## EXP-T

# a relativistic multireference coupled cluster program v 1.8.5

User manual

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## 1 Introduction

The EXP-T program package is designed for high-precision modeling of electronic structure of atoms and molecules using the relativistic Fock space coupled cluster method (FS-RCC). Features and program components are listed in Sec. 1.1 and 1.2.

The authors will be grateful for any comments or suggestions:

```
exp-t-program@googlegroups.com
https://groups.google.com/d/forum/exp-t-program
```

#### 1.1 Features of EXP-T

The EXP-T package does not include subroutines for solving the (Dirac-)Hartree-Fock equations and subsequent four-index transformation, so molecular integrals have to be imported from third party electronic structure packages. Currently EXP-T is interfaced to the DIRAC program package [1, 2], thus providing access to the wide variety of Hamiltonians and property operators implemented therein.

Electronic structure models available in EXP-T:

- single-point energy calculations with any point groups and (nearly) all Hamiltonians, implemented in DIRAC (4c-DC, X2Cmmf, two-component relativistic pseudopotential (RPP), scalar-relativistic or non-relativistic);
- ground state energy calculations (the 0h0p sector): CCSD, CCSD(T), CCSDT-n (n=1,2,3), CCSDT models;
- the FS-CCSD method for ionized and excited states is implemented for the 0h1p, 1h0p, 1h1p, 0h2p, 2h0p, 0h3p, 1h2p, 2h1p, 3h0p Fock space sectors;
- FS-CC models for excited states accounting for triples (CCSDT-n (n=1,2,3), CCSDT) are implemented for the 0h1p, 1h0p, 1h1p, 0h2p, 2h0p, 0h3p, 1h2p, 3h0p Fock space sectors.

The summary of all electronic structure models implemented in the EXP-T program system is given in the table:

FS sector	CCSD	CCSD+T(3)	CCSDT-1	CCSDT-2	CCSDT-3	CCSDT
$\overline{0h0p}$	+	+	+	+	+	+
1h0p	+	t	t	t	t	t
0h1p	+	t	t	t	t	t
2h0p	+	_	t	t	t	t
1h1p	+	t	t	t	t	t
0h2p	+	t	t	t	t	t
3h0p	t	_	t	t	t	t
2h1p	t	_	_	_	_	_
1h2p	t	_	_	_	_	t
0h3p	+	t	t	t	t	t

"+" - implemented, "-" - not implemented, "t" - currently being tested

Features for property calculations and analysis of wavefunctions:

• analytic density matrices and property calculations (including  $\mathcal{T}$ -odd operators) for the ground-state CCSD and CCSD(T) models;

- finite-field [3] and finite-order [4] approaches to calculation of transition properties;
- (unrelaxed) analytic density matrix and properties for ground-state CCSD and CCSD(T);
- quasi-diabatization of spin-orbit (SO) coupled states and extraction of spin-orbit interaction matrix elements [5].
- construction of model natural spinors (NS) and natural transition spinors (NTS) [6, 7] for non-trivial Fock space sectors;
- one-electron properties *via* the interface to either DIRAC or the OneProp program by L. V. Skripnikov [8].
- finite-field calculations of diagonal two-electron properties (assuming that integrals are provided).

Features to deal with the intruder state problem:

- "dynamic" shifts of energy denominators [5];
- Padé extrapolation to the zero-shift limit [9];
- intermediate Hamiltonian technique for incomplete main model spaces (IH-IMMS) [10].

Features for high-performance calculations:

- OpenMP parallelization (for shared-memory systems);
- parallel calculations on NVIDIA GPUs using the CUDA platform (to be tested) [11].

#### 1.2 Program components

expt.x Input processing + all CC calculations.

heffman.x Manipulations with effective Hamiltonian matrices, i. e. Padé extrapolations, finite-field transition property calculations, quasi-diabatization of spin-orbit-coupled states.

expt\_transform\_natorb.x

Transformation of model natural spinors to the basis of atomic orbitals.

expt\_diatomic.x

Rovibrational levels of diatomic molecules and calculations of expectation values and transition properties for them.

expt2pam.x A tool for specifying molecular geometry, basis sets and pseudopotentials in a convenient human-readable format; the program automatically generates input files for other quantum chemistry packages.

expt2spectrum.py

A tool for convenient extraction of calculated information about electronic states and transitions between them from EXP-T output files.

#### 1.3 Citation

We kindly ask you to acknowledge any use of the EXP-T program system that results in a published material using the following citation:

A. V. Oleynichenko, A. Zaitsevskii, E. Eliav. Towards high performance relativistic electronic structure modelling: the EXP-T program package. *Commun. Comput. Inf. Sci.* 1331, 375 (2020). doi: 10.1007/978-3-030-64616-5 33

In addition, credits must be given to the corresponding papers which describe the implementation and the underlying methodological developments:

- models including triple excitations in the low Fock space sectors: [12];
- Fock-space CC for the three-particle sector (0h3p): [13, 14];
- finite-field technique transition property calculations (general): [3];
- finite-field technique for transition matrix elements involving the  $0h0p \rightarrow 1h1p$  excitations: [15];
- spin-orbit coupling calculations: [5];
- dynamic shifts of energy denominators (real): [5];
- dynamic shifts of energy denominators (imaginary, or their real simulation): [16];
- Padé extrapolation of a series of effective Hamiltonians to the zero-shift limit: [9];
- Intermediate Hamiltonian technique for incomplete main model spaces (IH-IMMS): [10];
- finite-order method of transition matrix elements calculations: [4, 17].

#### 1.4 Credits

EXP-T is based on ideas and design solutions of the suite of FS-CC programs written by Ephraim Eliav and Uzi Kaldor [18, 19].

#### 2 Installation from the source code

## 2.1 Compiling EXP-T

EXP-T is currently oriented at Unix-like operating systems.

Tools required for compiling the EXP-T package from the source code:

• C and Fortran compilers. Currently supported compiler systems are

```
gcc/gfortran GNU compilers

icc/ifort Intel compilers

clang a C language family frontend for LLVM (for the C part)
```

• cmake (version 3.0.2 or higher) and make utilities:

```
https://cmake.org/
https://www.gnu.org/software/make/
```

#### Optional dependencies:

- Implementation of the BLAS/LAPACK linear algebra libraries (OpenBLAS [20] or Intel MKL [21] are recommended);
- NVIDIA drivers and CUDA Toolkit are required to perform parallel calculations on a GPU:

```
https://www.nvidia.com/Download/index.aspx?lang=en-ushttps://developer.nvidia.com/cuda-toolkit
```

Download the EXP-T source code as a \*.tar.xz file and unpack it:

```
tar xvf expt-X.Y.Z.tar.xz
```

Go to the EXP-T home directory, create the build directory and make it the current working dir:

```
mkdir build
cd build
Compilation (Intel compilers):
CC=icc FC=ifort cmake ..
make
Compilation (GNU compilers):
CC=gcc FC=gfortran cmake ..
make
```

We recommend specifying compilers explicitly (since the default C compiler may be other than gcc).

The cmake utility will try to locate BLAS/LAPACK libraries on your computer and link them to the project. If there are no pre-installed libraries on your machine, the internal OpenBLAS will be compiled and linked to EXP-T.

When using Intel MKL together with GNU compilers, you will probably need to explicitly specify the path to the MKL home directory by setting the environment variable MKLROOT. For example:

```
export MKLROOT=/opt/intel/mkl
```

If compilation was successful, the executable file expt.x will appear in the build directory. We strongly recommend that compilation is followed by testing (see Sect 2.4). Do not forget to add the directory containing binary files to the PATH environment variable!

#### 2.1.1 Compiling on macOS

The default C compiler for macOS is clang (https://clang.llvm.org). Thus one can use it instead of gcc (however, the gcc/gfortran pair is still acceptable):

```
CC=clang FC=gfortran cmake .. make
```

Note that by default gcc is just an alias for clang on macOS systems. We strongly recommend not to use GNU compilers installed with Homebrew. Instead, download the GNU compilers suitable for your version of macOS from:

```
http://hpc.sourceforge.net/
```

We recommend to use OpenBLAS attached to the EXP-T as a BLAS/LAPACK implementation.

## 2.2 Compiling with CUDA

Compilation of EXP-T does not require the user to have a CUDA-compatible GPU device. If the CUDA toolkit is missing, the CUDA code will simply be excluded from the compilation process.

EXP-T is adapted for parallel calculations on NVIDIA GPUs (the CUDA technology [11]). Notes:

- 1. At the moment, the OpenMP+CUDA hybrid model as well as multi-GPU support is not implemented. Only one GPU can be used for calculations.
- 2. non-professional GPUs for gaming are not suitable for CUDA calculations (except for Kepler cards).
- 3. CUDA code in EXP-T was tested only for CUDA 9.1 and 10.0.
- 4. Only Intel and GNU compilers were tested for compatibility with CUDA.

To compile EXP-T with CUDA:

- 1. Make sure your graphics card supports CUDA. CUDA-compatible GPUs are listed here: https://developer.nvidia.com/cuda-gpus
- 2. Install proprietary NVIDIA drivers (not the nouveau driver!). https://www.nvidia.com/Download/index.aspx?lang=en-us. For Linux you can find drivers in your distribution's repository.
- 3. Install CUDA toolkit https://developer.nvidia.com/cuda-downloads or from distribution's repository.
- 4. Make sure that all drivers and CUDA toolkit are installed and work correctly. Compile the test suite in the samples directory of the CUDA home directory and run some of them. The most important are the tests employing the CUBLAS linear algebra library.
- 5. Compile EXP-T as described in Sec. 2.1. All CUDA tools will be located automatically by CMake.
- 6. Make sure that the results produced by EXP-T do not depend on the use of CUDA.

## 2.3 Troubleshooting

This section describes the most common problems which can occur during the compilation process. In case of any questions or problems, please don't hesitate to contact us. Don't forget to attach the output file (in case of problems with calculations) or the file with the output of the make VERBOSE=1 command (in case of problems with compilation).

#### Google Groups:

https://groups.google.com/d/forum/exp-t-program

#### E-mail:

exp-t-program@googlegroups.com

#### 2.3.1 Outdated version of cmake

We highly recommend using cmake version 3.0.2 or higher, otherwise, problems may occur when building the internal OpenBLAS. The following command will show what version of cmake is installed:

cmake --version

If the cmake version is lower than 3.0.2, download the binary distribution from their official website https://cmake.org/download/. After installation do not forget to add the directory containing the cmake binary files to the PATH environment variable

Installation of CMake binary files does not require root privileges.

#### 2.3.2 Segmentation fault in heffman.x

The known problem is a possible segmentation fault inside the heffman.x program. It was discovered for the cases when EXP-T is compiled with the internal OpenBLAS and gfortran compilers (version 7.4 and 7.5). The segmentation fault occurs inside the zheev subroutine (diagonalization of a complex Hermitian matrix).

If this issue occurs, we recommend using a different version of the compiler or the BLAS/LAPACK implementation.

## 2.4 Testing

The test suite is located in the test directory. The testing system is written in the Python programming language (must be pre-installed on your machine). Either Python 2 or Python 3 can be used.

In order to run the test suite, run the following command:

make test

## 3 Running EXP-T

See Section 6 for a description of an additional utitily programs included in the EXP-T package.

A typical coupled cluster calculation consists of two sequential steps:

- 1. SCF and integral transformation within DIRAC (Section 3.1);
- 2. CC calculation within EXP-T (Section 3.2).

### 3.1 Precomputing spinors and molecular integrals: the DIRAC step

The interface to DIRAC was tested only for the DIRAC17, DIRAC18 and DIRAC19 releases.

DIRAC stores transformed molecular integrals in the following binary files:

MRCONEE information about SCF calculation, symmetry, occupation numbers, one-electron integrals;

MDCINT symmetry-unique nonzero two-electron integrals;

MDPROP integrals of one-electron property operators.

Run DIRAC and save the files containing the transformed integrals to the working directory using the --get option:

```
pam --noarch --mol=<mol-file> --inp=<inp-file> --get="MRCONEE MDCINT MDPROP"
```

If DIRAC starts in the parallel mode (MPI), it is necessary to choose the "scheme 4" transformation algorithm. By default, DIRAC uses the "scheme 6" algorithm which produces multiple MDCINT files [22] and currently EXP-T requires that all integrals are stored in a single file. In order to accomplish this, add the following lines to the DIRAC input file:

```
**MOLTRA
.SCHEME
4
```

To calculate properties, transition moments and intensities without resorting to the finite-field scheme property integrals transformed to the basis of molecular spinors are also required. We recommend explicitly asking DIRAC to calculate and transform these integrals by adding the .PRPTRA keyword together with the list of the required property operators to DIRAC input files:

```
**MOLTRA
.PRPTRA
.....
**PROPERTIES
.DIPOLE
```

If the molecular-mean-field X2C Hamiltonian (X2Cmmf) was used at the DIRAC stage, one **must** specify the **x2cmmf** flag in the EXP-T input file! (see also page 25).

## 3.2 Coupled cluster calculation: the EXP-T step

Run an EXP-T job using the following command:

```
expt.x [options] <input-file>
```

The EXP-T output (including error messages) is flushed to the UNIX standard output (stdout). Use the output redirection operator ">" in case of time-consuming calculations:

```
nohup expt.x <input-file> > <output-file> &
```

Examples of DIRAC and EXP-T input files are given in Section 5.

## 3.3 Command-line arguments

Use the expt.x --help command to print a list of available command-line arguments:

Usage: expt.x [OPTION...] <input-file>
expt -- A Fock-Space Multireference Relativistic Coupled-Cluster Program

-n, --no-clean Do not clean scratch directory on exit

(use this option to keep cluster amplitudes etc)  $% \frac{1}{2}\left( \frac{1}{2}\right) =\frac{1}{2}\left( \frac{1}{2}\right) +\frac{1}{2}\left( \frac{1}{2}\right)$ 

-s, --scratch=PATH Path to scratch directory (default: ./scratch)

-?, --help Print this help list and exit

--usage Print a short usage message and exit

-V, --version Print program version and exit

Mandatory or optional arguments to long options are also mandatory or optional for any corresponding short options.

Please report bugs to <exp-t-program@googlegroups.com>.

#### Input files syntax and keywords 4

See Section 6 for the description of additional utitily programs included in the EXP-T package.

The input file format may change in future versions.

The EXP-T input file contains a list of instructions. Each instruction is placed on a separate line and consists of a keyword and a list of its arguments. The order of instructions in the input file can be arbitrary. Single-line comments start with '#' and continue until the end of the line. Input is case insensitive. EXP-T input files for typical calculations are given and discussed in details in Section 5. Molecular geometry, basis sets and pseudopotentials are specified only once in DIRAC input files at the molecular integrals calculation stage (not in EXP-T input files).

The detailed description of the keywords used in the EXP-T input file is provided below. For convenience, keywords are grouped into several sections.

#### General settings

```
optional comment.
title
           Syntax: title <quoted-string>
           Default: no title
           Example: title "my first ccsd calculation"
           complex arithmetic will be used throughout the calculation.
arith
           Default: arithmetic depends on molecular symmetry (see [23] for details).
           Syntax: arith complex
           print level.
print
           Syntax: print (low || medium || high || debug || <quoted-string>)
           where <quoted-string> can be:
           "eff config"
                              analysis of electronic states in terms of effective configurations of active
                              spinors;
           "model vectors"
                              expansions of model vectors in terms of model determinants;
           "model space"
                              detailed information about model space determinants.
           One can also combine several print instructions.
           Default: print medium
           Example:
           print low
           print "eff config"
```

degen\_thresh energy levels are considered as degenerate if the energy gap between them is lower than degen\_thresh (in a. u.). It is recommended to decrease the threshold in case of finite-difference calculations of properties.

> Syntax: degen\_thresh <real thresh> [au | ev | cm | mhz] where allowed units are:

```
au atomic units (hartree; by default);

ev electronvolts (eV);

cm wavenumbers (cm<sup>-1</sup>);

mhz megahertz (MHz).

Default: degen_thresh 1e-8
(approximately 0.002 cm<sup>-1</sup>)
```

spinor\_labels manual assignment of symbolic names (labels) of specific atomic or molecular spinors (or spin-orbitals). Symbolic name is printed together with the symmetry of a spinor and its one-electron energy. This option can be useful to study composition of electronic states.

Syntax:

#### Model

sector

specifies the target Fock space sector. To the moment the 0h0p, 0h1p, 1h0p, 1h1p, 0h2p, 2h0p, 3h0p, 0h3p, 1h2p and 2h1p FS sectors are implemented (h - holes, p - particles).

