

Data Analysis and Preliminary Results of One-Ton WbLS Detector

detector response study

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NY-BNL
Spring, 2020



Data Analysis
of WbLS
Dector
R. Zhao

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about the
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waveform
analysis

PMT
Waveform
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Calibration of
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Problems
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Brief Introduction about the 1Ton Dector

waveform analysis

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backup

others



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about

- ▶ study the optical properties of WbLS and its performance in 1Ton dector.
- ▶ improve optical model of WbLS

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- ▶ run detector with water ↓
- ▶ calibrate detector response using data and MC simulation. ↓
- ▶ run detector with WbLS ↓
- ▶ analysis data with calculation factors ↓
- ▶ improve optical model of WbLS by tuning MC parameters.

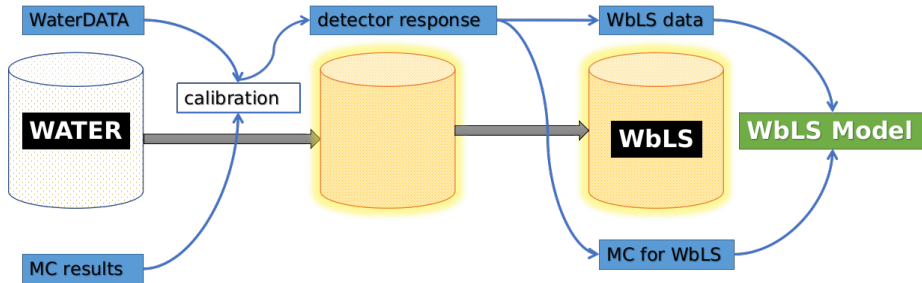


图: flowchart of analysis

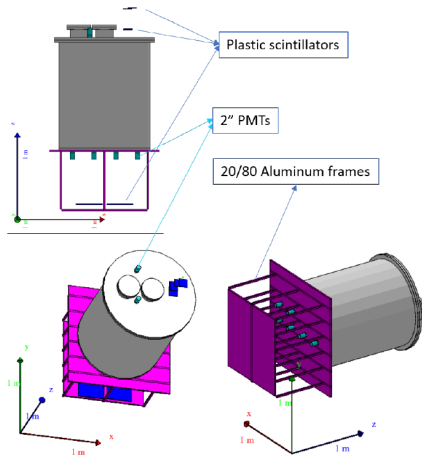


图: the 1Ton WbLS Dector

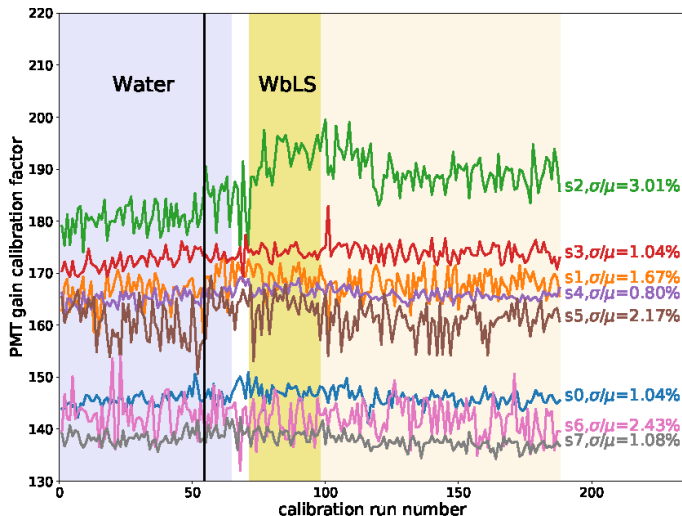
- ▶ ~ 1Ton WbLS is cylinder tank.
- ▶ 8 PMTs(2") for photon readout.
- ▶ 6 Hodoscope as trigger system.
- ▶ one 410nm LED for PMT calibration.

the Data Trigger Types :

1. hodoscope trigger:
 $(H0 \parallel H2) \&\& (H3 \parallel H1) \parallel H4 \parallel H5$
2. multiplicity trigger:
 $N_{fired-PMTs} \geq 6$
3. led trigger:
PMT single p.e. calibration



