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We know much about the arrangement of the orbital electron in an atom. This is large because of the force between the electron and the nucleus as well as between electron themselves is electrical in nature and mathematical treatment of such forces is well established. Unfortunately there is no satisfactory explanation for treating the strong nuclear interaction based on the concept of pion transfer. Consequently there is no simple theory of the detailed structure of the nucleus.

What physicist have done so for is to propose models which can be used to interpret certain aspects of the behaviour of nuclei. Few of them are

- (a) Why do nucleus emit  $\alpha$  and  $\beta$  particles. When there are known to contain only protons and neutrons.
- (b) Why is binding energy per nucleon almost constant?
- (c) Why are 4*u* nuclei particularly stable?
- (d) How one can explain the excited state of nuclei
- (e) How one can interpret the special properties of nucleus Eg: spin, stability magnetic moment etc.

Various models which have been preposed for the nucleus are the different collective models(of which the liquid drop model is the one), fermi gas model and the shell model with different types of coupling. One can see that the resemblence to a drop of liquid setves us basis for the liquid drop model and collective model and resemblence to a weakly interacting gas serves as the basis for the fermi gas model and the shell model. Liquid drop model and shell model are briefly reviewed in this discussion.

### 1.1 Liquid Drop Model

The binding energies and the volume of the nuclei are proportional to the number of nucleons present in nuclei. The proportionality however indicates the short range and saturation characters of the nuclear force. This follows that there is strong interactions amongst all the neighbouring nucleons inside the nucleus. This properties of a nucleus are analogous to the properties of the force which hold a liquid drop together. This analogy led to prepose the liquid drop model of the nucleus. This model is not considered about the individual characteristics of the nucleons and hence this is a statistical model. According to this model nucleus is regarded as an incompressible and uniformly charged liquid drop.

#### (a) Similarities Between a Nucleus and Liquid Drop

(i) There are large number of particle in a nucleus as in a drop of liquid, protons and neutrons in the former and molecule or atoms in the latter.

- (ii) Both of the liquid drop and nucleus exhibits homogeneity and incompressibility, for Eg: charge density is almost constant through out the drop and the nucleus. In both cases density is independent of diamensions
- (iii) Each nucleus in a nucleus interacts strongly with a small number of adjacent nucleons just as do the molecule in a liquid drop. This follows that the nucleon forces are of short range and of saturation character.
- (iv) Both liquid drop and nucleus show surface tension effect ie the surface energy of the nucleus is analogous to the surface tension of liquids
- (v) A part from the coloumb repulsion the forces between the n-n, n-p and p-p are the same in the nucleus just as the intermolecular forces in an ideal solution
  - (vi) Evaporation from a liquid is analogous to the loss of nucleus from a nucleus in the nuclear reaction.
- (vii) The fusion of small drops in to a bigger one, and the breaking up of large drop into small droplets are quite analogous to the fusion of light nuclei in to a heavy nucleus and the fusion of a heavy nucleus in to a light nuclei respectively. The fusion and fission in both cases are exoergic.
- (viii) The thermal agitation of the molecules in a drop is quite analogous to the kinetic energy of the nucleons in a nucleus.

#### (b) Bethe-Weizsacker Semi-Emperical Mass Formula

On the basis of the liquid drop model, Weiszacker and several others have attempted to express the mass of the nuclei in terms of nuclear characteristics in connection with their binding energy and stability. This formula is known as semi-emperical mass formula. Let m(A, Z) mass of the isotope of an element X of atomic number Z and mass number A. Then,

$$m(A,Z) = ZM_H + NM_n - E_b$$
Where

 $M_H = \text{Mass of Hydrogen}$ 
 $M_n = \text{Mass of Neutron}$ 
 $N = A - Z = \text{Number of Neutrons}$ 
 $E_b = \text{Binding energy}$ 

One can express the binding energy  $E_b$  as the sum of no of terms as bellow.

#### (I) Volume Energy

We start by assuming that the energy associated with each nucleon-nucleon bond has some value u. Because each bond energy u is shared by two nucleons. Each has a binding energy of  $\frac{1}{2}u$ . When an assembly of shares of same size is packet together in to a smallest volume, as we suppose is the case of nucleons with in a nucleus, each interior sphere 12 other spheres in contact with it. Hence each interior nucleon in a nucleus has a binding energy of  $12(\frac{1}{2}u) = 6u$ . If all A nucleons in a nucleus were in its interior. The total binding energy of the nucleus would be,

$$E_{v} = 6Au$$
  
Volume energy $E_{v} = a_{1}A$   
 $E_{v} \propto A$ 

#### (II) Surface Energy

Nucleons on the surface will have less than 12 neighbours in contacts. since we have included all nucleons (A) in volume energy we have to minus the energy of peripheral nucleons which is called surface energy. A nucleus

with radius *R* has surfice are  $4\pi R^2 (4\pi R_0^2 A^{\frac{2}{3}})$ . So surface energy is preportional to  $(4\pi R_0^2 A^{\frac{2}{3}})$ 

$$E_s \alpha - \frac{4}{3} \pi R_0^2 A^{\frac{2}{3}}$$
$$E_s = -a_2 A^{\frac{2}{3}}$$

For a nucleus to be stable binding energy should be maximum. So the surface energy term would be a small quantity. For a given volume sphere has the least surface area. So the nucleus size will be spherical and should exhibit the same surface tension effect as a liquid drop.

#### (III) Coloumb Energy

Electric repulsion between each pair of proton in a nucleus also contributes towards decreasing its binding energy. The columb energy Ec of a nucleus is the work that must be done to bring Z protons from infinity in to a spherical aggregate the size of the nucleus.

The potential energy of a pair of protons r a part,

$$V = \frac{-e^2}{4\pi\epsilon_0 r}$$
 Since there are  $\frac{Z(Z-1)}{2}$  pairs of protons 
$$E_c = \frac{Z(Z-1)}{2}V = \frac{-Z(Z-1)e^2}{4\pi\epsilon_0} \left(\frac{1}{r}\right)_{qq}$$

If the protons are uniformly distributed through out a nucleus of radius R

$$\left(\frac{1}{r}\right)_{av}$$
 is proportional to  $\left(\frac{1}{R}\right)$  and hence  $\frac{1}{A^{\frac{1}{3}}}$  Coloumb energy  $E_c=-a_3\frac{Z(Z-1)}{A^{\frac{1}{3}}}$ 

The coloumb energy is negative because it arises from the effect that opposes nuclear stability.

The total binding energy  $E_b$  of the nucleus is the sum of  $E_v$ ,  $E_s$ , and  $E_c$ 

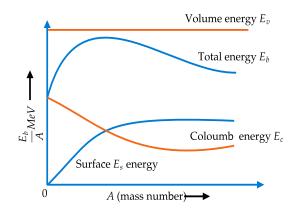
$$E_b = E_v + E_s + E_c$$

$$C = a_1 A - a_2 A^{\frac{2}{3}} - a_3 \frac{Z(Z-1)}{A^{\frac{1}{3}}}$$

Binding energy nucleon

$$\frac{E_b}{A} = a_1 - \frac{a_2}{A^{\frac{1}{3}} - a_3 \frac{z(z-1)}{A^{\frac{4}{3}}}}$$

When we plot,



#### (iv) Corrections to the Formula

The above binding energy formula can be improved by taking into account two effects that do not fit into the

simple liquid drop model but which make sense in terms of a model that provides for nuclear energy levels. The above result was improved by including two effects

- (a) Asymmetry Effect
- (b) Pairing Effect

#### (a) Asymmetry Effect(B<sub>a</sub>)

Asymmetry Energy Term,  $B_a$  depends on the neutron excess (N-Z) and decreases the nuclear binding energy. So far, we have neglected the quantization of energy states of individual nucleons in the nucleus and the application of the Pauli Exclusion Principle.

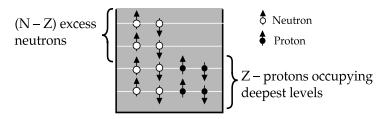


Figure 1.1

If we put Z protons and N neutrons into the nuclear energy shells, the lowest Z energy levels are filled first. By Pauli Exclusion Principle, the excess (N-Z) neutrons must go into previously unoccupied quantum states since the first Z quantum states are already filled up with protons and neutrons.

These (N-Z) excess neutrons are occupying higher energy quantum states and are consequently less tightly bound than the first 2Z nucleons which occupy the deepest lying energy levels. Thus neutron asymmetry gives rise to a disruptive term in nuclear binding energy. Excess energy per nucleon  $\alpha^{N-Z}$ 

Since the total number of excess neutrons is (N-Z), the total deficit in nuclear binding energy is proportional to product of these

$$\Rightarrow B_a = -a_a \frac{(N-Z)^2}{A} = -a_a \frac{(A-2Z)^2}{A},$$
 coefficient(19.0 MeV).

where  $a_a$ : asymmetric energy coefficient(19.0 MeV).

#### (b) Pairing Effect

Since all the previous terms have involved a smooth variation of B whenever Z or N changes and does not account for the kinks which show an evidence for favored pairing.

In liquid drop model we have omitted the intrinsic spin of the nucleons and shell effects. This is corrected by adding a pairing energy term  $B_p$  to the nuclear binding energy.

$$B_p = (\pm, 0) \frac{a_p}{A^{+3/4}}, a_p = \begin{cases} 0 \text{ for odd-even or even-odd} \\ -\text{ve for odd-odd} \\ +\text{ve for even-even} \end{cases} \text{ and } a_p = 33.5 \text{MeV}$$

The final expression for binding energy is

$$B = a_v A - a_s A^{2/3} - a_c \frac{Z(Z-1)}{A^{1/3}} - a_a \frac{(A-2Z)^2}{A} (\pm,0) \frac{a_p}{A^{3/4}}$$

Now, nuclear mass can be written as

$$M(Z,A) = AM_N - Z(M_N - M_P) - B/c^2 \qquad (M \& B \text{ in mass units})$$

$$M(Z,A) = AM_N - Z(M_N - M_P) + \left\{ -a_v A + a_s A^{2/3} + a_c \frac{Z(Z-1)}{A^{1/3}} + a_a \frac{(A-2Z)^2}{A} (\mp,0) a_p A^{-3/4} \right\} \frac{1}{c^2}$$

#### Most Stable Nuclei Among Members of Isobaric Family

For a given A, we have to find the value of Z for which the binding energy B is a maximum, which corresponds to maximum stability, we must show  $\left(\frac{dB}{dZ}\right)_{Z=z_0}=0$ 

Since 
$$B = a_v A - a_s A^{2/3} - a_c \frac{Z(Z-1)}{A^{1/3}} - a_a \frac{(A-2Z)^2}{A} (\pm, 0) \frac{a_p}{A^{3/4}}$$

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$$\Rightarrow \left(\frac{dB}{dZ}\right)_{Z=Z_0} = -\frac{a_c}{A^{1/3}} (2Z_0 - 1) - \frac{a_a}{A} 2(A - 2Z_0) (-2) = 0$$

$$\Rightarrow -\frac{a_c}{A^{1/3}} (2Z_0 - 1) + \frac{4a_a}{A} (A - 2Z_0) = 0 \quad \Rightarrow -2Z_0 \left(\frac{a_c}{A^{1/3}} + \frac{4a_a}{A}\right) = -\frac{a_c}{A^{1/3}} - 4a_a$$

$$\Rightarrow Z_0 = \frac{\left(4a_a + \frac{a_c}{A^{1/3}}\right)}{2\left(\frac{a_c}{A^{1/3}} + \frac{4a_a}{A}\right)} \Rightarrow Z_0 = \frac{4a_a + a_c A^{-1/3}}{2a_c A^{-1/3} + 8a_a A^{-1}}$$

**Exercise 1.1** The atomic mass of the zinc isotope  ${}^{64}_{30}$ Zn is 63.9294. Compare its binding energy with the prediction of the liquid drop model.