Syntax: sector <integer H>h<integer P>p

Default: sector OhOp (vacuum)

The code for the 1h2p and 2h1p sectors is experimental and is not well-tested!

model

CC model (approximation to the cluster operator).

```
Syntax: model ( ccs || ccd || ccsd || ccsd(t) || ccsd+t(3) || ccsdt-1 || ccsdt-2 || ccsdt-3 || ccsdt )
```

Default: model ccsd

Notes:

- Keywords ccsd(t) and ccsd+t(3) are equivalent.
- The codes implementing the CCSD+T(3), CCSDT-n and CCSDT models in the non-trivial sectors and all CC models in the 0h3p, 3h0p, 1h2p and 2h1p sectors are experimental (currently being tested), to be used with care.

- The current implementation of the CCSD(T) model for the 0h0p sector is not very computationally efficient.
- The CCS and CCD models are implemented simply by setting all  $T_2$  (CCS) or  $T_1$  (CCD) amplitudes to zero, this does not make computation faster, and exists primarily for the purposes of testing and debugging.

occ occupation numbers of individual spinors in a vacuum state. This keyword can be used to setup calculations with a "high-spin" vacuum state [24]. This keyword is alternative to the occ\_irreps keyword (see below).

Default: occupation numbers are read from files containing transformed molecular integrals

Syntax: occ <list of the 0 and 1 digits separated by spaces>

Example: 5 electrons and 10 spinors:

occ 1 1 0 0 0 1 1 1 0 0

CC calculations with high-spin vacuum states were tested for the 0h0p FS sector only.

It is rather convenient to use the dirac\_mointegral\_export routine [25] from the DIRAC package to generate string with occupation numbers.

occ\_irreps specification of a vacuum state using total numbers of electrons in each irreducible representation. Irrep occupation numbers are ignored if occupation numbers of individual spinors are given (by the occ keyword, see above). Names of irreducible representations adopted in EXP-T for some abelian groups are listed in Appendix D.

Syntax: occ\_irreps <pairs irrep:occ>

Example: occupation numbers for the ground triplet state of the  $O_2$  molecule (the  $\overline{D_{\infty h}}$  point group):

```
# for gerade irreps:
occ_irreps [1/2g+]:3 [1/2g-]:2 [3/2g+]:0 [3/2g-]:1
# for ungerade irreps:
occ_irreps [1/2u+]:2 [1/2u-]:2 [3/2u+]:1 [3/2u-]:1
```

Default: occupation numbers are read from MO integrals files.

## Active space

nacth/nactp active (i. e. valence) space specification (by the overall number of active spinors, regardless of their symmetry). This method of specifying an active space is recommended for non-experts.

- nacth number of highest occupied spinors considered as active (active holes)
- nactp number of lowest virtual spinors considered as active (active particles)

It is convenient to specify an active space using the nacth/nactp keywords when exploring electronic states of molecules in a range of geometries (potential energy surfaces, transition moment functions etc). Using the nacth/nactp keywords one can specify either the overall number of active spinors (holes or particles) or the number of active spinors in each irreducible representation.

```
nacth <pairs irrep:number-of-spinors>
nactp <pairs irrep:number-of-spinors>
Default: nacth 0
    nactp 0
    (no active spinors).
```

#### Example:

Let us consider low-lying electronic states of the La atom (see [14] for a discussion). Those are dominated by the  $6s^n5d^m$  configurations, generating even electronic states. So it seems reasonable to include the  $6s_{1/2}$ ,  $5d_{3/2}$  and  $5d_{5/2}$  atomic spinors of the La<sup>3+</sup> cation into the active space and obtain target electronic states in the 0h3p Fock space sector. In order to make the input file easy to read, one can use two separate nactp directives, one for the "negative" projections of  $m_i$ , and one for the "positive" ones:

```
nactp [1/2g+]:3 [3/2g+]:2 [5/2g+]:1 nactp [1/2g-]:3 [3/2g-]:2 [5/2g-]:1
```

One can further include the 6p and 4f spinors of  $La^{3+}$  to the active space to extend the manifold of achievable electronic states:

```
nactp [1/2u+]:4 [3/2u+]:3 [5/2u+]:2 [7/2u+]:1 nactp [1/2u-]:4 [3/2u-]:3 [5/2u-]:2 [7/2u-]:1
```

active

specifies the space of active (i. e. valence) spinors according to the energy criterion. Only spinors with one-electron energies lying in the range  $\varepsilon_i \in [\varepsilon_{min}; \varepsilon_{max}]$  will be considered as the active ones (regardless of their symmetry). Specifying an active space using the active energy instruction is convenient in atomic calculations.

```
Syntax: active energy <real eps_min> <real eps_max>
```

Default: no active spinors

Example: active energy -1.0 5.0

## Intermediate Hamiltonian for incomplete main model spaces (IH-IMMS)

Intermediate Hamiltonian technique formulated for incomplete main model spaces (IH-IMMS) can be used to achieve convergence of iterative solution of CC equations. The first paper [10] on the IH-IMMS method concerns only the special case of atomic systems. However, a structure of incomplete model spaces for molecules is typically simpler than for the special case of atoms; recent molecular applications with the detailed discussions on IH-IMMS for molecules can be found in [4] (the ThO molecule) and [26] (the AcF molecule).

The basic idea of the method is to split the whole model space into the "main" and "intermediate" subspaces. Slater determinants  $|\Phi_M\rangle$  which span the main model space must be chosen to describe the bulk of wavefunctions of target electronic states. FS-CC equations for amplitudes corresponding to excitations from main determinants remain unchanged. In contrast, equations for amplitudes associated with excitations from intermediate ("buffer") determinants  $|\Phi_I\rangle$  are solved with shifted energy denominators to

suppress the intruder state problem. For each cluster amplitude corresponding to the excitation  $I \to K$  from  $|\Phi_I\rangle$  a shift parameter  $S_{IK}$  is calculated as:

$$S_{IK} = \overline{E}_0 - E_I,$$

where  $E_I$  stands for the (zero-order) energy of the  $|\Phi_I\rangle$  model-space determinant and  $\overline{E}_0$  stands for the "frontier" energy.  $S_{IK}$  is used to calculate a shift for energy denominator corresponding to a cluster amplitude using one of the formulas (6)–(9). All model space determinants  $|\Phi_M\rangle$  with energies  $E_M \leq E_0$  are considered as main ones. Determinants  $|\Phi_I\rangle$  with energies  $E_I > E_0$  are treated as intermediate ones.

To set up the IH-IMMS calculation:

- 1. define Fock space sectors for which the IH method will be enabled;
- 2. split the set of active spinors into several subspaces. In the present version of EXP-T only spinor energies can be used to specify these subspaces;
- 3. define model Slater determinants which will be considered as the "main" ones. Sets of "main" determinants are defined via the numbers of particles (holes) in each spinor subspace;
- 4. (optional) define the "frontier" energy  $\overline{E}_0$ . By default, the frontier energy is calculated automatically.

The use of IH-IMMS for the simple case of the carbon atom is considered in details in Section 5.4. By default, IH-IMMS is disabled. IH settings are specified using the multi-line instruction ih\_imms:

The following keywords can be used inside the ih\_imms block:

sectors

a list of Fock space sectors for which the IH-IMMS technique will be applied. This keyword is mandatory.

Syntax: sectors <list-of-sector-labels>

Default: no target FS sectors. Example: sectors 0h1p 0h2p

subspace

defines a subspace of active spinors with energies  $\varepsilon \in [\varepsilon_{\min}, \varepsilon_{\max}]$ . Subspaces are further used to define composition of main model space determinants via the main\_occ keyword. One should use several subspace options in the input file to define more than one subspace of active spinors. This keyword is mandatory. At least one subspace must be specified.

Syntax: subspace energy <real emin> <real emax>

Default: no subspaces are defined.

Example: subspace energy -1.0 -0.5

main\_occ defines a subspace of main model space determinants. The argument of this option contains the number of particles (holes) in the 1st, 2nd, ... subspaces of active spinors. One should use several main\_occ options in the input file to define more than one subspace of main model

space determinants. This keyword is mandatory, at least one subspace of determinants must be specified.

Syntax: main\_occ <list-of-integers>

Default: no main model spaces are defined.

Example: imagine that we need to define model space determinants with 2 electrons on active spinors belonging to the 1st spinor subspace, 1 electron on spinors from the 2nd subspace and 0 electrons on the rest of active spinors (such a model determinant belongs to the 0h3p sector). The proper option will be as follows:

main\_occ 2 1 0

shift\_type specifies a formula for denominator shift to be used for cluster amplitudes associated with the "intermediate" model space determinants. See also the description for the shift\_type keyword (Page 24).

Syntax: shift\_type ( none || real || realimag || imag || taylor )

Default: shift\_type realimag

npower attenuation parameter used in the shifting formula (see Eq. (6)–(9)).

Syntax: npower <int n>

Default: npower 5

scale a scaling factor for the shift parameter  $S_{IK}$ :

$$S_{IK} \rightarrow S'_{IK} = \alpha S_{IK}$$

Syntax: scale <real factor>
Default: scale 1.0 (no scaling)

frontier\_energy "frontier" energy  $\overline{E}_0$  used to calculate shift for the given (intermediate) determinant.

Syntax: frontier\_energy <H>h<P>p [<real energy> | lower\_bound | upper\_bound ]

Default: frontier energy is evaluated automatically as the maximum energy of main model space determinants (frontier\_energy <H>h<P>p upper\_bound):

$$\overline{E}_0 = \max_{\Phi_M \in \text{main}} E_M, \qquad E_M = \langle \Phi_M | H_0 | \Phi_M \rangle.$$

Example: shifting to the minimum energy of model space determinants:

frontier\_energy Oh2p lower\_bound

For some pathological cases the automatically determined frontier energy is too high resulting in a poor convergence of an IH-IMMS calculation. This problem can be solved by manually defining frontier energy equal to the minimal energy of model space determinants (lower\_bound). In almost all practical cases studied to date answers (excitation energies and properties) were found to be pretty stable with respect to the particular value of the frontier energy.

#### Property calculations

analyt\_prop analytic calculation of coupled cluster expectation values of one-electron property operators (including  $\mathcal{T}$ -odd ones). Matrix elements of property operators must be stored in a DIRAC's MDPROP unformatted file or a text file (see Appendix B).  $\Lambda$ -equations, (unrelaxed) analytic density matrix and properties are currently available only for the 0h0p sector (vacuum state) and the CCSD and CCSD(T) models (see refs. [27, 28] for the theoretical background).

Syntax: analyt\_prop property-name>

Default: disabled.

Example: calculation of x,y,z-components of the dipole moment.

analyt\_prop XDIPLEN
analyt\_prop YDIPLEN
analyt\_prop ZDIPLEN

density analytic calculation of one-particle density matrix in the 0h0p sector (without contraction with a property operator matrix). Prints occupation numbers of CC natural orbitals. Analytic density matrices are currently implemented for the CCSD and CCSD(T) models only.

Syntax: density OhOp

Default: disabled.

hermit perform symmetric orthogonalization of model vectors. Orthogonalized model vectors are used only to calculate properties and approximate density matrices. Matrices of effective Hamiltonians flushed to HEFF files (see Section 6.1) remain non-Hermitian.

Syntax: hermit

Default: "hermitization" is disabled.

enable calculations of expectation values of one-electron property operators and their transition (off-diagonal) matrix elements (this option can be used for arbitrary one-electron operators, not only dipole moments). Property operator matrix elements transformed to the basis of molecular spinors are to be loaded from the DIRAC's MDPROP unformatted file. To avoid huge outputs, use options nroots or roots\_cutoff to restrict the set of target electronic states (or, alternatively, the expt\_spectrum.py utility to conveniently extract data for the specific transitions, see Section 6.4). Note that the currently available maximum truncation level is quadratic in cluster amplitudes, which might be not enough for systems with strong dynamic correlations (and thus relatively large cluster amplitudes).

The details on the finite-order method of transition matrix element calculations are given in [4]. The implementation of its connected version together with the first benchmark for transition dipole matrix elements is reported in [17]. Within this approach a property matrix element  $O_{if}$  for a pair of electronic states  $\psi_i$  and  $\psi_f$  is calculated via the relation [29] (scheme non-hermitian):

$$O_{if} = \langle \tilde{\psi}_i^{\perp \perp} | (\Omega^{\dagger} \Omega)_{closed} \Omega^{\dagger} O \Omega | \tilde{\psi}_f \rangle \cdot N_i N_f^{-1}, \tag{1}$$

where  $\tilde{\psi}_i^{\perp\perp}$  stands for the left model vector and the normalization factors are defined as

$$N_i = \langle \psi_i | \psi_i \rangle^{1/2} = \langle \tilde{\psi}_i | \Omega^{\dagger} \Omega | \tilde{\psi}_i \rangle^{1/2} \tag{2}$$

(the same for  $N_f$ ). This definition of the effective property operator leads to the non-Hermitian property matrix,  $O_{if} \neq O_{fi}^*$ . Any hermitization procedure can be further applied to restore

the hermiticity. The working equation contains disconnected diagrams, thus it can suffer from the lack of size-consistency.

There is also an alternative definition (scheme hermitian):

$$O_{if} = \langle \tilde{\psi}_i | \Omega^{\dagger} O \Omega | \tilde{\psi}_f \rangle \cdot N_i^{-1} N_f^{-1}$$
(3)

which is inherently Hermitian [29], but also contains disconnected diagrams.

Finally, the connected 2nd order expression for the non-hermitian effective property operator can be derived [17] (scheme connected):

$$O_{if} = \langle \tilde{\psi}_i^{\perp \perp} | \tilde{O} | \tilde{\psi}_f \rangle \cdot N_i N_f^{-1} \tag{4}$$

$$\tilde{O} \approx \left( O + T^{\dagger}O + OT + \frac{\{(T^{\dagger})^2\}}{2}O + T^{\dagger}OT + O\frac{\{T^2\}}{2} - (T^{\dagger}T)_{cl}O \right)_{closed,conn}. \tag{5}$$

Expressions  $\Omega^{\dagger}\Omega$  and  $\Omega^{\dagger}O\Omega$  used in Eqs. (3) – (5) are evaluated by direct substitution of the FS-CC Ansatz  $\Omega = \{e^T\}$  and have to be artificially truncated at some power of T (the truncation level is specified by the approx keyword). It is natural to use the same level of truncation for both these terms. Finally, the transformation properties with respect to the molecular point group operations of each specific property operator should be explicitly given (the sym keyword).

Syntax: mdprop "roperty-name>" [sym <irrep>] [approx <n> <m>] [scheme hermitian | non-hermitian | connected]

```
sym symmetry of the property operator under consideration;
```

approx truncation level for the numerator  $\Omega^{\dagger}O\Omega$  and the metric operator  $\Omega^{\dagger}\Omega$ :

0 for the full neglect of terms containing cluster amplitudes,

1 for the linear approximation,

2 for the quadratic approximation (recommended option);

scheme particular form of the effective operator to be used:

non-hermitian – Eq. (1),

hermitian – Eq. (3),

connected – Eqs. (4) and (5) (recommended option).

Default: disabled.

Example: calculation of the whole set of possible dipole transition moments for a linear diatomic heteronuclear molecule (the  $C_{\infty v}$  point group). Terms which are quadratic in cluster amplitudes are included (approx 2 2). The connected formulation is used. The list of instructions will be as follows:

```
mdprop "XDIPLEN" sym 1u+ approx 2 2 scheme connected mdprop "XDIPLEN" sym 1u- approx 2 2 scheme connected mdprop "YDIPLEN" sym 1u+ approx 2 2 scheme connected mdprop "YDIPLEN" sym 1u- approx 2 2 scheme connected mdprop "ZDIPLEN" sym 0u approx 2 2 scheme connected
```

Matrix elements of some operators calculated by DIRAC can be purely imaginary even if the molecular point group allows one to use real arithmetics. If matrix elements of such an operator are needed, it is recommended to solve amplitude equations in real arithmetics, then switch

on the complex one using the arith complex instruction, restart the task from converged amplitudes and calculate property matrix elements of interest. For this reason we recommend to thoroughly check the diagnostics in the EXP-T output.

The code for properties is currently experimental; we kindly ask you to report all failures and obviously incorrect results.

mstdm

enable calculations of the model-space (MS) approximations to transition dipole moments (wavefunctions are represented by model vectors) [30, 31, 15]). To avoid huge outputs, use options nroots or roots\_cutoff to restrict the set of target roots.

This option is equivalent to the mdprop instruction specified for all three directions of the dipole moment operator with the "approx 0 0" and "scheme non-hermitian" parameters.

