ORBITAL MECHANICS - MIDTERM

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Ex. 8

```
clear all
close all
%initial conditions
al=12000; %semi major axis [km]
                   %eccentricity
e1=.2;
i1=45*pi/180;
                  %inclination [rad]
                   %argument of perigee [rad]
omega1=0;
                   %RAAN [rad]
f1=0;
                    %true anomaly [rad]
p1=a1*(1-e1^2);
%final conditions
a2=13000;
e2=.1;
i2=45*pi/180;
w2 = 0;
omega2=0;
f2=120*pi/180;
p2=a2*(1-e2^2);
dt=60*60;
                    %transfer time [sec]
mu=398600;
                    %gravitational parameter [km<sup>3</sup>/s<sup>2</sup>]
```

Getting position and velocity from orbital elements

```
[r1vec,v1vec] = RVFromCOE( a1,i1,omega1,w1,e1,f1, mu );
[r2vec,v2vec] = RVFromCOE( a2,i2,omega2,w2,e2,f2, mu );
```

Solving Lambert's Problem

```
[v1t,v2t,dTheta] = LambertSolver( r1vec, r2vec, dt, mu );
```

Orbital Elements for the transfer orbit

[at,it,omegat,wt,et,ft,ht] = OrbitalElementsFromRV(rlvec, vlt, mu);

Calculating delta v's

```
dv1=norm(v1t-v1vec);
dv2=norm(-v2t+v2vec);
fprintf('deltaV 1 = %.2f km/s\ndeltaV 2 = %.2f km/s\n', dv1, dv2)
fprintf('Orbital elements of the transfer orbit\n')
fprintf('Semi major axis: %.0f km\n',at)
fprintf('Inclination: %.0fo\n',it*180/pi)
fprintf('Right ascension of the ascending node: %.0fo\n',omegat*180/
fprintf('Argument of periapsis: %.2fo\n',wt*180/pi)
fprintf('Eccentricity: %.2f\n',et)
deltaV 1 = 0.24 \text{ km/s}
deltaV 2 = 0.91 \text{ km/s}
Orbital elements of the transfer orbit
Semi major axis: 12993 km
Inclination: 45°
Right ascension of the ascending node: 0°
Argument of periapsis: 5.95°
Eccentricity: 0.26
```

COMMENTS

```
%{
The "LambertSolver" function provides a solution for Lambert's problem
using universal variables. More details are given in the
sub-functions. The method described in class is used, just normalized
to
universal variables so it is applicable to every conic section and it
takes
into account of particular cases and singularity. I attach every
sub-function used for that.

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%}
```

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