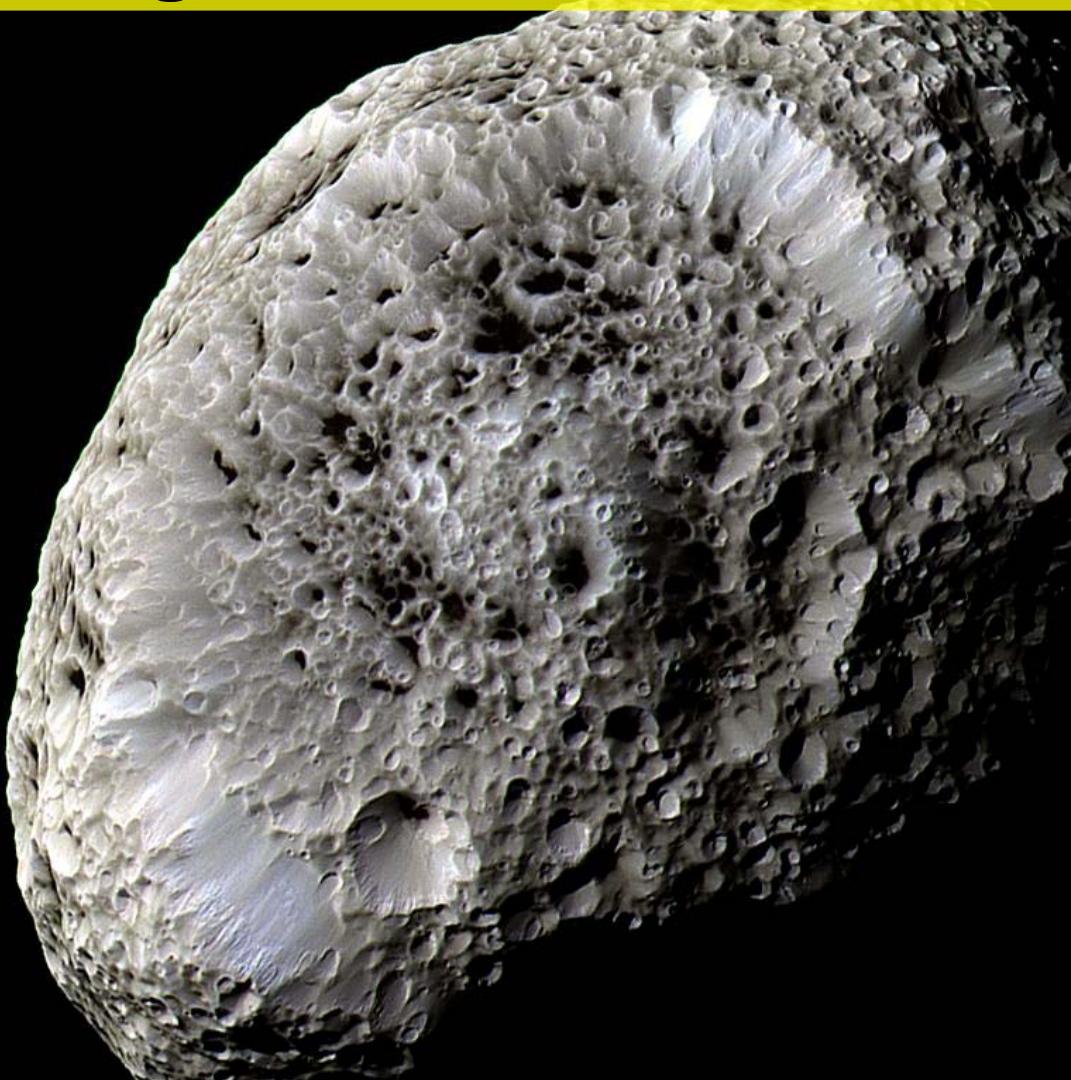


Rings and Moons of Saturn



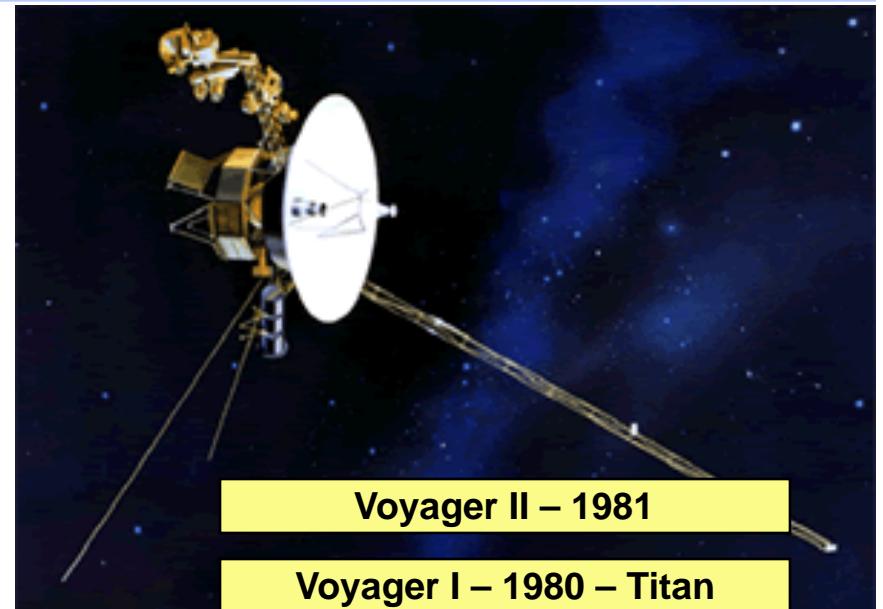
PTYS/ASTR 206 – The Golden Age of Planetary Exploration

Shane Byrne – shane@lpl.arizona.edu



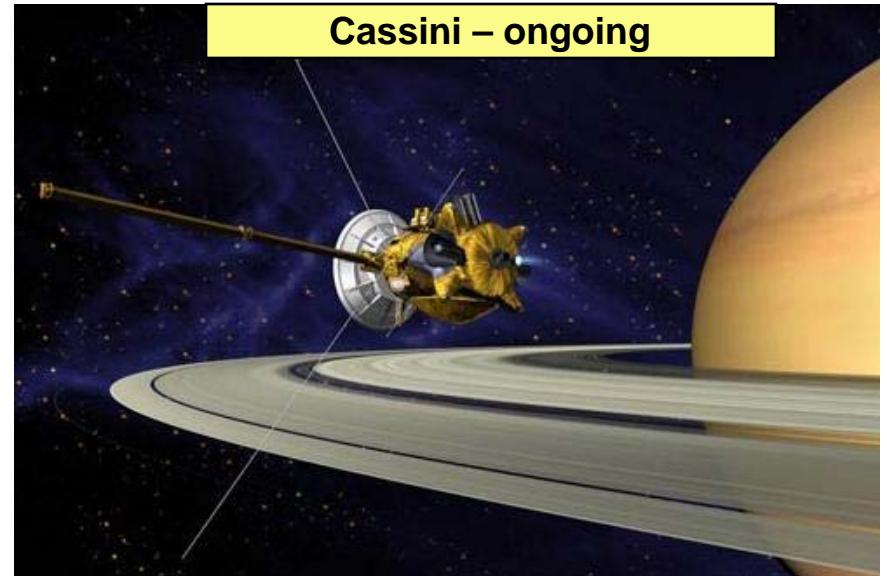
In this lecture...

- **Rings**
 - Discovery
 - What they are
 - How to form rings
 - The Roche limit
 - Dynamics
 - Gaps and resonances
 - Shepherd moons
- **Inner moons**
 - Tectonics and craters
 - Enceladus – a very special case
- **Outer Moons**
 - Captured Phoebe
 - Iapetus and Hyperion
 - Spray-painted with Phoebe debris



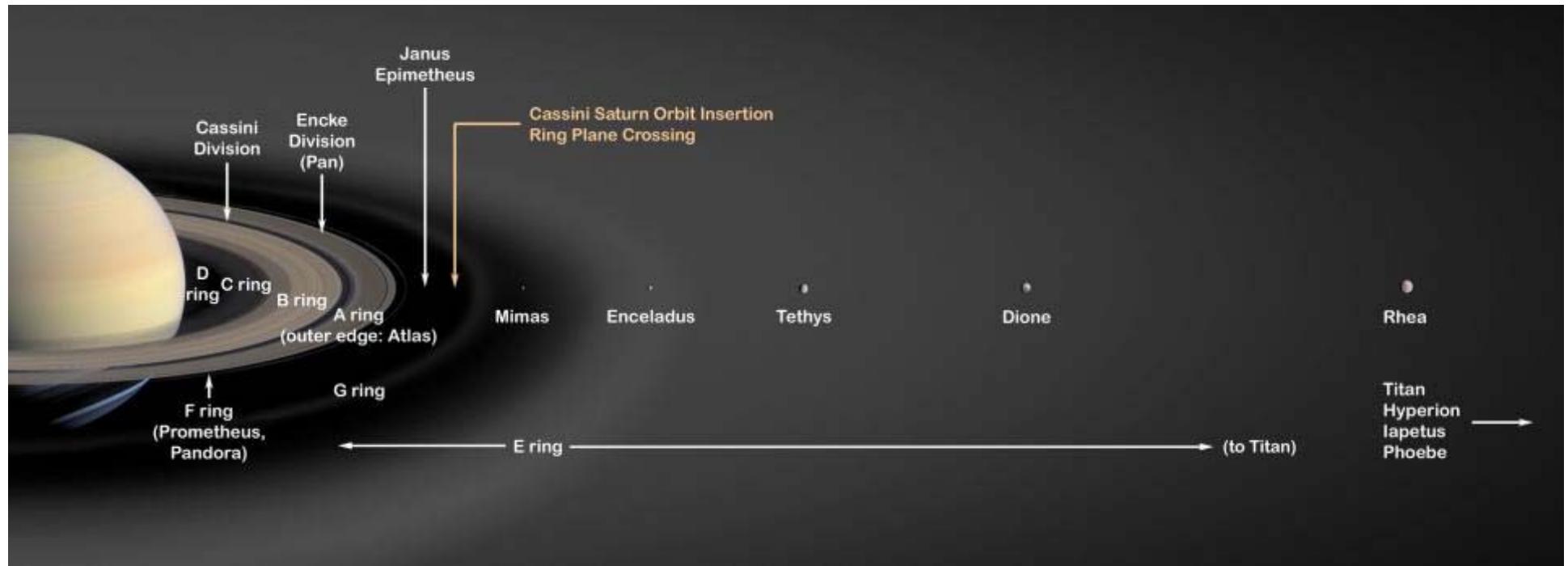
Voyager II – 1981

Voyager I – 1980 – Titan



Cassini – ongoing

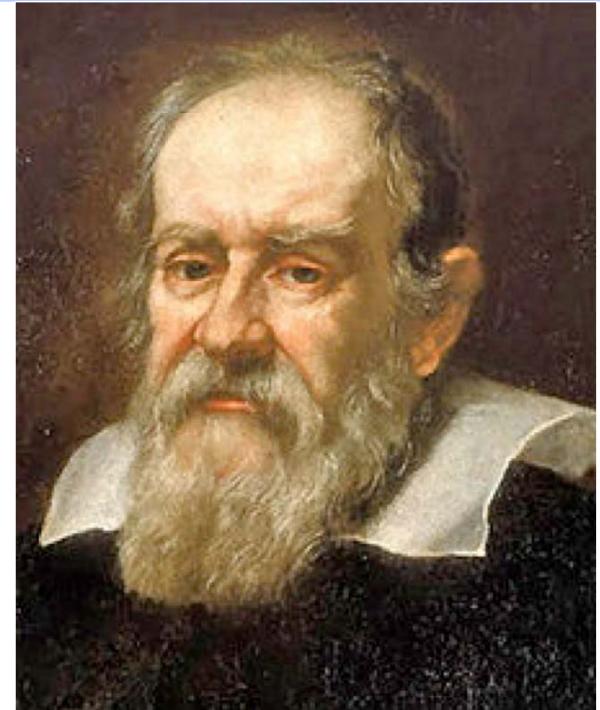
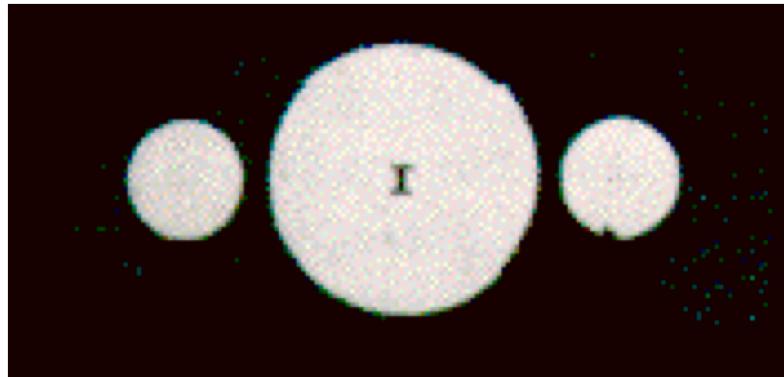
- We can divide Saturn's system into three main parts...
 - The A-D ring zone
 - ▶ Ring gaps and shepherd moons
 - The E ring zone
 - ▶ Ring supplies by Enceladus
 - ▶ Tethys, Dione and Rhea have a lot of similarities
 - The distant satellites Iapetus, Hyperion, Phoebe
 - ▶ Linked together by exchange of material





Discovery of Saturn's Rings

- Discovered by Galileo
- Appearance in 1610 baffled him

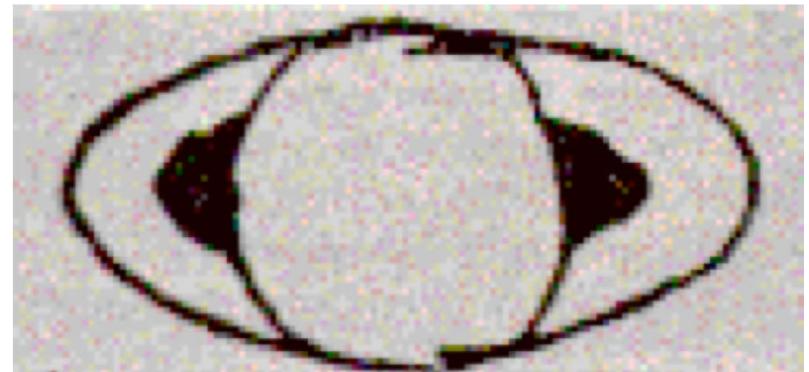


“...to my very great amazement Saturn was seen to me to be not a single star, but three together, which almost touch each other”

- It got more confusing...
 - In 1612 the extra “stars” had disappeared

“...I do not know what to say...”

- In 1616 the extra ‘stars’ were back
 - Galileo’s telescope had improved
 - He saw two “half-ellipses”
 - He died in 1642 and never figured it out



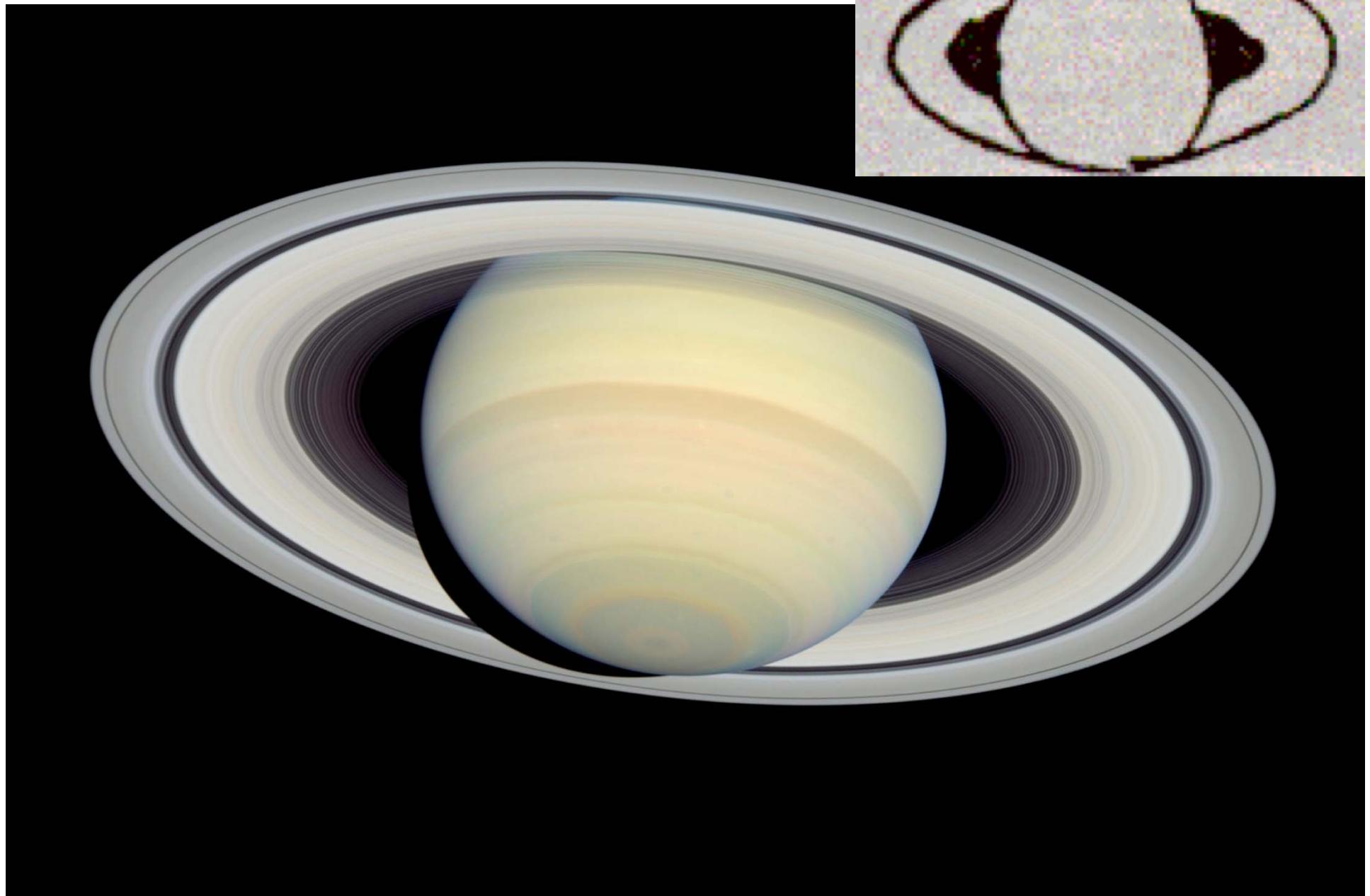
- In the 1650s Huygens figured out that
 - Saturn was surrounded by a flat disk
 - The disk disappears when seen edge on
 - He discovered Saturn's largest Moon – Titan
 - ...but he thought the disk was solid



- Other notable contributions
 - Cassini (1676) discovered there is more than one ring
 - Chapelain (1660) and Maxwell (1856) suggest the rings are a swarm of small particles

