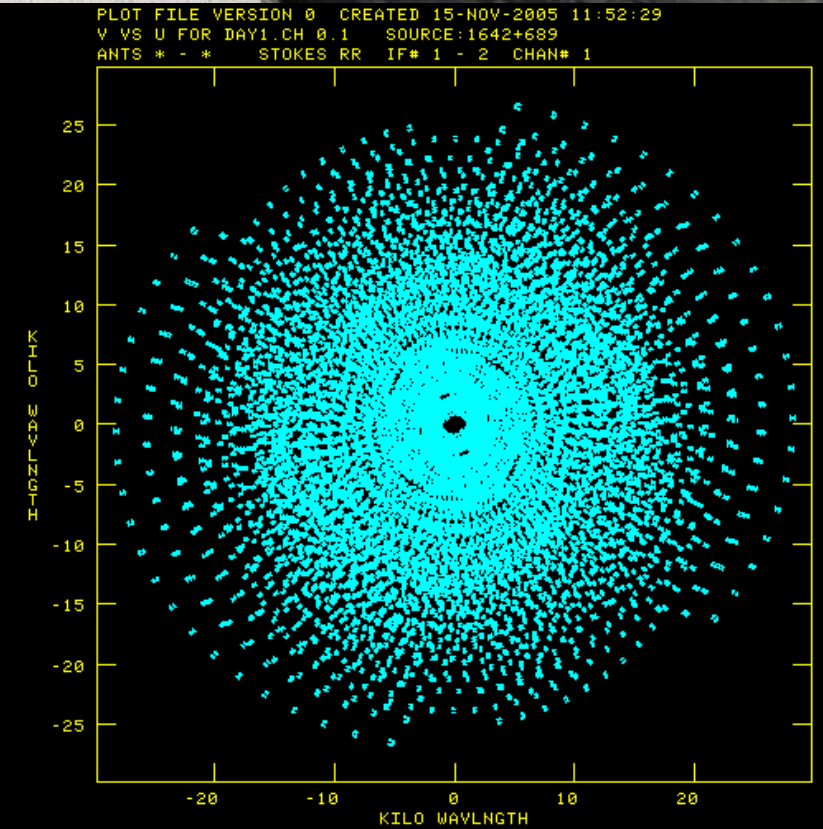




Cross correlators

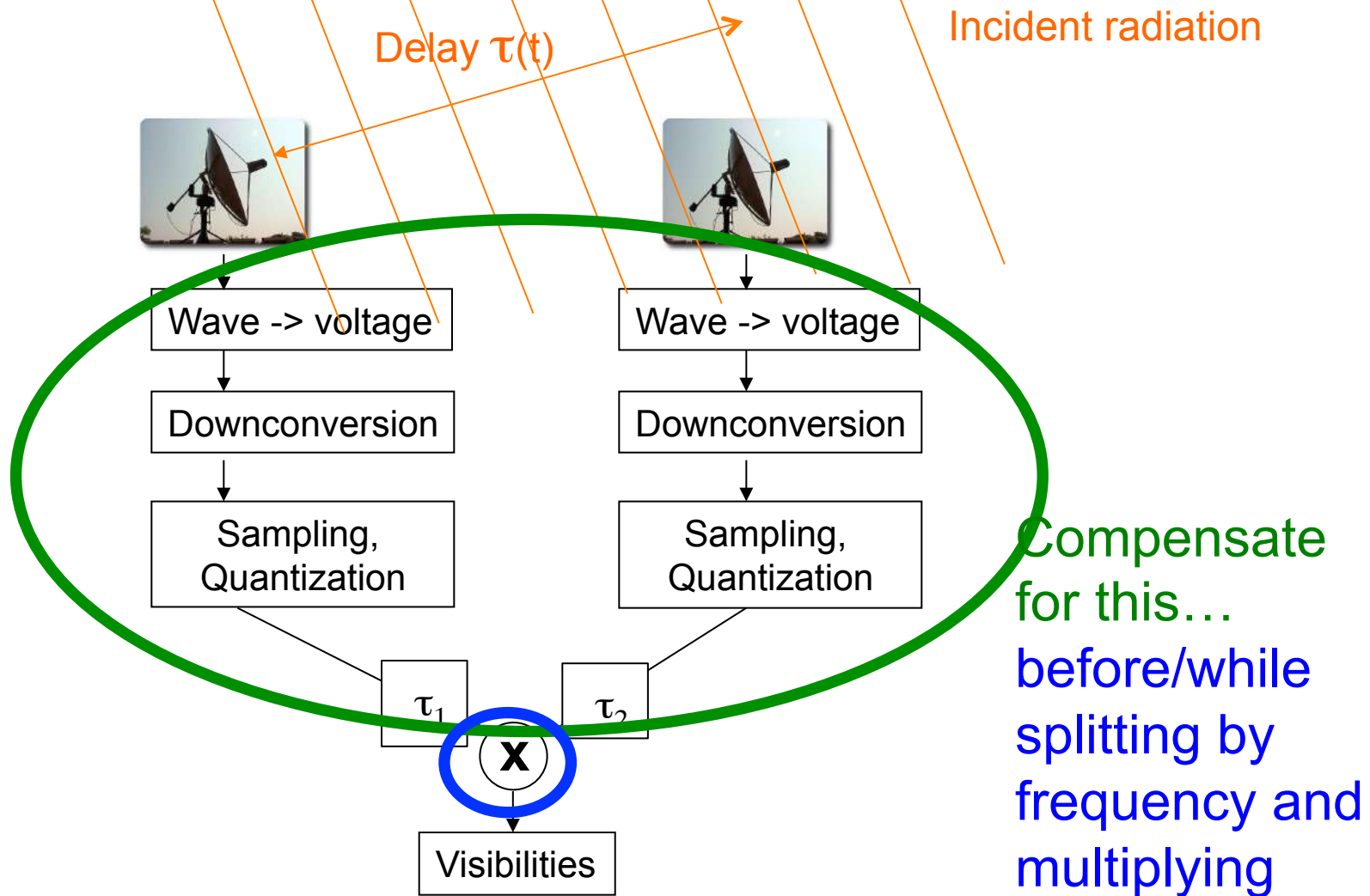
for radio astronomy



Adam Deller
May 16, 2018

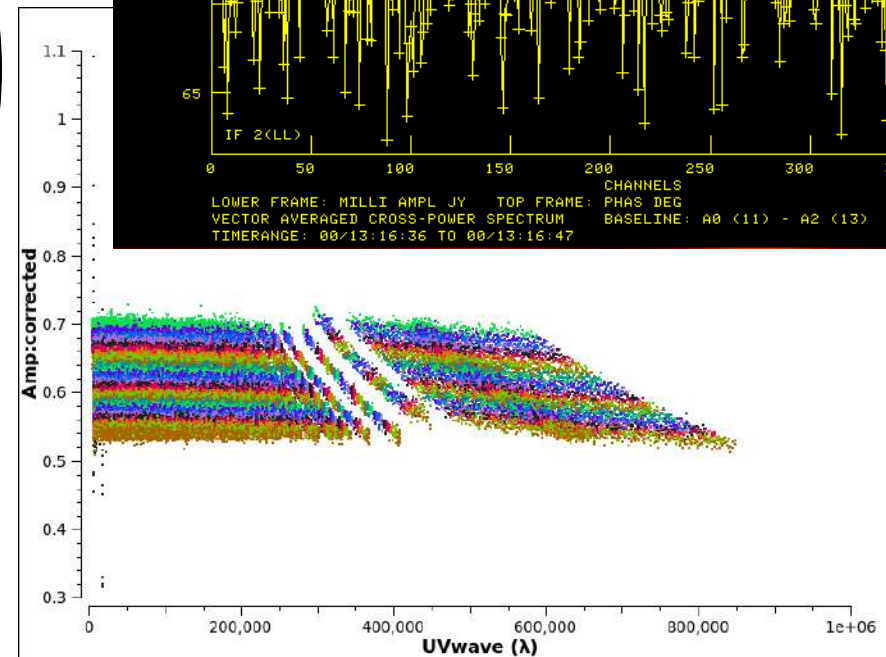
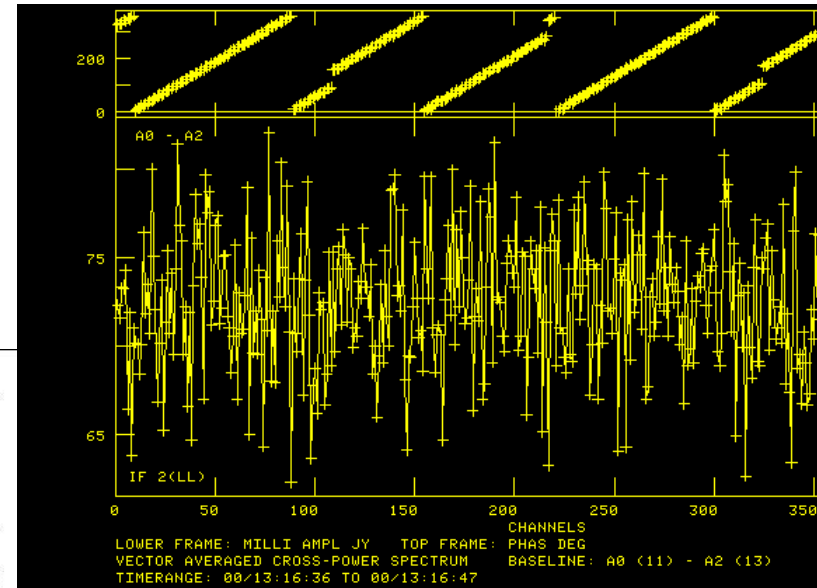
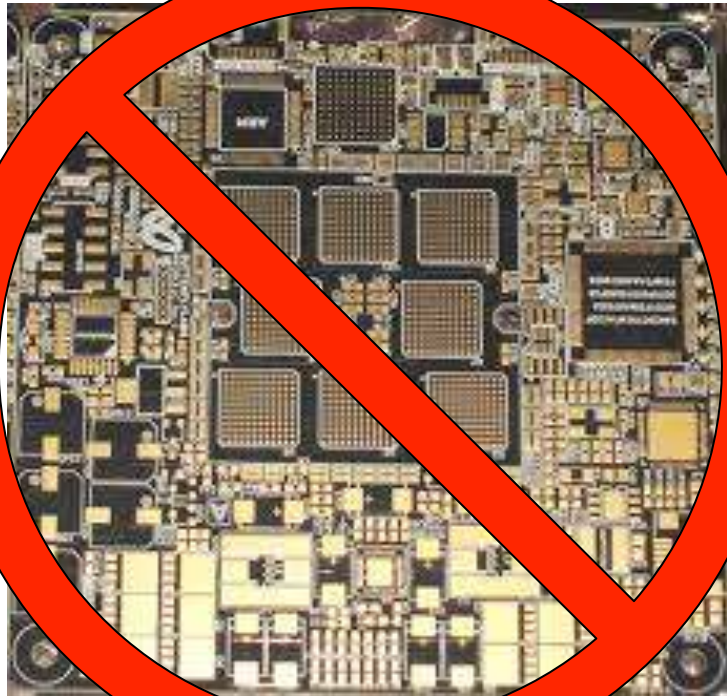


What is a correlator?





Why correlators matter to YOU



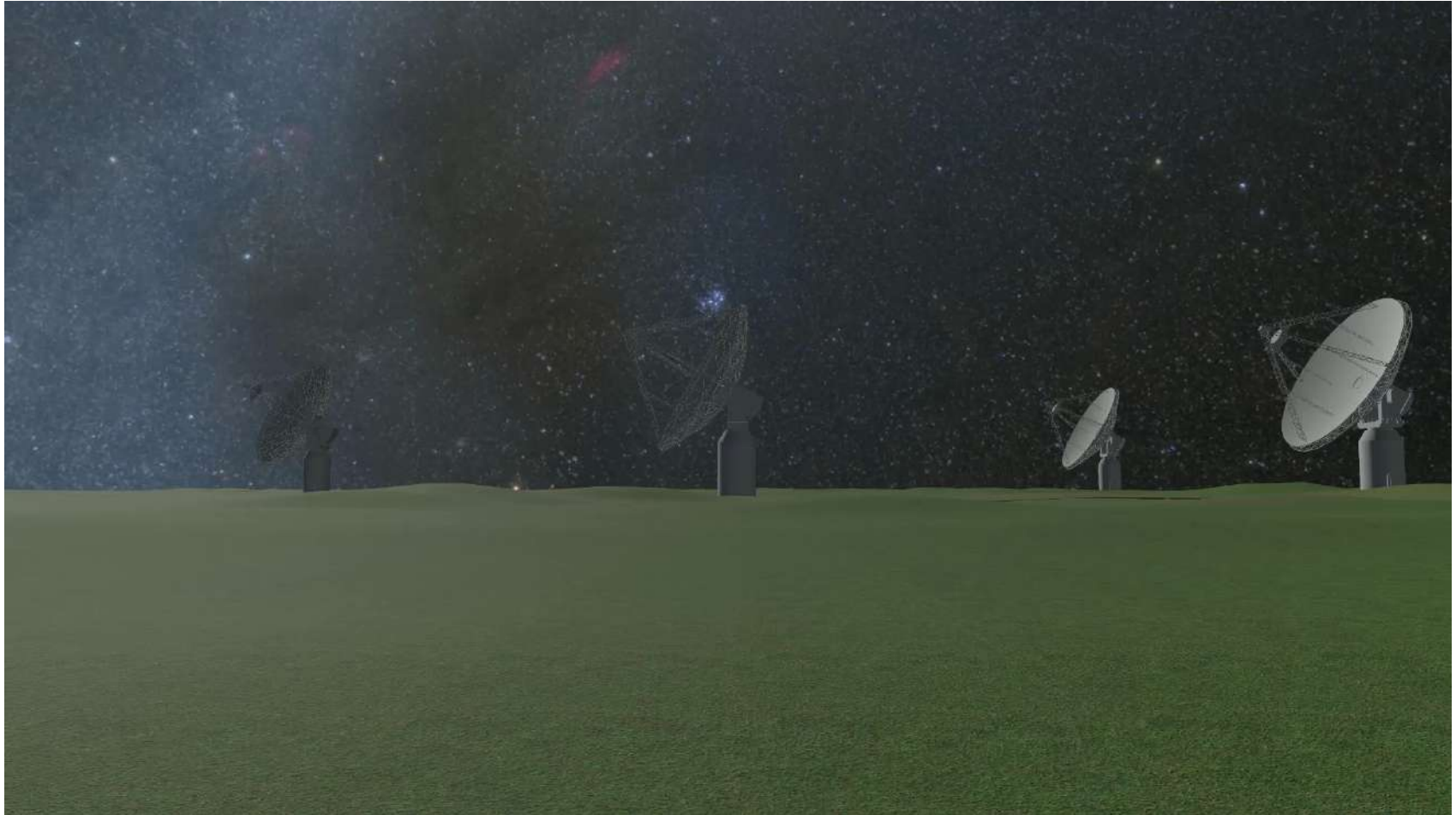
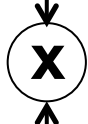
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Correlators and interferometry



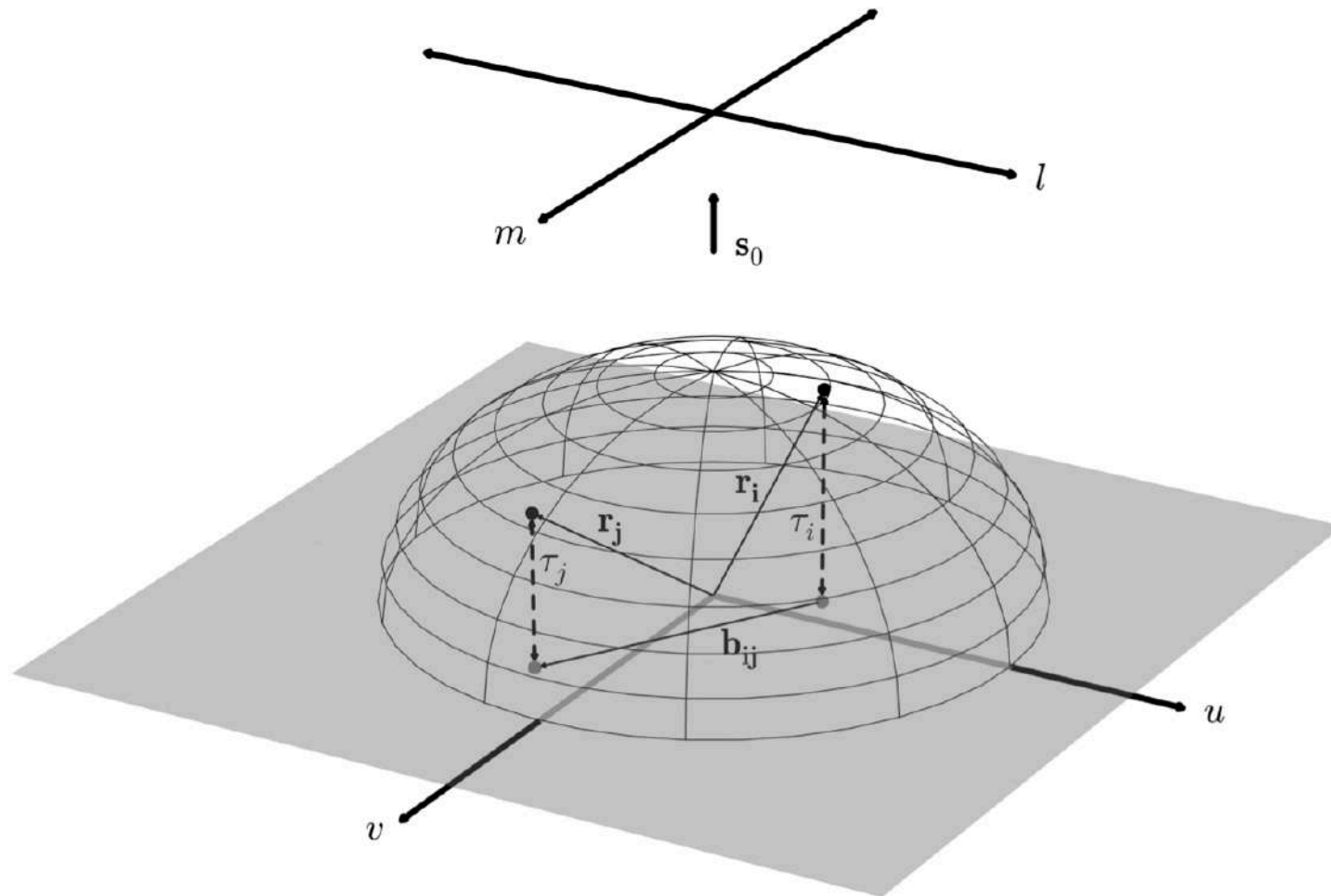
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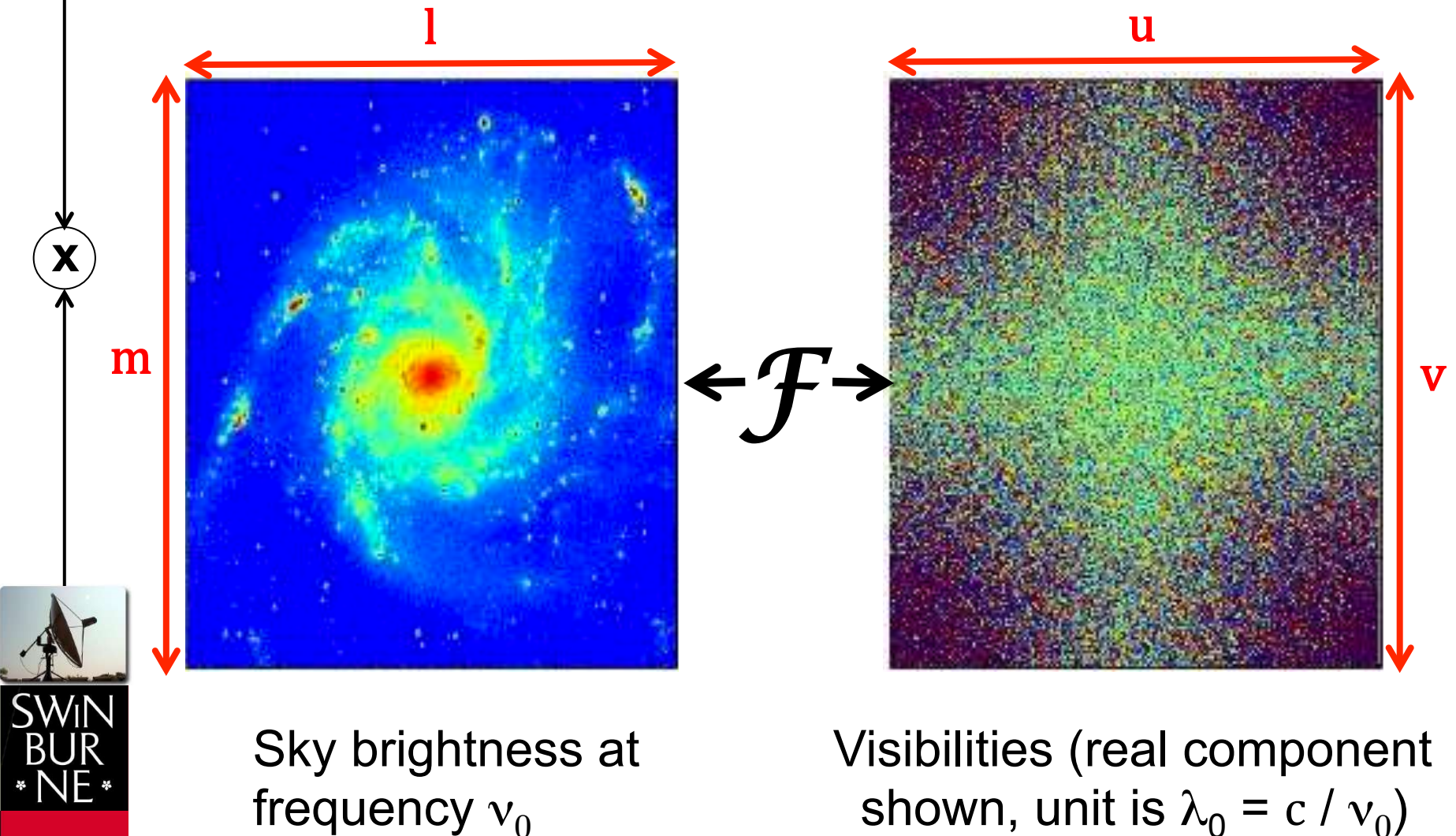