**Solution:** 

B.E. = 
$$[30 \times 1.007825 + 34 \times 1.008665 - 63.929] \times 931.49 = 559.1 \text{MeV}$$

From semi-empirical B.E. formula (Z = 30, N = 34, A = 64)

$$B = a_v A - a_c A^{2/3} - a_c \frac{Z(Z-1)}{A^{1/3}} - a_a \frac{(A-2Z)^2}{A} + \frac{a_p}{A^{3/4}} = 561.7 MeV$$

Thus percentage difference = 0.5%.

#### 1.2 Mass Parabola's

From the semi-empirical mass equation we have

$$M(Z,A) = AM_N - Z(M_N - M_p) - a_v A + a_s A^{2/3} + a_c \frac{Z(Z-1)}{A^{1/3}} + a_a \frac{(A-2Z)^2}{A} (\mp,0) a_p A^{-3/4}$$

$$M(Z,A) = A \left[ M_N - \left( a_v - a_a - \frac{a_s}{A^{1/3}} \right) \right] + Z \left[ (M_p - M_N) - \frac{a_c}{A^{1/3}} - 4a_a \right] + Z^2 \left[ \frac{a_c}{A^{1/3}} + \frac{4a_a}{A} \right] \pm E_p \text{ or }$$

$$M(Z,A) = \alpha A + \beta Z + \gamma Z^2 \pm \delta$$
where  $\alpha = M_N - \left( a_v - a_a - \frac{a_s}{A^{1/3}} \right), \beta = -4a_a - (M_n - M_p) - \frac{a_c}{A^{1/3}} \text{ and } \gamma = \left( \frac{4a_a}{A} + \frac{a_c}{A^{1/3}} \right)$ 

 $\delta$  is pairing energy  $= \left( E_p \right) = + \delta \text{for even Z even N}$ 

= 0 for odd Z even N or even N and odd Z

 $= -\delta$  for odd Z odd N

When A is constant, the equation  $M(Z,A) = \alpha A + \beta Z + \gamma Z^2 \pm \delta$  represents a parabola. Thus the plot of M and Z is parabolic with the "minimum" corresponding to that value of Z which gives the (hypothetical) "most stable" isobar in the isobaric family.

#### For Odd $A(\delta = 0)$

As either one of N or Z is even and the other one is odd (since odd + even = odd), so only one parabola implying that there is only one stable nucleus.

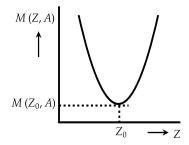


Figure 1.2

Consider the isobaric family for A,

$$\left(\frac{\delta M}{\delta Z}\right)_{A} = \beta + 2\gamma Z_{0} = 0$$

 $Z_0$  = Nuclear charge of "most stable nuclei

$$\therefore Z_0 = \frac{-\beta}{2\gamma} \Rightarrow (\beta = -2\gamma Z_0)$$

So mass of the "most stable" isobar is

$$M(Z_0, A) = \alpha A - 2\gamma Z_0 Z_0 + \gamma Z_0^2 (:: \beta = 2\gamma Z_0)$$

$$\therefore M(Z_0, A) = \alpha A - \gamma Z_0^2 \quad Also, M(Z, A) = \alpha A - 2\gamma Z_0 \cdot Z + \gamma Z^2$$

The difference in masses for odd A is:

$$M(Z, A) - M(Z_0, A) = -2\gamma Z_0 Z + \gamma Z^2 + \gamma Z_0^2 = \gamma (Z - Z_0)^2 = \gamma (Z - Z_0)^2$$

**Even A isobars**  $(\delta \neq \mathbf{0})$  Here pairing term  $\delta \neq 0$  since both odd-odd and even-even nuclei are included. So two parabolas,

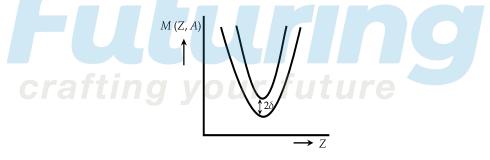


Figure 1.3

For odd-odd: 
$$M(Z_0, A) = \alpha A - \gamma Z_0^2 - \delta$$

For even-even: 
$$M(Z_0, A) = \alpha A - \gamma Z_0^2 + \delta$$

where 
$$Z_0 = -\frac{\beta}{2\gamma}$$

The vertical separation between two parabolas is  $2\delta$ 

$$M(Z, A) = \alpha A - 2\gamma Z_0 Z + \gamma Z^2 \pm \delta$$

#### 1.2.1 Shell Model

#### **Evidence of Shell Structure**

1. Just as intert gases, with  $Z = 2, 10, 18, 36, 54, \ldots$  electrons having closed shells show high chemical stability, nuclei with Z = 2, 8, 20, 28, 50, 82, and 126 nucleons - the so-called magic numbers of the same kind (either proton or neutron) are particularly stable. (Nuclei with Z = N, a magic number are said to be doubly magic and show exceptionally high stability).

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2. The number of stable isotopes (Z = const.) and isotones (N = const.) is larger with respective number of protons and neutrons and equal to either of magic numbers, e.g.  $S_n(Z = 50)$  has 10 stable isotopes, Ca(Z = 20) has 6; the biggest group of isotone is at N = 82, then at N = 50 and N = 20. The relative abundances of naturally occurring isotopes whose nuclei contain magic numbers of protons or neutrons, have greater relative abundances, e.g. the isotopes  ${}^{88}\text{Sr}(N = 50)$ ,  ${}^{138}\text{Ba}(N = 82)$  and  ${}^{140}\text{Ce}(N = 82)$  have relative abundances of  ${}^{82}\text{Se}(N = 50)$ ,  ${}^{138}\text{Ba}(N = 82)$  and  ${}^{140}\text{Ce}(N = 82)$  have relative abundances of  ${}^{82}\text{Se}(N = 50)$ ,  ${}^{138}\text{Ce}(N = 82)$  and  ${}^{140}\text{Ce}(N = 82)$  have

- 3. The three naturally occurring radioactive series decay to the stable end product the three isotopes of Pb;  $^{206}_{82}$  Pb(Z=82, N=124),  $^{207}_{8}$  Pb(Z=82, N=125) and  $^{208}_{82}$  Pb (with Z=82 and N=126) indicating extra stable configuration of magic nuclei.
- 4. The neutron absorption cross-section is low for nuclei with N = magic numbers, e.g. 50, 82 and 126, indicating reluctance of magic nuclei to accept extra neutrons in their completely filled shells.
- 5. Isotopes like  ${}^{17}_8O, {}^{87}_{36}$  K and  ${}^{137}_{54}$ Xe are spontaneous neutron emitters when excited by preceding  $\beta$ -decay. These isotopes have N=9,51 and 83 respectively, i.e., N(8+1),N(50+1) and N(82+1). We can interpret this loosely bound neutron as a valence neutron which the isotopes emit to assume some magic number N-value for stability.
- 6. Nuclei with magic numbers of neutrons or protons have their first excited states at higher energies than in cases of the neighbouring nuclei.
- 7. Electric quadrupole moment *Q* of magic nuclei is zero indicating spherical symmetry of nucleus for closed shells. When *Z*-value or *N*-value is gradually increases from one magic number to the next, *Q* increases from zero to a maximum and then decreases to zero at the next magic number.
- 8. The energy of  $\alpha$  or  $\beta$  particles emitted by magic radioactive nuclei is larger.
- 9. The asymmetry of the fission of the uranium nucleus could involve the sub-structure of nuclei, which is expressed in the existence of the magic numbers.

#### **Shell Theory**

The shell model of the nucleus is an attempt to account for the existence of magic numbers and certain other nuclear properties in terms of nucleon behavior in a common force field.

Because the precise form of the potential-energy function for a nucleus is not known, unlike the case of an atom, a suitable function U(r) has to be assumed. A reasonable guess on the basis of the nuclear density curves is a square well with rounded comers. Schrödinger's equation for a particle in a potential well of this kind is then solved, and it is found that stationary states of the system occur that are characterized by quantum numbers n,l, and  $m_l$  whose significance is the same as in the analogous case of stationary states of atomic electrons. Neutrons and protons occupy separate sets of states in a nucleus because the latter interact electrically as well as through the specifically nuclear charge. However, the energy levels that come from such a calculation do not agree with the observed sequence of magic numbers. Using other potential-energy functions, for instance that of the harmonic oscillator, gives no better results. Something essential is missing from the picture.

The problem was finally solved independently by Maria Goeppert-Mayer and J. H. D. Jensen in 1949. They realized that it is necessary to incorporate a spin-orbit interaction whose magnitude is such that the consequent splitting of energy levels into sublevels is many times larger than the analogous splitting of atomic energy levels.

#### Assumptions of Shell Theory

- 1. LS coupling holds only for the very lightest nuclei, in which the l values are necessarily small in their normal configurations. In this scheme, the intrinsic spin angular momenta  $S_i$  of the particles concerned (the neutrons form one group and the protons another) are coupled together into a total spin momentum S. The orbital angular momenta  $L_i$  are separately coupled together into a total orbital momentum L. Then S and L are coupled to form a total angular momentum J of magnitude  $\sqrt{J(J+1)}\hbar$ .
- 2. After a transition region in which an intermediate coupling scheme holds, the heavier nuclei exhibit jj coupling. In this case the  $S_i$  and  $L_i$  of each particle are first coupled to form a  $\mathbf{J}_i$  for that particle of magnitude  $\sqrt{j(j+1)}\hbar$ . The various  $\mathbf{J}_i$  then couple together to form the total angular momentum J. The jj coupling scheme holds for the great majority of nuclei.

