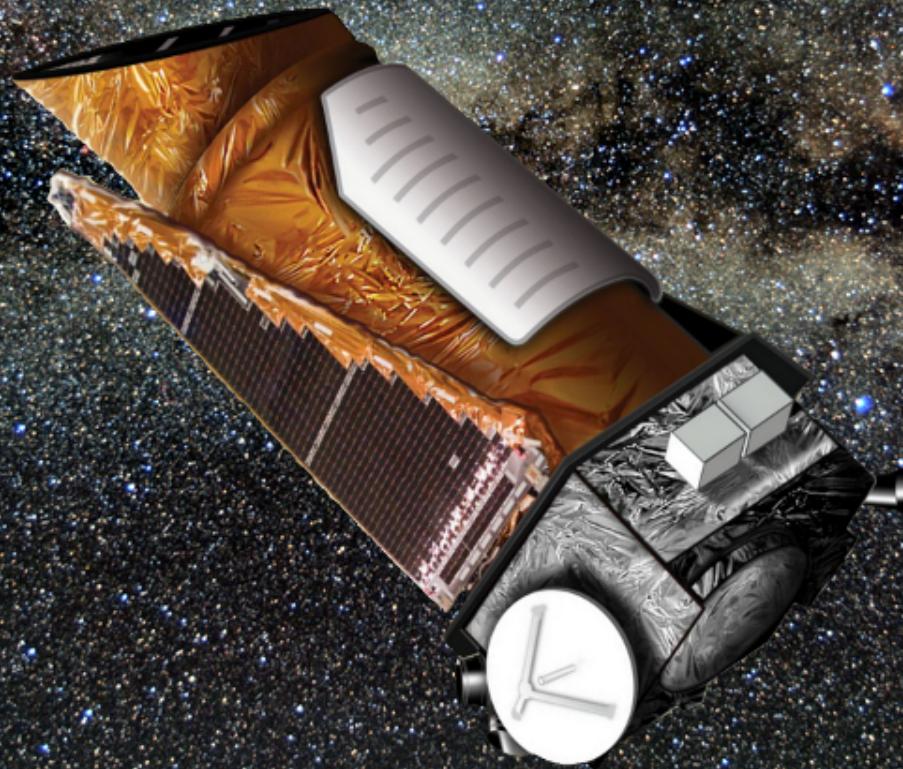


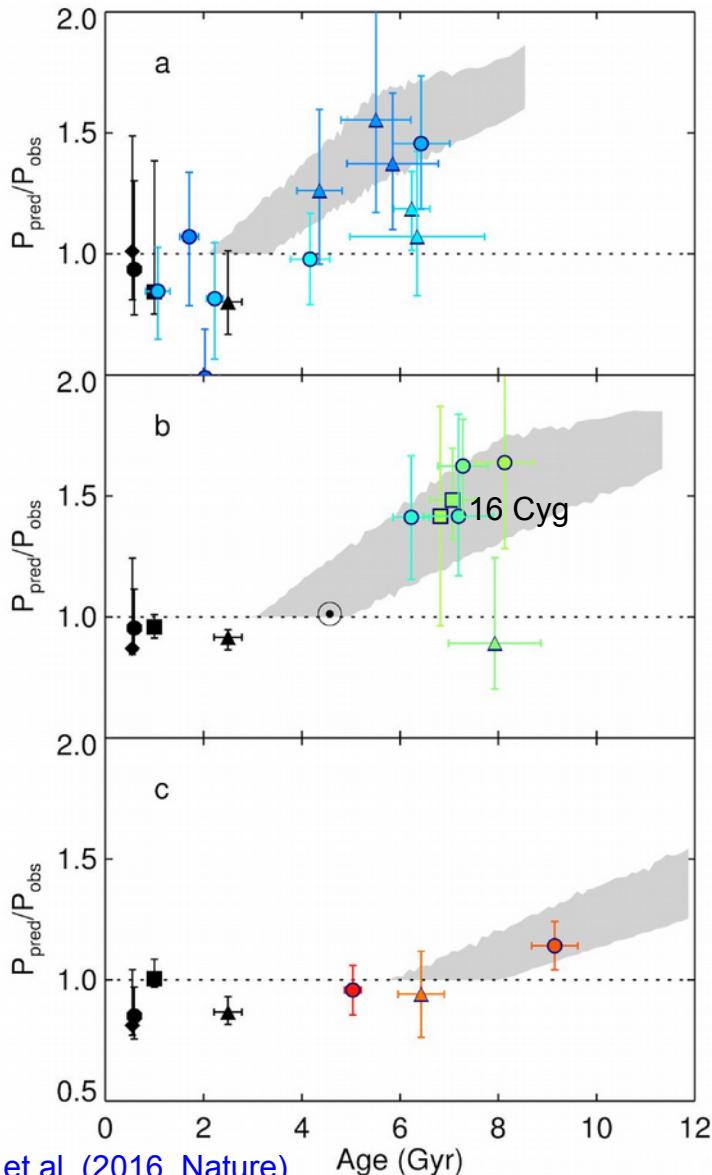
Stellar Evidence of a Solar Dynamo in Transition

