

# First results of GLOBUS-M2 ASCOT5 modelling

Bakharev N.N.

27.05.2019

# Motivation

---

Tests and benchmarks for future use in

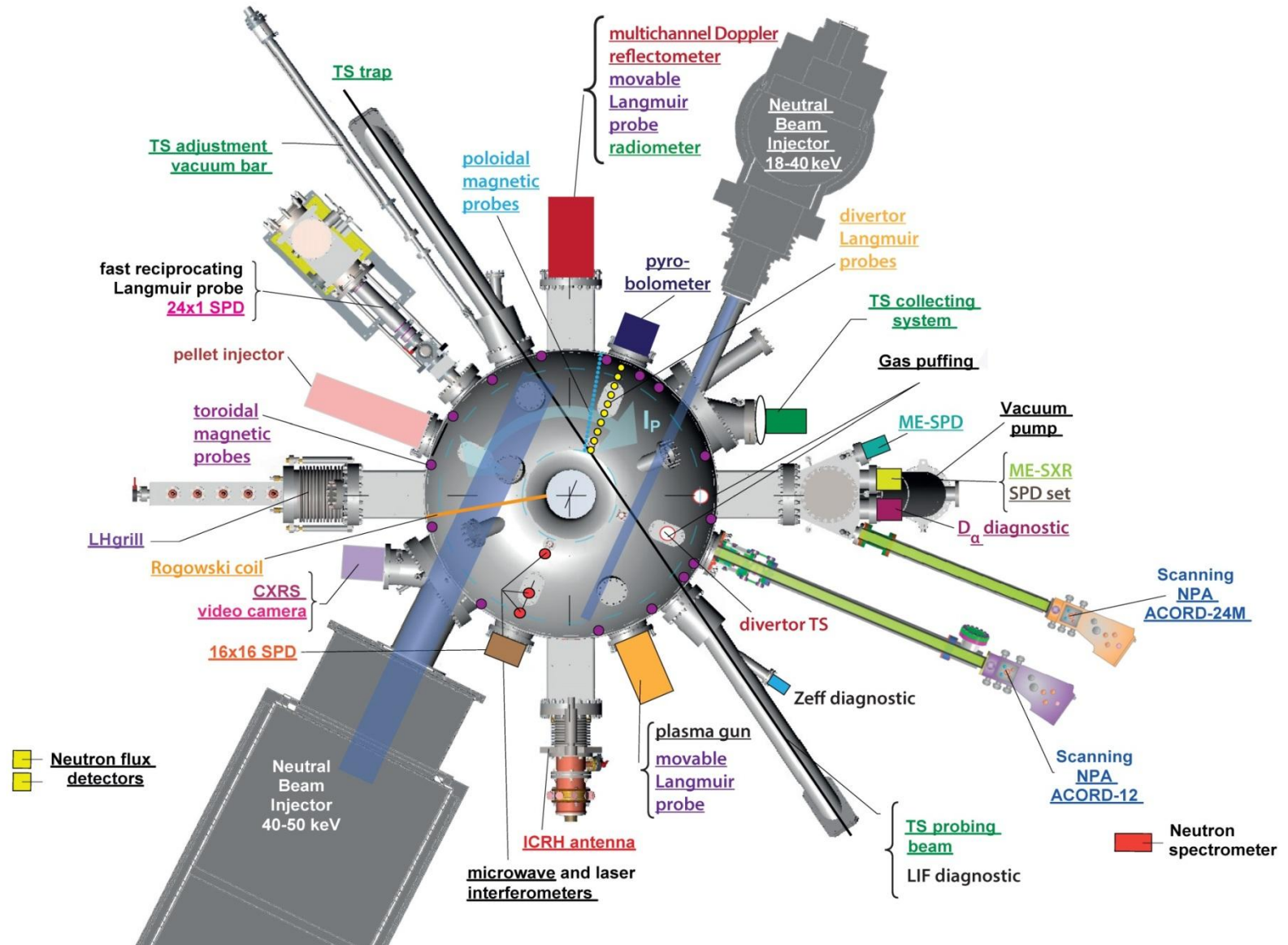
- Modeling of FI distribution in the NBI-heated experiments -calc. of  $P_{\text{abs}}$ , losses, neutron rate, NPA signals, CD etc. (CX required)
- Modeling of the TAE-induced losses/redistribution. (MHD required)

