

last time $\Omega_{DM} \approx 0.24$

DM timeline

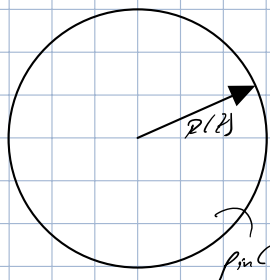
Ryden Ch 12

consider $\rho_{DM} \propto a^{-3}$ always

$$\rho \equiv \rho_{DM}$$

$$S \equiv \text{"Density Contrast"} = \frac{\rho - \bar{\rho}}{\bar{\rho}} \quad \begin{matrix} \nearrow \\ \text{at } \vec{x}, t \end{matrix} \quad \nwarrow \text{avg DM mass density}$$

only talk about linear term w/ small perturbation $|S| \ll 1$
 $\rightarrow O(S^2) \rightarrow 0$



$$M = \frac{4\pi}{3} \bar{\rho} R_0^3 = \text{constant}$$

Newton's 2nd law for $R(t)$

$$\ddot{R}(t) = -\frac{G}{R^2} (\Delta M) = -\frac{G}{R^2} \left(\frac{4\pi}{3} \bar{\rho} \cdot \delta \right) R^3$$

ex 1: no expansion, $\bar{\rho} = \text{const}$

$$\frac{\ddot{R}}{R} = -\frac{4\pi}{3} G \bar{\rho} \delta$$

mass conservation: $M = \frac{4\pi}{3} \rho(t) R(t)^3 = \text{const}$

$$\rightarrow R(t)^3 = \frac{K}{\rho(t)}$$

$$R(t) = \frac{K^{1/3}}{\bar{\rho}^{1/3} (1 + \delta(t))^{1/3}}$$

$$= \left(\frac{K}{\bar{\rho}} \right)^{1/3} (1 + \delta(t))^{-1/3}$$

$$= \left(\frac{K}{\bar{\rho}} \right)^{1/3} \left(1 - \frac{1}{3} \delta(t) + O(\delta^2) + \dots \right)$$

$$R(t) = R_0 \left(1 - \frac{1}{3} \delta(t) \right)$$

$$\ddot{R}(t) = -\frac{1}{3} R_0 \ddot{\delta}(t)$$

$$\ddot{R}(t) = -\frac{1}{3} R_0 \dot{\delta}(t)$$

$$\approx -\frac{1}{3} R \dot{\delta}(t) + O(\delta^2) + \dots$$

$$\frac{\ddot{R}(t)}{R(t)} = -\frac{1}{3} \ddot{\delta}(t)$$

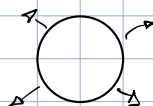
$$-\frac{1}{3} \ddot{\delta}(t) = -\frac{4\pi}{3} G \bar{\rho} \delta(t)$$

$$\star \ddot{\delta}(t) = 4\pi G \bar{\rho} \delta(t)$$

if $\delta < 0 \rightarrow$ collapse
 $\ddot{R} < 0$



if $\delta > 0 \rightarrow$ expand/dissipate
 $\ddot{R} > 0$



general solⁿ for $\delta(t)$?

$$\ddot{\delta} = 4\pi G \bar{\rho} \delta$$

$$= \left(\frac{1}{t_{\text{dyn}}^2}\right) \delta$$

$$t_{\text{dyn}} = \left(\frac{1}{4\pi G \bar{\rho}}\right)^{1/2}$$

$$\rightarrow \delta(t) = A e^{\frac{t}{t_{\text{dyn}}}} + B e^{-\frac{t}{t_{\text{dyn}}}}$$

grow decay

matter has pressure when collapsing, keeping air molecules from clumping together

Jeans Scale λ_J

pressure: $P = f(\rho)$ energy density
↘ equation of state

sound speed: $\underline{c_s}^2 = \frac{dP}{d\rho}$ (material specific)

When is pressure imp^t?

