# THE GPE CODE

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Date: 2010-06-28 12:33:51 +0100 (Mon, 28 Jun 2010)

Revision: 648

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# Chapter 1

# Introduction

The GPE code is a modular 3D Gross-Pitaevskii equation solver, written in Fortran 90 and parallelised using the Message Passing Interface (MPI).

The code can solve a range of problems for both homogeneous and non-homogeneous (trapped) condensates, including vortex dynamics (lines, rings), and non-equilibrium dynamics (condensate formation).

Time stepping is performed explicitly using one of the following methods:

- first-order Euler (E1);
- second-order Runge-Kutta (RK2);
- fourth-order Runge–Kutta (RK4);
- fourth/fifth-order adaptive Runge-Kutta-Fehlberg (RK45).

The spatial discretisation is performed with either second- or fourth-order accurate centred finite differences, and the boundary conditions can be either periodic or reflective.

The "kind" of floating point variables is parametrised, so real or double-precision arithmetic can be specified without relying on compiler switches.

Some publications that have used the code include Al-Amri et al. (2008) and White et al. (2010).

# 1.1 Preliminaries and required software

Since the code is parallelised using MPI, an MPI parallel environment must be available (currently, this is also true even when running on one processor). The code has been tested with the MPICH and OpenMPI implementations of MPI, but other implementations which adhere to the MPI standard should also be useable.

The code uses some Fortran 2003 features, including allocatable arrays within user-defined types, the protected keyword, and stream I/O, so a compiler which allows some

Fortran 2003 constructs is required. Compilers known to work include sunf95 (Sun Studio 12), ifort (Intel Fortran Compiler), and gfortran (GNU Fortran Compiler).

The code also requires the FFTw library (http://www.fftw.org) for some routines. This must be version 2 of the library (the latest is version 2.1.5), since FFTs in the code are performed in parallel, and FFTw version 3 does not yet support (distributed) parallel transforms.

To view the output of the code, any graphics program capable of reading space-separated, columnar text files will be able to produce time-series plots (gnuplot is a nice, easy to use program, http://www.gnuplot.info/). Routines written in IDL (Interactive Data Language, http://www.ittvis.com/ProductServices/IDL.aspx) for producing 2D contour and 3D isosurface plots are included with the code. The data for these plots are saved in binary format, so more work will be needed if a different graphics program is to be used. Volume renderings using VAPOR (http://www.vapor.ucar.edu/) are also possible.

### 1.1.1 Summary of required hardware and software

The following outlines the required hardware and software needed to run the code. Other hardware and software may work, but has not been tested.

- A Unix-like operating system, running on x86 or x86\_64 hardware.
- A standard development environment including make.
- A Fortran 90 compiler, supporting Fortran 2003 constructs (see above).
- An implementation of MPI, e.g. OpenMPI or MPICH.
- Version 2.x.y of the FFTw library.
- gnuplot and/or IDL for visualisation; VAPOR optional.

#### 1.1.2 IDL

If you intend to use IDL to produce 2D contour and 3D isosurface plots, then you will need to make sure that some environment variables are set in your startup files, e.g. .cshrc if you are using C Shell, or .profile if you are using Bash. These variables are:

- IDL\_DIR specifies the IDL install directory;
- IDL\_PATH specifies the directories in which IDL will look to find IDL procedures and functions;
- IDL\_STARTUP specifies where to find the IDL startup file .idlrc.

Two example files in the idl subdirectory of the main code directory, named idl.csh and idl.sh (for C Shell and Bash respectively), provide example commands to set these environment variables. They will likely need to be edited to suit your system, then added to your startup file.

An example .idlrc file is also provided (named dot.idlrc), which fixes some quirks with X displays, and sets 24-bit true colour output by default. If you wish to use this, then make sure that the IDL\_STARTUP environment variable points to this file.

### IDL command line

You might find that IDL's command line is broken, i.e. no command-line completion, no command-line history, inability to use the End or Home keys, etc. To remedy this, install the rlwrap package, and set up an alias alias -a -c idl for the idl command. This will provide readline functionality for IDL.

# 1.2 Organisation of this manual

The rest of this manual is organised as follows:

- Chapter 2 is a getting started guide. The chapter will guide you through the basics of setting up, compiling, and running the code, as well as visualising some of the results.
- Chapter 3 describes the governing equations, and the non-dimensionalisation used.
- Chapter 4 describes some of the numerical formulation, including time stepping schemes, spatial discretisation, and boundary conditions.
- Chapter 5 describes the files which make up the code, including source files, input and output files, and the IDL programs used for visualisation.

# Chapter 2

# Getting started

This chapter is intended as a quick start guide to help you compile, run, and view the results of the GPE code. Later chapters explain various aspects of the code in greater detail, and some amount of reading the (commented) source code should be expected, to gain a better understanding of what the code can do. Here, we shall outline the basics of going through a run cycle, by using the provided *ring* example, which propagates a vortex ring in a homogeneous condensate for a short time.

# 2.1 Setting the run parameters and initial condition

Change to the examples/ring directory. In here, you will find symbolic links to the source files, and in addition, three .in files. In general, it is these files — parameters.in, ic.in, and run.in — which you will need to edit, in order to set up a run.

# 2.1.1 parameters.in

For this example, simply change nyprocs and nzprocs, so that nyprocs\*nzprocs is equal to (or less than) the number of processors on your machine. If memory is an issue, then you may want to reduce any or all of nx, ny, or nz (and possibly xr, yr, and zr below). See §5.2.1 for a full description of this file.

### 2.1.2 ic.in

This file defines the initial condition. For this example, no changes are necessary. See §5.2.2 for a full description of this file.

#### 2.1.3 run.in

This file defines the main parameters for the run, as a set of Fortran namelists.

For the *ring* example, you might want to alter end\_time if the run takes too long. If you altered any of nx, ny, or nz in parameters.in, then you might also want to alter xr,

yr, or zr as appropriate. These parameters set the right-hand end of the physical extent of the computational box (the left-hand end is set to the same value with opposite sign). See §5.2.3 for a full description of this file.

# 2.2 Compiling the code

Choose a directory where your run will take place (we shall assume the directory run\_dir in what follows). This directory should not be on a network file system (NFS), which will likely be too slow, and possibly under quota restrictions (as might be the case for a network mounted /home area, for example). The *ring* example will require approximately 3GB of disk space with the default parameters.

Now compile the code with

```
./setup run_dir
```

This will compile the code<sup>1</sup>, create the directory run\_dir, and copy parameters.in, ic.in, run.in, and run.sh to the run directory. In addition, the executable gpe is moved to the run directory.

This run uses double-precision floating-point arithmetic. For other runs, if you require only single precision, then remember to set pr appropriately in parameters.in, and compile with

```
./setup run_dir single
```

Now change to the run directory, ready to run the code.

# 2.3 Running the code

To run the code simply type

```
./run.sh <nprocs>
```

where <nprocs> specifies on how many processes the code should run. Ideally, this should be less than or equal to the number of physical processors in your system. It should match nyprocs\*nzprocs in parameters.in, otherwise an error will result. The job is run in the background using nohup, so control immediately returns to the terminal, and logging out of the machine on which the job is running will not terminate the job.

The run.sh script has several options for controlling how the run should be started. Run it with no arguments to see usage instructions.

 $<sup>^1</sup>$ The default Makefile assumes the sunf95 compiler is available and in the PATH, and that 64-bit code should be generated. If this is not what you want, then you will need to edit the Makefile. See  $\S 5.3$  for more information.

# 2.4 During the run

As a first sanity check that the job is running, look for a file called ERROR in the run directory. If this does not exist, then that is a good sign. Also have a look at the file log.txt; this should say something like

```
Explicit fifth order Runge-Kutta-Fehlberg adaptive time stepping Homogeneous condensate, natural units non-dim.
-2i*dpsi/dt + 2iU*dpsi/dx = del^2(psi) + (1-|psi|^2)psi
```

and is something to check to make sure the correct time stepping scheme is being used, and that the correct form of the GPE is being solved.

