## photobiologyPlants Version 0.4.1 User Guide

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#### 1 Introduction

Package photobiologyPlants provides data sets and functions relevant to the photobiology of plants.

#### 2 Spectral data

See the User Guide for package photobiology for instructions on how to work with spectral data. In the examples below we use the solar spectral data included in package photobiology as a data frame in object sun.data.

### 3 Calculating the Phytochrome photoequilibrium

Photoequilibrium from a source\_spct object:

```
Pfr_Ptot(sun.spct)
## [1] 0.68341
```

Red:far-red ratio:

```
R_FR(sun.spct)

## Red.Smith10: FarRed.Smith10(q:q)
## 1.266704

## attr(,"radiation.unit")
## [1] "q:q ratio"
```

which is equivalent to:

```
q_ratio(sun.spct, Red("Smith"), Far_red("Smith"))

## Red.Smith10: FarRed.Smith10(q:q)
## 1.266704

## attr(,"radiation.unit")

## [1] "q:q ratio"
```

We can, and should whenever spectral data are available, calculate the photoequilibrium as above from them. It is possible to obtain and approximation in case of the solar spectrum and other broad spectra, using the red:far-red photon ratio. The calculation, however, is only strictly valid, for di-chromatic illumination with red plus far-red light.

```
Pfr_Ptot_R_FR(R_FR(sun.spct))
## Red.Smith10: FarRed.Smith10(q:q)
## 0.7051691
```

Here we calculated the R:FR ratio from spectral data, but in practice one would use this function only when spectral data is not available as when a R plus FR sensor is used. We can see that in such a case the photoequilibrium calculated is only a very rough approximation. For sunlight, in the example above when using spectral data we obtained a value of 0.683 in contrast to 0.705 when using the R:FR photon ratio. For other light sources differences can be much larger.

In the case of monochromatic light we can still use the same functions, as the defaults are such that we can use a single value as the 'w.length' argument, to obtain the Pfr:P ratio. For monochromatic light irradiance is irrelevant for the photoequilibrium.

```
Pfr_Ptot(660)
## [1] 0.869649

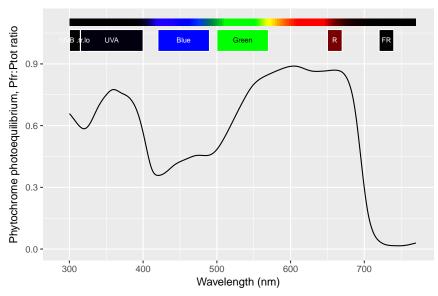
Pfr_Ptot(735)
## [1] 0.01749967

Pfr_Ptot(c(660, 735))
## [1] 0.86964902 0.01749967

Pfr_Ptot(435)
## [1] 0.3859998
```

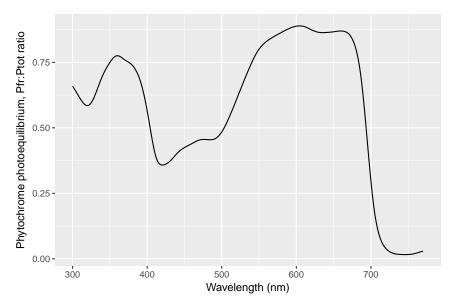
We can also plot Pfr:Ptot as a function of wavelength (nm) of monochromatic light. The default is to return a vector for short input vectors, and a response\_spct object otherwise, but this can be changed through argument spct.out.

```
plot(Pfr_Ptot(300:770), norm = NULL, unit.out = "photon",
    w.band = Plant_bands(),
    annotations = c("colour.guide", "labels", "boxes")) +
labs(y = "Phytochrome photoequilibrium, Pfr:Ptot ratio")
```



It is, of course, also possible to use base R plotting functions, or as shown here ggplot functions:

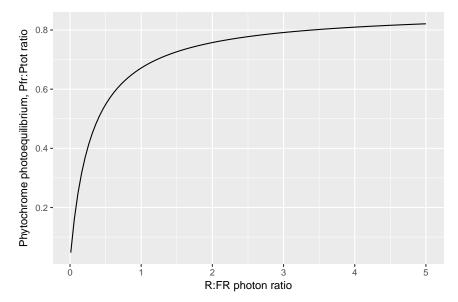
```
ggplot(data=Pfr_Ptot(300:770), aes(w.length, s.q.response)) +
  geom_line() +
  labs(x = "Wavelength (nm)",
    y = "Phytochrome photoequilibrium, Pfr:Ptot ratio")
```