Correlators and Interferometry





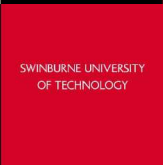
Correlators and Interferometry





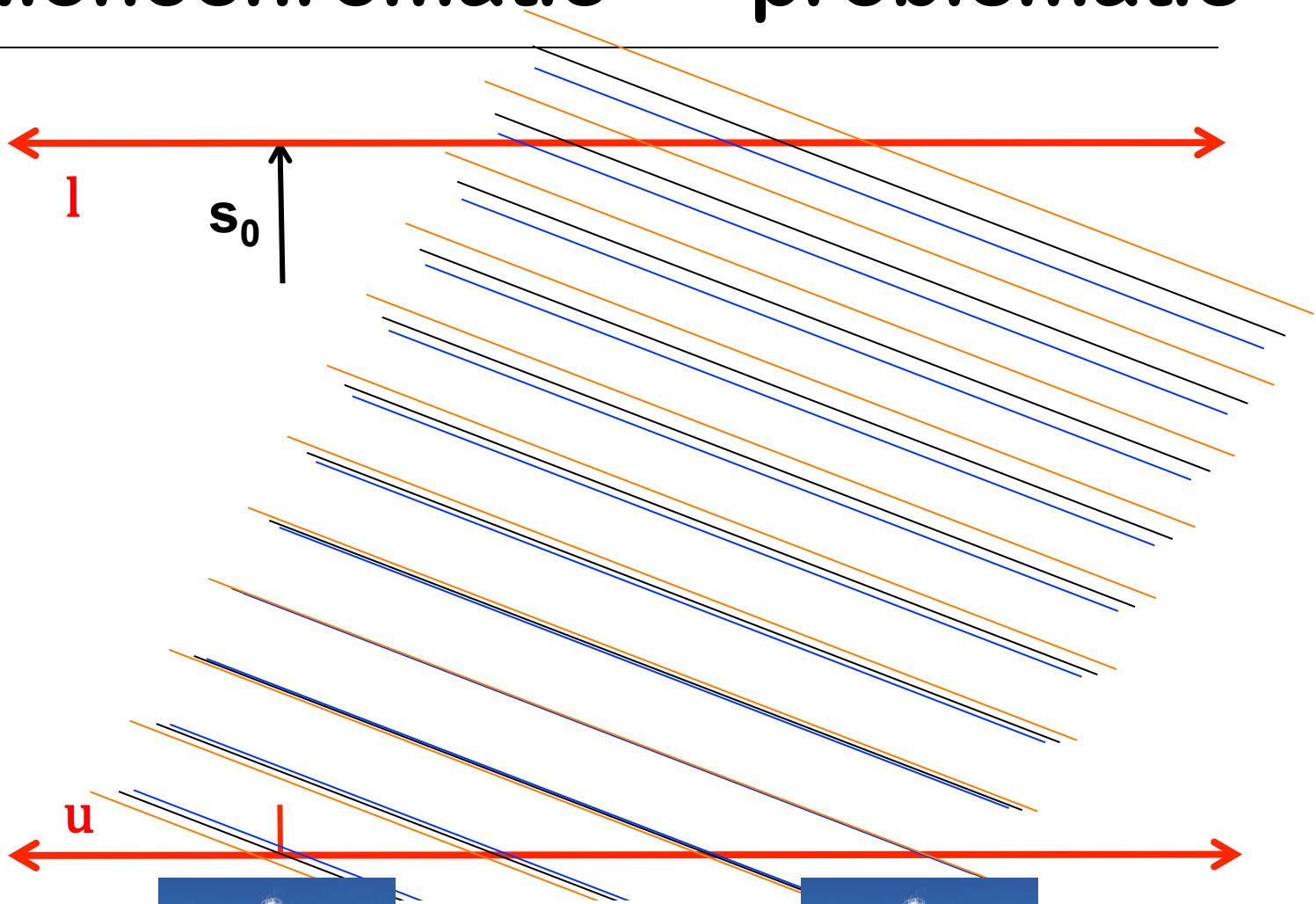
Monochromatic == problematic

- Mathematically: $u \times l + v \times m$ is supposed to be constant, but both u and v are $\propto \nu$
- No truly monochromatic radiation!
- Fortunately, “fairly narrow” band of $\Delta\nu$ (*quasi-monochromatic*) can suffice:
 - Real world viewpoint: different frequency components stay “in phase” as wavefront propagates from one antenna to the next





Monochromatic == problematic



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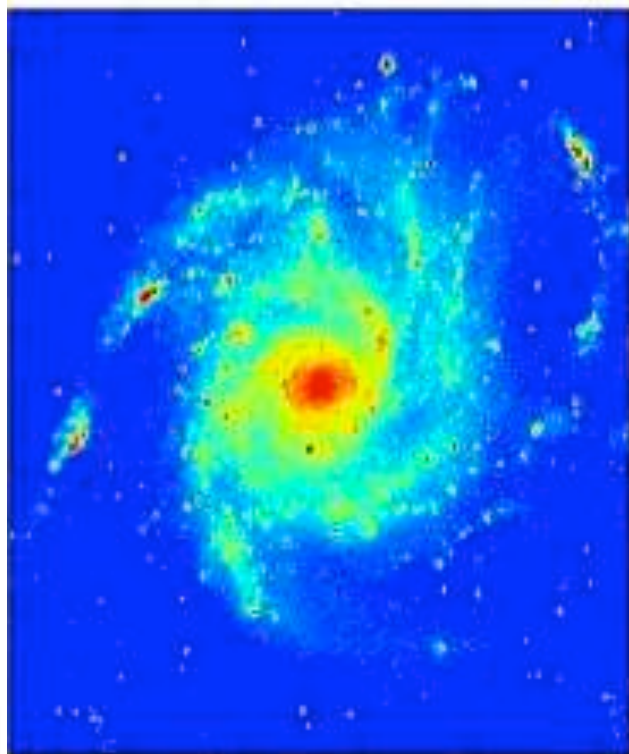
Monochromatic == problematic

- Mathematically: $u \times l + v \times m$ is supposed to be constant, but both u and v are $\propto \nu$
- No truly monochromatic radiation!
- Fortunately, “fairly narrow” band of $\Delta\nu$ (*quasi-monochromatic*) can suffice:
 - if $\Delta u \times l \ll 1$ and $\Delta v \times m \ll 1$ then the different frequency components stay in phase and we’re ok
 - Correlator needs to slice at least this finely



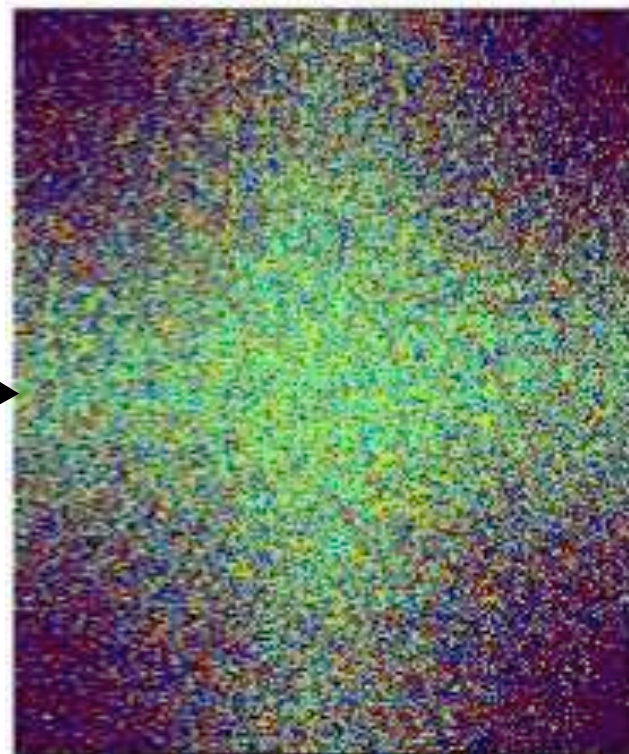


Correlators and Interferometry



Sky brightness at
frequency ν_0

$\leftarrow \mathcal{F} \rightarrow$



Visibilities (real component
shown, unit is $\lambda_0 = c / \nu_0$)

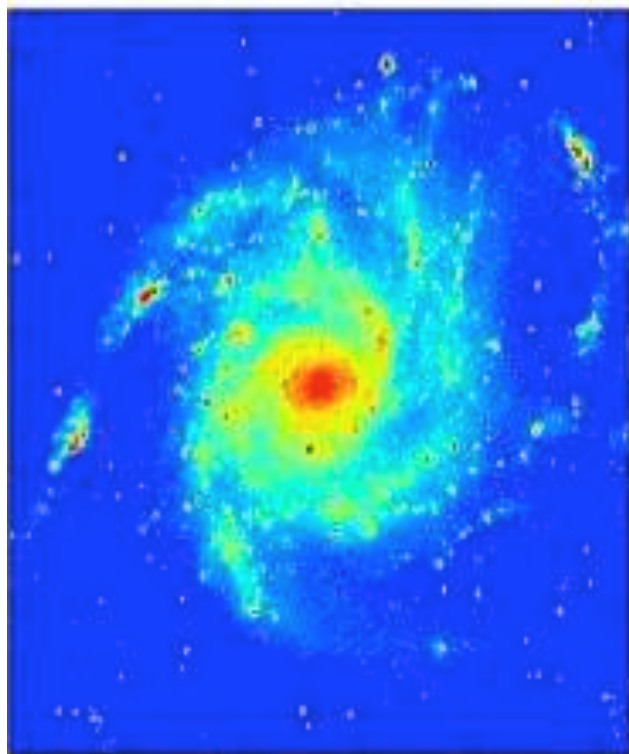
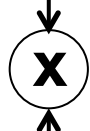


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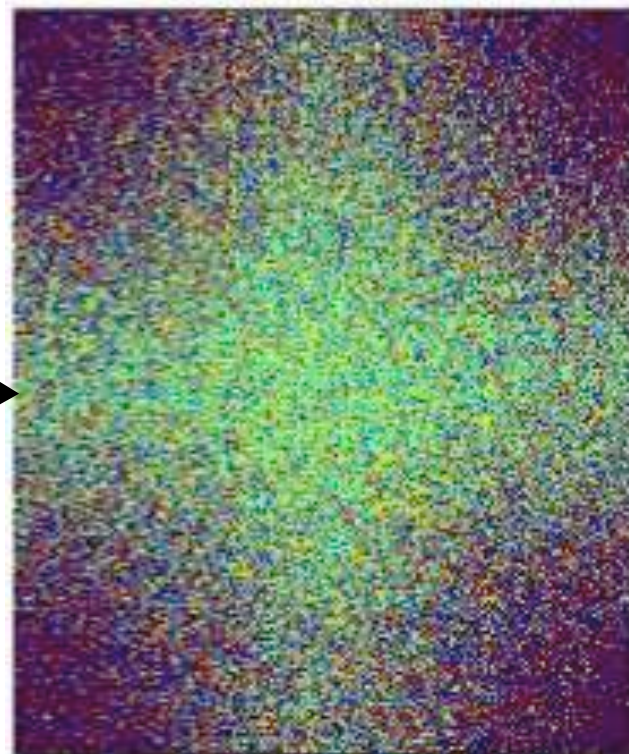
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Correlators and Interferometry



$\leftarrow \mathcal{F} \rightarrow$



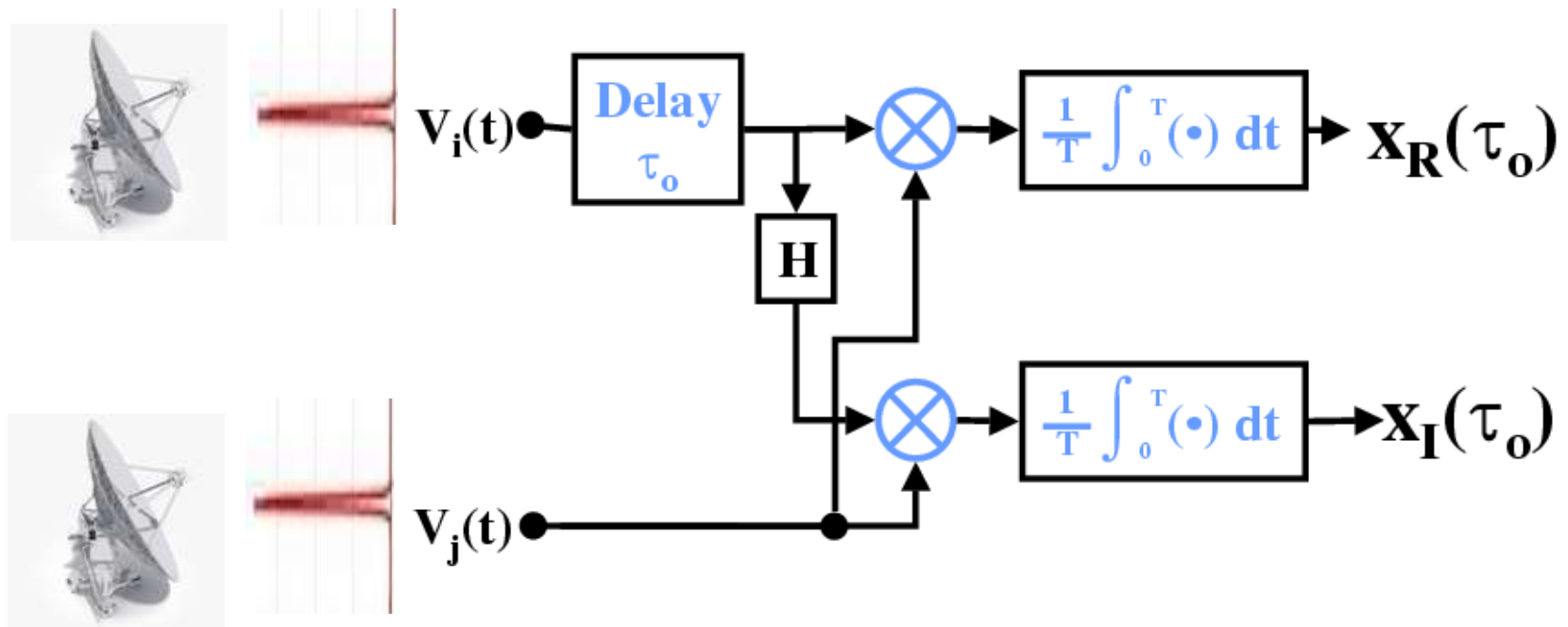
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Sky brightness at
frequency $\nu' = \nu_0 + \delta\nu$

Visibilities (real component
shown, unit is $\lambda' = c / \nu'$)

A “dumb” correlator

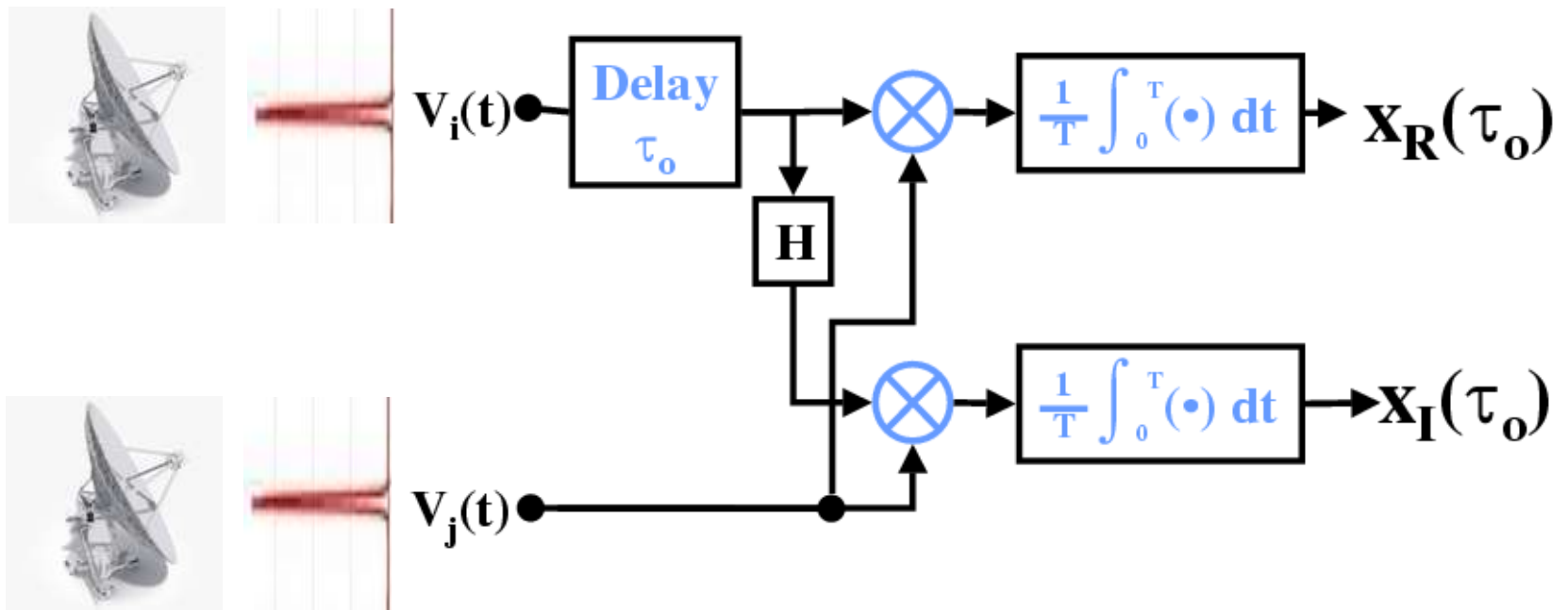
- Use many analog filters to make many narrow channels; correlate each one separately with a standard complex correlator:





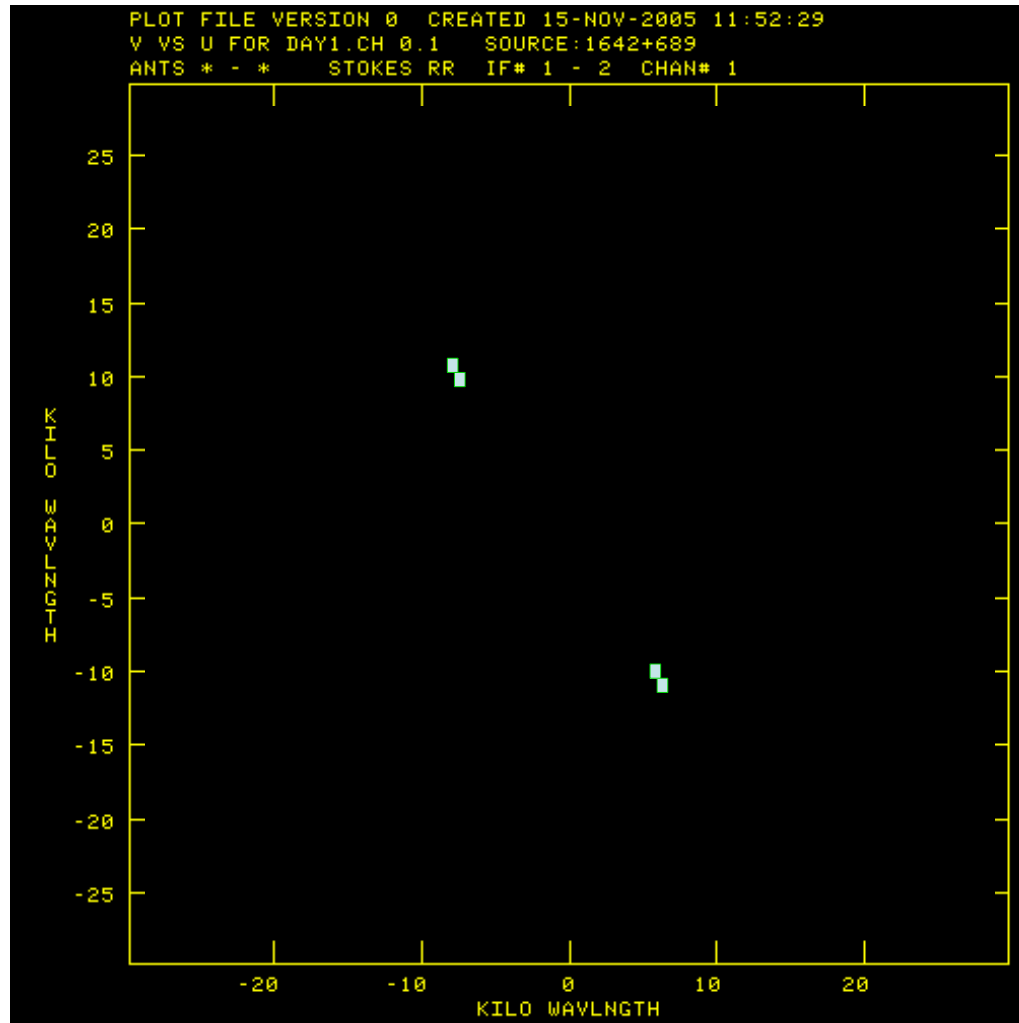
A “dumb” correlator

- Use many analog filters to make many narrow channels; correlate each one separately with a standard complex correlator:





The output



B
metres

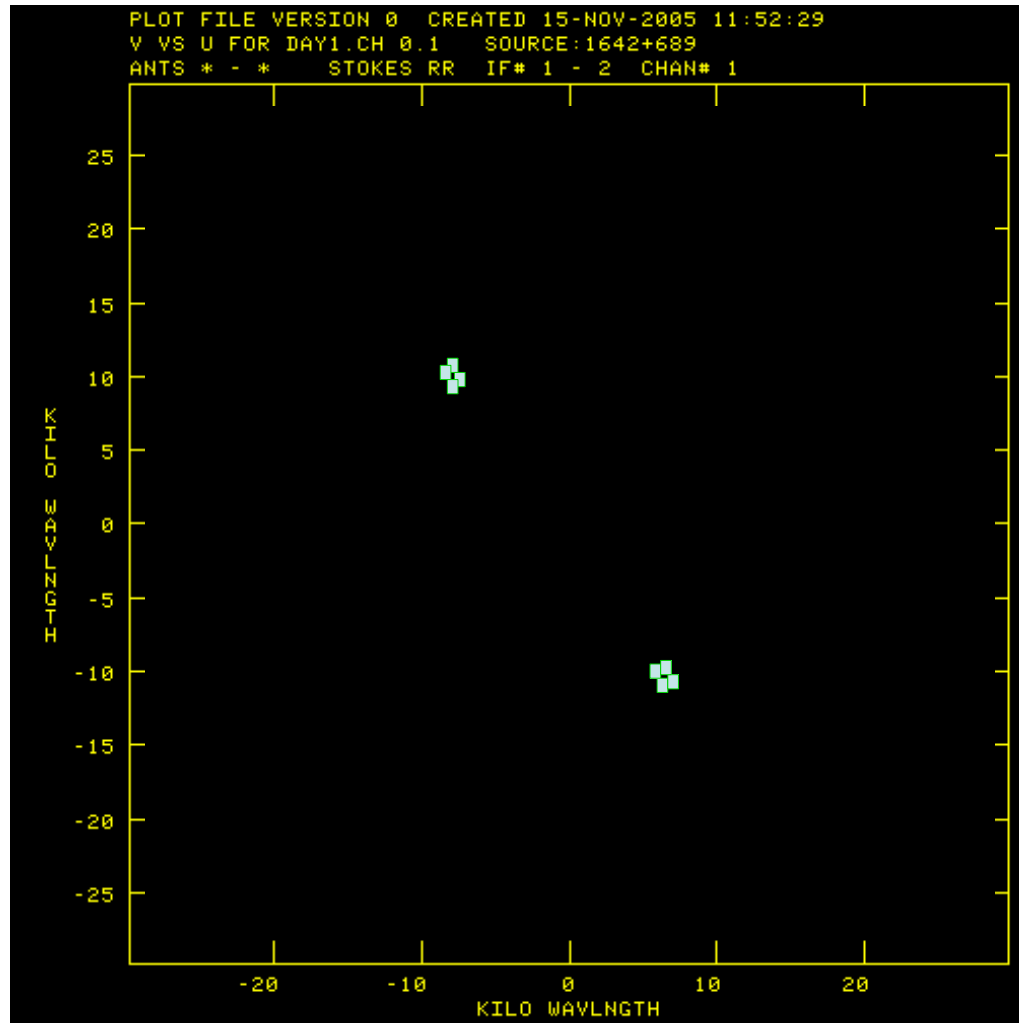


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The output



B'
metres



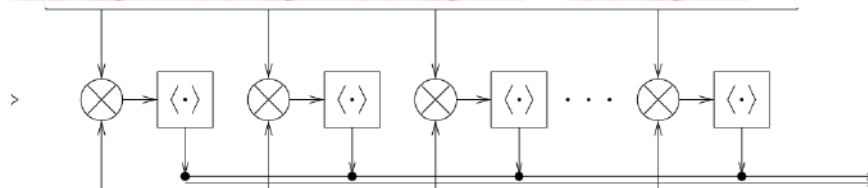
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Making it feasible