$$t_{\text{press}} \sim \frac{R}{\underline{c_s}}$$

size of system

Jeans length λ_J : $t_{\text{press}} = t_{\text{dyn}}$

$$\frac{\lambda_J}{c_s} = \frac{1}{(4\pi G \bar{\rho})^{1/2}}$$

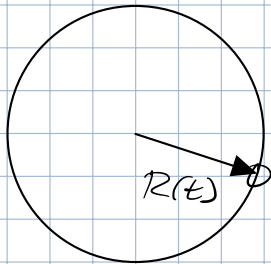
$$\lambda_J = \frac{c_s}{(4\pi G \bar{\rho})^{1/2}}$$

critical value for collapse or survival

if $R > \lambda_J \rightarrow$ collapse

if $R < \lambda_J \rightarrow$

ex 2: expansion! over/under dense sphere



what is acceleration of test mass?

$$\rho(t) = \underbrace{\bar{\rho}(t)}_{\propto a^{-3}} (1 + \delta(t))$$

$$\ddot{R} = -\frac{G}{R^2} M = -\frac{G}{R^2} \left(\frac{4\pi}{3} \rho R^3 \right) \\ = -G \frac{4\pi}{3} \rho(t) R(t)$$

$$\frac{\ddot{R}}{R} = -G \frac{4\pi}{3} \rho(t)$$

$$\star \left(\frac{\ddot{R}}{R} \right) = -G \frac{4\pi}{3} \bar{\rho}(t) (1 + \delta(t))$$

mass conservation: $M = \frac{4\pi}{3} \rho(t) R^3 = \frac{4\pi}{3} \underbrace{\bar{\rho}(t) (1 + \delta(t)) R(t)^3}_{\propto K} = \text{const} \cdot K$

$$R^3 = \frac{K}{\bar{\rho}(1 + \delta)}$$

$$R = \frac{K^{1/3}}{\bar{\rho}^{1/3} (1 + \delta)^{1/3}}$$

$$\bar{\rho} \propto a^{-3}$$

$$\bar{\rho}^{1/3} \propto a^{-1}$$

$$\bar{\rho}^{1/3} \propto a^{-1} \quad \bar{\rho}^{1/3} = \alpha \cdot a(t)^{-1} \quad K' \equiv \frac{K}{\alpha}$$

$$R = \frac{K'}{a^{-1} (1 + \delta)^{1/3}}$$

$$R(t) = K' a(t) (1 + \delta(t))^{-1/3}$$

expand to $O(\delta)$

$$R(t) = K' a (1 - \frac{1}{3} \delta)$$

$$\dot{R}(t) = K' \left(\dot{a} (1 - \frac{1}{3} \delta) + a (0 - \frac{1}{3} \dot{\delta}) \right)$$

$$\ddot{R}(t) = K' \left(\ddot{a} (1 - \frac{1}{3} \delta) + \dot{a} (0 - \frac{1}{3} \dot{\delta}) + \dot{a} (-\frac{1}{3} \dot{\delta}) + a (-\frac{1}{3} \ddot{\delta}) \right)$$

$$= K' \left(\ddot{a} (1 - \frac{1}{3} \delta) - \frac{2}{3} \dot{a} \dot{\delta} - \frac{1}{3} a \ddot{\delta} \right)$$

$$\frac{\ddot{R}}{R} = \frac{K' \left(\ddot{a} (1 - \frac{1}{3} \delta) - \frac{2}{3} \dot{a} \dot{\delta} - \frac{1}{3} a \ddot{\delta} \right)}{K' a (1 - \frac{1}{3} \delta)}$$

$$= \frac{1}{a (1 - \frac{1}{3} \delta)} \left(\ddot{a} (1 - \frac{1}{3} \delta) - \frac{2}{3} \dot{a} \dot{\delta} - \frac{1}{3} a \ddot{\delta} \right)$$

$$\frac{1}{a (1 - \frac{1}{3} \delta)} \approx \frac{1}{a} (1 + \frac{1}{3} \delta)$$

$$\frac{\ddot{R}}{R} = \frac{(1+\frac{1}{3}\delta)}{a} \left(\ddot{a} (1-\frac{1}{3}\delta) - \frac{2}{3} \dot{a} \dot{\delta} - \frac{1}{3} a \ddot{\delta} \right)$$

drop all δ^2 , $\dot{\delta}^2$, $\delta \cdot \dot{\delta}$, ...