The command top should show the gpe executable listed the same number of times as the number of processes on which you chose to run the job.

Finally, if numbers are appearing in the \*.dat files, then you can be sure that the job is running.

# 2.5 Important run-time files

This section describes some important files which may exist in the run directory, as a job is running.

### 2.5.1 The RUNNING file

When the job is running, there is an empty file called RUNNING in the run directory; when the job is finished this file is deleted. You can also delete this file at any time to immediately stop the run cleanly. The job should never be ended by using kill or killall, unless an unrecoverable error occurs.

#### 2.5.2 The ERROR file

If a serious error occurs, which the code can anticipate and handle, then the ERROR file is created in the run directory. Its contents will provide details of the error. The job will be terminated cleanly in this case, and the RUNNING file will be removed.

# 2.5.3 The log.txt file

Any output that the job would normally print to the screen, is instead redirected to the file log.txt. It is a good idea to periodically check this file to make sure that no unexpected run-time errors have occurred. Any errors which are written to ERROR are also written to this file. Any errors which the code could not anticipate, such as MPI communication errors, will appear here. In this case the job will most likely not terminate cleanly, and the RUNNING file might not be removed.

### 2.5.4 The SAVE file

As the job is running, data are periodically saved, according to the save\_rate parameters in run.in. Every time step, the code checks for the existence of a file called SAVE in the run directory. If this file exists, then 3D data (from which isosurfaces and contour plots can be produced) is immediately written to disk. The SAVE file is then automatically removed. This file can be created, for example, by using the touch command, i.e. touch SAVE.

# 2.6 Program output

Once a job has started, the code produces various output files containing various data. Exactly which files appear depends on which parameters have been set in the io\_params namelist in run.in.

The numbered **proc** directories contain data local to each process on which the job is running. This is where 3D isosurface data are stored. A full description of the output files is given in §5.4.

# 2.7 Viewing results

During and after a run, it is possible to plot various 1D graphs, 2D contour plots and 3D isosurfaces. For the 1D time-series graphs, gnuplot is a good program to use, although nearly any plotting program capable of understanding simple columnar text files will be sufficient. This section will describe how to produce some plots.

# 2.7.1 1D time-series plots

1D time-series data is produced in the \*.dat files at the top-level of the run directory, much of which is self-explanatory, given the names of the files, e.g. energy.dat, mass.dat. A full description of what can be found in these data files is given in §5.4, but note that the exact contents of files can be subject to change.

So, for the *ring* example, using gnuplot to plot the total energy in the system, you could do

```
p "energy.dat" u 1:3 w lp
```

which will plot column three (energy) versus column one (real time) in the file, and where the plot, using, with, and linespoints keywords have been abbreviated to the shortest non-ambiguous form, which is possible with any gnuplot command.

As it turns out, there are not many interesting 1D graphs for the ring example!

### 2.7.2 Contour plots

Contour and isosurface data are stored in binary form within the **proc** directories; the layout of the binary files is described in table 5.8 in §5.4.1.

If you have access to the Interactive Data Language (IDL) graphics program, then IDL routines are provided with the code to directly plot the data. These are located in the idl subdirectory of the main code directory. The main program is called gpe.pro, and is described in more detail in §5.6. The subsequent sections will guide you through the basics of using the program.

From the run directory, start IDL with the command idl. Providing the IDL environment is set up correctly (see §1.1.2) you should be able to type

which will show a contour plot of the density of the ring initial condition as a slice through the (x, y)-plane at z = 0. You should see something which looks like figure 2.1.

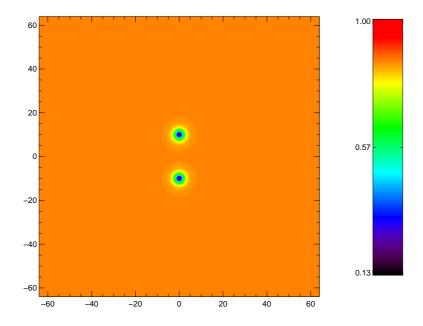


Figure 2.1: Contour plot of the density  $|\psi|^2$  of the condensate, showing a vortex ring of radius  $R_0 = 10$  in the (x, y)-plane at z = 0.

The word gpe represents the IDL program, which takes two non-optional arguments. The first is the index corresponding to the first file you want to plot, and the second is the index corresponding to the last file you want to plot. For output to the screen, these numbers should always be the same.

The numbers refer to the files within the proc directories. If you look in proc00, for example, you will see a number of files named dens\*\*\*\*\*.dat. The arguments to gpe are the numbers within the filenames, excluding the leading zeroes.

So for example, arguments of 0 and 0 would plot dens0000000.dat; arguments of 334 and 334 would plot dens0000334.dat, etc.

The remaining arguments are optional. The first, /db1, denotes that the binary data are double-precision. You must always provide this keyword when you have performed a double-precision run. In a single-precision run, this keyword is not necessary.

The second, /cntr, is a keyword which turns on contour output, rather than the default, which is an isosurface (see §2.7.3).

### 2.7.3 Isosurface plots

Now do exactly as in §2.7.2, but this time leave off the /cntr keyword.

You should see an isosurface plot in a window similar to that in figure 2.2, and a control panel with various buttons on it. Most of the control panel buttons can be ignored. The bbox, content, and axis buttons turn on or off the bounding box, the content, and the axes labels respectively. You will need to click the Redraw button if you make any changes. Left-click and hold in the white window and you can move the view around; right-clicking

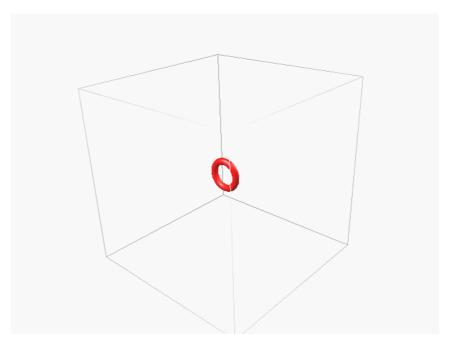


Figure 2.2: Isosurface plot of the density of the condensate at the level  $|\psi|^2 = 0.75$ , showing a vortex ring of radius  $R_0 = 10$ .

zooms in and out, and middle-clicking moves the centre viewpoint. Depending on the speed of your machine, and the complexity of the displayed image, zooming and moving the view might be a little slow, so move the mouse slowly, and avoid large, jerky movements.

The coloured bar in the control panel is a histogram of the density, and clicking in here will redraw the isosurface at the new density level, which is shown toward the right-hand side of the control panel.

Pressing auto will automatically redraw the isosurface when other boxes are checked, for example, the axes, or whether the surface is solid, wireframe, or a series of points. If the display is too slow, selecting points or wireframe instead of solid will speed it up.

#### The TMat

The TMat, or transformation matrix, controls the isosurface view, and is set in gpe.pro to give a perspective view by default. If you would like to change the default, find the view you would like by rotating and zooming the display, then click the TMat button on the control panel. This will print the transformation matrix for the view in the terminal. Copy and paste this into gpe.pro, overwriting the TMat that is currently defined.

### 2.7.4 Animations

By saving a series of snapshots of 2D or 3D data, it is possible to then combine them into an animation.

From the run directory, run the script create-links. This will create a directory links, which contains renamed symbolic links (shortcuts) pointing to the isosurface files within each process directory. This is to make sure that the files are numbered sequentially in steps of one. The create-links script is located in the scripts subdirectory of the main code directory. Giving it the argument --help will provide usage instructions.

#### Contour animations

Change to the links directory and create another directory called images, then start IDL again.