- 3. The levels are designated by a prefix equal to the total quantum number n, a letter that indicates l for each particle in that level according to the usual pattern (s, p, d, f, g, ... corresponding, respectively, to l = 0, 1, 2, 3, 4, ...), and a subscript equal to j.
- 4. The spin-orbit interaction splits each state of given j into 2j + 1 substates, since there are 2j + 1 allowed orientations of  $J_i$ .
- 5. The number of available nuclear states in each nuclear shell is, in ascending order of energy, 2, 6, 12, 8, 22, 32, and 44. Hence shells are filled when there are 2, 8, 20, 28, 50, 82, and 126 neutrons or protons in a nucleus.

When an appropriate strength is assumed for the spin-orbit interaction, the energy levels of either class of nucleon fall into the sequence shown in table.

Shell number	State in a Shell	Number of Particles in a Shell	Total Particles upto the Shell Closure
1	$1s_{\frac{1}{2}}$	2	2
2	$1p_{\frac{3}{2}}, 1p_{\frac{1}{2}}$	6	8
3	$1d_{\frac{5}{2}}, 2s_{\frac{1}{2}}, 1d_{\frac{3}{2}}$	12	20
3a	$1f_{rac{7}{2}}$	8	28
4	$2p_{\frac{3}{2}}, 1f_{\frac{5}{2}}, 2p_{\frac{1}{2}}, 1g_{\frac{9}{2}}$	22	50
5	$1g_{\frac{7}{2}}, 2d_{\frac{5}{2}}, 2d_{\frac{3}{2}}, 3s_{\frac{1}{2}}, 1h_{\frac{11}{2}}$	32	82
6	$1h_{\frac{9}{2}}, 2f_{\frac{7}{2}}, 2f_{\frac{5}{2}}, 3p_{\frac{3}{2}}, 3p_{\frac{1}{2}}, 1i_{\frac{13}{2}}$	44	126
7	$2g_{\frac{9}{2}}, 3d_{\frac{5}{2}}, 1i_{\frac{11}{2}}, 2g_{\frac{7}{2}}, 4s_{\frac{1}{2}}, 3d_{\frac{3}{2}}, 1j_{\frac{15}{2}}$	58	184

#### 1.2.2 Prediction of the Shell Model

The shell model accounts for several nuclear phenomena in addition to magic numbers. The shell model can also predict nuclear angular momenta. In even-even nuclei, all the protons and neutrons should pair off to cancel out one another's spin and orbital angular momenta. Thus even-even nuclei ought to have zero nuclear angular momenta, as observed. In even-odd and odd-even nuclei, the half-integral spin of the single "extra" nucleon should be combined with the integral angular momentum of the rest of the nucleus for a half-integral total angular momentum. Odd-odd nuclei each have an extra neutron and an extra proton whose half-integral spins should yield integral total angular momenta. Both these predictions are experimentally confirmed.

#### (i)Spin and Parity From Shell Model

Nuclear states have an intrinsic spin and a well defined parity,  $\eta = \pm 1$ , defined by the behaviour of the wavefunction for all the nucleons under reversal of their coordinates with the centre of the nucleus at the origin.

$$\Psi(-\mathbf{r}_1, -\mathbf{r}_2 \cdots - \mathbf{r}_A) = \eta \Psi(\mathbf{r}_1, \mathbf{r}_2 \cdots \mathbf{r}_A)$$

The spin and parity of nuclear ground states can usually be determined from the shell model. Protons and neutrons tend to pair up so that the spin of each pair is zero and each pair has even parity  $(\eta = 1)$ . Thus we have

- Even-even nuclides (both Z and A even) have zero intrinsic spin and even parity.
- Odd A nuclei have one unpaired nucleon. The spin of the nucleus is equal to the j value of that unpaired nucleon and the parity is  $(-1)^l$ , where l is the orbital angular momentum of the unpaired nucleon. For example  $^{47}_{22}$ Ti (titanium) has an even number of protons and 25 neutrons. 20 of the neutrons fill the shells up to magic number 20 and there are 5 in the  $1f_{\frac{7}{2}}$  state  $(l=3, j=\frac{7}{2})$  Four of these form pairs and the remaining one leads to a nuclear spin of  $\frac{7}{2}$  and parity  $(-1)^3=-1$ .

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• Odd-odd nuclei. In this case there is an unpaired proton whose total angular momentum is  $j_1$  and an unpaired neutron whose total angular momentum is  $j_2$ . The total spin of the nucleus is the (vector) sum of these angular momenta and can take values between  $|j_1 - j_2|$  and  $|j_1 + j_2|$  (in unit steps). The parity is given by  $(-1)^{(l_1+l_2)}$ , where  $l_1$  and  $l_2$  are the orbital angular momenta of the unpaired proton and neutron respectively.

For example  ${}_{3}^{6}$ Li (lithium) has 3 neutrons and 3 protons. The first two of each fill the 1s level and the thrid is in the  $1p_{\frac{3}{2}}$  level. The orbital angular mometum of each is l=1 so the parity is  $(-1)\times(-1)=+1$  (even), but the spin can be anywhere between 0 and 3.

#### **Examples**

1. The level configurations for the simple odd proton nuclei  ${}_{3}^{7}\text{Li}_{4}$  and  ${}_{7}^{15}$  N<sub>8</sub> in their ground states is as follows:

$$\frac{7}{3} \text{Li}_{4} : {\binom{1s_{\frac{1}{2}}}{2}}^{2}, {\left(1p_{\frac{3}{2}}\right)}^{1}$$

$$\frac{15}{7} \text{N}_{8} : {\binom{1s_{\frac{1}{2}}}{2}}^{2}, {\left(1p_{\frac{3}{2}}\right)}^{4}, {\binom{1p_{\frac{1}{2}}}{2}}^{1}$$

After putting pairs of particles in the earlier shells, one finds that they give rise to spins of lass odd protons as  $\frac{3}{2}$  and  $\frac{1}{2}$  respectively. One obtains the parities of both states as  $(-1)^l$  and is odd (-) since l=1 for the last odd protons in both the examples. Thus the states can be specified as  $I^{\pi}$  equals to  $\frac{3^-}{2}$  and  $\frac{1^-}{2}$  respectively.

2. We now consider another example of odd neutron nuclei, <sup>33</sup> S<sub>17</sub> and <sup>29</sup>Si<sub>15</sub> having following neutron level configurations:

$$\frac{{}^{33}_{16}\,\mathrm{S}_{17}: \left({}^{1s_{\frac{1}{2}}}\right)^{2} \left| \left(1p_{\frac{3}{2}}\right)^{4} \left({}^{1p_{\frac{1}{2}}}\right)^{2} \right| \left({}^{1d_{\frac{5}{2}}}\right)^{6} \left(2s_{1}\frac{1}{2}\right)^{2} \left({}^{29}d_{\frac{3}{2}}\right)^{1}}{2^{14}_{15}: \left({}^{1s_{1}}\frac{1}{2}\right)^{2} \left| \left({}^{1}p_{\frac{3}{2}}\right)^{4} \left({}^{1p_{\frac{1}{2}}}\right)^{2} \left| \left({}^{1d_{\frac{5}{2}}}\right)^{6} \left(2s_{1}\frac{1}{2}\right)^{1}}{2^{14}_{15}} \right|$$

and have spins  $\frac{3}{2}$  and  $\frac{1}{2}$  respectively.

3. Let us now take the example of mirror nuclei,  ${}^{13}C_7$  and  ${}^{13}N_6$ . The level configuration for 7 odd protons in  ${}^{13}N_6$  or 7 odd neutrons in  ${}^{13}C_7$  is the same and is expressed as

$$\left(1s_{1}\right)^{2},\left(1p_{\frac{3}{2}}\right)^{4},\left(1p_{\frac{1}{2}}\right)^{1}$$

Both have same spin  $\frac{1}{2}^-$  as predicted by the shell model and confirmed experimentally. There is another example of mirror nuclei  $^{17}_8O_9$  and  $^{17}_9F_8$ . The configuration for both these for the odd neutron or odd proton as

$$\left(\begin{array}{c} 1s_{\frac{1}{2}} \end{array}\right)^{2} \left| \left(1p_{\frac{3}{2}}\right)^{4} \left(1p_{\frac{1}{2}}\right)^{2} \right| \left(1d_{\frac{5}{2}}\right)^{1}$$

Shell model predicts spin  $\frac{5}{2}^+$  for both these mirror nuclei, which is in accordance with the experiment.

#### (ii) Magnetic Dipole Moments

Since nuclei with an odd number of protons and/or neutrons have intrinsic spin they also in general possess a magnetic dipole moment.

The unit of magnetic dipole moment for a nucleus is the "nuclear magneton" defined as

$$\mu_N = \frac{e\hbar}{2m_p}$$

which is analogous to the Bohr magneton but with the electron mass replaced by the proton mass. It is defined such that the magnetic moment due to a proton with orbital angular momentum is  $\mu_N$ . Experimentally it is found that the magnetic moment of the proton (due to its spin) is

$$\mu_p = 2.79 \mu_N = 5.58 \mu_N s, \quad \left(s = \frac{1}{2}\right)$$

and that of the neutron is

$$\mu_n = -1.91 \mu_N = -3.82 \mu_N s, \quad \left(s = \frac{1}{2}\right)$$

If we apply a magnetic field in the z-direction to a nucleus then the unpaired proton with orbital angular momentum  $\mathbf{l}$ , spin  $\mathbf{s}$  and total angular momentum  $\mathbf{j}$  will give a contribution to the z- component of the magnetic moment

$$\mu^z = (5.58s^z + l^z) \mu_N$$
.