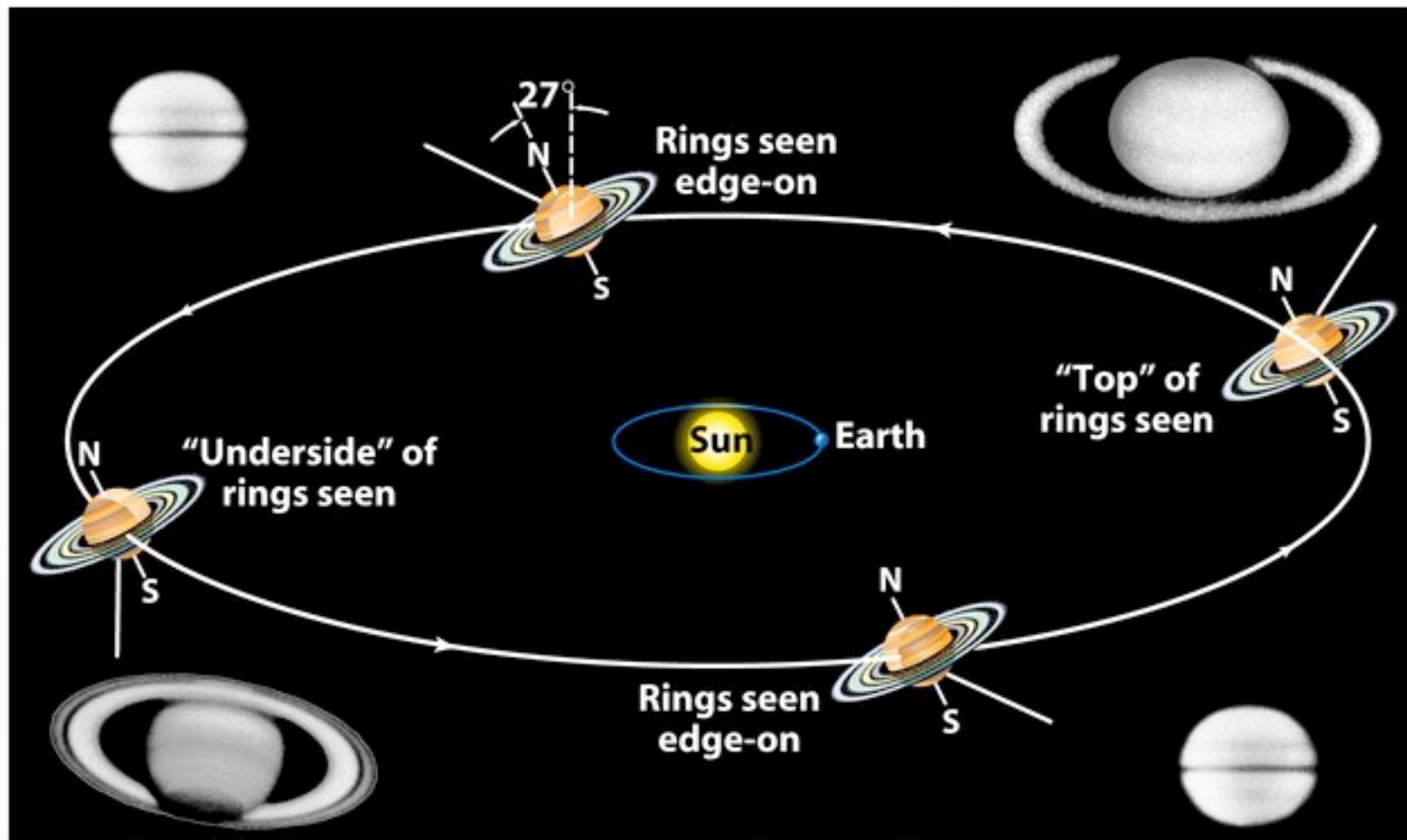


- 400 years makes a big difference



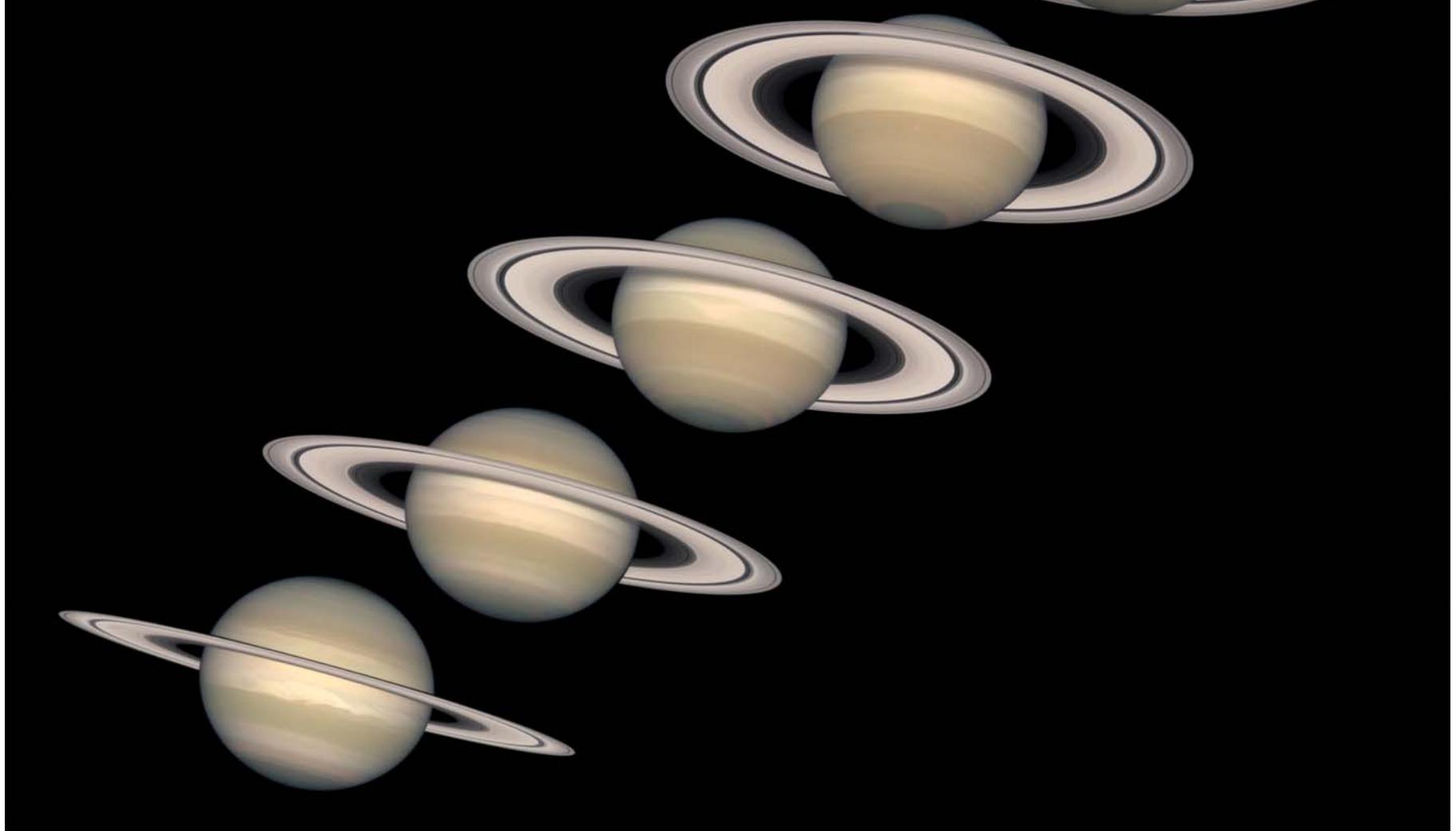
- Ring plane crossings

- Saturn's rings orbit around its equator
- Saturn is inclined to the ecliptic (where the Earth is) by 27°



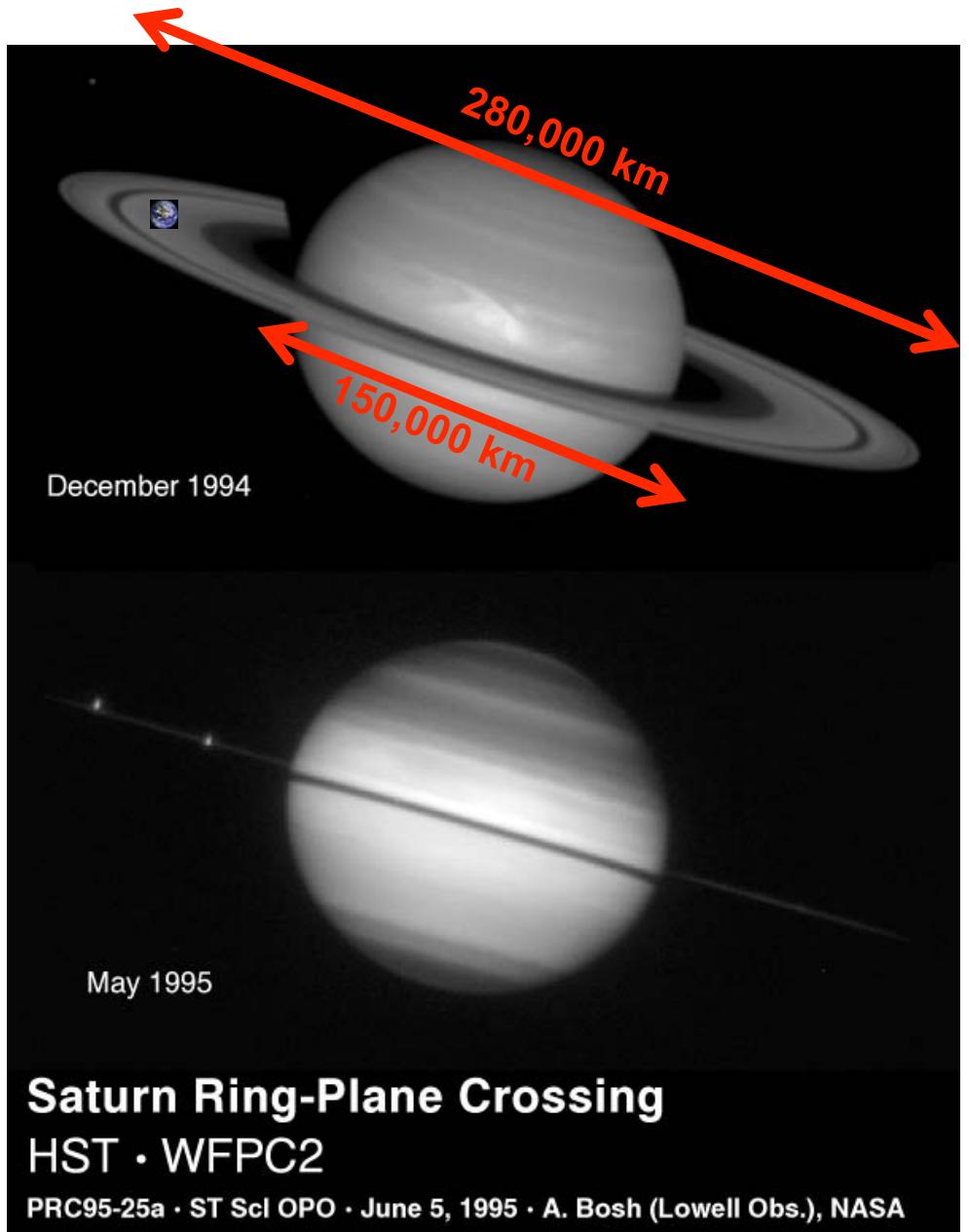


- This entire cycle takes 29.4 years –
Saturn's orbital period
 - Saturn's rings seen edge-on every 15 years (including in 2009)

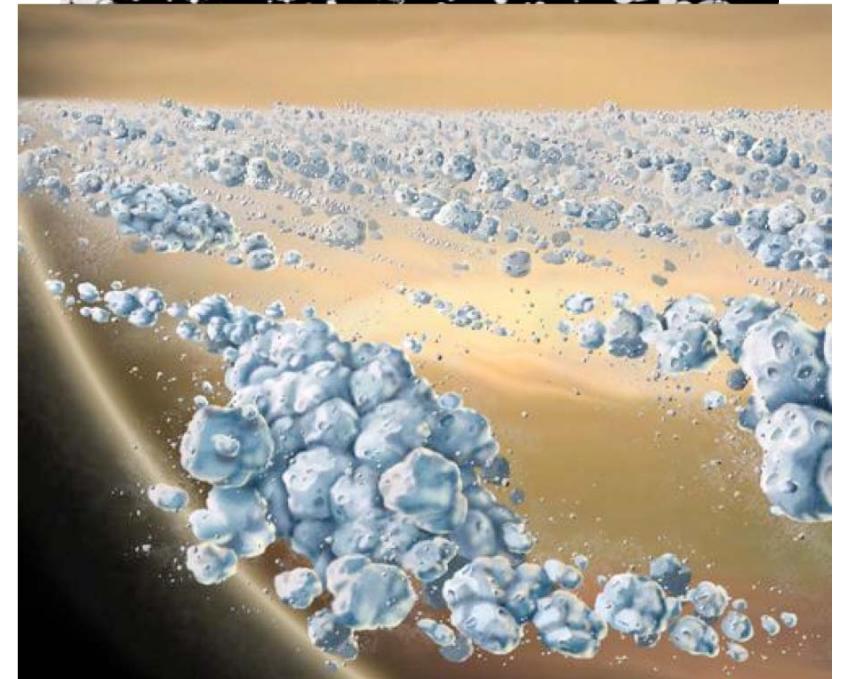
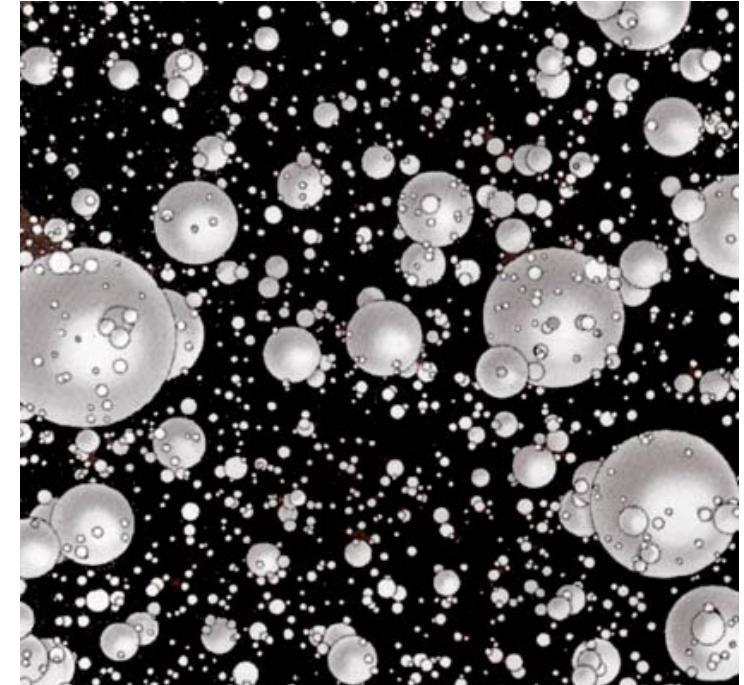




- Moons are easier to see during ring plane crossings
- Rings are **Very** thin
 - Less than 200m thick
 - Probably only a few 10s of meters thick
 - ...and 280,000 km across!
 - Much thinner than a sheet of paper scaled up
 - A standard piece of copier paper scaled up to the size of Saturn's rings would be about 100km thick
 - Rings are 1000 times thinner than this

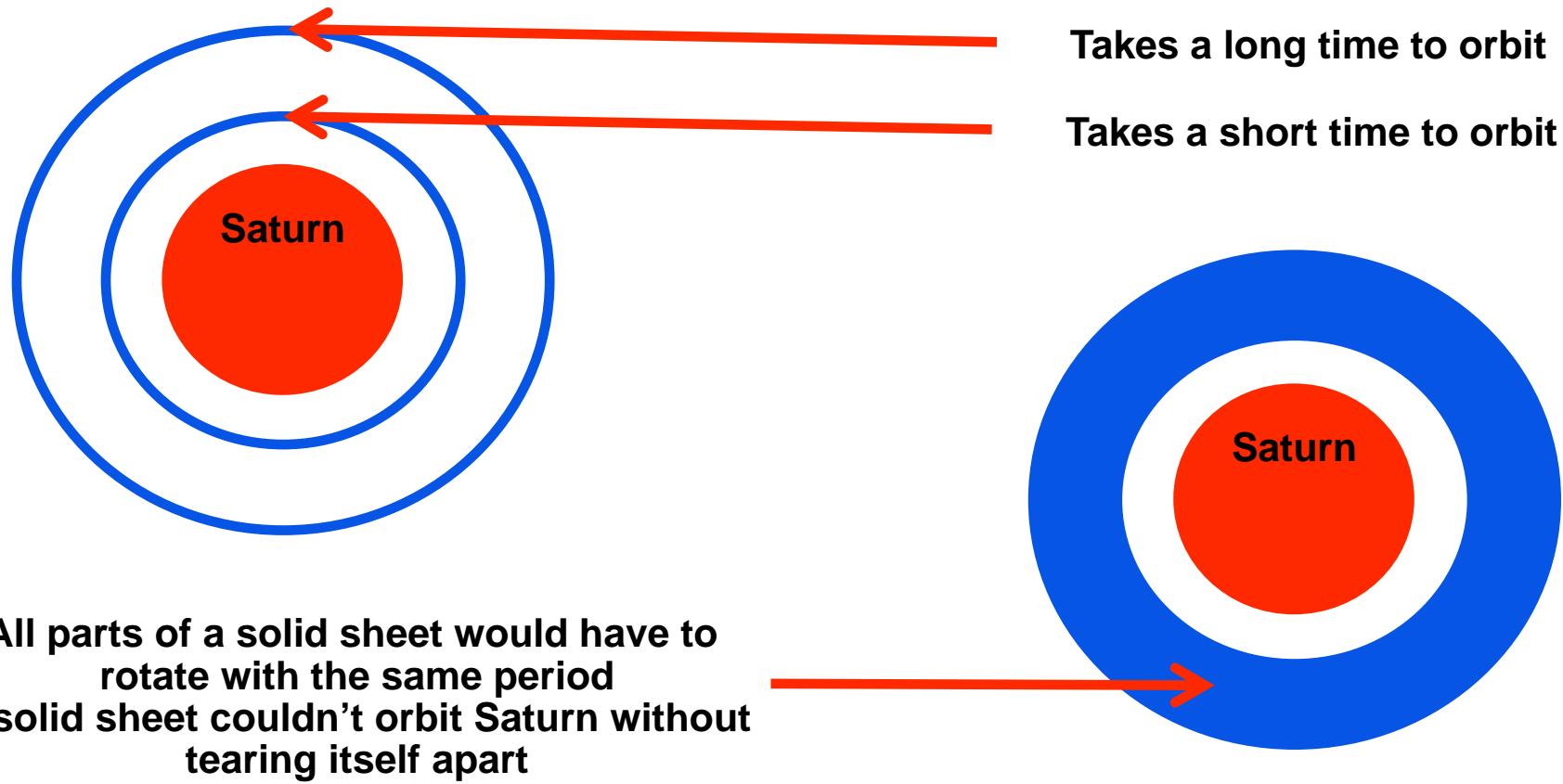


- Rings are composed of a swarm of small particles
 - Size range 1cm – 5m
 - Power law distribution – smaller particles more numerous than larger ones
 - ▶ Just like impact crater sizes
 - Particles are breaking up into smaller pieces due to collisions
- Spectra show these particles to be water ice
- Small particles can temporarily clump together