Now type

This will loop over all data files numbered 0 to 9 and save con\_dens\*\*\*\*\*.png files in the images directory. (If you left the *ring* example parameters at their defaults, then you should actually find 100 data files in each proc directory, so you could do plots from 0 to 99 if you wish.)

If you forget the /c\_anim keyword here, you will get all the plots shown to the screen. When you are plotting 100 or more files, this can be very bad and in some cases might run your machine out of memory.

Now change to the images directory and run the makemovie script, which is located in the scripts subdirectory of the main code directory:

```
makemovie -i png -p con_dens
```

which will create an AVI animation out of the PNG files, with 10 frames, saving the output as output.avi. The makemovie script has several options. Run the script with no arguments to see its help.

You can then play the animation with any media player, for example,

```
mplayer output.avi
```

The animation is very short; if it works you can try a longer animation by repeating the gpe IDL command with arguments 0 and 99.

#### Isosurface animations

Rename the images directory to something else, and recreate an empty images directory. From within IDL type

As for the contour plots before, this will save PNG files in the images directory, which you can then convert into an animation using makemovie.

If you forget the /png keyword, all the output will again go to the screen. This should be avoided if possible.

# 2.8 Restarting the code

After a run, you might decide that you would like to restart the code from where you left off, for example, maybe you started your run in imaginary time, and now want to switch to real time, or maybe you did not run for quite long enough.

If this is the case, create a new directory where the restarted run will take place, and copy parameters.in, ic.in, run.in, run.sh, and the executable gpe from the initial run directory to the restart directory.

Now edit run.in, change restart to .true., and make any other changes you feel necessary. Recompilation is not needed when making changes only to run.in.

Now type

./run.sh -r <initial run directory> <nprocs>

where <initial run directory> is the directory of the run which you want to continue. This will copy the end\_state.dat files from the proc directories to the restart directory (suitably renamed), and start the run. The log.txt file should then report that the code is

Getting restart conditions

and will restart from where it left off.

# Chapter 3

# Governing equations and non-dimensionalisation

The single-particle complex wavefunction  $\psi(\mathbf{r},t)$  for N bosons of mass m, obeys the 3-dimensional, time dependent, Gross-Pitaevskii (GP) equation (Gross, 1961; Pitaevskii, 1961)

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi + V_{\text{ext}} \psi + g|\psi|^2 \psi - \mu \psi, \tag{3.1}$$

where  $\hbar = h/(2\pi)$  is the reduced Planck constant,  $V_{\rm ext}$  is an external trapping potential, g is the strength of the interactions between the bosons, and  $\mu$  is the chemical potential. The wavefunction is normalised by the condition that

$$\int_{V} |\psi|^2 \mathrm{d}V = N.$$

The GPE describes the evolution of the ground state of a quantum system of weakly interacting bosons, in the limit of zero temperature, and when the number of bosons N is large.

# 3.1 Imaginary time

In making the transformation  $t \to -it$ , equation (3.1) becomes

$$-\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi + V_{\text{ext}} \psi + g|\psi|^2 \psi - \mu \psi,$$

which can be thought of as a modified diffusion equation. For certain initial conditions, it is sometimes necessary to run the code in imaginary time, because the initial conditions are not exact solutions of the GPE. By propagating in imaginary time, these solutions tend to the ground state, which is an exact solution of the GPE. The code can then be propagated in real time.

To propagate in imaginary time set real\_time equal to .true.; otherwise set real\_time equal to .false..

### 3.2 Non-dimensionalised GPE

The code solves a non-dimensionalised form of equation (3.1). There are two possibilities for this non-dimensionalisation, depending on whether the external trapping potential is present. Here, we show the two cases.

#### 3.2.1 Natural units

When the external trapping potential is absent, natural units are used to non-dimensionalise the GPE. In this case the scalings are

$$t \to \frac{\hbar}{2\mu}t,$$

$$r \to ar,$$

$$\psi \to \psi_{\infty}\psi,$$
(3.2)

where  $a = \hbar/(\sqrt{2m\mu})$  is the healing length, and  $\psi_{\infty} = \sqrt{\mu/g}$  is the bulk value of  $\psi$ . These scalings lead to the dimensionless form of the GPE

$$-2i\frac{\partial \psi}{\partial t} = \nabla^2 \psi + \left(1 - |\psi|^2\right)\psi.$$

See appendix A for a full derivation. To solve this form of the GPE, set eqn\_to\_solve to 1 in run.in.

### 3.2.2 Harmonic oscillator units

When the external trapping potential is present, harmonic oscillator units are used to non-dimensionalise the GPE. The scalings are

$$t \to \frac{t}{\overline{\omega}},$$

$$\mathbf{r} \to a_{\mathrm{OH}}\mathbf{r},$$

$$\psi \to a_{\mathrm{OH}}^{-\frac{3}{2}}\psi,$$

$$g \to a_{\mathrm{OH}}^{3}\hbar\overline{\omega}g,$$

$$\mu \to \hbar\overline{\omega}\mu,$$

$$V_{\mathrm{ext}} \to \hbar\overline{\omega}V_{\mathrm{ext}},$$

$$(3.3)$$

where  $\overline{\omega} = (\omega_x \omega_y \omega_z)^{1/3}$ ,  $\omega_i$  is the trap frequency along axis i, for i = x, y, z, and  $a_{\rm OH} = \sqrt{\hbar/(m\overline{\omega})}$  is the harmonic oscillator length.

These scalings lead to the alternative dimensionless form of the GPE

$$i\frac{\partial \psi}{\partial t} = -\frac{1}{2}\nabla^2 \psi + V_{\text{ext}}\psi + g|\psi|^2\psi - \mu\psi.$$

See appendix A for a full derivation. To solve this form of the GPE, set eqn\_to\_solve to 4 in run.in.

### 3.2.3 Alternative form

An alternative dimensionless form of the GPE can also be solved, specifically for the random phase approximation, as in Berloff & Svistunov (2002). This is

$$-2i\frac{\partial \psi}{\partial t} = \nabla^2 \psi + |\psi|^2 \psi,$$

and it can be selected by setting eqn\_to\_solve to 2 in run.in.

# 3.3 Dimensionless forms of other quantities

It is useful to know the dimensionless forms of other relevant variables and quantities. These are described in this section. Full derivations of each of the non-dimensionalisations are given in appendix A.

### 3.3.1 Harmonic trapping potential

The dimensional harmonic trapping potential is given by

$$V_{\text{ext}} = \frac{1}{2}m \left(\omega_x^2 x^2 + \omega_y^2 y^2 + \omega_z^2 z^2\right).$$
 (3.4)

Using the scaling  $\omega_i \to \overline{\omega}\omega_i$ , for  $i=x,\,y,\,z$ , and remembering to scale each coordinate, leads to the dimensionless form of the trapping potential

$$V_{\text{ext}} = \frac{1}{2} \left( \omega_x^2 x^2 + \omega_y^2 y^2 + \omega_z^2 z^2 \right).$$

### 3.3.2 Thomas–Fermi approximation

The dimensional Thomas–Fermi approximation is given by

$$\psi = \sqrt{\frac{\mu - V_{\text{ext}}}{g}}.$$

Since this approximation is only relevant for trapped condensates, we use harmonic oscillator units to non-dimensionalise. This leads to an identical dimensional expression, where each dimensionless variable is replaced by its dimensionless counterpart.

