Introduction to Quantum Mechanics

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1 The Wave Function

What we are looking for is the wave function Ψ .

Law 1.1 (Schrodinger Equation).

$$i\hbar\frac{\partial\Psi}{\partial t} = -\frac{\hbar^2}{2m}\frac{\partial^2\Psi}{\partial x^2} + V\Psi.$$

For simplicity, we always rewrite it as:

$$i\hbar\partial_t\Psi = -\frac{\hbar^2}{2m}\partial_x^2\Psi + V\Psi.$$

Born's statistical interpretation:

 $\int_a^b |\Psi(x,t)|^2 \, \mathrm{d}x = \text{probability of finding the particle between } a \text{ and } b \text{ at time } t.$

Law 1.2 (Normalization).

$$\int_{-\infty}^{\infty} |\Psi(x,t)|^2 \, \mathrm{d}x = 1.$$

Proposition 1.1. The wave function will always stay NORMALIZED.

$$\frac{\mathrm{d}}{\mathrm{d}t} \int_{-\infty}^{\infty} \left| \Psi(x,t) \right|^2 \mathrm{d}x = 0.$$

Proof. By Schrodinger EQ.,

LHS =
$$\frac{i\hbar}{2m} \left(\Psi^* \frac{\partial \Psi}{\partial x} - \frac{\partial \Psi^*}{\partial x} \Psi \right) \Big|_{-\infty}^{+\infty}$$
.

Definition 1.1.

$$\langle x \rangle \stackrel{def}{=} \int_{-\infty}^{\infty} x |\Psi|^2 dx$$

and

$$\langle p \rangle \stackrel{def}{=} m \frac{\mathrm{d} \langle x \rangle}{\mathrm{d}t}.$$

Theorem 1.1.

$$\langle x \rangle = \int \Psi^*(x) \Psi \, \mathrm{d}x$$

and

$$\langle p \rangle = \int \Psi^* \left(-i\hbar \frac{\partial}{\partial x} \right) \Psi \, \mathrm{d}x.$$

Remark 1.1 (Operator). We say that the operator x represents position, and the operator $-i\hbar \partial/\partial x$ represents momentum. Also,

$$\langle Q(x,p)\rangle = \int_{-\infty}^{\infty} \Psi^* \left[Q(x,-i\hbar \frac{\partial}{\partial x}) \right] \Psi \, \mathrm{d}x.$$

Property 1.1. Operators do **NOT**, in general, commute. For example, $\hat{x}\hat{p} \neq \hat{p}\hat{x}$, i.e.,

 \exists a function f, s.t. $(\hat{x}\hat{p})f \neq (\hat{p}\hat{x})f$.

Theorem 1.2 (de Broglie formula). The wave length is related to the momentum of the particle:

$$p = \frac{h}{\lambda} = \frac{2\pi\hbar}{\lambda}.$$

Theorem 1.3 (Heisenberg's uncertainty principle).

$$\sigma_x \sigma_p \ge \frac{\hbar}{2}.$$

2 Time-independent Schrodinger Equation

2.1 Stationary states

We look for solutions that are simple products,

$$\Psi(x,t) = \psi(x)\varphi(t).$$

Theorem 2.1. By the method of separation of variables,

$$-\frac{\hbar^2}{2m}\frac{\mathrm{d}^2\psi}{\mathrm{d}x^2} + V\psi = E\psi$$

and

$$\varphi(t) = e^{-iEt/\hbar}.$$

The first is called the **time-independent Schrodinger equation**.

Definition 2.1 (Hamiltonian). In classical mechanics, the total energy (kinetic plus potential) is called Hamiltonian:

$$H(x,p) = \frac{p^2}{2m} + V(x).$$

Now we introduce Hamiltonian operator:

$$\hat{H} = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V(x).$$

Thus the time-independent Schrodinger EQ. can be written

$$\hat{H}\psi = E\psi$$

which is **IMPORTANT**.