As in the case of the Zeeman effect, the vector model may be used to express this as

$$\mu^z = \frac{(5.58 < \mathbf{s} \cdot \mathbf{j} > + < \mathbf{l} \cdot \mathbf{j} >)}{< \mathbf{j}^2 >} j^z \mu_N$$

using

$$\langle \mathbf{j}^{2} \rangle = j(j+1)\hbar^{2}$$

$$\langle \mathbf{s} \cdot \mathbf{j} \rangle = \frac{1}{2} \left( \langle \mathbf{j}^{2} \rangle + \langle \mathbf{s}^{2} \rangle - \langle \mathbf{l}^{2} \rangle \right)$$

$$= \frac{\hbar^{2}}{2} (j(j+1) + s(s+1) - l(l+1))$$

$$\langle \mathbf{l} \cdot \mathbf{j} \rangle = \frac{1}{2} \left( \langle \mathbf{j}^{2} \rangle + \langle \mathbf{l}^{2} \rangle - \langle \mathbf{s}^{2} \rangle \right)$$

$$= \frac{\hbar^{2}}{2} (j(j+1) + l(l+1) - s(s+1))$$

We end up with expression for the contribution to the magnetic moment

$$\mu = \frac{5.58(j(j+1) + s(s+1) - l(l+1)) + (j(j+1) + l(l+1) - s(s+1))}{2j(j+1)}j\mu_N$$

and for a neutron with orbital angular momentum l' and total angular momentum j' we get (not contribution from the orbital angular momentum because the neutron is uncharged)

$$\mu = -\frac{3.82(j'(j'+1) + s(s+1) - l'(l'+1))}{2j'(j'+1)}j'\mu_N$$

Thus, for example if we consider the nuclide  ${}^{7}_{3}$ Li for which there is an unpaired proton in the  $2p_{\frac{3}{2}}$  state  $(l=1,j=\frac{3}{2}$  then the estimate of the magnetic moment is

$$\mu = \frac{5.58\left(\frac{3}{2} \times \frac{5}{2} + \frac{1}{2} \times \frac{3}{2} - 1 \times 2\right) + \left(\frac{3}{2} \times \frac{5}{2} + 1 \times 2 - \frac{1}{2} \times \frac{3}{2}\right)}{2 \times \frac{3}{2} \times \frac{5}{2}} = 3.79 \mu_N$$

The measured value is  $3.26\mu_N$  so the estimate is not too good. For heavier nuclei the estimate from the shell model gets much worse.

The precise origin of the magnetic dipole moment is not understood, but in general they cannot be predicted from the shell model. For example for the nuclide  $_9^{17}$  F (fluorine), the measured value of the magnetic moment is  $4.72\mu_N$  whereas the value predicted form the above model is  $-0.26\mu_N$ . !! There are contributions to the magnetic moments from the nuclear potential that is not well-understood.

#### 1.3 Liquid Drop Model

- Binding Energy Per nucleon is equivalent to latent heat of vaporiation of liquid.
- Both liquid drop and nucleus posses same density.

#### **Merits:**

Able to explain Binding energy, atomic masses and nuclear fission.

#### **Shell Model**

- Nucleons have freedom of motion inside nucleous
- Protons and neutrons have potential (average fields) created due to mutual interaction between other nucleons.
- Explains first few magic numbers. Eg.  ${}_{20}C^{40}(P=20,n=20) \rightarrow \text{Doubly magic nuclei}$

### 1.4 Application of shell model:

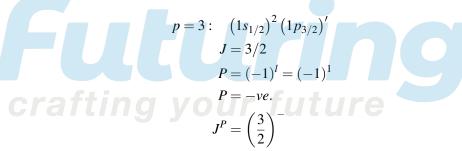
#### **Spin-Parity Prediction:**

a) For even-even nuclei:

For even-even nuclei both P and n are paired off and resultant to this rule, spin is zero.

#### b) Even-Odd or Odd-Even Nuclei:

Here, last unpaired nucleon will decide the nuclear spin.



**eg: ii**) 
$$C_6^{13}$$

$$p=6$$
 no role since it is even number  $n=7$ :  $\left(1s_{1/2}\right)^2\left(1p_{3/2}\right)^4\left(1p_{1/2}\right)^1$ .  $J^P=\left(\frac{1}{2}\right)^-$ 

#### **Magnetic Moment**

**Exercise 1.2** Find magnetic moments of  ${}_6C^{13}$  and  ${}_{11}Na^{23}$  nuclei by shell model.

#### **Solution:**

$$i_{6}^{13}C : n = 7 : (s_{1/2})^{2} (|_{p_{3}/2})^{4} (1p_{1/2})^{1}$$

$$j = 1/2$$

$$l = 1, s = 1/2$$

$$j = l - 1/2 = 1/2$$

$$\mu = 1.91 \frac{J}{J+1} \mu_{N}$$

$$= 1.91 \frac{1/2}{(1/2)+1} \mu_{N} = 0.63 \mu_{N}$$

$$i_{11}^{23}Na : p = 11 : (1s_{1/2})^{2} (1p_{3/2})^{4} (1p_{1/2})^{2} (1d_{5/2})^{3}$$

$$j = 5/2, l = 2, s = 1/2$$

$$J = l + 1/2 \text{ case}$$

$$\mu = (J+2.29) \mu_{N}$$

$$= \frac{5}{2} + 2.29 = 4.79 \mu_{N}$$

#### **Electric Quadrupole Moment:**

$$Q = -\frac{(2J-1)}{(2J+2)} \langle r^2 \rangle$$

$$r^2 = \text{mean square radius}$$

$$\langle r^2 \rangle = \frac{3}{5} R_0^2 \quad R_0 = 1.2 \times 10^{-15} \text{ m}$$

$$Q = -\frac{3}{5} \frac{(2J-1)}{(2J+2)} R_0^2 \text{ barn}$$

$$1 \text{ barm} = 10^{-28} \text{ m}$$

**Exercise 1.3** According to shell model, the ground state of  ${}_{8}^{15}0$  nucleus is:

**a.** 
$$(\frac{3}{2})^+$$

**b.** 
$$(\frac{1}{2})^+$$

**c.** 
$$(\frac{3}{2})^{-}$$

**d.** 
$$(\frac{1}{2})^{-}$$

**Solution:** 

$${}_{8}^{15}0 \rightarrow n = 7: (1s_{1/2})^{2} (1p_{3/2})^{4} (1p_{1/2})^{1}$$

$$J = 1/2, \quad P = (-1)^{1} = -\text{ve}$$

$$J^{P} = \left(\frac{1}{2}\right)^{-}$$

So the correct answer is **Option** (d)

1.5 Collective Model 13

**Exercise 1.4** For the  $O^{17}$  nucleus (A = 17, z = 8), the effective magnetic moment is given by

$$\vec{H}_{eff} = \frac{e\hbar}{2Mc} \vec{g} \vec{J}$$

 $\vec{H}_{eff} = \frac{e\hbar}{2Mc} \overrightarrow{gJ}$ Where g is equal to

**a.** 1.12

**b.** -0.77

c. -1.28

**d.** 1.28

**Solution:** 

$$I_{8}^{17}: \quad n = 9: (1s_{1/2})^{2} (1p_{3/2})^{4} (1p_{1/2})^{2} (1d_{5/2})'$$

$$J = 5/2, l = q, \quad S = 1/2$$

$$J = l + 1/2 \text{ case for neutron.}$$

$$\mu = -1.91 \mu_{N} = g \text{ J}\mu_{N}$$

$$g = \frac{-1.91 \times 2}{5} = -0.77$$

So the correct answer is **Option** (b)

**Exercise 1.5** The electric quadrupole moment of a odd proton nucleus is  $\frac{(2j-1)}{(2j+2)} \langle r^2 \rangle$ . What is the value, in barn of the quardrupole moment of  ${}^{27}Al$  nucleus in the shell model?

Solution:

$$27_{13}Al: \quad b = 13: (1s_{1/2})^2 (1p_{3/2})^4 (b_{1/2})^2 (1ds/2)^5$$

$$J = 5/2$$

$$Q = \frac{3}{5} \left[ \frac{2 \times \frac{5}{2} - 1}{2 \times \frac{5}{2} + 2} \right] (1.2 \times 10^{-15})^2 \cdot (27)^{2/3}$$

$$Q = 4.44 \times 10^{-30} \text{ m}$$

$$Q = 0.044 \times (10^{-28} \text{ m barn.}$$

So the correct answer is **Option** (a)

Condition to Posses Quadrupole Moment:  $I > \frac{1}{2}$ 

#### 1.5 **Collective Model**

Core consists of numer of nucleons equal to the magic number.

$$E_{\text{total}} = E_{R_{ot}} + E_{\text{vib}} + E_{\text{nuclear}}$$

#### **Rotational Spectra**

$$E_{Rot} = \frac{\hbar^2}{2I}J(J+1)$$

Nuclear spin J is restricted with even values with even parity.

i.e 
$$0^+, 2^+, 4^+, 6^+, \dots$$

$$E_{0^{+}}=0; \quad E_{2^{+}}=\frac{6\hbar^{2}}{2I}; E_{4^{+}}=\frac{20\hbar^{2}}{2I}; \ldots ...$$

**Exercise 1.6** The single particle energy difference between p- orbitals  $(p_{3/2}\&p_{1/2})$  of nucleus  ${}_{50}^{114}Sn$  is 3MeV The energy difference between the states in its 1f orbital is

a. -7MeV

**b.** 7MeV

c. 5 MeV

d. -5 MeV

**Solution:** 

$$E_{J} = \frac{\hbar^{2}}{2I}J(J+1)$$

$$\Delta E = \frac{\hbar^{2}}{2I} \left[ \frac{1}{2} \left( \frac{1}{2} + 1 \right) - \frac{3}{2} \left( \frac{3}{2} + 1 \right) \right]$$

$$3MEV = \frac{\hbar^{2}}{2I}(-3)$$

$$\Rightarrow \frac{\hbar^{2}}{2I} = -1 \text{ MeV}$$

$$\Delta E = f_{5/2} - f_{1/2} = \frac{\hbar^{2}}{2I} \left[ \frac{5}{2} \left( \frac{5}{2} + 1 \right) - \frac{7}{2} \left( \frac{7}{2} + 1 \right) \right]$$

$$= (-1)(-7)\text{MeV}$$

$$= 7\text{MeV}$$

Alternate:

$$1P_{1/2} - 1P_{3/2} = 3 \text{MeV}$$
  
 $1d3/2 - 1d5/2 = 5 \text{MeV}$   
 $1f5/2 - 1f7/2 = 7 \text{MeV}$ 

So the correct answer is **Option** (b)

**Exercise 1.7** A permanently deformed even-even nucleus with  $J^p = 2^+$  has rorational energy 93Kev. The energy of the next excited state

**a.** 372kev

**b.** 310keV

**c.** 273keV

**d.** 186keV

**Solution:** 

$$E_4^+ = \frac{\hbar^2}{2I} 2(2+1) = 93 \text{keV}$$

$$\frac{\hbar^2}{2I} = \frac{93}{6} \text{keV}$$

$$Eu^+ = \frac{\hbar^2}{2I} (4)(4+1) = 20 \times \frac{\hbar^2}{2I}$$

$$= 20 \times \frac{93}{6} \text{keV}$$

$$\Rightarrow 310 \text{keV}$$

So the correct answer is **Option** (b)

## **Practice Set-1**

1.	The root-mean-square (r.m.s) energy of a nucleon in a nucleus of atomic number A in its ground state
	varies as:

**a.**  $A^{4/3}$ 

**b.**  $A^{1/3}$ 

**c.**  $A^{-1/3}$ 

**d.**  $A^{-2/3}$ 

2. Let  $E_S$  denotes the contribution of the surface energy per nucleon in the liquid drop model. The ratio  $E_S(^{27}_{13}\text{Al}): E_S(^{64}_{30}\text{Zn})$  is

[NET/JRF (JUNE-2016)]

