Computational Physics

Homework 6

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1.

(a). Using Von Neumann stability analysis to show implicit FTSC is stable Implicit FTSC in Schrödinger equation is

$$\begin{split} \frac{i\hbar}{\tau} (\psi^{n+1} - \psi^n) &= H \psi^{n+1} \\ H &= -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V \\ \frac{i\hbar}{\tau} (\psi^{n+1}_j - \psi^n_j) &= \left(-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V \right) \psi^{n+1}_j \\ &= -\left(\frac{\hbar^2}{2m} \right) \frac{\psi^{n+1}_{j+1} + \psi^{n+1}_{j-1} - 2\psi^{n+1}_j}{h^2} + V \psi^{n+1}_j \end{split}$$

Von Neumann analysis is try to put $a_j^n = A^n e^{ikjh}$ into the function, and $x_j = jh$, $t_n = (n-1)\tau$. Therefore, we put it into Schrödinger equation.

$$\frac{i\hbar}{\tau} \left(A^{n+1} e^{ikjh} - A^n e^{ikjh} \right) = - \left(\frac{\hbar^2}{2m} \right) \frac{A^{n+1} e^{ik(j+1)h} + A^{n+1} e^{ik(j-1)h} - 2A^{n+1} e^{ikjh}}{h^2} + VA^{n+1} e^{ikjh}$$

We suppose that $\xi = \frac{A^{n+1}}{A^n}$, and if $|\xi| > 1$, it means that the function is unstable.

$$\frac{i\hbar}{\tau}(A^{n+1} - A^n) = -\left(\frac{\hbar^2}{2m}\right) \frac{A^{n+1}e^{ikh} + A^{n+1}e^{-ikh} - 2A^{n+1}}{h^2} + VA^{n+1}$$

$$\frac{i\hbar}{\tau}\left(1 - \frac{1}{\xi}\right) = -\left(\frac{\hbar^2}{2m}\right) \frac{e^{ikh} + e^{-ikh} - 2}{h^2} + V$$

$$\frac{1}{\xi} = \left(\frac{\hbar\tau}{2im}\right) \frac{e^{ikh} + e^{-ikh} - 2}{h^2} - \frac{\tau V}{i\hbar} + 1$$

$$\frac{1}{\xi} = \left(-i\frac{\hbar\tau}{mh^2}\right) \cos(kh) + i\frac{\hbar\tau}{mh^2} + i\frac{\tau V}{\hbar} + 1$$

$$\frac{1}{\xi} = i\left(\frac{\hbar\tau}{mh^2}\right) (1 - \cos(kh)) + \frac{\tau V}{\hbar} + 1$$

$$\frac{1}{|\xi|} = \sqrt{\left(\frac{\hbar\tau}{mh^2}(1 - \cos(kh)) + \frac{\tau V}{\hbar}\right)^2 + 1}$$

$$|\xi| = \frac{1}{\sqrt{\frac{\hbar\tau}{mh^2}(1 - \cos(kh)) + \frac{\tau V}{\hbar}}^2 + 1} < 1$$

Therefore, implicit FTSC in Schrödinger equation is stable.

(b). Using Tylor series to show the accuracy of Crank-Nicolson scheme.

Crank-Nicolson scheme is below:

$$\psi^{n+1} = \left(\frac{1 - \frac{i\tau}{2\hbar}H}{1 + \frac{i\tau}{2\hbar}H}\right)\psi^{n}$$

$$\psi^{n+1} + \frac{i\tau}{2\hbar}\left(-\frac{\hbar^{2}}{2m}\frac{\partial^{2}}{\partial x^{2}} + V\right)\psi^{n+1} = \psi^{n} - \frac{i\tau}{2\hbar}\left(-\frac{\hbar^{2}}{2m}\frac{\partial^{2}}{\partial x^{2}} + V\right)\psi^{n}$$

$$\psi^{n+1}_{j} - \left(\frac{i\tau\hbar}{4m}\right)\frac{\psi^{n+1}_{j+1} + \psi^{n+1}_{j-1} - 2\psi^{n+1}_{j}}{h^{2}} + \frac{i\tau}{2\hbar}V\psi^{n+1}_{j} = \psi^{n} + \left(\frac{i\tau\hbar}{4m}\right)\frac{\psi^{n}_{j+1} + \psi^{n}_{j-1} - 2\psi^{n}_{j}}{h^{2}} - \frac{i\tau}{2\hbar}V\psi^{n}_{j} \dots (1)$$