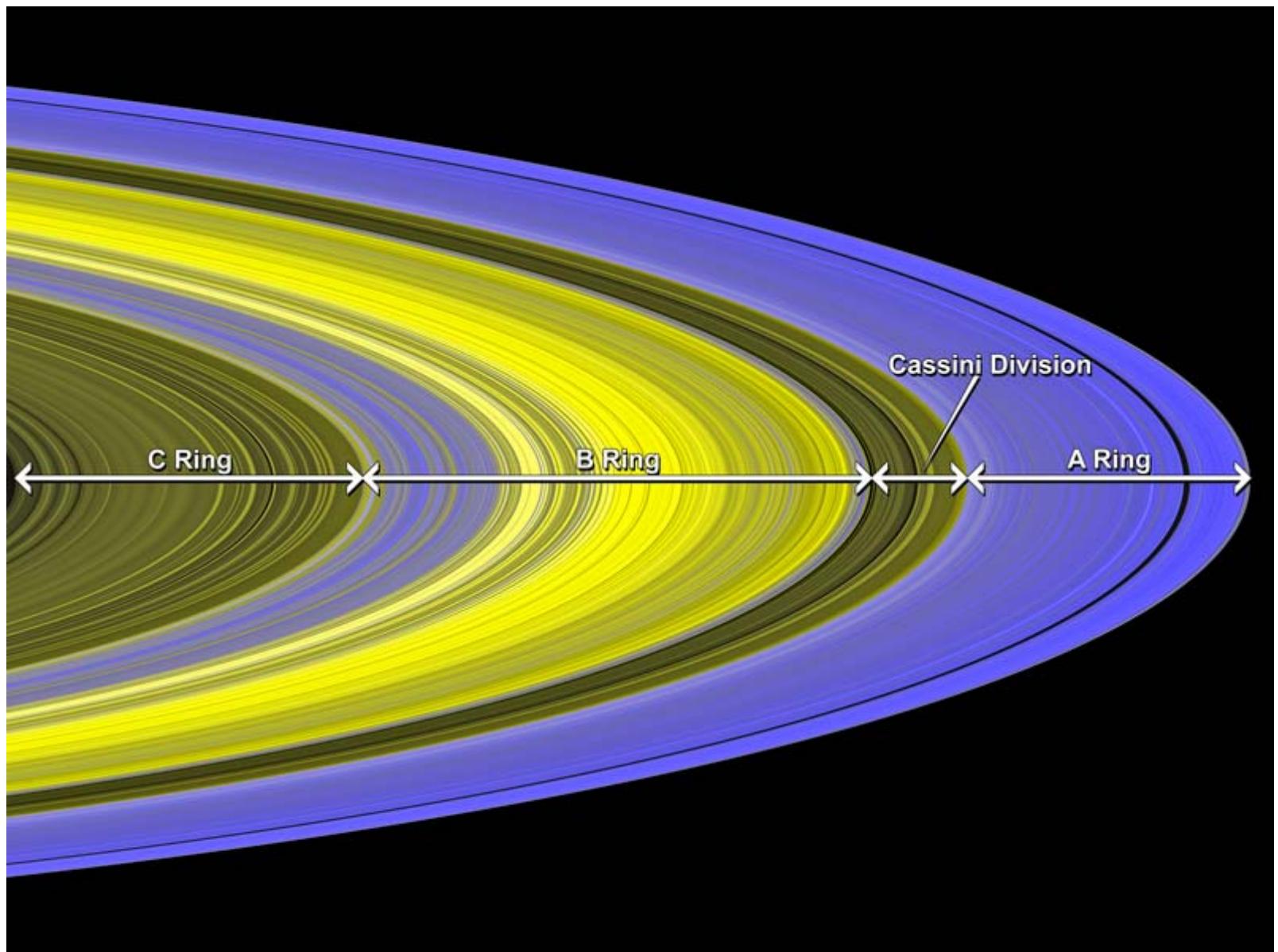


- Why can't Saturn's rings be solid sheets...
 - Remember Kepler's 3rd law
 - Objects nearer the Sun take less time to orbit
 - E.g. Earth takes 1 year, Mars takes 1.9 years
 - The same is true for objects orbiting Saturn

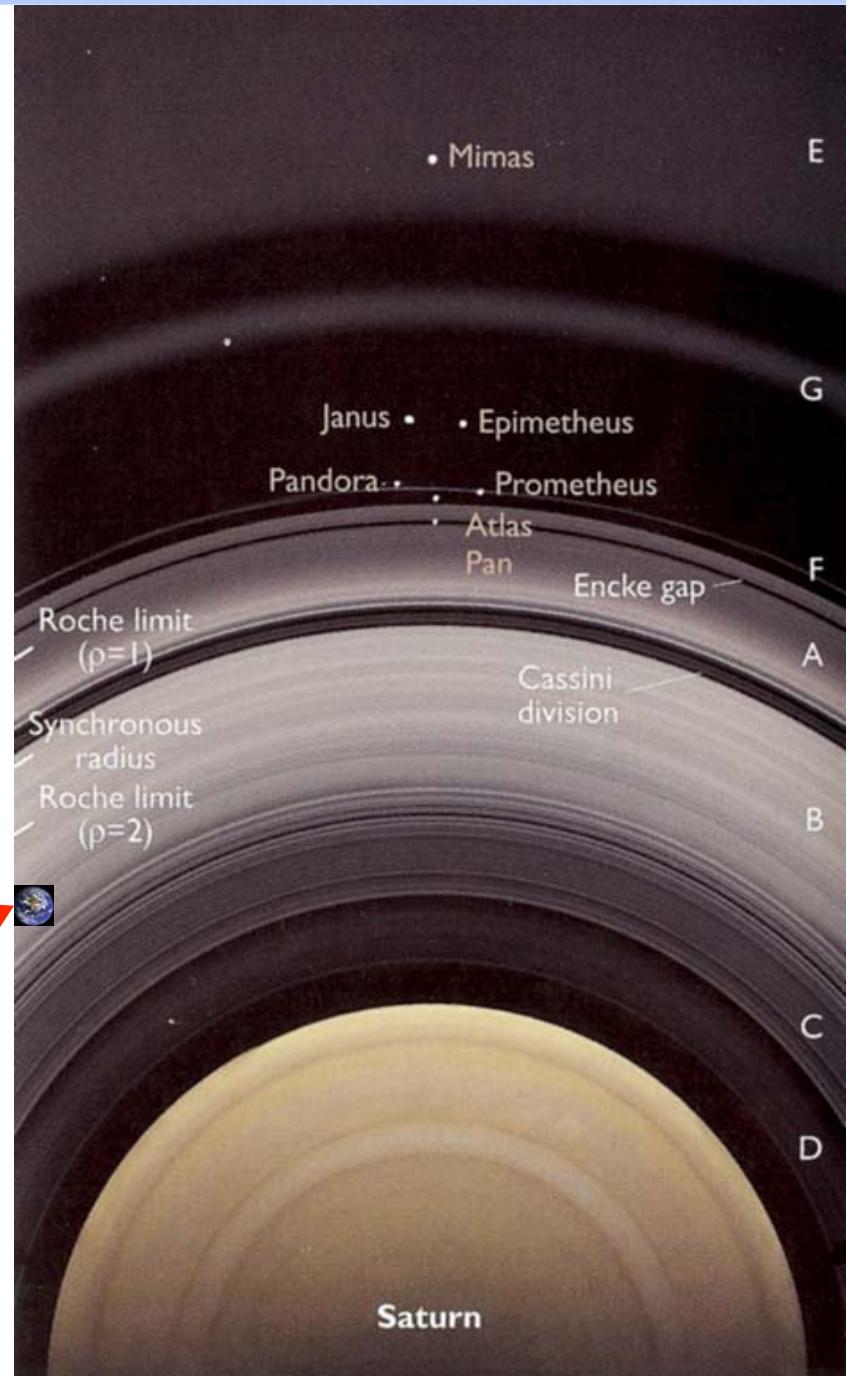
$$P^2 \propto a^3$$



- There's some compositional variation between rings
- Very little mixing of material from ring to ring – i.e. particles have very circular orbits

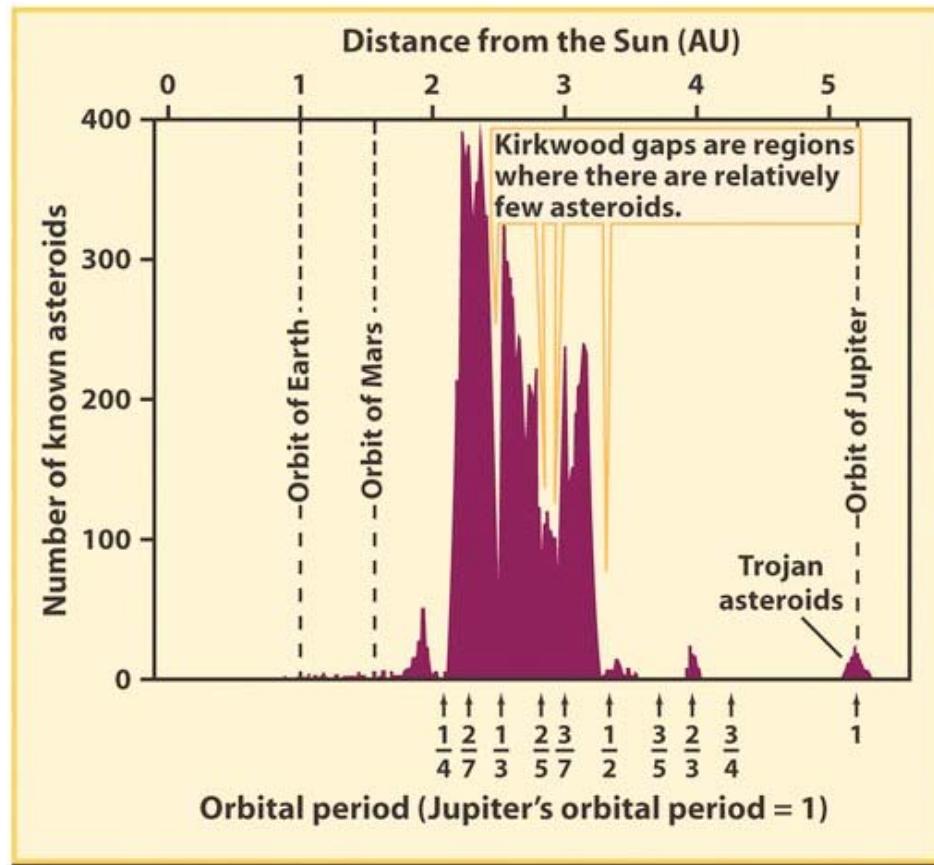


- Rings labeled by letters
 - A ring $200\text{-}500 \text{ Kg m}^2$
 - B ring $500\text{-}800 \text{ Kg m}^2$
 - C ring 10 Kg m^2
- Total mass of the rings
 - Similar to Mimas, $\sim 400\text{km}$ across
- Separated by gaps
 - The gaps are not totally empty
 - Cassini gap - 100 Kg m^2
- Ring albedo is very high, ~ 0.8
- Earth for scale

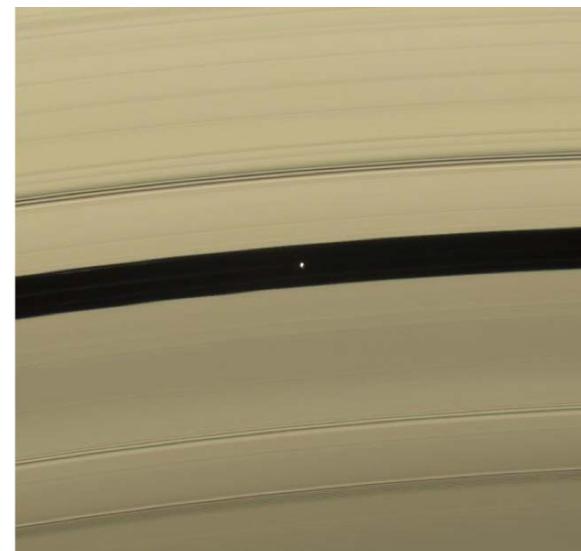




- Gaps in the rings caused by resonances with moons
 - Just like Kirkwood gaps in the asteroid belt – resonances with Jupiter
 - Periods are simple fractions of moon periods

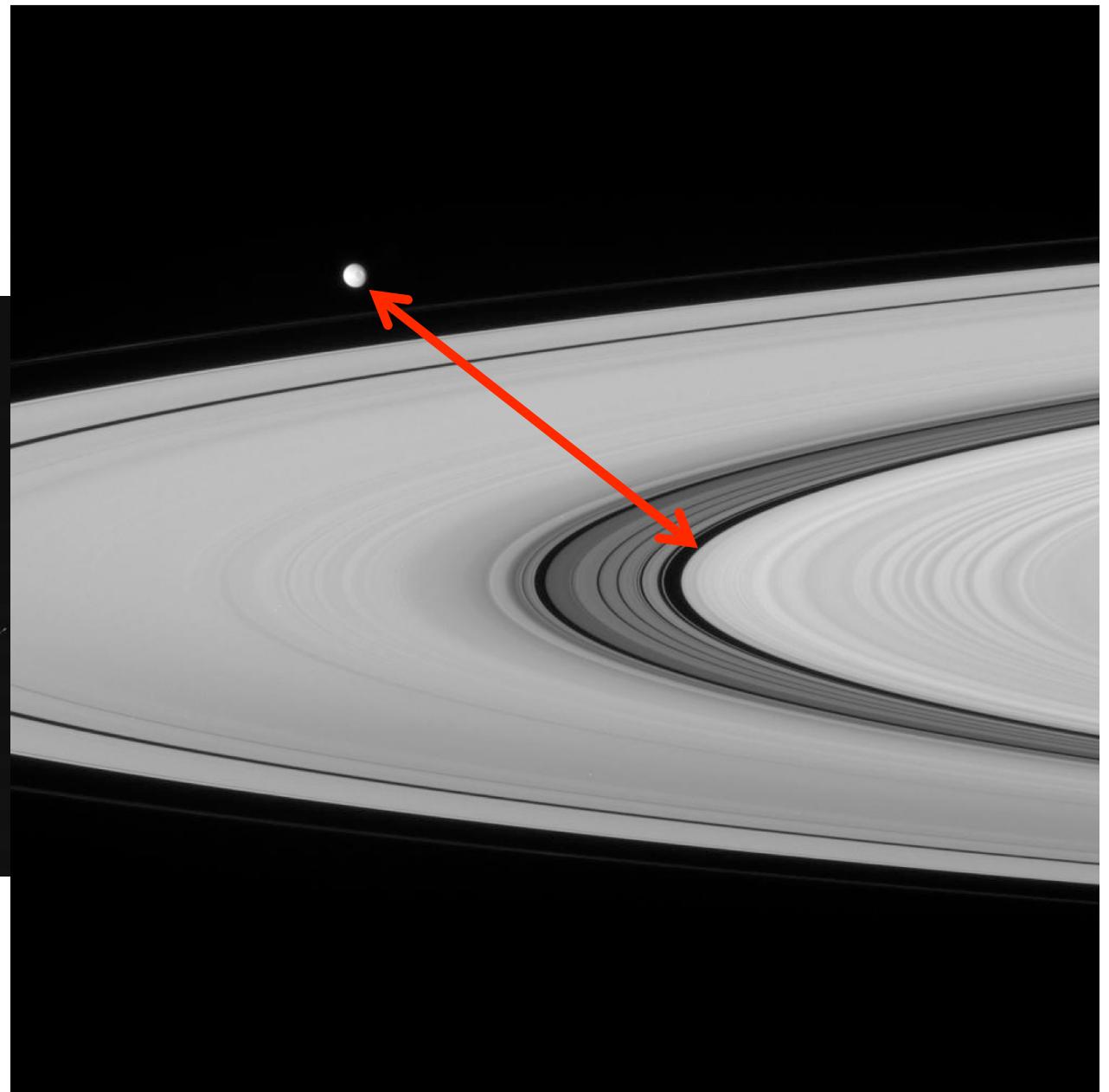


- ...but a little different
 - Asteroid orbits were eccentric
 - Asteroids criss-crossed the Kirkwood gaps
 - Ring particle orbits are very circular
 - Gaps really do look like gaps



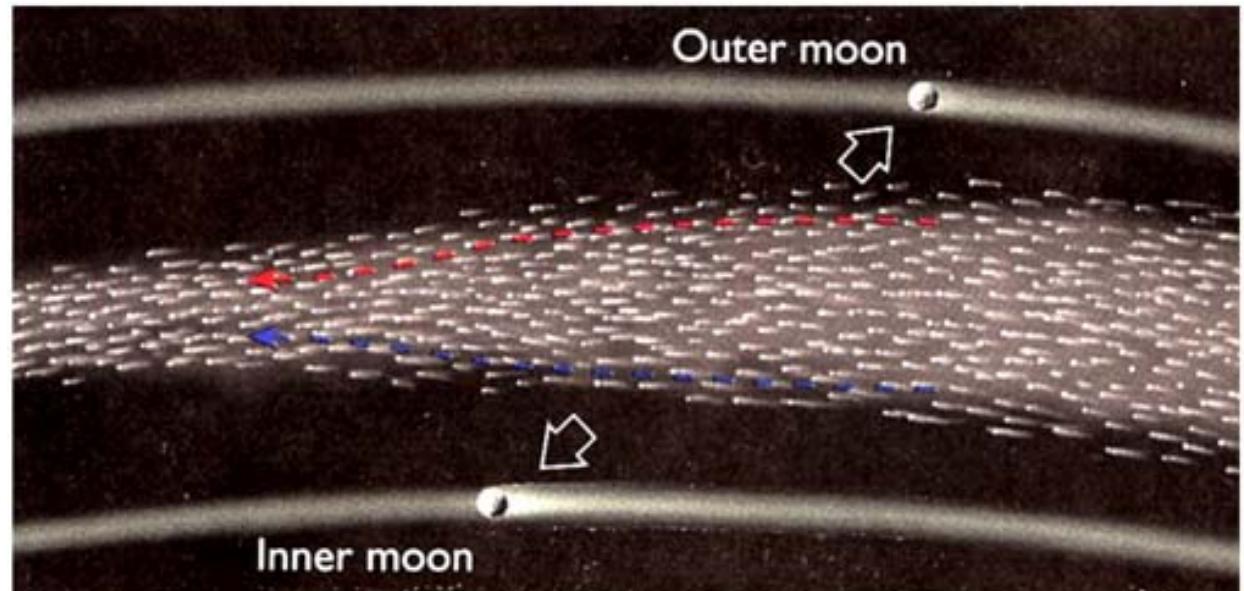
- Huygens gap

- 2:1 resonance with Mimas
- Mimas orbits in ~1 day

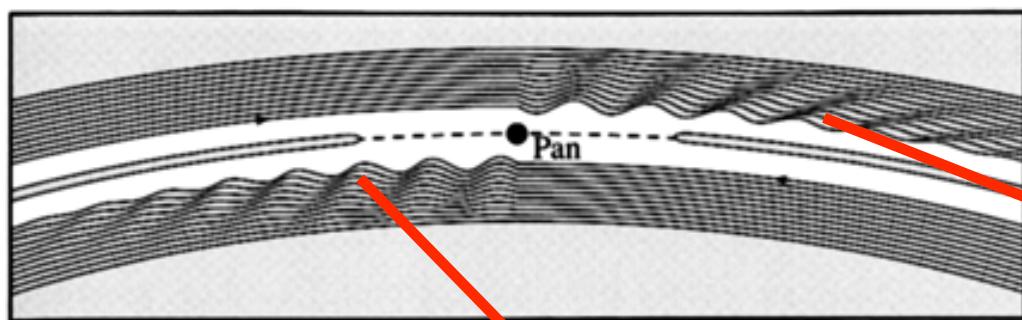


- Collisions between ring particles should cause the rings to spread out...
 - Interactions with Moons can keep the edges sharp
 - Swarm of ring particles is ‘herded’ by shepherd moons

- Moon ‘repel’ ring particles
- Moons speed up and slow down passing ring particles causing them to drift outwards/inwards
- Not a totally symmetric process
 - ▶ Inner moons cause particles to drift outwards
 - ▶ Outer moons cause particles to drift inwards



- Two types of shepherding...
- Gaps
 - Encke gap produced by Pan
 - Particles on either side repelled
 - Predicted density waves observed by Cassini