# Globus-M2

---

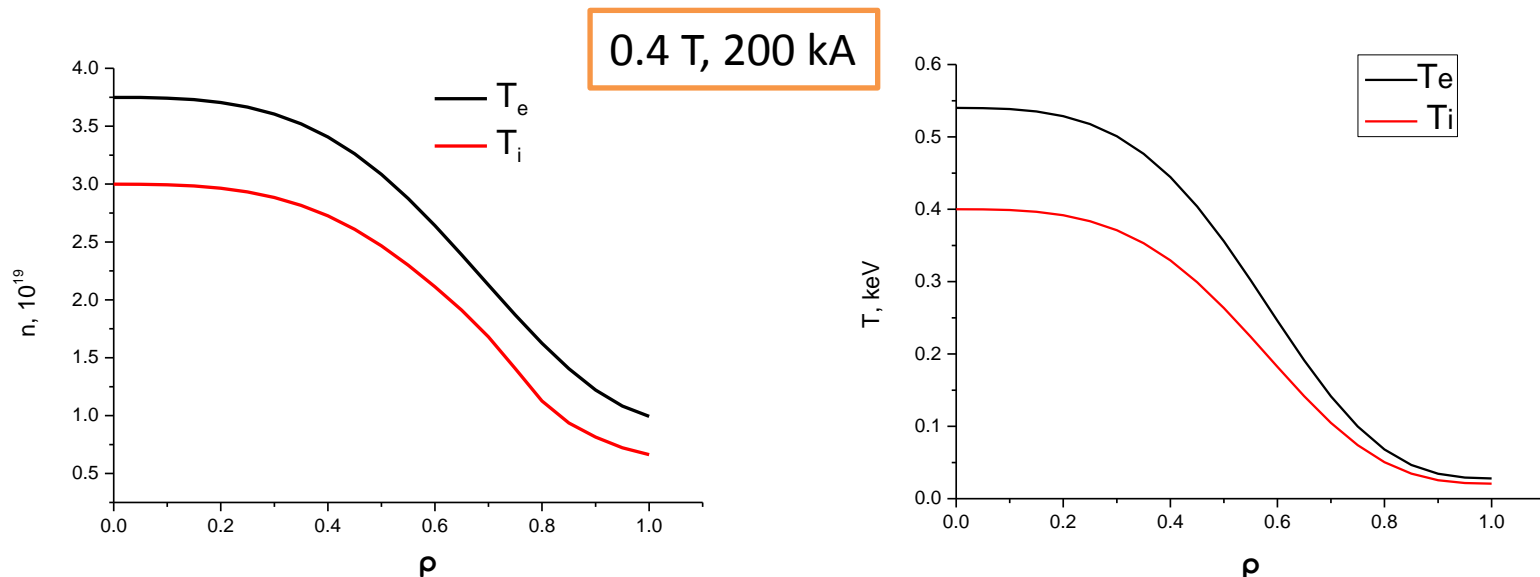
- New machine (first campaign - 2019)
- $R [\text{cm}]/a [\text{cm}] = 36/24 = 1.5$
- $B_T = 1\text{T}$ ,  $I_p = 500 \text{ kA}$
- Extreme  $P_{\text{heat}}/V = 6 \text{ MW/m}^3$
- Diverse diagnostics, heating and CD systems, including  $2 \times 1 \text{ MW NBI}$ ,  $\text{ICRH}$ , LHCD, plasma gun