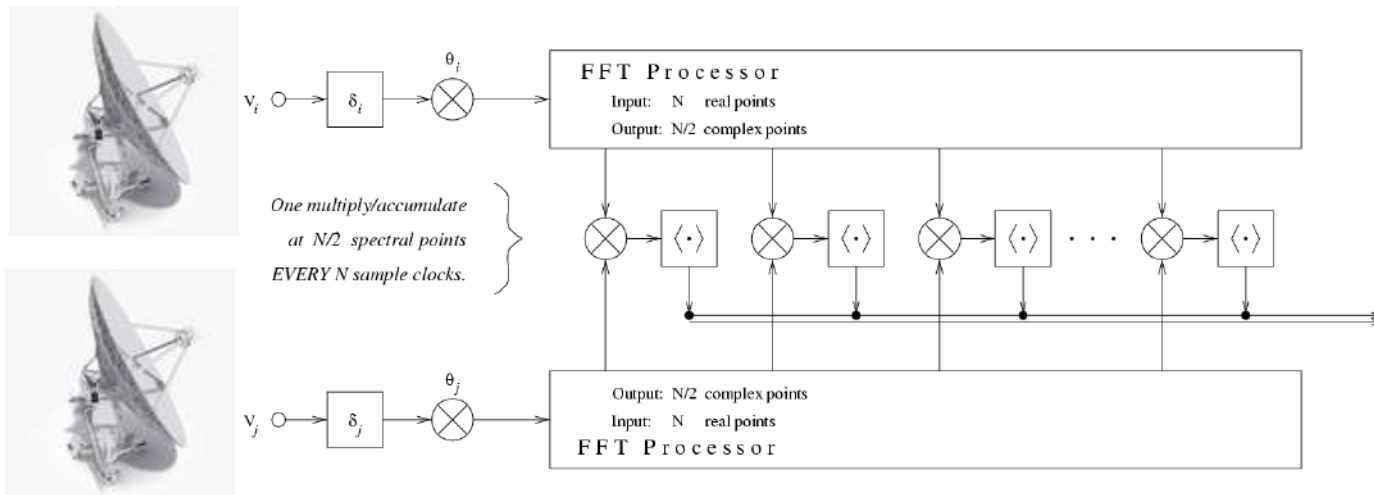
- Analog filters are costly & finnick; this would be expensive and temperamental





Making it feasible

- Analog filters are costly & finnick; this would be expensive and temperamental
- Fortunately, we can (and do) digitize the signal – meaning we can use a digital substitute: **digital filterbank**

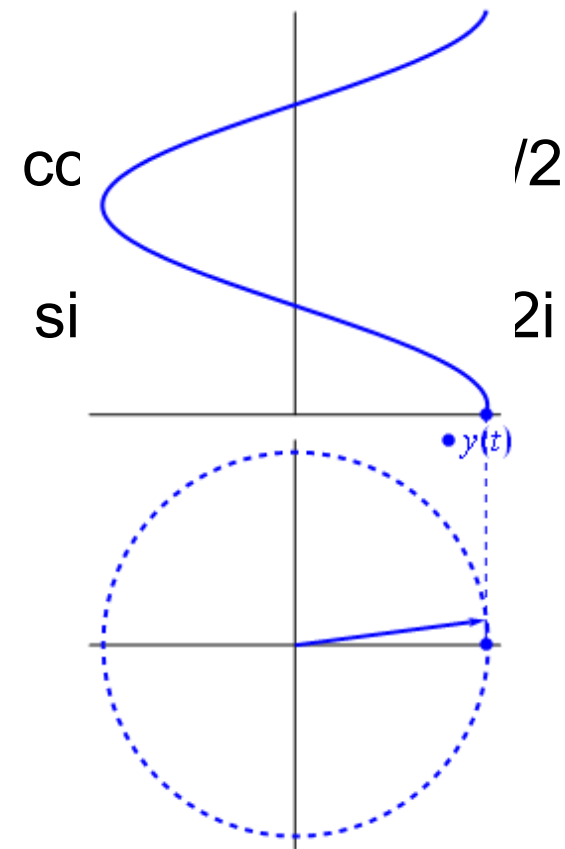




The advantage of going digital

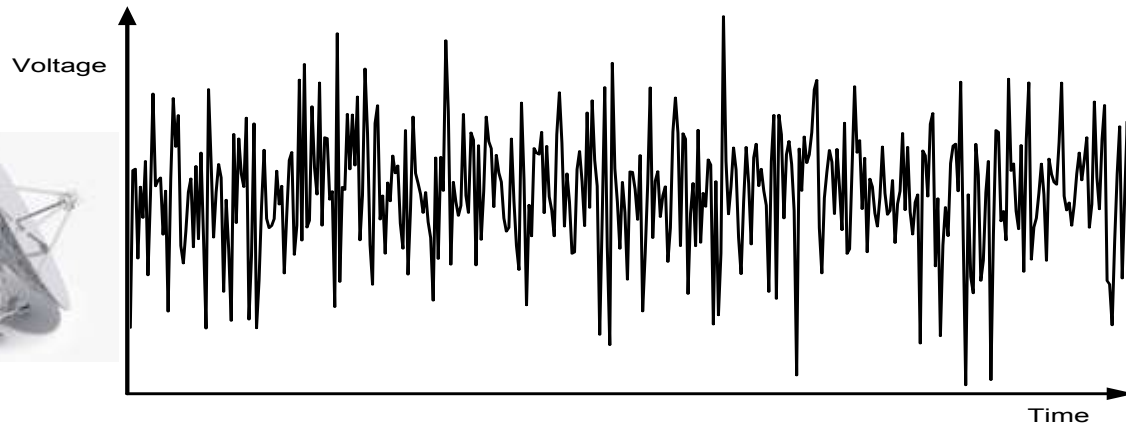
- Stable, cheap filters
- Produces complex output: when cross-multiplying, use 1 complex multiplier rather than 2 real multipliers and a phase shift

$$e^{i\phi} = \cos \phi + i \sin \phi$$





The “FX” correlator



X



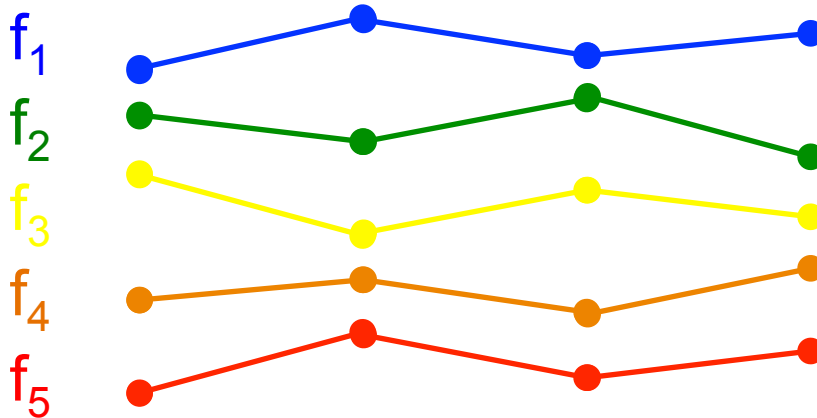
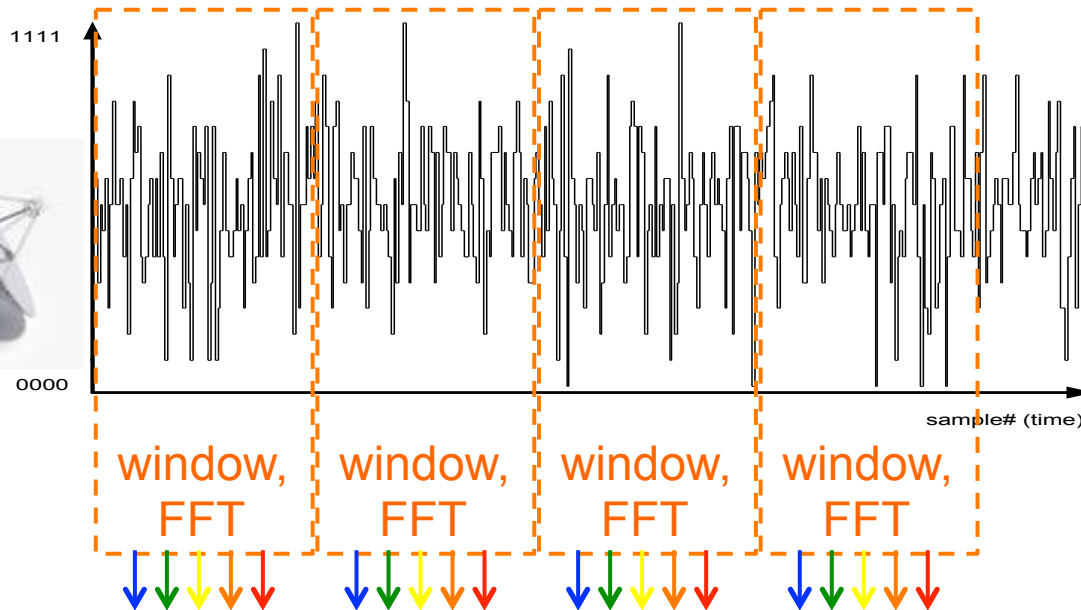
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The “FX” correlator

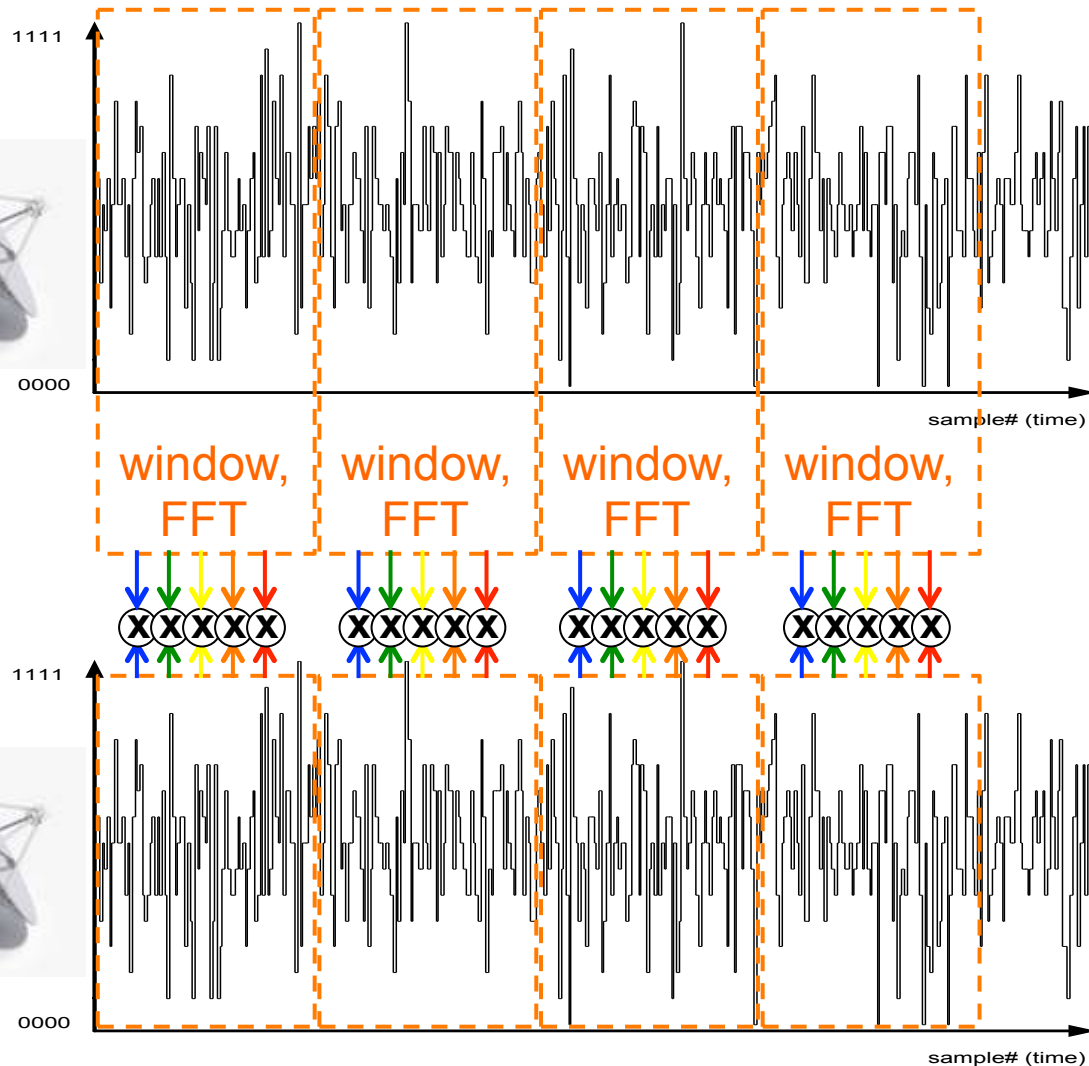


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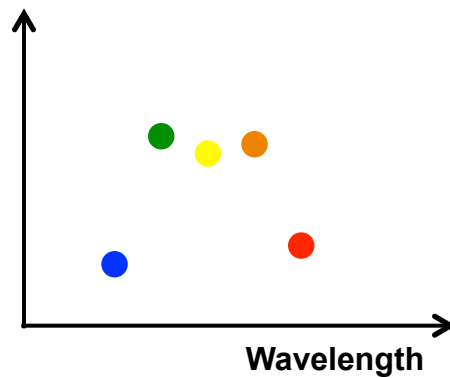
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The “FX” correlator

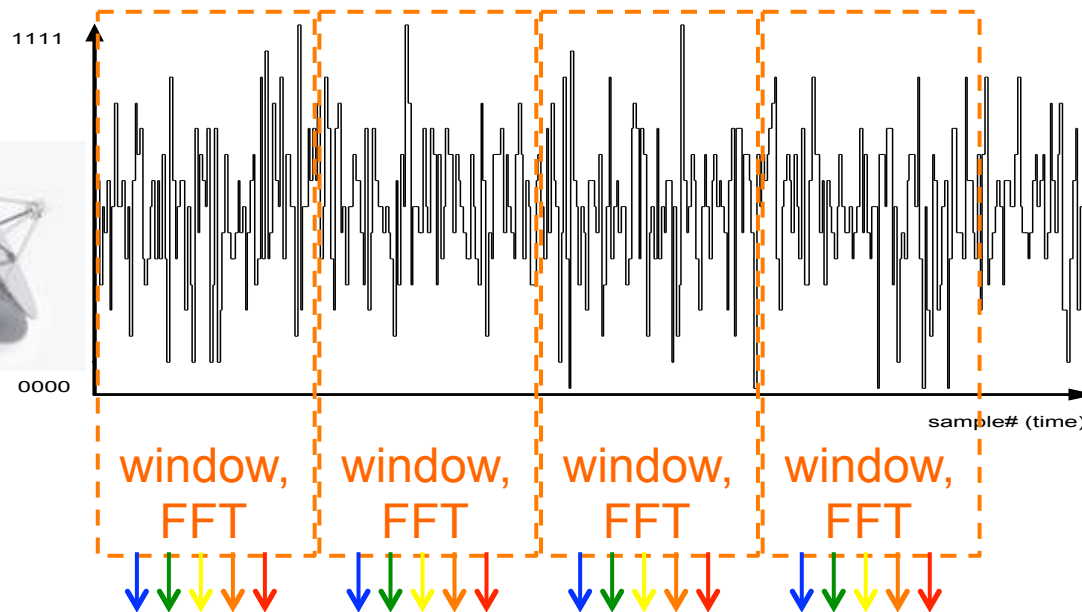


Visibility
amplitude





The “FX” correlator

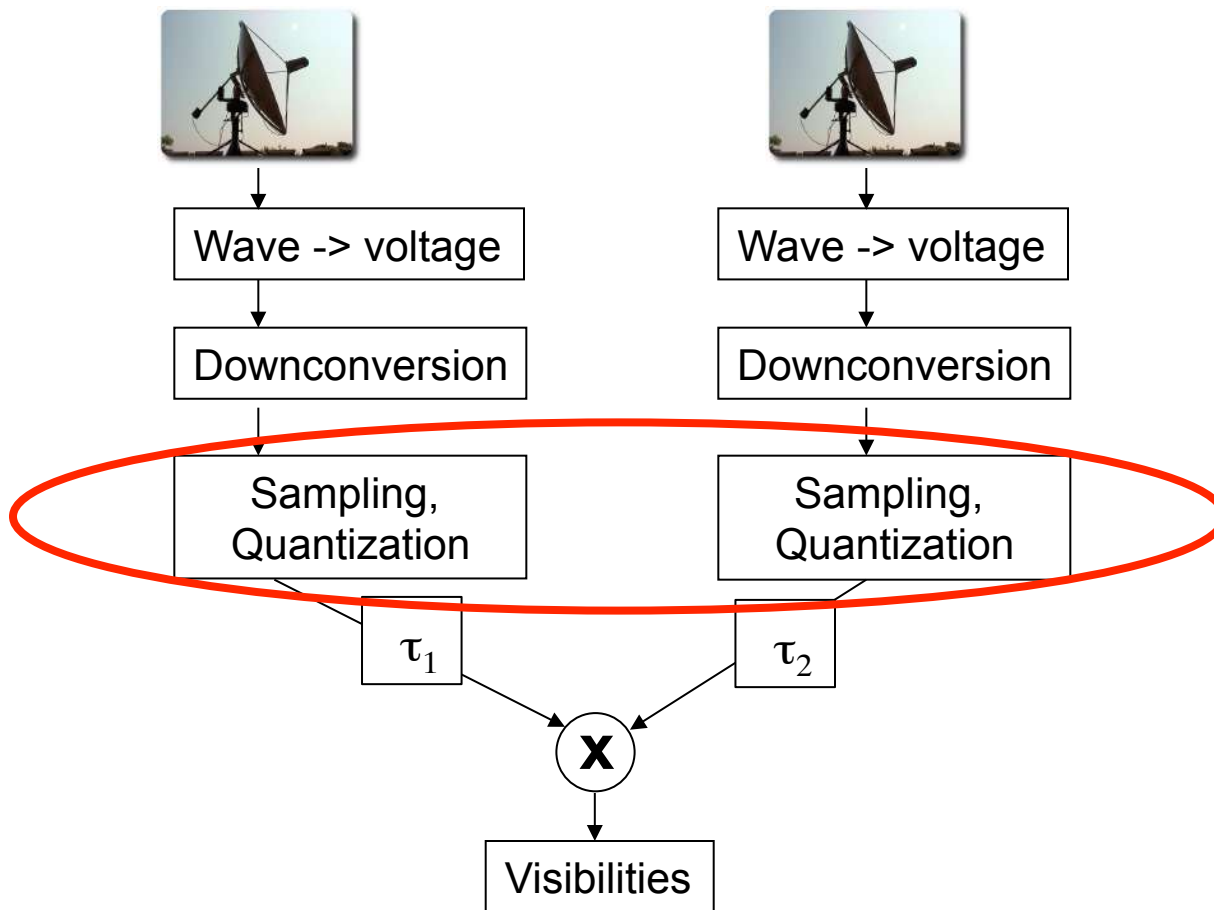


- Since this architecture consists of a Fourier transform (F) followed by cross-multiplication (X), we dub this the “FX” correlator





Righting the wrongs



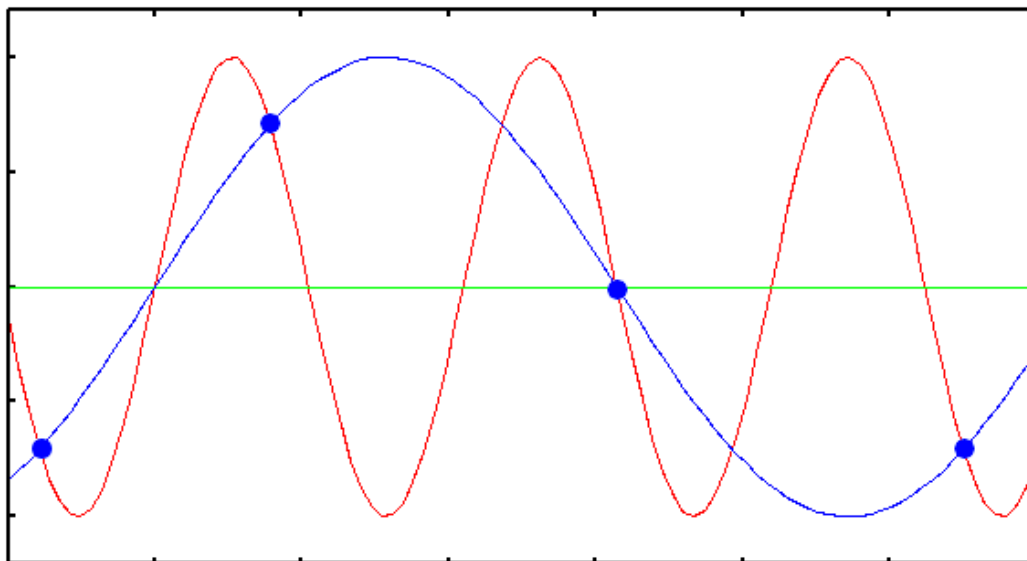
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Sampling

- Nyquist-Shannon sampling theorem:
 - real-valued signal is sampled every Δt sec
 - Original signal can be reconstructed perfectly so long as contains no power at frequencies $\geq 1 / (2 \Delta t)$ Hz (*band-limited*)



