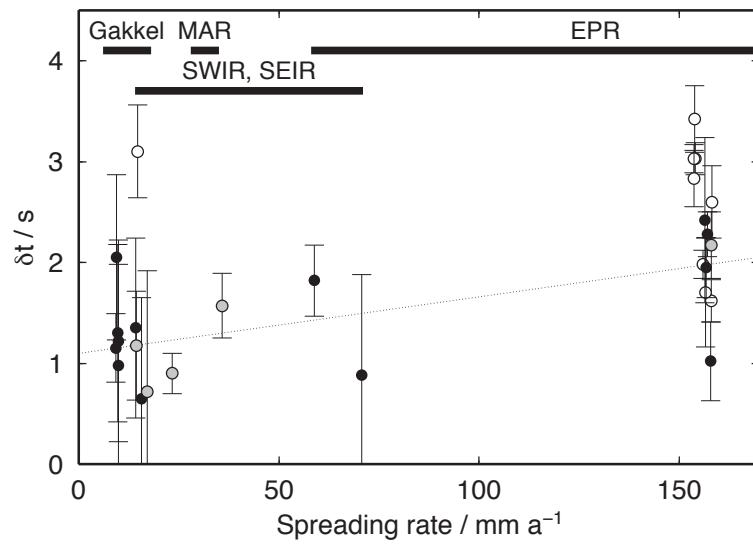


Figure 8: Splitting time versus spreading full-rate beneath all ‘ridge’ events. Filled circles show results less than 50 km from the ridge axis: black circles shows events with EHB locations; grey circles indicate ISC locations. Open circles indicate events >50 km from the axis (all EHB locations). Thick bars show range of spreading rates represented by events beneath each MOR in this study. Weighted linear fit to near-axis (filled circles) data is shown with thin dotted line.