**a.** 2:3

**b.** 4:3

**c.** 5:3

**d.** 3:2

3. The binding energy of a light nucleus (Z,A) in MeV is given by the approximate formula

$$B(A,Z) \approx 16A - 20A^{2/3} - \frac{3}{4}Z^2A^{-1/3} + 30\frac{(N-Z)^2}{A}$$

where N = A - Z is the neutron number. The value of Z of the most stable isobar for a given A is

[ NET/JRF (JUNE-2013)]

**a.**  $\frac{A}{2} \left( 1 - \frac{A^{2/3}}{160} \right)^{-1}$ 

**b.**  $\frac{A}{2}$ 

**c.**  $\frac{A}{2} \left( 1 - \frac{A^{2/3}}{120} \right)^{-1}$ 

**d.** 
$$\frac{A}{2} \left( 1 + \frac{A^{4/3}}{64} \right)^{-1}$$

4. If the binding energy B of a nucleus (mass number A and charge Z) is given by

$$B = a_V A - a_S A^{2/3} - a_{sym} \frac{(2Z - A)^2}{A} + \frac{a_c Z^2}{A^{1/3}}$$

where  $a_V = 16 \text{MeV}$ ,  $a_{SM} = 16 \text{MeV}$ ,  $a_{SMM} = 24 \text{MeV}$  and  $a_C = 0.75 \text{MeV}$ , then for the most stable isobar for a nucleus with A = 216 is

[ NET/JRF (DEC-2014)]

**a.** 68

**b.** 72

c. 84

**d.** 92

5. Of the nuclei of mass number A = 125, the binding energy calculated from the liquid drop model (given that the coefficients for the Coulomb and the asymmetry energy are  $a_c = 0.7 \text{MeV}$  and  $a_{sym} = 22.5 \text{MeV}$  respectively) is a maximum for

[ NET/JRF (DEC-2015)]

**a.**  $_{54}^{125}$ Xe

**b.**  $^{124}_{53}I$ 

**c.**  $^{125}_{52}$ Te

**d.**  $^{125}_{51}$ Sb

6. The Bethe-Weizsacker formula for the binding energy (in MeV ) of a nucleus of atomic number Z and mass number A is

$$15.8A - 18.3A^{2/3} - 0.714 \frac{Z(Z-1)}{A^{1/3}} - 23.2 \frac{(A-2Z)^2}{A}$$

The ratio Z/A for the most stable isobar of a A = 64 nucleus, is nearest to

**a.** 0.30

**b.** 0.35

**c.** 0.45

**d.** 0.50

7. Let us approximate the nuclear potential in the shell model by a three dimensional isotropic harmonic oscillator. Since the lowest two energy levels have angular momenta l=0 and l=1 respectively, which of the following two nuclei have magic numbers of protons and neutrons?

[ NET/JRF (JUNE-2015)]

**a.**  ${}_{2}^{4}$ He and  ${}_{8}^{16}$ O

**b.**  ${}_{1}^{2}D$  and  ${}_{4}^{8}Be$ 

c.  ${}_{2}^{4}$ He and  ${}_{4}^{8}$ Be

**d.**  ${}_{2}^{4}$ He and  ${}_{6}^{12}$ C

8. According to the shell model the spin and parity of the two nuclei  $^{125}_{51}$ Sb and  $^{89}_{38}$ Sr are, respectively, [ NET/JRF (DEC-2011)]

**a.**  $\left(\frac{5}{2}\right)^+$  and  $\left(\frac{5}{2}\right)^+$ 

**b.**  $(\frac{5}{2})^{+}$  and  $(\frac{7}{2})^{+}$ 

c.  $\left(\frac{7}{2}\right)^+$  and  $\left(\frac{5}{2}\right)^+$ 

**d.**  $(\frac{7}{2})^{+}$  and  $(\frac{7}{2})^{+}$ 

9. According to the shell model, the total angular momentum (in units of  $\hbar$ ) and the parity of the ground state of the  ${}_{3}^{7}Li$  nucleus is

[ NET/JRF (DEC-2013)]

**a.**  $\frac{3}{2}$  with negative parity

**b.**  $\frac{3}{2}$  with positive parity

c.  $\frac{1}{2}$  with positive parity

**d.**  $\frac{7}{2}$  with negative parity

10. According to the shell model, the nuclear magnetic moment of the  $^{27}_{13}$ Al nucleus is (Given that for a proton  $g_l = 1, g_s = 5.586$ , and for a neutron  $g_l = 0, g_s = -3.826$ )

[NET/JRF (JUNE-2016)]

**a.**  $-1.913 \mu_{N}$ 

**b.**  $14.414 \mu_N$ 

c.  $4.793 \mu_{\rm M}$ 

**d.** 0

11. The spin-parity assignments for the ground and first excited states of the isotope  ${}_{28}^{57}$ Ni, in the single particle shell model, are

[ NET/JRF (DEC-2017)]

**a.**  $(\frac{1}{2})^{-}$  and  $(\frac{3}{2})^{-}$ 

**b.**  $(\frac{5}{2})^+$  and  $(\frac{7}{2})^+$ 

**c.**  $(\frac{3}{2})^{+}$  and  $(\frac{5}{2})^{+}$ 

**d.**  $\left(\frac{3}{2}\right)^{-}$  and  $\left(\frac{5}{2}\right)^{-}$ 

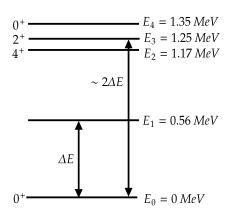
12. The first excited state of the rotational spectrum of the nucleus  $\frac{238}{92}U$  has an energy 45keV above the ground state. The energy of the second excited state (in keV) is

[ NET/JRF (DEC-2017)]

**a.** 150 **c.** 90 **ture** 

**d.** 60

13. The low lying energy levels due to the vibrational excitations of an even-even nucleus are shown in the figure below,



The spin-parity  $j^p$  of the level  $E_1$  is

[ NET/JRF (DEC-2018)]

**a.**  $1^+$ 

**b.** 1<sup>-</sup>

**c.** 2

**d.**  $2^+$ 

14. An excited state of a  $^8_4Be$  nucleus decays into two  $\alpha$ -particles which are in a spinparity 0+state. If the mean life-time of this decay is  $10^{-22}$  s, the spin-parity of the excited state of the nucleus is

1.5 Collective Model

**a.** 2<sup>+</sup>

**b.** 3<sup>+</sup>

**c.** 0<sup>-</sup>

**d.** 4<sup>-</sup>

Answer key			
Q.No.	Answer	Q.No.	Answer
1	С	2	b
3	a	4	с
5	С	6	С
7	a	8	d
9	a	10	c
11	d	12	
13	d	14	a
15			



# **Practice Set-2**

1.	•		0 01	tains a surface correction term
	This term depends on	the mass number A of t	ne nucleus as	[ GATE-2011
	<b>a.</b> $A^{-1/3}$	<b>b.</b> $A^{1/3}$	<b>c.</b> $A^{2/3}$	<b>d.</b> A
2.	The nuclear spin and j	parity of $^{40}_{20}$ Ca in its grou	and state is	
		20		[ GATE-2019
	<b>a.</b> $0^+$	<b>b.</b> 0 <sup>-</sup>	<b>c.</b> 1 <sup>+</sup>	<b>d.</b> 1 <sup>-</sup>
3.	In the nuclear shell me	odel, the spin parity of	$_{7}^{15}N$ is given by	
				[ GATE-2010
	<b>a.</b> $\frac{1^{-}}{2}$	<b>b.</b> $\frac{1^+}{2}$	<b>c.</b> $\frac{3^{-}}{2}$	<b>d.</b> $\frac{3^{+}}{2}$
4.		ear shell model, the resp	ective ground state spin-parit	ty values of ${}_{8}^{15}$ O and ${}_{8}^{17}$ O nucle
	are			[ GATE-2016
	<b>a.</b> $\frac{1^+}{2}, \frac{1^-}{2}$	<b>b.</b> $\frac{1}{2}^-, \frac{5^+}{2}$	<b>c.</b> $\frac{3^-}{2}, \frac{5^+}{2}$	<b>d.</b> $\frac{3^-}{2}, \frac{1^-}{2}$
5.	$J^P$ for the ground state	e of the ${}^{13}C_6$ nucleus is		
				[ GATE-2017
	a. 1 <sup>+</sup>	<b>b.</b> $\frac{3^{-}}{2}$	$c. \frac{3^{+}}{2}$	<b>d.</b> $\frac{1}{2}$
6.	According to the single	le particles nuclear shel	l model, the spin-parity of th	te ground state of ${}^{17}_{8}$ O is [ GATE-2011
	a. ½ C/a1	<b>b.</b> $\frac{3}{2}$	$U = \frac{3}{2} = $	<b>d.</b> $\frac{5^{+}}{2}$
7.		levels of $^{228}Th_{90}$ are sh		
		4+	187 keV	
		2+	57.5 keV	
		0+	0 keV	
	The expected spin-par	ity and energy of the ne	ext level are given by	[ GATE-2010
	<b>a.</b> $(6^+;400\text{keV})$		<b>b.</b> $(6^+;300\text{keV})$	
	<b>c.</b> $(2^+;400\text{keV})$		<b>d.</b> $(4^+;300 \text{keV})$	
8.	In the nuclear shell n	nodel, the potential is	modeled as $V(r) = \frac{1}{2}m\omega^2 r^2$	$2 - \lambda \vec{L} \cdot \vec{S}, \lambda > 0$ . The correct
	spin-parity and isospii	n assignments for the gr	ound state of 6°C 18	[ GATE-2015

9. The total angular momentum j of the ground state of the  ${}^{17}_{8}O$  nucleus is

**a.**  $\frac{1^-}{2}; \frac{-1}{2}$  **b.**  $\frac{1^+}{2}; \frac{-1}{2}$  **c.**  $\frac{3^+}{2}; \frac{1}{2}$ 

[ GATE- 2020]

**d.**  $\frac{3^{-}}{2}$ ;  $\frac{-1}{2}$ 

1.5 Collective Model

**a.**  $\frac{1}{2}$ 

**b.** 1

**c.**  $\frac{3}{2}$ 

**d.**  $\frac{5}{2}$ 

10. For nucleus  $^{164}$ Er, a  $J^{\pi}=2^+$ state is at 90keV. Assuming  $^{164}$ Er to be a rigid rotor, the energy of its  $4^+$ state is ——keV (up to one decimal place)

[ GATE-2018]

Answer key				
Q.No.	Answer	Q.No.	Answer	
1	С	2	a	
3	a	4	b	
5	d	6	d	
7	a	8	a	
9	d	10	300	
11		12		
13		14		
15				

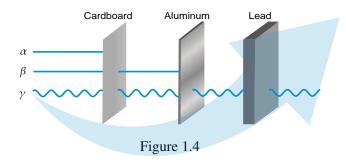


Many nuclides present in the universe are unstable and spondaneously change in to other nuclides by a process pulled radioactive decay. The phenomina is known as radioactivity and was discovered Antonine Bequerel. Following three aspects of radioactivity are extraordinary from the perspective of classical physics.