Tylor series can be express by these terms:

$$\begin{split} \psi_{j}^{n+1} &= \psi_{j}^{n} + \Delta t \left(\frac{\partial \psi_{j}^{n}}{\partial t} \right) + \frac{\Delta t^{2}}{2} \left(\frac{\partial^{2} \psi_{j}^{n}}{\partial t^{2}} \right) + \frac{\Delta t^{3}}{6} \left(\frac{\partial^{3} \psi_{j}^{n}}{\partial t^{3}} \right) + \frac{\Delta t^{4}}{24} \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial t^{4}} \right) + O\left(\Delta t^{5}\right) \\ \psi_{j+1}^{n} &= \psi_{j}^{n} + \Delta x \left(\frac{\partial \psi_{j}^{n}}{\partial x} \right) + \frac{\Delta x^{2}}{2} \left(\frac{\partial^{2} \psi_{j}^{n}}{\partial x^{2}} \right) + \frac{\Delta x^{3}}{6} \left(\frac{\partial^{3} \psi_{j}^{n}}{\partial x^{3}} \right) + \frac{\Delta x^{4}}{24} \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial x^{4}} \right) + O\left(\Delta x^{5}\right) \\ \psi_{j+1}^{n+1} &= \psi_{j}^{n} + \Delta x \left(\frac{\partial \psi_{j}^{n}}{\partial x} \right) + \frac{\Delta x^{2}}{2} \left(\frac{\partial^{2} \psi_{j}^{n}}{\partial x^{2}} \right) + \frac{\Delta x^{3}}{6} \left(\frac{\partial^{3} \psi_{j}^{n}}{\partial x^{3}} \right) + \frac{\Delta x^{4}}{24} \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial x^{4}} \right) + \Delta t \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial t^{2}} \right) + \frac{\Delta t^{2}}{2} \left(\frac{\partial^{2} \psi_{j}^{n}}{\partial t^{2}} \right) \\ &+ \frac{\Delta t^{3}}{6} \left(\frac{\partial^{3} \psi_{j}^{n}}{\partial t^{3}} \right) + \frac{\Delta t^{4}}{24} \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial t^{4}} \right) + \Delta x \Delta t \left(\frac{\partial^{2} \psi_{j}^{n}}{\partial x \partial t} \right) + \Delta x^{2} \Delta t \left(\frac{\partial^{3} \psi_{j}^{n}}{\partial x^{2} \partial t} \right) + \Delta x \Delta t^{2} \left(\frac{\partial^{3} \psi_{j}^{n}}{\partial x \partial t^{2}} \right) \\ &+ \Delta x^{3} \Delta t \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial x^{3} \partial t} \right) + \Delta x^{2} \Delta t^{2} \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial x^{2} \partial t^{2}} \right) + \Delta x \Delta t^{3} \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial x \partial t^{3}} \right) \end{split}$$