Note that left model vectors are used in bra:

$$d_{if,\eta} = \langle \tilde{\psi}_i^{\perp \perp} | d_{\eta} | \tilde{\psi}_f \rangle$$
  $\eta = x, y, z$ 

and if the "hermitization" of the effective Hamiltonian matrix (specified by the hermit keyword) is disabled, the TDM matrix will be (normally slightly) non-Hermitian  $(|d_{if}| \neq |d_{fi}|)$ .

This approach typically overestimates transition moments by 30-50% (see, for example, tests in [31, 32, 15]), hence we recommend using it only for semiquantitative estimates, e. g. the detection of the most intensive transitions and discerning between  $0^+$  and  $0^-$  states in diatomic molecules (such E1-transitions are strictly forbidden).

Syntax: mstdm

Default: disabled.

txtprop

enable calculations of expectation values and transition matrix elements of one-electron property operators whose matrix elements are specified in a text file. Property operator matrix elements are loaded from two separate formatted files. To avoid huge outputs, use options nroots and roots\_cutoff to restrict the set of target roots. Other options are the same as for the mdprop keyword.

Syntax: txtprop <string file\_re> <string file\_im> [sym <irrep>] [approx <n> <m>]
[scheme hermitian | non-hermitian | connected]

Default: disabled.

Example: atomic calculation of matrix elements of the electric field gradient (EFG):

txtprop efg\_re\_part.txt efg\_im\_part.txt sym 0g approx 2 2 scheme connected

(where the efg\_re\_part.txt and efg\_im\_part.txt files contain the real and imaginary parts of the EFG operator matrix, respectively).

The code for properties is currently experimental; we kindly ask you to report all failures and obviously incorrect results.

natorb

calculation of model natural spinors (NS) and transition spinors (NTS) [6, 7].

Note: (transition) density matrices  $d_{pq}^{IJ} = \langle \tilde{\psi}_I^{\perp \perp} | a_p^{\dagger} a_q | \tilde{\psi}_J \rangle$  are constructed with model vectors instead of true wavefunctions.

Syntax: natorb [<rep>:<state> | <rep1:state1>-<rep2:state2>] ...

Electronic states for which NS and NTS should be calculated are specified by the <rep>:<state> pair, where:

<state> sequential number of the state in this irreducible representation (1, 2, ...).

A pair of electronic states separated by a hyphen shows that NTS between these two states are to be calculated. You can ask EXP-T to calculate NS and NTS for arbitrary number of states (or pairs of states).

Default: calculation of NS and NTS is disabled.

Examples:

- NS for the ground  $(^2\Sigma^+_{1/2})$  and the first excited  $(^2\Sigma^+_{1/2})$  states of the RaCl molecule [33]: natorb [1/2+]:1 [1/2+]:2
- NTS for the  $(1)0^+ (2)0^+$  transition in the RbCs molecule [34]:

natorb [0]:1-[0]:3

(note that in the latter example the [0]:2 state is excluded, since it is the  $(1)0^-$  state which cannot be accessed by the direct transition from the ground state).

#### Iterative solution and convergence

conv specifies the requested convergence threshold (in amplitudes):

$$\left| |t_K^{(n+1)}| - |t_K^{(n)}| \right| < \text{conv} \quad \forall K \quad \Rightarrow \text{converged}$$

The convergence threshold must be made more stringent for finite-difference calculations of expectation values of property operators.

Syntax: conv <real thresh>

Default: 1e-9.

crop

enables the CROP extrapolation technique to accelerate convergence [35, 36]. It is recommended to use the least possible number of last steps used for extrapolation: crop 3.

Syntax: crop [ ( off || triples || <integer dim> ) ]

Disable CROP: crop off

Enable CROP for triples: crop triples

Default: CROP technique is disabled.

It is highly recommended to enable CROP for triples if the CCSDT-n or CCSDT models are employed.

damping enable damping in the given FS sector. Amplitudes obtained at the n-th step will be mixed with amplitudes from the (n-1)-th step using the damping factor  $\alpha$ :

$$t_K^{(n)'} = \alpha t_K^{(n-1)} + (1 - \alpha) t_K^{(n)}$$

Damping can aid the convergence of equations in some non-pathological cases (when the lack of convergence is caused by amplitude oscillations rather than intruder states). Note that in the other cases damping can significantly slow down the convergence.

Syntax: damping <H>h<P>p <integer last\_step> <real factor>

<H>h<P>pFS sector in which these damping parameters will be applied;

last\_step at which iteration damping will be disabled;

damping factor  $\alpha$  (can be negative). factor

Default: disabled in all sectors.

Example: damping 0h2p 100 0.5

enables the DIIS extrapolation technique [37, 38]. diis

Syntax: diis [ ( off || triples || <integer dim> ) ]

Disable DIIS: diis off

Enable DIIS for triples: diis triples

Default: diis 10

(DIIS is enabled for singles and doubles, disabled for triples).

It is highly recommended to enable DIIS for triples if the CCSDT-n or CCSDT models are employed.

div\_thresh specifies the divergence threshold (in amplitudes). Solution of amplitude equation is diverged if for some  $t_K$  holds:

$$\left||t_K^{(n+1)}| - |t_K^{(n)}|\right| > \texttt{div\_thresh} \quad \Rightarrow \text{diverged}$$

May be useful if few initial steps of the iterative solution are divergent. Note that this may also occur even in non-pathological cases.

Syntax: div\_thresh <real thresh>

Default: 1e+3.

indicates that cluster amplitudes must be flushed to disk every N iterations. flush

> Normally cluster amplitudes are written to disk only if the convergence was reached. Use the flush instruction to simplify the restarting of your tasks. This option can be useful for very time-consuming calculations with poor convergence. It also allows one to change the convergence acceleration strategy if needed.

Syntax: flush <integer N> iter

Default: no flushing. Example: flush 5 iter

nroots

specifies the number of roots (electronic states) of interest in each irreducible representation (in the target Fock space sector). Model vectors analysis will be performed only for the nroots lowest states.

Syntax: nroots <pairs irrep:number-of-roots>

Default: all roots will be printed and analyzed.

Example: consider the electronic states of the RbCs molecule, corresponding to the first three dissociation limits (see, for example, [5, 34] and references therein). In the Hund's case a these electronic states are  $X^1\Sigma^+$ ,  $a^3\Sigma^+$ ,  $b^3\Pi$ ,  $A^1\Sigma^+$ ,  $c^3\Sigma^+$ ,  $B^1\Pi$ . In the Hund's case c these states are classified according the projection of the total electronic angular momentum and can be rearranged as follows:  $(1-3)0^+$ ,  $(1-3)0^-$ , (1-4)1 and (1)2 states. The names of irreps used in EXP-T are given in Appendix D (the point group of RbCs is  $\overline{C_{\infty v}}$ ). All output and analysis of model vectors will be restricted to these 11 states using the following instruction:

```
nroots [0]:6 [1+]:4 [1-]:4 [2+]:1 [2-]:1
```

(note that in DIRAC the  $0^+$  and  $0^-$  states belong to the same irrep denoted as "0").

maxiter

maximum number of iterations allowed to solve the CC equations.

Syntax: maxiter <integer max>

Default: maxiter 50.

reuse

indicates that sorted integrals and cluster amplitudes from the previous calculation should be reused; thus the calculation will be restarted. Cluster amplitudes read from disk will be used as an initial guess. If EXP-T fails to locate the files containing amplitudes, an initial guess will be obtained using MBPT(2).

Syntax: reuse <list-of-arguments>

Possible values of arguments:

```
load one- and two-electron integrals (from a scratch directory);
integrals
1-integrals one-electron integrals only;
2-integrals two-electron integrals only;
amplitudes
              load amplitudes for all FS sectors occurring in the calculation;
0h0p, 0h1p, 1h0p, 2h0p, 1h1p, 0h2p, 3h0p, 2h1p, 1h2p, 0h3p
              load amplitudes for the given Fock space sector only.
```

Default: no reuse

Note: files with converged amplitudes are flushed to the scratch directory after each calculation (see also Appendix C). To prevent the deletion of files with amplitudes or integrals it is necessary to use the --no-clean command-line option (see Section 3.3).

Example: reuse integrals 0h0p 0h1p

In the current version of EXP-T one cannot reuse integrals specific for the given sector. If molecular integrals were not previously sorted for some of the sectors, then all the integrals have to be re-sorted and the "reuse integrals" instruction cannot be used. This issue is planned to be fixed in future releases.

roots\_cutoff specifies the upper energy bound for the electronic states to be analyzed and printed. Model vectors analysis and property calculations will be performed only for the states with energies lower than roots\_cutoff. Allowed units of energy are eV, cm<sup>-1</sup> or Hartrees.

```
Syntax: roots_cutoff <real max_energy> ( ev || cm || au )
```

Default: all roots will be printed and analyzed.

Example: roots\_cutoff 30000 cm

sets the denominator shift parameters in the  $n_h h n_p p$  Fock space sector. shift

Syntax:

```
shift <H>h<P>p <int n> <real shift_T1> <real shift_T2> [<real shift_T3>]
```

attenuation parameter (Eqs. (6) - (9); n

shift\_T1 shift value for the  $T_1^{n_h h}$   $n_p p$  cluster operator (singles). Must be omitted in the 2h0p and 0h2p FS sectors and in all sectors with three quasiparticles. It is recommended to set  $\mathtt{shift\_T1} = 0.0$  in the 1h1p sector, unless you really understand the consequences of using a nonzero value. Note that the  $\mathtt{shift\_T1}$  value in this sector does not affect the resulting energies, it is needed only for property calculations;

shift\_T2 shift value for  $T_2^{n_h h \ n_p p}$  (doubles). Must be omitted for the 3h0p and 0h3p sectors;

shift\_T3 shift value for  $T_3^{n_h h \ n_p p}$  (triples). Can be used only for CC models accounting for triples in a non-perturbative way (CCSDT-n, CCSDT).

Default: n = 3, all shifts = 0.0.

Examples:

shift 0h1p 3 -0.3 -0.6 -0.9

shift 0h1p 3 -0.6

shift 1h1p 4 0.0 -0.5

shift\_type indicates which formula for energy denominator shifts should be used [5, 9, 14].

shift\_type is the same for all FS sectors. If shift\_type is not equal to none, you must specify the shift parameters for at least one FS sector (see keyword shift; except for 0h0p).

Syntax: shift\_type ( none || real || realimag || imag || taylor )

Possible values:

none no shift

real real shift

$$D_K' = D_K + S \left(\frac{S}{D_K + S}\right)^n \tag{6}$$

realimag real simulation of an imaginary shift [16]

$$D_K' = D_K + \frac{S^2}{D_K} \left( \frac{S^2}{D_K^2 + S^2} \right)^n \tag{7}$$

imag imaginary shift [16]

$$D_K' = D_K + iS \left(\frac{|S|}{|D_K + iS|}\right)^n \tag{8}$$

taylor extrapolated intermediate Hamiltonian-like shift [39]

$$D_{K}^{'} = (D_{K} + S) \frac{\left(1 - \frac{S}{D_{K} + S}\right)}{1 - \left(\frac{S}{D_{K} + S}\right)^{n+1}}$$
(9)

Default:

- if there are no shifts specified using the shift instruction: shift\_type none
- otherwise (shifts are specified in some sectors): shift\_type realimag

#### Shifts are never applied in the 0h0p sector!

skip

indicates which sectors will be skipped in the FS-CC calculation. Can be useful in very costly calculations (like CCSDT) in order to avoid recalculation of the first steps in already converged sectors. To be used with the reuse option.

Syntax: skip <list-of-sector-symbols>

Default: disabled.

Example: skip 0h0p 0h1p

## Interfaces

integrals sets paths to files containing transformed molecular integrals.

Syntax: integrals <string "1-el int-s"> <string "2-el int-s">
 [<string "properties int-s">]

Arguments:

• one-electron integrals file;

• two-electron integrals file;

• properties transformed integrals file (optional, is needed only for property calculations).

Default: integrals MRCONEE MDCINT MDPROP

Example: integrals ../MRCONEE-Cinfv ../MDCINT-Cinfv /home/user/MDPROP

oneprop

imports one-electron property integrals from the MDPROP unformatted file. The one-electron operator  $\hat{O}_1$  is added to the Fock operator at the integral sorting stage:

$$\hat{f}' = \hat{f} + \lambda \hat{O}_1$$

where  $\lambda$  is some perturbation parameter. EXP-T does not assume Kramers symmetry, so both  $\mathcal{T}$ -odd and  $\mathcal{T}$ -even operators are allowed.

Syntax: oneprop <real L\_re> <real L\_im> <quoted-string property-name>

 $\texttt{L\_re}$ ,  $\texttt{L\_im}$  real and imaginary parts of the perturbation parameter  $\lambda$ ;

property-name string identifier used to encode the property matrix in the MDPROP file (can be also found in the DIRAC output file).

Default: disabled.

Example:

oneprop 1e-5 0.0 "ZDIPLEN"

(an interaction of a system with an external electric field is added to the Fock operator).

x2cmmf

since DIRAC's unformatted files do not contain any information about the relativistic Hamiltonian used, the x2cmmf flag must be used if the molecular-mean-field X2C Hamiltonian (X2Cmmf) was employed at the DIRAC stage.

Syntax: x2cmmf

#### Memory management

enables compression of all data written to disk. The LZ4 algorithm [40] is employed (it provides the fastest decompression to date). Compression can slow down calculations slightly (depends on the particular harddrive), but the disk space used can be reduced dramatically (up to 3 times).

Syntax: compress

Default: disabled (except for disk\_usage=4)

disk\_usage indicates which data should be stored on the disk. Note that very intensive disk usage will slow down calculations; however, disk usage is hardly avoidable in case of large tasks.

Syntax: disk\_usage <integer mode>

Possible values of <mode>:

- 0 all data is stored in RAM;
- 1 tensors of rank  $\geq 6$  are stored on disk ( $T_3$  amplitudes etc);
- 2 tensors of rank  $\geq 6$  and  $\langle pp||pp\rangle$  (4 virtual indices) two-electron integrals are stored on disk;
- 3 tensors of rank  $\geq 6$ ,  $\langle pp||pp\rangle$  and  $\langle *p||pp\rangle$  (3 virtual indices) two-electron integrals are stored on disk;
- 4 the same as 3, + compression of all data. Is suitable for extremely large tasks only.

Default: disk\_usage 2

 $(\langle pp||pp\rangle$  and tensors of rank  $\geq 6$  are stored on disk).

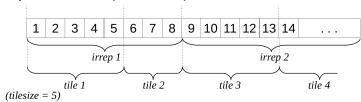
memory specifies the amount of memory that EXP-T can use for the job (just for the *dynamically* allocated space).

Syntax: memory <real size> ( mb || gb )

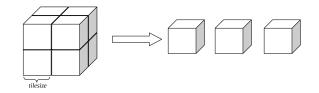
Min value: 10 Mb Default: memory 1 gb

tilesize the whole set of one-electron functions (spinors or spin-orbitals) is divided into blocks: (1) by irrep; (2) then into sub-blocks with size not exceeding tilesize (actually the same as "tiles" in NWChem-TCE):

sequentual numbers of one-electron functions:



The N-dimensional arrays of integrals and amplitudes (tensors) also turn out to be divided into blocks containing  $\sim (\mathtt{tilesize})^N$  elements:



tilesize must be large enough to place at least one N-dimensional array of size (tilesize)<sup>N</sup> in RAM. We recommend using as large tilesize as possible on your machine, otherwise overheads for the memory management can become enormous.

Syntax: tilesize <integer size>

Default: tilesize 100

#### Parallel execution

nthreads number of OpenMP threads to be used for parallel calculations. Note that the scaling with respect to the number of threads is much better for large tasks (> 400 spinors).

Syntax: nthreads <integer n\_omp\_threads>

Default: nthreads 1 (sequential execution).

At the moment parallelization is not used at the integral sorting stage, hence OpenMP can accelerate only the amplitude equations solution stage.

It is recommended to always enable OpenMP parallelization in calculations accounting for triple excitations.

openmp the same as nthreads.

openmp\_algorithm the approach to parallelization. Two parallelization algorithms are implemented in EXP-T:

external suits well for the problems with high granularity of tensors; it is recommended for all point groups except for  $C_1$ .

internal is recommended for the  $C_1$  symmetry only.

Syntax: openmp\_algorithm ( external | internal )

Default: external

enables parallel calculations on an NVIDIA GPU (currently, only one GPU card is supported in calculations). Requires nthreads 1 (the OpenMP+CUDA hybrid model is not implemented yet).

Syntax: cuda

Default: CUDA is disabled.

The CUDA code is outdated, to be used with care. This issue will be fixed in future releases.

## 5 Running typical calculations

The input and output files discussed in this section can be found in the examples directory of the EXP-T release.

