

# A Python API for OTTER

Noah Franz <sup>1</sup>, Kate D Alexander  <sup>1</sup>, and Sebastian Gomez  <sup>2</sup>

<sup>1</sup> Department of Astronomy and Steward Observatory, University of Arizona, 933 North Cherry Avenue,  
<sup>2</sup> Tucson, AZ 85721-0065, USA <sup>2</sup> Department of Astronomy, The University of Texas at Austin, 2515  
<sup>5</sup> Speedway, Stop C1400, Austin, TX 78712, USA

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## Summary

The Open multiWavelength Transient Event Repository (or “OTTER”) is a new catalog of published transient data. Here we present a thick Python wrapper on the REpresentational State Transfer (REST) application programming interface (API) built-in to the OTTER backend database (the “OTTER API”). Since the OTTER backend is built on the document database ArangoDB, using the REST API directly requires learning the Arango Query Language (AQL). Since AQL has a niche user base, OTTER users unfamiliar with it may face a roadblock to programmatic access of OTTER. To overcome this barrier, we created the OTTER Python API to make programmatic access easy and fast. In addition to wrapping the REST API, the OTTER Python API also provides additional methods for 1) Converting the stored photometry to standard units; 2) Helper methods for querying additional astronomy database services; and 3) Methods for quickly plotting the photometry stored in OTTER.<sup>17</sup>

## Statement of need

Transient astrophysical events provide a unique high energy laboratory that evolves on human timescales. Examples include supernovae, gamma ray bursts, tidal disruption events, and many other exotic transients ([Colgate & White, 1966](#); [Eichler et al., 1989](#); [Heger et al., 2003](#); [Hills, 1975](#); [Kouveliotou et al., 1993](#); [Maoz et al., 2014](#); [Narayan et al., 1992](#); [Norris et al., 1984](#); [Rees, 1988](#); [Smartt, 2009](#); [Woosley & Weaver, 1995](#)). Developing an understanding of the many astrophysical transients observed is not easy and typically requires detailed multiwavelength observations and analyses of the population of events ([Alexander et al., 2025](#); [Christy et al., 2024](#); [Costa et al., 1997](#); [Eftekhari et al., 2018, 2021](#); [Gezari et al., 2017](#); [Gomez et al., 2024](#); [Guolo et al., 2024](#); [Hajela et al., 2025](#); [Laskar et al., 2022, 2023](#); [Margutti et al., 2023](#); [Masterson et al., 2024](#); [Olivares E. et al., 2015](#); [Pasham & Velzen, 2018](#)). However, it can take years to gather the multiwavelength datasets from the literature, necessitating publicly available transient event data archives.<sup>19</sup>

Furthermore, transient classification is a non-trivial process that typically requires detailed spectroscopic and/or multiwavelength follow-up observations (e.g., [Arcavi et al., 2014](#); [Charalampopoulos et al., 2022](#)), a method that is only feasible with the current transient discovery rates. With the advent of Rubin Observatory’s Legacy Survey of Space and Time ([Ivezic et al., 2019](#)) the number of known transients will increase by at least an order of magnitude ([Bricman & Gomboc, 2020](#); [Sjoert van Velzen et al., 2011](#)). Therefore, additional methods for classifying transients, such as machine learning, are required ([Boesky et al., 2025](#); [Gomez et al., 2020, 2023](#); [Soto et al., 2024](#); [Stein et al., 2024](#); [V. A. Villar et al., 2019](#); [V. Ashley Villar et al., 2020, 2021](#)). However, machine learning classifiers require large training datasets that can be laborious to curate. This further motivates archival services for cataloging transient metadata and photometry, and will be necessary to maximize the scientific output of the Rubin time domain survey.<sup>20</sup>

43 For both of these reasons we created the Open multiWavelength Transient Event Repository  
44 (OTTER, [Franz et al., 2025](#)), a scalable catalog of transient event metadata and photometry.  
45 OTTER is a successor to the Open Astronomy Catalogs (OAC, [Auchettl et al., 2017; Guillochon](#)  
46 [et al., 2017](#))<sup>1</sup>, but designed and optimized for multiwavelength datasets. To store the various  
47 nuances of multiwavelength photometry (e.g., the model used to reduce and extract a flux from  
48 an X-ray observation), we chose to use a flexible document database management system as  
49 our backend: ArangoDB. The nested structure of the document database files also provides an  
50 intuitive way to store multiple values of a single measurement when different sources disagree.

51 One of our primary goals of OTTER is ease of access to the dataset, including a way to  
52 programmatically access it to make the curation of large transient samples easier. ArangoDB  
53 has a built-in REST API for programmatic access to the data. However, the API endpoints  
54 expect queries in the syntax of the “Arango Query Language” (AQL). Learning a new query  
55 language creates a barrier for user programmatic access to the indispensable dataset available  
56 in the OTTER catalog.