Travis Metcalfe (SSI & NSO)



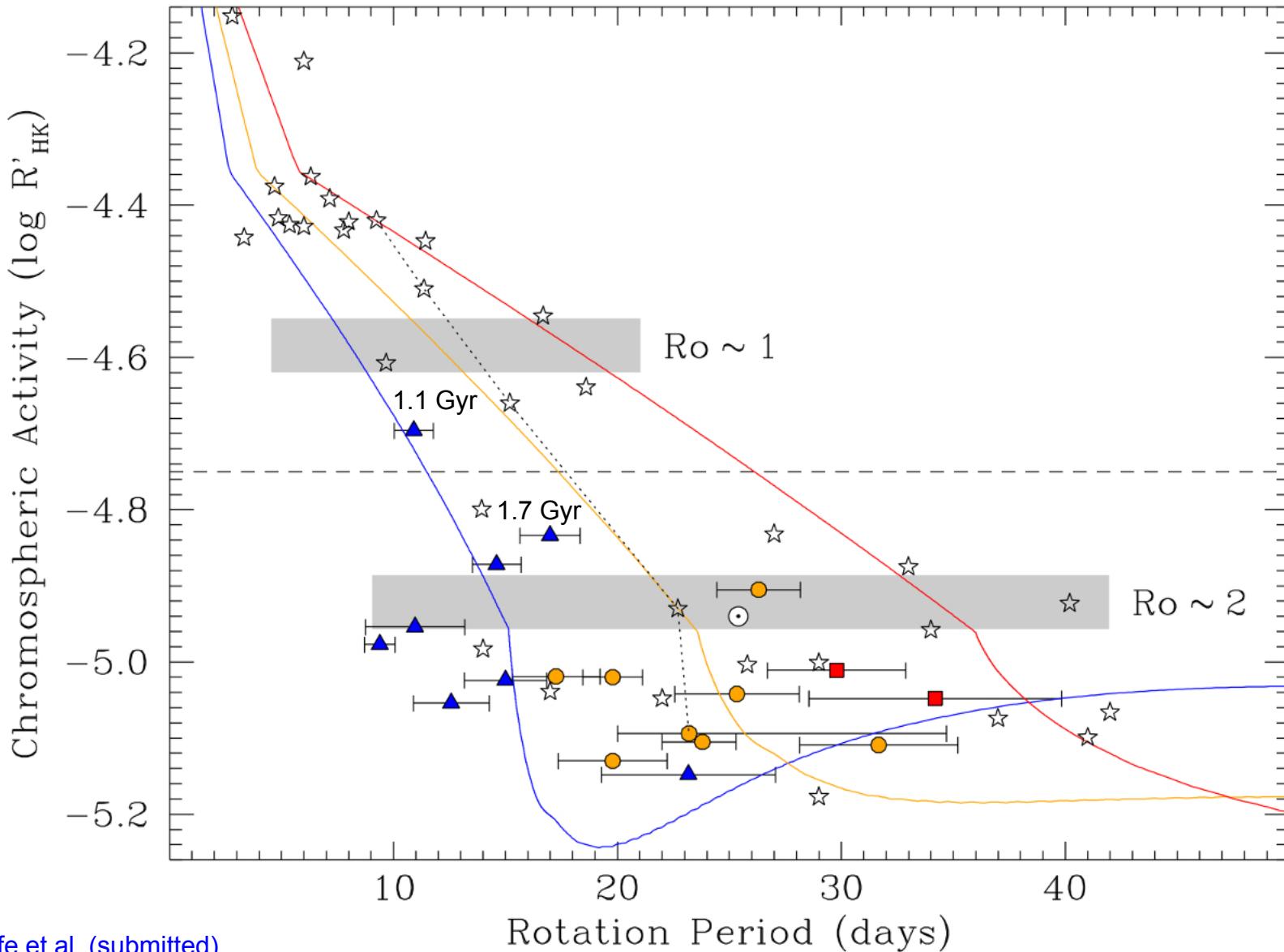
Collaborators: Ricky Egeland, Jennifer van Saders