## 5.1 Ground state energy with CCSD: the CO molecule

For input/output files see examples/CO\_ccsd.

Consider the calculation of the CCSD correlation energy of the CO molecule in the cc-pVDZ [41] basis set.

```
DIRAC input files:
   (CO.mol)
DIRAC
CO molecule, R = 2.132 bohrs
symmetry C1
    2
С
          0
        6.
               1
          0.0
C
                    0.0
                              0.0
LARGE BASIS cc-pVDZ
        8.
          0.0
                    0.0
                              2.132
LARGE BASIS cc-pVDZ
FINISH
   (moltra.inp)
**DIRAC
.TITLE
CO molecule
.WAVE FUNCTION
.4INDEX
**WAVE FUNCTION
.SCF
*SCF
. CLOSED
14
. ERGCNV
1E-12
**MOLTRA
.ACTIVE
all
*END OF INPUT
```

Use the following command to run DIRAC and obtain transformed molecular integrals:

```
pam --inp=moltra --mol=CO --get="MRCONEE MDCINT" --noarch
```

Note that the --get option is used so that pam exports the transformed molecular integrals (stored in the MRCONEE and MDCINT files) from DIRAC's scratch directory to the working directory. The -noarch option is used to avoid creation of a .tgz archive.

In some cases, it can be convenient to rename the files containing the integrals (this is not mandatory). For example,

```
my MRCONEE MRCONEE-C1
mv MDCINT MDCINT-C1
   Now the EXP-T input file ccsd.inp is set up:
# CO molecule, CCSD/cc-pVDZ
# non-relativistic, symmetry C1
# task title (of type "string")
title "CO/CCSD/cc-pVDZ"
# print level, default medium
print medium
# max number of iterations
maxiter 30
# convergence threshold (by cluster amplitudes)
conv 1e-9
# target Fock space sector
sector OhOp
# CC model: ccsd, ccsdt-1, etc
model ccsd
# import integrals from
integrals MRCONEE-C1 MDCINT-C1
Running EXP-T:
expt.x ccsd.inp | tee ccsd.out
```

The string with the correlation energy can be found at the end of the output. Note that by default EXP-T flushes all output simply to stdout.

```
SCF reference energy = -112.820480227130

CCSD correlation energy = -0.298117912040

Total CCSD energy = -113.118598139170
```

## 5.2 Simple atomic FS-CCSD calculation: the Pb atom

For input/output files see examples/Pb\_fs\_ccsd.

We now turn to simple atomic relativistic FS-CCSD calculations. Since wave functions of most lowlying electronic states of one-electron atoms and atomic ions are effectively single-reference, it is not very interesting to apply the multireference methodology to such cases. In contrast, electronic states of the "two-electron" atoms can hardly be simulated within any other approaches except those based on the multireference wave functions. Consider for example low-lying states of the Pb atom corresponding to the  $6s^26p^2$  configuration. The closed shell ground state  $6s^2$  of Pb<sup>2+</sup> can be considered as the Fermi vacuum, and one should select its 6p-spinors as the active ones (overall 6 active spinors). Pb atom is a purely relativistic system, and one should use either the four-component Hamiltonian or the relativistic pseudopotential (RPP) approach. The latter allows one to incorporate bulk of the Breit interaction [42] and (possibly) QED corrections [43, 44, 10]. Here the semilocal (valence) version of the N. S. Mosyagin's RPP for Pb is used (see [45] and references therein).

Input file for the FS-CCSD calculation will be as follows:

```
sector 0h2p
nactp 6
maxiter 500
memory 110 gb
conv 1e-7
disk_usage 0
nthreads 8
model ccsd
```

After the FS-CCSD calculation is finished, one can compare the calculated FS-CCSD excitation energies of Pb with their experimentally measured counterparts (7819, 10650, 21458 and 29467 cm<sup>-1</sup>, respectively [46]):

Le	evel	Re(eigenvalue)		•	Rel eigv, eV	Rel eigv, cm-1	deg	symmetry
			 ٠	•				
0	1	-0.8258538771			0.0000000000	0.000000	1	0g
@	2	-0.7916687782			0.9302239316	7502.761992	3	0g 1g+ 1g-
@	3	-0.7784232250			1.2906537953	10409.824894	5	0g 1g+ 1g- 2g+ 2g-
@	4	-0.7298818976			2.6115306047	21063.414836	5	0g 1g+ 1g- 2g+ 2g-
0	5	-0.6892150215			3.7181326781	29988.762480	1	0g

One can also find in the output file the calculated values of sequential ionization energies corresponding to the  $Pb^+ \to Pb^{2+}$  and  $Pb^0 \to Pb^{2+}$  processes, respectively:

```
Ionization potential 0h1p \rightarrow 0h0p = ... = 15.0036 \text{ eV} = 121012.35 \text{ cm}^-1 Ionization potential 0h2p \rightarrow 0h1p = ... = 22.4726 \text{ eV} = 181253.98 \text{ cm}^-1
```

The derived first ionization potential of the neutral lead atom would be:

$$IP_1 = E(Pb^+ \to Pb^{2+}) - E(Pb^0 \to Pb^{2+}) = 181254 - 121012 = 60242 \text{ cm}^{-1}$$

(the experimental value is  $59819 \text{ cm}^{-1}$ ).

Note that the obtained results are consistent with an observation that the typical error of the FS-CCSD model is of order 300–400 cm<sup>-1</sup>. The remaining error can be significantly reduced by inclusion of triple excitations [12].

## 5.3 Electronic states and potential energy curves of a diatomic molecule: HgH

For input/output files see examples/HgH\_fs\_ccsd.

Let us consider a very simple (but typical) example of FS-CCSD calculation of a system of electronic states and corresponding potential energy curves of a heavy diatomic molecule of mercury hydride (HgH). Some time ago this molecule was considered as a suitable system for experimental studies of different parity non-conservation effects (see, for example, papers [47, 48, 49, 50, 51, 52] and references therein).

The exhaustive picture of the HgH electronic states was obtained using the MRCI method in [52]; potential energy curves calculated in this paper for 4 lowest-lying electronic states are shown in dashed

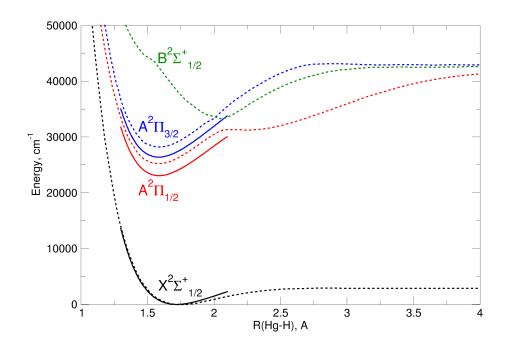


Figure 1: Potential energy curves of low-lying electronic states of the HgH molecule calculated using the FS-CCSD method (solid lines; see text) and multireference configuration interaction [52] (dashed lines). Avoided crossing of the  $A^2\Sigma_{1/2}^+$  and  $B^2\Sigma_{1/2}^+$  states cannot be described within the model space comprising only 6 active spinors.

on Fig. 1. We will try to reproduce three states  $(X^2\Sigma_{1/2}^+, A^2\Pi_{1/2}, A^2\Pi_{3/2})$ , thus the minimal active space describing these target states comprises 6 virtual spinors of HgH<sup>+</sup>. The basis set employed here was borrowed from Ref. [47] and cropped to [4s5d3d] in order to drastically reduce the computational cost; the cc-pVDZ basis set [41] was placed on the hydrogen atom. Thus one shouldn't expect excellent agreement with either experimental data or high-quality calculations reported previously, but all the basic features of the considered problem are still preserved. 60 core electrons of Hg were substituted by the semi-local version of the small-core shape-consistent relativistic pseudopotential of N. S. Mosyagin et al [53].

We first calculate vertical excitation energies of HgH at the internuclear distance r(Hg-H) = 1.74 Å which is close to the experimental one. The first stage implying the relativistic Hartree-Fock calculation and transformation of molecular integrals, is performed using DIRAC:

```
pam --inp=moltra --mol=HgH --noarch --get="MRCONEE MDCINT"
```

Molecular geometry, basis set and RPP are specified in the HgH.mol file, while options for SCF and four-index transformation are given in moltra.inp.

It can be more convenient to use the special NWChem-like format to specify geometry, basis set and RPP data instead of using DIRAC .mol files (see Section 6.3 for details). In this case just convert the HgH.exp file attached to this example to the .mol format using the following commands:

```
expt2pam.x HgH.exp
mv dirac.mol HgH.mol
```

Input file for FS-CCSD calculation using EXP-T will be as follows (ccsd.inp):

```
# Test:
```

- # (1) HgH molecule, FSCC scheme: HgH+ -> HgH
- # (2) hamiltionian: 2-comp gatchina ECP
- # (3) symmetry Cinfv
- # (4) sector (0h,1p)

```
title "Test HgH+ -> HgH/gatchina ECP/Cinfv"
sector 0h1p
nactp 6
conv 1e-9
maxiter 200
diis 100
   Running EXP-T:
expt.x ccsd.inp > ccsd.out
```

Information about electronic states is summarized in the end of an output file and is marked by the '@' symbol in the first position of a line:

```
Level Re(eigenvalue)
                                        Abs energy . . Rel eigv, cm-1 deg symmetry
        -0.2843267762
                            -153.90380203063062936
                                                               0.000000
                                                                           2
                                                                               1/2+ 1/2-
        -0.1742398284
                            -153.79371508284035031
                                                           24161.292284
                                                                           2
                                                                               1/2+ 1/2-
        -0.1588816394
                            -153.77835689385861428
                                                           27532.025149
                                                                               3/2+ 3/2-
```

Vertical excitation energies are given in the column named "Rel eigv, cm-1". Projections of the total angular momentum  $\Omega$  on the z-axis are given in the last column "symmetry". In this particular case there's not much point in comparing vertical excitation energies with their experimental counterparts since equilibrium distances in excited states are significantly displaced with respect to  $r_e$  in the ground state. However, by varying the Hg-H internuclear separation and collecting absolute energies given in a. u. (the "Abs energy" column) one can obtain potential energy curves in the range  $r \in [1.3, 2.1]$  Å (see Fig. 1). Note that for larger r there is an avoided crossing of the  $A^2\Sigma_{1/2}^+$  and  $B^2\Sigma_{1/2}^+$  states. The latter state cannot be obtained within the model space comprising only 6 active spinors considered here.

To extract spectroscopic constants  $(r_e, T_e, \omega_e, ...)$  one should solve the rotational-vibrational problem for each of the curves. It is convenient to use the expt\_diatomic.x program (Section 6.5) or any other software (we also recommend the VIBROT code [54]) to do it. For example, the input file (vib\_rot\_X2Sigma+.inp) for expt\_diatomic.x needed to calculate spectroscopic constants of the  $X^2\Sigma_{1/2}^+$ state will be as follows (the format is self-explanatory):

Running expt\_diatomic.x:

```
expt_diatomic.x vib_rot_X2Sigma+.inp | tee vib_rot_X2Sigma+.out
```

Output file contains analysis of spectroscopic constant in three different approximations, (a) harmonic oscillator, (b) Morse oscillator, (c) exact numerical solution of the radial Schrödinger equation:

> equilibrium parameters

```
e(min) = -153.903855470153 a.u. = -33777991.944687 cm^-1
re = 3.2440 a.u. = 1.7167 angstrom
```

> analysis in the harmonic oscillator + rigid rotor approximation

```
we = 1446.5254 cm<sup>-1</sup>
Be = 5.704258 cm<sup>-1</sup>
De = 0.000355 cm<sup>-1</sup>
```

> analysis in the morse approximation

```
(parameters are estimated at the minimum point only) alpha = 1.2462 bohr^-1
De = 0.025565 a.u. = 5611 cm^-1
```

we =  $1446.53 \text{ cm}^{-1}$ wexe =  $93.23 \text{ cm}^{-1}$ 

vmax = 7

> rovibrational energy levels

	J	v	energy (cm^-1)	< r > (Ang)	< B_v > (cm^-1)	<prop></prop>
0	0	0	730.0496	1.746630	5.570576	0.000000
@	0	1	2249.5311	1.764043	5.560544	0.000000

Potential energy curves and spectroscopic constants can be further refined by:

- using of high-quality basis sets designed specifically for the problem under consideration (using the ANO approach [55], for example);
- using of full (non-local) versions of generalized relativistic pseudopotentials [56];
- extending the model space, possibly involving the intermediate Hamiltonian approach to bypass the intruder state problem. In principle, this can help to describe the avoided crossing  $A^2\Sigma_{1/2}^+ \sim B^2\Sigma_{1/2}^+$ ;
- accounting for triple excitations at some level of theory (even the FS-CCSDT-1 model in the 0h1p sector would be useful [12]).

## 5.4 Intermediate Hamiltonian for incomplete main model spaces: the C atom

For input/output files see examples/C\_atom\_ih-imms.

Intermediate Hamiltonian formulation of the FS-CC method allows one to mitigate intruder state problem. The recent formulation for incomplete main model spaces [10] works in a nearly black-box fashion.

Let us consider low-lying electronic states of the C atom. Using the NIST ASD database [57] we find that for the states below 80000 cm<sup>-1</sup> the following configurations dominate in wave functions:

- $2s^22p^2$ ;
- $2s^12p^3$ ;
- $2s^22p^13s^1$ ;
- $2s^22p^13p^1$ ;
- $2s^22p^13d^1$ ;
- $2s^22p^14s^1$ .

Within the Fock space coupled cluster formalism, electronic states dominated by the  $2s^12p^3$  configuration (e. g. the  $^5S_2^o$  state,  $33735 \text{ cm}^{-1}$ ) can be accessed only in 1h3p or 0h4p sectors. Implementation of the corresponding FS-CC models is not available to date, thus we will not consider these states in the further discussion. In contrast, other states of the carbon atom with closed  $2s^2$  subshell can be easily obtained in the 0h2p Fock space sector provided that the  $2s^2$  state of the  $C^{2+}$  ion is considered as the Fermi vacuum. For the sake of simplicity we will not also consider electronic states involving shells with principal quantum number n=4 (the generalization of the IH-IMMS procedure to such states is straightforward).

At this point we have several useful observations:

- active space should comprise at least 2p, 3s, 3p, 3d spinors of the  $C^{2+}$  ion. These active spinors are needed to describe the manifold of electronic states below 80000 cm<sup>-1</sup>;
- only the 2p shell can be doubly occupied in some target electronic states  $(2s^22p^2)$ ;
- all target electronic states involve configurations with at least one 2p-electron;
- we can expect that electronic states dominated by the  $3s^2$ ,  $3s^13p^1$ , etc configurations will be very high in energy (and even lie well above the IP = 90820 cm<sup>-1</sup>). These states arise inevitably in the 0h2p sector since the "classical" FS-CC is formulated for quasicomplete model spaces. Moreover, these *intruder states* will destroy convergence of amplitude equations;
- the linear span of these "malicious" model determinants can be considered as the *intermediate space*;
- the main model space should include only determinants belonging to the  $2p^2$ ,  $2p^13s^1$ ,  $2p^13p^1$ ,  $2p^13d^1$  configurations. These determinants will be dominating in the model vectors representing target electronic states of the carbon atom.

We start actual calculation with the molecular integral transformation in DIRAC: (C.mol)

```
C atom doubly augmented triple-zeta basis set (d-aug-cc-pVTZ) symmetry Dinfh C \, 1
```

```
1
        6.
C
     0.000000000
                      0.000000000
                                      0.000000000
LARGE EXPLICIT
                  4
                        2
                             2
                                  1
                                        1
# s functions
F 10 2
       8236.00000000
                         0.00053100
                                       -0.00011300
       1235.00000000
                         0.00410800
                                       -0.00087800
        280.80000000
                         0.02108700
                                       -0.00454000
         79.27000000
                         0.08185300
                                       -0.01813300
         25.59000000
                         0.23481700
                                       -0.05576000
          8.99700000
                         0.43440100
                                       -0.12689500
          3.31900000
                         0.34612900
                                       -0.17035200
          0.90590000
                         0.03937800
                                        0.14038200
          0.36430000
                        -0.00898300
                                        0.59868400
          0.12850000
                         0.00238500
                                        0.39538900
F
   4 0
          0.90590000
          0.12850000
          0.04402000
          0.01510000
# p functions
   5
     1
         18.71000000
                         0.01403100
          4.13300000
                         0.08686600
          1.20000000
                         0.29021600
          0.38270000
                         0.50100800
          0.12090000
                         0.34340600
F
   4 0
          0.38270000
          0.12090000
          0.03569000
          0.01050000
 d functions
   4
      0
          1.09700000
          0.31800000
          0.1000000
          0.03140000
# f functions
F
  3 0
          0.76100000
          0.26800000
          0.09440000
FINISH
   (moltra.inp)
**DIRAC
.TITLE
Carbon atom
```

.4INDEX

```
.WAVE FUNCTION
```

- \*\*HAMILTONIAN
- .X2Cmmf
- . GAUNT
- \*\*WAVE FUNCTION
- .SCF
- \*SCF
- .CLOSED

4 0

. ERGCNV

1E-12

\*\*MOLTRA

.SCHEME

4

.ACTIVE

all

\*END OF INPUT

Note that the .X2Cmmf keyword is used, thus it also must be used in the EXP-T input file.