57 To help overcome this barrier, we present a Python API for access to the OTTER dataset.  
58 This API acts as a thick wrapper on the AQL-based API, with many additions that make it  
59 easier to access and analyze the dataset. Some of these features include:

- 60     ▪ In OTTER we store photometry as close to the actual published value as possible to  
61       make the data more reproducible. However, this also means that the data is not stored  
62       in consistent units (but the unit of the photometry point is stored). In the OTTER API  
63       we automatically convert the photometry into the user-requested units. Specifically, the  
64       conversion is done in the `Otter.get_phot` and `Transient.clean_photometry` methods  
65       which use astropy ([Astropy Collaboration et al., 2013, 2018, 2022](#)) and synphot ([STScI](#)  
66       [Development Team, 2018](#)).
- 67     ▪ The same raw data from an astronomical observation may be reduced<sup>2</sup> by multiple,  
68       distinct, teams. Depending on the differences in the reduction methodology this may  
69       produce different flux measurements. If this is the case, we store both flux measurements  
70       in the OTTER database to allow the user to choose their preferred reduction. However,  
71       to help users de-duplicate these datasets while curating large samples, we provide an  
72       (optional) automated algorithm for finding duplicates and choosing only one of the  
73       multiple reductions. This is done in the `Transient.deduplicate_photometry` method.
- 74     ▪ Sometimes users want to quickly view the photometry for a specific transient event as  
75       either a light curve (flux as a function of time) or a spectral energy distribution (flux as  
76       a function of wavelength, frequency, or energy; i.e., an “SED”). In the plotter module  
77       of the OTTER API we provide numerous methods for quickly and automatically plotting  
78       the photometry ([Hunter, 2007; Inc., 2015](#)).
- 79     ▪ Identifying the host galaxy of a transient event can be difficult and it is important to  
80       store identifying information (e.g., name and coordinates) for a host galaxy, if it is  
81       known. However, there are numerous existing astronomical databases that store galaxy  
82       properties and we do not want to duplicate their efforts. We therefore provide methods  
83       as part of the `Host` object to query other public services for host photometry or spectra.  
84       These other services include Simbad ([Wenger et al., 2000](#)), ATLAS ([Tonry et al., 2018](#)),  
85       ZTF ([Sjoert van Velzen et al., 2019](#)), iPTF ([Blagorodnova et al., 2017](#)), ASAS-SN  
86       ([Hart et al., 2023; Kochanek et al., 2017; Shappee et al., 2014](#)), Vizier ([Ochsenbein et](#)  
87       [al., 2000](#)), WISE ([Hwang & Zakamska, 2020](#); [Mainzer et al., 2011, 2014](#); [Wright et](#)  
88       [al., 2010](#)), FIRST ([White et al., 1997](#)), NVSS ([Condon et al., 1998](#)), HEASARC, and  
89       Sparcl ([Juneau et al., 2024](#)) — most of which are queried using the astropy-affiliated  
90       astroquery package ([Ginsburg et al., 2019](#)).  
91     ▪ Users may want to compare new observations stored locally with the publicly available

<sup>1</sup>The OAC was an indispensable resource but has not been maintained since 2022, further necessitating a successor like OTTER.

<sup>2</sup>By “reduced” we mean that the proper calibrations are applied and a flux, flux density, or magnitude is extracted from the raw data.

92 data in OTTER. As part of the OTTER API we make this very easy as long as their  
93 data is stored in a well-documented CSV file format (see the OTTER web application  
94 upload form or the example jupyter notebook titled “Interfacing with Private Data”).  
95 When the data is stored like this a user is able to use the `Otter.from_csvs` method to  
96 construct an `Otter` object that will pass their queries to both the public OTTER dataset  
97 and the one locally stored and return all relevant information in a consistent format.  
98 ■ We allow for the storage of different measurements (e.g., redshift, discovery date, etc.)  
99 associated with the same property of the transient. The OTTER API will automatically  
100 choose a default value if multiple measurements are present for a single property.

## 101 Software Impact and Conclusions

102 Moving forward, OTTER, in its entirety, will be a useful infrastructure tool for time domain  
103 science. Even more, the OTTER API described here will make access to that dataset easier for  
104 users by lowering the API learning curve. Evidence of this is the impact of the Open Astronomy  
105 Catalogs, which has > 500 citations (Guillochon et al., 2017) and is still used today, despite  
106 being deprecated.

107 There are already multiple astronomers using the software for their research, spanning from  
108 undergraduate students to faculty. There are currently two papers citing OTTER (Alexander  
109 et al., 2025; Christy et al., 2025) and at least another in preparation (Farley et al., in prep.).  
110 Additionally, our immediate research groups has already used OTTER for writing successful  
111 telescope observing proposals. We presented this work at the Kavli Institute for Theoretical  
112 Physics: Towards a Physical Understanding of Tidal Disruption Events session, Astronomical  
113 Data Analysis Software and Systems: 2025 Monsoon Workshop, and the 2025 X-ray Quasi-  
114 Periodic Eruptions and Repeating Nuclear Transients Conference. It was positively received at  
115 all of these conferences and we accrued ~ 15 beta testers who provided invaluable feedback.  
116 We welcome GitHub issues with comments and feedback (or even pull requests!) from the  
117 community on our [GitHub repository](#).

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