Adequately sampled

Undersampled,
cannot be
reconstructed



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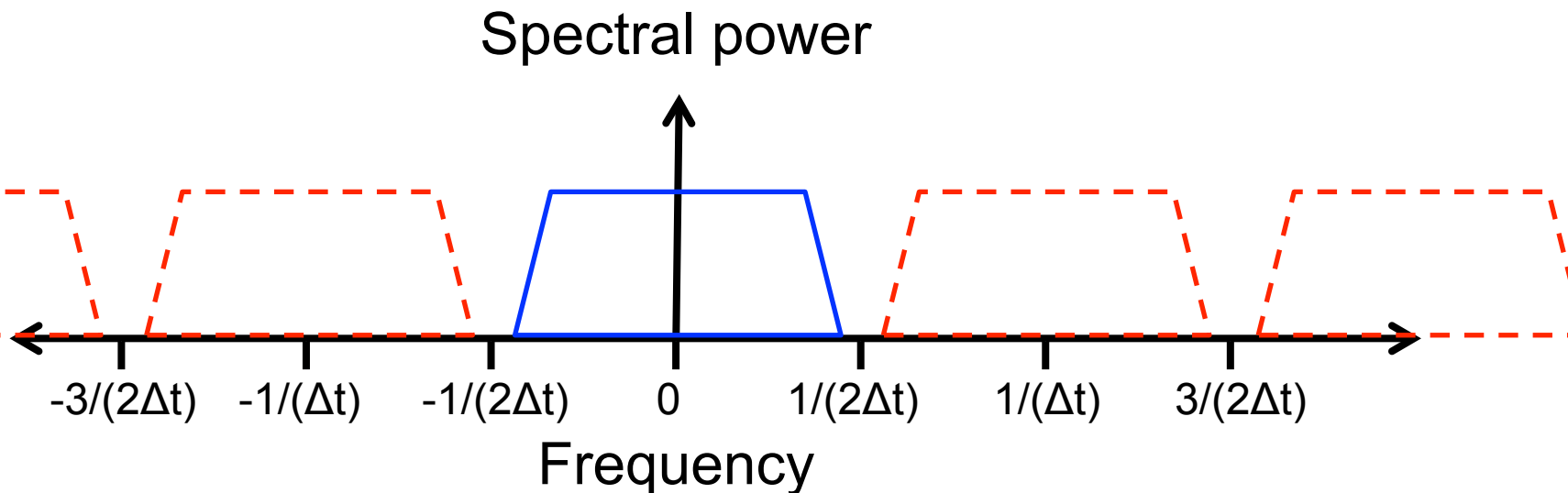
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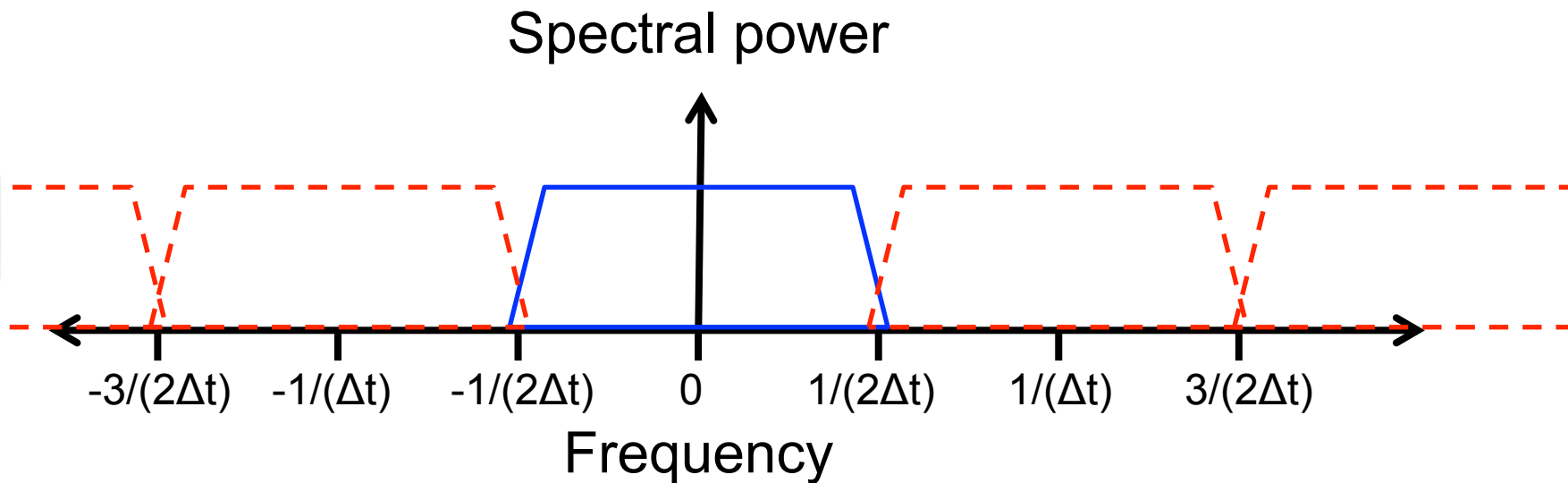
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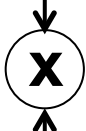
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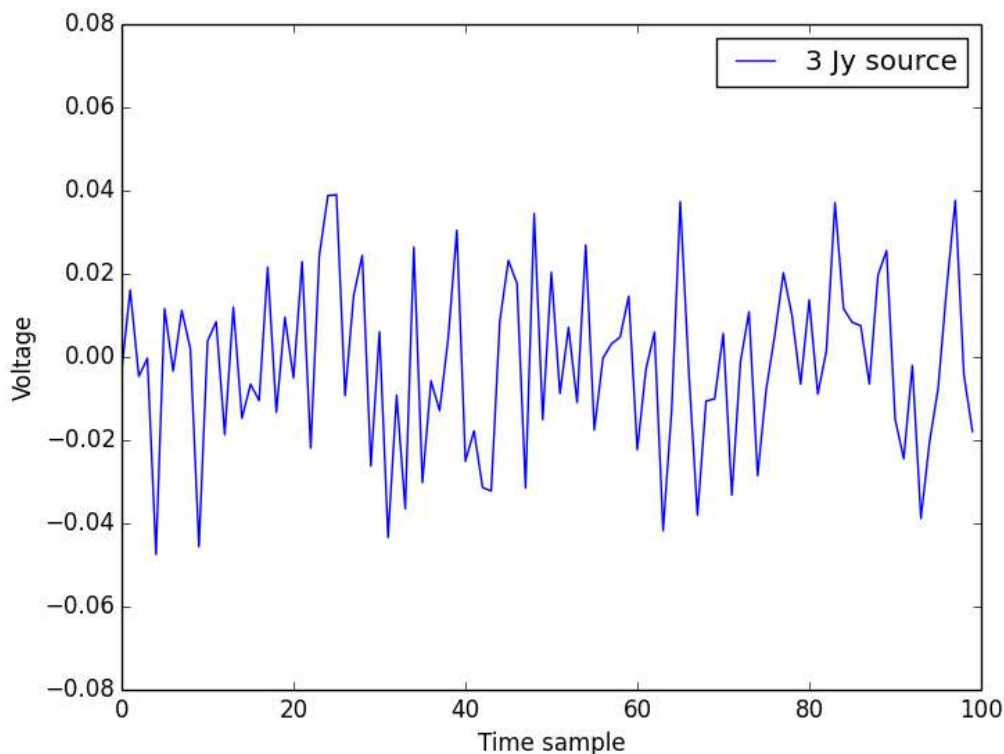
Quantization

- When correlation is low (almost always) even very coarse quantization is ok!



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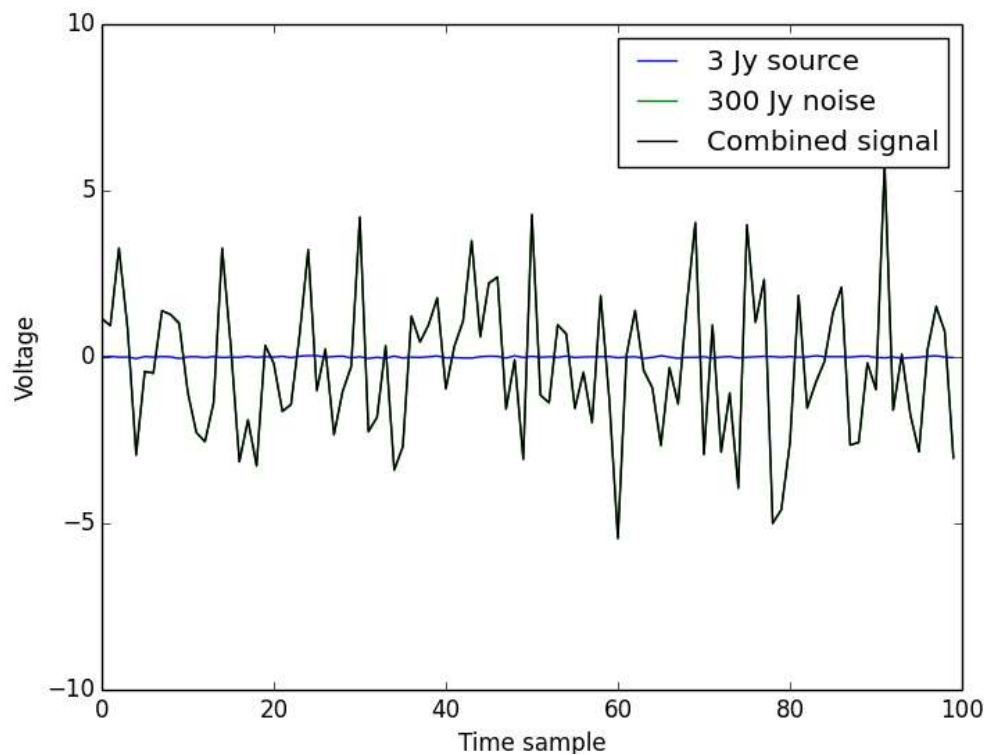
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Quantization

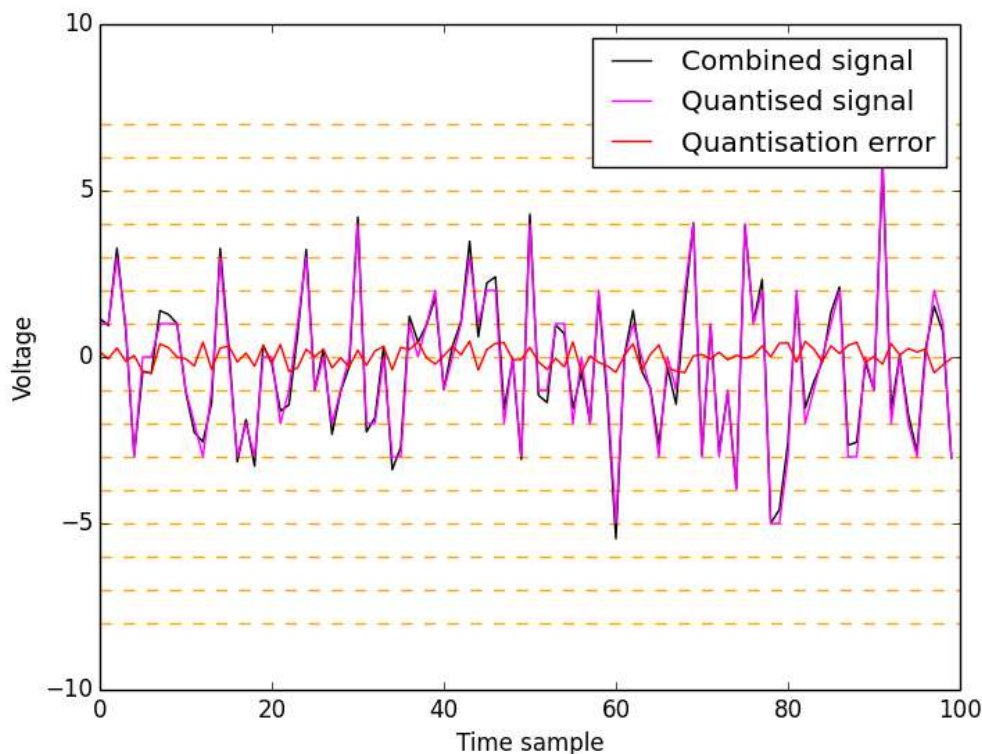
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Quantization

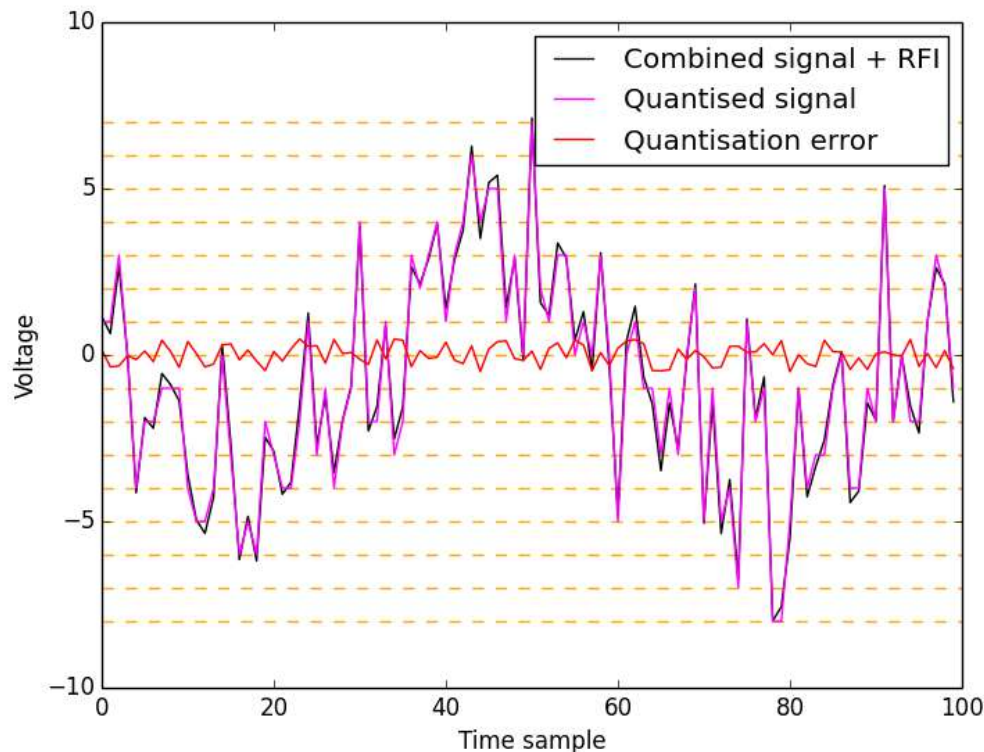
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Quantization

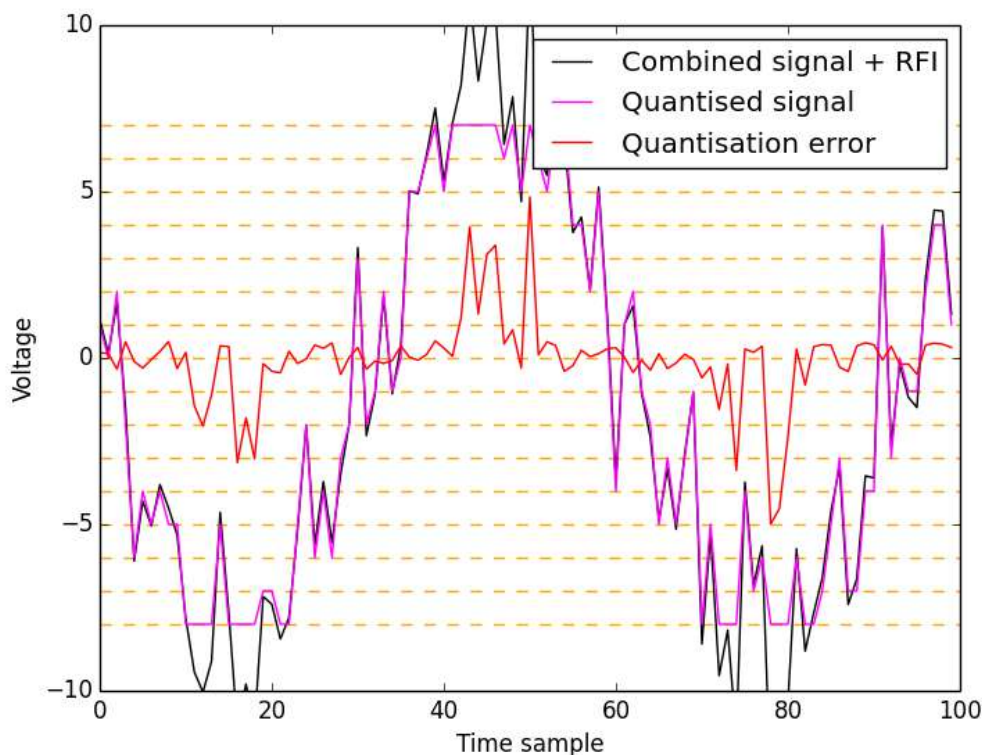
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Quantization

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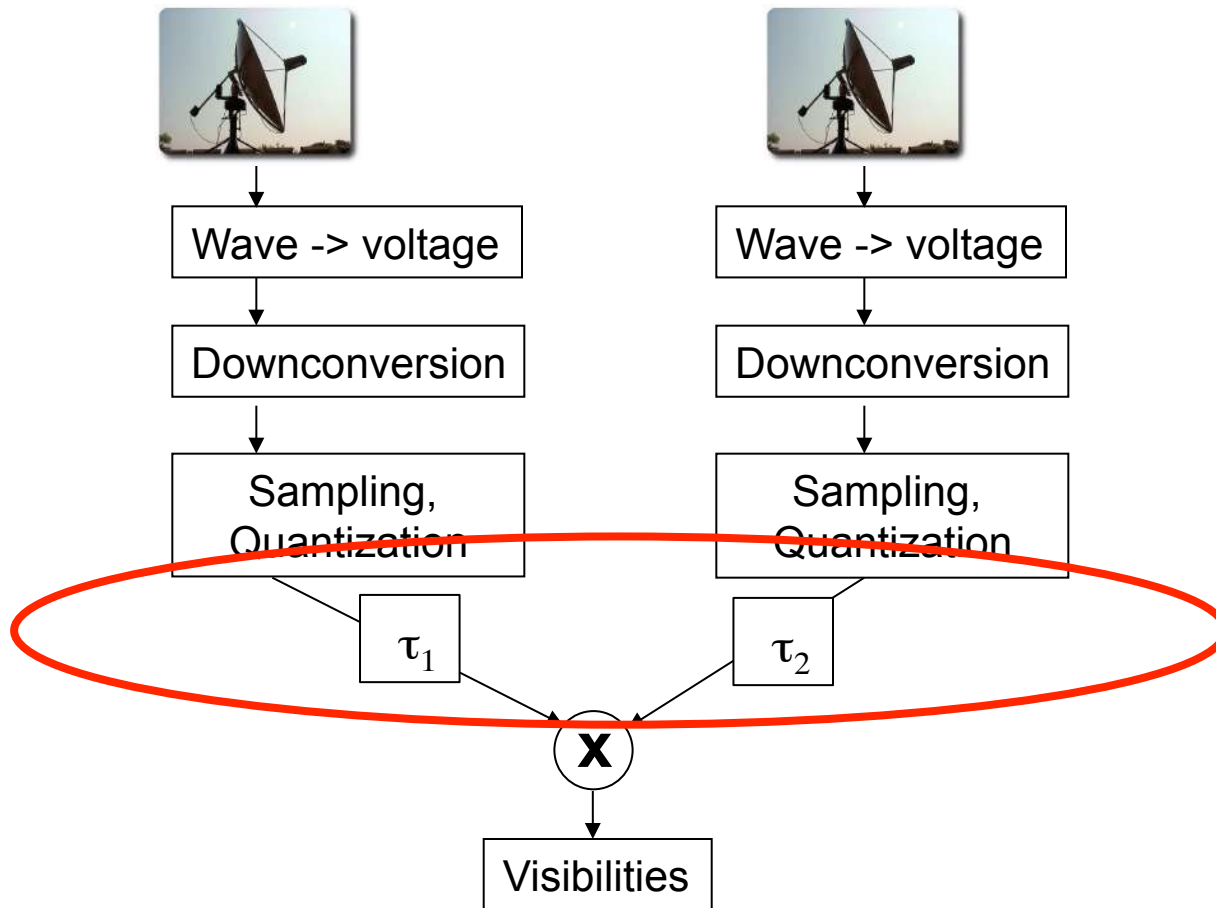


Quantization

- When correlation is low (almost always) even very coarse quantization is ok!
- Sensitivity loss due to quantisation:
 - 8 bit: 0.1%
 - 4 bit: 1.3%
 - 2 bit: 12%
 - 1 bit: 36%
- Correct visibility amplitudes for this sensitivity loss



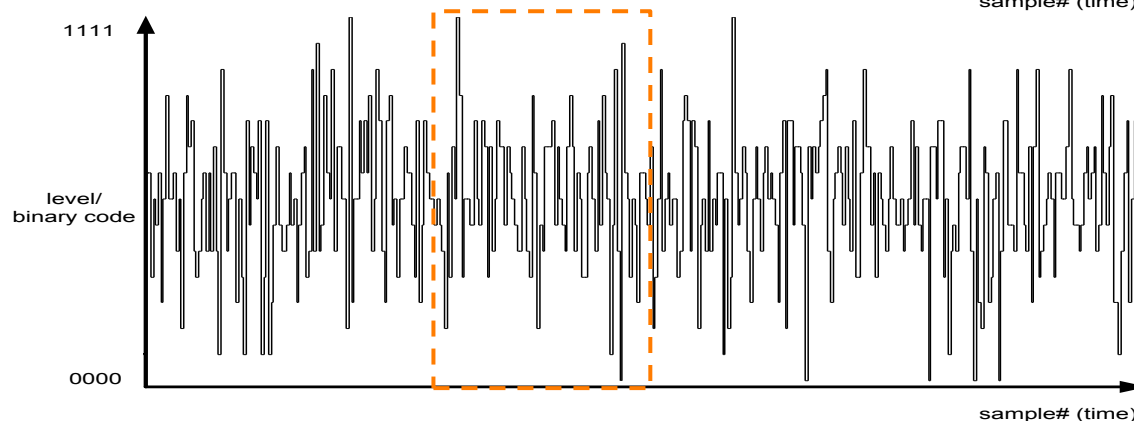
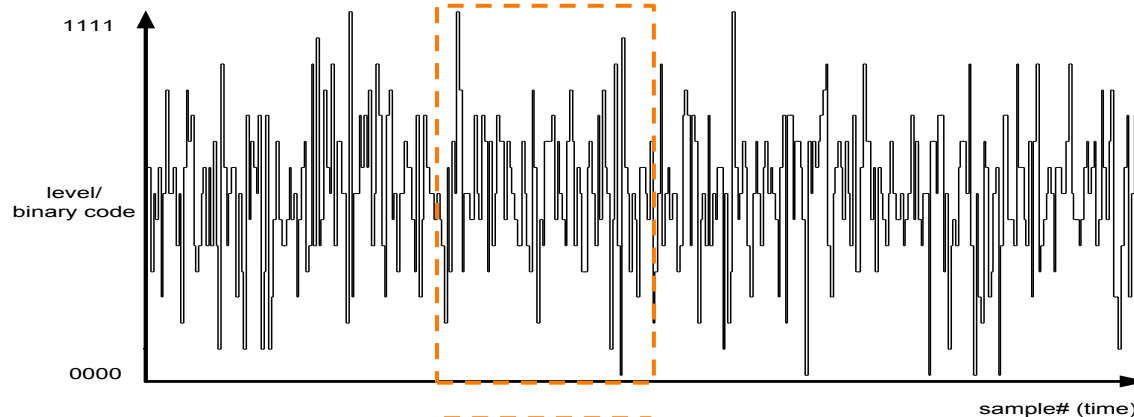
Righting the wrongs





Delay compensation

- Delay to the nearest sample is easy:



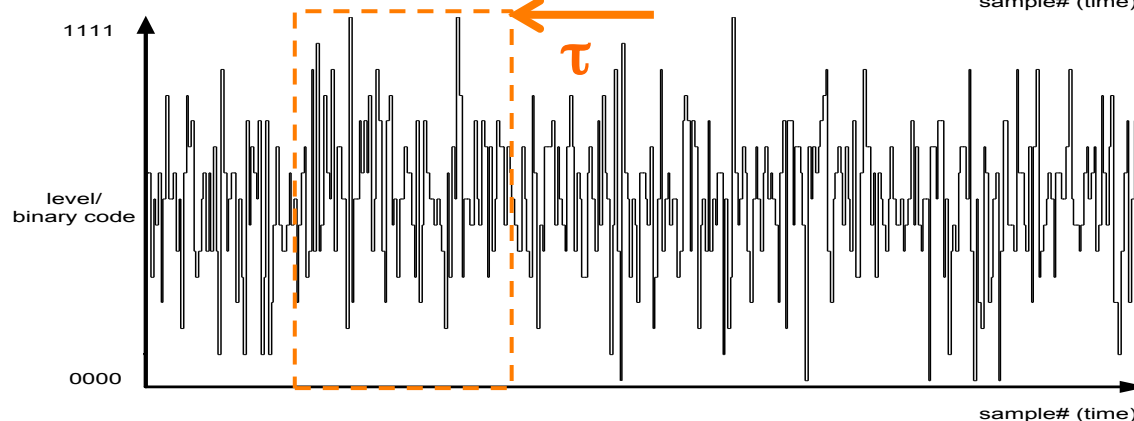
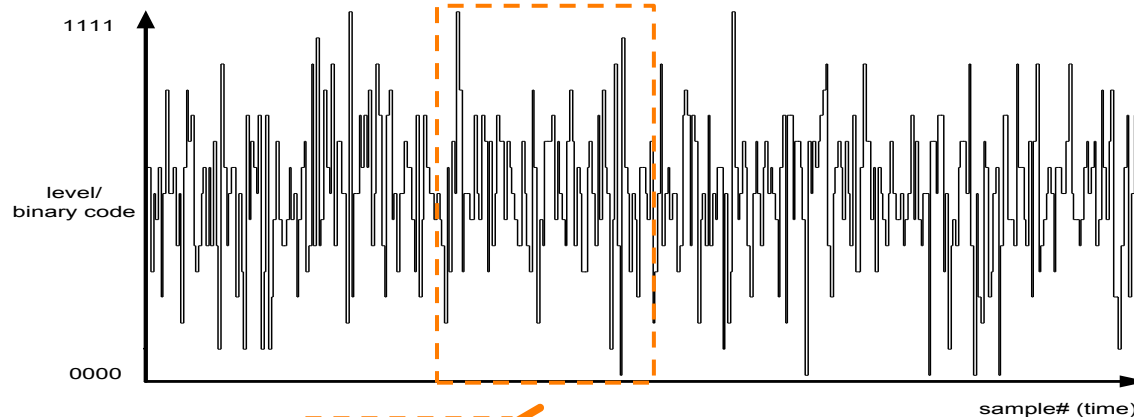
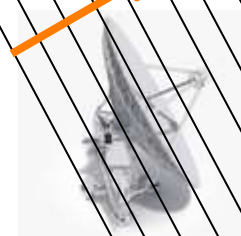