Remark 2.1. Intriguingly and intuitively,

$$\langle H \rangle = E.$$

Also, if the equation yields an infinite collection of solutions $(\psi_1(x), \psi_2(x), \cdots)$, each with its associated value of the separation constant $(E1, E2, \cdots)$; thus the wave function is:

$$\Psi(x,t) = \sum_{n=1}^{+\infty} c_n \psi_n(x) e^{-iE_n t/\hbar}.$$

Particularly,

$$E_n \geq 0$$
 for all n

2.2 The infinite square well

Suppose

$$V(x) = \begin{cases} 0 & \text{if } 0 \le x \le a \\ \infty & \text{otherwise} \end{cases}.$$

Theorem 2.2. Inside the well, we have

$$E_n = \frac{n^2 \pi^2 \hbar^2}{2ma^2}$$

and

$$\psi_n(x) = \sqrt{\frac{2}{a}} \sin\left(\frac{n\pi}{a}x\right).$$

Property 2.1. $\psi_n(x)$ has some interesting and important porperties:

- 1. They are alternately even and odd, with the respect to the center of the well.
- 2. They are mutually orthogonal (i.e., $\int \psi_m(x)^* \psi_n(x) dx = \delta_{mn}$) where δ_{mn} is **Kronecker delta**:

$$\delta_{mn} = \begin{cases} 0, & \text{if } m \neq n \\ 1, & \text{if } m = n \end{cases}.$$

3. They are complete by Dirichlet's theorem.

2.3 The harmonic oscillator

Let

$$V(x) = \frac{1}{2}m\omega^2 x^2.$$

Here I will introduce 2 entirely different approaches to this problem. The first is a diabolically clever algebraic technique and the second is a straitforward "brute force" solution.

2.3.1 Algebraic method

To begin with, let's rewrite the EQ. in a more suggestive form:

$$\frac{1}{2m} \left[\left(-i\hbar \frac{\mathrm{d}}{\mathrm{d}x} \right)^2 + \left(m\omega x \right)^2 \right] \psi = E\psi.$$

The idea is to factor the term in square brackets:

$$u^{2} + v^{2} = (u - iv)(u + iv).$$

Definition 2.2 (Ladder operator).

$$\hat{a}_{\pm} = \frac{1}{\sqrt{2\hbar m\omega}} (\mp i\hat{p} + m\omega x).$$

Definition 2.3 (Commutator). The commutator of operators \hat{A} and \hat{B} is

$$\left[\hat{A}, \hat{B}\right] \stackrel{def}{=\!\!\!=\!\!\!=} \hat{A}\hat{B} - \hat{B}\hat{A}.$$

Property 2.2.

$$[\hat{a}_{-}, \hat{a}_{+}] = 1.$$

Theorem 2.3. If ψ satisfies the Schrodinger's EQ. with energy E, then $\hat{a}_+\psi$ satisfies the Schrodinger's EQ. with energy $E + \hbar\omega$:

$$\hat{H}\psi = E\psi \Longrightarrow \hat{H}(\hat{a}_+\psi) = (E + \hbar\omega)(\hat{a}_+\psi).$$

Similarly,

$$\hat{H}\psi = E\psi \Longrightarrow \hat{H}(\hat{a}_-\psi) = (E - \hbar\omega)(\hat{a}_-\psi).$$

Proof.

$$\hat{H} = a_+ a_- + \frac{1}{2}\hbar\omega.$$

Here, then, is a wonderful machine for generating new solutions—if we could just find one solution. Thus, we call \hat{a}_+ raising operator and \hat{a}_- lowering operator.

But what if I apply the lowering operator **repeatly**? We will reach a state with energy less than zero. By 2.1, there is **NO** guarantee that it will be normalized.

Proposition 2.1. Thus, there occurs a "lowest rung" ψ_0 such that

$$\hat{a}_{-}\psi_{0}=0.$$

Theorem 2.4.

$$\psi_0(x) = A_0 e^{-m\omega/2\hbar x^2}$$

and

$$E_0 = \frac{1}{2}\hbar\omega.$$

Thus we could get

$$\psi_n(x) = A_n(a_+)^n e^{-m\omega/2\hbar x^2}$$
, with $E_n = \left(n + \frac{1}{2}\right)\hbar\omega$

where A_n are used for normalization.

Theorem 2.5. ψ_n and ψ_{n+1} should satisfy:

$$\begin{cases} a_+\psi_n = i\sqrt{(n+1)\hbar\omega} \\ a_-\psi_n = -i\sqrt{n\hbar\omega}\psi_{n-1} \end{cases}.$$

Proof.

$$\int_{-\infty}^{\infty} |a_+ \psi_n|^2 dx = (n+1)\hbar\omega$$

and

$$\int_{-\infty}^{\infty} |a_{-}\psi_{n}|^{2} \, \mathrm{d}x = n\hbar\omega.$$

Ultimately,

$$A_n = \left(\frac{m\omega}{\pi\hbar}\right)^{1/4} \frac{(-i)^n}{\sqrt{n!(\hbar\omega)^n}}.$$

2.3.2 Analytic method

Things look a little cleaner if we introduce the dimensionless variables

$$\xi = \sqrt{\frac{m\omega}{\hbar}}x$$
 and $K = \frac{2E}{\hbar\omega}$.

