

Helio 6 ppm Helio.

$$v_{RMS} = \sqrt{\langle v^2 \rangle} \approx 1100 \text{ m/s.}$$

$$\langle v^2 \rangle = \frac{3RT}{M}$$

$$v_e = 11,4 \text{ km/s}$$

$$f(v) = \left( \left[ \frac{M}{2\pi RT} \right]^{3/2} 4\pi \right) \underbrace{v^2}_{\text{}} e^{-\frac{Mv^2}{2RT}} \quad \text{f.v.}$$

Google Class  
Excel.  $\uparrow$   $\text{exp}(\text{---})$



$p, n, T, V \rightarrow$  Estado.

$$\boxed{pV = nRT} \Rightarrow R = \frac{pV}{nT}$$

$$A, B, C, \dots, Z) \quad T = \frac{PV}{nR}$$

A:  $p_A, V_A, T_A \rightarrow \underline{V_A}$

transformaciones  $\rightarrow 1, 2, 3, \dots$

$p = \text{cte}$  isobárico

$T = \text{cte}$  isotermo

$V = \text{cte}$  isocora.

$n = \text{cte}$  isomáxico

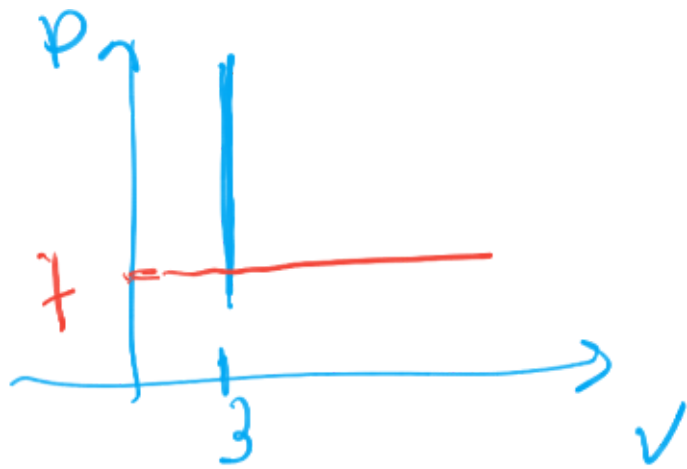
$pV = \text{cte} \Rightarrow pV = (nRT)_{\text{cte}}$