The stability of detector during data taking from LED data:





the Average Waveform of 3 Trigger Types

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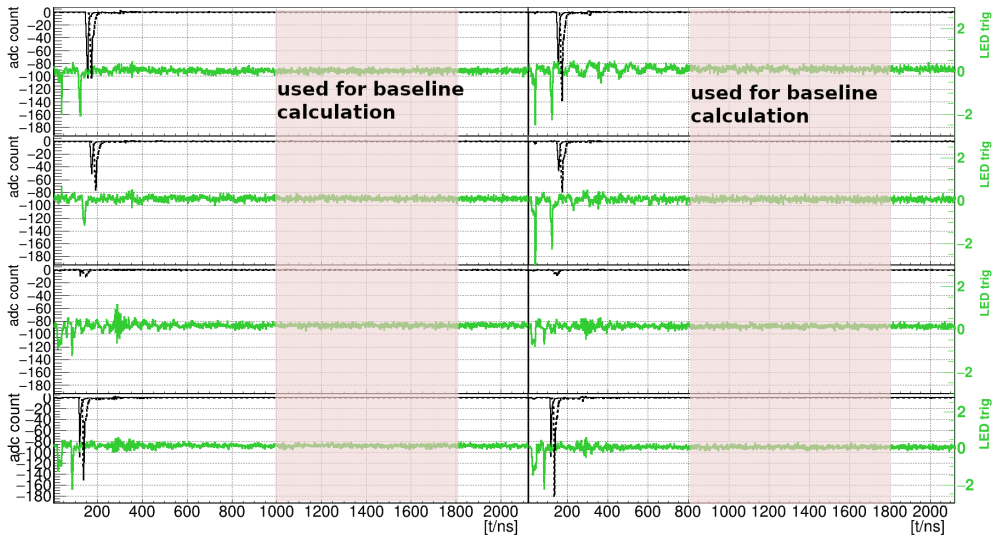


图: the 2560ns average waveform of PMTs, run 23456

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the Average Waveform of 3 Trigger Types

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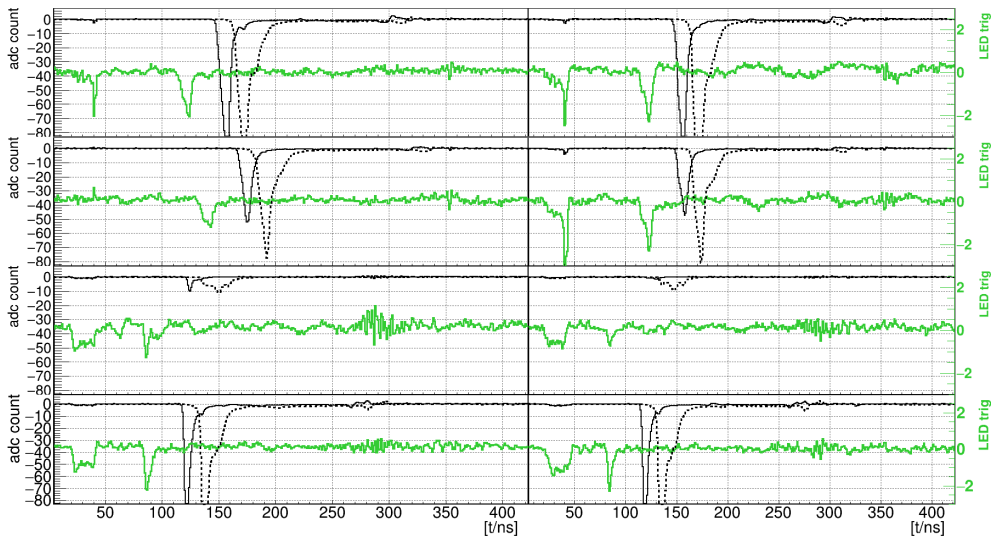