# Globus-M2

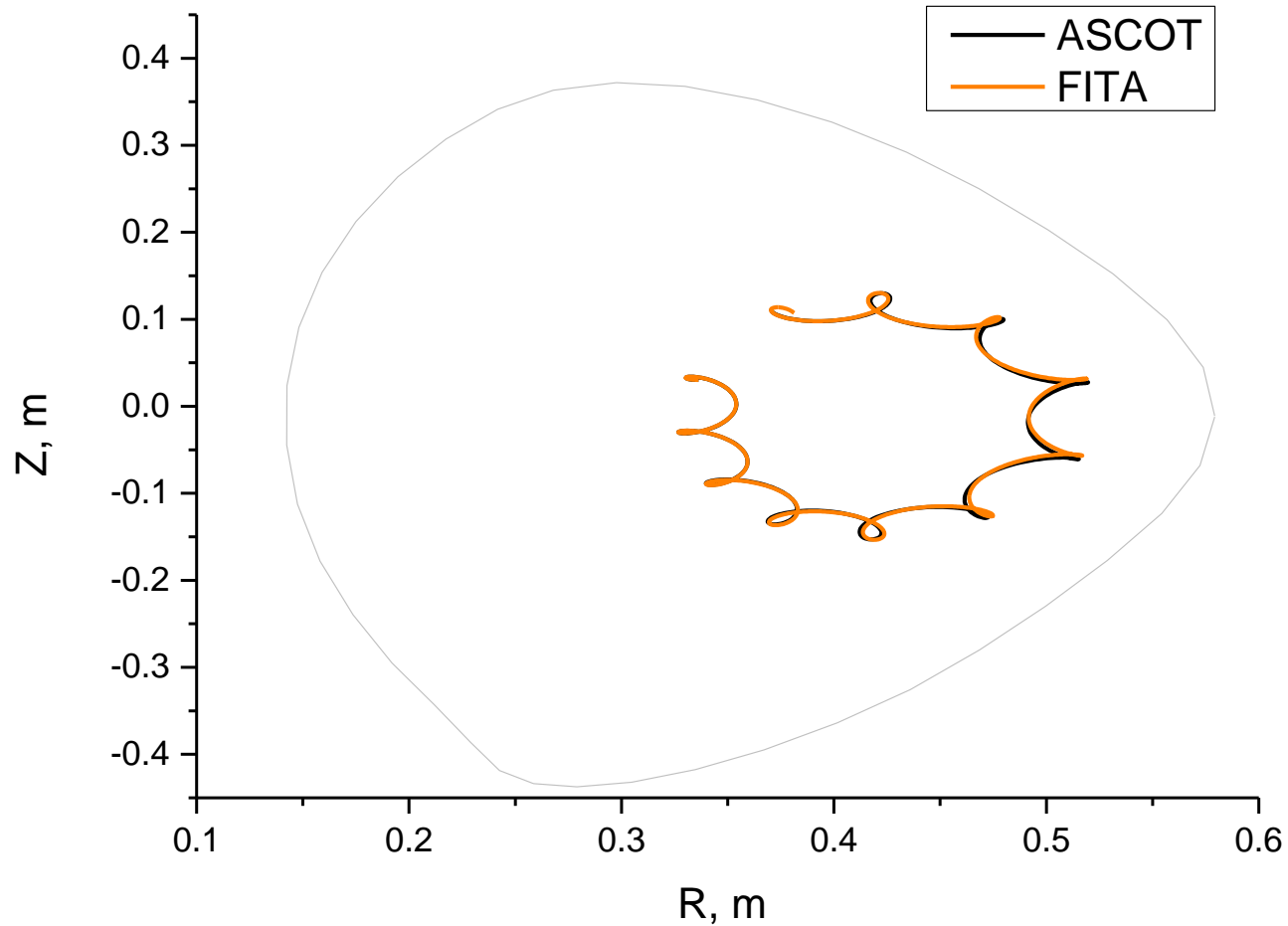


# Modelling overview

- 18 keV H (#32994-98) and D (#32952-57) NBI is considered.
- MHD-stable, neutron and NPA measurements exist (CX needed).
- Comparison with NUBEAM and homegrown GO simulations – fast in tracking algorithm(FITA).