Ley de Boyle

$p(V) \Rightarrow p(V) = \frac{nRT}{V}$

Hiperbolas

Isocora  $V = \text{cte}$



Isobare  $p = \text{cte}$

$$\Delta V = V_B - V_A > 0$$



$$W = F \Delta h$$

$$\uparrow$$

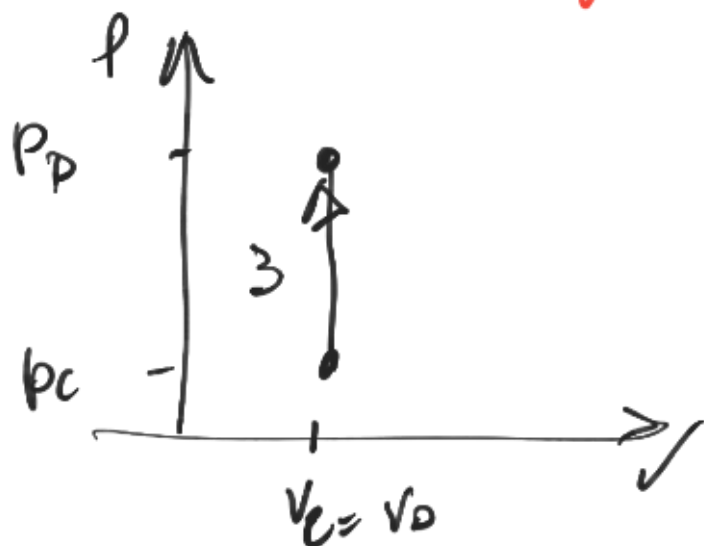
$$p \cdot A$$

$$W = p \cdot A \cdot \Delta h$$

$$\underbrace{(W)}_{>0} = p \cdot \underbrace{(\Delta V)}_{>0}$$

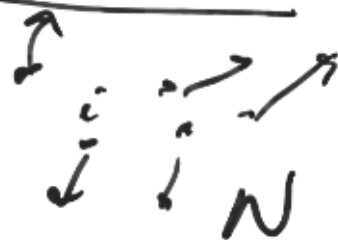


1.1 - 1.1 1.1



Energia

$$\langle E_K \rangle = \frac{1}{N} \sum_{i=1}^N E_{Ki}$$



$N \langle E_K \rangle = \cancel{N} \cancel{\frac{1}{N}} \sum_{i=1}^N E_{Ki}$  Energia total

$N \langle E_K \rangle \stackrel{\text{def}}{=} U$

— = 0 1 —

$$1 = \frac{2}{3} \frac{1}{k} \langle E_k \rangle$$

$$\langle E_k \rangle = \frac{3}{2} kT$$

$$\frac{3}{2} N k T = U$$

$$U = \frac{3}{2} N \frac{N_A}{N_A} k T$$

$$U = \frac{3}{2} n R T$$

$$n = cV$$

$$dU = \frac{3}{2} R d(nT)$$

$$dU = \frac{3}{2} R [dnT + n dT]$$

$$\text{Si } n = cV$$

$$\Delta U = \frac{3}{2} R n \Delta T$$

1

$$Q = C n \Delta T$$

$$C \stackrel{\text{def}}{=} \frac{Q}{n \Delta T}$$

Color especifica

Gas ideal  $V = \text{cte}$   
 $P = \text{cte}$

i)  $V = \text{cte}$   $W = 0$

$$Q = C n \Delta T$$

↓

$$Q = \Delta U = \frac{3}{2} n R \Delta T$$

$$C \cancel{n} \cancel{\Delta T} = \frac{3}{2} \cancel{n} R \cancel{\Delta T}$$

$$C_V = \frac{3}{2} R$$

Color esp.  
 Gas ideal  
 Macroscopic

$$\Delta V \neq 0 \quad p = \text{cte} \Rightarrow W \neq 0$$

$$W = p \Delta V$$

$$pV = nRT \Rightarrow V = \frac{nR}{p} T$$

$$\Delta V = \left( \frac{nR}{p} \right) \Delta T$$

$$W = \cancel{p} \cdot \frac{nR}{\cancel{p}} \Delta T$$

$$W = nR \Delta T$$

$$\Delta U = \frac{3}{2} nR \Delta T$$

$$Q = \Delta U + W$$

$$Q = \frac{3}{2} n R \Delta T + n R \Delta T$$

$$Q = \left( \frac{3}{2} + 1 \right) n R \Delta T$$

$$Q = \frac{5}{2} n R \Delta T$$

$$Q = C_p n \Delta T$$

$$\cancel{C_p n \Delta T} = \cancel{\frac{5}{2} n R \Delta T}$$

$$C_p = \frac{5}{2} R$$

Calor specific  
at constant  
pressure

En un tubo est. de un gas ideal



Quantum de temperatura.