#### Command line:

```
pam --inp=moltra --mol=C --noarch --get="MRCONEE MDCINT MDPROP"
```

Let us now look at the DIRAC output. We find that one-electron energies of our active spinors are (in a. u.):

 $\varepsilon_{2p} \sim -0.86$ 

 $\varepsilon_{3s} \sim -0.36$ 

 $\varepsilon_{3n} \sim -0.28$ 

 $\varepsilon_{3d} \sim -0.23$ 

We split these active spinors into two subsets. Spinors from the first subset can be doubly occupied in the main model space determinants, and spinors from the second subset can be only singly occupied. In some cases it can be useful to define the third subset of spinors which must be unoccupied in all main determinants. This procedure resembles that used to define an active space in the RASSCF method. For the C atom these subsets of active spinors are:

- 1. 2p;
- 2. 3s, 3p, 3d.

In terms of these two subspaces of active spinors the main model space determinants can be defined using the following directives inside the ih\_imms block in the EXP-T input file:

- main\_occ 2 0: determinants in which both electrons occupy spinors from the first subspace, the  $2p^2$  configuration;
- main\_occ 1 1: one electron occupy a spinor from the first subspace (2p), the other one occupy a spinor from the second subspace (3s, 3p or 3d). These determinants correspond to the  $2p^13s^1$ ,  $2p^13p^1$  and  $2p^13d^1$  configurations.

We will rely on the automatic determination of the frontier\_energy parameter (by default, the uppermost energy of main model space determinants is used). Finally, the EXP-T input file will be as follows:

```
# relativistic CC model
sector Oh2p
model ccsd
nactp 24
x2cmmf
# intermediate Hamiltonian parameters.
# spinor subspaces are defined via energy ranges (in a. u.)
ih_imms
  sectors Oh2p
  subspace energy -1.0 -0.5 # 2p spinors
 subspace energy -0.5 -0.2 # 3s, 3p, 3d spinors
                             # 2p^2 determinants
 main_occ 2 0
                             # 2p3s, 2p3p, 2p3d determinants
 main_occ 1 1
end
# memory
memory 4 gb
disk_usage 0
# convergence
maxiter 300
diis 300
conv 1e-9
# print compositions of model vectors in terms of
# effective occupation numbers of active spinors
print "eff config"
```

Let us take a look at the table with calculated energy levels in the 0h2p sector in the EXP-T output file:

Le	evel	Re(eigenvalue)R	el eigv, cm-1	% main	deg	symmetry	
0	1	-1.2955562077	0.000000	100.0	1	0g	
@	2	-1.2954943219	13.582372	100.0	3	0g 1g+ 1g-	
0	3	-1.2953712703	40.589069	100.0	5	0g 1g+ 1g- 2g	g+ 2g-
0	4	-1.2503086362	9930.694080	99.9	5	0g 1g+ 1g- 2g	g+ 2g-
0	5	-1.1923569864	22649.611052	99.8	1	0g	
@	6	-1.0208724411	60286.118415	99.9	1	0u	
@	7	-1.0207910577	60303.980008	99.9	3	0u 1u+ 1u-	
@	8	-1.0206248742	60340.453071	99.9	5	0u 1u+ 1u- 2u	ı+ 2u-
@	9	-1.0089659280	62899.295993	99.9	3	0u 1u+ 1u-	
@	10	-0.9902128248	67015.126416	100.0	3	0g 1g+ 1g-	
@	11	-0.9824896577	68710.165670	100.0	3	0g 1g+ 1g-	
@	12	-0.9824017917	68729.450018	100.0	5	0g 1g+ 1g- 2g	g+ 2g-
@	13	-0.9822669729	68759.039338	100.0	7	0g 1g+ 1g- 2g	g+ 2g- 3g+ 3g-
@	14	-0.9751409822	70323.013514	100.0	3	0g 1g+ 1g-	
@	15	-0.9723223910	70941.622775	100.0	1	0g	
@	16	-0.9722734205	70952.370551	100.0	3	0g 1g+ 1g-	
@	17	-0.9721863073	70971.489684	100.0	5	0g 1g+ 1g- 2g	g+ 2g-
@	18	-0.9616074654	73293.277111	99.8	5	0g 1g+ 1g- 2g	g+ 2g-

```
19
          -0.9497229655 ...
                               75901.623363
                                                99.9
0
                                                        1
                                                            0g
0
    20
          -0.9322863070 ...
                               79728.527553
                                               100.0
                                                        5
                                                            0u 1u+ 1u- 2u+ 2u-
0
    21
          -0.9307044048 ...
                               80075.714943
                                               100.0
                                                        5
                                                            0u 1u+ 1u- 2u+ 2u-
@
    22
          -0.9306318565 ...
                               80091.637473
                                               100.0
                                                        7
                                                            0u 1u+ 1u- 2u+ 2u- 3u+ 3u-
0
    23
                                                            0u 1u+ 1u- 2u+ 2u- 3u+ 3u-
          -0.9305012392 ...
                               80120.304651
                                               100.0
                                                        9
0
                                                        3
                                                            0u 1u+ 1u-
    24
          -0.9303011549 ...
                               80164.218067
                                               100.0
                                                            0u 1u+ 1u- 2u+ 2u-
@
    25
          -0.9302612163 ...
                               80172.983578
                                               100.0
                                                        5
@
          -0.9302131715 ...
                               80183.528205
                                                        7
                                                            0u 1u+ 1u- 2u+ 2u- 3u+ 3u-
    26
                                               100.0
@
    27
          -0.9264080659 ...
                               81018.652356
                                               100.0
                                                        7
                                                            0u 1u+ 1u- 2u+ 2u- 3u+ 3u-
@
    28
          -0.9244722663 ...
                              81443.511260
                                               100.0
                                                        5
                                                            0u 1u+ 1u- 2u+ 2u-
@
    29
          -0.9243950805 ...
                               81460.451580
                                               100.0
                                                        3
                                                            0u 1u+ 1u-
@
    30
          -0.9243547369 ...
                              81469.305985
                                               100.0
                                                        1
                                                            0u
@
    31
          -0.9218925156 ...
                               82009.701090
                                                99.9
                                                        3
                                                            0u 1u+ 1u-
0
    32
          -0.5545617414 ... 162629.487352
                                                 9.6
                                                        1
                                                            0g
@
    33
          -0.5088229529 ... 172667.991098
                                                 0.0
                                                        1
                                                            0u
0
    34
          -0.5088105565 ... 172670.711776
                                                        3
                                                            0u 1u+ 1u-
                                                  0.0
0
    35
          -0.5087857581 ... 172676.154402
                                                 0.0
                                                        5
                                                            0u 1u+ 1u- 2u+ 2u-
 Ionization potential Oh2p -> Oh1p = ... =
                                                10.9819 \text{ eV} =
                                                                 88575.16 cm<sup>-1</sup>
 Ionization potential Oh1p -> Oh0p = ... =
                                                24.2720 \text{ eV} =
                                                                195766.56 cm<sup>-1</sup>
```

The column named "% main" contains total weights of main model space determinants in model vectors describing the given electronic states. It can be seen that these weights are very close to 100% for all target (or main) electronic states. Intermediate states (with weighs of main model space determinants not exceeding several percents) are also obtained as a by-product. However, they can hardly be considered as physical ones (compare their energies with the  $IP_1 \sim 90000~cm^{-1}$ ). The greater the weight of main model space determinants, the less the dependence on the parameters of the applied denominator shifting scheme.

It is worth noting that in some cases convergence of amplitude equations can be achieved only by the introduction of additional active spinors, which are unoccupied in all main model space determinants. These spinors serves as buffer ones. In calculations accounting for triples such an extension of active space can result in a dramatic increase of computational cost. So it is recommended to thoroughly choose an upper bound of an active space to make it as small as possible. However, this is not a problem for calculations employing the CCSD model only. Introduction of additional buffer states also greatly improves stability of energy levels with respect to the shifting parameters (see Figures 3 and 4 in [10] for the case of electronic states of the Lu<sup>+</sup> ion).

### 5.5 Finite-field transition dipole moments calculations: the Rb atom

For input/output files see examples/Rb\_atom\_TDMs\_finite\_field.

See [31, 58, 3] for detailed discussions on the finite field (FF) technique for transition dipole moments. The algorithm of a typical FF calculation is as follows:

- solve the SCF problem for the non-perturbed Hamiltonian (no external fields) and obtain one-electron functions (spinors);
- transform SCF molecular integrals to the basis of "non-perturbed" one-electron functions;
- perform two FS-CC calculations with these integrals at external fields strengths -F and +F;

• estimate TDMs or other off-diagonal matrix elements using the finite-difference-type formula:

$$\boldsymbol{d}_{\eta,if}^{FF} = (E_i - E_f) \lim_{\substack{F_{\eta} \to 0 \\ F_{\eta'} = 0, \eta' \neq \eta}} \frac{\langle \tilde{\psi}_i^{\perp \perp} (-F_{\eta}) | \tilde{\psi}_f (+F_{\eta}) \rangle}{2F_{\eta}} \quad (\eta = x, y, z), \tag{10}$$

where  $\tilde{\psi}_i^{\perp\perp}$  and  $\tilde{\psi}_f$  are the left and right model vectors corresponding to the **i**nitial and **f**inal electronic states, respectively.

Consider the FF calculation of TDMs for the  ${}^2S_{1/2} \leftrightarrow {}^2P^o_{1/2}$  and  ${}^2S_{1/2} \leftrightarrow {}^2P^o_{3/2}$  transitions in the Rb atom. These states can be obtained in the 0h1p sector starting from the Rb<sup>+</sup> closed shell state:

$$Rb^+ (0h0p) \rightarrow Rb^0 (0h1p)$$

The simplest active space required to describe the low-lying  ${}^2S_{1/2}$ ,  ${}^2P_{1/2}^o$  and  ${}^2P_{3/2}^o$  states of Rb comprised 5s-and 5p-spinors of Rb<sup>+</sup> (overall 8 spinors, or 4 Kramers pairs). The basis set and relativistic pseudopotential used here were taken from [58] and reduced to [7s7p5d3f2g].

Firstly, we perform the Hartree-Fock calculation of Rb<sup>+</sup> and subsequent transformation of molecular integrals. Note that transformed integrals over one-electron property operators (in this case, the dipole moment operator) are stored in the MDPROP unformatted file which should be copied from DIRAC scratch to the working directory:

```
pam --noarch --inp=moltra.inp --mol=Rb.mol --get="MRCONEE MDCINT MDPROP"
```

Other integrals are saved to the MRCONEE (one-electron Fock) and MDCINT (two-electron Coulomb) unformatted files. Note that the .PRPTRA option as well as the \*\*PROPERTIES/.DIPOLE specification must be explicitly specified in the moltra.inp input file:

```
**MOLTRA
.PRPTRA
.SCHEME
4
.ACTIVE
energy -100.0 20.0 0.5
**PROPERTIES
.DIPOLE
```

These lines are used to ask DIRAC to calculate and transform dipole moment integrals (these integrals are used later to construct the perturbation contribution to the Fock matrix due to the external field). Also note that the external uniform electric field lowers spherical symmetry of the atom to  $C_{\infty v}$ .

Now we are ready to run an FS-CC calculation. At this step we use the oneprop directive to add the contribution of the an external field  $\mathbf{F}$  contribution to the Fock operator:

$$f'_{pq} = f_{pq} + (d_x)_{pq} F_x + (d_y)_{pq} F_y + (d_z)_{pq} F_z$$

Here we use only the  $d_z$  component, stored in the MDPROP file in the ZDIPLEN record. The operator which is added to the Fock matrix must be fully symmetric. The first EXP-T input file (for the field strength -F):

```
# title for the task (of type "string")
title "Rb atom -- relativistic EA-CCSD calculation"
# max number of iterations
maxiter 50
```

```
# convergence threshold (by cluster amplitudes)
conv 1e-9

# target Fock space sector
sector Oh1p

# active space specification: 8 lowest virtual spinors (5s,5p1/2,5p3/2)
nactp 8

# add the external field contribution to the Fock matrix:
# Fock = Fock - d_z * F_z
# (F_z = - 1e-5)
oneprop -1e-5 0.0 "ZDIPLEN"
```

For the other field strength +F it is convenient to reuse two-electron integrals, which have already been sorted (the reuse option). One-electron (Fock) integrals depend on the finite perturbation parameter and thus have to be recomputed. Cluster amplitudes converged at -F can also serve as a good initial guess for the calculation at +F. This will lead to significant computational savings. Use the --no-clean option to prevent the deletion of files containing sorted two-electron integrals and converged amplitudes.

The second EXP-T input file (for the field strength +F):

```
title "Rb atom -- relativistic EA-CCSD calculation"
maxiter 50
conv 1e-9
sector 0h1p
nactp 8

reuse 2-integrals amplitudes
oneprop +1e-5 0.0 "ZDIPLEN"
```

Run EXP-T and save formatted files HEFF with effective Hamiltonian matrices to the working directory:

```
expt.x --no-clean ccsd_F-.inp | tee ccsd_F-.out
mv scratch/HEFF HEFF1
expt.x --no-clean ccsd_F+.inp | tee ccsd_F+.out
mv scratch/HEFF HEFF2
```

The electronic spectrum is given at the end of the EXP-T output file. Corresponding lines begin with the @ symbol:

```
Level Re(eigenvalue)
                      . . . Rel eigv, cm-1 deg symmetry
        -0.1531234146
                                  0.000000
                                             2
                                                 1/2+ 1/2-
        -0.0959179975
                              12555.137832
                                             2
                                                 1/2+ 1/2-
    3
                              12786.558570
                                                 1/2+ 1/2-
        -0.0948635671
                                             2
                                                 3/2+ 3/2-
        -0.0948635506
                              12786.562182
                                             2
```

(compare with experimental values 12578.950 cm<sup>-1</sup> and 12816.545 cm<sup>-1</sup> [46]; note that the  ${}^{2}P_{3/2}^{o}$  state is split in the external electric field).

Finally, we turn to the calculation of transition moments. The input file for the heffman.x utility is as follows (see Section 6.1 for details and explanation):

file: HEFF1 file: HEFF2 sector: Oh1p

rep: 1

step: 0.00002

print: 1

Run heffman.x:

heffman.x < ff\_tdm.inp | tee ff\_tdm.out

Transition moments and Einstein coefficients can be found in the end of the output file ff\_tdm.out:

tran	siti	on	energy, cm^	-1		d ^2	d	osc str	A,10^6 s-1
1	->	2	12555.13	88	 	2.627905	1.62108158	0.100220	10.537579
1	->	3	12786.55	59		5.173064	2.27443710	0.200921	21.911672
2	->	3	231.42	21	 	0.000000	0.00001167	0.000000	0.000000

Note: this is only the  $|d_z|$  component of the full transition moment  $\mathbf{d}$ . It should be mentioned that the small, but non-zero value of  $|\mathbf{d}|$  in the last line is an artifact of the FF approach.

Taking into account the degeneracy of the  ${}^2P^o_{3/2}$  state and the spherical symmetry of the Rb atom (one can use the Wigner-Eckart theorem to restore the other components of d), the full transition moments |d| and corresponding Einstein coefficients can be calculated:

$$A(^{2}S_{1/2} \leftarrow {^{2}P_{1/2}^{o}}) = 31.6 \text{ s}^{-1}$$
  
 $A(^{2}S_{1/2} \leftarrow {^{2}P_{3/2}^{o}}) = 32.9 \text{ s}^{-1}$ 

(compare with experimental values  $34.0 \text{ s}^{-1}$  and  $37.0 \text{ s}^{-1}$  [46]).

Results can be improved significantly by extending the basis set and active space [58].

## 5.6 Finite-order approach to transition dipole moment calculations: the Rb atom

For input/output files see examples/Rb\_atom\_TDMs\_finite\_order.