Therefore, we can bring these terms into equation(1). On the left hand side:

$$\begin{split} LHS = \ \psi_{j}^{n} + \Delta t \left(\frac{\partial \psi_{j}^{n}}{\partial t} \right) + \frac{\Delta t^{2}}{2} \left(\frac{\partial^{2} \psi_{j}^{n}}{\partial t^{2}} \right) + \frac{\Delta t^{3}}{6} \left(\frac{\partial^{3} \psi_{j}^{n}}{\partial t^{3}} \right) + \frac{\Delta t^{4}}{24} \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial t^{4}} \right) \\ - \left(\frac{i\tau\hbar}{4mh^{2}} \right) \left[\Delta x^{2} \left(\frac{\partial^{2} \psi_{j}^{n}}{\partial x^{2}} \right) + \frac{\Delta x^{4}}{12} \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial x^{4}} \right) + 2\Delta x^{2} \Delta t \left(\frac{\partial^{3} \psi_{j}^{n}}{\partial x^{2} \partial t} \right) + 2\Delta x^{2} \Delta t^{2} \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial x^{2} \partial t^{2}} \right) \right] \\ + \frac{i\tau}{2\hbar} V \left[\psi_{j}^{n} + \Delta t \left(\frac{\partial \psi_{j}^{n}}{\partial t} \right) + \frac{\Delta t^{2}}{2} \left(\frac{\partial^{2} \psi_{j}^{n}}{\partial t^{2}} \right) + \frac{\Delta t^{3}}{6} \left(\frac{\partial^{3} \psi_{j}^{n}}{\partial t^{3}} \right) + \frac{\Delta t^{4}}{24} \left(\frac{\partial^{4} \psi_{j}^{n}}{\partial t^{4}} \right) \right] \end{split}$$

On the right hand side:

RHS =
$$\psi_j^n + \left(\frac{i\tau\hbar}{4mh^2}\right) \left[\Delta x^2 \left(\frac{\partial^2 \psi_j^n}{\partial x^2}\right) + \frac{\Delta x^4}{12} \left(\frac{\partial^4 \psi_j^n}{\partial x^4}\right)\right] - \frac{i\tau}{2\hbar} V \psi_j^n$$

Also, $\Delta t = \tau$, $\Delta x = h$:

$$\begin{split} 0 &= \tau \left(\frac{\partial \psi_j^n}{\partial t} \right) + \frac{\tau^2}{2} \left(\frac{\partial^2 \psi_j^n}{\partial t^2} \right) + \frac{\tau^3}{6} \left(\frac{\partial^3 \psi_j^n}{\partial t^3} \right) + \frac{\tau^4}{24} \left(\frac{\partial^4 \psi_j^n}{\partial t^4} \right) \\ &- \left(\frac{i\tau\hbar}{4m} \right) \left[2 \left(\frac{\partial^2 \psi_j^n}{\partial x^2} \right) + \frac{h^2}{6} \left(\frac{\partial^4 \psi_j^n}{\partial x^4} \right) + 2\tau \left(\frac{\partial^3 \psi_j^n}{\partial x^2 \partial t} \right) + 2\tau^2 \left(\frac{\partial^4 \psi_j^n}{\partial x^2 \partial t^2} \right) \right] + \frac{i\tau}{2\hbar} V[2\psi_j^n] \\ &+ \tau \left(\frac{\partial \psi_j^n}{\partial t} \right) + \frac{\tau^2}{2} \left(\frac{\partial^2 \psi_j^n}{\partial t^2} \right) + \frac{\tau^3}{6} \left(\frac{\partial^3 \psi_j^n}{\partial t^3} \right) + \frac{\tau^4}{24} \left(\frac{\partial^4 \psi_j^n}{\partial t^4} \right) \right] \end{split}$$