- Keeler gap is another example

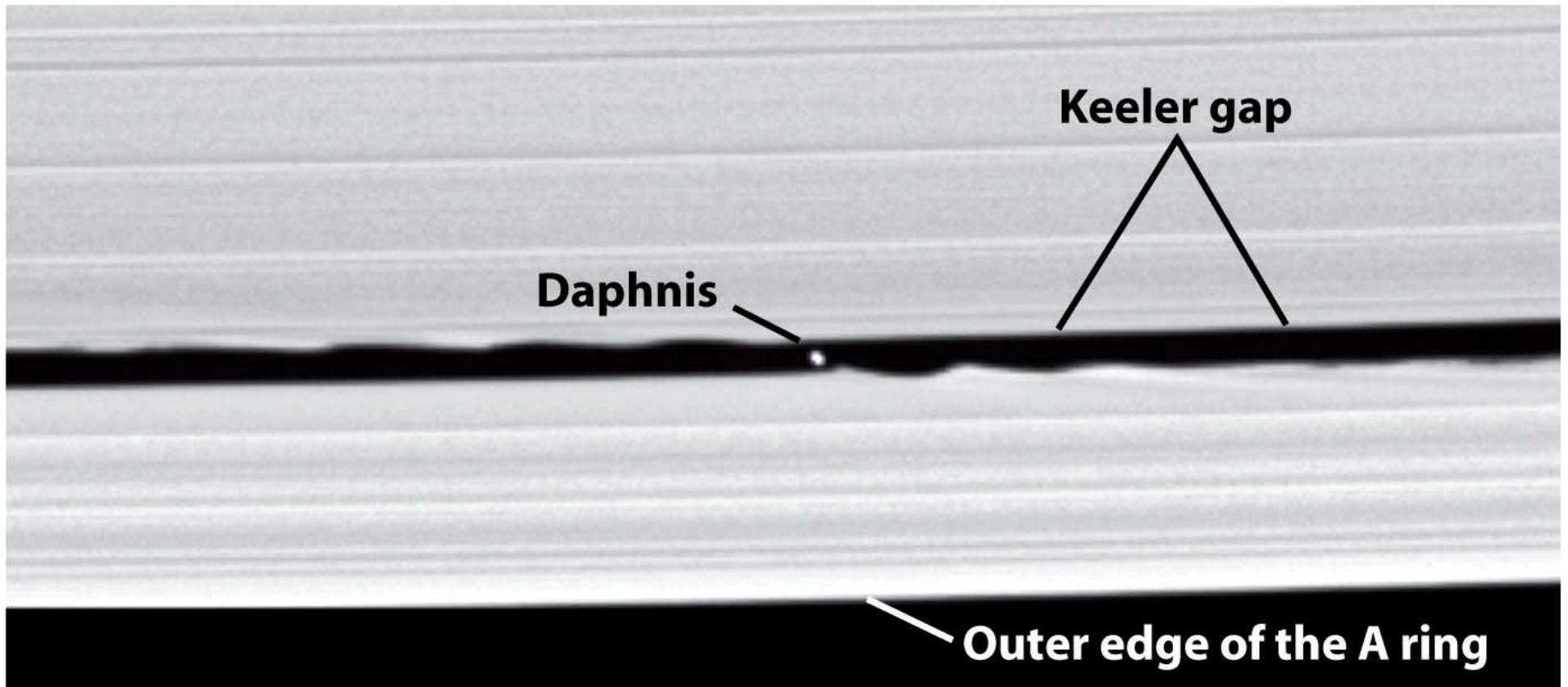
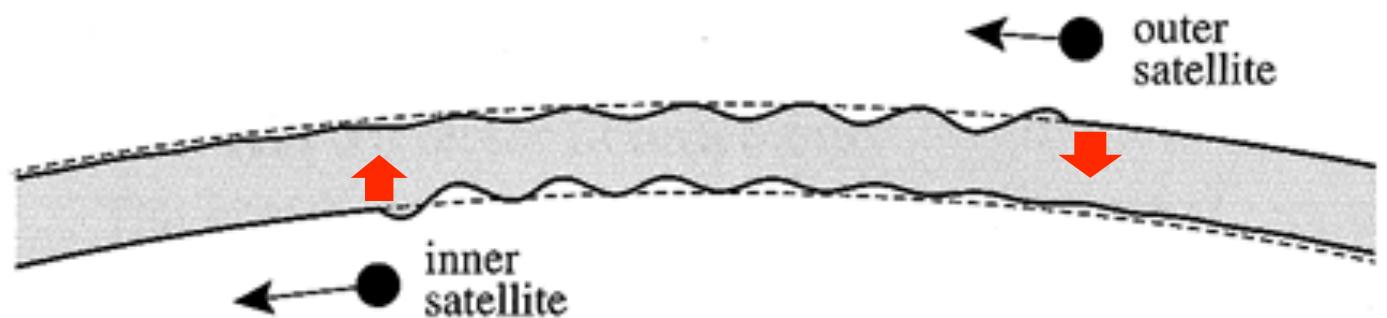
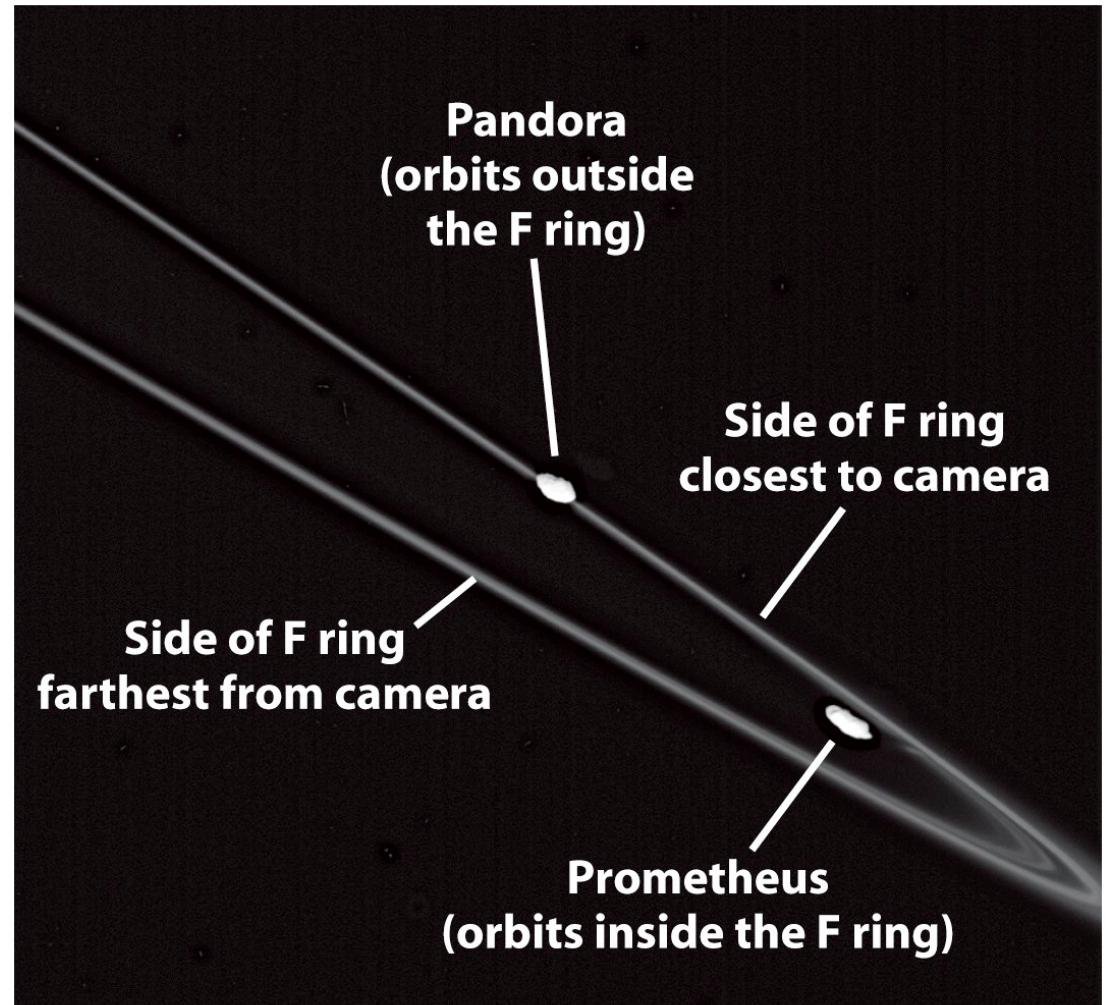


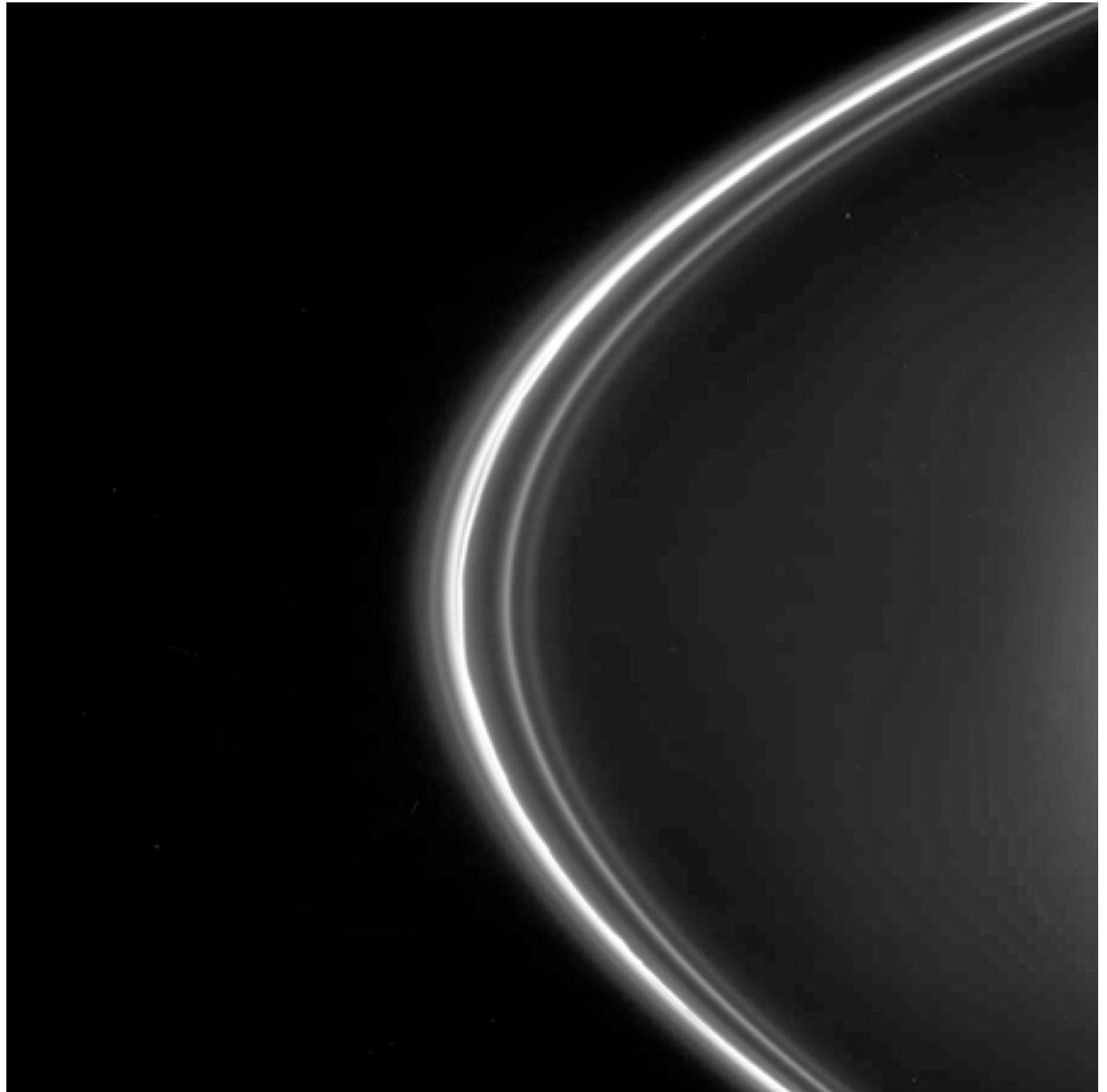
Figure 12-25

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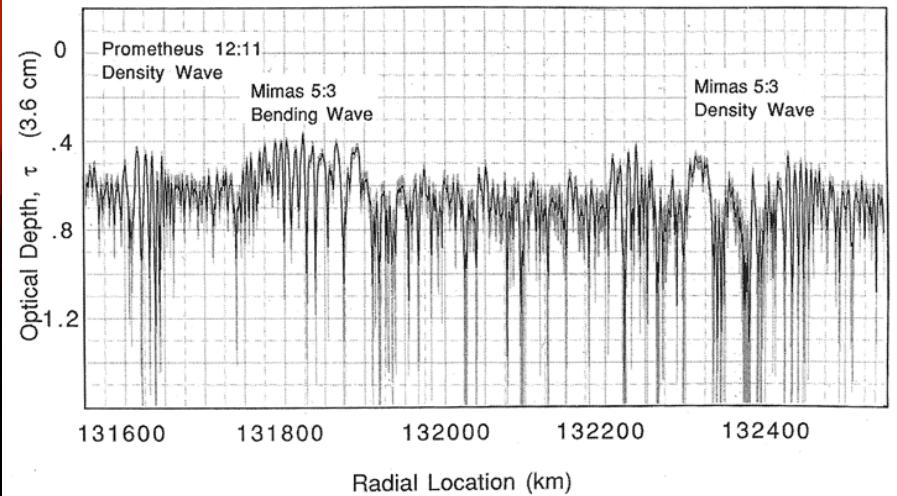
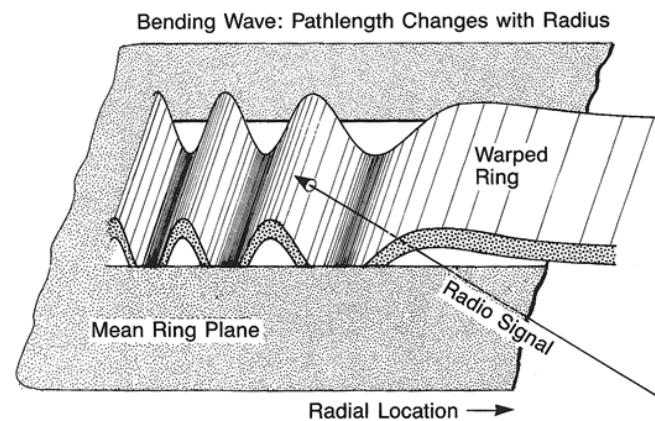
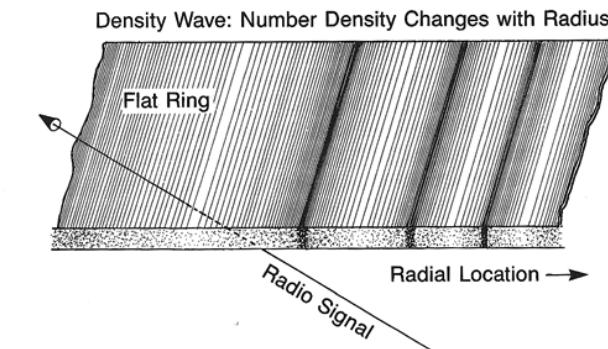
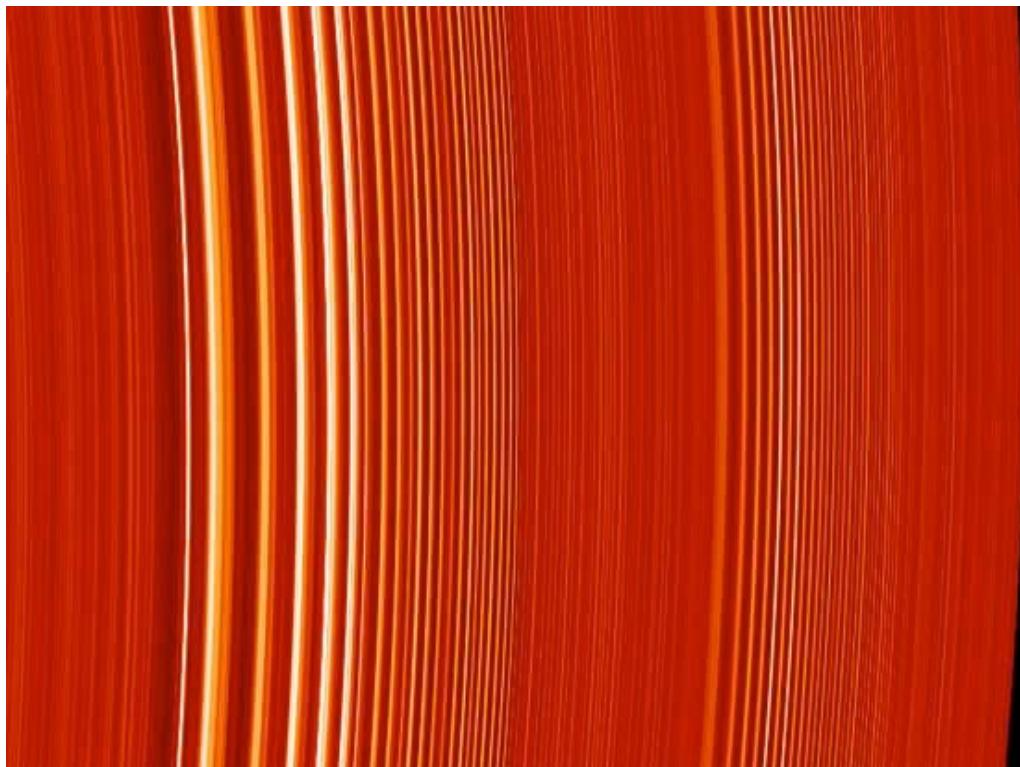
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- Shepherd moons also produce confined rings
 - E.g. F ring confined by Pandora and Prometheus
 - Pandora pushes particles inwards
 - Prometheus pushes particles outwards
 - Confines the F ring





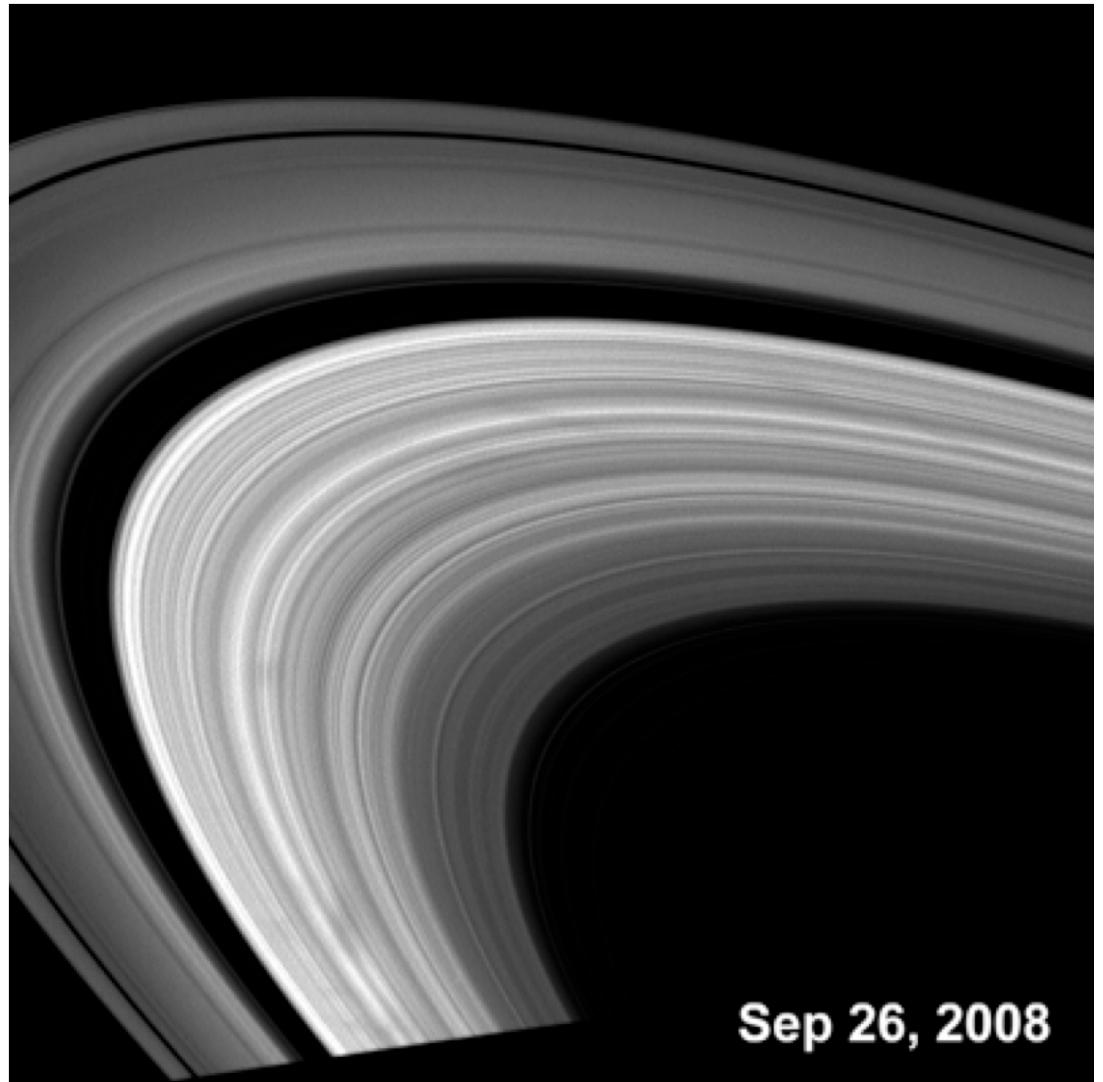
- Moons set up bending and density waves in the ring plane
 - We detect these tracking how the radio signal from the spacecraft is attenuated
 - Dips in the signal correspond to dense rings