#### Condensate extent

The extent of the condensate  $R_i$ , i = x, y, z, in the Thomas-Fermi limit, is given by

$$R_i^2 = \frac{2\mu}{m\omega_i^2}, \quad i = x, y, z.$$

Non-dimensionalising using harmonic oscillator units yields

$$R_i^2 = \frac{2\mu}{\omega_i^2}, \quad i = x, y, z.$$

#### Number of atoms

The number of atoms within the condensate N, under the Thomas–Fermi approximation, is given by

$$N = \frac{8\pi}{15} \left( \frac{2\mu}{m\overline{\omega}^2} \right)^{\frac{3}{2}} \frac{\mu}{q}.$$

Again, non-dimensionalising using harmonic oscillator units, leads to

$$N = \frac{16\sqrt{2}\pi}{15} \frac{\mu^{\frac{5}{2}}}{g}.$$

### 3.3.3 Circulation

The dimensional circulation  $\kappa$ , around a vortex is defined to be

$$\kappa = \oint_C \boldsymbol{u} \cdot d\boldsymbol{l}.$$

Then, using the fact that  $\mathbf{u} = (\hbar/m)\nabla\phi$ , where  $\phi$  is the phase, and also noting that the circulation is quantised, such that the phase differs by  $2\pi n$  around the vortex, where n is the winding number, we obtain

$$\kappa = \frac{\hbar}{m} \oint_C \nabla \phi \cdot d\mathbf{l} = \frac{2\pi\hbar}{m} n.$$

Using natural units with the scaling  $\kappa \to (2\mu a^2/\hbar)\kappa$ , or harmonic oscillator units with the scaling  $\kappa \to a_{\rm OH}^2 \overline{\omega} \kappa$ , in both cases leads to a dimensionless circulation of  $\kappa = 2\pi n$ .

### 3.4 Initial conditions

The code implements three basic initial conditions, any combination of which can theoretically be multiplied together. This section outlines the mathematical description of the initial conditions; for a description of the numerical implementation see §5.2.2.

#### 3.4.1 Vortex line

A vortex line is given by

$$\psi_0 = f(r) \exp{(i\theta)},$$

in polar coordinates  $(r, \theta)$ , where f(r) is some function which models the vortex core, and  $\theta$  is the phase. In the code, we use the vortex core model proposed by Berloff & Roberts (2001), so that

$$f(r) = [1 - \exp(-0.7r^{1.15})] \exp(i\theta).$$

The vortex line initial condition can also support planar or helical perturbations along its length. For a vortex line oriented along the z-direction, this is achieved by imposing a sinusoidal perturbation of the form

$$\sin\left(\frac{2\pi z}{l}\right)$$
,

in the x-direction, and the same perturbation shifted by  $\pi/2$  in the y-direction. These perturbations can be set independently; doing so will result in a planar perturbation, while setting both will result in a helical perturbation.

Cyclically permuting x, y, and z results in a vortex line along another direction, and the initial position and circulation of the line can also be defined. See §5.2.1 for a description of the parameters which control the vortex line initial condition.

### 3.4.2 Vortex ring

A vortex ring in the (y, z)-plane, travelling in the x-direction takes the form (Berloff, 2004)

$$\psi_0 = \Psi(x, s + R_0)\Psi^*(x, s - R_0),$$

where  $\Psi(x,s)$  is given by

$$\Psi(x,s) = f\left(\sqrt{x^2 + s^2}\right) \exp(i\theta),$$

f(r) (which again models the vortex core) is given by

$$f(r)^{2} = \frac{r^{2} (a_{1} + a_{2}r^{2})}{1 + b_{1}r^{2} + b_{2}r^{4}},$$

where  $a_i$  and  $b_i$  are constants,  $s = \sqrt{y^2 + z^2}$ , and  $R_0$  is the radius of the ring.

The vortex ring initial condition also supports planar and helical perturbations, in a manner similar to that of the vortex line. The perturbed initial condition then takes the form

$$\psi_0 = \Psi \left\{ x - A_1 \cos(m_1 \theta), s + R_0 - A_2 \cos(m_2 \theta) \right\} \times \Psi^* \left\{ x - A_1 \cos(m_1 \theta), s - R_0 - A_2 \cos(m_2 \theta) \right\},$$

where  $\Psi(x,s)$  is given by

$$\Psi(x,s) = f\left(\sqrt{x^2 + s^2}\right)(x + is).$$

The function f(r) is as above,  $s = \sqrt{y^2 + z^2} - A_1 \sin(m_1\theta)$ ,  $A_1$  and  $m_1$  are the amplitude and wavenumber of a purely planar perturbation, and  $A_2$  and  $m_2$  are the amplitude and wavenumber of a purely helical perturbation.

The direction of motion of the vortex ring in the x-direction can be defined, but note that it is not yet possible to set up rings moving in the y- or z-directions. See §5.2.1 for a description of the parameters which control the vortex ring initial condition.

### 3.4.3 Random phase

The GPE can model the formation of a condensate, starting from a non-equilibrium initial condition. The random phase initial condition describes a weakly interacting Bose gas, where the particles remain in a strongly non-equilibrium state. Evolving this state leads to the formation of a quasi-condensate, consisting of a tangle of quantised vortices. This tangle decays as the system reaches thermal equilibrium, with a certain number of particles in the zero-momentum (genuine condensate) state. See, for example, Berloff & Svistunov (2002), Connaughton et al. (2005), and Berloff & Youd (2007) for more details on the random phase approximation.

To model this, the initial condition is set to

$$\psi_0 = \sum_{\mathbf{k}} a_{\mathbf{k}} \exp\left(\mathrm{i}\mathbf{k} \cdot \mathbf{r}\right),$$

where the ks are the wavenumbers in momentum space, and the phases of the complex amplitudes  $a_k$  are distributed randomly.

The mass and kinetic energy density must be set in run.in, prior to performing a random phase simulation. These are nv and enerv respectively.

# Chapter 4

# Numerical formulation

This chapter outlines the numerical formulation which is used to solve the GPE. Time stepping is performed explicitly using a choice of schemes, and spatial discretisation is performed with either second- or fourth-order centred finite differences.

# 4.1 Time stepping

As briefly mentioned in the introduction, a choice of time stepping schemes is possible. Each of these is explained in this section.

### 4.1.1 Euler's method

For a general partial differential equation of the form

$$\frac{\partial \psi}{\partial t} = f(\psi, \boldsymbol{r}, t),$$

where f is some function representing the right hand side of the equation, Euler's method is

$$\psi^{p+1} = \psi^p + \Delta t f^p + \mathcal{O}(\Delta t^2),$$

where  $\Delta t$  is the time step,  $\psi^p = \psi(x, y, z, t^p)$ ,  $t^p = p\Delta t$ , and  $f^p = f(\psi^p)$ .

This scheme is only first order accurate in  $\Delta t$ , and is not recommended for use, other than for very rough testing. It can be selected in the code by setting the parameter scheme equal to euler in run.in.

# 4.1.2 Second-order Runge-Kutta method

The second-order Runge–Kutta method (also known as the midpoint method) is second order accurate in time, and takes the form

$$\psi^{p+1} = \psi^p + \Delta t k_2 + \mathcal{O}(\Delta t^3),$$

where

$$k_1 = f(t^p, \psi^p),$$
  
 $k_2 = f\left(t^{p+\frac{1}{2}}, \psi^p + \frac{1}{2}\Delta t k_1\right).$ 

This scheme can be selected by setting scheme equal to rk2.

### 4.1.3 Fourth-order Runge–Kutta method

The fourth-order Runge–Kutta method is fourth order accurate in time and takes the form

$$\psi^{p+1} = \psi^p + \Delta t \left( \frac{k_1}{6} + \frac{k_2}{3} + \frac{k_3}{3} + \frac{k_4}{6} \right) + \mathcal{O}(\Delta t^5),$$

where

$$k_{1} = f(t^{p}, \psi^{p}),$$

$$k_{2} = f\left(t^{p+\frac{1}{2}}, \psi^{p} + \frac{1}{2}\Delta t k_{1}\right),$$

$$k_{3} = f\left(t^{p+\frac{1}{2}}, \psi^{p} + \frac{1}{2}\Delta t k_{2}\right),$$

$$k_{4} = f\left(t^{p+1}, \psi^{p} + \Delta t k_{3}\right).$$

This scheme can be selected by setting scheme equal to rk4.