# Orbit modelling



✓ **Sanity check of Magnetic eq. and marker inputs**

# Comparison with NUBEAM

---

## NUBEAM

- uses GC approx. + Finite Larmor radius adjustment
- Slightly different options

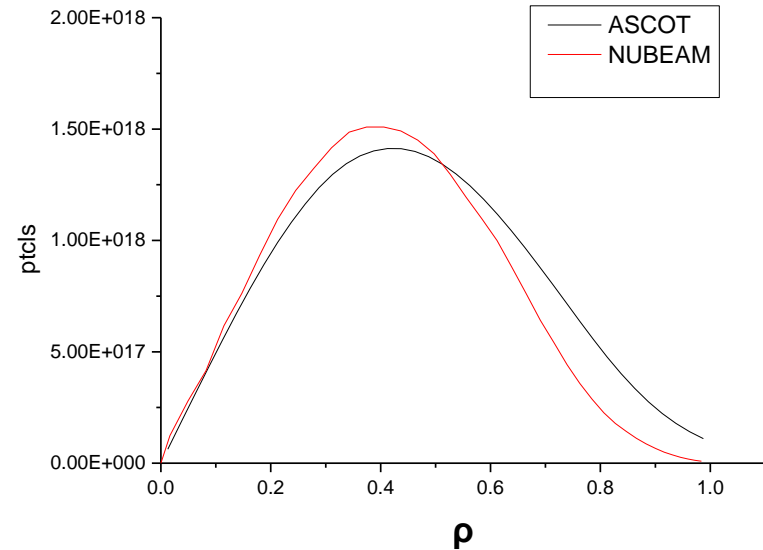
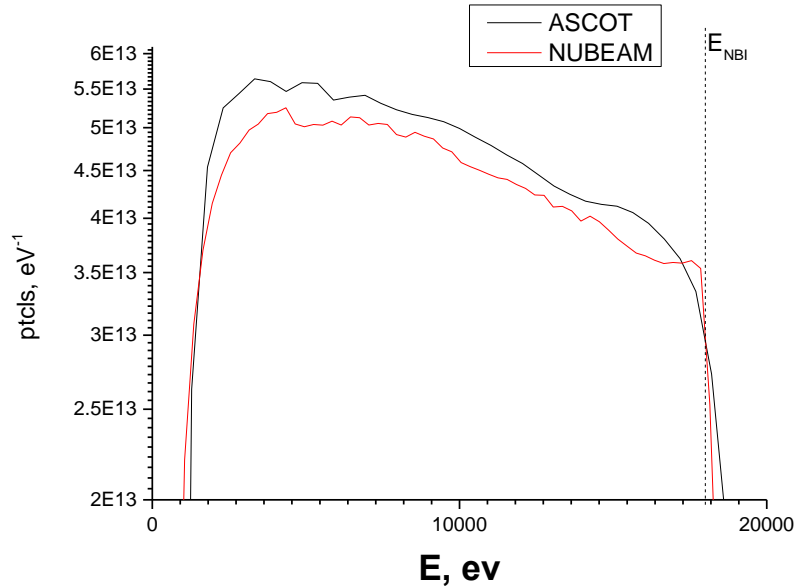
## NUBEAM markers -> ASCOT

Problems:

- Represented in GC approximation
- Not completely compatible with ASCOT:  
e. g. first order GCTRANS on leads to  $\mu < 0$  error  
and other errors (~3% of markers)
- NUBEAM uses irregular grid and increased central statistics.
- Not very precise  $\sigma_{CX}$  for such low energies.

# Comparison with NUBEAM

We will consider monoenergetic 18 keV beam



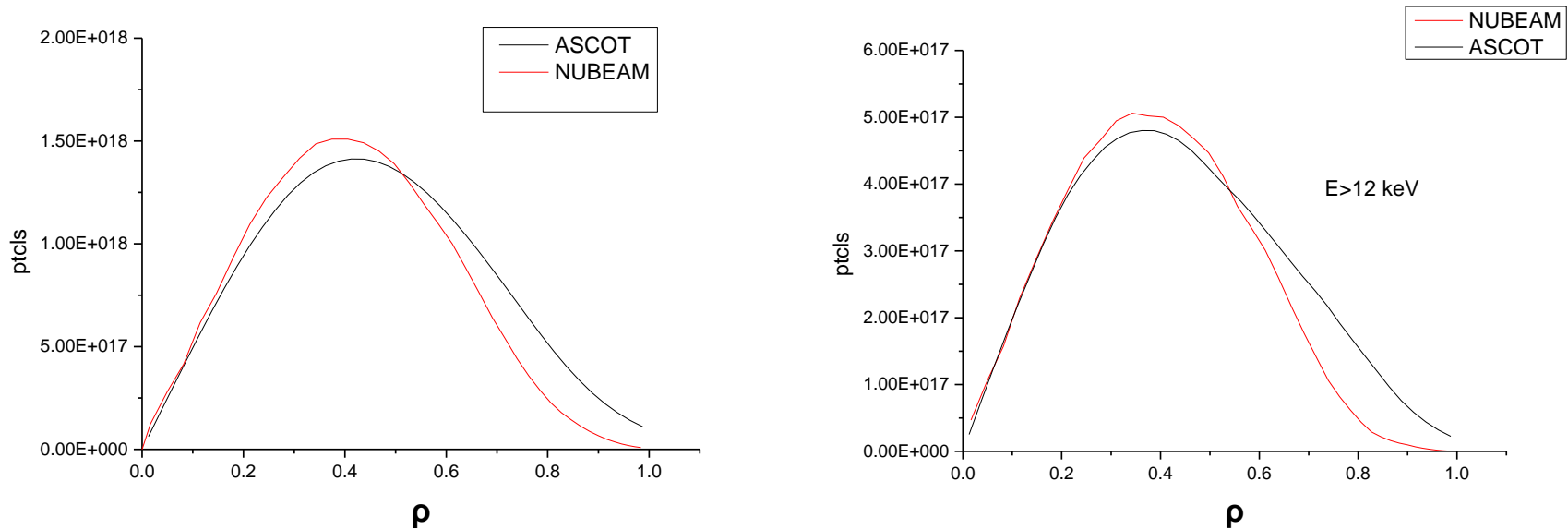
- ✓ Almost similar slowing-down.
- ✓ Density, predicted by NUBEAM is shifted inwards.