图: the first 400ns average waveform of PMTs, run 23456

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here the example wave

1. **baseline calculation:** choose a fix window from 1000ns to 1800ns, get the mean value as baseline of one waveform.
2. **signal seek threshold:** $baseline - 3 \cdot \sigma_{baseline}$
3. **charge integration:** from risetime-2 to risetime+2
4. **time coincidence:** if the rise time of a group of signals are inside a ns window, if the arrive time match \Rightarrow one effective group of signal; merge to peaks if the arrive time difference $< 25ns$

.....
a fixed-window of 40ns width is selected for LED spe calibration and fitting with convolution of poisson and gaussianfunction:

$$S(x) = P(n; \mu) \otimes G_n(x) = \sum_{n=0}^{\infty} \frac{\mu^n e^{-\mu}}{n!} \frac{1}{\sigma_1 \sqrt{2n\pi}} \exp\left(-\frac{(x - nQ_1)^2}{2n\sigma_1^2}\right) \quad (1)$$



PMT SPE Calibration using LED Trigger Data

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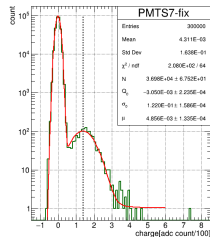
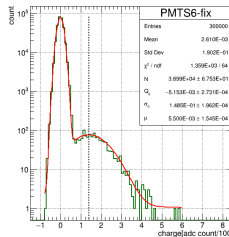
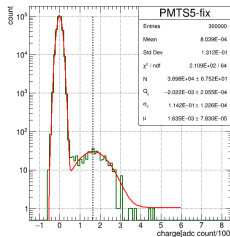
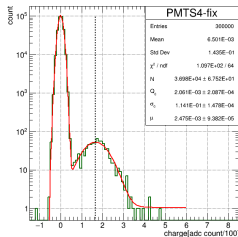
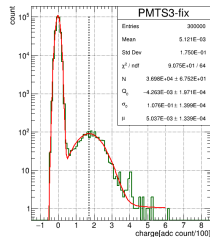
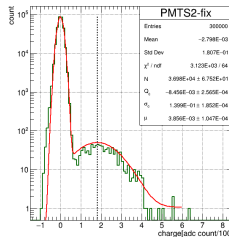
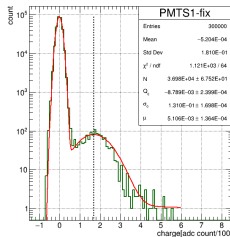
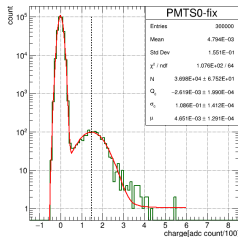


图: the LED single p.e. spectrum of PMTs for gain calibration

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calibrate the "effective" efficiency factors:

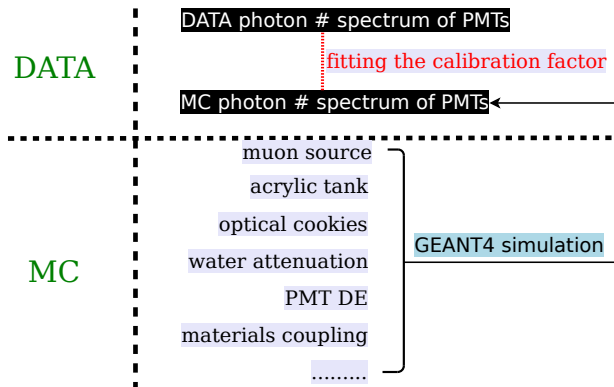


图: the LED single p.e. spectrum of PMTs for gain calibration



- ▶ CRY muon generator, @1m above the dector,
- ▶ Rat-Pac as framework
- ▶ use spe response of PMT data.
- ▶ the npe spectrum of each PMT.
- ▶ use nominal values of optical parameters for acrylic,cookie, PMT.
- ▶ 20m water attenuation by default.

Due to the limitation from muon generator, some high energy events can not be properly simulated,so choose a threshold of counts for the last bin of histogram.