Breaking magnetic braking

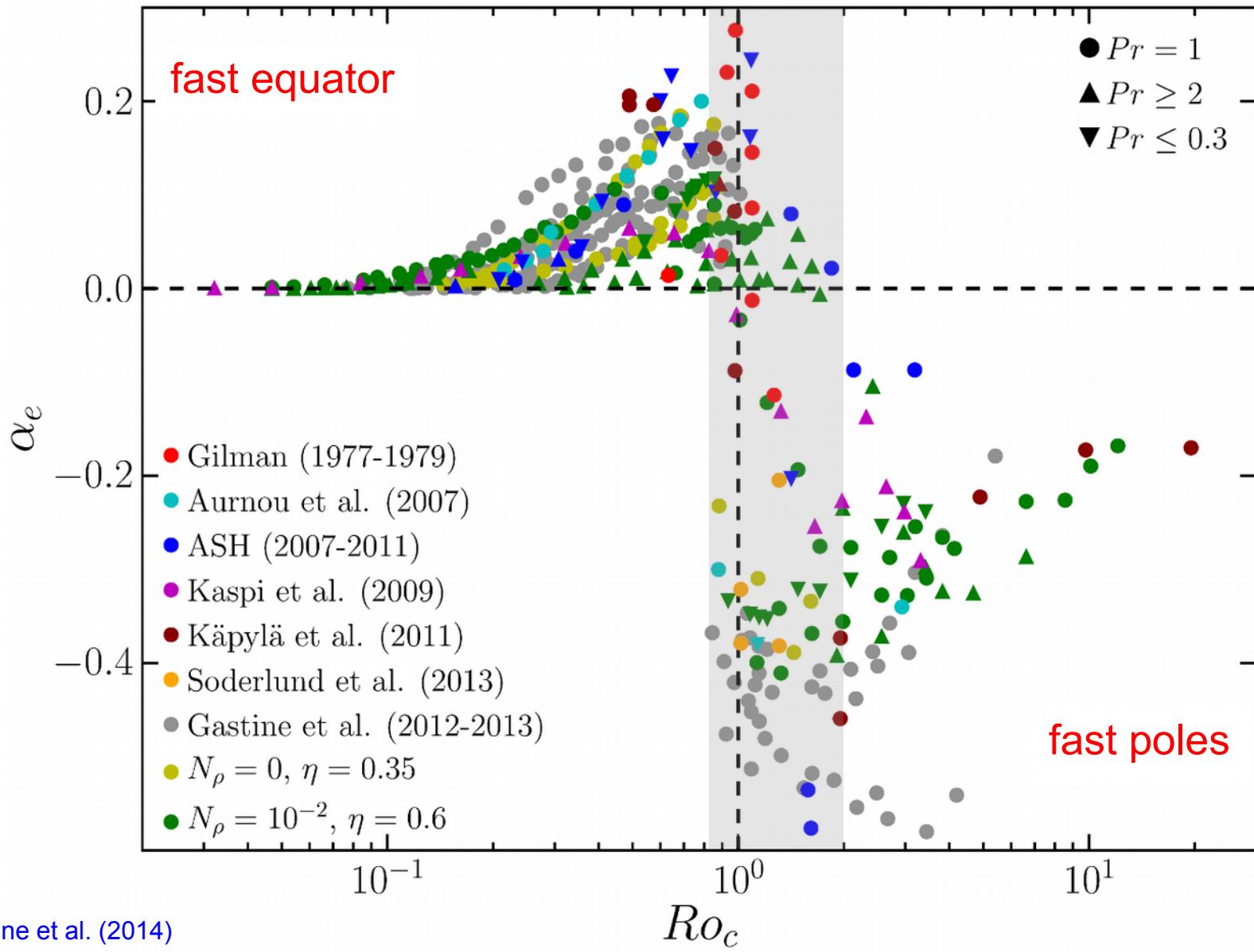


- Stars beyond middle-age rotate more quickly than gyrochronology predicts
- Onset is earlier in F-type, near solar age in G-type, later in K-type stars
- Effect seems to appear beyond a critical Rossby number ($\text{Ro} \sim 2$)