# 4.1.4 Runge-Kutta-Fehlberg method

The Runge–Kutta–Fehlberg method is a hybrid fourth/fifth order adaptive time stepping scheme. In the code, the fifth-order formula is

$$k_{1} = \Delta t f (t^{p}, \psi^{p}),$$

$$k_{2} = \Delta t f (t^{p} + a_{2} \Delta t, \psi^{p} + b_{21} k_{1}),$$

$$k_{3} = \Delta t f (t^{p} + a_{3} \Delta t, \psi^{p} + b_{31} k_{1} + b_{32} k_{2}),$$

$$k_{4} = \Delta t f (t^{p} + a_{4} \Delta t, \psi^{p} + b_{41} k_{1} + b_{42} k_{2} + b_{43} k_{3}),$$

$$k_{5} = \Delta t f (t^{p} + a_{5} \Delta t, \psi^{p} + b_{51} k_{1} + b_{52} k_{2} + b_{53} k_{3} + b_{54} k_{4}),$$

$$k_{6} = \Delta t f (t^{p} + a_{6} \Delta t, \psi^{p} + b_{61} k_{1} + b_{62} k_{2} + b_{63} k_{3} + b_{64} k_{4} + b_{65} k_{5}),$$

$$\psi^{p+1} = \psi^{p} + c_{1} k_{1} + c_{2} k_{2} + c_{3} k_{3} + c_{4} k_{4} + c_{5} k_{5} + c_{6} k_{6} + \mathcal{O}(\Delta t^{6}).$$

The embedded fourth-order formula is given by

$$\psi_*^{p+1} = \psi^p + c_1^* k_1 + c_2^* k_2 + c_3^* k_3 + c_4^* k_4 + c_5^* k_5 + c_6^* k_6 + \mathcal{O}(\Delta t^5),$$

so that an error estimate can be obtained with

$$\Delta \equiv \psi^{p+1} - \psi_*^{p+1} = \sum_{i=1}^6 (c_i - c_i^*) k_i.$$

i	$a_i$			$b_{ij}$			$c_i$	$c_i^*$
1							$\frac{37}{378}$	$\frac{2825}{27648}$
2	$\frac{1}{5}$	$\frac{1}{5}$					0	0
3	$\frac{3}{10}$	$\frac{3}{40}$	$\frac{9}{40}$				$\frac{250}{621}$	$\frac{18575}{48384}$
4	<u>3</u> 5	$\frac{3}{10}$	$-\frac{9}{10}$	$\frac{6}{5}$			$\frac{125}{594}$	$\frac{13525}{55296}$
5	1	$-\frac{11}{54}$	$\frac{5}{2}$	$-\frac{70}{27}$	$\frac{35}{27}$		0	$\frac{277}{14336}$
6	$\frac{7}{8}$	$\frac{1631}{55296}$	$\frac{175}{512}$	$\frac{575}{13824}$	$\frac{44275}{110592}$	$\frac{253}{4096}$	$\frac{512}{1771}$	$\frac{1}{4}$
j	=	1	2	3	4	5		

Table 4.1: The Cash-Karp constants for the Runge-Kutta-Fehlberg scheme.

The values of the constants are those given by Cash & Karp (1990), and are reproduced here in table 4.1. Full details of the algorithm can be found in Numerical Recipes (§16.2, p.708, Press *et al.*, 1992).

This scheme can be selected by setting scheme equal to rk45.

# 4.2 Spatial discretisation

Centred finite differences are used to spatially discretise the GPE; these can be either second or fourth order.

For a general partial differential equation of the form

$$\frac{\partial \psi}{\partial t} = f(\psi, \boldsymbol{r}, t),$$

where f is some function representing the right hand side of the equation, the spatial discretisation takes the form

$$\frac{\partial \psi_{ijk}}{\partial t} = f_{ijk},$$

where  $\psi_{ijk} = \psi\left(x_i, y_j, z_k, t\right)$ ,  $x_i = i\Delta x$ ,  $y_j = j\Delta y$ ,  $z_k = k\Delta z$ , and  $f_{ijk} = f\left(\psi_{ijk}\right)$ .

### 4.2.1 Second-order finite differences

The second-order finite-difference approximation to the first derivative, with respect to x, is given by

$$\frac{\partial \psi}{\partial x} \approx \frac{\psi_{i+1} - \psi_{i-1}}{2\Delta x},$$

where we have dropped the j and k indices for clarity.

Similarly, the second-order approximation to the second derivative is

$$\frac{\partial^2 \psi}{\partial x^2} \approx \frac{\psi_{i+1} - 2\psi_i + \psi_{i-1}}{\Delta x^2}.$$

Second-order finite differences can be chosen by setting order equal to 2 in run.in.

#### 4.2.2 Fourth-order finite differences

The fourth-order finite-difference approximation to the first derivative, with respect to x, is given by

$$\frac{\partial \psi}{\partial x} \approx \frac{-\psi_{i+2} + 8\psi_{i+1} - 8\psi_{i-1} + \psi_{i-2}}{12\Delta x}.$$

Similarly, the fourth-order approximation to the second derivative is

$$\frac{\partial^2 \psi}{\partial x^2} \approx \frac{-\psi_{i+2} + 16\psi_{i+1} - 30\psi_i + 16\psi_{i-1} - \psi_{i-2}}{12\Delta x^2}.$$

Fourth-order finite differences can be chosen by setting order equal to 4.

# 4.3 Boundary conditions

The code implements both periodic and reflective boundary conditions. Since both of these conditions are behavioural, rather than numerical, no boundary conditions are explicitly set on  $\psi$  itself. Note that it is not possible to mix boundary conditions yet.

# 4.3.1 Periodic boundary conditions

Suppose the x-coordinate runs from  $x_0$  to  $x_n$ . Then clearly the values of  $\psi$  at  $x_{-1}$  and  $x_{n+1}$  are needed to compute a first derivative, for example.

Periodic boundary conditions are implemented such that

$$x_{-1} = x_n,$$
  
$$x_{n+1} = x_0.$$

These conditions are also applied in the y- and z-directions, and can be naturally extended for higher order approximations.

To use periodic boundary conditions set bcs equal to 1 in run.in.

# 4.3.2 Reflective boundary conditions

The physical meaning of reflective boundary conditions is that structures within the computational box see images of themselves across the box boundaries. This would cause a vortex ring to annihilate on approaching a boundary, for example, as it sees an image of

itself approaching the boundary from the opposite direction (outside the computational box).

These boundary conditions are implemented such that

$$x_{-1} = x_1, x_{n+1} = x_{n-1}.$$

To use reflective boundary conditions set bcs equal to 2.

# Chapter 5

# File reference

This chapter is intended as a reference for the main files associated with the code, such as the source files, input and output files, and the IDL gpe.pro program.

# 5.1 Program source files

Table 5.1 lists the .f90 source code files which make up the code, and describes their functionality.

File	Description
constants.f90	This is a module which defines the constants needed for
	FFTw.
derivs.f90	Routines to do with derivatives.
error.f90	Error-handling routines.
gpe.f90	Main program file.
ic.f90	Routines to do with setting up the initial condition and gen-
	eral initialisation.
parameters.f90	Compile-time parameters and global variables.
solve.f90	Routines to do with actually solving the equation, for exam-
	ple, the time stepping algorithms are defined in this file.
variables.f90	User-defined types, and other general routines to do with the
	equation variables.

Table 5.1: Program source files which make up the GPE code.

# 5.2 Program input files

This section describes the .in input files which will generally need to be edited to set up a run.