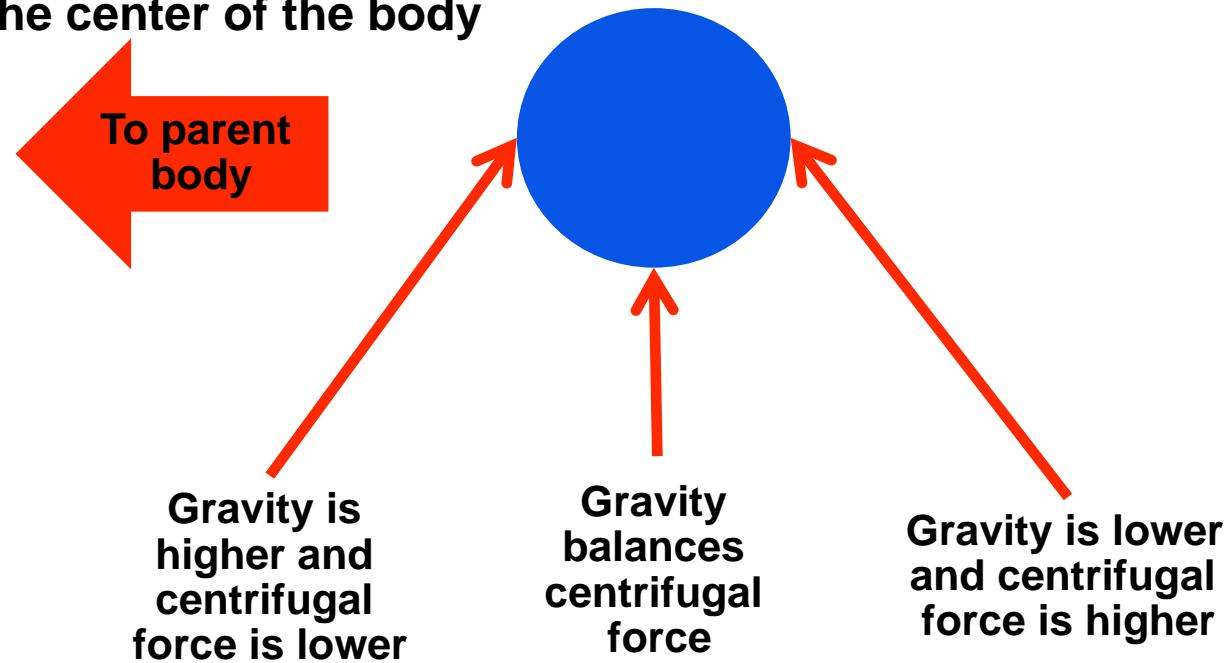


- Spokes

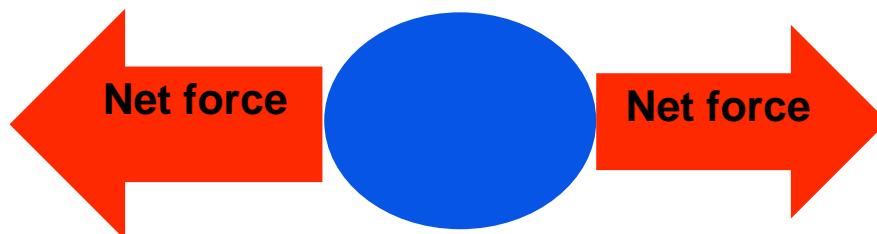
- Probably electrostatically charged dust-sized particles
- Sweep around with Saturn's magnetosphere
- Visible only at certain seasons
 - Near equinoxes – like now



- Where do the rings come from?
- Recap on tides
 - A body is in orbit when gravitational forces balance centrifugal
 - Only works at the center of the body

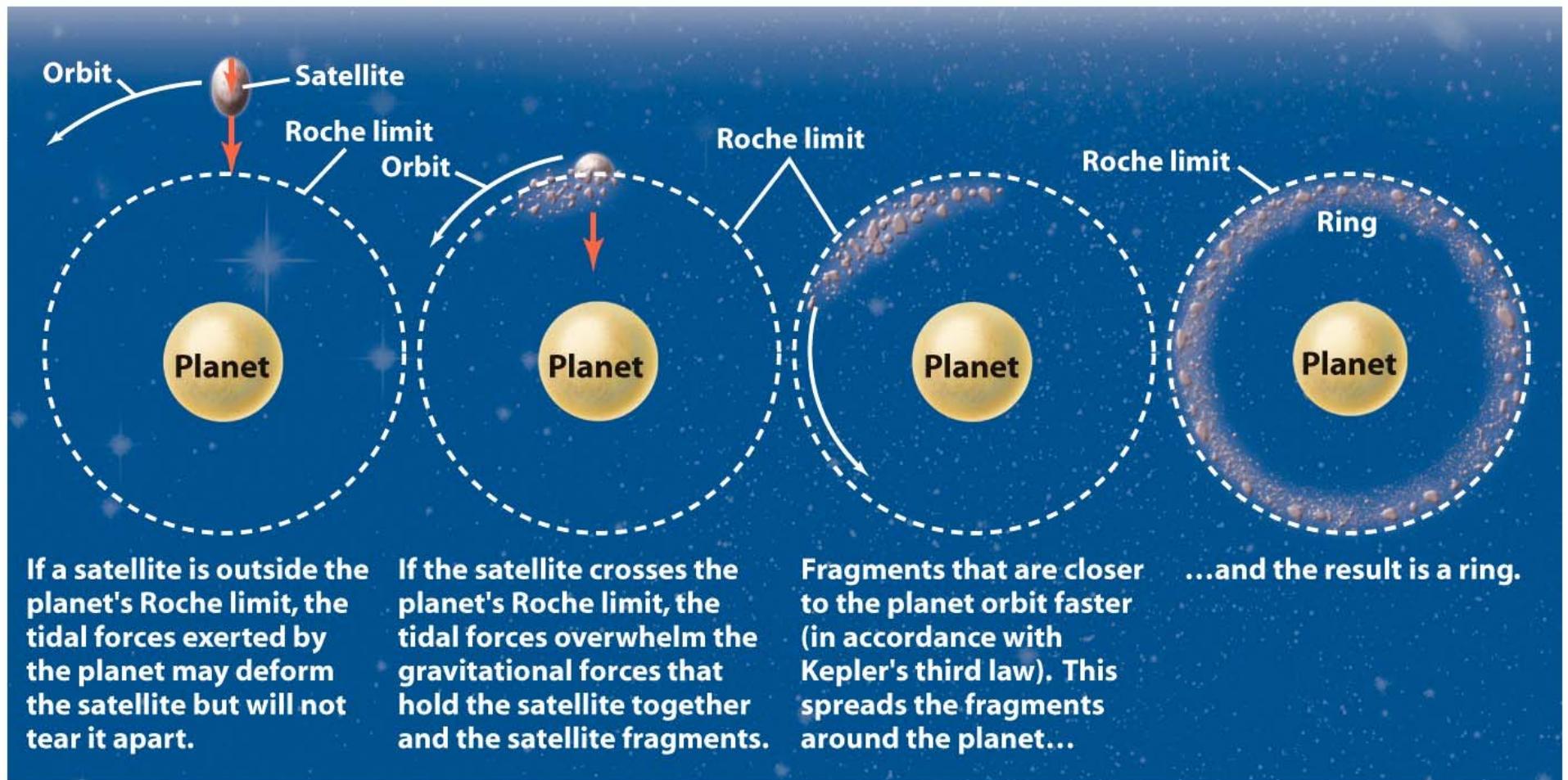


- Difference in forces elongates the moon – i.e. raises tides

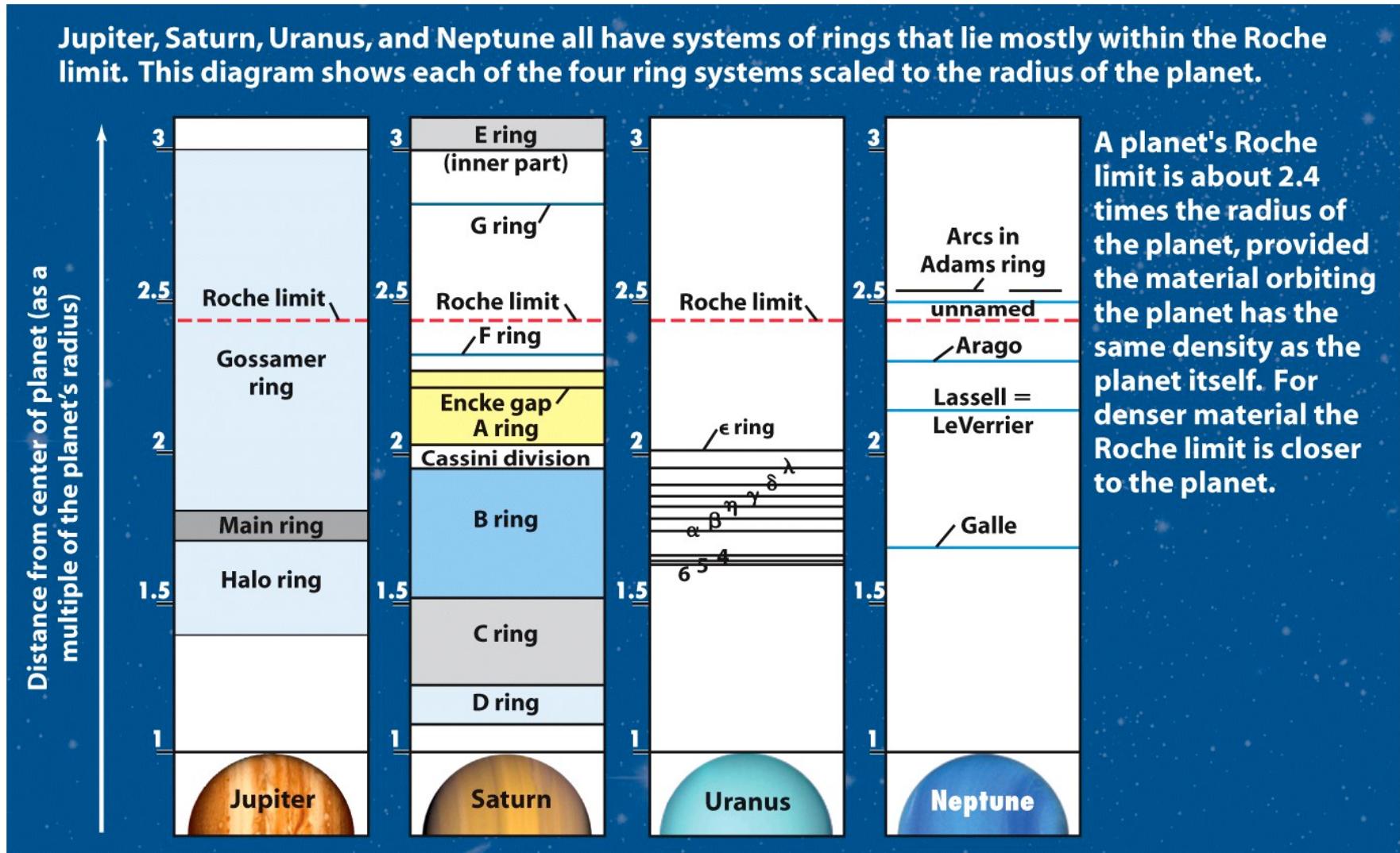


- What if the tidal force is very strong?

- Stronger than the moon's self-gravity...
- Tidal forces can tear a moon apart
 - Much like comet Shoemaker-Levy 9

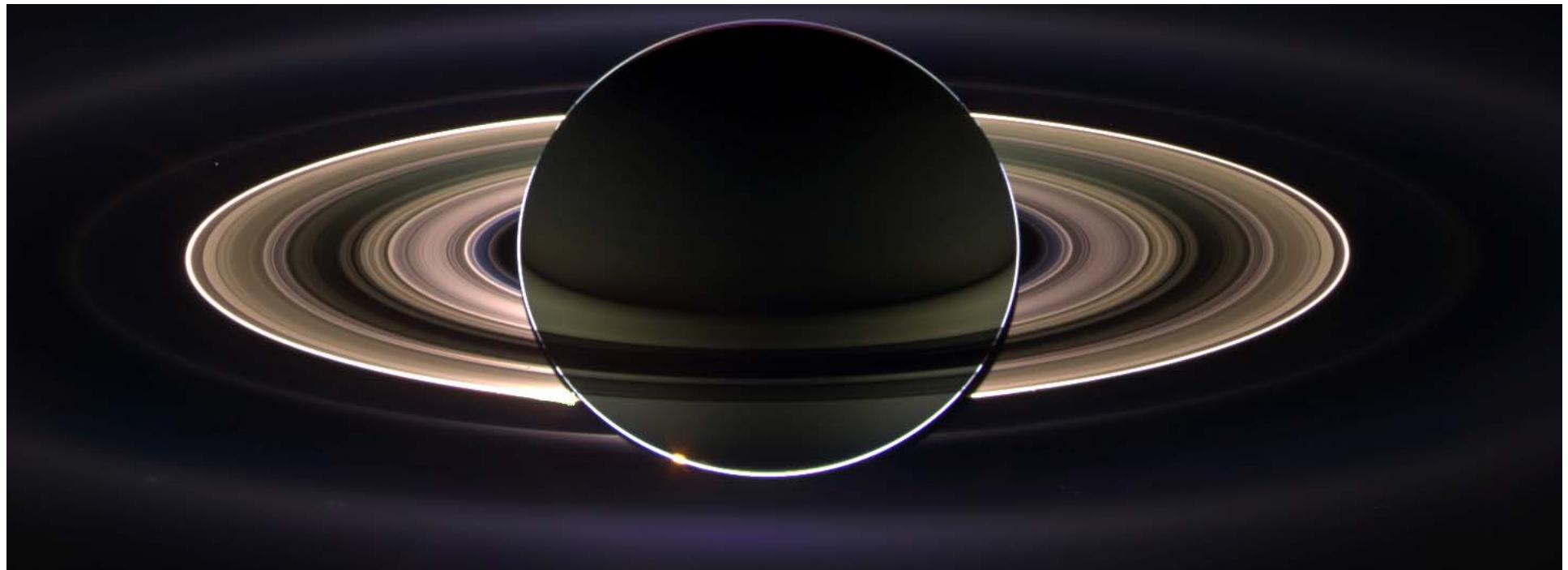


- Planetary rings are almost all within the Roche limit
 - Stops ring particles from sticking together into a bigger moon

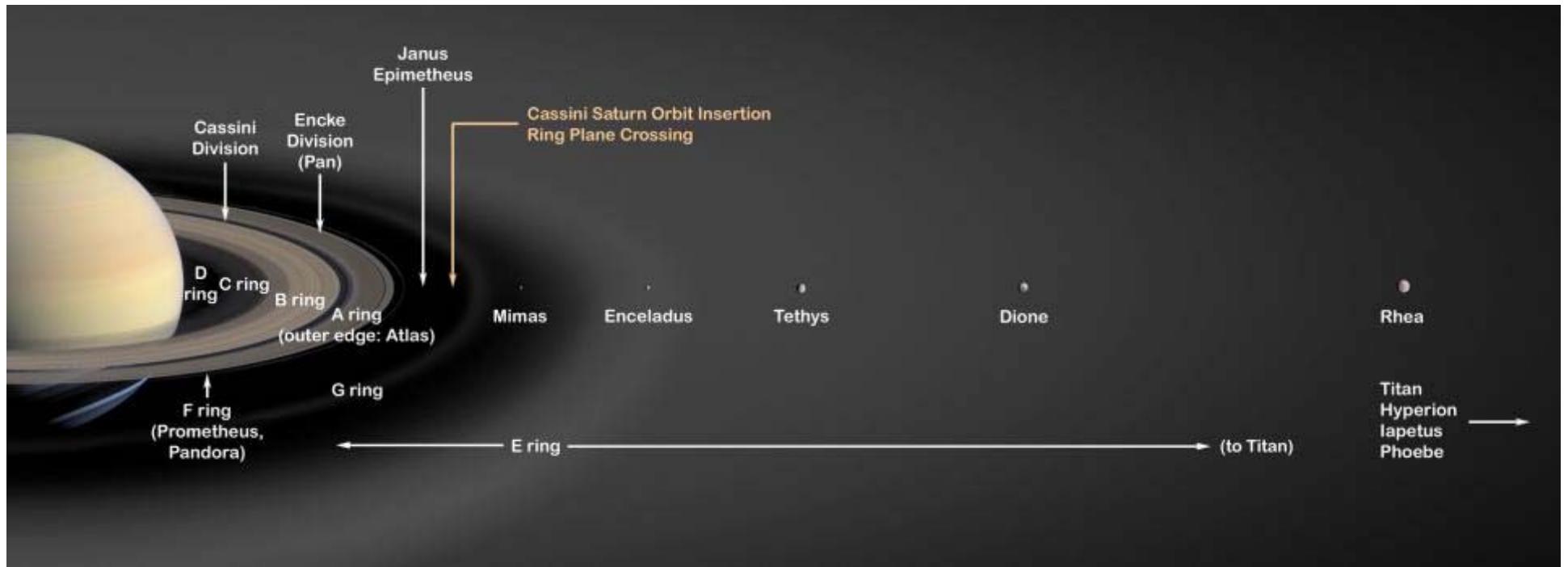




- Brightness of Saturn's rings indicates that they're young
 - <100 million years old
 - Probably due to the destruction of a small moon
- Rings are short-lived
 - Collisions grind down the ring particles to dust
 - Probably many sets of rings that have come and gone over the age of the solar system
- What a stroke of luck for us....



