# Precipitation sequence in Niobium-alloyed ferritic stainless steel <sup>1</sup>

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<sup>&</sup>lt;sup>1</sup>Fujita, Nobuhiro, H. K. D. H. Bhadeshia, and Masao Kikuchi. "Precipitation sequence in niobium-alloyed ferritic stainless steel." Modelling and Simulation in Materials Science and Engineering 12.2 (2004): 273.

### Niobium in ferritic stainless steels

- ullet Automobile industry becoming energy-efficient o increase in temperature of exhaust gas
- Material in exhaust:
  - Better high temperature strength
  - Resistance to thermal fatigue
- Solution: Nb based ferritic stainless steel
- Disadvantages: Precipitates at high temperatures and long time scales → Decrease in performance.

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## Experiments

- Nominal composition: 19Cr-0.8Nb wt.%
- Procedure followed:
  - Vacuum melting  $\rightarrow$  heating at 1250°C for 30 minutes in an Ar atmosphere  $\rightarrow$  hot-rolled and normalized to 900°C  $\rightarrow$  Annealing at 1000°C for 10 minutes  $\rightarrow$  water quenched  $\rightarrow$  Machining  $\rightarrow$  isothermal heat treatments at 950°C and 1000°C for 500 hours.
- Characterization done:
  - XRD-precipitates formed were noted.
  - TEM and EDS with carbon extraction replicas
     → microstructures and particle sizes.

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## Experimental results

#### Phase transformation

$$\alpha \rightarrow \alpha + Nb(C,N) + Fe_2Nb + Fe_3Nb_3C \rightarrow \alpha + NbN + Fe_3Nb_3C.$$

- Equilibrium has Fe<sub>3</sub>Nb<sub>3</sub>C
- NbC and Laves phase dissolves.
- ullet Precipitates o more spherical than needle-like.

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Table: Presence of precipitates in samples at different experimental conditions: VW, W,S,VS stand for very weak, weak, strong and very strong X-ray intensities

Aging conditions		Precipitates detected		
Temperature ( $^{\circ}$ C)	Time (h)	Nb(C,N)	$Fe_3Nb_3C$	$Fe_2Nb$
As annealed		S	VS	W
at $1000^{\circ}\text{C}$				
	1	VS	VS	VS
	8	S	VS	W
950	20	S	VS	W
	50	S	VS	VW
	100	VW	VS	-
1000	20	W	VS	-

#### Models used: Nucleation

## Classical nucleation theory

$$I = \left(1 - rac{V^{eta}}{V^{lphaeta}}
ight) N_0 rac{kT}{h} exp \left(-rac{G^* + Q^*}{RT}
ight)$$

$$G^* = \frac{16\pi\sigma^3}{3\Delta G_V^2}$$

- ullet  $V^{eta}$  and  $V^{lphaeta}$  are instantaneous and equilibrium volume fractions of the precipitate.
- $N_0$  and  $Q^*$  are the number density of nucleating sites and activation energy respectively.
- ullet  $\sigma$  and  $\Delta G_V$  are interfacial energy and volume Gibbs energy change.

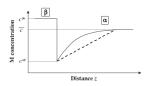
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# Growth: Binary

### Growth: governing equation

$$v(c^{\beta\alpha}-c^{\alpha\beta})=-D\frac{\partial c}{\partial z}$$

- D: diffusion coefficient.
- v: growth velocity
- $c^{etalpha}$ ,  $c^{lphaeta}$ : concentrations at equilibrium
- $\frac{\partial c}{\partial z}$ : concentration gradient.



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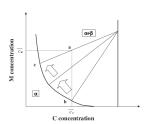


# Growth: MC precipitate

#### Growth: MC governing equation

$$v(c_X^{\beta\alpha} - c_X^{\alpha\beta}) = -D_X \nabla c_X$$

- C: interstitial atom  $\rightarrow D_C >> D_{\rm Nb}$ .
- Flux of C and Nb should almost match.
- Movement from b to c and a to c
  - b → initial C composition such that C gradient minimized and M gradient maximised.
  - As solute from matrix depleted, the average composition moves from a to c.
  - c → Equilibrium



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## Capillarity

#### Capillarity equation

$$c_{r,\mathrm{M}}^{lphaeta} = \left(1 + rac{\sigma}{kT} rac{v^{eta}}{r} rac{1 - c_{\mathrm{M}}^{lphaeta}}{c_{\mathrm{M}}^{etalpha} - c_{\mathrm{M}}^{lphaeta}}
ight) c_{\mathrm{M}}^{lphaeta}$$

- Phase boundaries → modified
- Reduces amount of small sized precipitates
- ullet Larger the precipitate, lower is the solute content at its interface ightarrow drives coarsening

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#### Calculations and results

- CALPHAD description and solubility products→ Volume Gibbs energy
- Fe-Nb-C system considered
- Parameters varied:  $N_0$  and  $\sigma$ .
- Close agreement with experimental data.
- Interfacial energy 15 times stronger impact than number of nucleation sites.

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## Parameters obtained

Table: Results from modeling

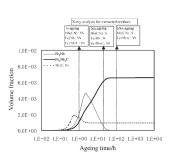
	AL	`
	Number density of sites: $N_0$ (m <sup>-3</sup>	)
NbN		$2 \times 10^{12}$
$Fe_2Nb$		$3 \times 10^{11}$
$Fe_3Nb_3C$		$3 \times 10^{12}$
	Interfacial energy: $\sigma(\ {\sf J}\ {\sf m}^{-2})$	
NbN		0.230
$Fe_2Nb$		0.280
$Fe_3Nb_3C$		0.330

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## Volume fraction

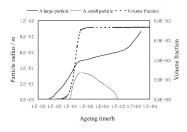
- Precipitation sequence verified
- Increases and saturates after some time
  - $\rightarrow$  Simultaneous dissolution of smaller particles + coarsening of the larger particles.



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## Precipitate sizes

- Mean radius increases and stabilizes after some time
- Larger particle: radii increases
   → Coarsening
- Smaller particle: radii decreases
   → Dissolution
- The number density of precipitates, decreasing during coarsening.



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## Summary

- Fe-Nb-C system considered
- Parameters varied:  $N_0$  and  $\sigma$ .
- Close agreement with experimental data.
- Interfacial energy 15 times stronger impact than number of nucleation sites.
- ullet Larger particle: radii increases o Coarsening
- Smaller particle: radii decreases → Dissolution

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