### 5.2.1 parameters.in

Edit this file to set the floating-point precision, the number of processes, and the grid dimensions of the run. If the initial condition consists of a vortex line or vortex ring, then the line/ring parameters can also be set here. Note that any changes to this file will require a recompilation of the code. See tables 5.2, 5.3, and 5.4 for a description of the parameters.

Parameter	Type	Description
pr	integer	The precision of real variables is parametrised. Choose the
		desired line for either real or double precision.
nyprocs	integer	The number of processes in the $y$ -direction.
nzprocs	integer	The number of processes in the $z$ -direction.
nx	integer	The number of grid points in the x-direction.
ny	integer	The number of grid points in the $y$ -direction.
nz	integer	The number of grid points in the $z$ -direction.

Table 5.2: Compile-time parameters to set.

There are no restrictions on the nyprocs and nzprocs parameters, other than that they be  $\geq 1$ . In general, it is recommended that nzprocs  $\geq$  nyprocs for best performance, so that fewer non-contiguous data transfers are performed. They can both be set to 1, in which case the job is simply run on one process (an MPI parallel environment is still required though).

Parameter	Type	Description
x0	real	x-position of the line.
yО	real	y-position of the line.
z0	real	z-position of the line.
amp1	real	Amplitude of a sinusoidal disturbance in one direction
		along the line.
amp2	real	Amplitude of a sinusoidal disturbance in the other di-
		rection along the line.
11	real	The wavelength of the above disturbances.
sgn	real	The sign of the argument of the line (i.e. circulation
		direction).
dir	character	The direction $(x, y, \text{ or } z)$ in which the line should
		extend.
$imprint_phase$	logical	Whether only the phase should be imprinted, i.e. no
		vortex core should be modelled.

Table 5.3: Compile-time parameters for a vortex line.

Parameter	Type	Description
x0	real	x-position of the ring.
yО	real	y-position of the ring.
z0	real	z-position of the ring.
amp	real	Amplitude of a planar disturbance around the ring.
mm	integer	Wavenumber of a planar disturbance.
r1	real	Amplitude of a helical disturbance around the ring.
kk	integer	Wavenumber of a helical disturbance.
dir	real	Ring propagation direction $(\pm 1)$ .

Table 5.4: Compile-time parameters for a vortex ring.

New vortex lines and rings can be defined simply by copying an existing definition in parameters.in, and renaming, so for example, the *ring* example defines

```
type (ring_param), parameter :: &
  vr1 = ring_param(0.0_pr, 0.0_pr, 0.0_pr, 10.0_pr, &
      0.0_pr, 5, 0.0_pr, 10, -1.0_pr)
```

To create another vortex ring definition, say with a radius of 20, and situated at x = 5, define

```
type (ring_param), parameter :: &
    vr2 = ring_param(5.0_pr, 0.0_pr, 0.0_pr, 20.0_pr, &
        0.0_pr, 5, 0.0_pr, 10, -1.0_pr)
```

Then an initial condition consisting of these two vortex rings could be set up as described in the next section.

#### 5.2.2 ic.in

This file defines the initial condition. The initial condition must be defined in the form init\_cond = function(), where function() is some function in the code which defines a possible component of an initial condition. Components can be multiplied together to form any number of different initial conditions. Look at ic.in in the src directory to see some examples.

In the *ring* example, the initial condition is set to vortex\_ring(vr1), where vr1 is a type parameter declared in parameters.in (see above).

As another example, if you define vr2 as above, then you could construct an initial condition consisting of two rings with

```
init_cond = vortex_ring(vr1) * vortex_ring(vr2)
```

In this way, any number of initial conditions can be constructed, simply by multiplying functions together.

As with parameters.in, any changes to this file will require the code to be recompiled.

### 5.2.3 run.in

This file defines the main parameters for the run, as a set of Fortran namelists. Each namelist loosely collects together related parameters. The namelists are:

- run\_params these are parameters to do with the run itself, such as time step, time stepping scheme, when the run should end, which equation to solve, etc.;
- eqn\_params these parameters set properties of the equation, such as whether it should be solved in a moving reference frame, trap parameters, random phase parameters, etc.;
- io\_params these parameters control input/output, for example, what data should be saved and how often;
- misc\_params miscellaneous parameters which do not fit in the other categories.

Table 5.5 describes these parameters in detail.

Parameter	Type	Description
tau	real	Time step (initial time step for RK45). Valid for both
		imaginary and real time.
end_time	real	The final (dimensionless) time.
xr	real	The x-coordinate of the right-hand-side of the compu-
		tational box (the left-hand-side is set to -xr).
yr	real	As above but for the $y$ -coordinate.
zr	real	As above but for the $z$ -coordinate.
scheme	character	The time stepping scheme to use. This must be set
		to one of euler (for explicit second-order Euler time
		stepping), rk2 (for explicit second-order Runge–Kutta
		time stepping), rk4 (for explicit fourth-order Runge-
		Kutta time stepping), or rk45 (for explicit adaptive
		fourth-order Runge–Kutta–Fehlberg time stepping).
eqn_to_solve	integer	The form of the GPE to solve. See §3.2 for the possible
		values to use.
bcs	integer	The boundary conditions to use. Set to 1 for periodic
		BCs; set to 2 for reflective BCs.
order	integer	The order of the derivatives to use. Set to 2 for second-
		order; set to 4 for fourth-order.
restart	logical	Set to .true. to do a restart of a previous run. See
		§2.8.
saved_restart	logical	Set to .true. if using filtered data from a previous run
		to multiply with the initial condition of a new run.
renorm	logical	Set to .true. if the wavefunction should be renor-
		malised at every time step in imaginary time.

imprint_vl	logical	Set to .true. if the wavefunction should be multiplied by a vortex line at each time step in imaginary time.
stop_imag	logical	Set to .true. if the run should stop at the end of imaginary time, i.e. when the relative norms of successive time steps are deemed to be sufficiently close (currently $10^{-12}$ ).
real_time	logical	Set to .true. if the run should be started in real time.
Urhs	real	Set non-zero to solve the equation in a moving reference frame.
diss_amp	real	Set non-zero to include dissipation of this amplitude at the boundaries.
scal	real	Set non-zero to scale vortex rings/lines.
nv	real	For random phase approximation, the total mass per unit volume.
enerv	real	For random phase approximation, the total kinetic energy per unit volume.
g	real	Interaction parameter for trapped condensate.
mu	real	Chemical potential for trapped condensate.
nn	real	Number of atoms in a trapped condensate. Should currently be left to 1.0, and instead tune mu and g to specify nn.
omx	real	Frequency of trap in $x$ -direction.
omy	real	Frequency of trap in $y$ -direction.
omz	real	Frequency of trap in $z$ -direction.
save_rate	integer	The rate at which time-series data should be saved (roughly corresponding to the number of time steps).
save_rate2	real	How often (in terms of actual time units) isosurface data should be saved (3D isosurfaces, 2D surfaces).
save_rate3	real	How often (in terms of actual time units) data to do with condensed particles and PDFs should be saved.
p_save	real	How often (in terms of actual time units) the code should save its own state, so that it can be restarted in the event of a machine failure, for example.
save_contour	logical	Should 2D contour data be saved?
save_3d	logical	Should 3D isosurface data be saved?
save_filter	logical	Should 3D filtered isosurfaces of the density be saved?
filter_kc	real	The cutoff wavenumber used to filter the isosurfaces.
save_average	logical	Should 3D time-averaged isosurfaces of the density be saved?
save_spectrum	logical	Should various spectra be saved (mainly for random phase approximation)?
save_pdf	logical	Should PDFs of the velocity components be saved?
save_vcf	logical	Should the velocity correlation function be saved?

save_ll	logical	Should the vortex line length be saved?
save_zeros	logical	Should the points of zero density be saved? (Not reli-
		able!)

Table 5.5: Run-time parameters to set.