? Density, predicted by NUBEAM is lower.

? Different FI acceleration near  $E_{\text{NBI}}$

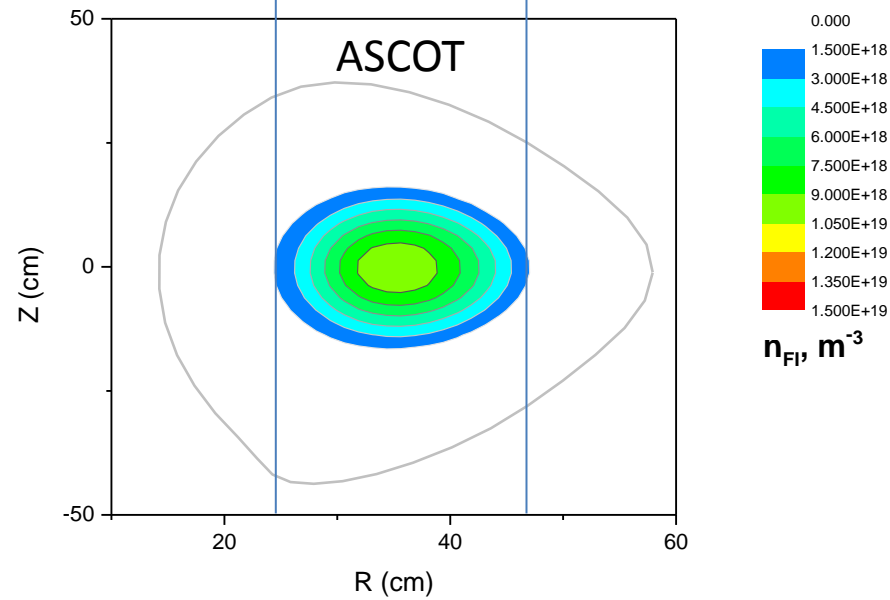
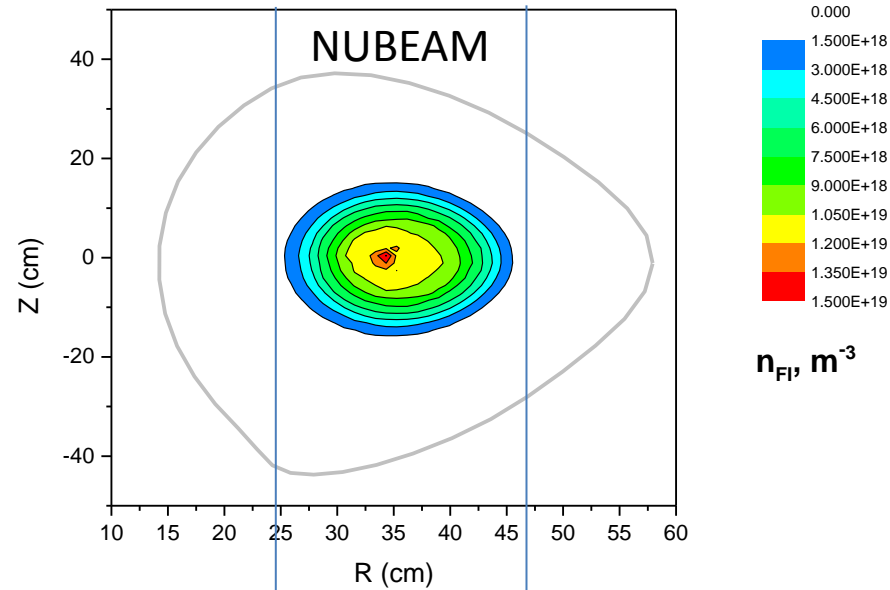


