

## ABSTRACT ORAL CONTRIBUTION

### KINEMATICS AND TOTAL MASS OF THE LOCAL GROUP

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Recent kinematic studies of the Local Group show some inconsistency with one and other. Some observations suggest the tangential velocity of M31 with respect to the Milky Way is around  $v_{tan} \sim 30 \text{ km s}^{-1}$  while other studies claim values of  $v_{tan} \sim 100 \text{ km s}^{-1}$ . This leads to the question of what is the most likely kinematic configuration for the Local Group. In this work we use cosmological simulations in order to find the most common kinematic configurations in a dark matter dominated Universe. Our results for the kinematics are presented as a function of different cosmological parameters and prior values for the LG total mass. In turn, assuming the most accepted constraints for the LG total mass and the cosmological parameters we find that the observed LG has atypical kinematics once it is considered in a broad cosmological context.

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