R for Photobiology: a handbook

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Part I Getting ready

CHAPTER

Introduction

Abstract

In this chapter we explain the physical basis of optics and photochemisatry.

1.1 Radiation and molecules

Part II

Cookbook

Unweighted irradiance

Abstract

In this chapter we explain how to calculate unweighted energy and photon irradiances from spectral irradiance.

2.1 Task: calculate (energy) irradiance

The task to be completed is to calculate the (energy) irradiance (I) in W m⁻² from spectral (energy) irradiance (I_{λ}) in W m⁻² nm⁻¹ and the corresponding wavelengths (λ) in nm.

$$I_{\lambda_1 < \lambda < \lambda_2} = \int_{\lambda_1}^{\lambda_2} I_{\lambda} \, d\lambda$$
 (2.1)

Let's assume that we want to calculate photosynthetically active radiation (PAR) energy irradiance. In this example we will use example data for sunlight.

```
with(sun.data, energy_irradiance(w.length, s.e.irrad, PAR()))
## PAR
## 196.7
```

PAR() is predefined in package photobiologyVIS as a convenience function. It is also possible to define our own range of wavelengths or *waveband* to use for the intergartion. If $\lambda_1 = 380$ nm and $\lambda_2 = 780$ nm we can use the following code to calculate $I_{380\,\mathrm{nm}<\lambda<780\,\mathrm{nm}}$:

2.2 Task: calculate photon irradiance

The task to be completed is to calculate the photon irradiance (Q) in $\,\mathrm{mol}\,\mathrm{m}^{-2}\,\mathrm{s}^{-1}$ from spectral (energy) irradiance (I_{λ}) in $\,\mathrm{W}\,\mathrm{m}^{-2}\,\mathrm{nm}^{-1}$ and the corresponding wavelengths (λ) in nm.

The energy of a quantum of radiation in a vacuum, q, depends on the wavelength, λ , or frequency¹, ν ,

$$q = h \cdot v = h \cdot \frac{c}{\lambda} \tag{2.2}$$

with the Planck constant $h = 6.626 \times 10^{-34}$ Js and speed of light in vacuum $c = 2.998 \times 10^8$ m s⁻¹. When dealing with numbers of photons, the equation (2.2) can be extended by using Avogadro's number $N_A = 6.022 \times 10^{23}$ mol⁻¹. Thus, the energy of one mole of photons, q', is

$$q' = h' \cdot \nu = h' \cdot \frac{c}{\lambda} \tag{2.3}$$

with $h'=h\cdot N_{\rm A}=3.990\times 10^{-10}~{\rm J\,s\,mol^{-1}}$. Example 1: red light at 600 nm has about 200 kJ mol⁻¹, therefore, 1 µmol photons has 0.2 J. Example 2: UV-B radiation at 300 nm has about 400 kJ mol⁻¹, therefore, 1 µmol photons has 0.4 J. Equations 2.2 and 2.3 are valid for all kinds of electromagnetic waves.

Combining equations 2.1 and 2.3 we obtain:

$$Q_{\lambda_1 < \lambda < \lambda_2} = \int_{\lambda_1}^{\lambda_2} I_{\lambda} \, \frac{h' \cdot c}{\lambda} \, \mathrm{d} \, \lambda \tag{2.4}$$

Let's assume that we want to calculate photosynthetically active radiation (PAR) photon irradiance. In this example we will use example data for sunlight.

```
with(sun.data, photon_irradiance(w.length, s.e.irrad, PAR()))
## PAR
## 0.0008938
```

If we want to have Q_{PAR} expressed in the usual units of μ mol m⁻² s⁻¹, we need to multiply the result above by 10⁶:

 $^{^1}$ Wavelength and frequency are related to each other by the speed of light, according to $\nu=c/\lambda$ where c is speed of light in vacuum. Consequently there are two equivalent formulations for equation 2.2.

2.3. TASK: CALCULATE ENERGY AND PHOTON IRRADIANCES FROM SPECTRAL PHOTON IRRADIANCE 9

```
with(sun.data, photon_irradiance(w.length, s.e.irrad, PAR())) *
    1e+06

## PAR
## 893.8
```

PAR() is predefined in package photobiologyVIS as a convenience function, see section 2.1 for an example with arbitrary values for λ_1 and λ_2 .