The finite-order approach to transition moment calculations consists in the substitution of the FS-CC wave functions  $|\psi_i\rangle = \{e^T\} |\tilde{\psi}_i\rangle$  and  $|\psi_f\rangle = \{e^T\} |\tilde{\psi}_f\rangle$  into the expression for the matrix element  $O_{if} = \langle \psi_i | O | \psi_f \rangle$ . Resulting infinite summations are truncated at some power of T. Detailed description of the theoretical background is given in [4, 17]. Actually the same approach was previously developed for atomic applications, see the papers [59, 60, 61] (and references therein).

Let us apply the finite-order techique to calculate transition dipole moments in Rb. Since property matrix elements are evaluated analytically, there is no need to reduce the symmetry from  $D_{\infty h}$  to  $C_{\infty v}$ . We again begin with the DIRAC step to obtain unformatted files with transformed molecular integrals:

```
pam --inp=moltra --mol=Rb --noarch --get="MRCONEE MDCINT MDPROP"
```

The MDPROP file contains three entries representing matrices of the  $d_x$ ,  $d_y$  and  $d_z$  components of the dipole moment operator (XDIPLEN, YDIPLEN and ZDIPLEN, respectively). For some reasons the XDIPLEN and ZDIPLEN matrices are purely imaginary (this was found in the series of preliminary numerical experiments), appealing to complex arithmetics in CC calculation. Such a situation is quite typical.

Thus we recommend to perform an FS-CC calculation of properties in two steps. At the first step we run the usual FS-CC calculation with the real arithmetics employed. The EXP-T input file (ccsd.inp) looks like:

```
title "Rb atom -- relativistic EA-CCSD calculation"
maxiter 50
conv 1e-9
sector 0h1p
nactp 8
disk_usage 0
```

At the second step we calculate transition moments (the input file tdm.inp is listed below). We should turn on complex arithmetics (arith complex) and restart the calculation from the pre-sorted integrals and converged amplitudes (they will be automatically expanded to the complex arithmetics):

```
title "Rb atom -- relativistic EA-CCSD calculation"
maxiter 50
conv 1e-9
sector Oh1p
nactp 8
disk_usage 0

reuse 2-integrals amplitudes
arith complex
print "model vectors"

mdprop "XDIPLEN" sym 1u+ approx 2 2 scheme connected
mdprop "XDIPLEN" sym 1u- approx 2 2 scheme connected
mdprop "YDIPLEN" sym 1u+ approx 2 2 scheme connected
mdprop "YDIPLEN" sym 1u+ approx 2 2 scheme connected
mdprop "YDIPLEN" sym 1u- approx 2 2 scheme connected
mdprop "ZDIPLEN" sym 0u approx 2 2 scheme connected
```

Note that the instruction print "model vectors" was added. It prints additional information about electronic energy levels and is mandatory for further successful processing of the output file with the expt\_spectrum.py script (for more details on this auxiliary tool, see Section 6.4).

Running EXP-T:

```
expt.x --no-clean ccsd.inp > ccsd.out
expt.x --no-clean tdm.inp > tdm.out
```

The second run of EXP-T typically produces huge output files since all possible pairs of transition matrix elements are calculated. It is recommended to restrict the manifold of target electronic states to be printed and analyzed using either the nroots or roots\_cutoff keywords. The required TDM values can be found in the output file and then extracted manually or using the expt\_spectrum.py script:

```
expt_spectrum.py tdm.out -irep 1/2g+ -H
```

where the -irep option defines the irrep to which the initial states belong, and -H enables hermitization of transition dipoles. The output will be as follows:

```
initial irreps: ['1/2g+']
initial states: all
final irreps : all
final states : all
transitions 1/2g+ -> 1/2u+
```

```
|d|,a.u.
                e_i,cm^-1
                             e_f,cm^-1
                                                                osc.str.
   1 ->
           1
                   0.0000
                            12555.1431
                                                   1.756964
                                                                0.117726
   1 ->
            1
               12555.1431
                                0.0000
                                                   1.725401
                                                               -0.113534
                                                   1.741111
                                                                0.115611
           2
   1 ->
                   0.0000
                            12786.5663
                                                   2.480954
                                                                0.239065
               12786.5663
   2 ->
                                0.0000
                                                   2.433780
                                                               -0.230060
                                                   2.457254
                                                                0.234519
transitions
              1/2g+ -> 1/2u-
                                                   |d|,a.u.
                e_i,cm^-1
                             e_f, cm^-1
                                                                osc.str.
                   0.0000
                            12555.1431
                                                   2.484722
   1 ->
           1
                                                                0.235452
               12555.1431
                                0.0000
                                                   2.440085
                                                               -0.227068
                                                   2.462303
                                                                0.231222
   1 ->
           2
                   0.0000
                            12786.5663
                                                   1.754299
                                                                0.119532
   2 ->
           1
               12786.5663
                                0.0000
                                                   1.720942
                                                               -0.115030
                                                   1.737541
                                                                0.117260
              1/2g+
transitions
                             e_f, cm^-1
                e_i,cm^-1
                                                   |d|,a.u.
                                                                osc.str.
   1 ->
            1
                   0.0000
                            12786.5663
                                                   3.038535
                                                                0.358597
   1 ->
            1
               12786.5663
                                0.0000
                                                   2.980759
                                                               -0.345090
                                                   3.009509
                                                                0.351779
```

One can see that the finite-order technique in the quadratic approximation (approx 2 2) overestimates transition moments by  $\sim 9\%$  (compared to the finite-field approach). It seems reasonable to recalculate transition moments in the linear approximation (approx 1 0); for the particular case of the Rb atom it gives nearly the same answers. This is not a surprising result since cluster amplutudes are relatively small (< 0.07). Thus even the linear approximation works correctly.

### 5.7 "High-spin" CCSD for open shell systems

For input/output files see examples/02\_highspin.

The example considered in this subsection is borrowed from the DIRAC program manual [24]. Let us calculate the energy of the ground (triplet) state of the  $O_2$  molecule.

The DIRAC input file contains directives for the Hartree-Fock calculation and subsequent transformation of molecular integrals:

(02.mol)

```
test of open-shell CCSD calculation
molecular oxygen at eq distance taken from NIST
automatic symmetry detection: DIRAC will identify the Dinfh group
C
    1
                    Α
       8.
             2
0
       0.0
                  0.0
                           0.60376
                  0.0
                          -0.60376
0
       0.0
LARGE BASIS cc-pVDZ
FINISH
```

```
(moltra.inp)
! calculation of the triplet ground state of molecular oxygen
! (Dirac-Coulomb Hamiltonian)
**DIRAC
.TITLE
molecular oxygen -- ground state
.WAVE F
.4INDEX
**WAVE FUNCTIONS
.SCF
*SCF
.CLOSED SHELL
 6 8
.OPEN SHELL
1
2/4,0
*END OF
```

DIRAC writes occupation numbers of molecular spinors to the MRCONEE unformatted file, but only spinors belonging to doubly occupied shells (the .CLOSED\_SHELL keyword) are labeled as occupied. This is why EXP-T will calculate the energy for the  $O_2^{2+}$  cation without manually specified occupation numbers. EXP-T provides two ways to set the occupation numbers of molecular spinors:

• by setting the total number of occupied spinors in each irrep (the occ\_irreps keyword). For example, occupation numbers for the open-shell triplet O<sub>2</sub> molecule can be specified using the following two instructions:

```
occ_irreps [1/2g+]:3 [1/2g-]:2 [3/2g+]:0 [3/2g-]:1 occ_irreps [1/2u+]:2 [1/2u-]:2 [3/2u+]:1 [3/2u-]:1
```

Note that the 1s-orbitals of oxygen atoms are frozen here, and the corresponding electrons should not be accounted for in the occ\_irreps instructions. For the example discussed:

- 3 electrons occupy the lowest spinors belonging to the 1/2g+ irrep;
- $\frac{1}{2g}$ ,  $\frac{1}{2u}$ ,  $\frac{1}{2u}$ : 2 electrons in each irrep;
- -3/2g-, 3/2u+, 3/2u-: 1 electron in each irrep.
- you can also specify the occupation number for *each* spinor explicitly:

To identify spinors which should be populated in the electronic state under consideration, we again recommend to start the calculation with the default occupation numbers, then to interrupt the task and find information about spinors and their symmetry in the output. Then the calculation is restarted with the correct occupation numbers.

We recommend to use the dirac\_mointegral\_export program from the DIRAC package to process the MRCONEE unformatted file and get such a string of occupation numbers.

Due to the fact that non-canonical orbitals are used, the Fock matrix is not diagonal and the energy of the vacuum determinant  $\langle \Phi_0 | H | \Phi_0 \rangle$  is not equal to the SCF energy obtained before (in the Hartree-Fock calculation with fractional occupation numbers for open-shells):

```
-149.686661451842
Total SCF energy =
sorting one-electron integrals ...
 Fock matrix reconstruction ...
 SCF energy (energy of reference determinant) =
                                                 -149.718144633810 a.u.
 SCF energy (energy of reference determinant) was updated:
   old energy =
                 -149.686661451842 a.u.
   new energy =
                    -149.718144633810 a.u.
        SCF reference energy =
                                  -149.718144633810
      CCSD correlation energy =
                                     -0.366958682652
            Total CCSD energy =
                                   -150.085103316462
```

## 6 Utility programs

#### 6.1 heffman.x - manipulations with effective Hamiltonian matrices

The heffman.x program (written by A. Zaitsevskii) is used to perform manipulations with FS-CC effective Hamiltonian matrices. Implemented features are:

- the finite-field technique for transition moment calculations (e. g. TDMs) [31, 58, 3];
- extraction of spin-orbit interactions from full-relativistic models by projection [5];
- Pade extrapolation of series of effective Hamiltonians to the limit of zero energy denominator shifts [9].

heffman.x reads effective Hamiltonian matrices of effective Hamiltonians from formatted (text) files supported by both EXP-T and the latest versions of DIRAC.

- to compute finite-field transition moments two files with effective Hamiltonians must be provided, calculated at "field 1" and "field 2";
- to extract spin-orbit interaction matrix elements by the projection technique, two files with effective Hamiltonians must be provided, the first one for spin-orbit (nearly) switched off and the second one for spin-orbit switched on.

In this case you cannot freeze spinors at the FS-CC step!

• to perform Pade extrapolation of the series of effective Hamiltonians, you should put a sequence of effective Hamiltonians to a single file using the cat command:

```
cat heff_n_eq_1 heff_n_eq_2 heff_n_eq_3 > some_file_name
```

Recall that each effective Hamiltonian matrix is calculated at its own attenuation parameter n. You can also perform extrapolation before transition moments or SO calculation, just use such "composite" files with effective Hamiltonians.

To run heffman.x, an input file containing a list of instructions must be redirected to stdin:

#### heffman.x < input\_file

heffman.x input files are very similar to EXP-T input files, but keywords are separated from their arguments with the ':' symbol, and the symbol '#' at the beginning of a comment must be placed in the first position of the line.

List of instructions:

file: file with the first effective Hamiltonian or with a sequence of Hamiltonian matrices to be extrapolated, e. g.

- Hamiltonian for the first field value:
- Hamiltonian for the spin-orbit interaction (nearly) switched off;
- Hamiltonian for which you want to get the eigenstates;
- sequence of effective Hamiltonian matrices you want to Padé-extrapolate.

#### Syntax: file: <string path-to-first-heff>

If you are doing finite-field transition moment calculations or SO separation, you have to supply the second file (see the next item).

heff for another field value, or heff for the spin-orbit switched on, respectively. If there is no second file, the effective Hamiltonian eigenvectors and eigenvalues are written to the formatted file HEFFEVF.

Syntax: file: <string path-to-second-heff>

sector: the Fock space sector you are interested in (heff files for higher sectors contain several heff

matrices).

Syntax: sector: <H>h<P>p
Example: sector: 0h2p

rep: irreducible representation you are interested in. As a rule, rep: 1 stands for the fully symmetrical representation

metrical representation.

Syntax: rep: <integer n>

Numbering of representations starts from 1!

main: — for the FF TDM calculations or the Padé extrapolation: the number of the lowest states or the range of states of the given sector and symmetry you want to see in the output file:

- for the SO extraction or the projected approximant construction [9] this keyword defines the subspace of states used for projection.

Syntax: main: (<integer number> | <integer number1>-<integer number2>).

Default: all states.

Example: main: 10 (or, equivalently, main: 1-10)

[L/M] (optional) to perform Padé extrapolation, you have to indicate which matrix Padé approximant [L/M] should be constructed ([0/1], [1/1], [0/2] etc). You should have L+M+2 effective Hamiltonian matrices to build the [L/M] approximant. You can cut off all states above the subspace done by the main keyword (see above); to do this, add the cut keyword.

Syntax: [L/M] [cut]

Default: Padé extrapolation disabled

Example: [0/1] cut

step: (only for FF calculations of transition moments.) Difference of field strengths (a. u.) between the two finite-field calculations. The presence of this line activates the FF TDM calculation mode (do not use this keyword in other types of calculations).

Syntax: step: <real value>

scale: (only for SO calculations.) Fraction of effective spin-orbit retained in "scalar" calculations to preserve good relativistic symmetries (typically of order  $10^{-4}$ ). Omit this line if you do not want to diabatize anything.

Syntax: scale: <real value>

ground: ground state energy (a. u.). This option is useful if one needs to compute energies with respect to some state (normally ground) which belongs to another symmetry or Fock space sector.

Syntax: ground: <real energy>

Default: ground: 0.0

print: print level. "print: 0" or "print: 1" are normally ok; enlarging this number, you can get large amount of superfluous information listed.

Syntax: print: <integer level>

Default: print: 1

activates the computation of finite-field transition moments with Bloch effective Hamiltonian eigenvectors (note: transition property matrices will be non-Hermitian, this has no physical meaning). Otherwise, FF computations are performed with des Cloizeaux effective Hamiltonian eigenvectors ("hermitization").

Syntax: Bloch

Default: "hermitization" is enabled in finite-field calculations

When using any of the techniques implemented in the heffman.x program, please, cite the corresponding paper.

## 6.2 expt\_transform\_natorb.x - transformation of model natural spinors to the AO basis

Having model vectors  $|\tilde{\psi}_i\rangle$  one can construct approximate density  $d_{pq} = \langle \tilde{\psi}_i | a_p^{\dagger} a_q | \tilde{\psi}_i \rangle$  and transition density matrices  $d_{pq} = \langle \tilde{\psi}_i | a_p^{\dagger} a_q | \tilde{\psi}_j \rangle$  and then calculate model natural (transition) molecular spinors (NS and NTS, respectively) [6, 7]. These operations can be performed by EXP-T itself (see Section 4, keyword natorb). EXP-T stores NS and NTS expanded in the basis of molecular spinors in formatted files named NATORB\*.dat. To use visualization tools one has to transform NS or NTS to the AO basis. This transformation can be performed by the expt\_transform\_natorb.x utility.

expt\_transform\_natorb.x requires two command-line arguments:

- path to the DIRAC output file containing molecular spinors expanded in the basis of atomic orbitals;
- path to the NATORB\*.dat formatted file produced by EXP-T and stored in its scratch directory (contains natural spinors expanded in the basis of molecular spinors).

To ask DIRAC to print molecular spinors (vectors) to the output, enable the .ANALYZE module in the \*\*DIRAC section and then add the following text to the \*\*ANALYZE section of the DIRAC .inp input file [62]:

```
**ANALYZE
```

.PRIVEC

\*PRIVEC

.AOLAB

. VECPRI

all

(.AOLAB = "print vectors in the AO basis"). For point groups with inversion one must add the second string specifying the range of molecular spinors:

.VECPRI

all

all

Running expt\_transform\_natorb.x (for example):

```
expt_transform_natorb.x dirac.out NATORB_Oh1p_1:1.dat > nts.out
```

Transformed spinors are flushed to stdout again in the DIRAC four-column format and can be redirected to a file.

#### 6.3 expt2pam.x - input file format converter

EXP-T does not provide subroutines for solution of Hartree-Fock equations and integral transformation, it completely relies on molecular integrals calculated with the help of any other quantum chemistry software. Thus EXP-T input files do not contain any information on molecular geometry, basis set and pseudopotentials. However, to simplify interaction with the third-party software and make the preparation of input files more convenient, the special file format was developed. It is designed to be quite flexible and free of any format restrictions typical for software written in Fortran.

The EXP-T-style input files containing information on molecular geometry, basis sets and pseudopotentials are recognized by the expt2pam.x program, which can produce input files for the following software:

```
DIRAC [1, 2]: option --dirac;
CFOUR [63]: option --cfour;
OneProp [8]: option --oneprop.
```

The input file format closely resembles that of NWChem [64, 65, 66, 67], but is substantially simplified. The most important points on the format to be emphasized are:

- only Cartesian coordinates for molecular geometry definition are available to the moment;
- primitive Gaussian functions can be defined without the superfluous contraction coefficient 1.0;
- primitive Gaussian functions representing the uncontracted part of the basis set with the given angular momentum L can be placed into the same block.