$$\frac{\ddot{R}}{R} = \frac{1}{a} \left(\cancel{\frac{1}{3} \delta \ddot{a}} + \ddot{a} (1 - \cancel{\frac{1}{3} \delta}) - \frac{2}{3} \dot{a} \dot{\delta} - \frac{1}{3} a \ddot{\delta} \right)$$

$$\frac{\ddot{R}}{R} = \frac{\ddot{a}}{a} - \frac{2}{3} \left(\frac{\dot{a}}{a} \right) \dot{\delta} - \frac{1}{3} \ddot{\delta}$$

$$-G \frac{4\pi}{3} \bar{\rho} (1+\delta) = \frac{\ddot{a}}{a} - \frac{2}{3} H \dot{\delta} - \frac{1}{3} \ddot{\delta}$$

In matter domination $\bar{\rho} = \rho_{tot} \propto a^{-3}$

Friedmann: $H = \frac{\dot{a}}{a} = \left(\frac{8\pi}{3} G \rho_{tot} \right)^{1/2}$

$$\rho_{tot} = \rho_i a^{-3}$$

$$\frac{\dot{a}}{a} = \left(\frac{8\pi}{3} G \rho_i \right)^{1/2} a^{-3/2}$$

$$\dot{a} = \left(\frac{8\pi}{3} G \rho_i \right)^{1/2} a^{-1/2}$$

$$\frac{d}{dt}$$

$$\ddot{a} = \left(\frac{8\pi}{3} G \rho_i \right)^{1/2} \left(-\frac{1}{2} \right) a^{-3/2} \dot{a}$$

$$= \left(\frac{8\pi}{3} G \rho_i a^{-3} \right)^{1/2} \left(-\frac{1}{2} \dot{a} \right)$$

$$= \left(\frac{8\pi}{3} G \rho_{tot} \right)^{1/2} \left(-\frac{1}{2} \dot{a} \right)$$

$$\frac{\ddot{a}}{a} = \left(\frac{8\pi}{3} G \rho_{tot} \right)^{1/2} \left(-\frac{1}{2} \right) \left(\frac{\dot{a}}{a} \right)$$

$$= H \left(-\frac{1}{2} \right) H$$

$$= -\frac{1}{2} H^2$$

$$\frac{\ddot{a}}{a} = -\frac{1}{2} \frac{8\pi}{3} G \bar{\rho} = -\frac{4\pi}{3} G \bar{\rho}$$

sub in ① & get

$$\ddot{\delta} + 2H\dot{\delta} = \frac{4\pi}{3} G \bar{\rho} \delta$$

$$\frac{\ddot{a}}{a} - \frac{2}{3} H \dot{\delta} - \frac{1}{3} \ddot{\delta} = -G \frac{4\pi}{3} \bar{\rho} (1+\delta)$$

$$-\cancel{\frac{4\pi}{3} G \bar{\rho}} - \frac{2}{3} H \dot{\delta} - \frac{1}{3} \ddot{\delta} = -\cancel{G \frac{4\pi}{3} \bar{\rho}} - G \frac{4\pi}{3} \bar{\rho} \delta$$

$$\frac{2}{3} H \dot{\delta} + \frac{1}{3} \ddot{\delta} = G \frac{4\pi}{3} \bar{\rho} \delta$$

↓ × 3

$$\ddot{\delta} + 2H\dot{\delta} = G4\pi\bar{\rho}\delta$$

General form: $\Omega_M = \bar{\rho} / \rho_{tot}$

$$\bar{\rho}_M = \Omega_M \rho_{tot}$$

$$H^2 = \frac{8\pi}{3} G \rho_{tot}$$

$$\rightarrow \rho_{tot} = \frac{3H^2}{8\pi G}$$

$$\bar{\rho}_M = \Omega_M \cdot \frac{3H^2}{8\pi G}$$

$$\ddot{\delta} + 2H\dot{\delta} = 4\pi G \left(\Omega_M \frac{3H^2}{8\pi G} \right) \delta$$

$$\ddot{\delta} + 2H\dot{\delta} = \frac{3}{2} G \Omega_M H^2 \delta$$