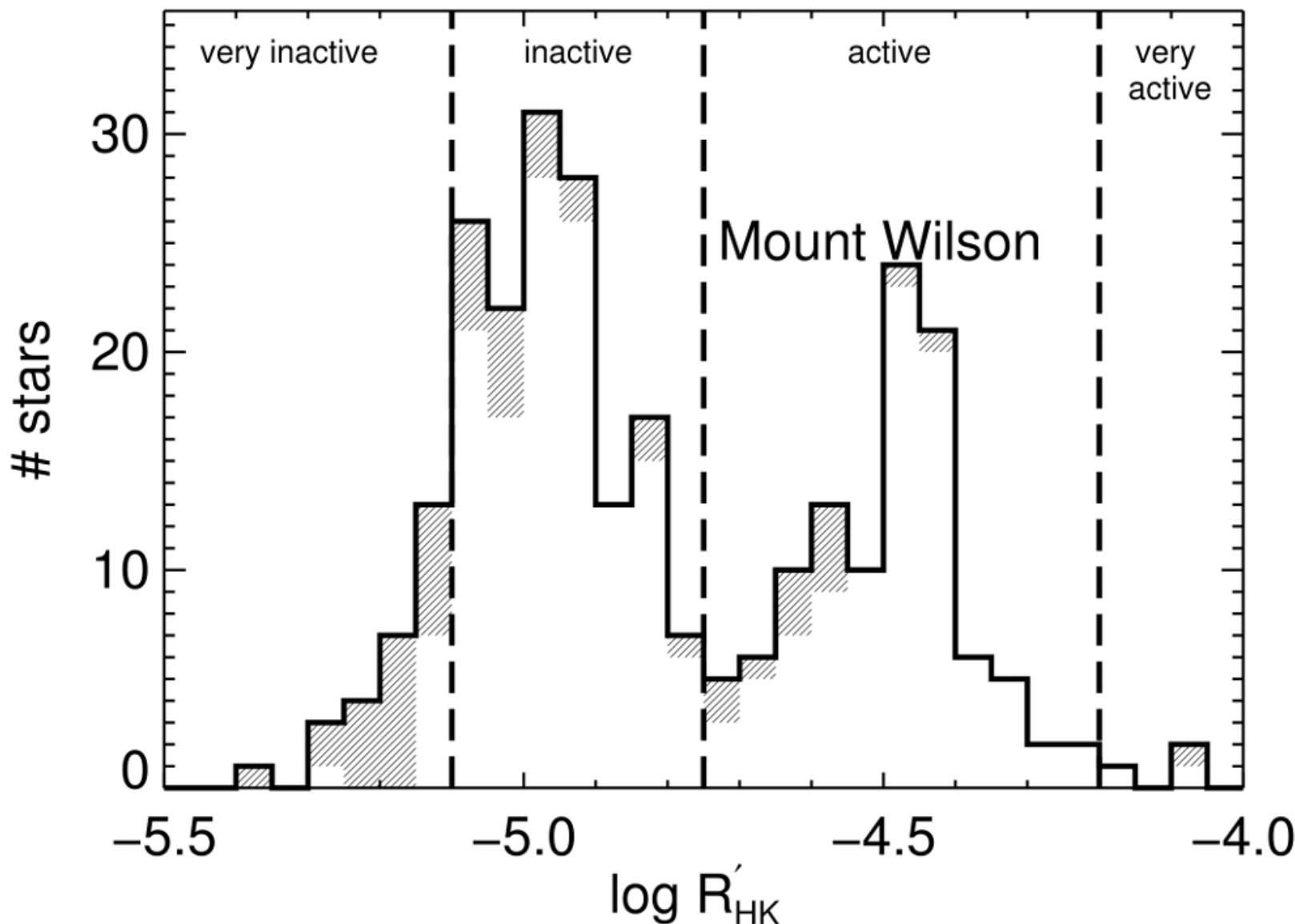
Rotation-activity relation



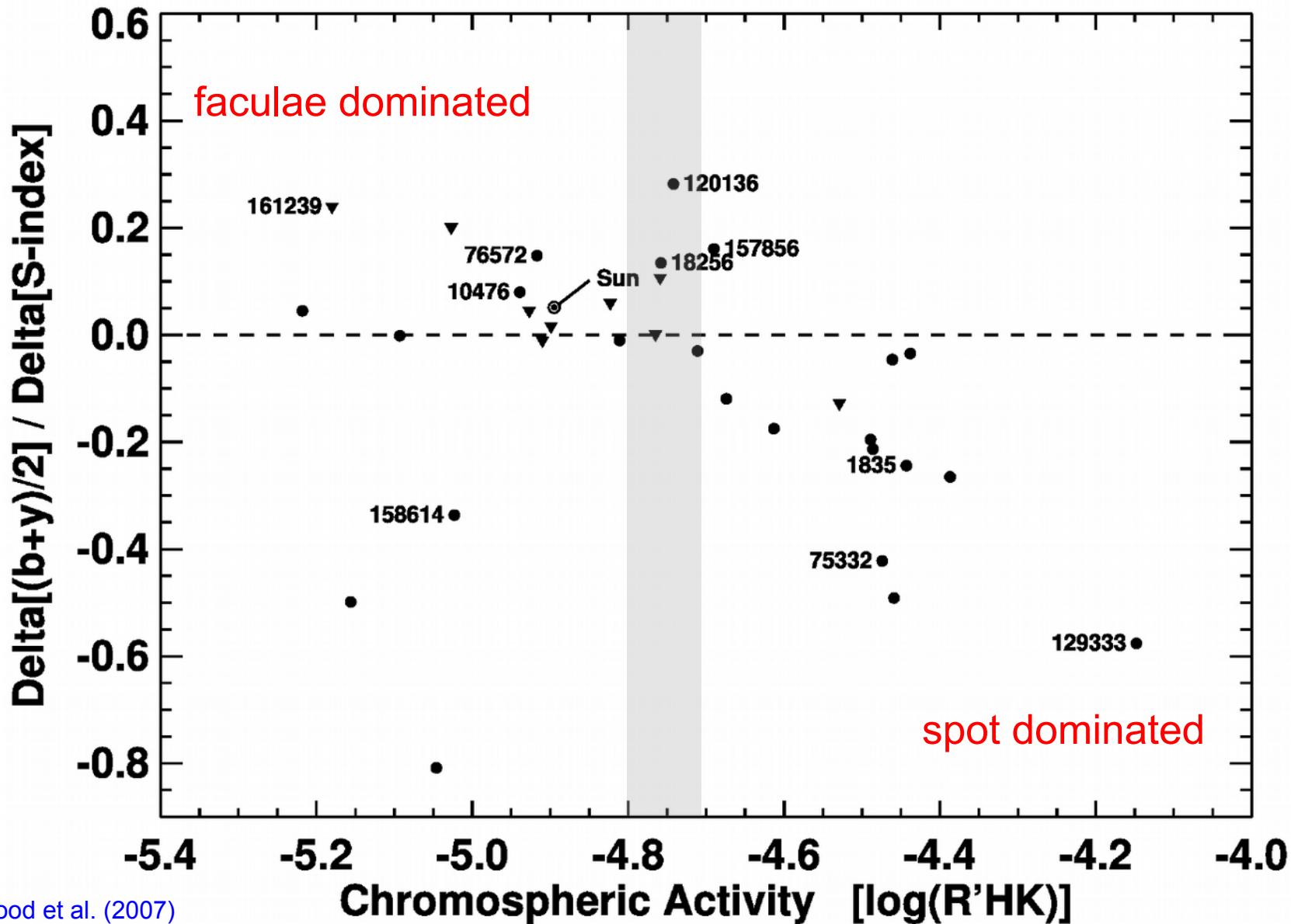