$$\begin{split} &=-i\hbar\left(\frac{\partial\psi_{j}^{n}}{\partial t}\right)+\frac{-i\hbar\tau}{2}\left(\frac{\partial^{2}\psi_{j}^{n}}{\partial t^{2}}\right)+\frac{-i\hbar\tau^{2}}{6}\left(\frac{\partial^{3}\psi_{j}^{n}}{\partial t^{3}}\right)+\frac{-i\hbar\tau^{3}}{24}\left(\frac{\partial^{4}\psi_{j}^{n}}{\partial t^{4}}\right)\\ &-\left(\frac{\hbar^{2}}{4m}\right)\left[2\left(\frac{\partial^{2}\psi_{j}^{n}}{\partial x^{2}}\right)+\frac{\mathbf{h}^{2}}{6}\left(\frac{\partial^{4}\psi_{j}^{n}}{\partial x^{4}}\right)+2\tau\left(\frac{\partial^{3}\psi_{j}^{n}}{\partial x^{2}\partial t}\right)+2\tau^{2}\left(\frac{\partial^{4}\psi_{j}^{n}}{\partial x^{2}\partial t^{2}}\right)\right]+\frac{1}{2}V[2\psi_{j}^{n}+\tau\left(\frac{\partial\psi_{j}^{n}}{\partial t}\right)\\ &+\frac{\tau^{2}}{2}\left(\frac{\partial^{2}\psi_{j}^{n}}{\partial t^{2}}\right)+\frac{\tau^{3}}{6}\left(\frac{\partial^{3}\psi_{j}^{n}}{\partial t^{3}}\right)+\frac{\tau^{4}}{24}\left(\frac{\partial^{4}\psi_{j}^{n}}{\partial t^{4}}\right)] \end{split}$$

So,

$$\begin{split} i\hbar\left(\frac{\partial\psi_{j}^{n}}{\partial t}\right) + \left(\frac{\hbar^{2}}{2m}\right)\left(\frac{\partial^{2}\psi_{j}^{n}}{\partial x^{2}}\right) - V\psi_{j}^{n} &= 0 \\ &= \frac{-i\hbar\tau}{2}\left(\frac{\partial^{2}\psi_{j}^{n}}{\partial t^{2}}\right) + \frac{-i\hbar\tau^{2}}{6}\left(\frac{\partial^{3}\psi_{j}^{n}}{\partial t^{3}}\right) + \frac{-i\hbar\tau^{3}}{24}\left(\frac{\partial^{4}\psi_{j}^{n}}{\partial t^{4}}\right) \\ &- \left(\frac{\hbar^{2}}{4m}\right)\left[\frac{h^{2}}{6}\left(\frac{\partial^{4}\psi_{j}^{n}}{\partial x^{4}}\right) + 2\tau\left(\frac{\partial^{3}\psi_{j}^{n}}{\partial x^{2}\partial t}\right) + 2\tau^{2}\left(\frac{\partial^{4}\psi_{j}^{n}}{\partial x^{2}\partial t^{2}}\right)\right] \\ &+ \frac{1}{2}V\left[\tau\left(\frac{\partial\psi_{j}^{n}}{\partial t}\right) + \frac{\tau^{2}}{2}\left(\frac{\partial^{2}\psi_{j}^{n}}{\partial t^{2}}\right) + \frac{\tau^{3}}{6}\left(\frac{\partial^{3}\psi_{j}^{n}}{\partial t^{3}}\right) + \frac{\tau^{4}}{24}\left(\frac{\partial^{4}\psi_{j}^{n}}{\partial t^{4}}\right)\right] \\ &\approx \frac{-i\hbar\tau}{2}\left(\frac{\partial^{2}\psi_{j}^{n}}{\partial t^{2}}\right) - \frac{(\hbar\hbar)^{2}}{24m}\left(\frac{\partial^{4}\psi_{j}^{n}}{\partial x^{4}}\right) + \frac{V\tau}{2}\left(\frac{\partial\psi_{j}^{n}}{\partial t}\right) + O(\tau^{3}) + O(\hbar^{5}) \end{split}$$

We can see that Crank-Nicolson scheme accurate to the second order.

(2). Find the retarded time

We suppose that the charge orbit is $\vec{r_1}(t) = (ct, 0, v_0 t)$, and the observer is on $\vec{r}(t) = (ct, 1, r_v)$.