Below the example of such an input file demonstrating the main features of the expt2pam format is presented (the UO<sub>2</sub> molecule [56]):

```
# Uranium dioxide molecule UO2
geometry units angstrom
    0.0 0.0
                0.000
    0.0 0.0
                1.775
  0
     0.0 0.0
              -1.775
end
 Contracted basis set adapted for the 28-in-core\ GRPP by N.S. Mosyagin
#
basis
US
 11931.71300000
  6539.04200000
  3678.23270000
  2139.54940000
  1288.45330000
   792.07718000
   484.69932000
   298.00127000
   186.73959000
   120.33238000
```

```
75.85031400
```

- 49.22312100
- 30.97937000
- 19.33898600
- 11.90917900
- 7.28108690
- 4.52381560
- 2.69147870
- 1.62933420
- 1.02810670
- 1.02010010
- 0.60588439
- 0.35564698
- 0.20180019
- 0.12481356
- 0.06383616

#### UΡ

- 8256.41640000
- 4196.77840000
- 2191.89420000
- 1174.82830000
  - 644.61478000
- 361.41388000
- 206.27034000
- 119.33670000
  - 70.14308600
  - 41.70407000
  - 24.70288000
  - 14.49639800
  - 8.57587000
  - 4.98086080
  - 2.84040830
  - 1.59903720
  - 0.86476597
  - 0.45724205
  - 0.23893352
  - 0.11879300
  - 0.05700662

#### U D

- 6012.79360000
- 2754.36670000
- 1351.96790000
  - 699.79805000
- 377.01141000
- 209.73941000
- 119.30191000
  - 69.11171400
  - 40.43241500
  - 23.65423900
  - 13.80639900
    - 8.06231670

```
4.66375400
     2.63565040
     1.45796000
     0.77481216
     0.37248472
     0.17195534
     0.07613361
UF
  1163.36020000
   573.50499000
   285.71163000
   136.47329000
    69.43427500
    36.84465700
    19.84804900
    10.72168500
     5.70450120
     2.92433440
     1.43960400
     0.67385930
     0.29434970
     0.11323985
U G
    28.78919130
                   0.00068974
                                 -0.00118120
                                                -0.00737416
                                                                0.00494647
                                                                             -0.10867784
    14.01320380
                  -0.00898795
                                 -0.05421964
                                                 0.17284128
                                                               -0.36594951
                                                                             -0.33163329
     6.82095857
                   0.07975524
                                  0.18134441
                                                -0.50920321
                                                                1.20716202
                                                                              1.79274020
     3.32011698
                    0.29993353
                                  0.65729862
                                                -0.50617914
                                                               -0.67702022
                                                                             -2.65141994
     1.61607443
                   0.39302861
                                  0.04007231
                                                 1.01049051
                                                               -0.71780643
                                                                              2.66362756
     0.78662787
                   0.34151054
                                 -0.44799971
                                                 0.05077223
                                                                1.59628888
                                                                             -2.01035588
     0.38289289
                   0.18576214
                                 -0.38383279
                                                -0.74551549
                                                               -1.12509753
                                                                              0.90492997
UH
    14.01320380
                  -0.02449353
                                 -0.07337288
                                                -0.15906456
                                                               -0.23132505
     6.82095857
                   0.10594999
                                  0.30442614
                                                 0.71623431
                                                                1.22980099
     3.32011698
                   0.35765256
                                  0.65468802
                                                 0.09614021
                                                               -1.24473080
                                                -0.95514471
     1.61607443
                   0.43019757
                                 -0.18764249
                                                                0.32540556
     0.78662787
                   0.31957043
                                 -0.48353526
                                                 0.41946211
                                                                0.73545263
     0.38289289
                   0.10197395
                                 -0.22548628
                                                 0.48672026
                                                               -0.90679010
UΙ
     6.82095857
                   0.12524957
                                 -0.37984959
                                                 0.81416792
     3.32011698
                   0.42961205
                                 -0.61109425
                                                -0.35518331
     1.61607443
                   0.42313291
                                                -0.73501069
                                  0.38916709
     0.78662787
                   0.31396902
                                  0.50454838
                                                 0.88347742
end
# Relativistic pseudopotential (RPP) for Uranium by N.S. Mosyagin
# 28e in core, QED+Breit are included
#
еср
```

```
U nelec 28
U UI.
 1
     573.3050904771354
                              -11.40084159918683
 1
     192.8309334143903
                              -7.097553073552547
 1
     59.69944003435170
                              -28.01857233638748
 1
     41.19441327874497
                               42.95299466064716
 1
     27.48466038227223
                              -39.29896548592660
                               21.07523831226354
 1
     17.41575927233871
 1
     10.74156012298913
                              -8.338244516015038
 1
     6.516110723607048
                               3.069168416643151
 1
     4.072460983953957
                              -1.269035430883378
 1
     2.593752889495458
                              0.4038040550003643
 1
     1.642108539873937
                             -0.7510043731679161E-001
    0.5908313234915240
                              0.3890633153407208E-002
 1
US
 0
     197.0298350128211
                               6.000000000000000
 1
     129.5661840079536
                              -355.1867566615911
 1
     93.02818620674168
                               1191.419958549814
 1
     66.13198912807648
                              -1657.083906766264
 1
     45.36994786350133
                               1304.226236566319
 1
     31.03973237770222
                              -448.9389445476668
 1
     21.15323065428634
                               86.95891078559328
 1
     8.842323178530615
                              -7.379538028367257
 1
     2.503929576228269
                               3.473180786461057
 2
     2.925942781684598
                              -5.455412335112918
UР
 0
     8.311409755029183
                               5.000000000000000
     126.5274319867533
                              -10.06000092823524
 1
 1
     85.18677663478607
                               70.55297376279486
 1
     51.21216731114048
                              -221.8073735731159
 1
     29.49108937084788
                               312.1302683559188
 1
     17.84600190857478
                              -299.1627976468299
                               229.1321036359295
 1
     5.463491362893393
 2
     8.266178701884957
                              -440.5446582608856
 2
     5.564984596440512
                              -230.2801015020814
 2
     3.621496981299310
                              0.6144684865466843
U D
 0
     63.75445017844640
                               3.000000000000000
     35.83829800943973
 1
                              -244.0643648227415
 1
     25.77961295248244
                               369.6289272300128
 1
     16.62718251626653
                              -66.05286509307552
 2
     21.10083097720674
                              -511.0304327796694
 2
     13.00689068383810
                               209.7715531529684
 2
     7.652250767129301
                              -56.31696437943174
 2
     4.429977591227489
                               19.86068552036704
 2
     2.777850799855992
                              -8.087414881056247
 2
     1.804331049757116
                               2.539779591468687
 2
     1.174834421468355
                             -0.5000680379316691
 2
    0.7571835654831253
                              0.4510869608679968E-001
end
```

```
so
U P
     126.5274319867533
                            -60.8272897577457000
 1
 1
     85.18677663478607
                             144.881442506022000
 1
     51.21216731114048
                            -174.959781119600000
 1
     29.49108937084788
                             113.392635098503000
     17.84600190857478
                            -60.8426035265438000
 1
 1
     5.463491362893393
                             45.8248263173737000
 2
     8.266178701884957
                            -82.3448048559949000
 2
     5.564984596440512
                            -47.9408659655601000
 2
     3.621496981299310
                              0.4919304400803720
U D
 1
     35.83829800943973
                             -0.4757330698863370
     25.77961295248244
                             -8.8286299074417900
 1
 1
     16.62718251626653
                              9.6090737588470800
 2
     21.10083097720674
                              14.829424758465000
 2
     13.00689068383810
                             -30.085848351789200
 2
     7.652250767129301
                              10.782019233677400
 2
     4.429977591227489
                             -3.9646231214668700
 2
     2.777850799855992
                              1.4116480556389400
 2
                             -0.3371713502943750
     1.804331049757116
 2
     1.174834421468355
                              0.0452391741906467
 2
    0.7571835654831253
                             -0.0020991859739297
UF
     573.3050904771354
                              1.9478068936173300
 1
 1
     192.8309334143903
                              0.3189229591513530
 1
     59.69944003435170
                              1.1230953284521100
     41.19441327874497
                             -2.2059715944841200
 1
 1
     27.48466038227223
                              2.4430555431733500
     17.41575927233871
                             -1.7928190693741500
 1
 1
     10.74156012298913
                              1.0519241531739200
 1
     6.516110723607048
                             -0.5026634876599670
 1
     4.072460983953957
                              0.2165908956875290
 1
     2.593752889495458
                             -0.0639500309090571
     1.642108539873937
                              0.0107075321136888
 1
    0.5908313234915240
                              0.0000807064226547
end
#
 basis for oxygen: aug-cc-pVQZ-DK
basis
0 S
      61420.00000000
                         0.00019890
                                      -0.00004520
```

spin-orbit part of RPP for Uranium

#

```
9199.00000000
                         0.00088840
                                       -0.00020260
       2091.00000000
                         0.00394250
                                       -0.00089480
        590.90000000
                         0.01560720
                                       -0.00360580
        192.30000000
                         0.05294610
                                       -0.01230330
         69.32000000
                         0.14643820
                                       -0.03646840
         26.97000000
                         0.30541480
                                       -0.08316710
         11.10000000
                         0.39799790
                                       -0.15217400
          4.68200000
                         0.21643540
                                       -0.11462710
          1.42800000
                         0.01753570
                                        0.29002590
          0.55470000
                        -0.00249230
                                        0.58584370
          0.20670000
                         0.00095070
                                        0.27673420
0 S
          1.42800000
          0.55470000
          0.20670000
          0.06959000
0 P
         63.42000000
                         0.00611070
         14.66000000
                         0.04189570
          4.45900000
                         0.16123290
          1.53100000
                         0.35667520
          0.53020000
                         0.44818370
          0.17500000
                         0.24505420
0 P
          1.53100000
          0.53020000
          0.17500000
          0.05348000
0 D
          3.77500000
          1.30000000
          0.44400000
          0.15400000
0 F
          2.66600000
          0.85900000
          0.32400000
0 G
          1.84600000
          0.71400000
end
#
 Empty-core semilocal RPP for oxygen by N.S. Mosyagin
#
еср
O nelec O
O UL
          2.443097724280e+02
                                    1.369170316080e-01
   1
```

```
8.520362677383e+01
                                  -2.852592197455e-01
   1
   2
          1.054731223008e+02
                                   2.684494987739e+00
   2
          1.825657247757e+01
                                  -9.433884948610e-03
   2
          4.806589859822e+00
                                  -1.366233123930e-03
   2
          1.753580551235e+00
                                   4.456314085999e-04
   2
          6.566864682586e-01
                                  -2.781240176237e-04
          4.686645649504e-01
                                   3.558106708362e-04
0 S
   1
          1.376748432210e+02
                                   5.591554374814e-02
   1
          1.378552088359e+01
                                   3.868632083368e-03
   2
          6.768869843542e-01
                                   5.419543370664e-04
   2
          4.618614187585e-01
                                  -3.882304906071e-04
end
SO
0 P
   1
          2.443097724280e+02
                                  -9.027082481378e-02
   1
          8.520362677383e+01
                                   2.802454585393e-01
   2
          1.054731223008e+02
                                  -2.610156015857e+00
   2
          1.825657247757e+01
                                   1.266811741578e-02
   2
          4.806589859822e+00
                                   2.570019532778e-03
   2
                                   2.973424726019e-04
          1.753580551235e+00
   2
          6.566864682586e-01
                                   6.007087622745e-05
   2
          4.686645649504e-01
                                  -5.673502079255e-05
end
```

## 6.4 expt\_spectrum.py – a program for extraction of transition moments from EXP-T output files

EXP-T calculations of transition matrix elements can produce huge output files. To simplify their analysis and extraction of subsets of matrix elements needed for a particular problem, the expt\_spectrum.py script is provided. It is written in Python 3 and requires the pre-installed NumPy library.

Options of the expt\_spectrum.py script are as follows:

```
$ expt_spectrum.py --help
usage: expt_spectrum.py [-h] [-u UPPER] [-irep INITIAL_IRREP] [-frep FINAL_IRREP]
       [-i INITIAL_STATE INITIAL_STATE] [-f FINAL_STATE FINAL_STATE]
       [-H] [-s] [-ev] filename
parses exp-t output files and collects information about electronic transitions
and their dipole moments
positional arguments:
 filename
optional arguments:
  -h, --help
                        show this help message and exit
  -u UPPER
                        upper energy limit for the electronic states considered
                        (cm^-1 or ev)
                        irrep of initial electronic states
 -irep INITIAL_IRREP
```

Names of irreducible representations used here are common for the EXP-T package and are listed in Appendix D.

For example, let us consider allowed E1-transitions in the AcF molecule. Its electronic states and E1-transitions between them were thoroughly studied in [26]. Suppose that calculated transition moments are stored in the EXP-T output file named tdms.out. The following command should be used to show all possible allowed transitions from the ground state ( $\Omega = 0^+$ ):

```
expt_spectrum.py tdms.out -i 0 1
```

Note that DIRAC does not discern  $0^+$  and  $0^-$  states, they both are denoted as 0. The ground state is the first (1) state in the irrep with name 0.

To calculate radiation lifetime of an excited state and estimate branching ratios during its decay, one should obtain information on transition dipole moments between this excited state and all states lying below it. For the particular case of the AcF molecule, the most strong transition from the  $X0^+$  ground state is the transition to the (8)1 state. Thus this excited state seems to be the most easily accessed in experiment. To extract all information about transitions from (8)1 to the manifold of low-lying states, one can use the following command:

```
python expt_spectrum.py tdms.out -i 1+ 8 -u 26200
```

Here the -u option is used to restrict the range of final states to those lying below (8)1 ( $\approx 26170 \text{ cm}^{-1}$ ). The expt\_spectrum.py routine can be combined with Unix grep and awk utilities to easily extract transition dipole moment functions or even surfaces from a large set of EXP-T output files.

## 6.5 expt\_diatomic.x - rotational-vibrational levels and properties of diatomic molecules

If one has a potential energy curve U(R) for an electronic state of a diatomic molecule, one can solve the radial Schrödinger equation and obtain rovibrational levels and wavefunctions as well as any other properties (molecular constants, transition moments, Franck-Condon factors). The radial Schrödinger equation is written as:

$$\left[ -\frac{\hbar^2}{2\mu} \frac{d^2}{dr^2} + \frac{J(J+1)}{2\mu r^2} + U(r) \right] \psi_{vJ}(r) = E_{vJ} \psi_{vJ}(r), \tag{11}$$

where

- $\mu = \frac{M_1 M_2}{M_1 + M_2}$  reduced mass,
- $\bullet$  v vibrational quantum number,
- J rotational quantum number.

Normally this second-order differential equation is solved numerically using some mesh of radial points (see, for example, [68, 69, 70]). Within the current implementation, the U(R) function is interpolated using the cubic natural spline. The matrix version of either the Numerov method [71] or the 2nd order finite-difference method [68] is used to perform integration. Only bound rovibrational states can be treated to the moment.

#### 6.5.1 Input file syntax

The syntax of input files for the solver is quite similar to that of the coupled cluster program. Single line comments start with the '#' symbol. The following directives are allowed:

masses of isotopes  $M_1$  and  $M_2$  (in amu).

Syntax: masses <float M1> <float M2>

charge net charge of the molecule q. For charged molecules the "charge-modified reduced mass" formula [72, 70] is used:

 $\mu = \frac{M_1 M_2}{M_1 + M_2 - m_e q},\tag{12}$ 

where  $m_e$  stands for the electron mass. Can be useful for light molecular ions.

Syntax: charge <integer q>

viblevels range of vibrational levels of interest  $[v_{min}, v_{max}]$ .

Syntax: viblevels <integer vmin> <integer vmax>

Default: viblevels 0 0

rotlevels range of rotational levels of interest  $[J_{min}, J_{max}]$ .

Syntax: rotlevels <integer Jmin> <integer Jmax>

Default: rotlevels 0 0

grid\_size number of radial mesh points.