Delay compensation

- Delay to the nearest sample is easy:



τ



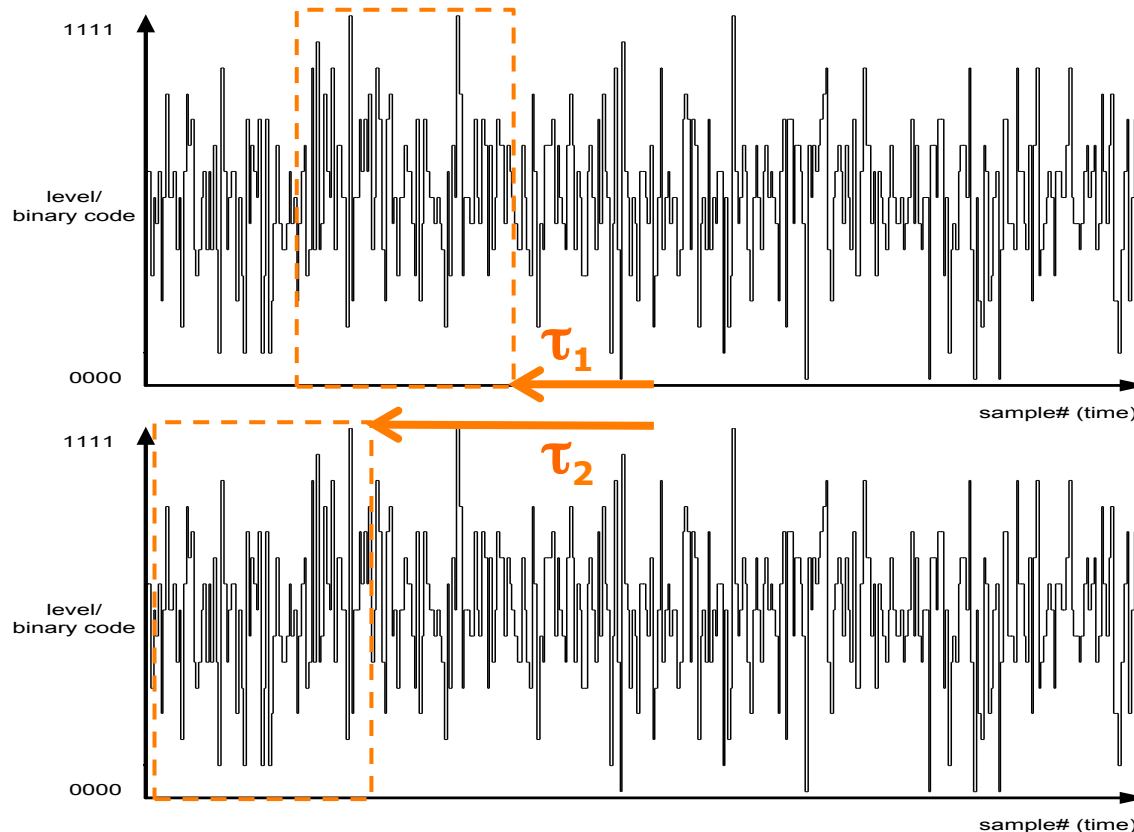
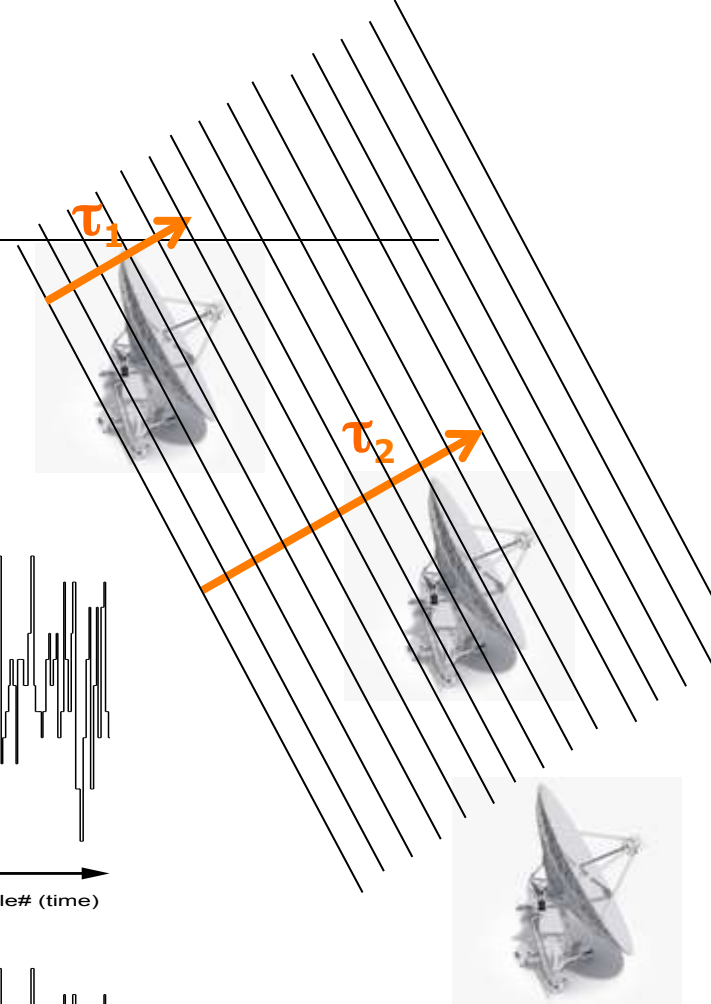
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Delay compensation

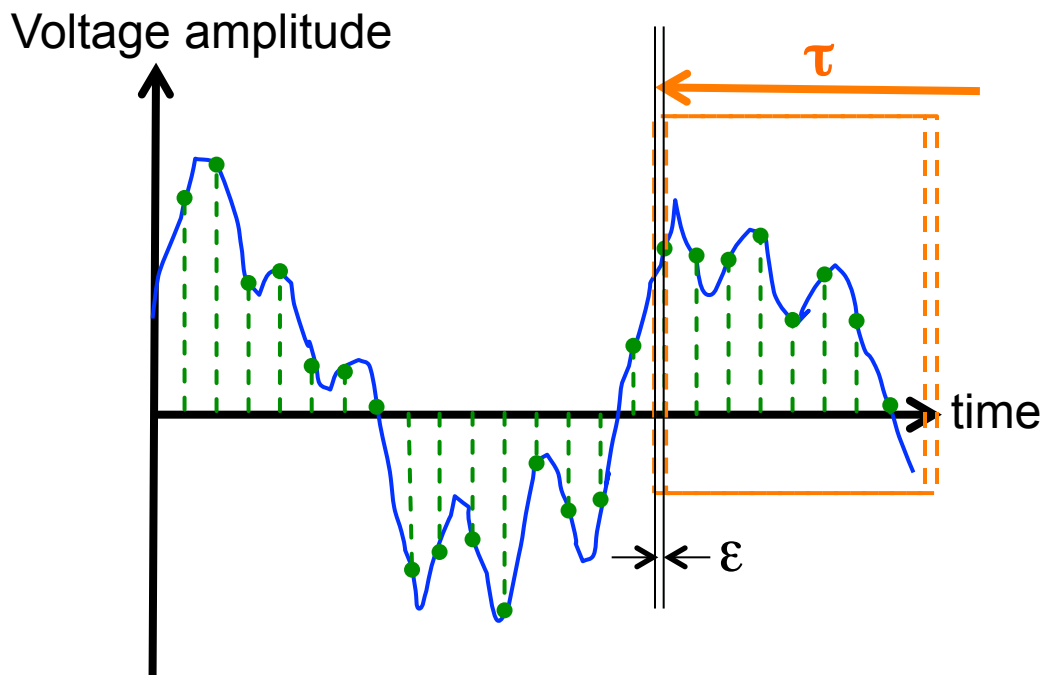
- In practise, delay all to common reference





Fractional-sample correction

- Sampling prevents perfect alignment of datastreams; always a small error



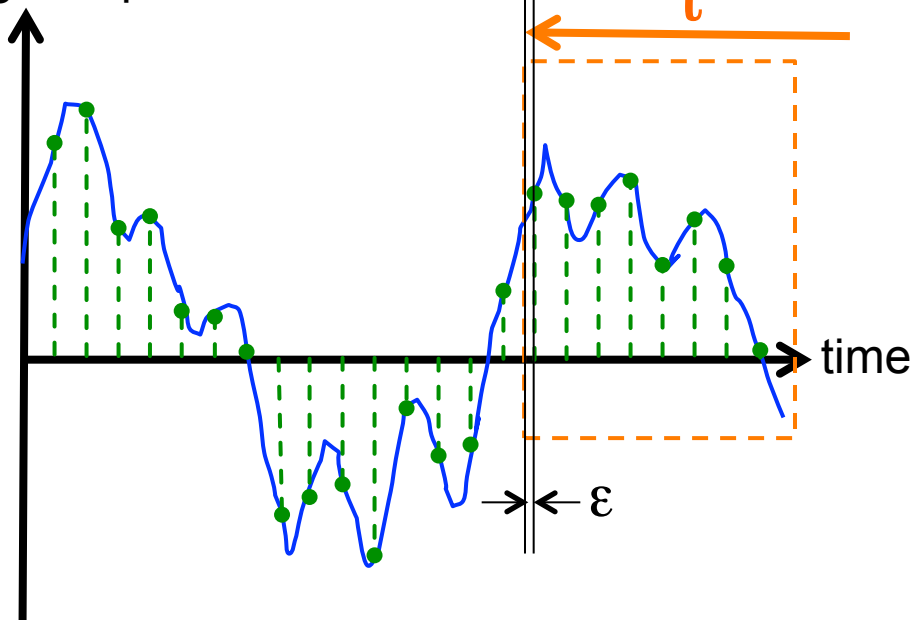


Fractional-sample correction

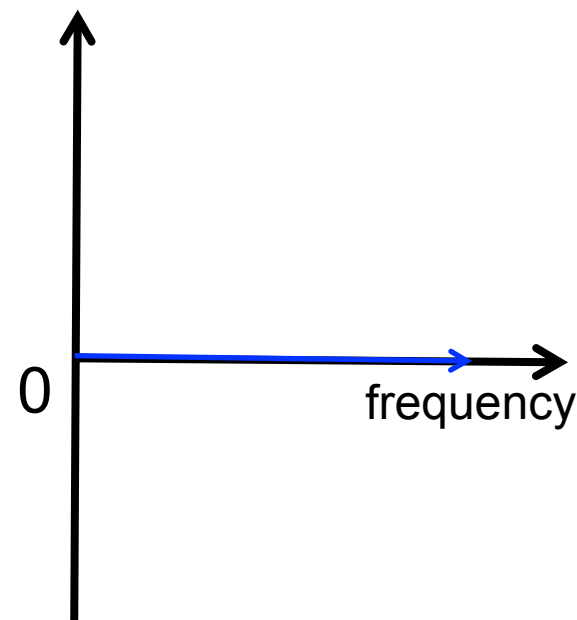
- Sampling prevents perfect alignment of datastreams; always a small error



Voltage amplitude



Visibility phase



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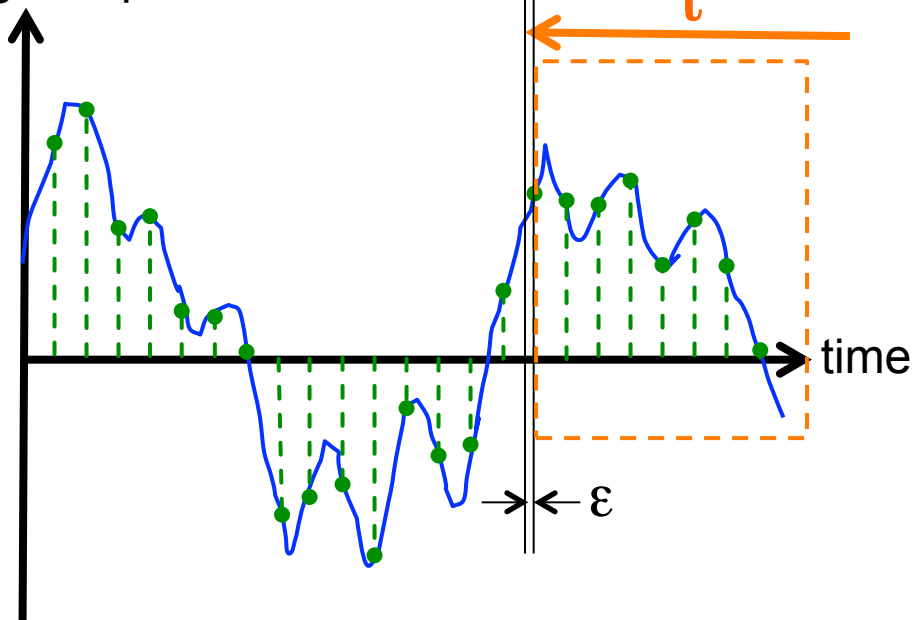


Fractional-sample correction

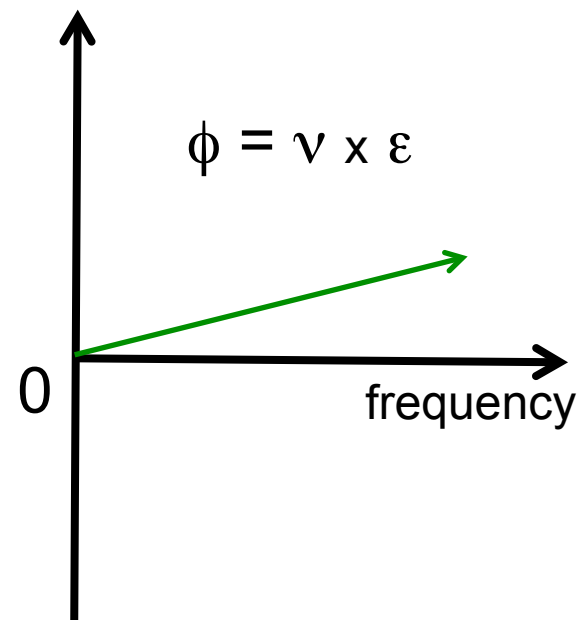
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Voltage amplitude



Visibility phase



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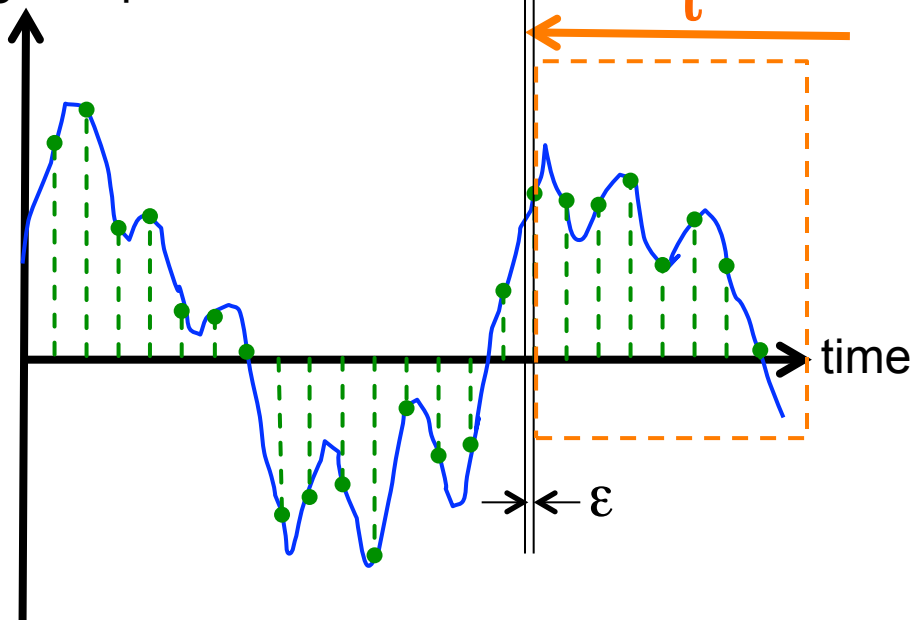


Fractional-sample correction

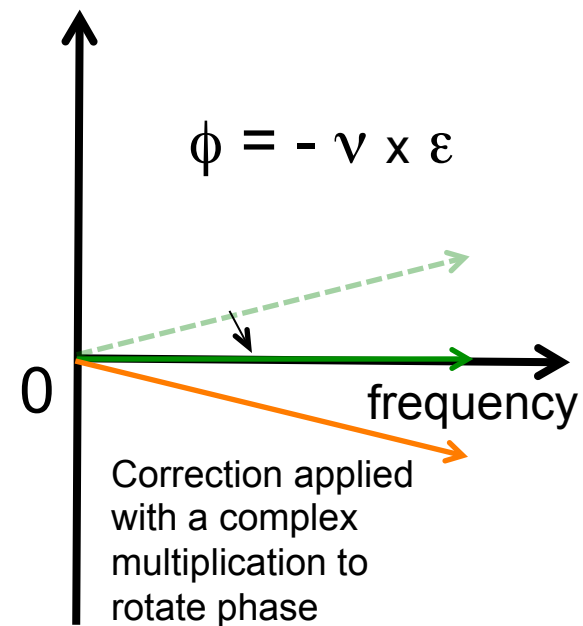
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Voltage amplitude



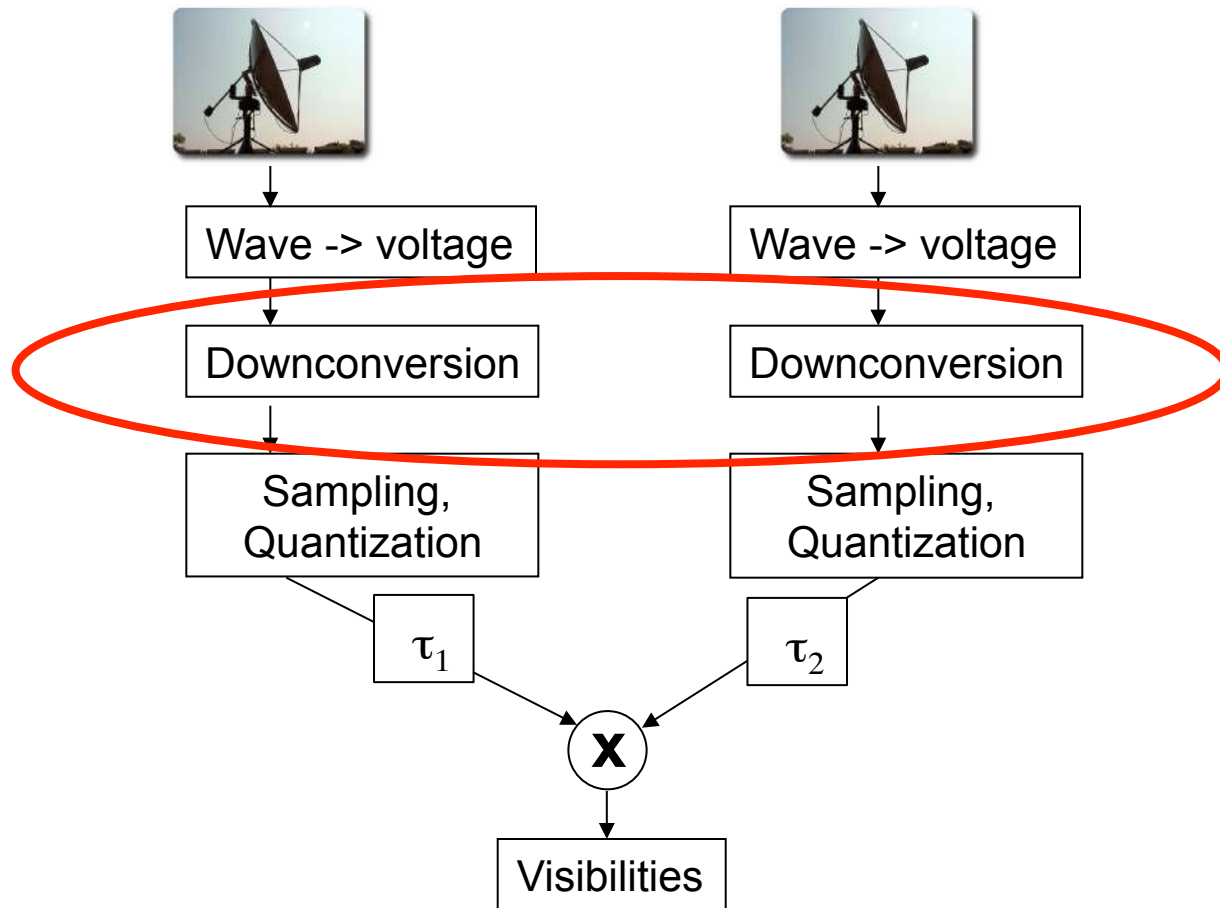
Visibility phase



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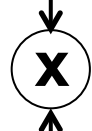
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Righting the wrongs



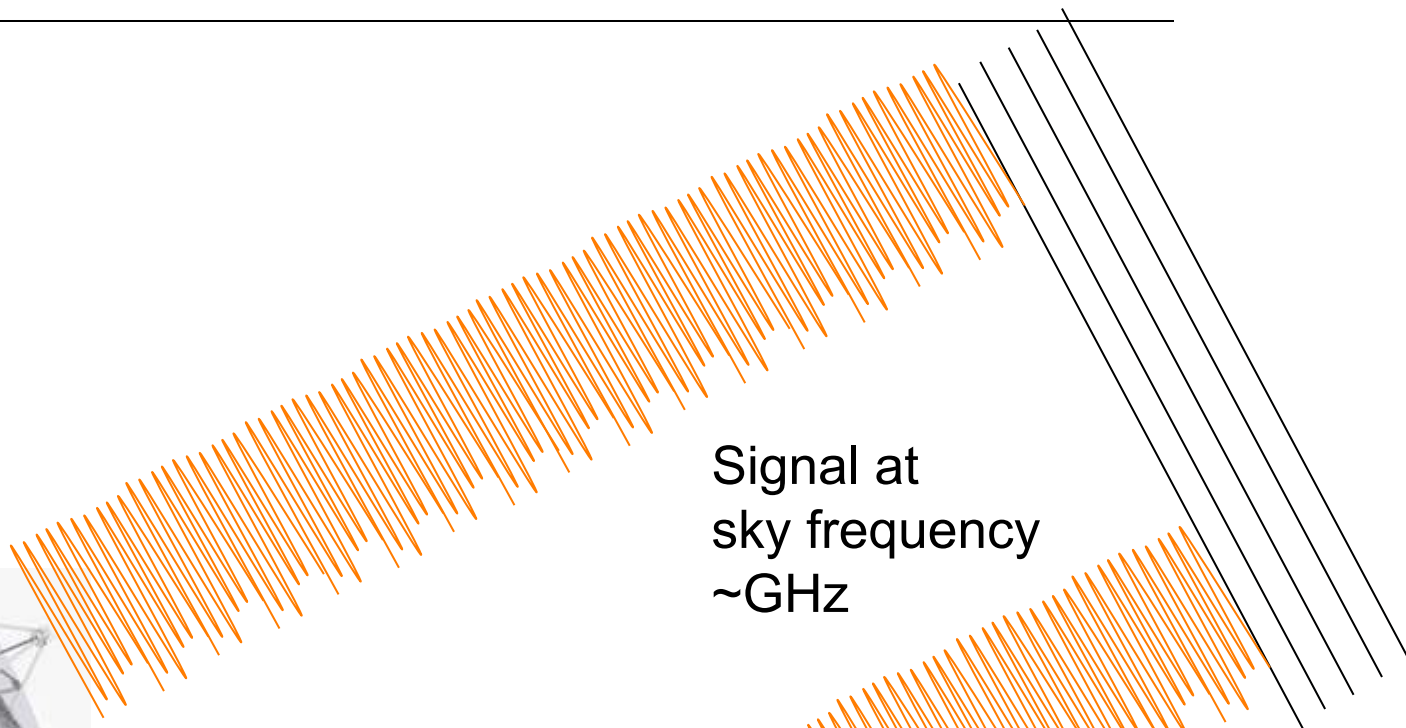


Fringe rotation



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Signal at
sky frequency
 \sim GHz