Differential rotation



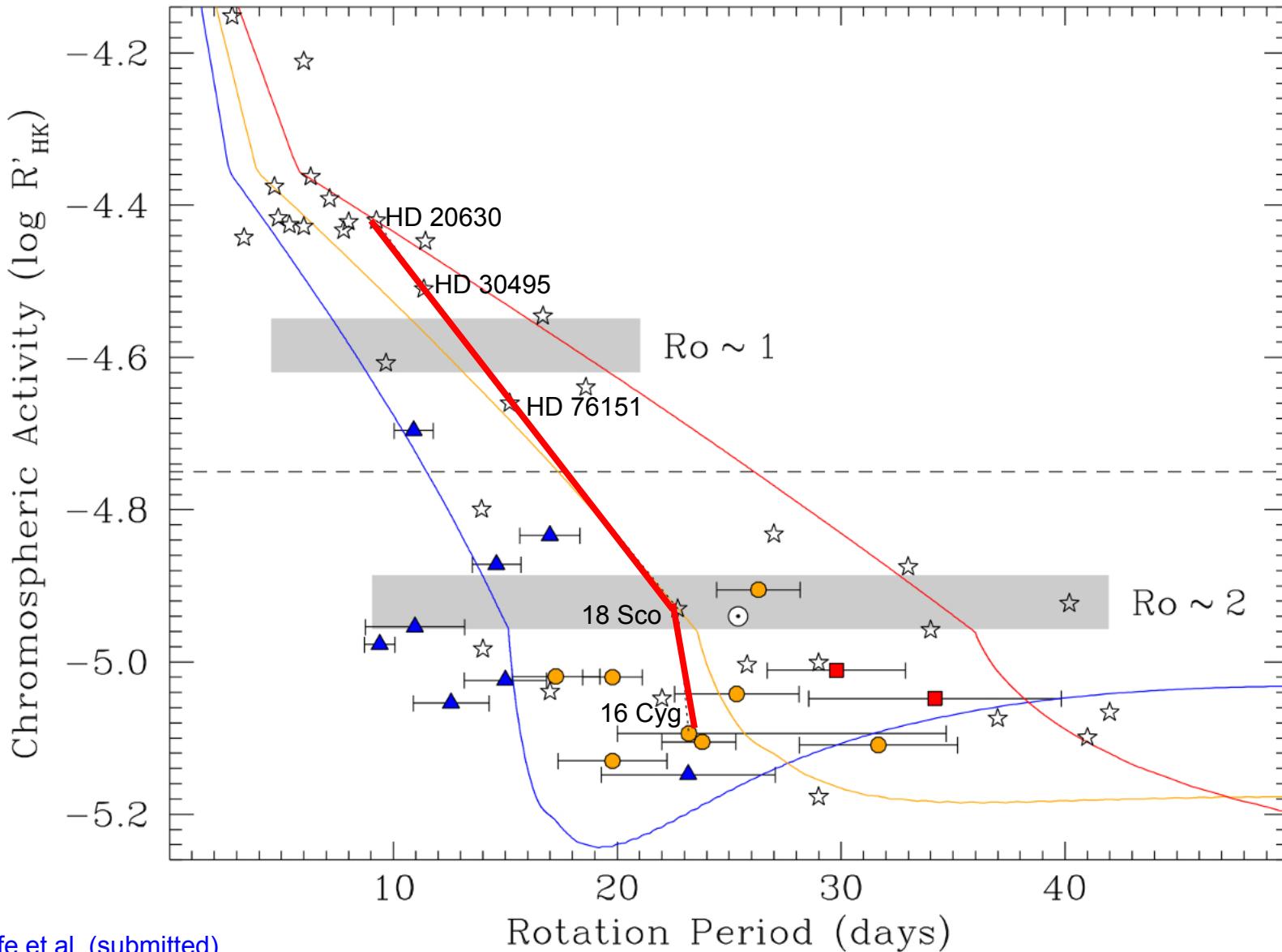
Vaughan-Preston gap



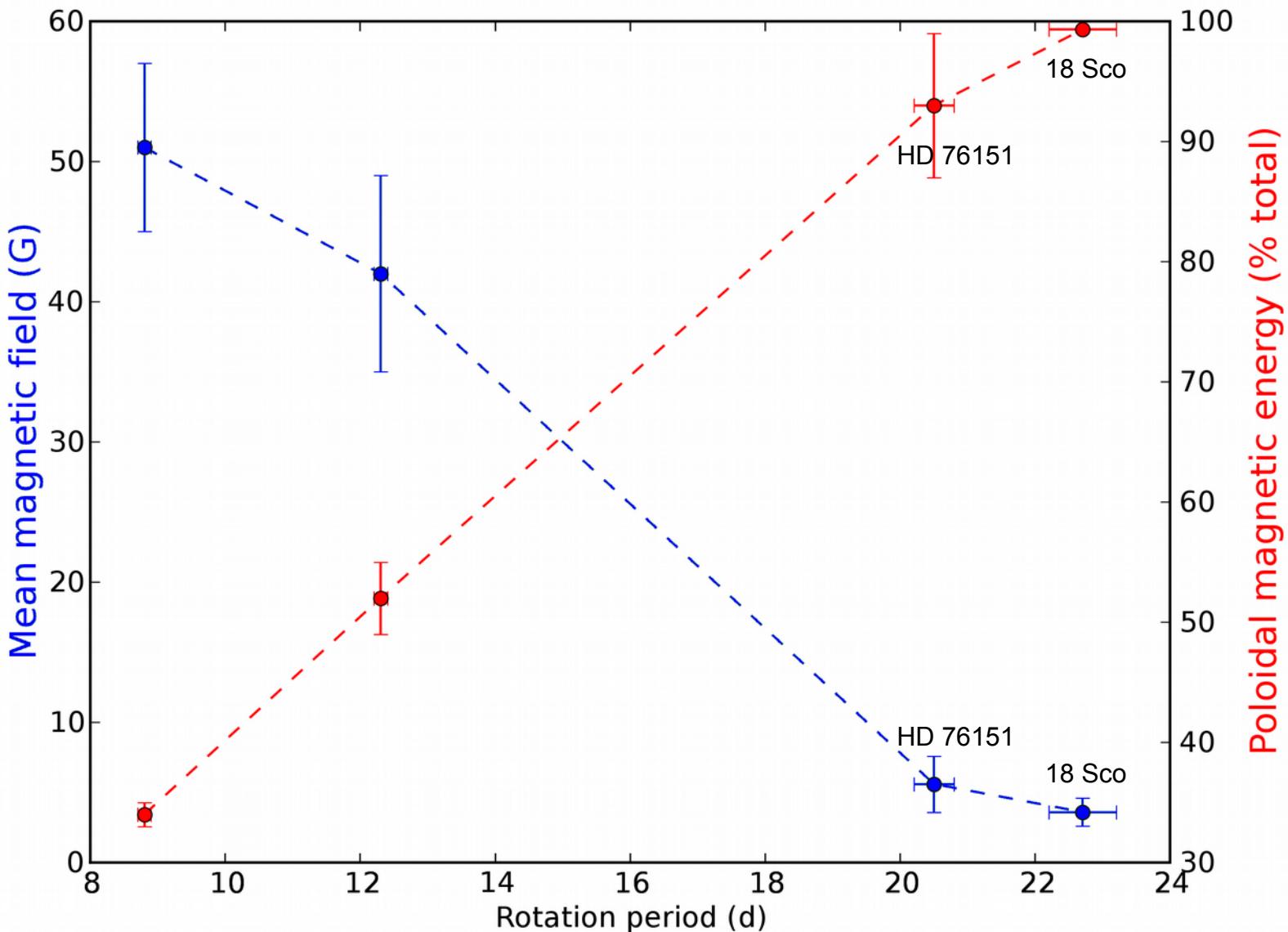