In terms of ξ and K, the Schrodinger equation reads

$$\frac{\mathrm{d}^2 \psi}{\mathrm{d}\xi^2} = (\xi^2 - K)\psi.$$

To begin with, consider that at very large ξ , ξ^2 completely dominates over the constant K, so in this regime $\mathrm{d}^2\psi/\mathrm{d}\xi^2=\xi^2\psi$, which means that $\psi\Longrightarrow Ae^{\xi^2/2}+Be^{-\xi^2/2}$. Thus we let $\psi=h(\xi)e^{-\xi^2/2}$. Plugging ψ into Schordinger EQ., we have

$$h(\xi) = \sum_{n=0}^{\infty} a_n \xi^n$$
 and $a_{n+2} = \frac{2n+1-K}{(n+1)(n+2)}$.

For physically acceptable solutions (normalizable solutions), then, we must have K = 2n + 1. Finally,

$$\psi_n(x) = \left(\frac{m\omega}{\pi\hbar}\right)^{1/4} \frac{1}{\sqrt{2^n n!}} H_n(\xi) e^{-\xi^2/2}$$

where H_n is the **Hermite polynomials**.

-2

0

2

4

6

1 0.4 0.8 0.2 0.60 0.4-0.20.2-0.40 -22 0 2 -20 1 4 2 0 0 -1-2

The first four stationary states of the harmonic oscillator are as follows.

2.4 The Free Particle

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We turn next to what should have been the simplest case of all: the free particle. The time Schrodinger Eq. reads:

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$$-\frac{\hbar^2}{2m}\frac{\mathrm{d}^2\psi}{\mathrm{d}x^2} = E\psi.$$

Let $k \equiv \sqrt{2mE/\hbar}$, we have

$$\Psi_k(x,t) = Ae^{i(kx-\hbar k^2t/2m)}.$$

Remark 2.2. The speed of these waves is:

0

2

$$v_{\rm quantum} = \sqrt{E/2m} = 0.5 v_{\rm classical}$$

And

$$\int_{-\infty}^{\infty} \Psi_k^*(x,t) \Psi_k(x,t) \, \mathrm{d}x = +\infty,$$

which means that a free particle cannot exist in a stationart state.

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Theorem 2.6. The general solution to the time-independent Schrodinger EQ. is still a linear combination of separable solutions:

$$\Psi(x,t) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \phi(k)e^{i(kx-\hbar k^2t/2m)} dk.$$

Now this wave function can be normalized for appropriated $\phi(k)$. We call it a wave packet.

Definition 2.4 (phase velocity and group velocity). For the wave function:

$$\Psi(x,t) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \phi(k) e^{i(kx - \omega t)} \, \mathrm{d}k.$$

We define:

$$v_{\mathrm{phase}} = \frac{\omega}{k}, \ v_{\mathrm{group}} = \frac{\mathrm{d}\omega}{\mathrm{d}k}.$$

3 Formalism

3.1 Gerneralized Statistical Interpretation

First we assume the spectrum of the wave funtion is discrete, we have

$$\langle Q \rangle = \sum_{n'} \sum_{n} c_{n'}^* c_n q_n \langle f_{n'} | f_n \rangle = \sum_{n} |c_n|^2 q_n$$

where q_n is the eigenvalue of operator \hat{Q} and $\Psi(x,t) = \sum_n c_n(t) f_n(x)$. What about momentum?

$$\Phi(p,t) = \frac{1}{\sqrt{2\pi\hbar}} \int_{-\infty}^{\infty} e^{-ipx/\hbar} \Psi(x,t) dx$$

and

$$\Phi(x,t) = \frac{1}{\sqrt{2\pi\hbar}} \int_{-\infty}^{\infty} e^{-ipx/\hbar} \Phi(p,t) dp.$$

3.2 Uncertainty Principle

Theorem 3.1 (generalized uncertainty principle).

$$\sigma_A^2 \sigma_B^2 \ge \left(\frac{1}{2i} \left\langle \left[\hat{A}, \hat{B}\right] \right\rangle \right)^2.$$

How to interpret Δt ?