Downconversion

Signal at
baseband ~ 0 Hz





Fringe rotation

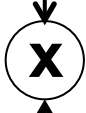
- Implementation: rotate phase using complex multiplier
- $\Delta\phi = 2\pi \nu_{lo} \tau_g$ ν_{lo} = local oscillator frequency,
 τ_g = applied delay
- Update rate of $\Delta\phi$ depends on how fast τ_g changes:
 - If τ_g is changing fast, update every sample in the time domain
 - For shorter baseline / low frequency instruments, can do post-channelisation or even post-accumulation





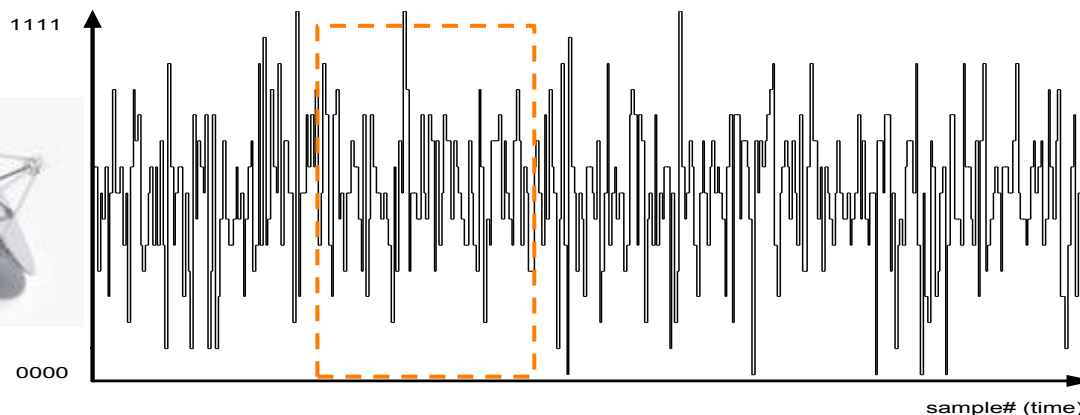
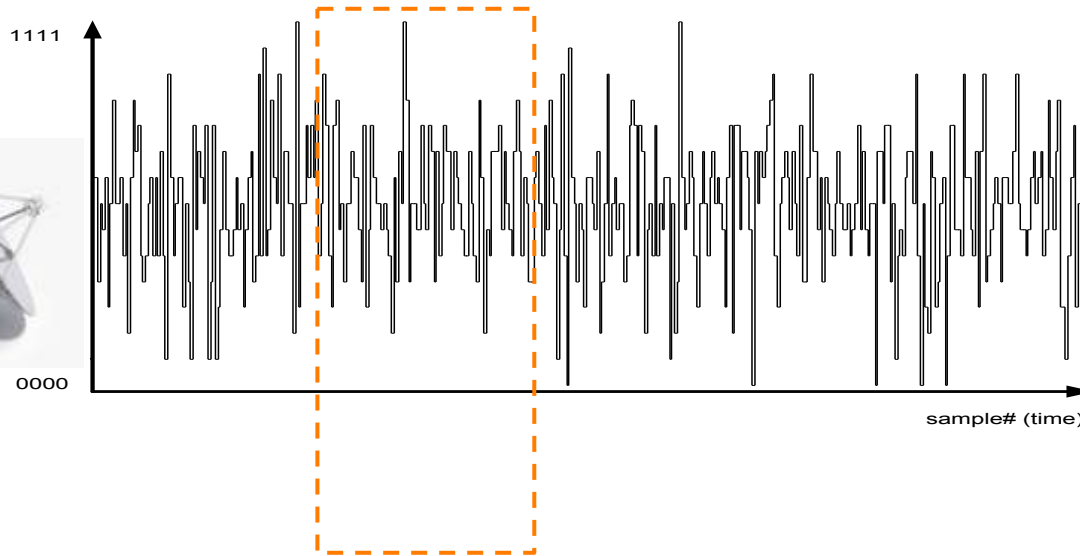
Alternate implementation

- We have shown how to build a practical FX correlator, which first Fourier transforms and then multiplies
- Convolution theorem: **Multiplication** in the frequency domain is equivalent to **convolution** in the time domain
- It is mathematically equivalent to convolve the two signals in the time domain and then Fourier transform



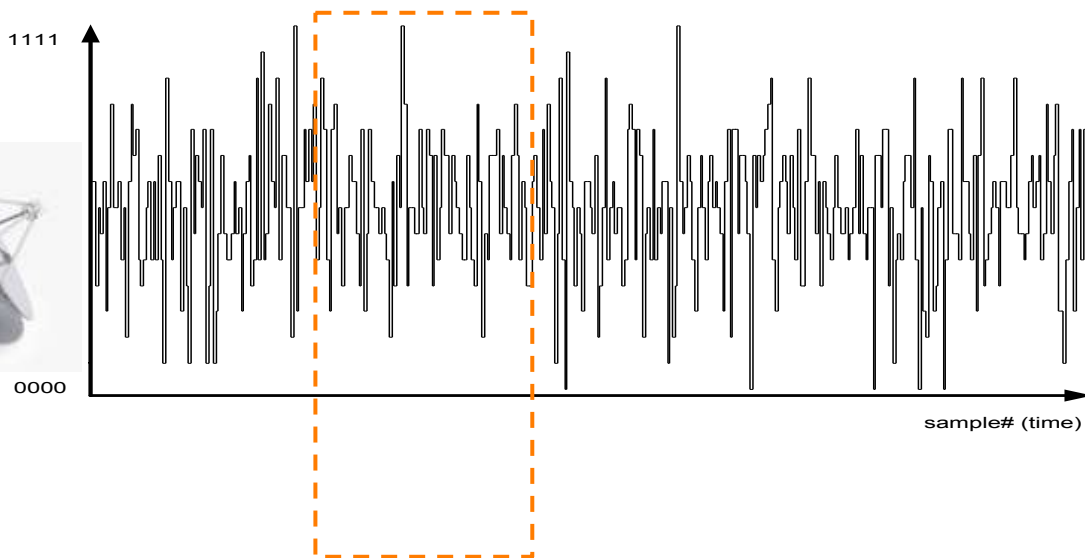


An equivalent “XF” correlator

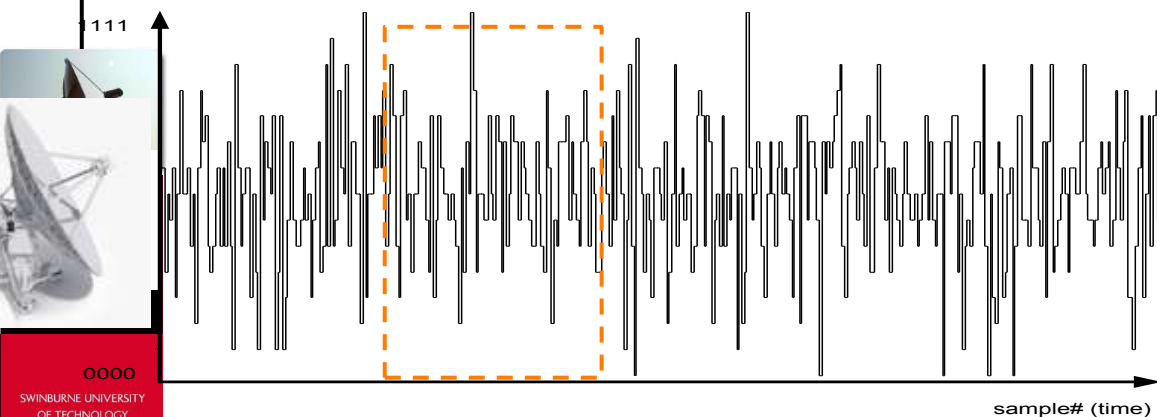




An equivalent “XF” correlator



X

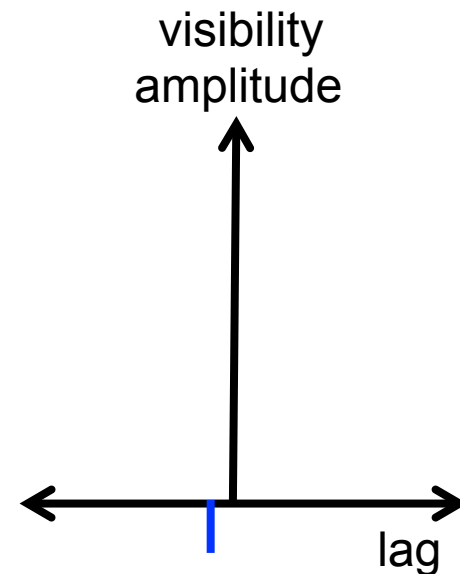
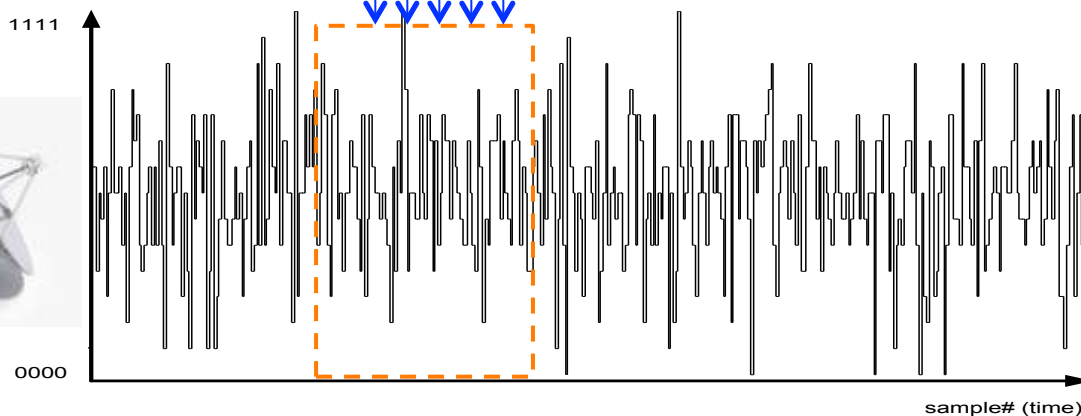
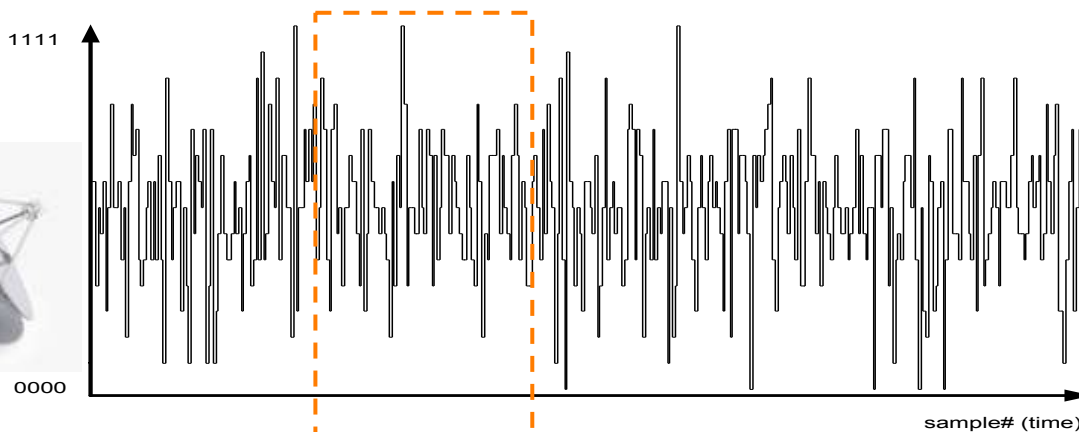


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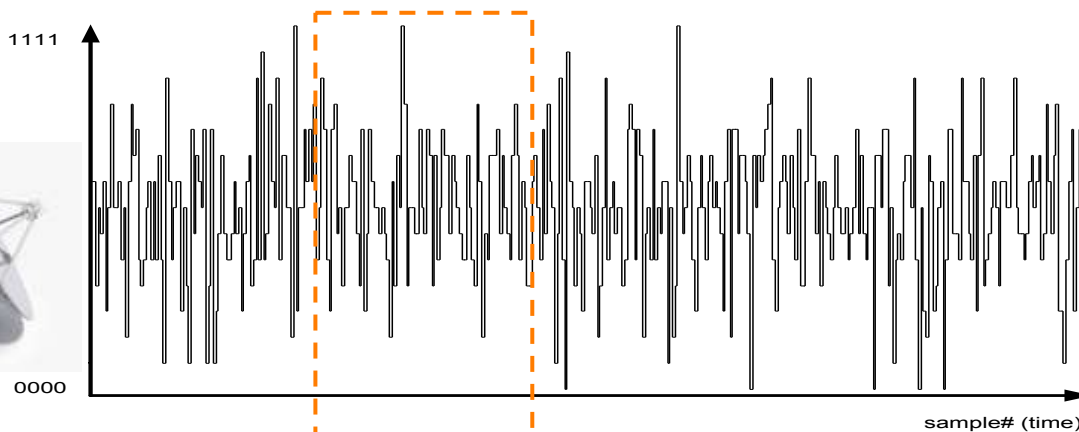


An equivalent “XF” correlator

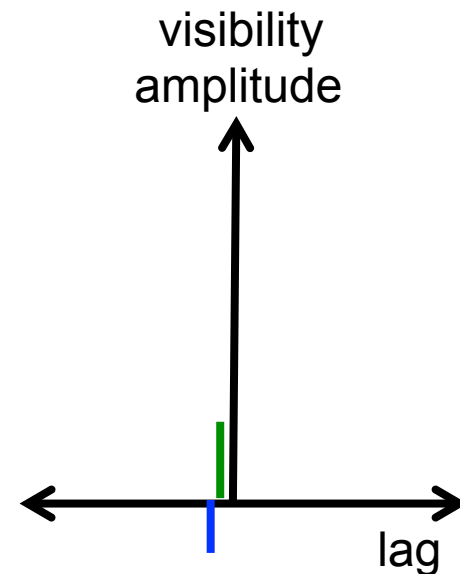
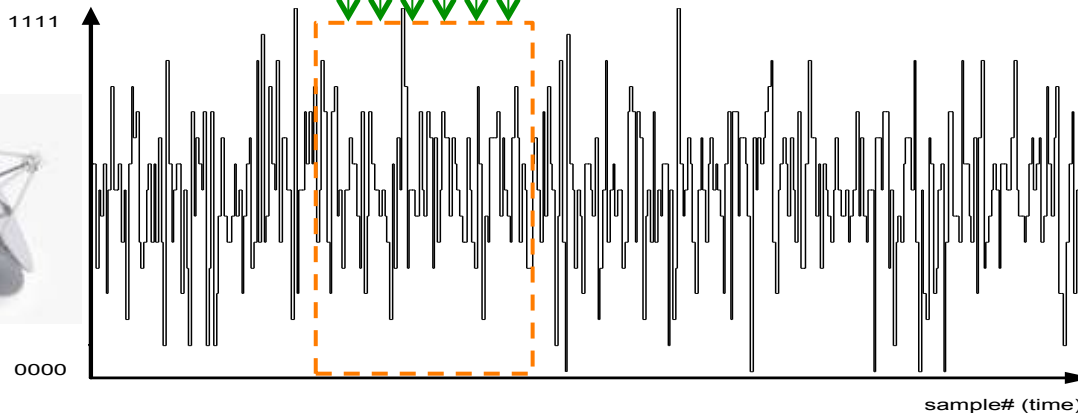




An equivalent “XF” correlator



Multiply
& accum.

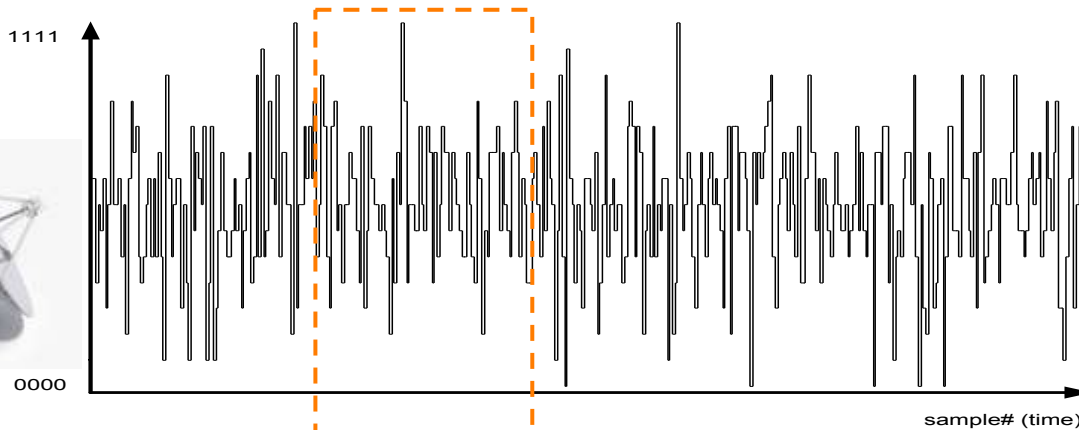


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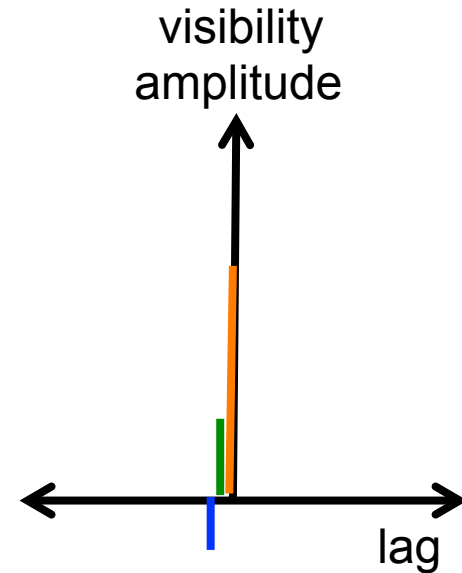
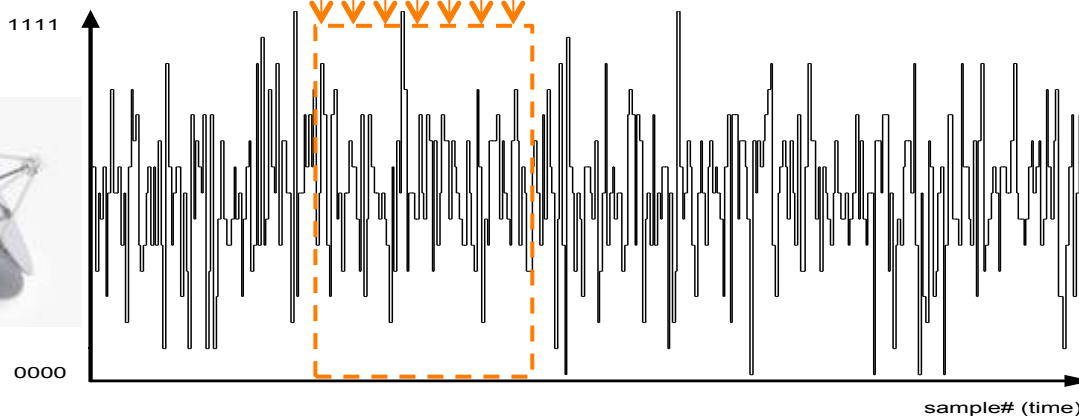
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An equivalent “XF” correlator

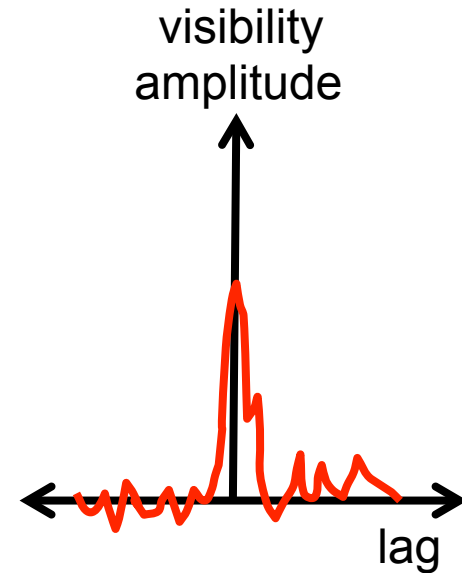
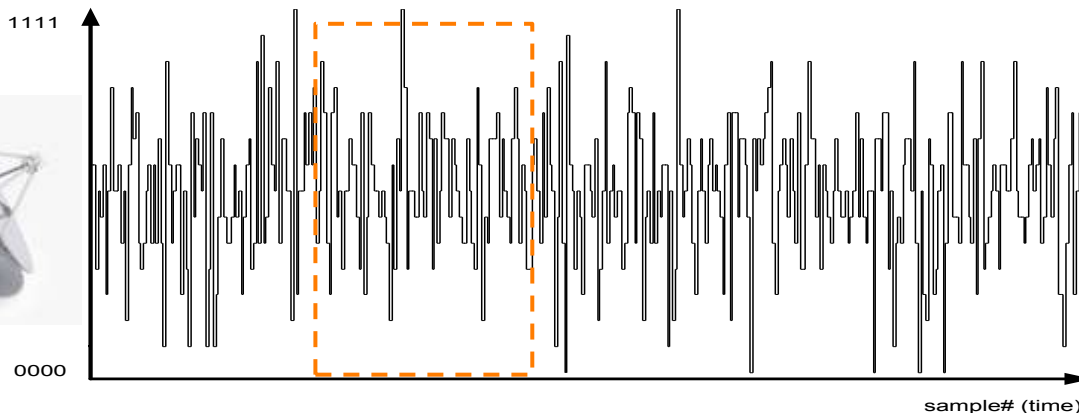
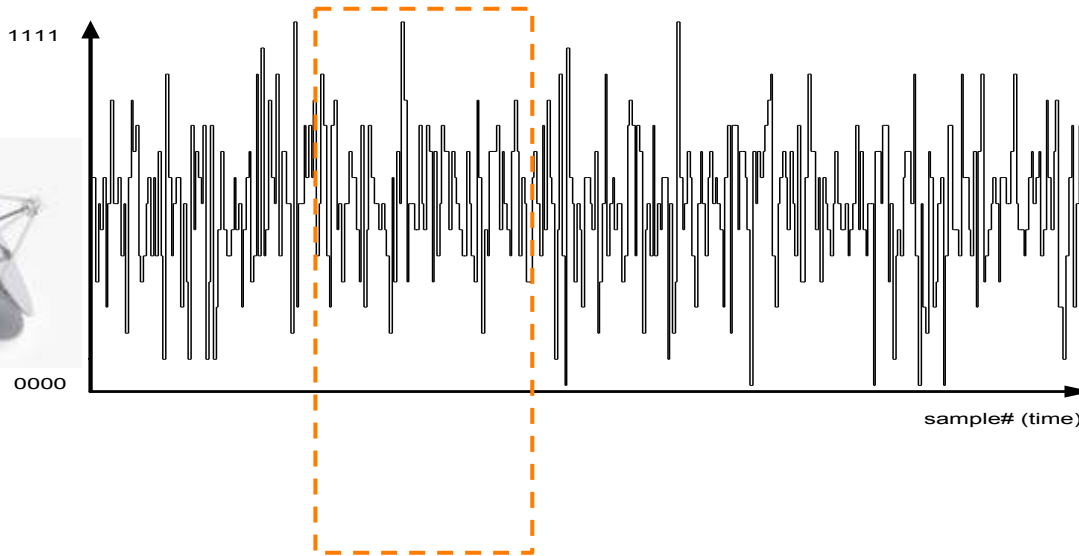


Multiply
& accum.