$$V=cte \quad Q_v = \frac{3}{2} R n \Delta T$$

$$p=cte \quad Q_p = \frac{5}{2} R n \Delta T$$

$$C_p = \frac{5}{2} R \Rightarrow C_p = \left( \frac{3}{2} + 1 \right) R$$

$$= \underbrace{\frac{3}{2} R}_{C_v} + R$$

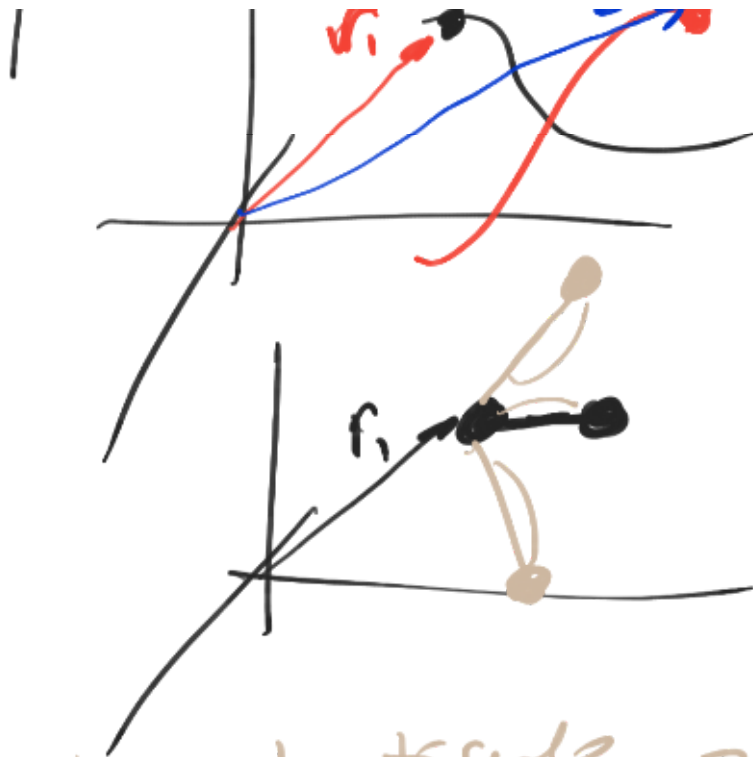
$$\Rightarrow \boxed{C_p = C_v + R} \quad \text{ley del  
para  
trabaja  
en gas  
ideal.}$$

He  
↑

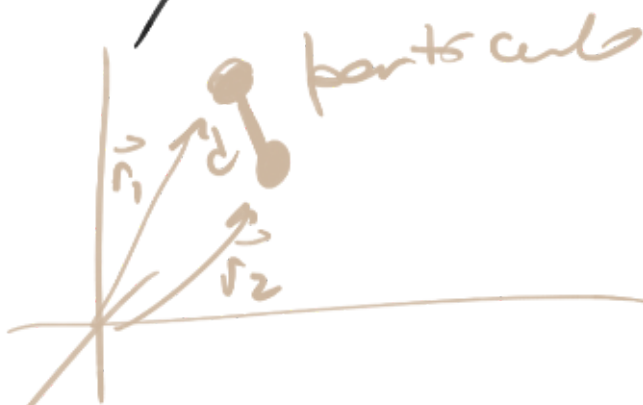
$$O = O$$

$$O = C = O$$





Grados  
de  
Libertad



partículas 3 Grados

Grados de libertad  
por part

$$cte: d = \sqrt{(x_1 - x_2)^2 + (y_1 - y_2)^2 + (z_1 - z_2)^2}$$

bicromica 5 Grados

mono  $z = 3$

$$C_v = \frac{3}{2} R$$

bi  $z = 5$

$$C_v = \frac{5}{2} R$$

tri  $z = 6$

$$C_v = 3R$$

tema i. Gradi de libertate

$$C_v = \frac{5}{2} R$$

$$C_p = C_v + R$$

$$\gamma \stackrel{\text{def}}{=} C_p / C_v$$

$$\text{mono: } \gamma_{\text{mono}} = \frac{5/2 R}{3/2 R} = \frac{5}{3}$$

$$\gamma_{\text{bi}} = \frac{7}{5}$$

$$\gamma_{\text{tri}} = \frac{4}{3}$$

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$$\text{Monatomic } C_v = \frac{3}{2} R \quad v = \text{etc}$$

$$\text{Biatomic } C_v = \frac{5}{2} R \quad v = \text{etc}$$

$$Q = C_v n \Delta T$$

$$Q_v^{\text{molar}} = \frac{3}{2} R n \Delta T$$

$$Q_v^{\text{bi}} = \frac{5}{2} R n \Delta T$$