# Comparison with NUBEAM



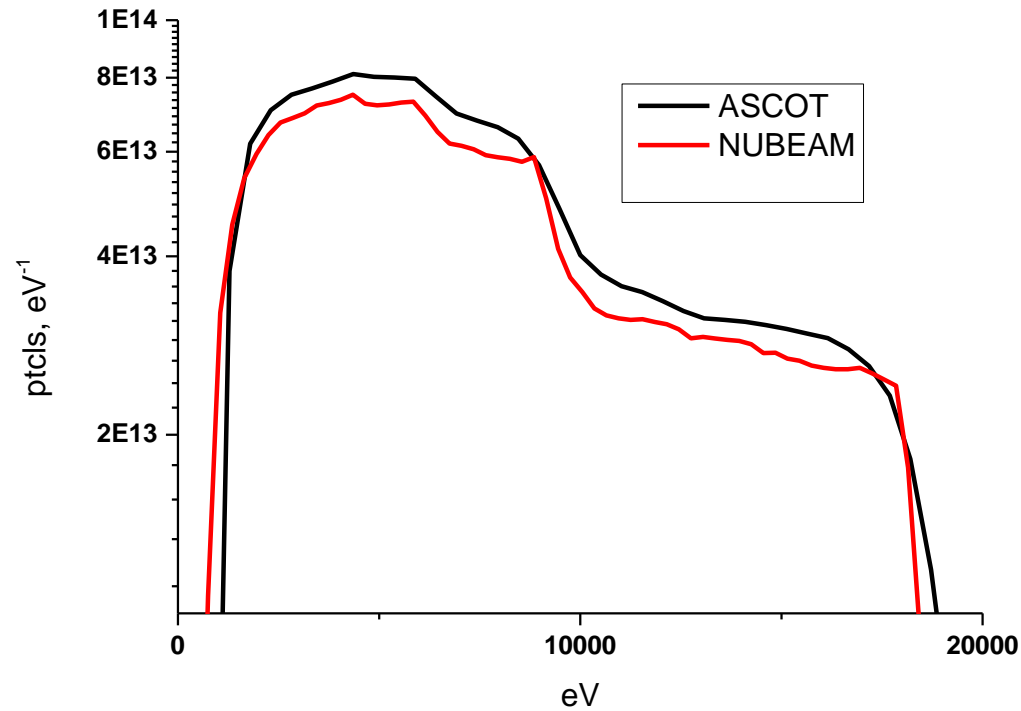
The difference near LCFS is even higher  
-> GC approx. plays the main role

# Comparison with NUBEAM distributions



# 3-component beam

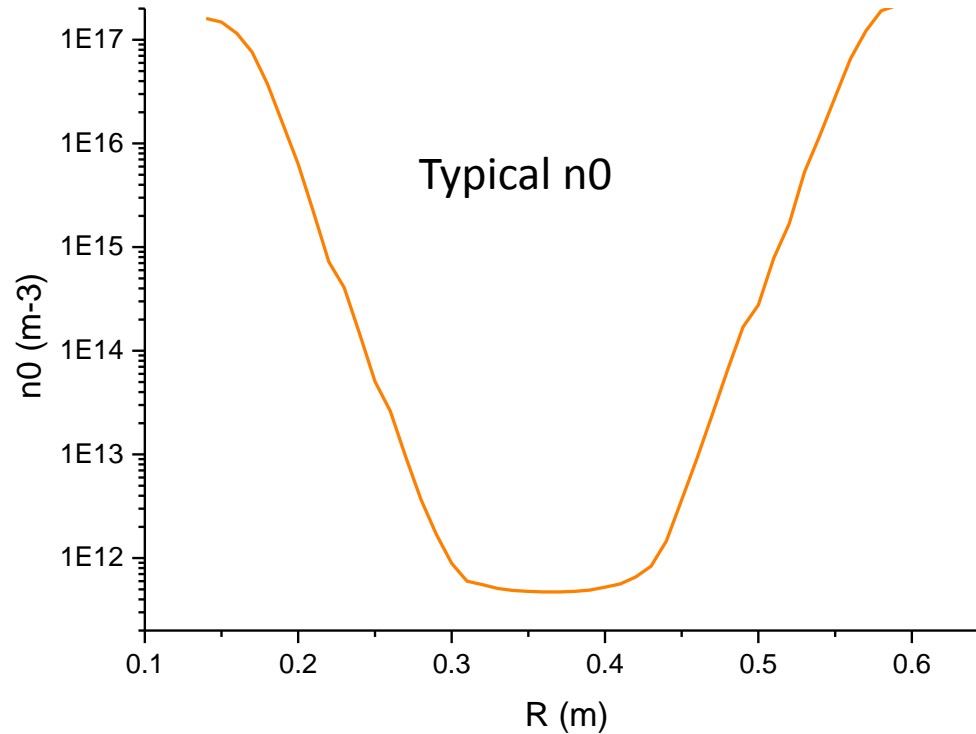
---



- Same features.