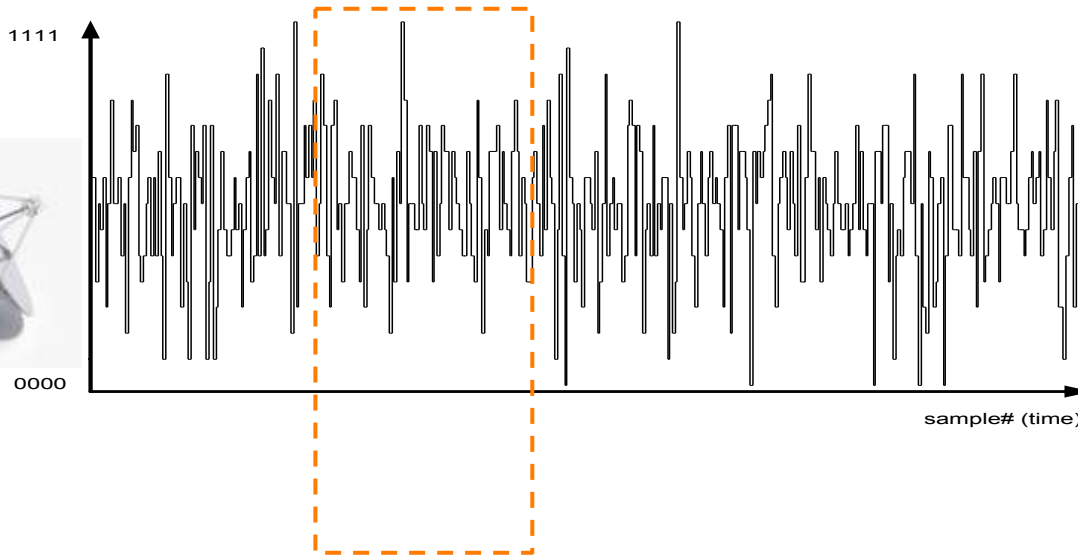


An equivalent “XF” correlator

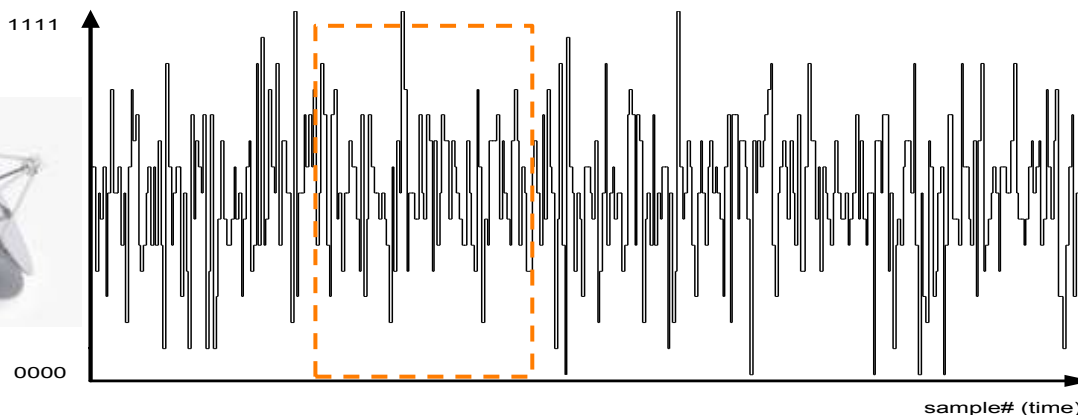
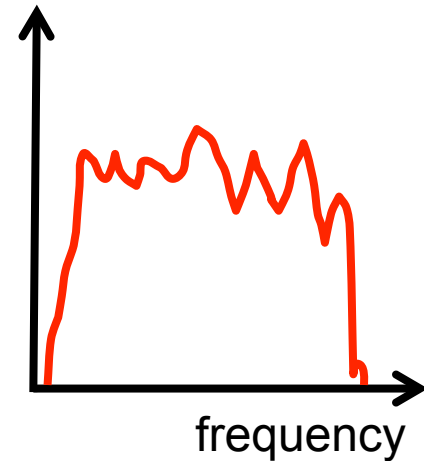




An equivalent “XF” correlator

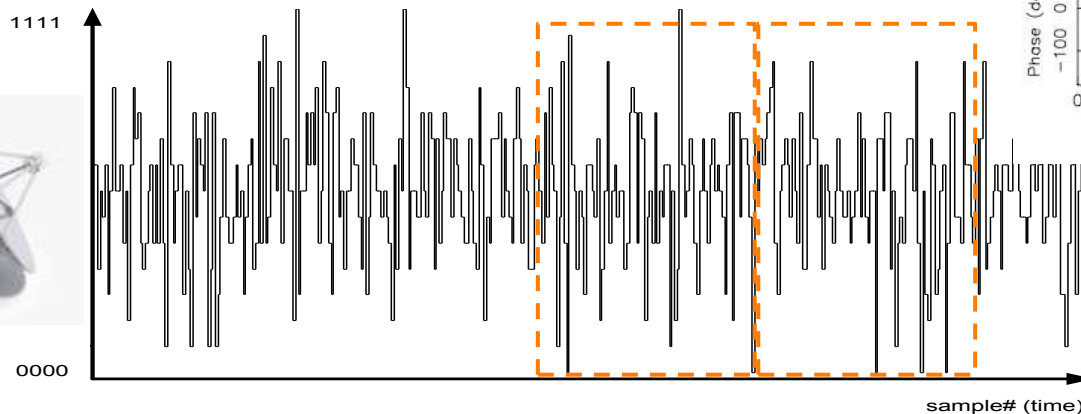
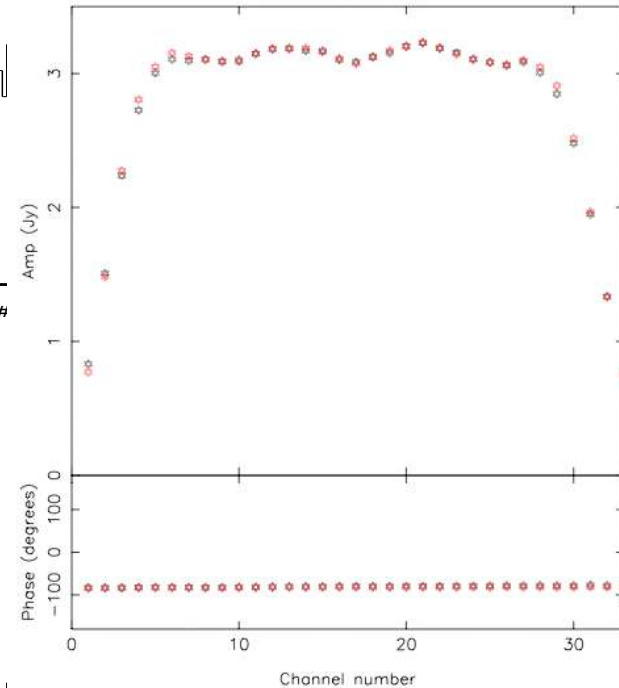
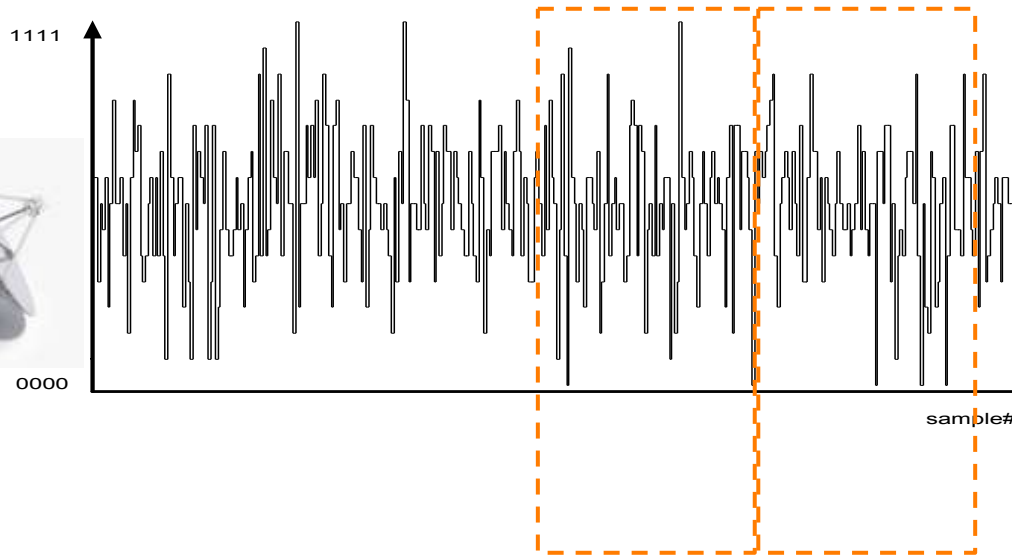


visibility
amplitude





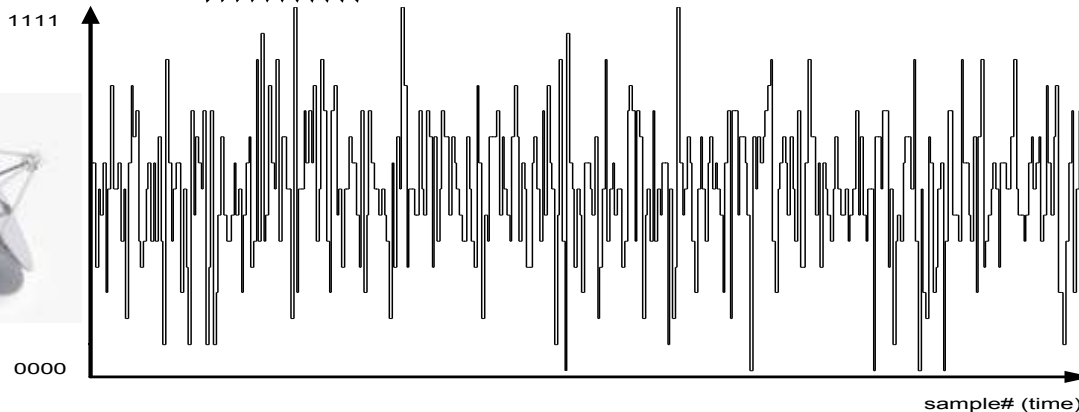
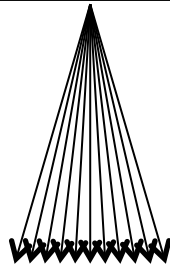
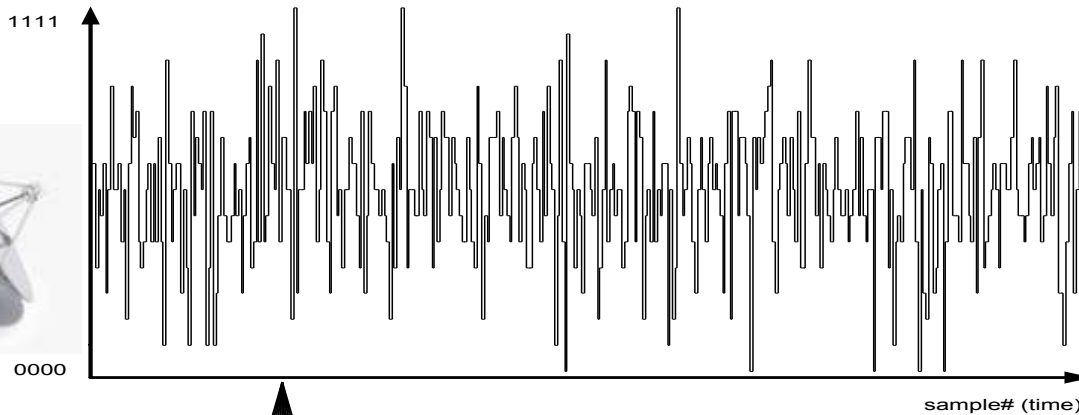
An equivalent “XF” correlator



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A realistic XF correlator



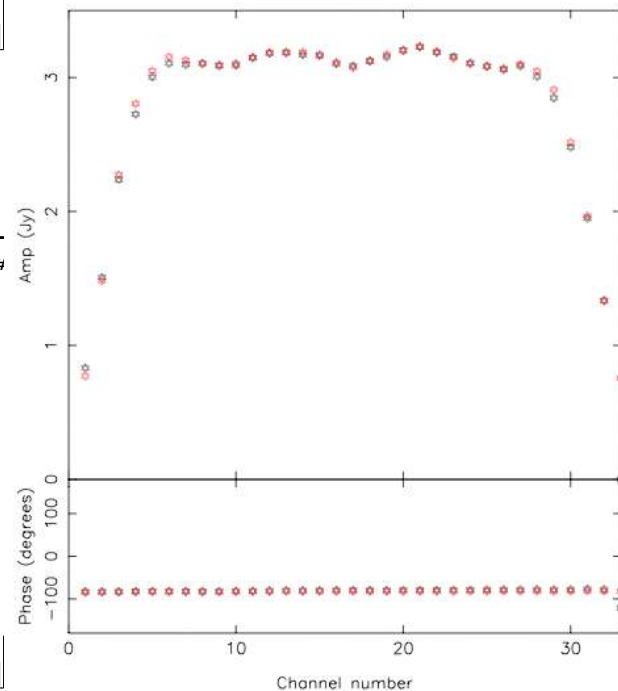
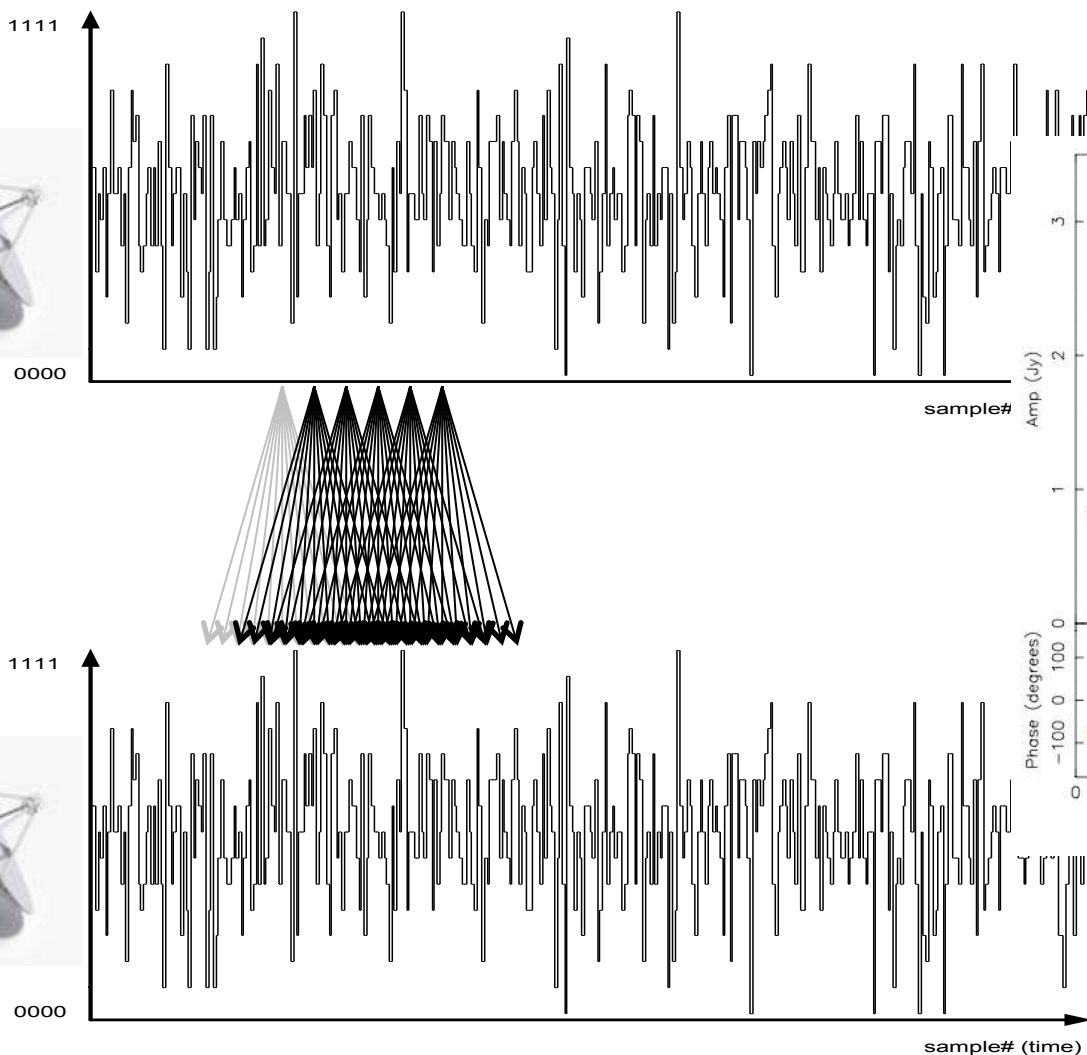
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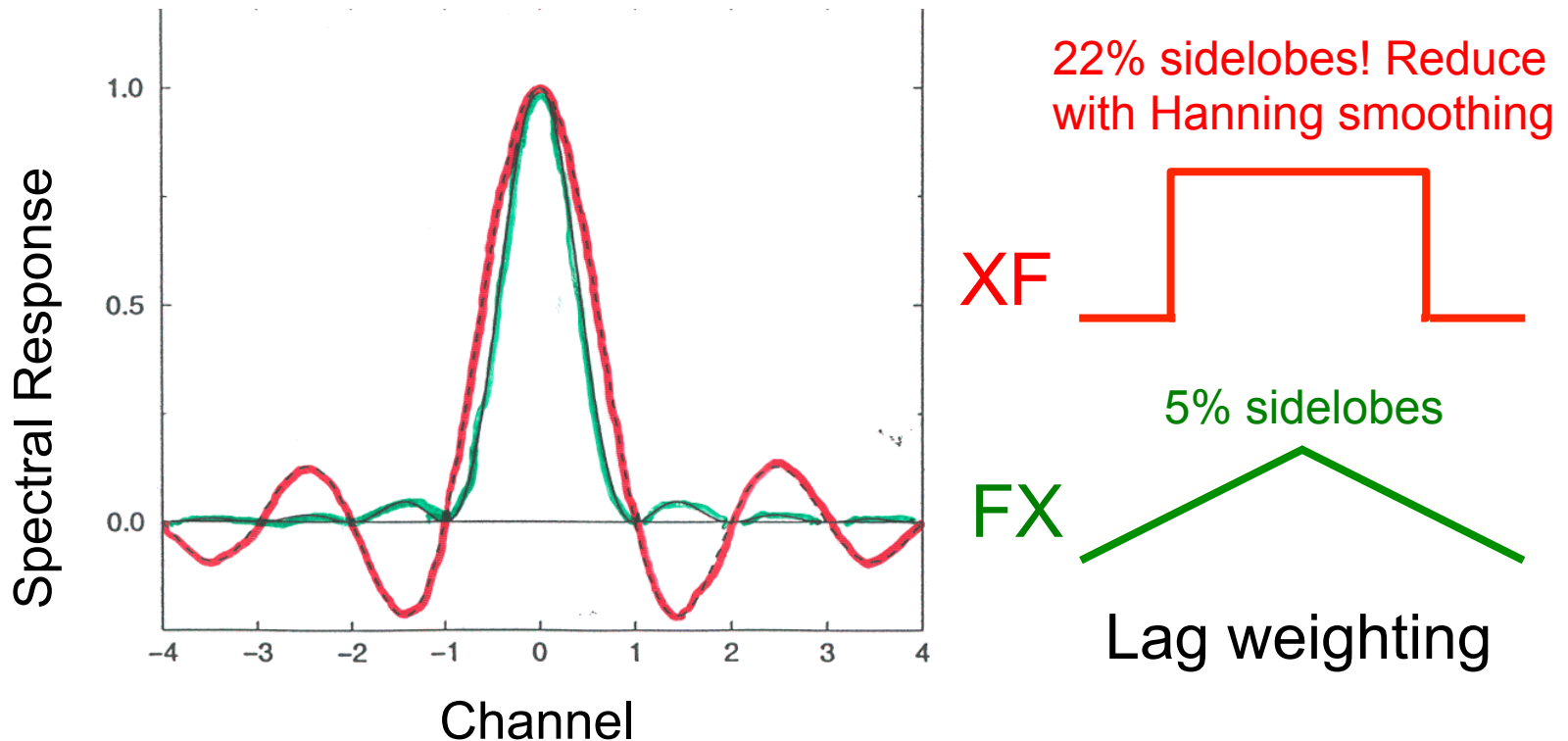
A realistic XF correlator





XF vs FX

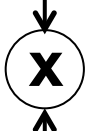
- Different windowing in time domain gives different spectral response





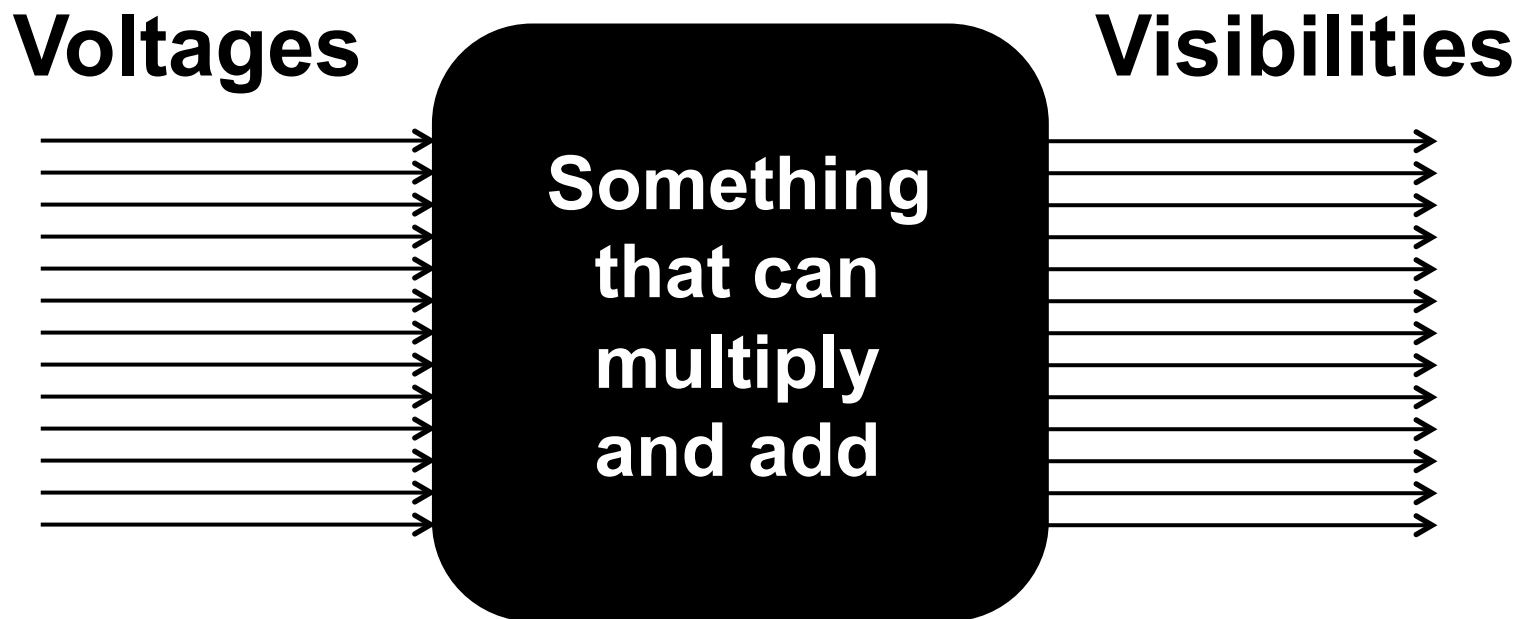
XF vs FX: which is better?