2.3 Task: calculate energy and photon irradiances from spectral photon irradiance

In the case of the calcualtion of energy irradiance from spectral photon irradiance the calculation is:

$$I_{\lambda_1 < \lambda < \lambda_2} = \int_{\lambda_1}^{\lambda_2} Q_{\lambda} \, \frac{\lambda}{h' \cdot c} \mathrm{d} \, \lambda \tag{2.5}$$

And the code²:

```
with(sun.data, energy_irradiance(w.length, s.q.irrad, PAR()),
    unit.in = "photon")

## PAR
## 0.0008938
```

The calculation of photon irradiance from spectral photon irradiance, is a simple integration, analogous to that in equation 2.1, and the code is:

```
with(sun.data, photon_irradiance(w.length, s.q.irrad, PAR()),
    unit.in = "photon")

## PAR
## 4.158e-09
```

2.4 Task: calculate irradiances for more than one waveband

It is possible to calculate the irradiances for several wavebands with a single function call by suppying a list of wavebands as argument:

 $^{^2}$ The dataframe sun.data contains both spectral energy irradiance vales in 'column' s.e.irrad and spectral photon irradiance in 'column' s.q.irrad

```
Q.RGB <- with(sun.data, photon_irradiance(w.length, s.e.irrad,
    list(Red(), Green(), Blue()))) * 1e+06
signif(Q.RGB, 3)
##
     Red.ISO Green.ISO Blue.ISO
             220
##
        452
                           149
Q.RGB[1]
## Red.ISO
##
    452.2
Q.RGB["Green.ISO"]
## Green.ISO
##
   220.2
```

A named list can be used to override the use as names for the output of the waveband names:

Even when using a single waveband:

2.5 Task: use wavebands

Please, consult the packages' documentation for a list of predefined functions for creating wavebands. Here we will present just a few examples of their use. We usually associate wavebands with colours, however, in many cases there are different definitions in use. For this reason, the functions provided accept an argument that can be used to select the definition to use. In general, the default, is to use the ISO standard whenever it is applicable. Towards the end of this section we will consider here the case of UV-B radiation.

We can use a predifined function to create a new waveband object, which as any other R object can be assigned to a variable:

```
uvb <- UVB()
uvb
```

```
## UVB.ISO
## low (nm) 280
## high (nm) 315
```

As seen above, there is a specialized print function for wavebands. Functions available are min, max, range, center_wl, labels, and color.

```
red <- Red()
red

## Red.ISO
## low (nm) 610
## high (nm) 760

range(red)

## [1] 610 760

labels(red)

## $label
## [1] "Red.ISO"

center_wl(red)

## [1] 685

color(red)

## Red.ISO CMF Red.ISO CC
## "#900000" "#FF0000"</pre>
```

Here we demonstrate the use of an argument to choose a certain definition:

```
UVB()
## UVB.ISO
## low (nm) 280
## high (nm) 315

UVB("ISO")
## UVB.ISO
## low (nm) 280
## high (nm) 315

UVB("CIE")
## UVB.CIE
## low (nm) 280
## high (nm) 315
UVB("CIE")

## UVB.CIE
## low (nm) 280
## high (nm) 315
```

```
## UVB.medical
## low (nm) 290
## high (nm) 320

UVB("none")

## UVB.none
## low (nm) 280
## high (nm) 320
```

Here we demonstrate the importance of complying with standards, and how much the photon irradiance calculated can depend on the definition used.

2.6 Task: define wavebands

Here we briefly introduce new_waveband, and only in chapter ?? we describe it in full, including the use of sspectral weighting functions (SWFs).

Definig a new waveband based on extreme wavelengths expressed in nm.

```
with(sun.data, photon_irradiance(w.length, s.e.irrad, wb2)) *
    le+06

## my.colour
## 314.1
```

2.7 Task: calculate photon ratio

In photobiology sometimes we are interested in calculation the photon ratio beween two wavebands.