The defination of χ^2 function in spectrum fitting:

$$\chi^2(f) = \sum_{i=1}^{i=\maxbin-1} \frac{N_i^{data} - N(f)_i^{MC}}{\sigma(f)_i} \quad (2)$$

where

- ▶ N_i^{data} is count in the i th bin of data N.p.e spectrum;
- ▶ f is the calibration factor;
- ▶ $N(f)_i^{MC}$ is count in the i th bin of MC N.p.e spectrum with f applied (multiplied) to each event;
- ▶ $\sigma(f)_i = \sqrt{N(f)_i^{MC}/r^2 + N_i^{data}}$ is the uncertainty of i th bin; $r=10$ is the ratio of total event number from MC and data ¹.

¹the MC histogram is scaled with $r = 10$ in the fitting

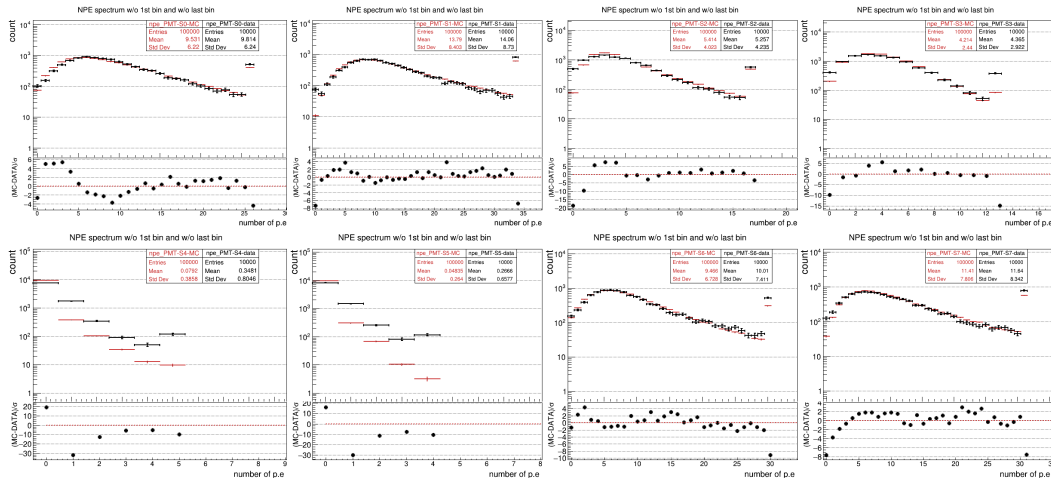


图: the N.p.e spectrum fitting of 8 PMTs, with 40 for last bin threshold, fitting without 1st bin without last bin



fitting χ^2 results N.p.e

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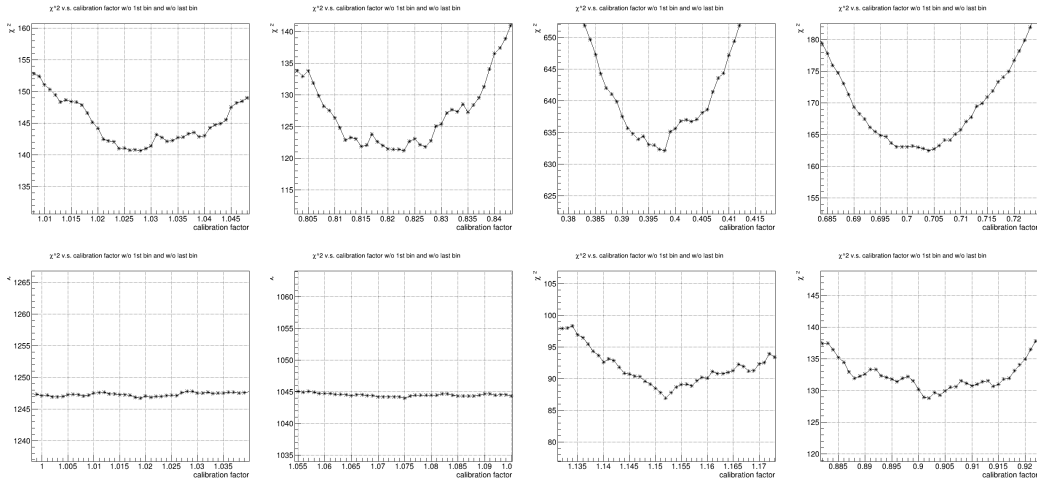