- We can divide Saturn's system into three main parts...
 - The A-D ring zone
 - ▶ Ring gaps and shepherd moons
 - The E ring zone
 - ▶ Ring supplies by Enceladus
 - ▶ Tethys, Dione and Rhea have a lot of similarities
 - The distant satellites Iapetus, Hyperion, Phoebe
 - ▶ Linked together by exchange of material



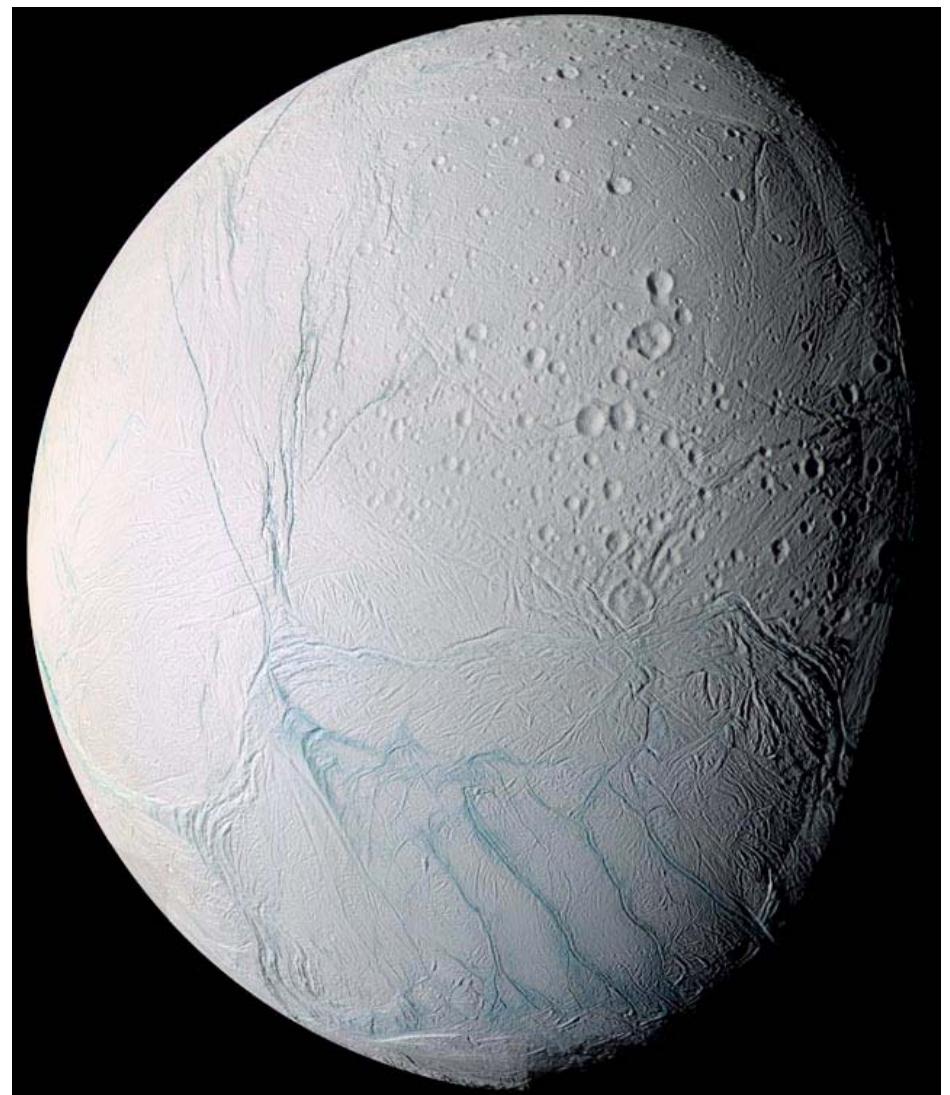


E ring Moons

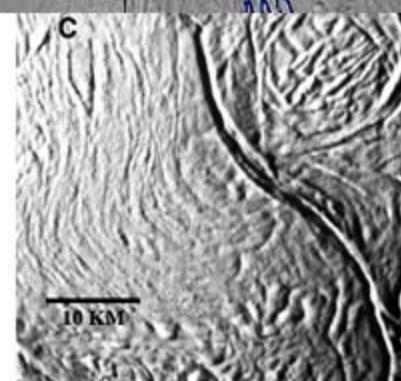
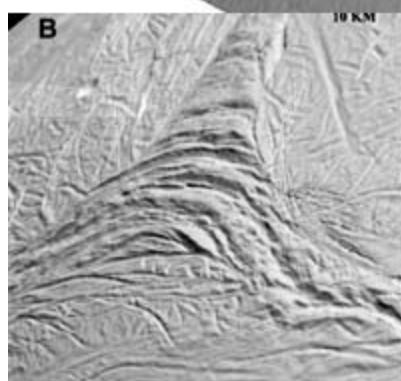
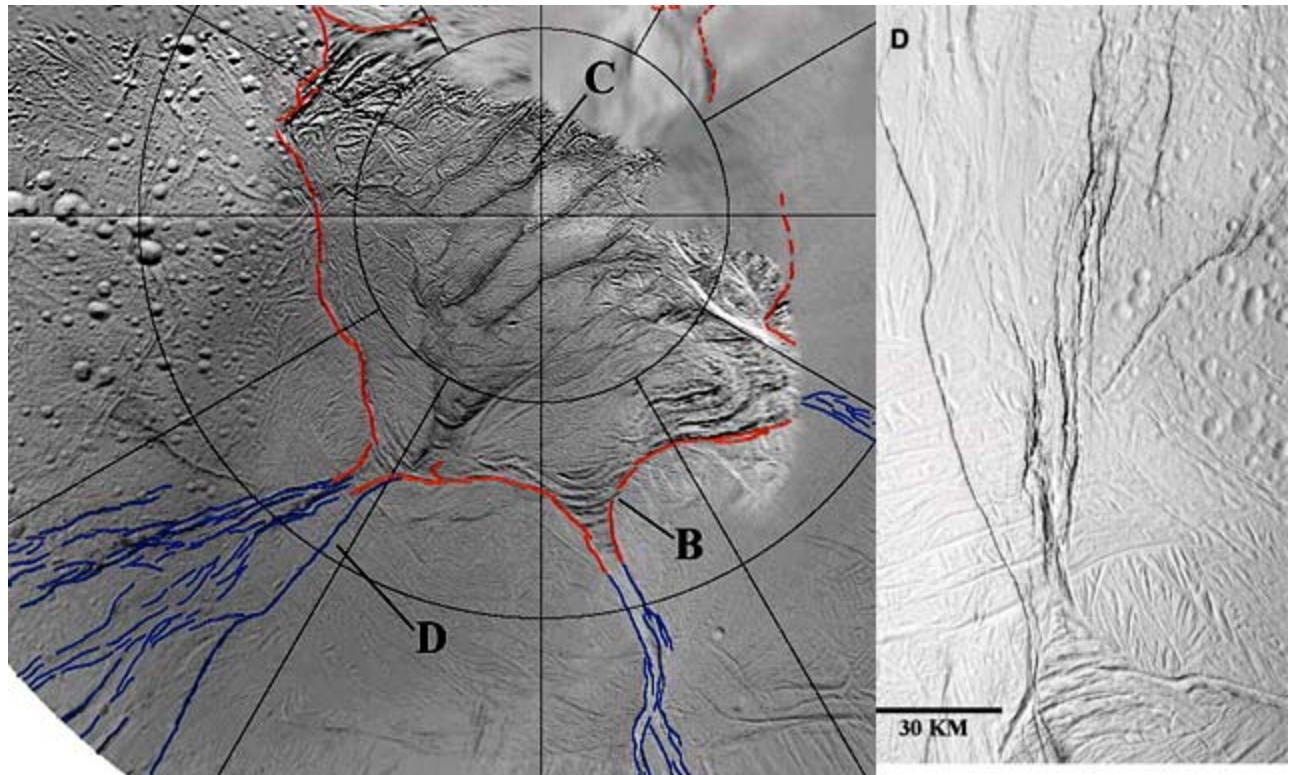
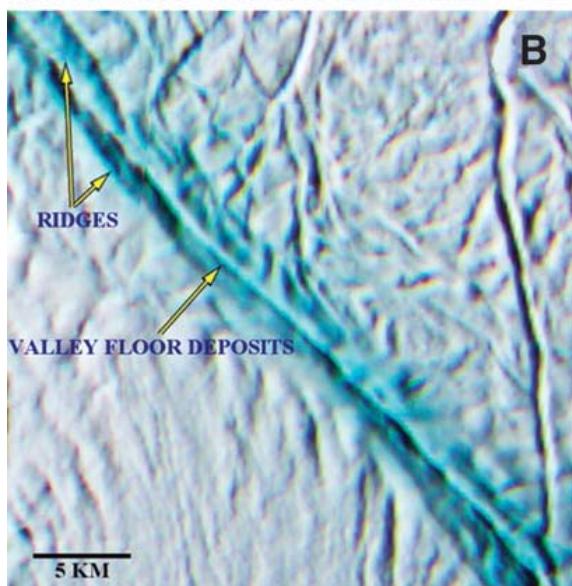
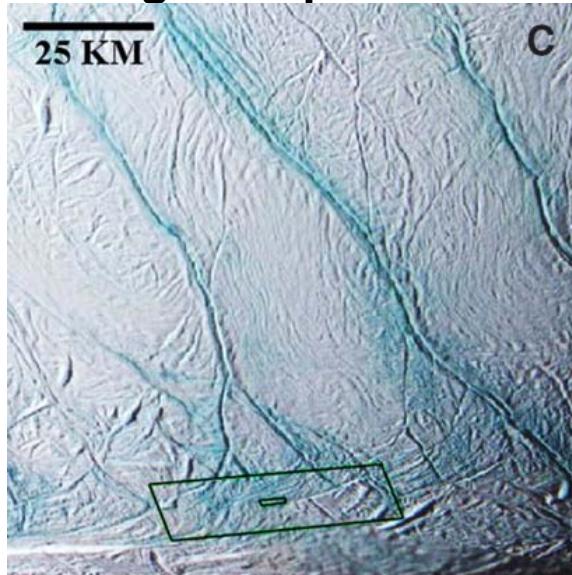
- Source of the E ring is Enceladus
 - Long suspected but now known from Cassini results

- Enceladus

- Brightest object in the solar system – albedo ~99%
 - Small – radius 250 km
 - i.e. not much internal heat expected
 - Old terrain – heavily cratered
 - Young terrain
 - Tectonically very active
 - Lots of extension in the south polar area

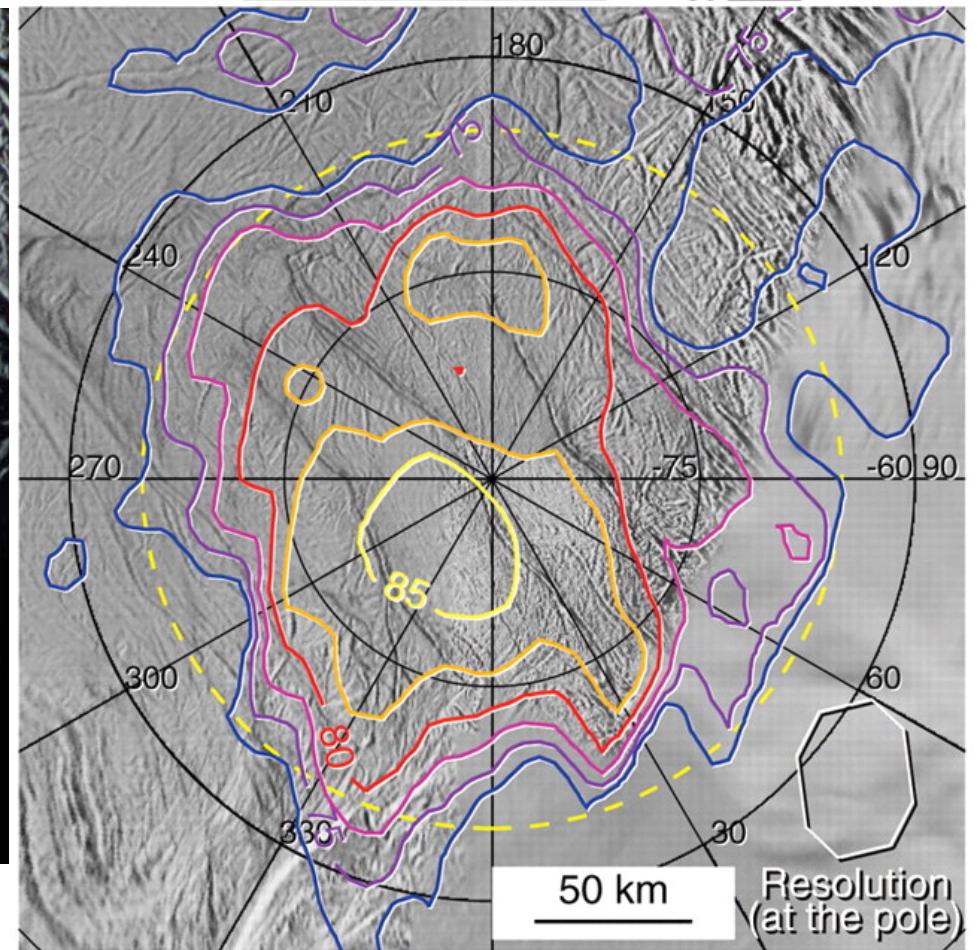


- South polar terrain
 - Tectonically confined
 - Tiger stripes



Porco et al., 2006

- Hot-spot over the south pole
 - Highest temperatures over the tiger stripes themselves





- Dust impact detector
 - Stream of particles emanating from south polar region
 - Explains the E ring
- Imaging in forward scattered light discovers plumes
- UVIS observation of a stellar occultation showed plume had H₂O composition
- Magnetometer results indicate liquid water below the surface

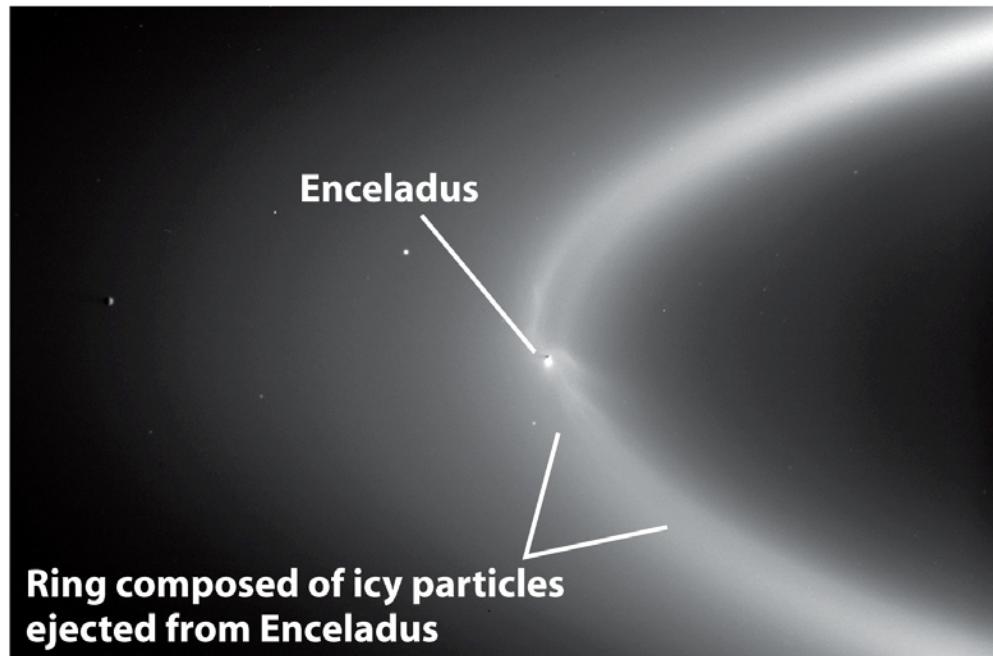
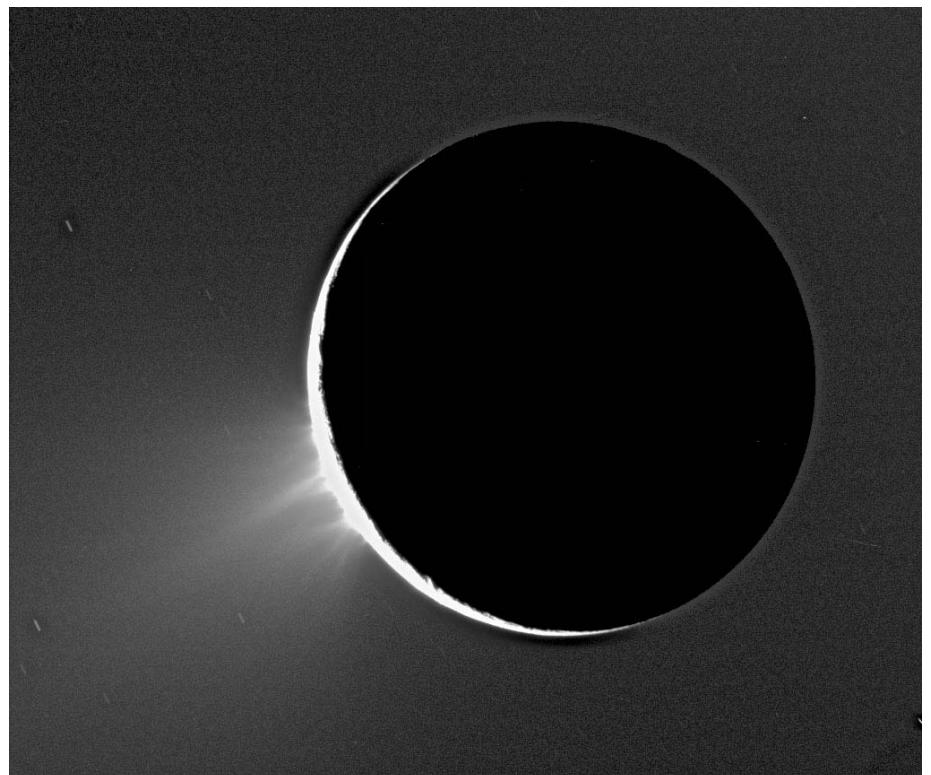
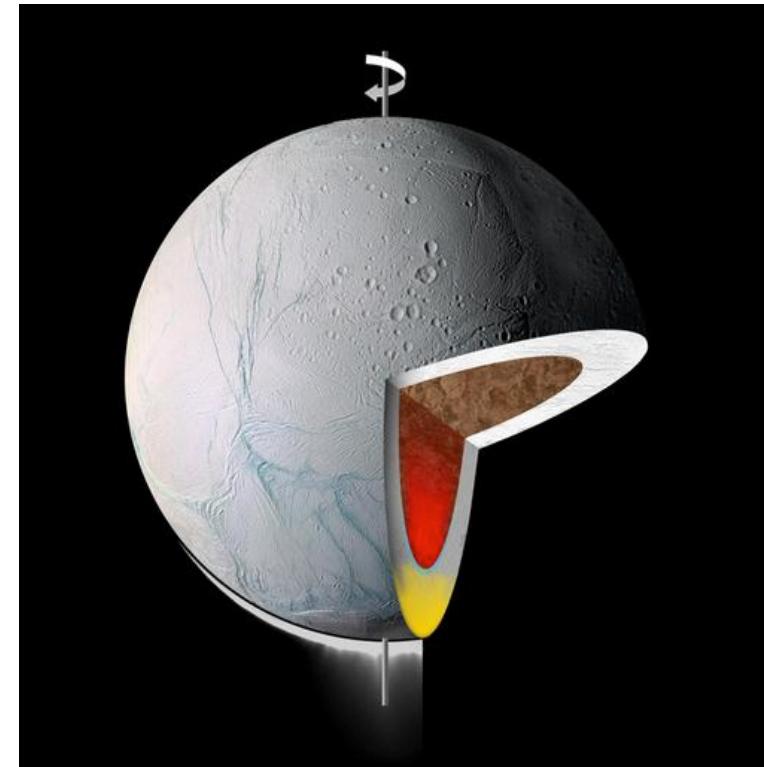
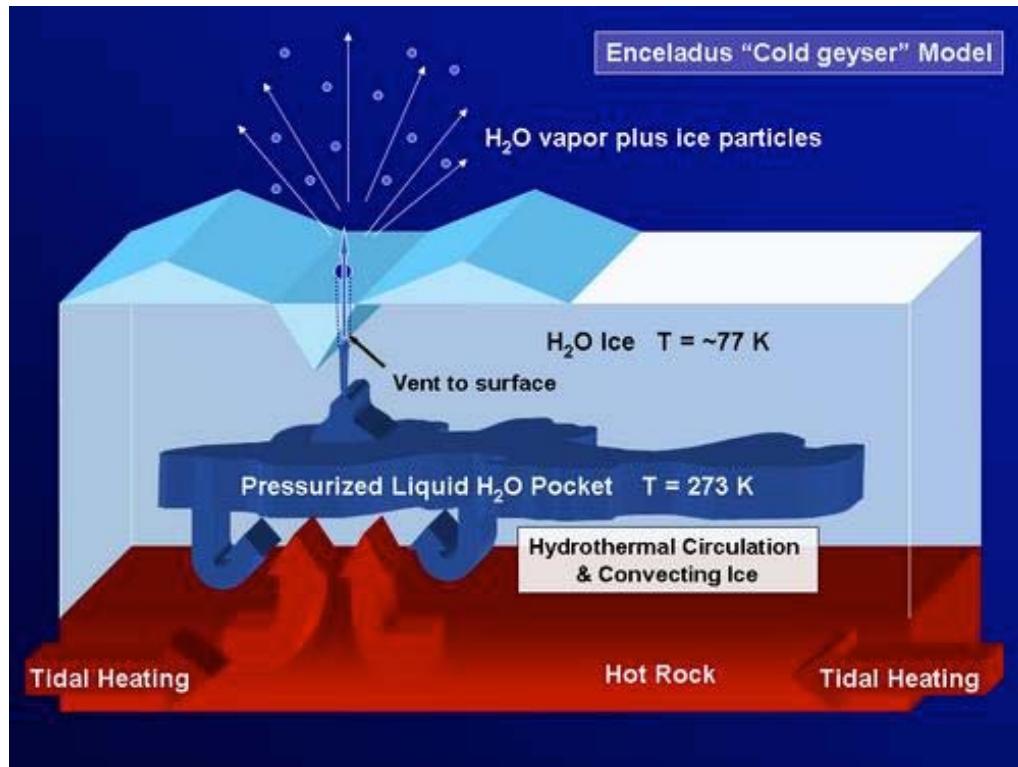