- Advantages and disadvantages to both
 - FX many fewer operations overall
 - XF can make use of very efficient low-precision integer multipliers up-front (but need special-purpose hardware)
 - FX: access to frequency domain at short timescale allows neat tricks and higher precision correction of delay effects
 - Modern correlators mostly FX-style, but use digital filterbank to channelise rather than FFT (shape channel response, contain RFI)





Correlator platforms



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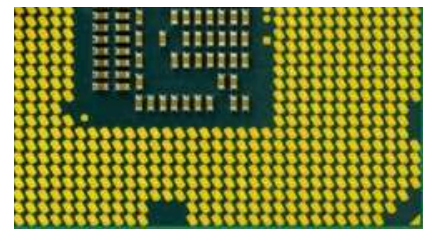
Correlators on CPUs



```
status = vectorFFT_CtoC_cf32(complexunpacked, fftd, pFFTSpecC, fftbuffer);  
if(status != vecNoErr)  
    csevere << startl << "Error doing the FFT!!!" << endl;
```

...

```
status = vectorAddProduct_cf32(vis1, vis2, &(scratchspace->threadcrosscorrs[result
```





Correlators on CPUs

- Many positive points:
 - Can implement in “normal” code (e.g., C++); maintainable, many skilled coders
 - Development effort transferrable across generations of hardware
 - Incremental development is trivial
 - Natively good at floating point (good for FX), no cost to do high precision
- One major disadvantage:
 - CPUs not optimised for correlation; big system like JVLA would take **many** CPUs.



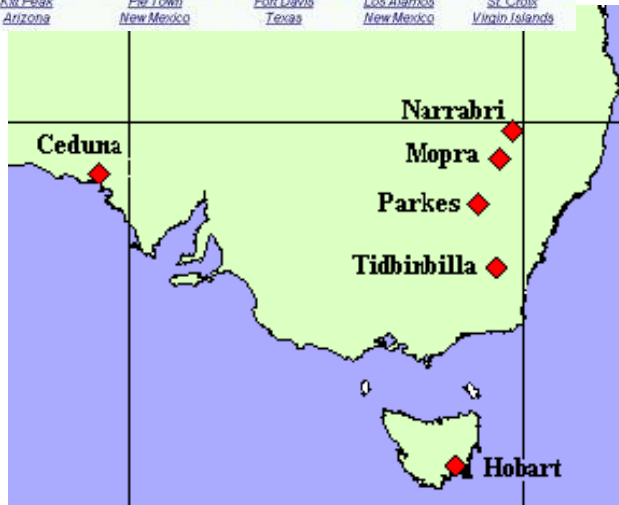
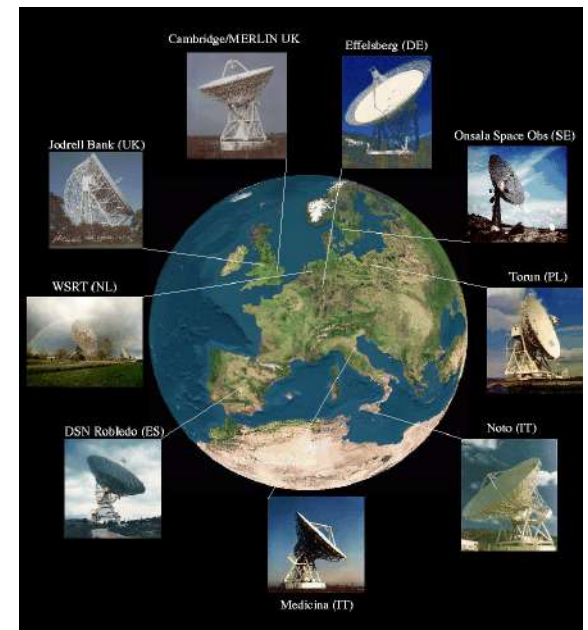


Correlators on CPUs



The Very Long Baseline Array,
10 stations

The European
VLBI Network,
~20 stations



The Long Baseline Array,
Australia, ~6 stations

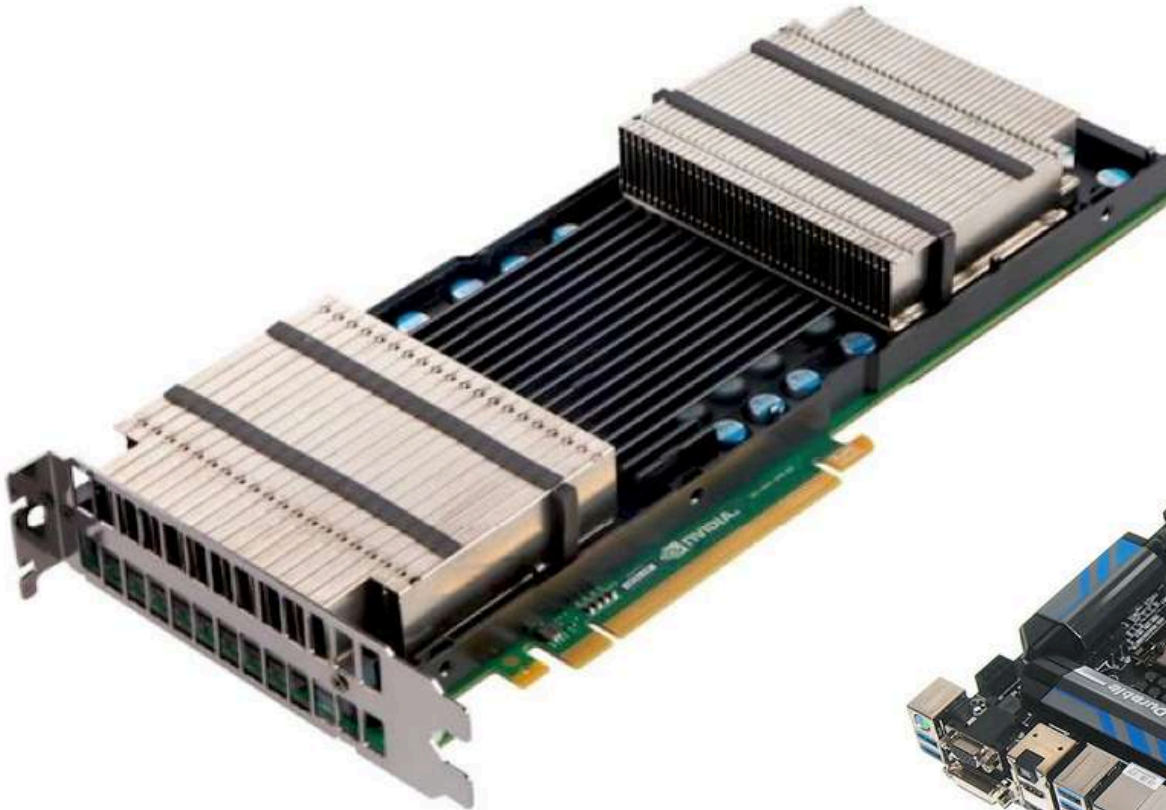
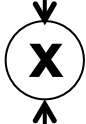


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Correlators on GPUs



Like CPUs, GPUs are mounted on a standard motherboard



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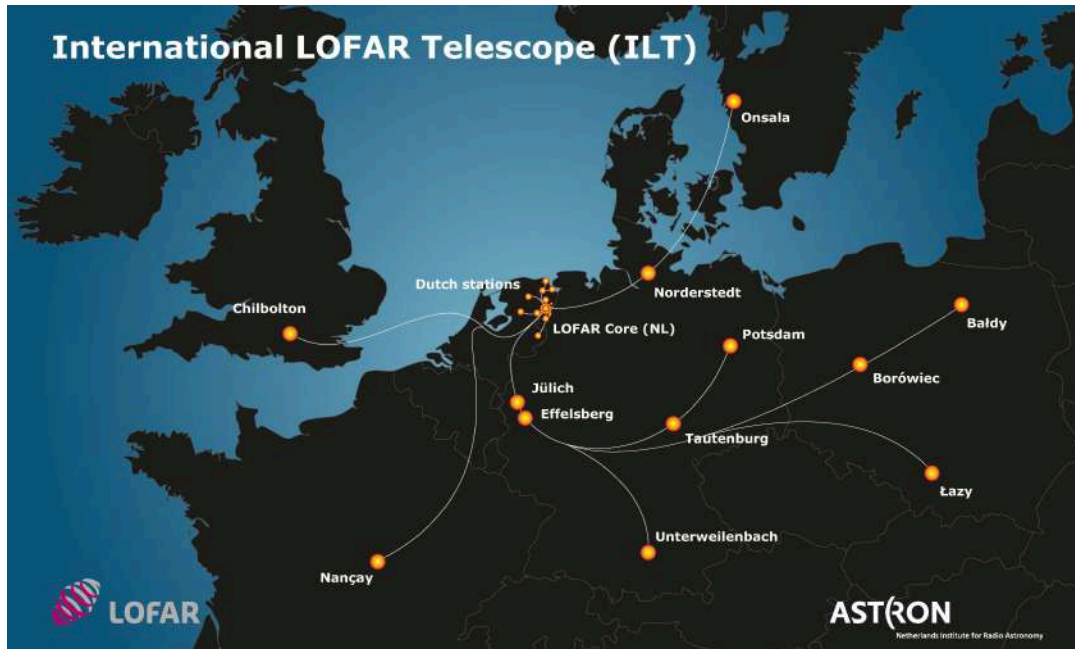
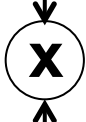
Correlators on GPUs

- Advantages:
 - More powerful and more efficient than CPUs
 - Also good at floating point
- Disadvantages:
 - Writing code is more difficult (GPUs are more specialized, less flexible: need to carefully manage data transfers)
 - Fewer trained GPU programmers available
 - Transfer-ability of code across hardware generations harder (capabilities change faster, need new code to use)





Correlators on GPUs



GMRT, India,
30 stations

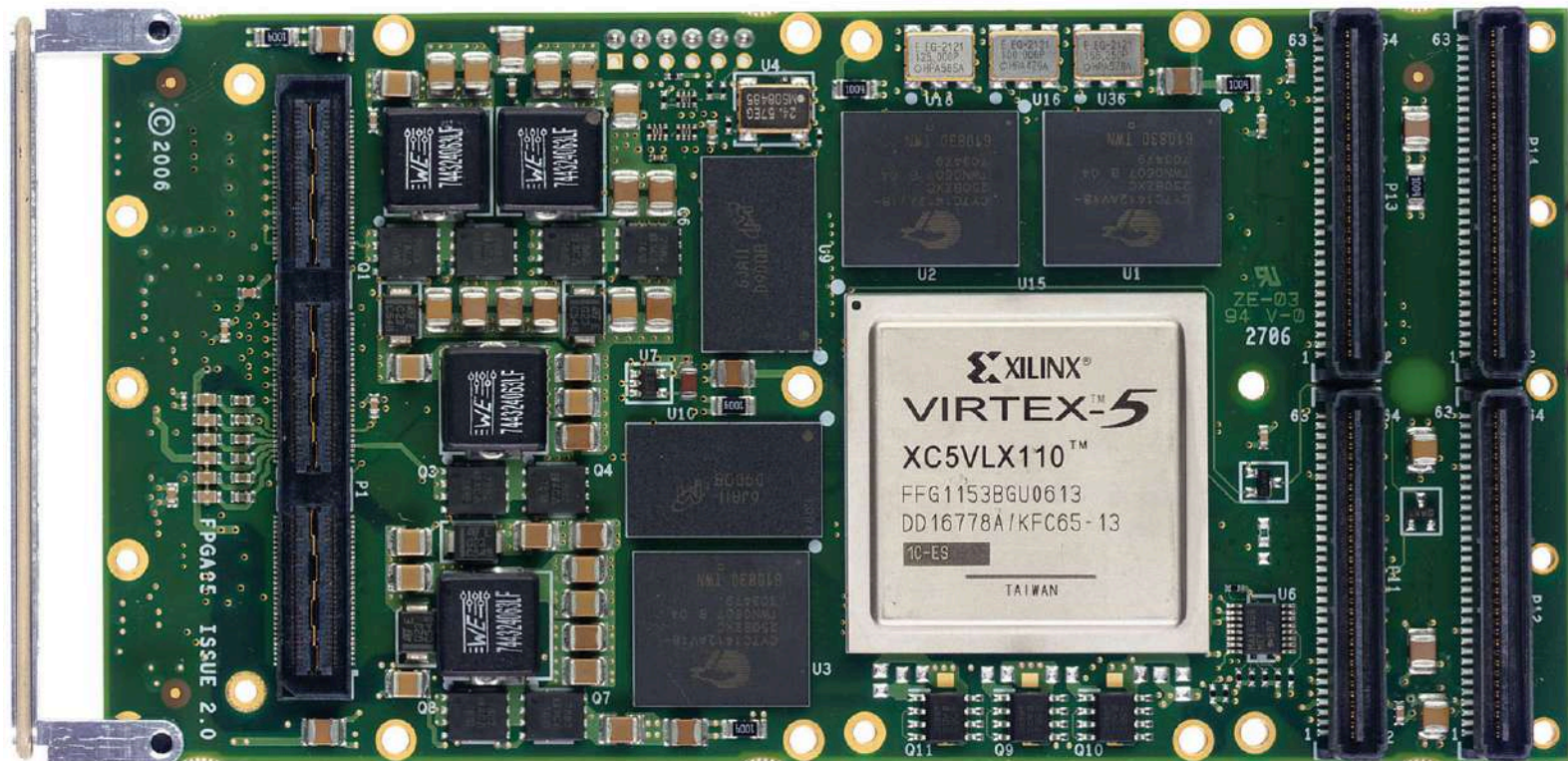
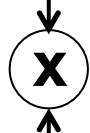


The Low Frequency Array
(LOFAR), 73 stations





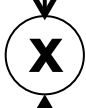
Correlators on FPGAs





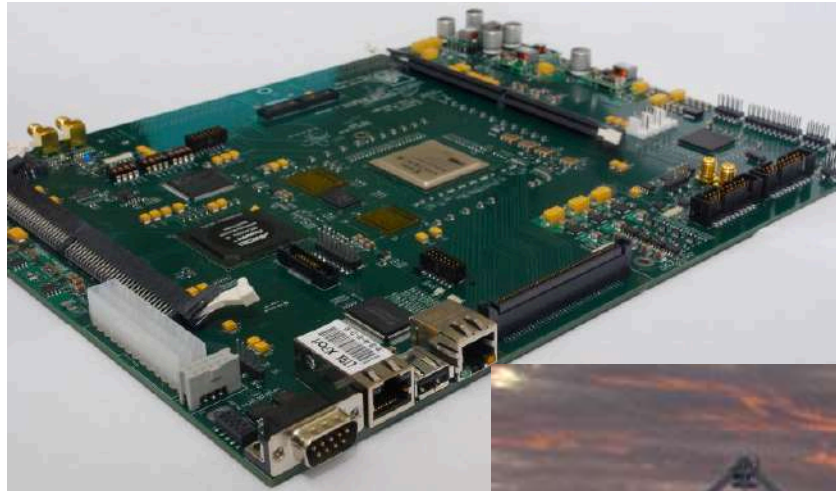
Correlators on FPGAs

- Advantages:
 - More efficient than CPUs or GPUs, particularly for integer multiplication
- Disadvantages:
 - Programming is harder again (especially debugging), yet fewer trained people
 - Transfer-ability across hardware generations even more limited
 - Synchronous (clocked) system, less robust to perturbations c.f. CPUs/GPUs





Correlators on FPGAs



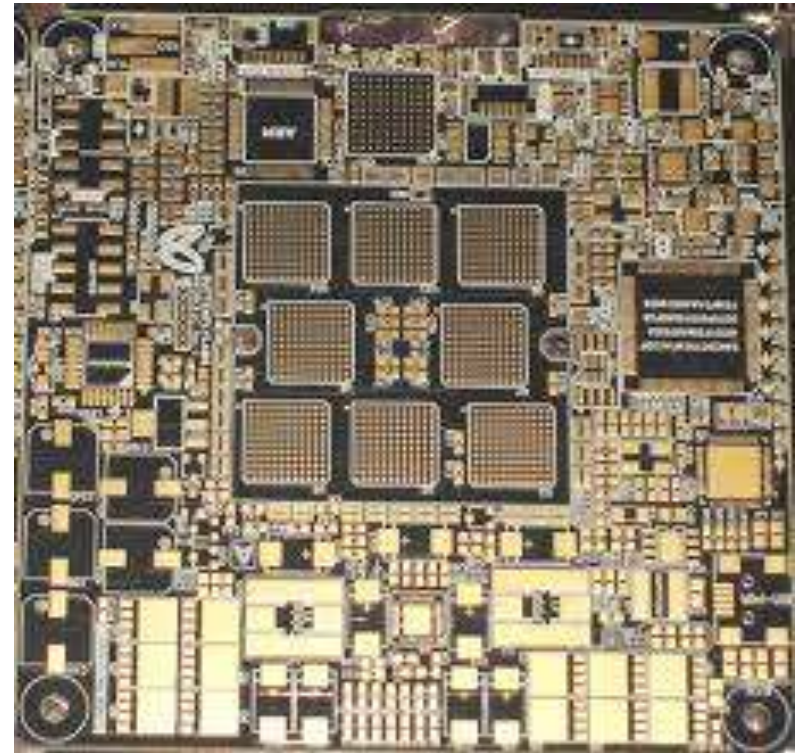
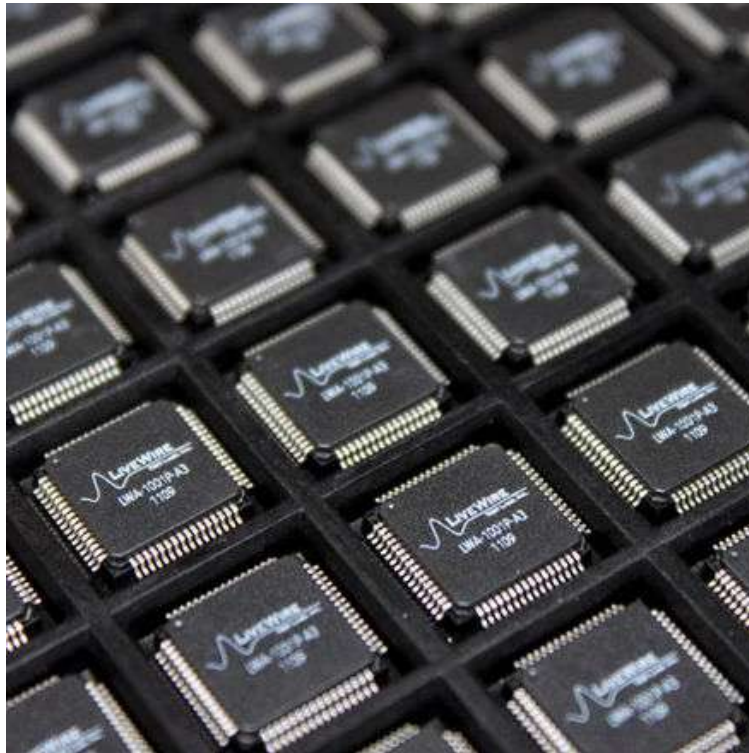
“Roach” reconfigurable
FPGA board used for
correlation



MeerKAT, 64 dishes, under construction



Correlators on ASICs



As with FPGAs, ASICs are mounted on boards





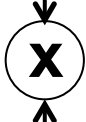
Correlators on ASICs

- Advantages:
 - Highest possible efficiency, low per-unit cost
- Disadvantages:
 - Highest development cost (time and manufacturing setup)
 - Specialized knowledge required
 - Can't be changed / very difficult to upgrade during lifetime





Correlators on ASICs



The Very Large Array,
New Mexico

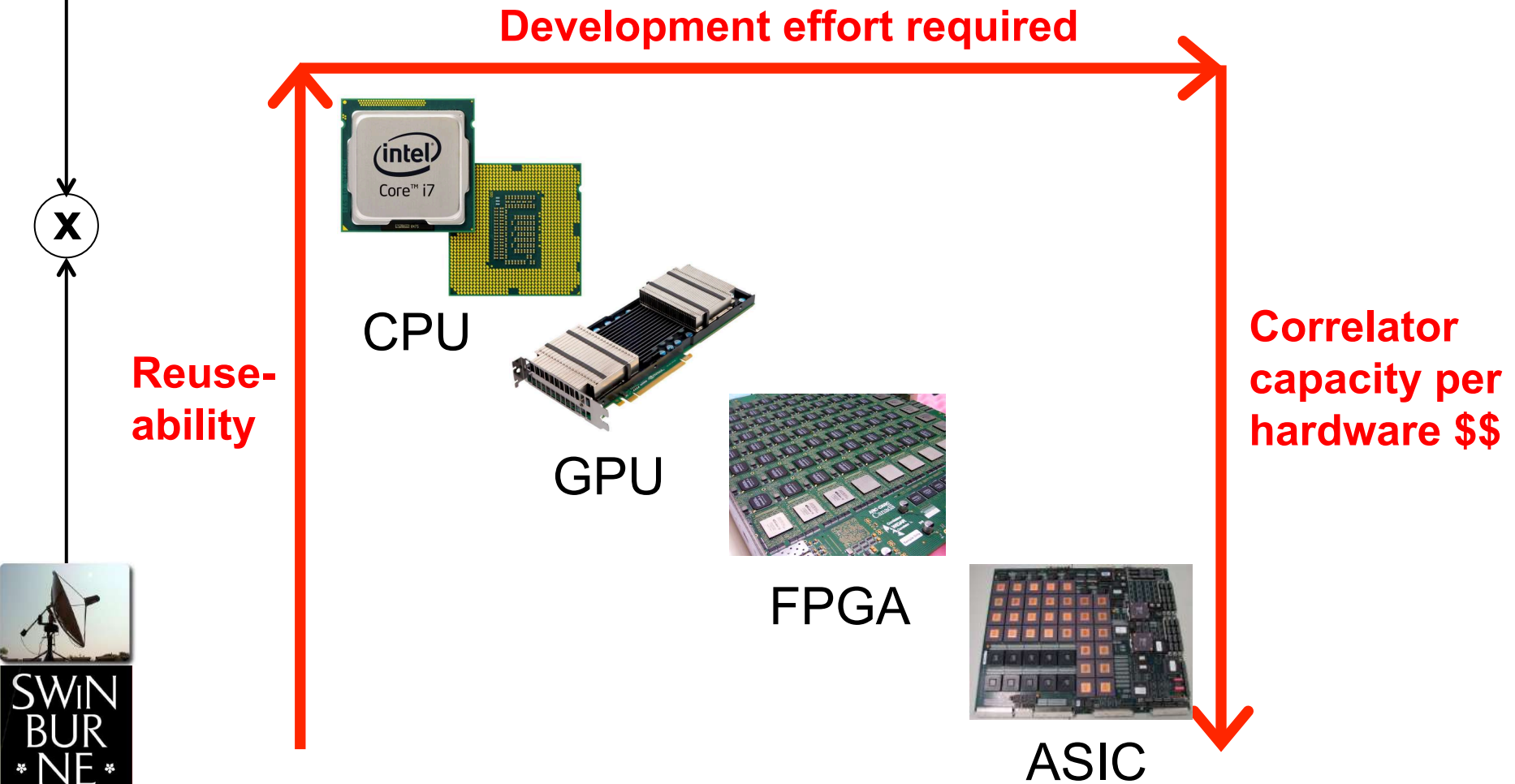


The Westerbork Synthesis
Radio Telescope, Netherlands





Correlator platform overview



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The end