The distance between is:

$$R^{\mu} = \vec{r}(t) - \overrightarrow{r_1}(t_{ret}) = (c(t - t_{ret}), 1, r_v - v_0 t_{ret})$$

We know electromagnetic field propagate by speed of light c = 1, so R is lightlike.

$$R^{\mu}R_{\mu} = 0 = -(t - t_{ret})^2 + 1 + (r_{\nu} - v_0 t_{ret})^2$$

We rewrite the formula, and $v_0 = \frac{1}{2}$.

$$\frac{3}{4}t_{ret}^2 + (r_y - 2t)t_{ret} + t^2 - r_y^2 - 1 = 0$$

$$t_{ret} = \frac{2}{3} \left((2t - r_y) \pm \sqrt{(2t - r_y)^2 + 3(1 + r_y^2 - t^2)} \right)$$

There is a constraint that $t_{ret} < t$, for we cannot receive the message from future.

In another way, we also use Newton method to find retarded time. Newton method is that to use first derivation to get closer the real answer. We suppose x_{ans} is real answer of f(x), and x_i is the answer what we guess.

$$f(x_{ans}) = 0 = f(x_0) - \delta x_0 f'(x_0)$$
$$\delta x_0 = \frac{f(x_0)}{f'(x_0)}$$
$$x_1 = x_0 - \delta x_0$$

This process will continue until $\delta x_i < \Delta_{accuracy}$. Therefore, in this case, δt_{ret} will become this terms:

$$\delta t_{ret} = \frac{\frac{3}{4}t_{ret}^2 + \left(r_y - 2t\right)t_{ret} + t^2 - r_y^2 - 1}{\frac{3}{2}t_{ret} + r_y - 2t}$$

First, I create the two functions in the code.

```
import matplotlib.pyplot as plt
import math
acc = 0.01 #In the second question, the answer of accuracy are 0.01

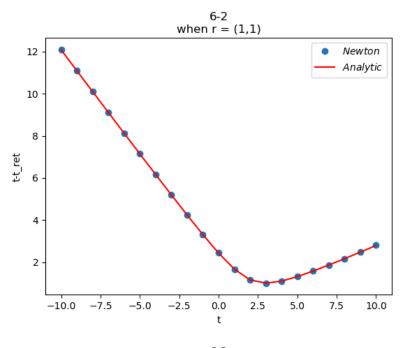
def analytic(ry,t): #analytic solution
    ans = ((2*t-ry)+math.sqrt(4*pow(ry,2)+pow(t,2)-4*ry*t+3))/1.5
    if ans >t: #retarded time cannot be bigger than observer time
        ans = ((2 * t - ry) - math.sqrt(4 * pow(ry, 2) + pow(t, 2) - 4 * ry * t + 3)) / 1.5
    return ans

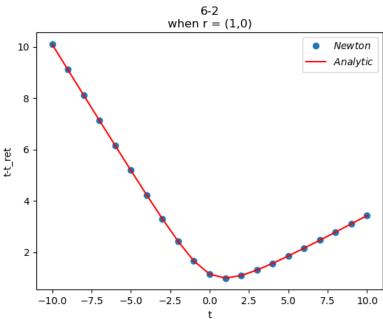
def delta(ry,t,gu): #Newton method; This function is create the delta
    ans = (0.75*pow(gu,2)+(ry-2*t)*gu+pow(t,2)-pow(ry,2)-1)/(1.5*gu+ry-2*t)
    return ans
```

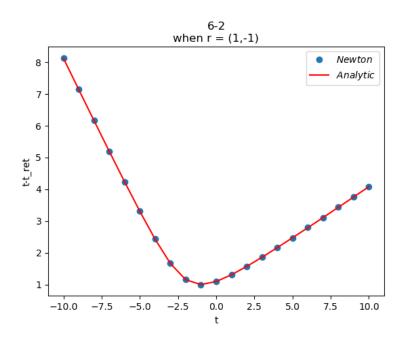
Next, we will try to plot the retarded time from t=-10 $^{\sim}$ 10, with condition r_y =1,0,-1.

```
test_3()
```

Result:







(3). Charge in circular motion

(a) Calculating retarded time

In this question, the condition is following: $q=1, c=1, \omega=0.1 \ or \ 0.9$, and charge orbit $\overrightarrow{r_q}(t)=(\cos(\omega t)$, $\sin(\omega t))$, calculating in the grid $\overrightarrow{r_{i,j}}=\left(-10+0.4(i-1),-10+0.4(j-1)\right)$ at t = 0.