Figure 13-23b
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- Enceladus is in a 2:1 tidal resonance with Dione
 - Akin to Io's resonance with Europa
 - Tidal heating can only account for about 10% of the energy needed
 - Enceladus is still a bit of a mystery



- Other moons close to the E ring

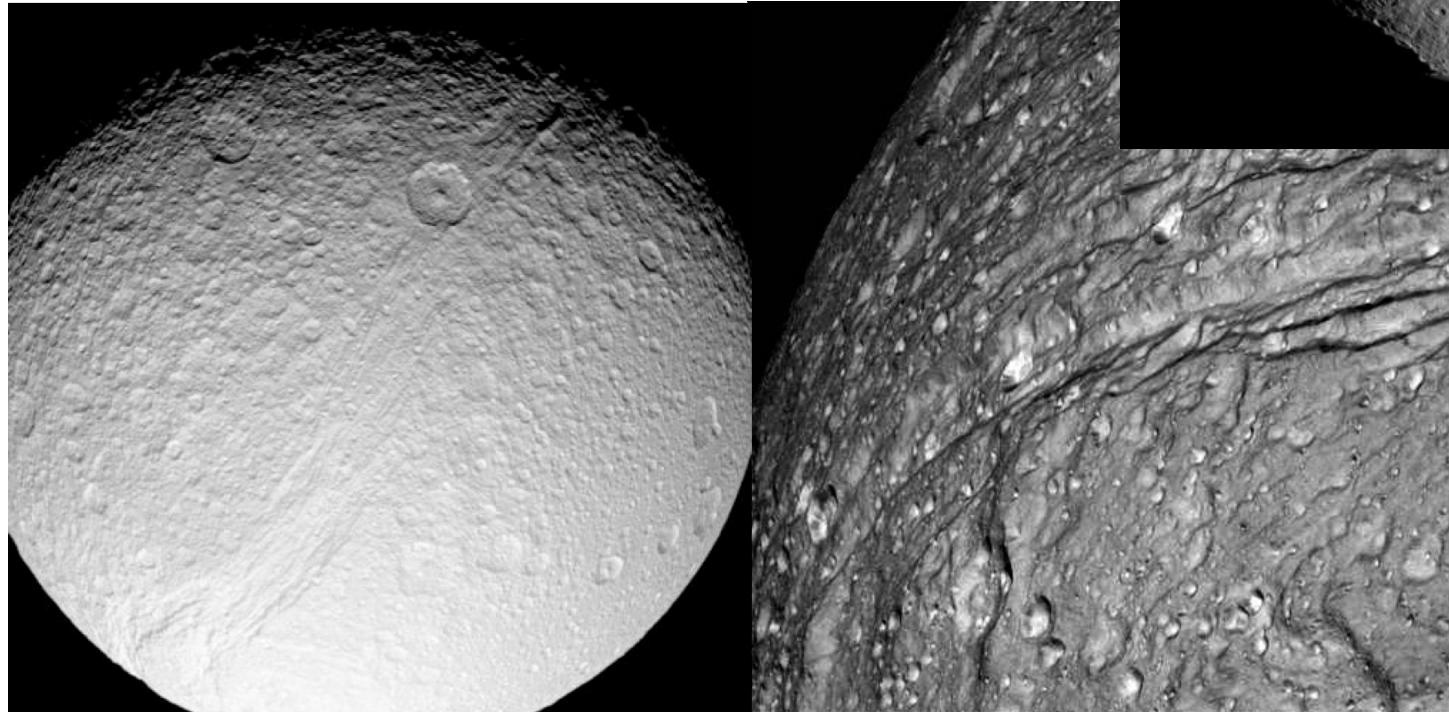
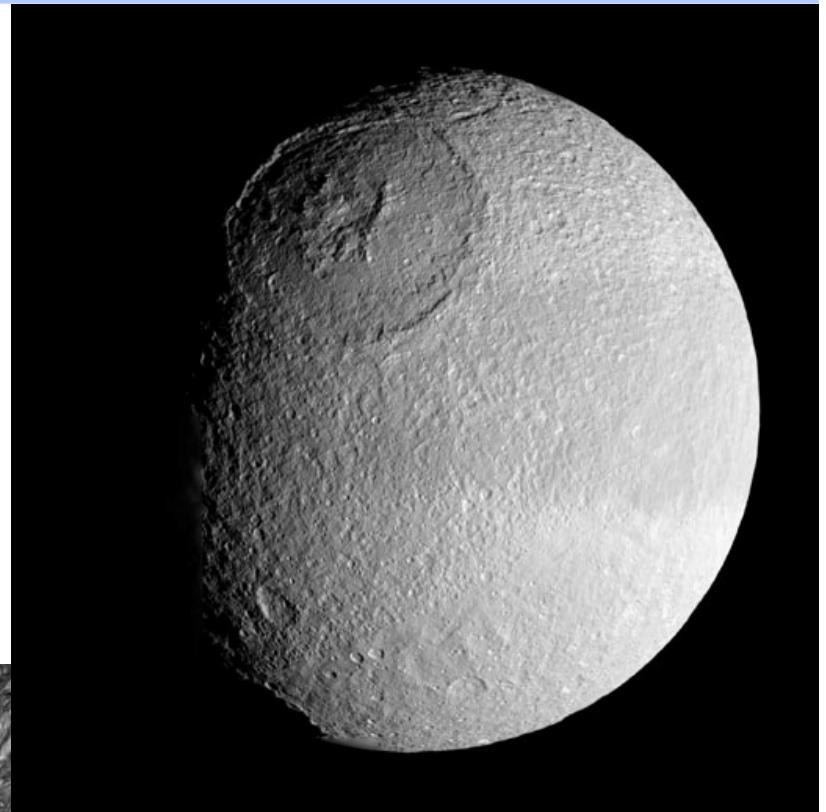
- Tethys
- Dione
- Rhea



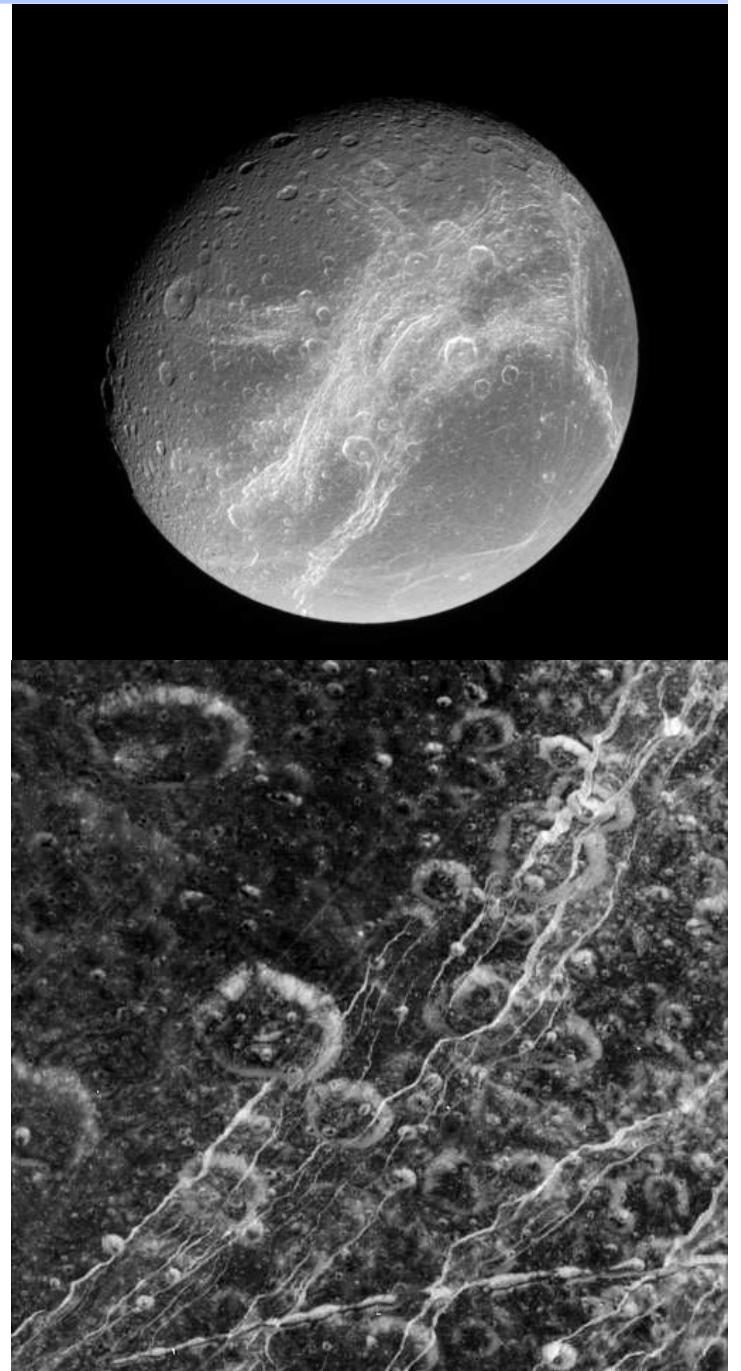
	Radius	Density
Tethys	530 km	970 kg m^{-3}
Dione	560 km	1480 kg m^{-3}
Rhea	764 km	1230 kg m^{-3}

- Smaller than the Galilean satellites of Jupiter
 - Probably no sub-surface oceans here
- Less dense & very bright – dominated by water ice composition
- All heavily cratered with little geologic activity for billions of years
- Tidally locked to Saturn – like our moon is to Earth

- Tethys has two major features
 - Odysseus
 - 400km impact basin
 - Ithaca Chasma
 - A pole-to-pole canyon created by stretching the crust



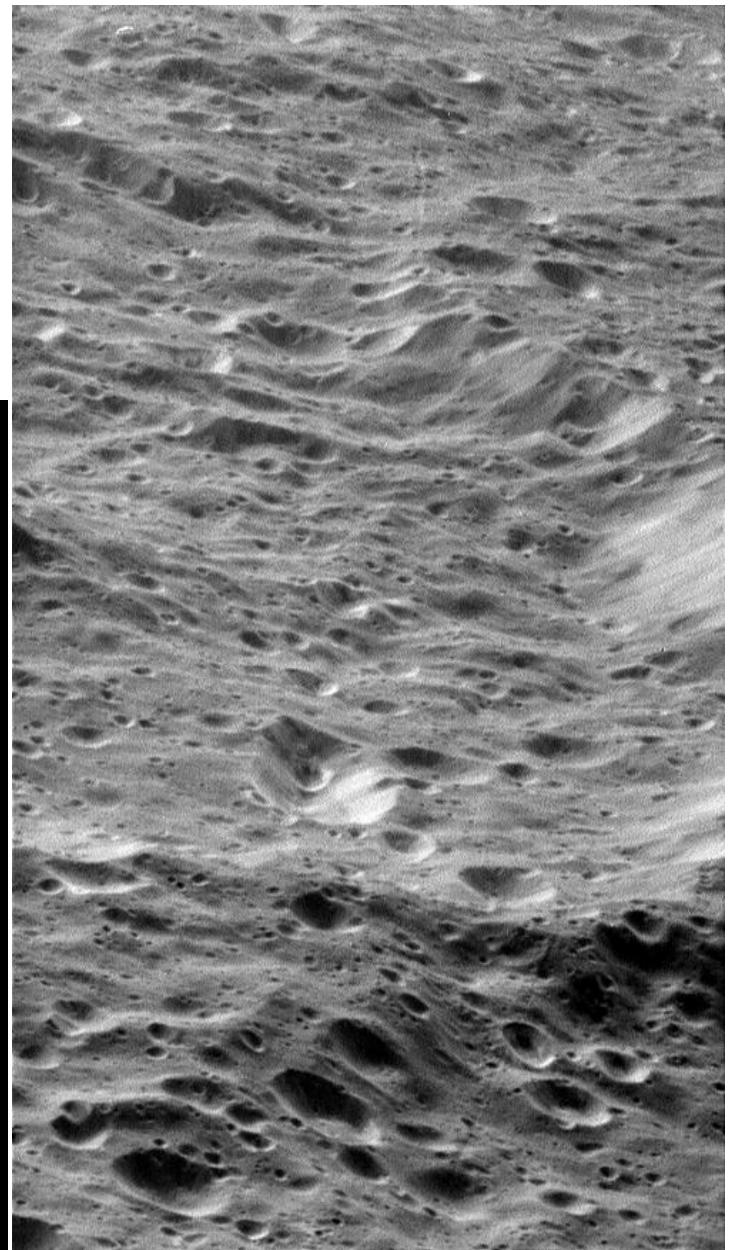
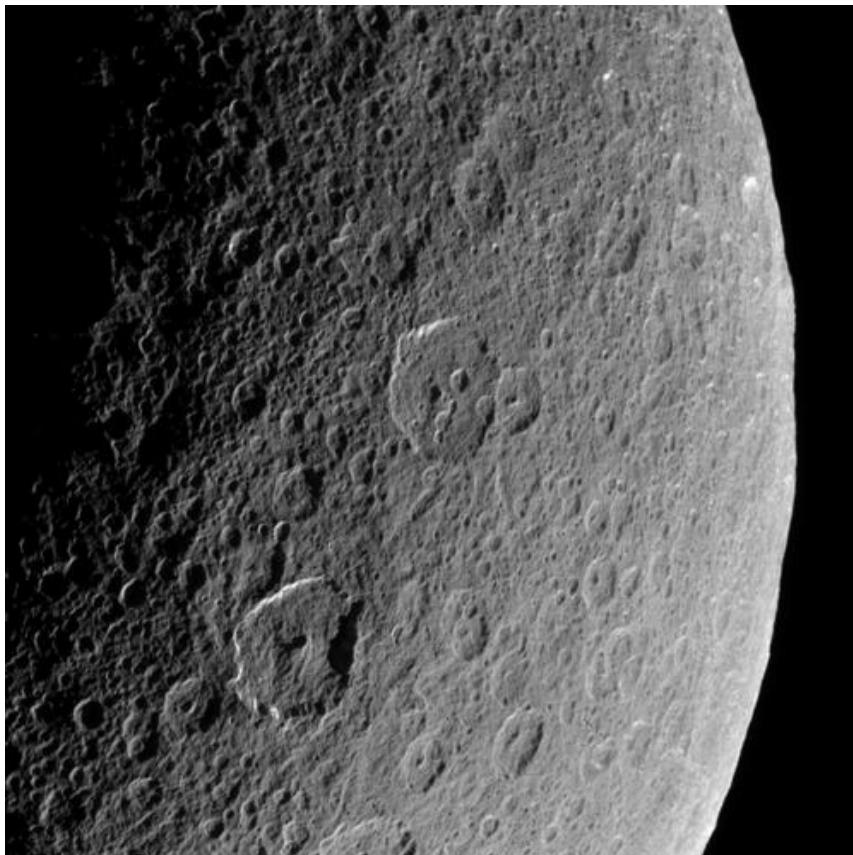
- **Dione**
 - Density 1430 kg m^{-3} – $\frac{1}{3}$ rocky core
- **Bright wispy features**
 - Tectonic – many sun-facing scarps
- **More craters on the trailing hemisphere**
 - Opposite to that expected
 - Leading hemisphere should collect more impacts
 - Dione can be spun by craters only 35km in diameter





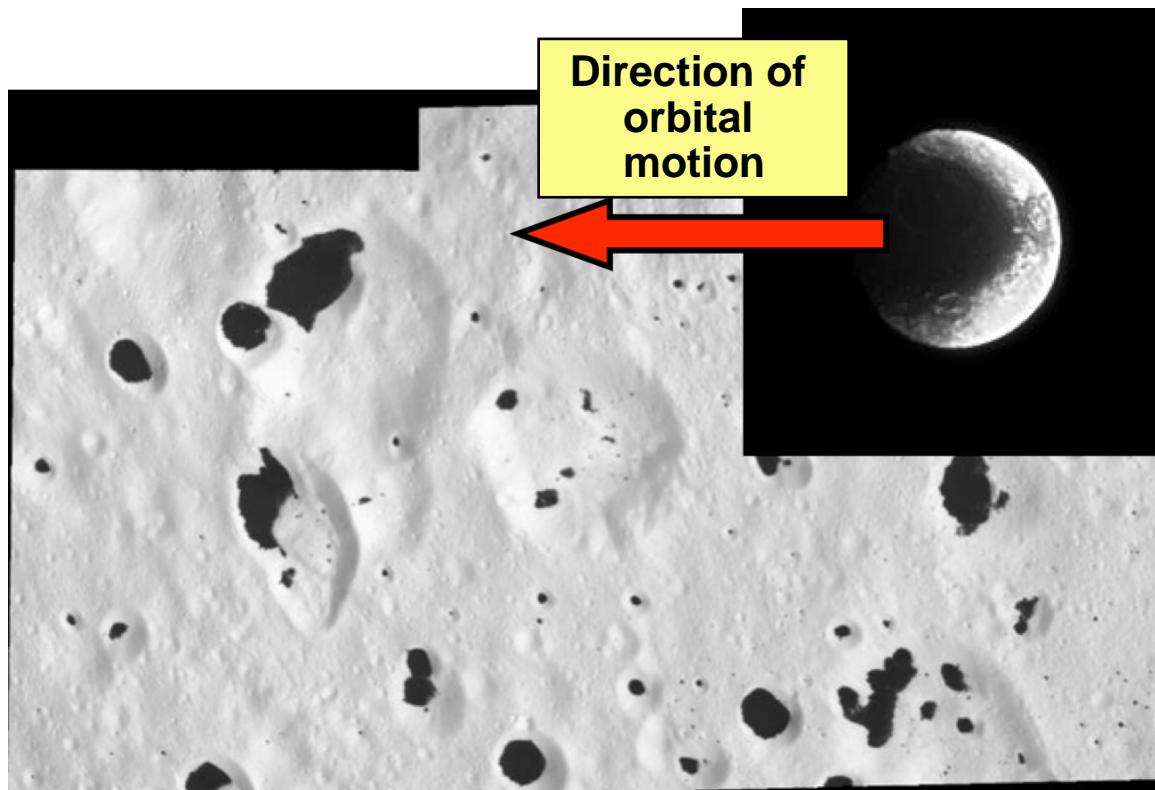
- **Rhea**

- Larger than Tethys/Dione
- Further from Saturn
 - Less tidal heating
- No surface activity
- Interior probably undifferentiated
 - Cassini gravity results