# CX losses

---



- CX and increase in  $E_{\text{beam}}$  will result in even higher differences with NUBEAM which fails to reproduce the experiments.

# FI losses

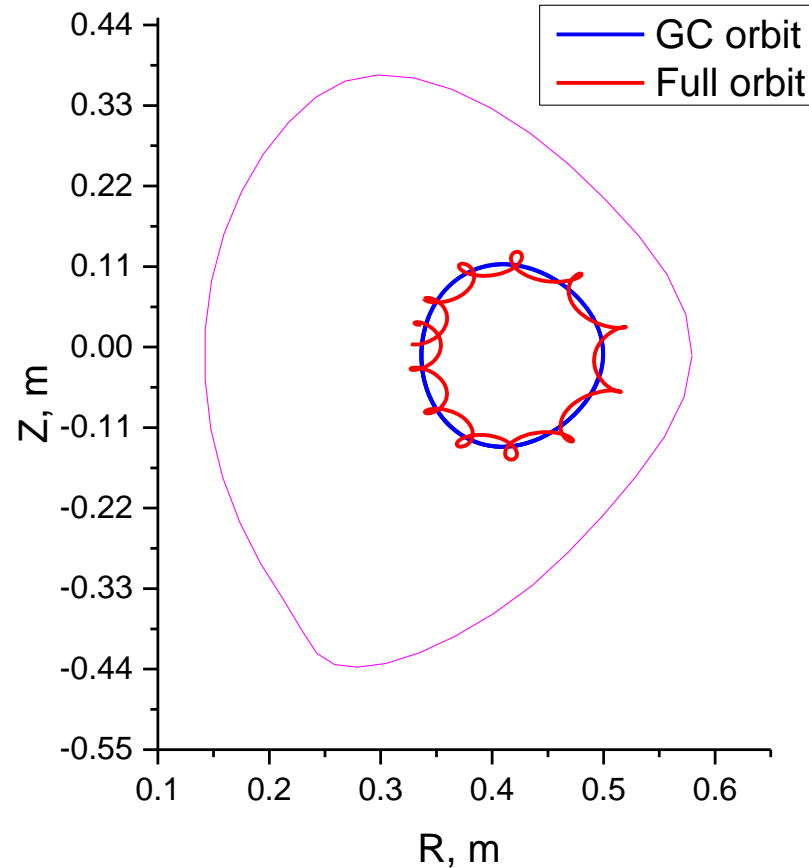
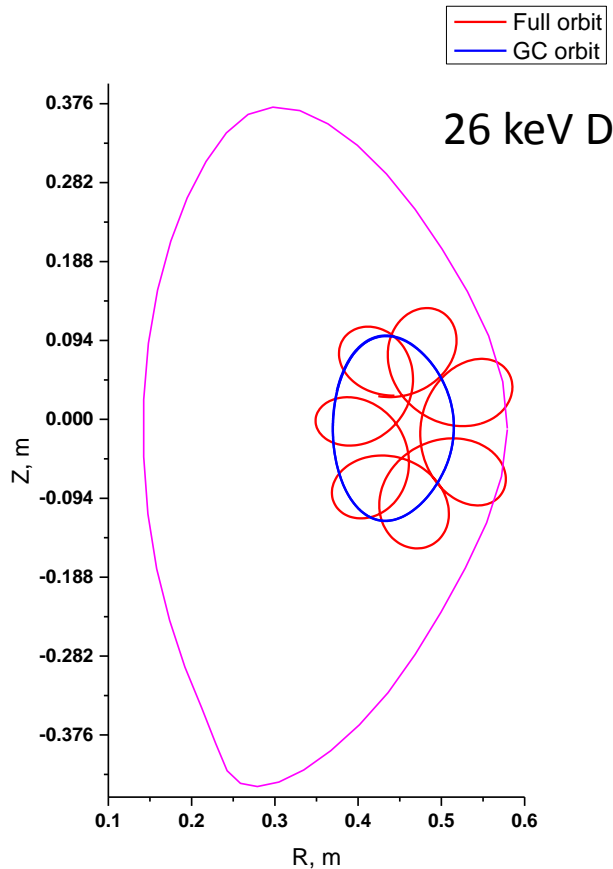
---