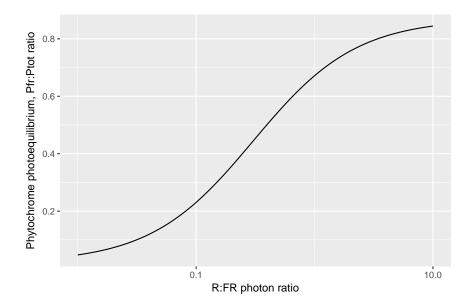
In the case of dichromatic illumination with red (660 nm) and far-red (730 nm) light, we can use a different function that takes the R:FR photon ratio as argument.

```
Pfr_Ptot_R_FR(1.15)
## [1] 0.6919699
Pfr_Ptot_R_FR(0.01)
## [1] 0.04747996
Pfr_Ptot_R_FR(c(1.15,0.01))
## [1] 0.69196990 0.04747996
```

Of course it is also easy to plot Pfr:P ratio as a function of R:FR photon ratio. However we have to remember that such values are exact only for dichromatic light, and only a rough approximation for wide-spectrum light sources. For wide-spectrum light sources, the photoequilibrium should, if possible, be calculated from spectral irradiance data.



Below we approximate figure 11 from Mancinelli (1994), where he uses a logarithmic scale for the R:FR ratio.



### 4 Calculating reaction rates

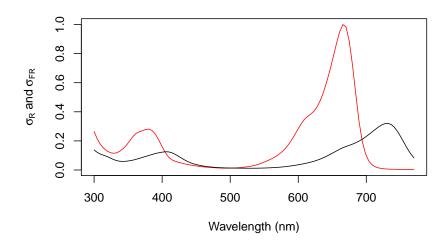
```
with(sun.data, Phy_reaction_rates(w.length, s.e.irrad))
## $k1
## [1] 1.259332
##
## $k2
## [1] 0.5833862
##
## $nu
## [1] 1.842718
```

# 5 Calculating the absorption cross section at given wavelengths

The phytochrome photoequilibrium cannot be calculated from the absorptance spectra of Pr and Pfr, because Pr and Pfr have different quantum yields for the respective phototransformations. We need to use action spectra, which in this

context are usually called 'absorption cross-sections'. They can be calculated as the product of absorptance and quantum yield. The values in these spectra, in the case of Phy are called 'Sigma'.

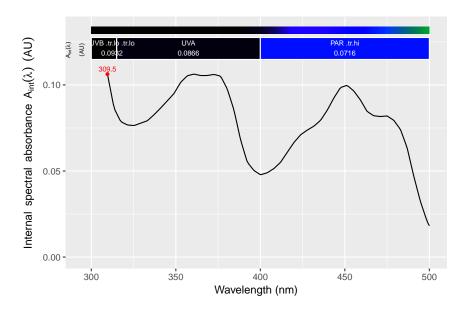
Here we reproduce Figure 3 in Mancinelli (1994), which gives the 'Relative photoconversion cross-sections' of Pr  $(\sigma_R)$  and Pfr  $(\sigma_{FR})$ . The values are expressed relative to  $\sigma_R$  at its maximum at  $\lambda = 666$  nm.



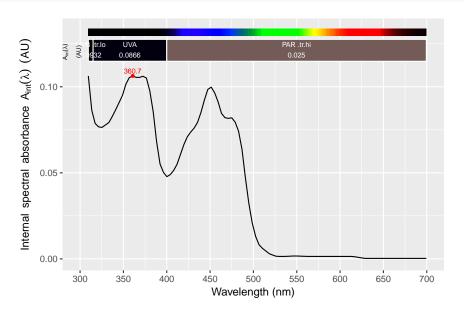
# 6 Calculating the CRY absorbance at given wavelengths

Here we approximate Figure 1.B from Banerjee et al. (2007).

```
plot(interpolate_wl(CRY2.mspct$dark_adapted, 300:500),
    plot.qty = "absorbance")
```







## 7 UVR8