Spots and faculae



Solar analogs

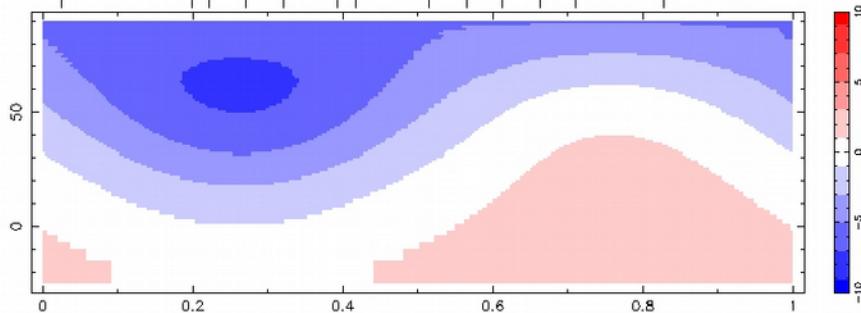


Zeeman Doppler imaging

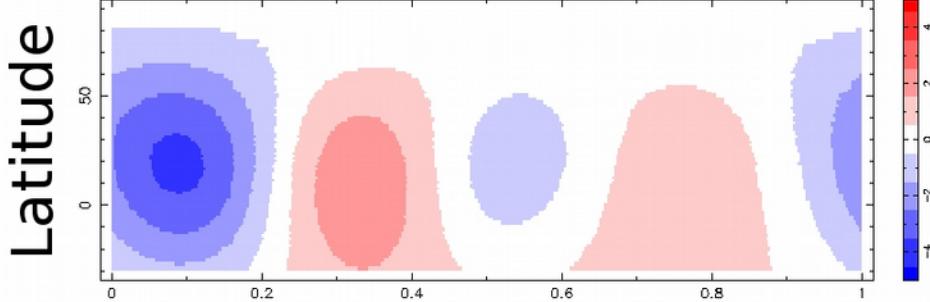


Zeeman Doppler imaging

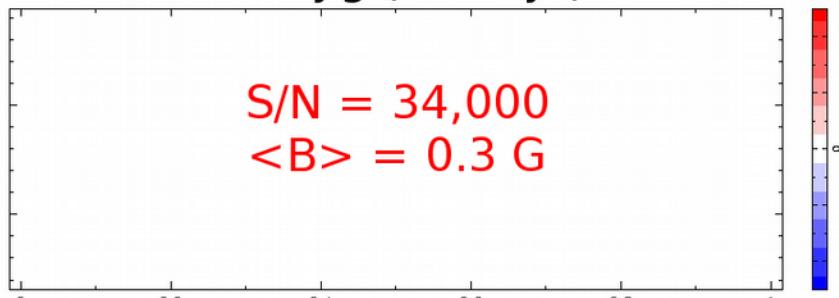
HD 76151 (2.5 Gyr)