Losses	ASCOT (NO n0)	FITA	NUBEAM (NO n0)
Shine-through, %	-	8	6
wall	5% of markers	-	6% of P
CX, %	-	-	
FO( $\rho > 1$ ), %	13(NUBEAM markers) 12 (FITA markers)	16	15

**FO losses in ASCOT are a bit lower than in other codes**

# GC approximation in ASCOT

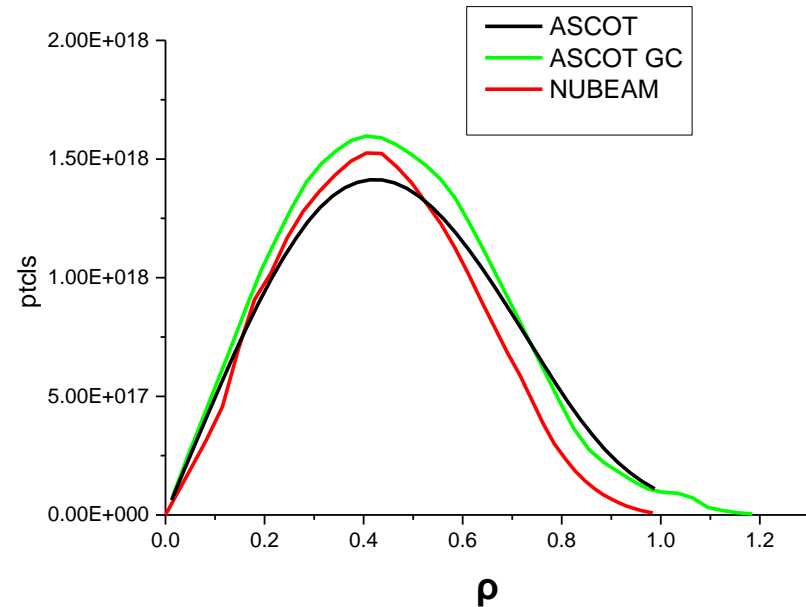
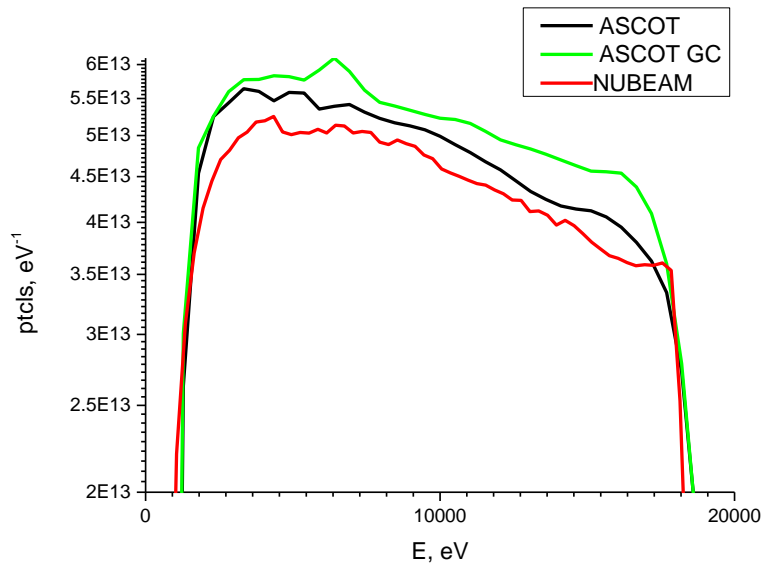
Good opportunity to compare GO and GC, implemented in one code



Not all orbits are easy to describe with GC

# GC approximation in ASCOT

Good opportunity to compare GO and GC, implemented in one code



- ✓ Higher FI density with GC approximation.
- ✓ NUBEAM is somewhere between ASCOT's GO and GC.

# SUMMARY

---

- **Comparison with NUBEAM and FITA was performed.**
- **Modelling shows reasonable results, however some peculiarities are unclear**
  - > **“smoke” test of input scripts is needed.**
- **After that we are ready for the real-world simulations:**
  - FI distribution (CX is needed)
  - TAE or other MHD(MHD is needed)
- **Consultations are needed.**

ASCOT user-experience:

Great design – transparent (almost a “white-box”) and flexible. :3

No backward compatibility. (° ° □ °) (—)

Thanks to the ASCOT team and especially Konsta for his time





Thank you for your attention!