- (i) When a nucleus undergo  $\alpha$  or  $\beta$  decays its atomic number Z changes and it becomes the nucleus of a different form. Obviously the elements are not immutable. (not transformed to the original one naturally)
- (ii) The energy liberated during radioactive decay comes from with in the individual nuclei without external excitation.
- (iii) Radioactive decay is statistical process.

#### 1.6.1 Radioactive Decay

There are five different ways in which a radio action nuclide can decay. That are by emitting an alpha  $\binom{4}{2}He$  nuclei), beta(electrons),gamma(high energy photons) particles or by positron emission and electron capture. In which alpha particle is +vely charged and P particles are -vely charged and the gamma rays are natural. The penitrating power is maximum for gamma rays. The penitrating power of  $\alpha, \beta, \gamma$ , particle can be picturized as



	J	
Decay	Transformation	Example
Alpha decay	$_{2}^{A}X \rightarrow _{Z-2}^{A-4}Y + _{2}^{4}He$	$^{238}_{29}\text{U} \rightarrow ^{234}_{90}\text{Th} + ^{4}_{2}\text{He}$
Beta decay	${}_{2}^{A}X \rightarrow {}_{Z+1}^{A}Y + e^{-}$	$^{14}_{6}\text{C} \rightarrow ^{14}_{7}\text{N} + e^{-}$
Positron emission	${}_2^AX \rightarrow {}_{-1}^AY + e^+$	$^{64}_{29}{ m Cu}  ightarrow ^{64}_{28}{ m Ni} + e^+$
Electron capture	$_{Z}^{A}X+e^{-}\rightarrow_{z-1}^{A}Y$	$^{A}_{64}\mathrm{Cu}+e^{-} ightarrow ^{64}_{28}\mathrm{Ni}$
Gamma decay	$_{Z}^{A}X^{st} ightarrow _{Z}^{A}X+\gamma$	$^{87}_{87}$ Cu + $^{8}_{38}$ Sr* $\rightarrow$ $^{87}_{38}$ Sr + $\gamma$

The \* denotes an excited nuclear state and γ denotes a gamma-ray photon.

#### 1.6.2 Activity

The activity of a sample of any radioactive nuclide is the rate at which the nuclei of its constituent atoms decay. If N is the number of nuclei present in the sample at a certain time, its activity R is given by

$$R = \frac{-dN}{dt}$$

The minus sign is used to make R a positive quantity since  $\frac{dN}{dt}$  is negative. The time variation of activity is followed by the formula,

$$R = R_0 e^{-\lambda t}$$

Where  $\lambda$  is called decay constant or disintegration constant.

#### 1.6.3 **Radioactive Decay Law**

On emission of  $\alpha$  or  $\beta$  particle which is usually but not invariably accompanied by  $\gamma$ -ray emission, the emitting parent nuclide transforms in to a new daughter element. The daughter element again is radioactive. So that the process of successive disintegration continues till the original active parent nuclid get transformed in to a stable one.

(2) The rate of radioactive disintegration that is the number of atoms that break up at any instant of time t is directly preportional to the number N of active nuclides present in the sample at that instant.

In other words "The probability per unit time that a nucleus will decay is a constant and is independent of time".  $\lambda$  is the probability per unit time. Which is a constant

#### **Decay Equation**

The mathematical representation of the law of radioactive decay is

$$-\frac{dN}{dt}\alpha N\tag{1.1}$$

$$\frac{dN}{dt} = \lambda N, \lambda \text{decay constant}$$
 (1.2)

$$\frac{dN}{N} = -\lambda dt \tag{1.3}$$

$$\int \frac{dN}{N} = -\lambda \int dt \tag{1.4}$$

$$ln N = -\lambda t + A$$
(1.5)

at 
$$t = 0$$
  $N = N_0$  the initial number of unclides. (1.7)

$$A = \ln N_0 \tag{1.8}$$

$$\therefore \ln N = -\lambda t + \ln N_0 \tag{1.9}$$

$$\ln \frac{N}{N_0} = -\lambda t \tag{1.10}$$

$$\frac{N}{N_0} = e^{-\lambda t} \tag{1.11}$$

$$\frac{N}{N_0} = e^{-\lambda t} \tag{1.11}$$

$$N = N_0 e^{-\lambda t}$$
 which is the equation form of the law of radioactive decay (1.12)

we know, 
$$(1.13)$$

$$-\frac{dN}{dt} = R \tag{1.14}$$

and 
$$-\frac{dN}{dt} = \lambda N$$
 (1.15)

$$So.R = \lambda N \tag{1.16}$$

### **1.6.4** Half Life $(T_{\frac{1}{2}})$

Half life is the time  $(t = T_{\frac{1}{2}})$  at which the activity R drops to  $\frac{1}{2}R_0$ 

$$R = R_0 e^{-\lambda t}$$

$$\frac{1}{2} R_0 = R_0 e^{-\lambda T_{\frac{1}{2}}}$$

$$e^{\lambda} T_{\frac{1}{2}} = 2$$

$$\lambda T_{\frac{1}{2}} = \ln 2 = .693$$

$$T_{\frac{1}{2}} = \frac{\ln 2}{\lambda} = \frac{0.693}{\lambda}$$

From half-life the decay constant  $\gamma$  of a radioactive nuclei can be found out

$$\lambda = \frac{-693}{T_{\frac{1}{2}}}$$

The decay constant of radionuclide whose half life is 5h is

$$\lambda = \frac{0.693}{T_{\frac{1}{2}}} = \frac{0.693}{5 \times 3600 \text{ S}} = 3.85 \times 10^{-5} 8^{-1}$$

The larger the decay constant, the greater the chance the given nucleus will decay in a certain period of time

### 1.6.5 Mean Lifetimer $\langle \bar{T} \rangle$ or Average Life

The mean life time of a nuclide is the resiprocal of its decay probability per unit time.

$$ar{T} = rac{1}{\lambda}$$
Hence  $ar{T} = rac{T_{rac{1}{2}}}{0.693} = 1.44 T_{rac{1}{2}}$ 

 $ar{T}$  is nearly half again more than  $T_{rac{1}{2}}\left[\left(1+rac{1}{2}
ight)T_{rac{1}{2}}
ight]$ 

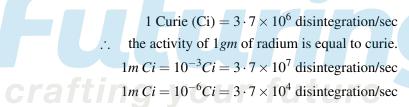
The mean life time of a radionuclide whose half life is 5hr is

$$\bar{T} = 1.44 \ T_{\frac{1}{2}} = 1.44 \times 5 = 7.2 hr$$

#### 1.6.6 Units of Radioactivity

1. Curie (Ci)

The traditional units of activity is curie (Ci). Which can be defined as the activity of 1g of radium. 1g radius have  $3 \cdot 7 \times 10^{10}$  disintegration per second . so



2. Rutherford (rd)

1 Rutherford 
$$(rd) = 10^6$$
 disintegration/sec  

$$1m \ rd = 10^{-3} \ rd$$

$$1m \ rd = 10^{-6} \ rd$$

3. Bequerel (Bq)

Bequerel is the SI unit of radioactivity

$$1Bq=1$$
 disintegration/sec  $1Ci=3\cdot7\times10^{10}Bq=37GB_q=37\times10^9Bq$   $MB_q=10^6Bq$   $GB_q=10^9Bq$ 

### **Practice Set 1**

1. The radius of a  $^{64}_{29}$ Cu nucleus is measured to be  $4.8 \times 10^{-13}$  cm.

(A) The radius of  $a_{12}^{27}Mg$  nucleus can be estimated to be

[NET/JRF(JUNE-2011)]

**A.**  $2.86 \times 10^{-13}$  cm

**B.**  $5.2 \times 10^{-13}$  cm

**C.**  $3.6 \times 10^{-13}$  cm

**D.**  $8.6 \times 10^{-13}$  cm

**(B)** The root-mean-square (r.m.s) energy of a nucleon in a nucleus of atomic number A in its ground state varies as:

**A.**  $A^{4/3}$ 

**B.**  $A^{1/3}$ 

 $C. A^{-1/3}$ 

**D.**  $A^{-2/3}$ 

2. The difference in the Coulomb energy between the mirror nuclei  $^{49}_{24}$ Cr and  $^{49}_{25}$ Mn is 6.0MeV. Assuming that the nuclei have a spherically symmetric charge distribution and that  $e^2$  is approximately 1.0MeV – fm, the radius of the  $^{49}_{25}$ Mn nucleus is

[NET/JRF(DEC-2011)]

**A.**  $4.9 \times 10^{-13}$  m

**B.**  $4.9 \times 10^{-15}$  m

**C.**  $5.1 \times 10^{-13}$  m

**D.**  $5.1 \times 10^{-15}$  m

3. The ground state of  ${}^{207}_{12}Pb$  nucleus has spin-parity  $J^p = \frac{1}{2}$ , while the first excited state has  $J^p = \frac{5}{2}$ . The electromagnetic radiation emitted when the nucleus makes a transition from the first excited state to ground state are

[NET/JRF(JUNE-2012)]

A. E2 and E3

**B.** M2 or E3

**C.** E2 or M3

**D.** M2 or M3

4. According to the shell model, the total angular momentum (in units of  $\hbar$ ) and the parity of the ground state of the  ${}_{3}^{7}Li$  nucleus is

[NET/JRF(DEC-2013)]

A.  $\frac{3}{2}$  with negative parity

**B.**  $\frac{3}{2}$  with positive parity

C.  $\frac{1}{2}$  with positive parity

**D.**  $\frac{7}{2}$  with negative parity

5. The recently-discovered Higgs boson at the LHC experiment has a decay mode into a photon and a Z boson. If the rest masses of the Higgs and Z boson are  $125 \text{GeV}/c^2$  and  $90 \text{ GeV}/c^2$  respectively, and the decaying Higgs particle is at rest, the energy of the photon will approximately be

[NET/JRF(JUNE-2014)]

**A.**  $35\sqrt{3}$  GeV

**B.** 35 GeV

**C.** 30 GeV

**D.** 15 GeV

6. The reaction  ${}^2_1D + {}^2_1D \rightarrow {}^4_2\text{He} + \pi^0$  cannot proceed via strong interactions because it violates the conservation of

[NET/JRF(JUNE-2015)]