We need to calculate δx because we will use Newton method to solve this question. First, we need to find the distance R:

$$R = \overrightarrow{r_{t,j}} - \overrightarrow{r_q}(t) = (-t_{ret}, r_x - \cos(\omega t_{ret}), r_y - \sin(\omega t_{ret}))$$

R is lightlike, so:

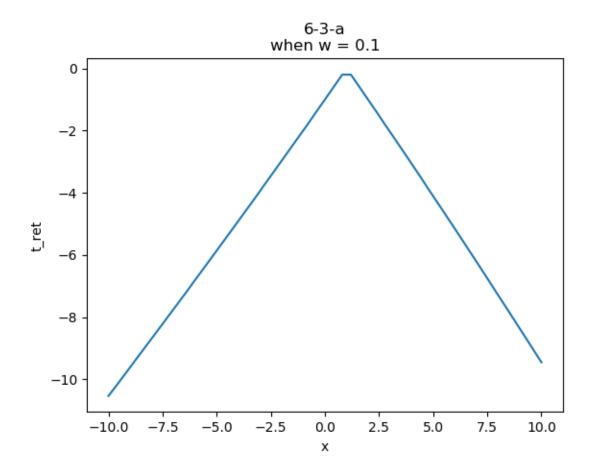
$$R^{\mu}R_{\mu} = -t_{ret}^{2} + (r_{x} - \cos(\omega t_{ret}))^{2} + (r_{y} - \sin(\omega t_{ret}))^{2} = 0$$
$$= -t_{ret}^{2} + r_{x}^{2} + r_{y}^{2} - 2r_{x}\cos(\omega t_{ret}) - 2r_{y}\sin(\omega t_{ret})$$

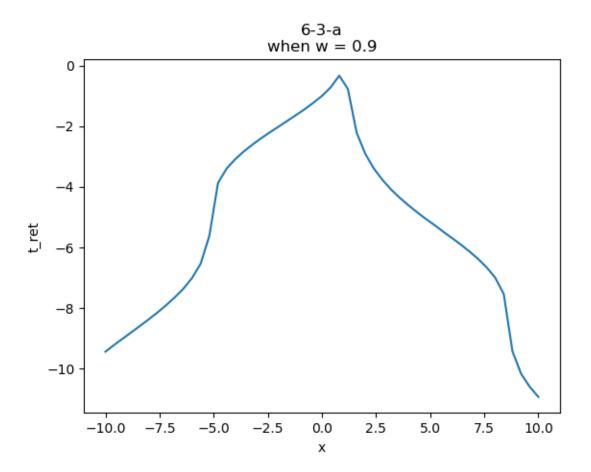
And the delta is:

$$\delta t_{ret} = \frac{-t_{ret}^2 + r_x^2 + r_y^2 - 2r_x \cos(\omega t_{ret}) - 2r_y \sin(\omega t_{ret})}{-2t_{ret} + 2r_x \omega \sin(\omega t_{ret}) - 2r_y \omega \cos(\omega t_{ret})}$$

```
s = math.sin(w*gu)
c = math.cos(w*gu)
ans = (pow(gu,2)+2*x*c+2*y*s-pow(x,2)-pow(y,2)-1)/(2*gu-2*w*x*s+2*w*y*c)
return ans
```

Result:





(b) Finding the potential

The retarded potential is like this:

$$A_{\mu}(x) = -\frac{u_{\mu}(t_{ret})}{R \cdot u(t_{ret})}$$

$$u_{\mu}(t_{ret}) = \frac{1}{\gamma} \left(\frac{\partial \overrightarrow{r_q}}{\partial t}\right)_{t=t_{ret}}$$

$$= (1 - \omega^2)(1, -\omega \sin(\omega t_{ret}), \omega \cos(\omega t_{ret}))$$

$$R \cdot u(t_{ret}) = (1 - \omega^2)(t_{ret} - r_x \omega \sin(\omega t_{ret}) + r_y \omega \cos(\omega t_{ret}))$$

Therefore,

$$A_{\mu}(x) = \frac{(-1, \omega \sin(\omega t_{ret}), -\omega \cos(\omega t_{ret}))}{t_{ret} - r_x \omega \sin(\omega t_{ret}) + r_y \omega \cos(\omega t_{ret})}$$

```
gu -= degu
```

```
plt.quiver(Ax,Ay)

plt.xticks([0, 5, 10, 15, 20, 25, 30, 35, 40, 45, 50], [-10, -8, -6, -4, -2, 0, 2, 4, 6, 8, 10])

plt.yticks([0, 5, 10, 15, 20, 25, 30, 35, 40, 45, 50], [-10, -8, -6, -4, -2, 0, 2, 4, 6, 8, 10])

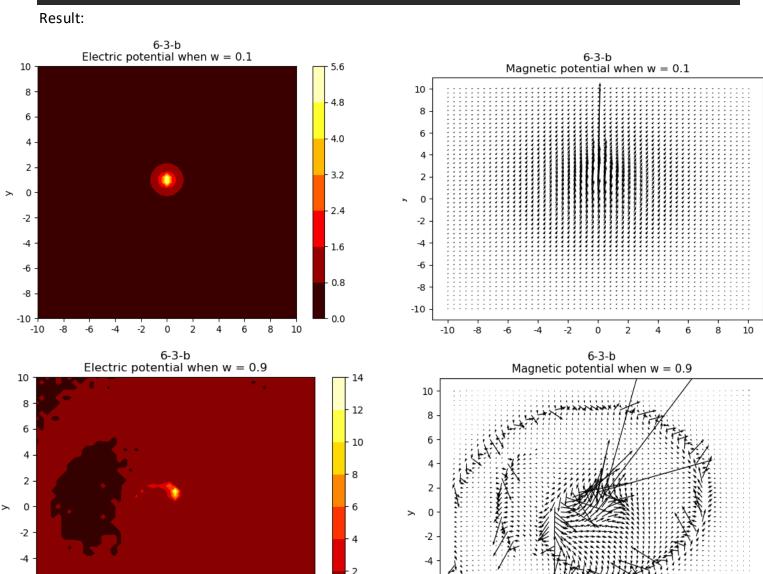
plt.title("6-3-b\nMagnetic potential when w = 0.9")

plt.xlabel("x")

plt.ylabel("y")

plt.show()

test_4()
```



-8

-10

-10

10

0

6

0

8

10

-6

-8

-10 - -10

(c). Calculating retarded electric field

Retarded electric field

$$\vec{E} = \underbrace{\left(\frac{1}{\gamma^{2}} \frac{(\hat{R} - \beta)}{R^{2} (1 - \beta \cdot \hat{R})^{3}}\right)}_{Comlomb (near field)} + \underbrace{\frac{\hat{R} \times \left[(\hat{R} - \beta) \times \dot{\beta}\right]}{|R|(1 - \beta \cdot \hat{R})^{3}}}_{Radiation (far field)}$$

$$= \frac{1}{\gamma^{2}} \frac{(\hat{R} - \beta)}{R^{2} (1 - \beta \cdot \hat{R})^{3}} + \frac{(\hat{R} \cdot \dot{\beta})(\hat{R} - \beta)}{|R|(1 - \beta \cdot \hat{R})^{3}} - \frac{\dot{\beta}}{|R|(1 - \beta \cdot \hat{R})^{2}}$$

And we know that $|R|=t_{ret}$, $\frac{1}{\gamma^2}=1-\omega^2$, $\beta=\gamma u$.