Syntax: grid\_size <integer npoints>

Default: grid\_size 300

solver integration algorithm:

solver numerov Numerov method [71, 70] (default)

solver fd2 2nd order finite-difference method [68]

mapping meshkov08 <float ra in Angstroms> <float beta>

mapping formula for the radial variable (disabled by default). The reduced variable mapping was first proposed by V. V. Meshkov and co-authors [68]:

$$y(r) = \frac{r^{\beta} - r_a^{\beta}}{r^{\beta} + r_a^{\beta}}, \qquad y \in [-1, 1]$$
(13)

Mapping is available only for the finite-difference solver (solver fd2).

write\_psi flush wavefunctions  $\psi_{Jv}(R)$  to the formatted text files (disabled by default).

potential point-wise potential energy curves and/or property matrix element. The first (r) and the second  $(U_1(r))$  columns are mandatory. The third (property) and the fourth  $(U_2(r))$  columns are optional. If two potentials are specified, transition matrix elements will be calculated. To calculate Franck-Condon factors, use M(r) = 1 in the third column.

Syntax:

```
potential <units of distance> <units of energy>
  <float r> <float U1(r)> [<float M(r)> [<float U2(r)>] ]
   . . .
end

Allowed units of distance are: angstrom, atomic, bohr, au.
Allowed units of energy are: cm-1, cm, atomic, hartree.
```

#### 6.5.2 Example: Franck-Condon factors for the AcOH<sup>+</sup> molecular ion

To estimate suitability of a molecule for direct laser cooling one should calculate vibrational branching ratios and Franck-Condon factors (FCFs) for the working electronic transition [73, 74]. Recently the AcOH<sup>+</sup> molecular ion was proposed as a candidate for direct laser cooling [75].

This molecule is linear in both the ground (1)1/2 and the first excited (1)3/2 electronic states. If we neglect the anharmonic couplings between stretching and bending modes, the total FCF between the  $(v'_1, v'_2, v'_3)$  and the  $(v''_1, v''_2, v''_3)$  vibrational states can be factorized:

$$FCF = |\langle v_1' v_3' | v_1'' v_3'' \rangle|^2 \times |\langle v_2' | v_2'' \rangle|^2$$

$$\tag{14}$$

The latter factor is the one-dimensional bending mode FCF and it is expected to be very close to unity since the corresponding sections of potential energy surfaces are perfectly parallel to each other. The overlap intergal over the  $v_3$  stretch mode is expected to be nearly equal to unity and, which is more important, this factor will be the same for the  $(v'_1, v'_2, 0)$  and  $(v''_1, v''_2, 0)$  vibrational states analyzed here since this stretching mode is not excited in the states under consideration. Thus the total FCF can be approximated by the squared one-dimensional  $\langle v'_1|v''_1\rangle$  overlap integral. To calculate this integral one should solve the radial Schrödinger equation, considering the OH group as a pseudoatom.

The input file for the expt\_diatomic.x program:

```
#
# calculation of Frank-Condon factors for the (1)1/2-(1)3/2 transition in AcOH+
 potential curves were calculated by the relativistic FS-CCSD(0h1p) method,
# for more details, see:
# A. Oleynichenko, L. V. Skripnikov, A. V. Zaitsevskii, V. V. Flambaum
# Laser-coolable AcOH+ ion for CP-violation searches (2022)
 arXiv:2112.02307 [physics.atom-ph]
#
#
           Aс
                      OH
masses 227.028000 17.003000
rotlevels 0 0
viblevels 0 5
                      property(r)=1.0
# r
      potential1(r)
                                          potential2(r)
# units: dist
                 energy
potential atomic atomic
0.340150702E+01
                -0.451045595E+03
                                    1.0 -0.451004746E+03
0.349599333E+01 -0.451064998E+03
                                    1.0 -0.451025470E+03
0.359047964E+01 -0.451079018E+03
                                    1.0 -0.451040834E+03
0.368496594E+01 -0.451088543E+03
                                    1.0 -0.451051696E+03
0.377945225E+01 -0.451094322E+03
                                    1.0
                                        -0.451058791E+03
0.387393856E+01 -0.451096994E+03
                                    1.0
                                        -0.451062746E+03
```

```
0.396842486E+01
                -0.451097103E+03
                                     1.0
                                         -0.451064095E+03
0.406291117E+01
                 -0.451095111E+03
                                    1.0
                                         -0.451063297E+03
0.415739747E+01
                 -0.451091413E+03
                                     1.0
                                         -0.451060740E+03
0.425188378E+01
                 -0.451086341E+03
                                     1.0
                                         -0.451056758E+03
0.434637009E+01
                 -0.451080177E+03
                                     1.0
                                         -0.451051631E+03
0.444085639E+01
                 -0.451073161E+03
                                    1.0
                                         -0.451045598E+03
                                     1.0 -0.451038858E+03
0.453534270E+01
                -0.451065491E+03
end
```

Run the job:

```
expt_diatomic.x AcOH+_FCF.inp | tee AcOH+_FCF.out
```

At the end of the output file one can find the table with transition moments (and their squares). We select only the FCFs with the wavenumber of the upper state v'' = 0:

J,	J,,	ν,	v',		delta e,cm^-1	prop	prop ^2
\$ 0	0	0	0		7279.8734	-0.942947	0.889149
\$ 0	0	1	0		6627.6841	0.325405	0.105888
\$ 0	0	2	0		5979.2731	0.069678	0.004855
\$ 0	0	3	0		5334.3122	0.010337	0.000107
\$ 0	0	4	0		4692.7001	0.001070	0.00001
\$ 0	0	5	0		4054.6587	-0.000079	0.000000

The sum of first three FCFs (see the last column) is equal to 0.9999, thus allowing one to conclude that the closed optical cycle can exist for the AcOH<sup>+</sup> molecular cation [75].

## Appendix A Alphabetical index of keywords

```
active 15
analyt_prop 18
arith 12
compress 26
conv 21
crop 21
cuda 27
damping 21
degen_thresh 12
density 18
diis 22
disk_usage 26
div_thresh 22
flush 22
hermit 18
ih_imms 15
ih_imms : frontier_energy 17
ih_imms: main_occ 16
ih_imms: npower 17
ih_imms : scale 17
ih_imms : sectors 16
ih_imms : shift_type 17
ih_imms : subspace 16
integrals 25
maxiter 23
mdprop 18
memory 26
model 13
mstdm 20
nacth 14
nactp 14
natorb 20
nroots 22
nthreads 27
{\tt occ}\ 14
occ_irreps 14
oneprop 25
openmp 27
openmp_algorithm 27
print 12
reuse 23
roots_cutoff 23
sector 13
shift 23
shift_type 24
skip 25
spinor_labels 13
tilesize 26
```

 $\begin{array}{c} \mathtt{title} \ 12 \\ \mathtt{txtprop} \ 20 \\ \mathtt{x2cmmf} \ 25 \end{array}$ 

# Appendix B Formatted files for storing property operator matrices

EXP-T can be used to calculate matrix elements (expectation values and transition moments) of an arbitrary one-electron property operator not implemented in DIRAC. To carry out such a calculation EXP-T requires two text formatted files containing real and imaginary parts of a property operators matrix expressed in the basis of molecular spinors. Each of these two files consists of a list of lines storing indices (i, j) of a matrix element and an element  $A_{ij}$  itself (its real or imaginary part):

```
...
<integer i> <integer j> <real A_ij>
...
```

Numbering of indices starts from 1. Note that since one-electron property operators are Hermitian, the real part of a matrix must be symmetric while the imaginary part is always antisymmetric.

## Appendix C Temporary files

During the calculation, EXP-T operates with temporary files stored in the *scratch* directory. We recommend to remove these temporary files after every calculation (except for those planned to be used in the future, like the HEFF files or files with cluster amplitudes). All types of temporary files generated by EXP-T are listed and described below.

- HINT one-electron integrals: core Hamiltonian matrix + one-particle part of perturbation (if presented). Recommended for removal after calculation.
- VINT-\*-\*-\* two-electron integrals. These files are generated during the integral sorting stage. Recommended for removal after calculation.

Approximate number of files:  $\sim N_b^4$  ( $N_b$  – number of blocks into which the whole set of spinors is divided by symmetry and the tilesize parameter)

\*.sb a file containing single symmetry block of molecular integrals/cluster amplitudes. It is just a part of a diagram. During the calculation each file is mapped to some structure in RAM. Recommended for removal after calculation.

Approximate number of .sb files:  $\sim 10 \times N_b^4$  for CCSD,  $\sim N_b^6$  for CCSDT-n and CCSDT models  $\Rightarrow$  can be very large!

- \*.dg a file containing a diagram. Diagram files containing cluster amplitudes (t1c.dg, t2c.dg etc) can be used in subsequent calculations when constructing the initial approximation to amplitudes (see Section 4, keyword reuse).
- HEFF a formatted (text) file containing effective Hamiltonian matrices. These files can be read by the heffman.x utility program used to perform Padé extrapolation of series of effective Hamiltonians [9], finite-field transition moment calculations [58, 3], spin-orbit coupling calculations [5].
- \*DIPLEN formatted files with dipole moment integrals (just its electronic part) in the basis of molecular spinors.
- MVCOEF\*\* binary files containing model vectors expanded in the basis of model determinants. These files are used to organize data flow inside EXP-T in the most convenient and logical manner; model vectors are used for calculation of model-space approximations of TDMs, model natural spinors, properties within the finite-order approach, etc.
- NATORB\*.dat formatted text files containing model natural spinors expanded in the basis of molecular spinors. These expansions can be transformed to the AO basis with the expt\_transform\_natorb.x utility program (see Section 6.2).

#### density\_0h0p.dat

a formatted text file containing one-electron CC density matrix (in the basis of molecular spinors).

## Appendix D Symmetry and irreducible representations

This section lists the names of irreducible representations (irreps) used in EXP-T.

- For non-relativistic groups, an irrep is determined by the spatial symmetry (Mulliken notation is used) and the  $M_s$  value (for  $M_s >= 2$  only symbol 2 is used);
- For relativistic double groups we use the Mulliken-type notation (completely inherited from DIRAC);
- For  $U(1) = \overline{C_{\infty v}}$  and  $U(1) \times C_i = \overline{D_{\infty h}}$  irreps are determined by  $\Omega$  values.

 $C_1$ 

 $A_a$ 

A\_b

 $A_{-}3/2$ 

A\_+3/2

A\_0

 $A_4$ 

 $A_{+1}$ 

 $A_{-}1$ 

 $C_2$ 

A\_a B\_a

A\_b B\_b

A\_-3/2 B\_-3/2

A\_+3/2 B\_+3/2

A\_0 B\_0

A\_2 B\_2

A\_+1 B\_+1

A\_-1 B\_-1

 $C_s$ 

A'\_a A"\_a

A'\_b A"\_b

A'\_-3/2 A"\_-3/2

A'\_+3/2 A"\_+3/2

A'\_O

A"\_O"

A'\_2 A"\_2"

A'\_+1 A"\_+1

A'\_-1 A"\_-1

### $C_i$

Ag\_a Au\_a
Ag\_b Au\_b
Ag\_-3/2 Au\_-3/2
Ag\_+3/2 Au\_+3/2
Ag\_0 Au\_0
Ag\_2 Au\_2
Ag\_+1 Au\_+1
Ag\_-1 Au\_-1

#### $C_{2v}$

A1\_a B2\_a A2\_a B1\_a A1\_b B1\_b B2\_b A2\_b A1\_-3/2 B2\_-3/2 B1\_-3/2 A2\_-3/2 A1\_+3/2 B2\_+3/2 B1\_+3/2 A2\_+3/2 A1\_0 B2\_0 B1\_0 A2\_0 A1\_2 B2\_2 B1\_2 A2\_2 B2\_+1 B1\_+1 A2\_+1 A1\_+1 A1\_-1 B2\_-1 B1\_-1 A2\_-1

#### $C_{2h}$

Ag\_a Bg\_a Bu\_a  $Au_a$ Ag\_b Bg\_b Bu\_b Au\_b Ag\_-3/2 Bg\_-3/2 Bu\_-3/2 Au\_-3/2 Ag\_+3/2 Bg\_+3/2 Bu\_+3/2 Au\_+3/2 Ag\_0 Bg\_0 Bu\_0 Au\_0 Ag\_2 Bg\_2 Bu\_2 Au\_2  $Bu_+1$ Au\_+1 Ag\_+1 Bg\_+1 Bg\_-1  $Bu_--1$  $Au_-1$ Ag\_-1

#### $D_2$

 $A_a$ B3\_a B1\_a B2\_a  $A_b$ B3\_b B1\_b B2\_b A\_-3/2 B3\_-3/2 B1\_-3/2 B2\_-3/2 A\_+3/2 B3\_+3/2 B1\_+3/2 B2\_+3/2 A\_0 B3\_0 B1\_0 B2\_0 A\_2 B2\_2 B3\_2 B1\_2 A\_+1 B1\_+1 B2\_+1 B3\_+1 A\_-1 B3\_-1 B1\_-1 B2\_-1

```
D_{2h}
                                   Ag_a
                                                                             B1u_a
                                                                                                                             B2u_a
                                                                                                                                                                            B3g_a
                                                                                                                                                                                                                            B3u_a
                                                                                                                                                                                                                                                                           B2g_a
                                                                                                                                                                                                                                                                                                                            B1g_a
                                                                                                                                                                                                                                                                                                                                                                           Au_a
                                                                                                                             B2u_b
                                                                                                                                                                                                                            B3u_b
                                                                                                                                                                                                                                                                                                                            B1g_b
                                    Ag_b
                                                                             B1u_b
                                                                                                                                                                             B3g_b
                                                                                                                                                                                                                                                                           B2g_b
                                                                                                                                                                                                                                                                                                                                                                           Au_b
                                    \label{eq:ag_3/2} \verb| B1u_-3/2 | B2u_-3/2 | B3g_-3/2 | B3u_-3/2 | B2g_-3/2 | B1g_-3/2 | Au_-3/2 | B3g_-3/2 | 
                                    Ag_+3/2 B1u_+3/2 B2u_+3/2 B3g_+3/2 B3u_+3/2 B2g_+3/2 B1g_+3/2 Au_+3/2
                                    Ag_0
                                                                             B1u_0
                                                                                                                             B2u_0
                                                                                                                                                                             B3g_0
                                                                                                                                                                                                                            B3u_0
                                                                                                                                                                                                                                                                           B2g_0
                                                                                                                                                                                                                                                                                                                            B1g_0
                                                                                                                                                                                                                                                                                                                                                                           Au_0
                                    Ag_2
                                                                             B1u_2
                                                                                                                             B2u_2
                                                                                                                                                                             B3g_2
                                                                                                                                                                                                                            B3u_2
                                                                                                                                                                                                                                                                           B2g_2
                                                                                                                                                                                                                                                                                                                            B1g_2
                                                                                                                                                                                                                                                                                                                                                                           Au_2
                                    Ag_+1
                                                                             B1u_+1
                                                                                                                             B2u_+1
                                                                                                                                                                             B3g_+1
                                                                                                                                                                                                                            B3u_+1
                                                                                                                                                                                                                                                                           B2g_+1
                                                                                                                                                                                                                                                                                                                            B1g_+1
                                                                                                                                                                                                                                                                                                                                                                           Au_+1
                                    Ag_-1
                                                                             B1u_-1
                                                                                                                             B2u_-1
                                                                                                                                                                            B3g_-1
                                                                                                                                                                                                                            B3u_-1
                                                                                                                                                                                                                                                                           B2g_-1
                                                                                                                                                                                                                                                                                                                           B1g_-1
                                                                                                                                                                                                                                                                                                                                                                           Au_-1
\overline{C_1}
                                    Α
                                    a
\overline{C_2}
                                    1E 2E
                                    a b
\overline{C_s}
                                    1E 2E
                                    a b
\overline{C_i}
                                    AG AU
                                    ag au
\overline{C_{2v}}
                                    1E 2E
                                    a b
\overline{D_2}
                                    1E 2E
                                    a b
\overline{D_{2h}}
```

1Eg 2Eg 1Eu 2Eu ag bg au bu  $\overline{C_{\infty v}}$ 

```
1/2+ 1/2- 3/2+ 3/2- 5/2+ 5/2- 7/2+ 7/2-
 9/2+ 9/2- 11/2+ 11/2- 13/2+ 13/2- 15/2+ 15/2-
17/2+ 17/2- 19/2+ 19/2- 21/2+ 21/2- 23/2+ 23/2-
25/2+ 25/2- 27/2+ 27/2- 29/2+ 29/2- 31/2+ 31/2-
   0
         1+
               1-
                     2+
                           2-
                                  3+
                                        3-
   4-
                                        7-
         5+
               5-
                     6+
                           6-
                                  7+
                                              8+
         9+
               9-
                          10-
                                 11+
                                             12+
   8-
                    10+
                                       11-
  12-
        13+
              13-
                    14+
                          14-
                                 15+
                                       15-
                                             16+
```

 $\overline{D_{\infty h}}$ 

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