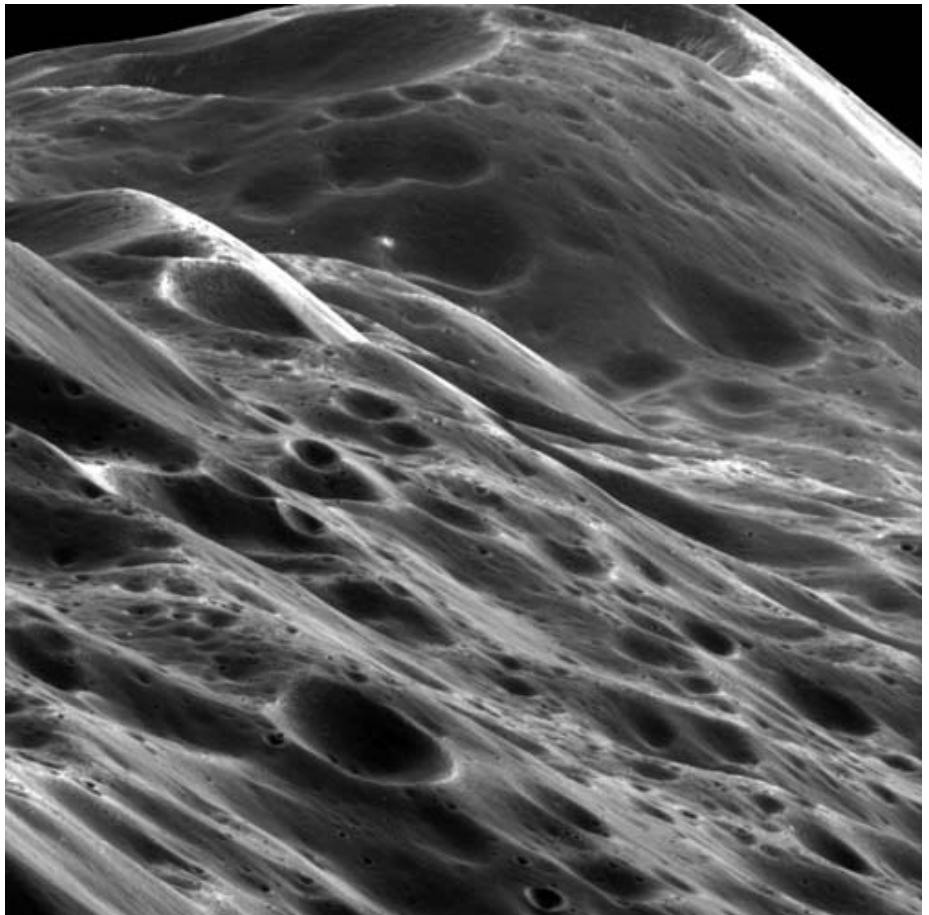
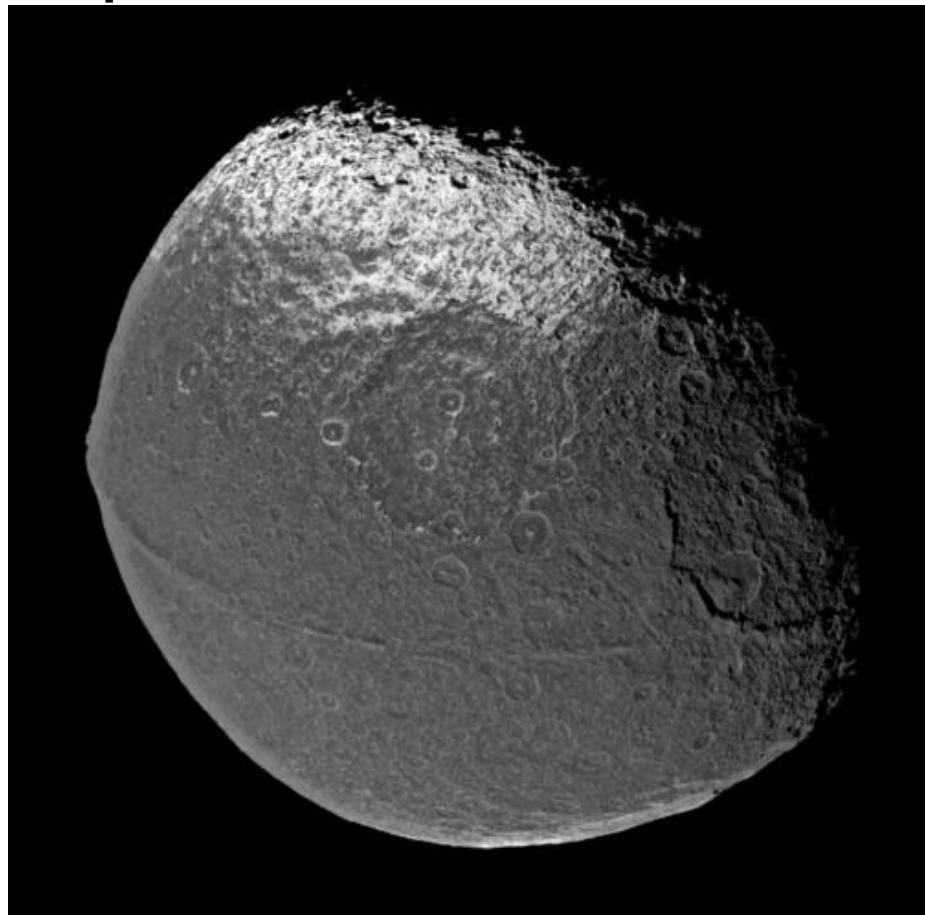
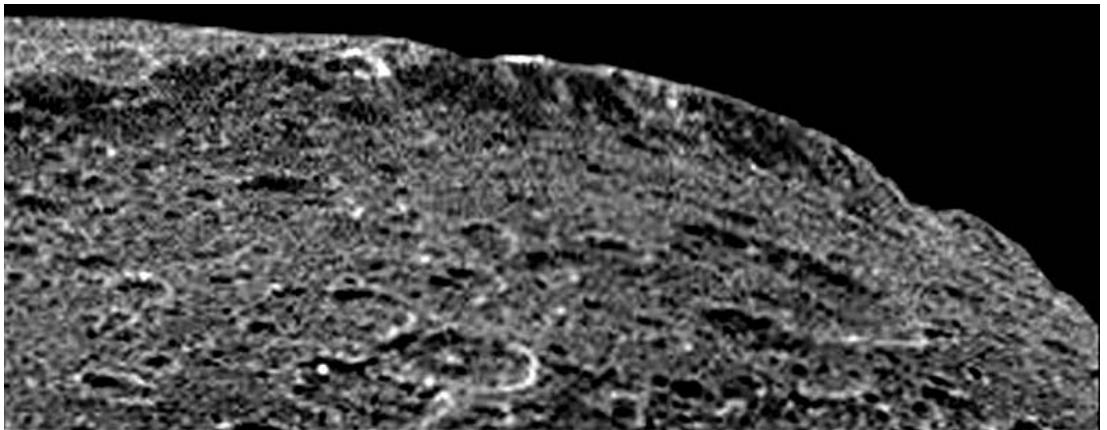


Moons beyond the E ring

- Iapetus: Two-toned moon
 - Dark organic material ~ 3-4% albedo
 - Bright icy material ~ 50% albedo
 - Radius ~ 720km
- Dark material concentrated on the leading hemisphere
 - Some in floors of dark terrain craters
- Density ~ 1034 kg m^{-3} indicates icy bulk composition

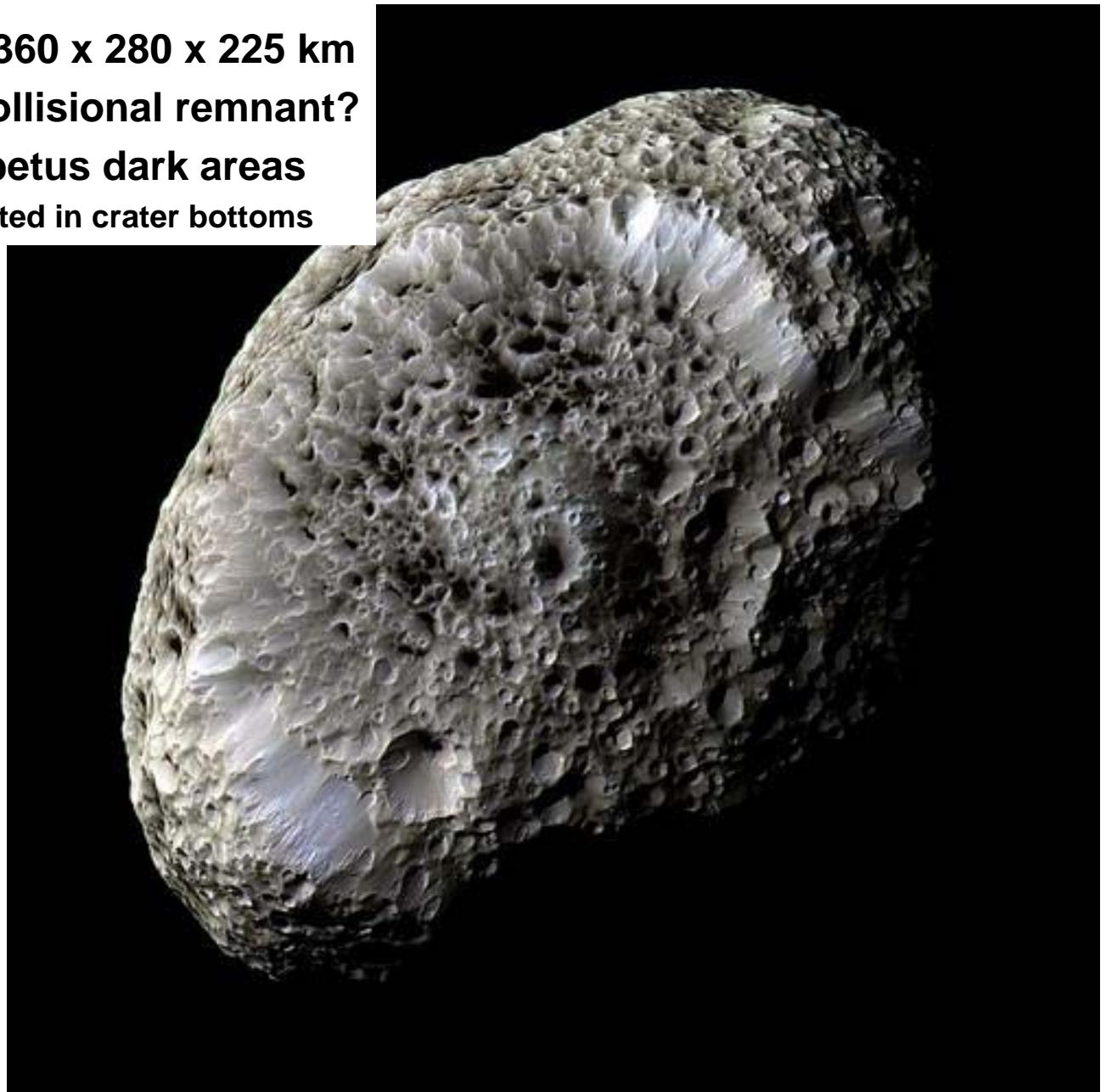


- Ancient equatorial bulge
 - Heavily cratered
 - Covered with dark material
 - Up to 20km in height
- Some sort of change in spin?



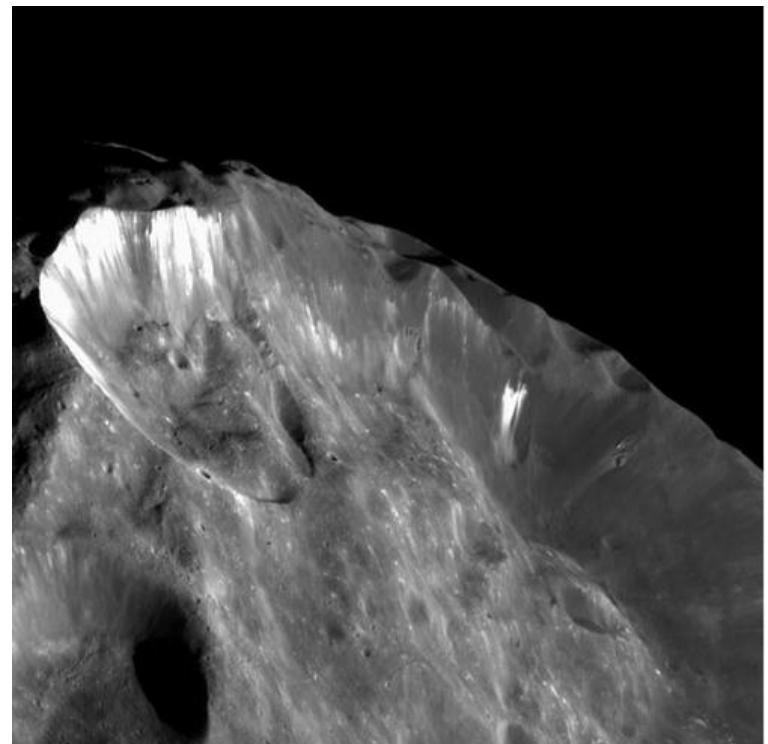
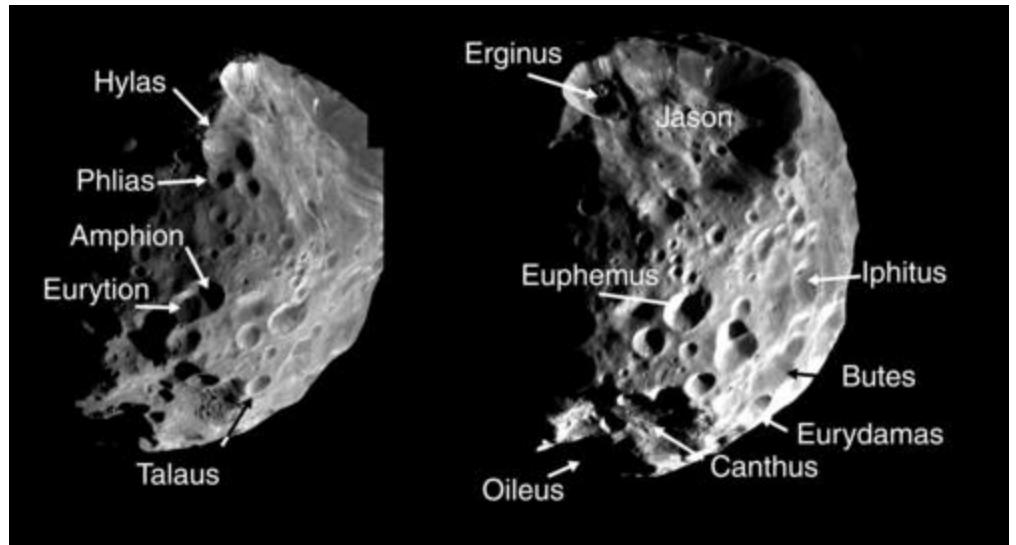
- Hyperion

- Equatorial diameter: $360 \times 280 \times 225$ km
- Chaotic tumbling – collisional remnant?
- Reddish color like Iapetus dark areas
 - ▶ Dark material concentrated in crater bottoms

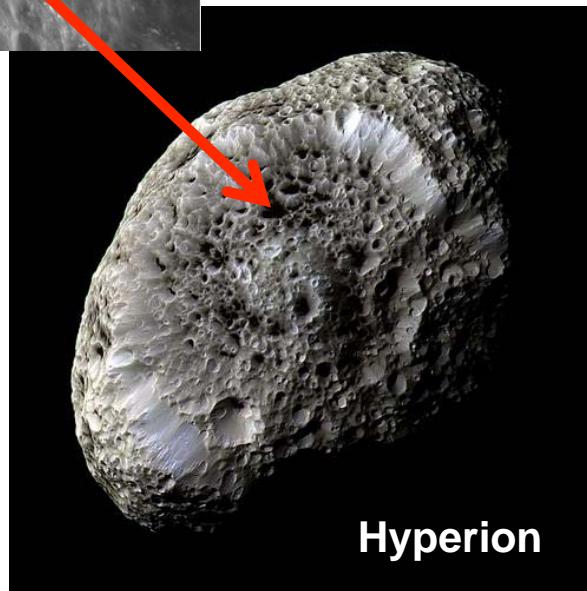
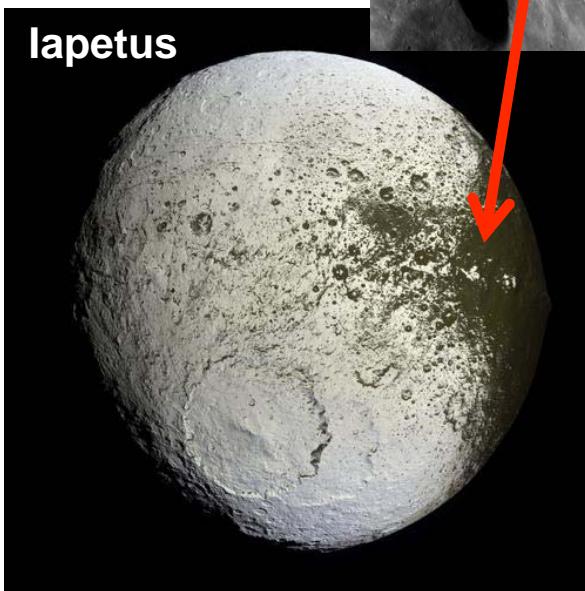
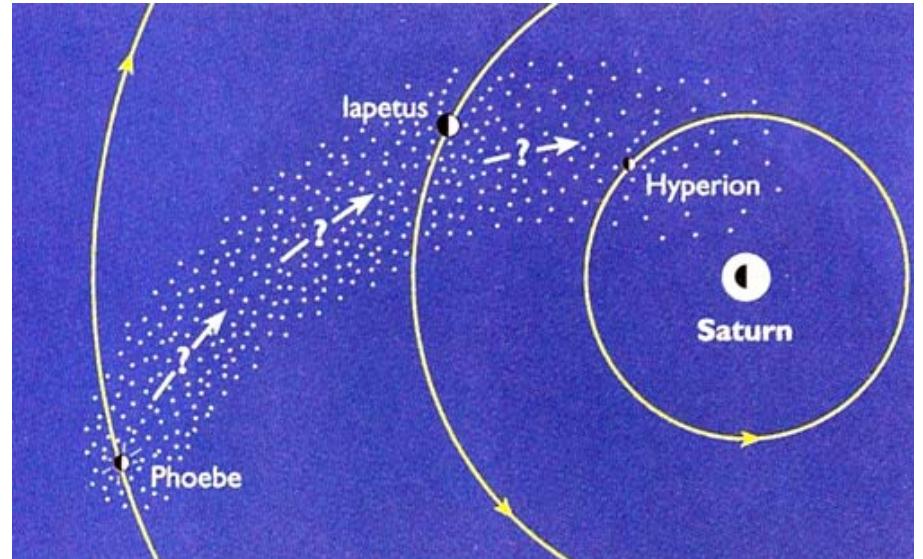
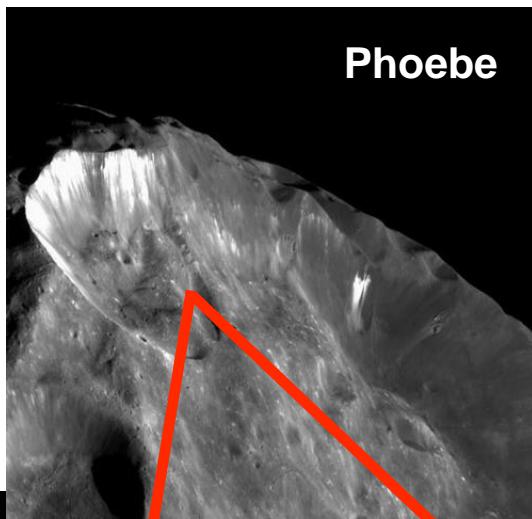


Phoebe

- Retrograde eccentric orbit
 - Implies a captured body
 - Radius ~110 km
- Compositionally distinct
 - Spectrally close to C-type asteroid
 - Minimum density ~ 1600 kg m^{-3}
 - ...but CO₂ detection implies outer solar system origin
- Albedos 7-30%
 - Ice outcrops are not that clean
 - Ice is shallow – but darkens quickly
- Small impacts can eject dark stuff
 - Escape velocity is only ~1 ms⁻¹



- Phoebe is the odd one out here
 - Dark rocky object among bright icy ones
 - Probably supplying the dark material that coats Iapetus and fills craters on Hyperion





In this lecture...

- **Rings**
 - A planar swarm of icy particles
 - How to form rings
 - ▶ The Roche limit
 - Dynamics
 - ▶ Gaps and resonances
 - ▶ Shepherd moons
- **Inner E-ring moons**
 - Mid-size icy satellites: Tectonics and craters
 - Enceladus – water geysers from subsurface liquid!
- **Outer Moons**
 - Captured Phoebe
 - Iapetus and Hyperion
 - ▶ Spray-painted with Phoebe debris

Next: Titan

- **Reading**
 - Chapter 12 & 13 to revise this lecture
 - Chapter 13, part 8 for next lecture