So,

$$\begin{split} \hat{R} &= \frac{R}{|R|} = (\frac{r_x - \cos(\omega t_{ret})}{t_{ret}}, \frac{r_y - \sin(\omega t_{ret})}{t_{ret}}) \\ \dot{\beta} &= (-\omega^2 \cos(\omega t_{ret}), -\omega^2 \sin(\omega t_{ret})) \\ 1 - \beta \cdot \hat{R} &= 1 - \frac{\omega}{t_{ret}} (-r_x \sin(\omega t_{ret}) + r_y \cos(\omega t_{ret})) \\ \hat{R} - \beta &= (\frac{r_x - \cos(\omega t_{ret})}{t_{ret}} - \omega \sin(\omega t_{ret}), \frac{r_y - \sin(\omega t_{ret})}{t_{ret}} + \omega \cos(\omega t_{ret})) \end{split}$$

Therefore, Electric field is following

$$\begin{split} E_{\chi} &= \frac{(1-\omega^2) \left(\frac{r_{\chi} - \cos(\omega t_{ret})}{t_{ret}} - \omega \sin(\omega t_{ret})\right)}{t_{ret}^2 \left(1 - \frac{\omega}{t_{ret}} \left(-r_{\chi} \sin(\omega t_{ret}) + r_{y} \cos(\omega t_{ret})\right)\right)^3} \\ &+ \frac{\left(\frac{\omega}{t_{ret}} \left(-r_{\chi} \sin(\omega t_{ret}) + r_{y} \cos(\omega t_{ret})\right)\right) \left(\frac{r_{\chi} - \cos(\omega t_{ret})}{t_{ret}} - \omega \sin(\omega t_{ret})\right)}{t_{ret} \left(1 - \frac{\omega}{t_{ret}} \left(-r_{\chi} \sin(\omega t_{ret}) + r_{y} \cos(\omega t_{ret})\right)\right)^3} \\ &+ \frac{\omega^2 \cos(\omega t_{ret})}{t_{ret} \left(1 - \frac{\omega}{t_{ret}} \left(-r_{\chi} \sin(\omega t_{ret}) + r_{y} \cos(\omega t_{ret})\right)\right)^2} \\ E_{y} &= \frac{\left(1 - \omega^2\right) \left(\frac{r_{y} - \sin(\omega t_{ret})}{t_{ret}} + \omega \cos(\omega t_{ret})\right)}{t_{ret} \left(1 - \frac{\omega}{t_{ret}} \left(-r_{\chi} \sin(\omega t_{ret}) + r_{y} \cos(\omega t_{ret})\right)\right)^3} \\ &+ \frac{\left(\frac{\omega}{t_{ret}} \left(-r_{\chi} \sin(\omega t_{ret}) + r_{y} \cos(\omega t_{ret})\right)\right) \left(\frac{r_{y} - \sin(\omega t_{ret})}{t_{ret}} + \omega \cos(\omega t_{ret})\right)}{t_{ret} \left(1 - \frac{\omega}{t_{ret}} \left(-r_{\chi} \sin(\omega t_{ret}) + r_{y} \cos(\omega t_{ret})\right)\right)^3} \\ &+ \frac{\omega^2 \sin(\omega t_{ret})}{t_{ret} \left(1 - \frac{\omega}{t_{ret}} \left(-r_{\chi} \sin(\omega t_{ret}) + r_{y} \cos(\omega t_{ret})\right)\right)^2} \end{split}$$

```
Exn[idx,jdx] = tempxn
```

Result:

We can see that near field is just Coulomb part, so it will like a charge in the middle, and because Coulomb part is $\sim \frac{1}{r^2}$, it vanished very quickly. Far field is radiation part, so it seems like a dipole in the middle, and also because radiation part is $\sim \frac{1}{r'}$, it will seems like affect all the space.