A. Angular momentum

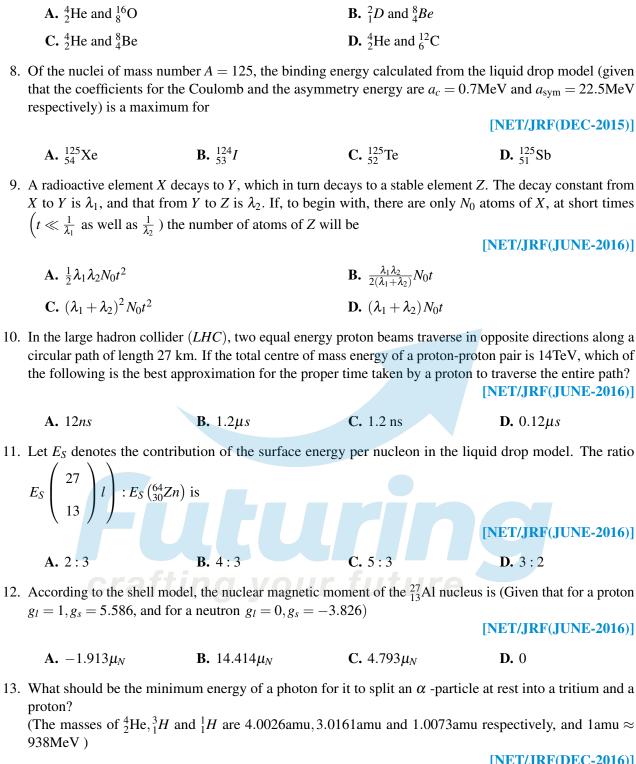
B. Electric charge

C. Baryon number

D. Isospin

7. Let us approximate the nuclear potential in the shell model by a three dimensional isotropic harmonic oscillator. Since the lowest two energy levels have angular momenta l = 0 and l = 1 respectively, which of the following two nuclei have magic numbers of protons and neutrons?

[NET/JRF(JUNE-2015)]



[**NET/JRF(DEC-2016)**]

A. 32.2MeV

B. 3MeV

C. 19.3MeV

**D.** 931.5MeV

14. If in a spontaneous  $\alpha$  - decay of  $\frac{232}{92}U$  at rest, the total energy released in the reaction is Q, then the energy carried by the  $\alpha$  - particle is

[NET/JRF(JUNE-2017)]

**A.** 570/58

**B.** *O*/57

**C.** O/58

**D.** 23*Q*/58

15. The range of the nuclear force between two nucleons due to the exchange of pions is 1.40fm. If the mass of pion is  $140 \text{MeV}/c^2$  and the mass of the rho-meson is  $770 \text{MeV}/c^2$ , then the range of the force due to exchange of rho-mesons is

[NET/JRF(JUNE-2017)]

- **A.** 1.40fm
- **B.** 7.70fm
- C. 0.25fm
- **D.** 0.18fm
- 16. The first excited state of the rotational spectrum of the nucleus  ${}^{238}_{92}U$  has an energy 45keV above the ground state. The energy of the second excited state (in keV) is

[**NET/JRF(DEC-2017)**]

- **A.** 150
- **B.** 120
- **C.** 90

- **D.** 60
- 17. The reaction  $^{63}Cu_{29} + p \rightarrow ^{63}Zn_{30} + n$  is followed by a prompt  $\beta$ -decay of zinc  $^{63}Zn_{30} \rightarrow ^{63}Cu_{29} + e^+ + v_e$ . If the maximum energy of the position is 2.4MeV, the Q value of the original reaction in MeV is nearest to [Take the masses of electron, proton and neutron to be  $0.5 \text{MeV/c}^2$ , 938MeV/c<sup>2</sup> and 939.5MeV/c<sup>2</sup>, respectively.]

[NET/JRF(JUNE-2018)]

- A. -4.4
- **B.** -2.4
- $\mathbf{C.} -4.8$
- **D.** -3.4
- 18. A nucleus decays by the emission of a gamma ray from an excited state of spin parity  $2^+$  to the ground state with spin-parity  $0^+$  what is the type of the corresponding radiation?

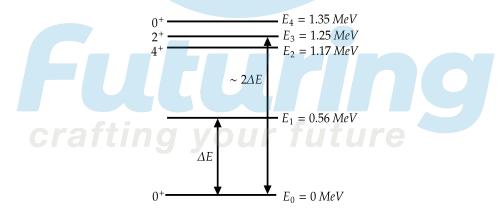
[**NET/JRF**(**DEC-2018**)]

**A.** Magnetic dipole

B. Electric quadrupole

C. Electric dipole

- D. Magnetic quadrupole
- 19. The low lying energy levels due to the vibrational excitations of an even-even nucleus are shown in the figure below.



The spin-parity  $j^p$  of the level  $E_1$  is

[NET/JRF(DEC-2018)]

**A.**  $1^{+}$ 

**B.** 1<sup>-</sup>

 $C. 2^{-}$ 

- **D.**  $2^{+}$
- 20. The Bethe-Weizsacker formula for the binding energy (in MeV) of a nucleus of atomic number *Z* and mass number *A* is

$$15.8A - 18.3A^{2/3} - 0.714\frac{Z(Z-1)}{A^{1/3}} - 23.2\frac{(A-2Z)^2}{A}$$

The ratio Z/A for the most stable isobar of a A = 64 nucleus, is nearest to

[NET/JRF(DEC-2019)]

- **A.** 0.30
- **B.** 0.35
- **C.** 0.45
- **D.** 0.50
- 21. A negative muon, which has a mass nearly 200 times that of an electron, replaces an electron in a Li atom. The lowest ionization energy for the muonic Li atom is approximately

[NET/JRF(DEC-2019)]

- **A.** The same as that of He
- **B.** The same as that of normal *Li*
- C. 200 times larger than that of normal Li
- **D.** The same as that of normal Be
- 22. The binding energy B of a nucleus is approximated by the formula  $B = a_1A a_2A^{2/3} a_3Z^2A^{-1/3} a_4(A-2Z)^2A^{-1}$ , where Z is the atomic number and A is the mass number of the nucleus. If  $\frac{a_4}{a_2} \simeq 30$ . The atomic number Z for naturally stable isobars (constant value of A) is

[NET/JRF(JUNE-2020)]

**A.** 
$$\frac{30A}{60+A^{2/3}}$$

**B.** 
$$\frac{30A}{30+A^{2/3}}$$

C. 
$$\frac{60A}{120+A^{2/3}}$$

**D.** 
$$\frac{120A}{60+A^{2/3}}$$

23. The magnetic moments of a proton and a neutron are  $2.792\mu_N$  and  $-1.913\mu_N$ , where  $\mu_N$  is the nucleon magnetic moment. The values of the magnetic moments of the mirror nuclei  $^{19}_{9}F_{10}$  and  $^{19}_{10}Ne_9$ , respectively, in the Shell model, are closest to

[NET/JRF(JUNE-2020)]

**A.** 23.652
$$\mu_N$$
 and  $-18.873\mu_N$ 

**B.** 26.283
$$\mu_N$$
 and  $-16.983\mu_N$ 

**C.** 
$$-2.628\mu_N$$
 and  $1.887\mu_N$ 

**D.** 
$$2.628\mu_N$$
 and  $-1.887\mu_N$ 



# **Practice Set 2**

1	In the nuclear shell r	model the spin parity of $\frac{1}{7}$	<sup>5</sup> N is given by		
1.	in the nacion shen i	noder the spin parity of 7	iv is given by	[GAT	TE 2010]
	<b>A.</b> $\frac{1^{-}}{2}$	<b>B.</b> $\frac{1^+}{2}$	<b>C.</b> $\frac{3^{-}}{2}$	<b>D.</b> $\frac{3^{+}}{2}$	
2.	The ground state was	vefunction of deuteron is	in a superposition of s and	d states. Which of the fe	ollowing
	is NOT true as a con	sequence?		[GAT	TE 2010]
	A. It has a non-ze	ero quadruple moment			
	<b>B.</b> The neutron-p	roton potential is non-cer	ntral		
	C. The orbital wa	vefunction is not spheric	ally symmetric		
	<b>D.</b> The Hamilton	ian does not conserve the	total angular momentum		
3.	The first three energy	y levels of $^{228}Th_{90}$ are sh	own below		
		4+	187 keV		
		2+	57.5keV		
		0+	0 keV		
	The expected spin-particle A. (6+;400keV)	arity and energy of the ne	ext level are given by $\mathbf{B.}\ (6^+;300\text{keV})$	[GAT	TE 2010]
	<b>C.</b> (2 <sup>+</sup> ;400keV)		<b>D.</b> (4 <sup>+</sup> ;300keV)		
4.	The semi-empirical	mass formula for the bin	ding energy of nucleus co	ntains a surface correct	ion term
	This term depends o	n the mass number A of the	he nucleus as	[GAT	TE 2011]
	<b>A.</b> $A^{-1/3}$	<b>B.</b> $A^{1/3}$	<b>C.</b> $A^{2/3}$	<b>D.</b> A	
5.	According to the sin	gle particles nuclear shell	model, the spin-parity of		is T <b>E 2011</b> ]
	<b>A.</b> $\frac{1}{2}$	<b>B.</b> $\frac{3}{2}$	<b>C.</b> $\frac{3^{+}}{2}$	<b>D.</b> $\frac{5^{+}}{2}$	
6.	In the $\beta$ decay proce	ess, the transition $2^+ \rightarrow 3$	<sup>+</sup> , is	[GAT	TE 2013]
	<b>A.</b> Allowed both	by Fermi and Gamow-Te	ller selection rule		
	<b>B.</b> Allowed by Fe	ermi and but not by Gamo	ow-Teller selection rule		

C. Not allowed by Fermi but allowed by Gamow-Teller selection rule

**D.** Not allowed both by Fermi and Gamow-Teller selection rule

7. Which one of the following is a fermions'?

[GATE 2014]

**A.**  $\alpha$ -particle

**B.**  $_4Be^7$  nucleus

C. Hydrogen atom

**D.** Deuteron

8. The atomic masses of  ${}^{152}_{63}Eu, {}^{152}_{62}\text{Sm}, {}^1_1H$  and neutron are 151.921749, 151.919756, 1.007825 and 1.008665 in atomic mass units (amu), respectively. Using the above information, the Q - value of the reaction  ${}^{152}_{63}Eu+n \rightarrow {}^{152}_{62}Sm+p$  is————  $\times 10^{-3}$  amu (upto three decimal places)

[GATE 2015]

9. In the nuclear shell model, the potential is modeled as  $V(r) = \frac{1}{2}m\omega^2r^2 - \lambda\vec{L}\cdot\vec{S}, \lambda > 0$ . The correct spin-parity and isospin assignments for the ground state of  $^{13}_6C$  is

[GATE 2015]

**A.** 
$$\frac{1^{-}}{2}$$
;  $\frac{-1}{2}$ 

**B.** 
$$\frac{1^+}{2}$$
;  $\frac{-1}{2}$ 

C. 
$$\frac{3^+}{2}$$
;  $\frac{1}{2}$ 

**D.** 
$$\frac{3^{-}}{2}$$
;  $\frac{-1}{2}$ 

10. According to the nuclear shell model, the respective ground state spin-parity values of  ${}_{8}^{15}O$  and  ${}_{8}^{17}O$  nuclei are

[GATE 2016]

**A.** 
$$\frac{1^+}{2}, \frac{1^-}{2}$$

**B.** 
$$\frac{1}{2}^-, \frac{5^+}{2}$$

$$C. \frac{3^-}{2}, \frac{5^+}{2}$$

**D.** 
$$\frac{3^{-}}{2}, \frac{1^{-}}{2}$$

[GATE 2017]

[GATE 2018]

Answer key			
Q.No.	Answer	Q.No.	Answer
41119	В	2	
3	A	4	C
5	D	6	C
7	В	8	2.833
9	A	10	В
11	29.84	12	300

## **Practice Set-3**

1. Using the liquid drop model, find the expression of the most stable isobar for a given A. Find the stable atom with A = 77. Here  $a_c = 0.58$ MeV and  $a_a = 19.3$ MeV.