### 5.3 Makefile

Settings in the Makefile may need to be changed, depending on the architecture on which the code is run, and what compilers are available to you. Table 5.6 describes the Makefile variables.

Variable	Description
OBJECT	The name of the executable produced on compilation.
OBJS	The object files that should be linked.
FC	The Fortran compiler to be used. This could be the MPI
	wrapper compiler mpif90, but the underlying Fortran com-
	piler must be able to compile the code (e.g. sunf95, ifort,
	gfortran).
FFLAGS	Compiler flags. See the compiler's manual. Sunf95 works well
	with -fast. If nonsense results are produced -fsimple=0
	might be required.
LDFFTW	FFTw libraries to link. Compiling with
	make precision=single will link the single precision
	libraries.
LDFLAGS	Any extra flags required by the linker. If mpif90 is not the
	compiler, then all the MPI libraries will need to be linked.
INCLUDE	Include path (e.g. for MPI header files).

Table 5.6: Description of Makefile variables.

# 5.4 Program output files

The output that the program produces depends on the parameters set in the io\_params namelist. Table 5.7 briefly describes these files.

# 5.4.1 Binary layout

If you do not have access to IDL, but still want to be able to view contour and isosurface plots, then you will need to know how this data is arranged in the numbered dens\*\*\*\*\*\*.dat files, within each proc directory. Table 5.8 describes the data which are saved to these files, and how much space (in terms of bytes) is needed to store the data.

### 5.5 run.sh

This script can be used to start a run on a shared memory machine, such as those with the latest multi-core processors, or older multi-processor machines. Usage instructions are provided with the script itself; run with no arguments to see the help.

# 5.6 The IDL gpe.pro program

If you have access to IDL, then all of the contour and isosurface visualisation can be done through the gpe.pro program. This section will briefly describe its use. Detailed examples of some aspects of the program are given in §2.7.

### 5.6.1 Floating point precision

It is important to note the precision of the run for which you want to view results. If you have performed a double-precision run, then all gpe commands must include the /dbl keyword.

# 5.6.2 Fortran unformatted I/O

The GPE code used to write binary data using Fortran 77 unformatted, record-based I/O. With the advent of Fortran 2003 stream I/O, this is no longer necessary, and record markers are not part of the data. It might still be necessary to view the old data, however, in which case the /f77 keyword should be used in all gpe commands.

# 5.6.3 Isosurface plots

The program will produce an isosurface plot of the density  $|\psi|^2$  if no keywords or options are provided, e.g.

To change the default surface colour, set the index option, e.g.

The index option must be in the range 0 to 255, and corresponds to the index into the currently loaded IDL colour table. The isosurface level can also be controlled, by using the level option, e.g.

### 5.6.4 Contour plots

A contour plot of the density in the (x, y)-plane at z = 0 is displayed with the addition of the /cntr keyword, e.g.

Plots of the phase or velocities can also be produced by adding the /phase, /vx, /vy, or /vz keywords, e.g.

The position of the contour slice can be controlled by using the xpos, ypos, or zpos options, e.g.

The position is given in real units (as opposed to grid units). The plane in which the contour slice sits can be controlled with the dir option, e.g.

This will produce a contour plot in the (x, z)-plane.

### 5.6.5 Slice plots

One-dimensional slices through the data can also be generated, by using the /slice keyword, and the position and direction controlled as for contour plots, e.g.

### 5.6.6 Contour animations

A series of contour snapshots can be generated, by using the /c\_anim keyword, e.g.

Snapshots are saved to the directory images, from the directory under which IDL is started, so this directory must exist before attempting to create snapshots, otherwise an error will result.

### 5.6.7 Isosurface animations

A series of isosurface snapshots can be generated, by using the /png keyword, e.g.

Currently, it is only possible to generate isosurfaces of the density.

# 5.6.8 EPS output

High quality EPS figures of all 1D, 2D, and 3D plots can be produced, by using the /eps keyword. Figures are saved in the images directory (as for the snapshots in the animations above), so this directory must exist prior to attempting to save as EPS.

File	Description			
energy.dat	Saves the energy at each time.			
linelength.dat	Saves the total line length of vortices in the condensate.			
mass.dat	Saves the mass at each time.			
minmax_*.dat	Saves the minimum and maximum of the density, filtered			
	density, and time-averaged density, over the duration of th			
	run.			
norm.dat	Saves the norm at each time.			
misc.dat	Any miscellaneous data can be sent to this file.			
momentum.dat	Saves the three components of the momentum.			
p_saved.dat	The values of the time index p when the code saved its own			
	state.			
save.dat	The parameters for the most recently saved state.			
timestep.dat	The imaginary and real time step at each time, if adaptive			
	time stepping is chosen.			
psi_time.dat	Saves the real and imaginary time, the real and imaginary			
	parts of the wavefunction, the density, and the phase.			
proc**	Numbered directories corresponding to each process involved			
	in the run. Each of these directories contains the binary files			
	listed below.			
end_state.dat	The saved state of the run.			
im_zeros******.dat	The coordinates where the imaginary part of the wavefunc-			
	tion goes to zero.			
re_zeros******.dat	The coordinates where the real part of the wavefunction goes			
	to zero.			
dens******.dat	If 3D isosurfaces are requested, the data for them are saved			
	in these files.			
filtered*****.dat	As above, but for filtered data.			
ave******.dat	As above, but for time-averaged data.			
spectrum******.dat	The spectrum $(n_k \text{ vs. } k)$ .			
zeros******.dat	The coordinates where the real and imaginary parts of the			
	wavefunction simultaneously go to zero.			

Table 5.7: Output files from the GPE code.

Variable	Type	Size (bytes)		Description	
	0.1	Single	Double	•	
t+im_t	real	4	8	The total time elapsed.	
nx	integer	4	4	Number of grid points in the $x$ -	
				direction.	
ny	integer	4	4	Number of grid points in the $y$ -	
				direction.	
nz	integer	4	4	Number of grid points in the $z$ -	
				direction.	
nyprocs	integer	4	4	Number of processes in the $y$ -	
				direction.	
nzprocs	integer	4	4	Number of processes in the $z$ -	
				direction.	
js	integer	4	4	Starting index of data in $y$ -direction,	
				local to each process.	
je	integer	4	4	Ending index of data in $y$ -direction,	
				local to each process.	
ks	integer	4	4	Starting index of data in $z$ -direction,	
				local to each process.	
ke	integer	4	4	Ending index of data in $z$ -direction,	
				local to each process.	
psi	complex	8nx*nyl*nzl	16nx*nyl*nzl	Complex wavefunction $\psi$ , local to	
				each process.	
X	real	$4\mathtt{nx}$	8nx	The grid array in the $x$ -direction.	
У	real	4ny	8ny	The grid array in the $y$ -direction.	
Z	real	4nz	8nz	The grid array in the $z$ -direction.	

Table 5.8: Data saved to the dens\*\*\*\*\*\*.dat files, with sizes in bytes for both single and double precision (assuming x86 or x86-64 architectures). The number of grid points local to each process in the y- and z-directions is given by nyl=je-js and nzl=ke-ks.

# Appendix A

# Derivations of non-dimensionalised equations and variables

This chapter derives the non-dimensionalisations of the GPE, and related quantities, in more detail.

# A.1 Physical units

The following table lists the physical units for various quantities related to the GPE and Bose–Einstein condensation.

Quantity	Description	Unit
$\overline{a}$	Healing length	m
$a_{\mathrm{OH}}$	Harmonic oscillator length	m
$\hbar$	Reduced Planck constant	Js
m	Mass	$\rm Js^2m^{-2}$
$\mu$	Chemical potential	J
g	Interaction parameter	$\mathrm{Jm}^3$
$\psi$	Complex wavefunction	$m^{-\frac{3}{2}}$
$\omega$	Trap frequency	$s^{-1}$
$V_{ m ext}$	Trapping potential	J
$\kappa$	Circulation	$\mathrm{m}^2\mathrm{s}^{-1}$

Table A.1: Physical units for quantities related to the GPE and Bose–Einstein condensation.