图: the fitting χ^2 of 8 PMTs, with 40 for last bin threshold, fitting without 1st bin without last bin



we considered several different factors in the fitting:

- ▶ the threshold of overbin
- ▶ uncertainty of water attenuation length
- ▶ first bin position
- ▶ influence of first and last bin in fitting
- ▶ PMT gain calibration uncertainty
- ▶ combinations top hodoscopes

表: mean value and uncertainty of calibration factors

pmt#	s0	s1	s2	s3	s4	s5	s6	s7
cf mean	1.032	0.823	0.395	0.708	0.984	1.089	1.152	0.902
uncertainty	0.051	0.052	0.025	0.057	0.041	0.016	0.043	0.03



the uncertainty evaluation of fitting results

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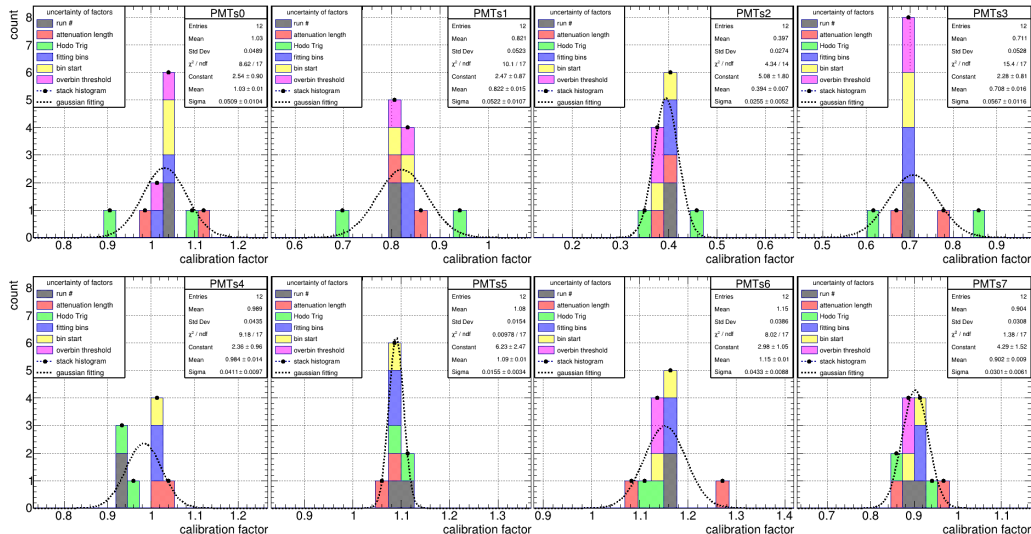


图: the stack histogram of calibration factor from different tests



calibration factors for 9 trigger combinations

|trigger

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Hodoscope trigger patterns

hodoscope#	1010	1001	0110	0101	1110	1101	0111	1011	1111
hodoscope 2	1	1	0	0	1	1	0	1	1
hodoscope 0	0	0	1	1	1	1	1	0	1
hodoscope 3	1	0	1	0	1	0	1	1	1
hodoscope 1	0	1	0	1	0	1	1	1	1