#### **Solution:**

$$BE = a_{\nu}A - a_{s}A^{\frac{2}{3}} - a_{c}Z^{2}A^{-\frac{1}{3}} - a_{a}(A - 2Z)^{2}A^{-1} \pm \frac{\varepsilon}{a^{3/4}}$$

For most stable isobar (A=constant) is the one with the maximum binding energy.

$$\frac{d(BE)}{dZ} = -2a_c A^{-\frac{1}{3}} Z + 4a_a (A - 2Z) A^{-1} = 0$$

$$Z = \frac{4a_a}{2a_c A^{-\frac{1}{3}} + 8a_a A^{-1}} = \frac{A}{\left(\frac{a_c}{2a_a}\right) A^{2/3} + 2}$$

Using  $a_c = 0.58 \text{MeV}$  and  $a_a = 19.3 \text{MeV}$ 

$$Z = \frac{A}{0.015 \, A^{2/3} + 2}$$

For 
$$A = 77$$
,

$$Z = \frac{77}{0.015(77)^{2/3} + 2} = 33.9$$

thus  $_{34}^{77}Se$  is stable while  $_{33}^{77}As$  and  $_{33}^{77}Br$  are unstable

2. Explain which is the most stable among <sub>2</sub>He<sup>6</sup>, <sub>4</sub>Be<sup>6</sup> and <sub>3</sub>Li<sup>6</sup>.

#### Solution:

He, Be and Li are all light nuclei for which  $0.015A^{2/3}$  is negligible or  $2 \gg 0.015A^{2/3}$ .

Hence for most stable nuclei  $Z \sim \frac{A}{2}$ 

Hence out of three nuclei, 3Li<sup>6</sup> is most stable

- 3. Show by way of computation, which nuclei you would expect to be more stable.
  - (a)  $_3Li^7$  or  $_3Li^8$ (b)  $_4Be^9$  or  $_4Be^{10}$

#### **Solution:**

(a) for 
$$A = 7, Z = \frac{7}{2 + 0.015(7)^{2/3}} = 3.4$$

and for 
$$A = 8, Z = \frac{8}{2 + (0.015)(8)^{2/3}} = 3.88$$

3.4 is more closer to 3 . So,  $_3Li^7$  is more stable.

(b) For 
$$A = 9$$
,  $Z = \frac{9}{2 + (0.015)^2 (9)^{2/3}} = 4.36$ 

and 
$$A = 10, Z = \frac{10}{2 + (0.015)(10)^{2/3}} = 4.80$$

4.36 is more close to 3, so 4 Be 9 is more stable.

4. Using the semi- empirical binding energy formula, calculate the binding energy of 20Ca<sup>40</sup>.

Here  $a_v = 15.75 \text{MeV}$ ,  $a_s = 17.80 \text{MeV}$ ,  $a_c = 0.71 \text{MeV}$ ,  $a_n = 22.7 \text{MeV}$  and delta = 34 MeV.

**Solution:** 

$$BE = a_v A - a_s A^{\frac{2}{3}} - a_c \frac{z(z-1)}{A^{1/3}} - a_n \frac{(4-2z)^2}{A} + \frac{\delta}{A^{3/4}}$$

$$= 680 - 208.2 - 84.3 - 0 + 2.14$$

$$= 339.64 \text{MeV}$$

5. Establish the relation  $\mathbf{A} = 2\mathbf{Z}$  for light nuclei using the semi-empirical mass formula, given  $a_c = \mathbf{0.71} \text{MeV}$ ,  $a_n = \mathbf{22.7} \text{MeV}$ , mass of  $_1\text{H}^1 = 1.0078$  units and mass of neutron = 1.0086 units.

$$M = ZM_H + (A - Z)M_n - \frac{1}{c^2} \left[ a_v A - a_s A^{\frac{2}{3}} - a_c \frac{Z^2}{A^{1/3}} - a_n \frac{(A - 2Z)^2}{A} \pm \frac{\delta}{A^{3/4}} \right]$$

The most stable nuclei for a given A (Isobars)

$$\left(\frac{\partial M}{\partial Z}\right)_{A} = 0$$

$$(M_{H} - M_{n})c^{2} + 2a_{c}\frac{Z}{A^{1/3}} - 4a_{n}\frac{(A - 2Z)}{A} = 0$$

$$\frac{2Z}{A}\left[a_{c}A^{2/a} + 4a_{n}\right] = (M_{n} - M_{H})c^{2} + 4a_{n}$$

$$\frac{2Z}{A}\left[\frac{a_{c}A^{\frac{2}{3}}}{4a_{n}} + 1\right] = \frac{(M_{n} - M_{H})c^{2}}{4a_{n}} + 1$$

$$Z - \frac{A}{2} \left[ \frac{1 + \frac{M_n - M_H)c^2}{4a_n}}{1 + \frac{a_c A^{2/3}}{4a_n}} \right]$$

$$Z = \frac{A}{2} \left[ \frac{1 + 0.0086}{1 + (0.0078)A^{2/3}} \right] = \frac{A}{1.98 + 0.0156A^{2/3}} \sim \frac{A}{2}$$

6. Using the semi- empirical mass formula, show that the neutron excess is approximately proportional to  $A^{5/3}$  for the families of isobars having odd **A** nuclei. Given  $a_c = \mathbf{0.71MeV}$ ,  $a_n = 22.7 \text{MeV}$ , mass of  $_1\text{H}^1 = 1.0078$  units and mass of neutron = 1.0086 units.

#### **Solution:**

$$M = ZM_H + (A - Z)M_n - \frac{1}{c^2} \left[ a_v A - a_s A^{\frac{2}{3}} - a_c \frac{Z^2}{A^{\frac{1}{3}}} - a_n \frac{(A - 2Z)^2}{A} \pm E_{\delta} \right]$$

For odd *A* nuclei,  $E_{\delta} = 0$ 

For most stable nucleus in a family of isobars

$$\left(\frac{\partial M}{\partial Z}\right)_A = 0$$

$$\left(\frac{\partial M}{\partial Z}\right)_A = (M_H - M_n)c^2 + 2a_c \frac{Z}{A^{1/3}} - 4a_n \frac{(A - 2Z)}{A} = 0$$

Adding and subtracting  $a_c A^{2/3}$ 

$$(M_H - M_n)c^2 + a_c A^{2/3} - a_c \frac{A}{A^{1/3}} + 2a_c \frac{Z}{A^{1/3}} - \frac{4a_n(A - 2Z)}{A} = 0$$

$$(A - 2Z) \left[ \frac{a_c A^{\frac{2}{3}}}{A} + \frac{4a_n}{A} \right] = a_c A^{2/3} - (M_n - M_H)c^2$$

$$= \frac{A \left[ \frac{a_c}{4a_n} A^{2/3} - \frac{(M_n - M_H)c^2}{4a_n} \right]}{1 + \frac{a_c}{4a_n} A^{2/3}}$$

$$= \frac{A0.0078A^{2/3} - 0.0086}{1 + 0.0078A^{2/3}}$$

$$\frac{A^{5/3} \times 00078 \left[ 1 + \frac{1.1}{A^{2/3}} \right]}{1 + 0.0078A^{2/3}}$$
For A varying from 60 to 210
$$(A - 2Z) \sim 0.006A^{5/3}$$

- 7. (a). For mirror nuclei which have N and Z differing by one unit, determine the mass difference. Consider A to be odd.
  - (b). The masses of  $_7N^{15}$  and  $_8O^{15}$  are 15.000108u and 15.003070u respectively. Using this data, determine the Coulomb coefficient  $a_c$  in the semi-empirical mass formula. Here  $\mathbf{M}_{\rm n} \mathbf{M}_{\rm p} = 0.000844u$

#### **Solution:**

(a) 
$$M_{Z+1} - M_Z = (M_p - M_n)[(Z+1) - 2]$$
  
 $+ \frac{a_c}{A^{1/3}}[(Z+1)^2 - Z^2)$   
 $= (M_p - M_n) + a_c \frac{(2Z+1)}{A^{1/3}}$   
 $= M_p - M_n + a_c \frac{A}{A^{2/3}}$   
 $M_{Z+1} - M_2 = (M_p - M_n) + a_c A^{2/3}$   
(b)  $M_{Z+1} - M_Z = (15.003070 - 15.000108)u$   
 $= 2.962 \times 10^{-3}u$   
 $M_p - M_n = -0.000844u$   
 $(2.962 \times 10^{-3})u = (-0.000844)u + a_c (15)^{2/3}$   
 $a_c (6.08) = (2.962 + 0.844) \times 10^{-3}u$   
 $= 3.806 \times 10^{-3} \times 931 \text{MeV}$   
 $a_c = \frac{3.542}{6.08} = 0.58 \text{MeV}$ 

8. Deduce that with  $a_c = 0.72 \text{MeV}$  and  $a_s = 23 \text{MeV}$  the ratio  $\mathbf{Z}_{\min}/\mathbf{A}$  is approximately 0.5 for light nuclei and 0.4 for heavy nuclei.

#### **Solution:**

Using the mass formula one can deduce the atomic number of the most stable isobar. It is given by

$$Z_{min} = \frac{A}{2 + \left(\frac{a_c}{2a_s}\right) A^{2/3}}$$
$$\frac{Z_{min}}{A} = \frac{1}{2 + 0.0156A^{2/3}}$$

For light nuclei, say A = 10, the second term in the denominators is small, and  $Z_{min}/A = 0.48$ .

For heavy nuclei, say  $A = 200, Z_{\min}/A = 0.4$ 