### A.2 Natural units non-dimensionalisation

Using the scalings given by (3.2), the dimensional GPE (3.1) becomes (ignoring the trapping potential  $V_{\text{ext}}$ )

$$i\hbar \frac{2\mu}{\hbar} \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{1}{a^2} \nabla^2 \psi + g\psi_{\infty}^2 |\psi|^2 \psi - \mu \psi.$$

Each term in the GPE involves a term in  $\psi$ . The factor of  $\psi_{\infty}$  used to non-dimensionalise  $\psi$  is therefore cancelled throughout.

Cancelling  $\hbar$  on the left, and dividing through by  $\mu$  leads to

$$2\mathrm{i}\frac{\partial\psi}{\partial t} = -\frac{\hbar^2}{2\mu m}\frac{1}{a^2}\nabla^2\psi + \frac{g}{\mu}\psi_\infty^2|\psi|^2\psi - \psi.$$

Then

$$-\frac{\hbar^2}{2\mu m}\frac{1}{a^2} = -\frac{\hbar^2}{2\mu m}\frac{2\mu m}{\hbar^2} = -1,$$

and

$$\frac{g}{\mu}\psi_{\infty}^2 = \frac{g}{\mu}\frac{\mu}{g} = -1,$$

and so the non-dimensionalised GPE using natural units is

$$-2i\frac{\partial \psi}{\partial t} = \nabla^2 \psi + (1 - |\psi|^2) \psi.$$

# A.3 Harmonic oscillator units non-dimensionalisation

Using the scalings given by (3.3), the dimensional GPE (3.1) becomes

$$i\hbar\overline{\omega}\frac{\partial\psi}{\partial t} = -\frac{\hbar^2}{2m}\frac{1}{a_{\rm OH}^2}\nabla^2\psi + \hbar\overline{\omega}V_{\rm ext}\psi + a_{\rm OH}^3\hbar\overline{\omega}g\frac{1}{a_{\rm OH}^3}|\psi|^2\psi - \hbar\overline{\omega}\mu\psi,$$

where again, the scaling factor for  $\psi$  is cancelled throughout.

Then

$$-\frac{\hbar^2}{2m}\frac{1}{a_{\mathrm{OH}}^2} = -\frac{\hbar^2}{2m}\frac{m\overline{\omega}}{\hbar} = -\frac{1}{2}\hbar\overline{\omega},$$

and

$$a_{\mathrm{OH}}^{3}\hbar\overline{\omega}g\frac{1}{a_{\mathrm{OH}}^{3}}=\hbar\overline{\omega}g.$$

Then, dividing through by  $\hbar \overline{\omega}$  leads to the GPE, non-dimensionalised using harmonic oscillator units

$$i\frac{\partial \psi}{\partial t} = -\frac{1}{2}\nabla^2 \psi + V_{\text{ext}}\psi + g|\psi|^2\psi - \mu\psi.$$

# A.4 Harmonic trapping potential

Using the scalings  $V_{\rm ext} \to \hbar \overline{\omega} V_{\rm ext}$ ,  $\omega_i \to \overline{\omega} \omega_i$ , for i = x, y, z, and  $\mathbf{r} \to a_{\rm OH} \mathbf{r}$ , the dimensional trapping potential (3.4) becomes

$$\hbar \overline{\omega} V_{\text{ext}} = \frac{1}{2} m a_{\text{OH}}^2 \overline{\omega}^2 \left( \omega_x^2 x^2 + \omega_y^2 y^2 + \omega_z^2 z^2 \right)$$
$$= \frac{1}{2} m \frac{\hbar}{m \overline{\omega}} \overline{\omega}^2 \left( \cdots \right)$$
$$= \frac{1}{2} \hbar \overline{\omega} \left( \cdots \right),$$

and therefore, on dividing through by  $\hbar \overline{\omega}$ , the dimensionless trapping potential is

$$V_{\text{ext}} = \frac{1}{2} \left( \omega_x^2 x^2 + \omega_y^2 y^2 + \omega_z^2 z^2 \right).$$

# A.5 Thomas–Fermi approximation

Starting from

$$|\psi|^2 = \frac{\mu - V_{\text{ext}}}{g},$$

and scaling with harmonic oscillator units, leads to

$$\frac{1}{a_{\rm OH}^3} |\psi|^2 = \frac{\hbar \overline{\omega} \mu - \hbar \overline{\omega} V_{\rm ext}}{a_{\rm OH}^3 \hbar \overline{\omega} g},$$

so that the dimensionless form of the Thomas–Fermi approximation is

$$|\psi|^2 = \frac{\mu - V_{\text{ext}}}{g}.$$

### A.5.1 Condensate extent

Starting from

$$R_i^2 = \frac{2\mu}{m\omega_i^2},$$

for  $i=x,\,y,\,z,$  and again scaling using harmonic oscillator units, gives

$$a_{\rm OH}^2 R_i^2 = \frac{2\hbar \overline{\omega} \mu}{m \overline{\omega}^2 \omega_i^2}$$
$$= \frac{2a_{\rm OH}^2 \overline{\omega} \mu}{\overline{\omega} \omega_i^2},$$

and, on dividing through by  $a_{\rm OH}^2$ , we have the dimensionless condensate extent

$$R_i^2 = \frac{2\mu}{\omega_i^2}.$$

### A.5.2 Number of atoms

The number of atoms within the condensate, under the Thomas–Fermi approximation, is given by

$$N = \frac{8\pi}{15} \left( \frac{2\mu}{m\overline{\omega}^2} \right)^{\frac{3}{2}} \frac{\mu}{g}.$$

Non-dimensionalising gives

$$N = \frac{8\pi}{15} \left( \frac{2\hbar \overline{\omega} \mu}{m \overline{\omega}^2} \right)^{\frac{3}{2}} \frac{\hbar \overline{\omega} \mu}{a_{\text{OH}}^3 \hbar \overline{\omega} g}$$

$$= \frac{8\pi}{15} \left( 2a_{\text{OH}}^2 \mu \right)^{\frac{3}{2}} \frac{\mu}{a_{\text{OH}}^3 g}$$

$$= \frac{8\pi}{15} \left( 2\mu \right)^{\frac{3}{2}} \frac{\mu}{g}$$

$$= \frac{8\pi}{15} 2^{\frac{3}{2}} \frac{\mu^{\frac{5}{2}}}{g}$$

$$= \frac{16\sqrt{2}\pi}{15} \frac{\mu^{\frac{5}{2}}}{g}.$$

### A.6 Circulation

The dimensional circulation is given by

$$\kappa = \frac{h}{m} = \frac{2\pi\hbar}{m}.$$

Letting  $\kappa \to (2\mu a^2/\hbar)\kappa$ , and scaling with natural units, gives

$$\frac{2\mu a^2}{\hbar} \kappa = \frac{2\pi\hbar}{m}$$

$$\implies \kappa = \frac{2\pi\hbar^2}{2\mu a^2 m}$$

$$= \frac{2\pi a^2}{a^2}$$

$$= 2\pi.$$

Simlarly, letting  $\kappa \to a_{\rm OH}^2 \overline{\omega} \kappa$ , and scaling with harmonic oscillator units, gives

$$a_{\mathrm{OH}}^{2}\overline{\omega}\kappa = \frac{2\pi\hbar}{m}$$

$$\Longrightarrow \kappa = \frac{2\pi\hbar}{a_{\mathrm{OH}}^{2}\overline{\omega}m}$$

$$= \frac{2\pi a_{\mathrm{OH}}^{2}}{a_{\mathrm{OH}}^{2}}$$

$$= 2\pi.$$

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