

At t = 0 we have a spherical cloud of radius R and total charge Q, comprising N point-like particles. Each particle has charge q = Q/N and mass m. The particle density is uniform, and all particles are at rest.

a) Evaluate the electrostatic potential energy of a charge located at a distance r < R from the center at t = 0.

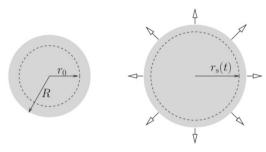


Fig. 1.4

sion, the cloud begins to expand radially, keeping its spherical symmetry. Assume that the particles do not overtake one another, i.e., that if two particles were initially located at  $r_1(0)$  and  $r_2(0)$ , with  $r_2(0) > r_1(0)$ , then  $r_2(t) > r_1(t)$  at any subsequent time t > 0. Consider the particles located in the infinitesimal spherical shell

b) Due to the Coulomb repul-

 $r_0 < r_s < r_0 + dr$ , with  $r_0 + dr < R$ , at t = 0. Show that the equation of motion of the layer is

$$m\frac{\mathrm{d}^2 r_{\mathrm{S}}}{\mathrm{d}t^2} = k_{\mathrm{e}} \frac{qQ}{r_{\mathrm{s}}^2} \left(\frac{r_0}{R}\right)^3 \tag{1.16}$$

**c**) Find the initial position of the particles that acquire the maximum kinetic energy during the cloud expansion, and determinate the value of such maximum energy.

their final kinetic energy. Compare the total kinetic energy with the potential energy initially stored in the electrostatic field.

e) Show that the particle density remains spatially uniform during the expansion.

a)

As N is large, it can be approximated to a continuous charge distribution.

$$V = -\int E dI + -\int E dI$$

Field outside the sphere = 
$$1 \times 2$$
 $V = -\int_{E}^{R} E \cdot dI + -\int_{E}^{R} E \cdot dI$ 

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$$\frac{-Q}{4\pi\epsilon_0} \int_{82}^{1} d8 \frac{Q}{4\pi\epsilon_0} \int_{82}^{1} \frac{1}{8} \frac{Q}{4\pi\epsilon_0} R$$

Field inside the sphere

$$E \times 4 \pi \sigma^2 = Q \times \frac{\chi^3}{R^3}$$
 Chauss's Law 
$$\mathcal{E}_0$$

$$\frac{R3}{\varepsilon_0} = \frac{R3}{\varepsilon_0} = \frac{R3}{4\pi\varepsilon_0 R^3} = \frac{R3$$

$$\frac{\sqrt{4\pi\xi_0}R^3}{\sqrt{4\pi\xi_0}R^3} \left[\frac{R^2-\chi^2}{Z}\right]$$

Total potential = 
$$QR^2 + Q \left[R^2 - 8^2\right]$$
  
at the point  $hIE_0R^3 + IIE_0R^3 \left[\frac{R^2 - 8^2}{2}\right]$ 

$$h_{5} \frac{Q}{R^3} \left[ \frac{3}{3} R^2 - \frac{\chi^2}{2} \right]$$

field at the infinitesmal layer at rs

charge enclosed = constant throughout

time for the layer =  $Q\left(\frac{r_0}{R}\right)^3$ at some later

time =  $Q\left(\frac{r_0}{R}\right)^3$   $Q\left(\frac{r_0}{R}\right)^3$   $Q\left(\frac{r_0}{R}\right)^3$   $Q\left(\frac{r_0}{R}\right)^3$   $Q\left(\frac{r_0}{R}\right)^3$   $Q\left(\frac{r_0}{R}\right)^3$   $Q\left(\frac{r_0}{R}\right)^3$ For one particle =  $\frac{Q}{d+2}\left(\frac{r_0}{R}\right)^3$   $Q\left(\frac{r_0}{R}\right)^3$   $Q\left(\frac{r_0}{R}\right)^3$ 

 $F \cdot dr = \Delta KE$   $= \frac{Qq}{4\pi z_0} \left(\frac{r_0}{R}\right)^3 \int_{S_2^2}^{1} \cdot ds$   $= \frac{Qq}{4\pi z_0} \left(\frac{r_0}{R}\right)^3 \times 1 = \frac{2}{\pi z_0}$ 

 $= \underbrace{Qq}_{h \Im \mathcal{E}_{D}} \underbrace{\left( \underbrace{x_{D}}_{R} \right)^{3} \times \underline{l}}_{h \Im \mathcal{E}_{D}} = \underbrace{Qq}_{h \Im \mathcal{E}_{D}} \underbrace{x_{D}}_{R 3}$   $KE \max = \underbrace{Qq}_{h \Im \mathcal{E}_{D}} \underbrace{R^{2}}_{h \Im \mathcal{E}_{D}} = \underbrace{Qq}_{h \Im \mathcal{E}_{D}} \underbrace{R}_{A}$ 

PE for outer layer = OG

HIROR

Outer layer only -> initial PE = final KE

not true for other layers, as final potentia

not true for other layers, as final potential will not be D Innumment point does not move at all d) Posticle density remains spatially uniform through time introduce a new variable x = Ys  $m \cdot \frac{d^2x}{d+2} = \underbrace{q \cdot Q \times^2}_{4\pi \epsilon_0 R^3} = \underbrace{q \cdot Q \times^2}_{4\pi \epsilon_0 R^3}$  $\frac{d^2r_0}{dt^2} = r_0 \cdot \frac{d^2x}{dt^2} = \frac{q_1Q_1}{dt^2} \times \frac{2}{dt^2} + \frac{q_1Q_1}{dt^2} \times \frac{2}{dt^2} \times \frac{2}{dt^2} + \frac{q_1Q_1}{dt^2} \times \frac{2}{dt^2} \times \frac{2}{d$ number of particles contained between them =  $N \times (3_{10}^3 - 3_{20}^3)$ this remains constant, because layers do not over  $R^3$  each other So at a time t, the particle density =  $\frac{1}{20^3}$ 50 x (+) is independent  $\frac{4}{2}J(x_{1s}^3 - x_{2s}^3)$   $x_{1s} = n(t)x_{20}$ of ro - valid for all parricles in the cloud  $\frac{457(7)^3-823)\times(1)^3}{3}$  =) independent of choice of shells  $\Rightarrow$  spanally uniform  $\Rightarrow$ 

=) Independent of choice of shells 
$$\Rightarrow$$
 spanally the layers with more particles are further apart