Definition 3.1.

$$\Delta t \equiv \frac{\sigma_Q}{|\operatorname{d}\langle Q\rangle/\operatorname{d}t|},$$

where

$$\frac{\mathrm{d}\left\langle Q\right\rangle }{\mathrm{d}t}=\frac{i}{\hbar}\left\langle \left[\hat{H},\hat{Q}\right]\right\rangle +\left\langle \frac{\partial\hat{Q}}{\partial t}\right\rangle .$$

I recommend you to learn Hilbert space and Dirac notation.

4 Quantum Mechanics in Three Dimensions

4.1 The schrodinger Equation

The generalization oto three dimensions is straitforward.

$$i\hbar\frac{\partial\Psi}{\partial t} = -\frac{\hbar^2}{2m}\nabla^2\Psi + V\Psi$$

where

$$\nabla^2 \equiv \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}$$

is the **Laplacian**. Also the normalization conditions reads $\int \Psi d^3 \mathbf{r} = 1$. If V is independent of time, there will be a complete set of stationary states

$$\Psi_n(\mathbf{r},t) = \psi_n(\mathbf{r})e^{-iE_nt/\hbar}$$

Now we adopt spherical coordinates

Lemma 4.1 (Laplacian in spherical coordinates).

$$\nabla^2 = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \left(\frac{\partial^2}{\partial \phi^2} \right).$$

If $\Psi = R(r)Y(\theta, \phi)$ and $Y = \Theta(\theta)\Phi(\phi)$, we could separate r, θ and ϕ into three equations with important separation constants.

4.1.1 The angular Equation

The ϕ equation is easy

$$\frac{\mathrm{d}^2 \Phi}{\mathrm{d}\phi^2} = -m^2 \Phi \implies \Phi = e^{im\phi}.$$

When ϕ advances by 2π , we return to the same point in space, so it is natural to require that $\Phi(\phi+2\pi) = \Phi(\phi)$. From this it follows that m must be an integer:

$$m = 0, \pm 1, \pm, 2, \cdots$$

The θ equation reads

$$\sin\theta \frac{\mathrm{d}}{\mathrm{d}\theta} \left(\sin\theta \frac{\mathrm{d}\Theta}{\mathrm{d}\theta}\right) + \left[l(l+1)\sin^2\theta - m^2\right]\Theta = 0.$$

Lemma 4.2 (Legendre function). The solution of Θ is

$$\Theta(\theta) = AP_l^m(\cos\theta).$$

where

$$P_l^m(x) \triangleq (-1)^m (1-x^2)^{m/2} \left(\frac{\mathrm{d}}{\mathrm{d}x}\right)^m P_l(x)$$

is the associated Legendre function, defined by the Rodrigues formula

$$P_l(x) \triangleq \frac{1}{2^l l!} \left(\frac{\mathrm{d}}{\mathrm{d}x}\right)^l (x^2 - 1)^l.$$

Remark 4.1. Notice that l must be a non-negative integer, for Rodrigues formula to make sense; moreover, if m > l, we cwill have $P_l^m(x) = 0$. For any given l, then there are 2l + 1 possible values of m:

$$l = 0, 1, 2 \cdots$$
 and $m = -l, -l + 1, \cdots, l - 1, l$.

By normalization condition

$$\int_0^\pi \int_0^{2\pi} \left| Y \right|^2 \sin\theta \, \mathrm{d}\theta \, \mathrm{d}\phi = 1,$$

we deduce that

$$Y_l^m(\theta,\phi) = \sqrt{\frac{2l+1}{4\pi} \frac{(l-m)!}{(l+m)!}} e^{im\phi} P_l^m(\cos\theta).$$

4.1.2 The Radial Equation

Theorem 4.1 (Radial equation).

$$-\frac{\hbar^2}{2m}\frac{\mathrm{d}^2 u}{\mathrm{d}r^2} + \bigg[V + \frac{\hbar^2}{2m}\frac{l(l+1)}{r^2}\bigg]u = Eu$$

where $u(r) \equiv rR(r)$.

Remark 4.2 (Effective potential).

$$V_{\text{eff}} = V + \frac{\hbar^2}{2m} \frac{l(l+1)}{r^2}$$

and the latter term is the so-called centrifugal potential.

4.2 The Hydrogen Atom

To be continued.

4.3 Angular Momentum