18 Sco (4.0 Gyr)



16 Cyg (7.0 Gyr)

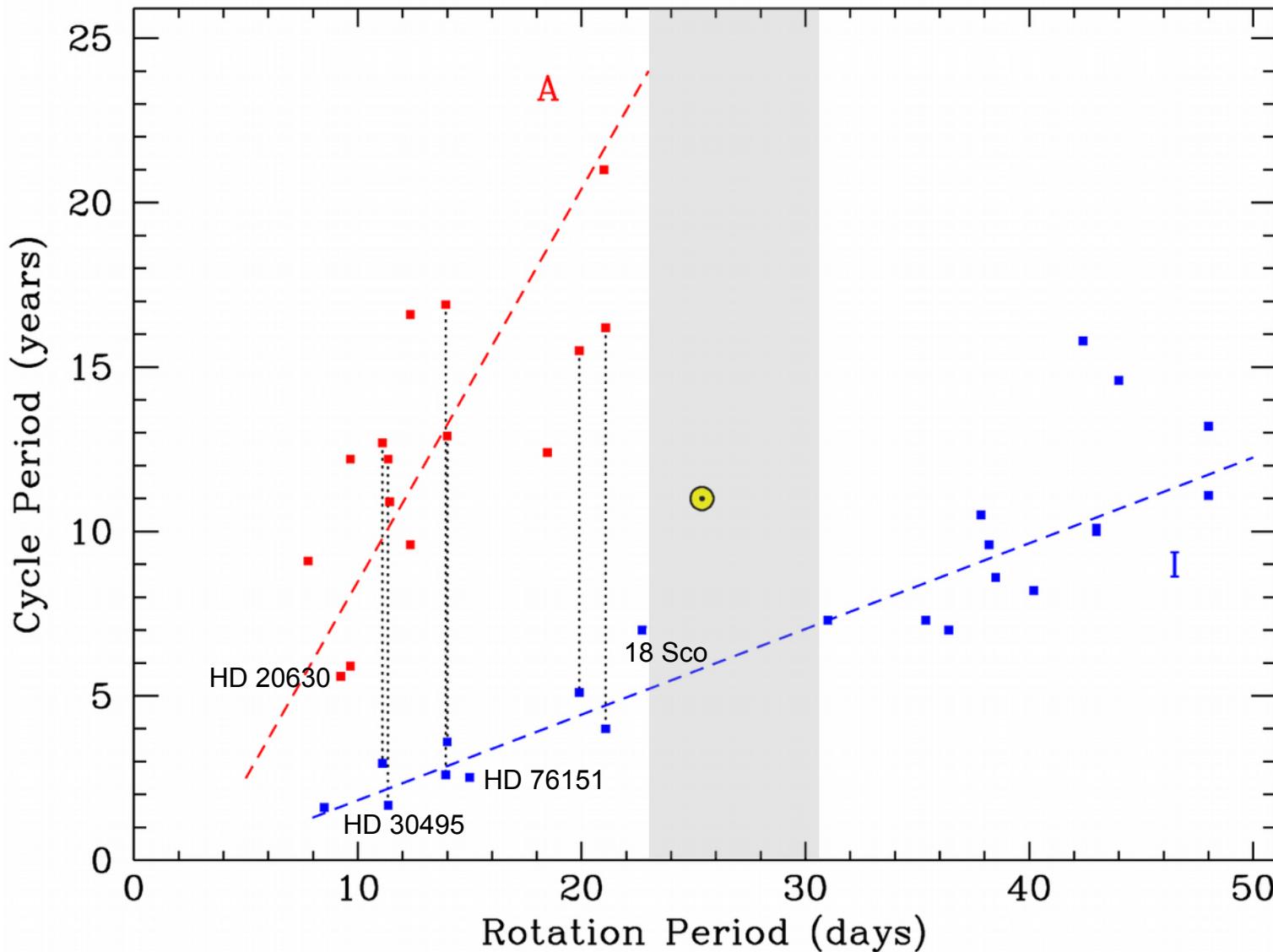


Rotational Phase

S/N = 34,000
 $\langle B \rangle = 0.3$ G

- Young solar analog is dominated by a dipole (80% of poloidal field)
- Dipole (35%) is already disappearing in 18 Sco, with 55% in quadrupole
- Old solar analog has no Zeeman signatures, weak average line-of-sight field

Stellar activity cycles



Summary

- A change in differential rotation at $\text{Ro} \sim 1$ pushes stars across the V-P gap, rapidly decreases spot area, then disrupts magnetic braking at $\text{Ro} \sim 2$
- The Sun is in a transitional evolutionary phase, and its 11-year activity cycle may represent a special case of stellar dynamo theory
- Future ZDI measurements, more constraints on differential rotation, and asteroseismology of the Mount Wilson sample will help test this scenario

More details on arXiv:1606.01926