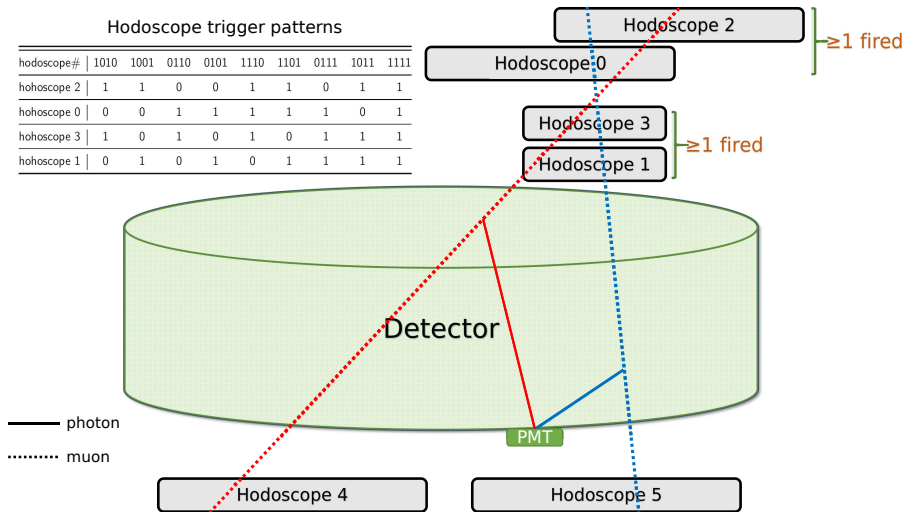


图: the trigger system



we expect the "calibration factors" to be independent of trigger pattern(the muon track), a pure linear efficiency factor.



short summary:

- ▶ a group of fitted effective efficiency of PMTs
- ▶ uncertainty estimation of these "calibration factors"
- ▶ the water attenuation length (not sensitive enough)



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tuning parameters in the model by fitting

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the projection view of detector

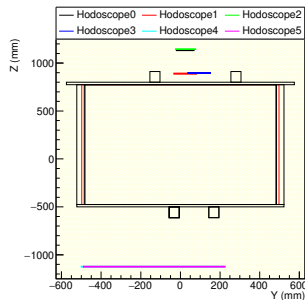
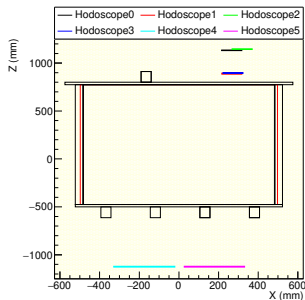
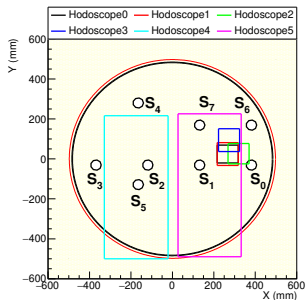


图: the geometry of 1Ton detector



the 9 combinations table:

表: mean value and uncertainty of calibration factors

hodoscope#	1010	1001	0110	0101	1110	1101	0111	1011	1111
hohoscope 2	1	1	0	0	1	1	0	1	1
hohoscope 0	0	0	1	1	1	1	1	0	1
hohoscope 3	1	0	1	0	1	0	1	1	1
hohoscope 1	0	1	0	1	0	1	1	1	1



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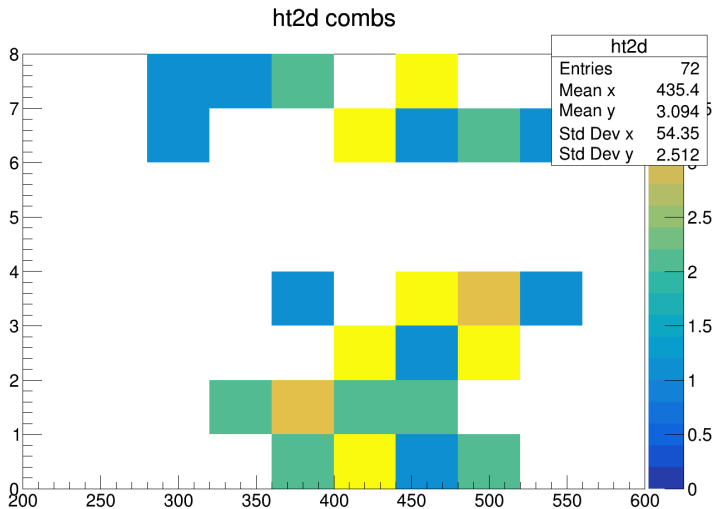
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photon acceptance model : water, acrylic, optical cookie, pmt solid angle
the model of attenuation and geometry



2d histogram, update this





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Thank You

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