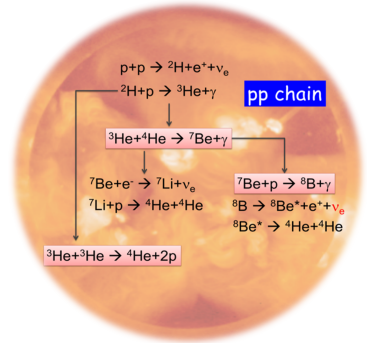


# *Ab initio* calculation of the ${}^3\text{He}(\alpha, \gamma){}^7\text{Be}$ astrophysical $S$ factor

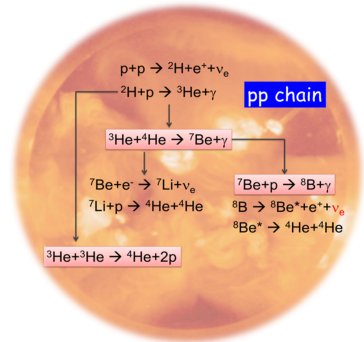
Mack C. Atkinson



# ${}^3\text{He}(\alpha, \gamma){}^7\text{Be}$ important for solar-model predictions

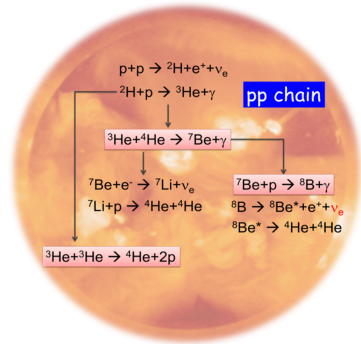
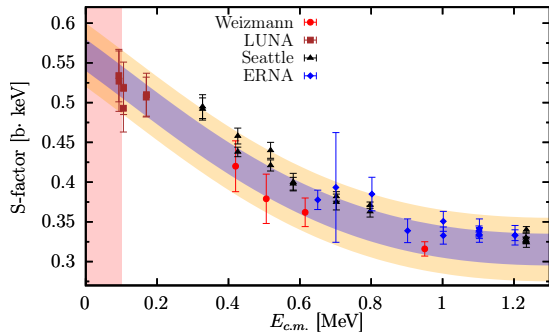


# $^3\text{He}(\alpha, \gamma)^7\text{Be}$ important for solar-model predictions



$$\sigma(E) = \frac{S_{34}(E)}{E} \exp \left\{ -\frac{2\pi Z_1 Z_2 e^2}{\hbar \sqrt{2E/m}} \right\}$$

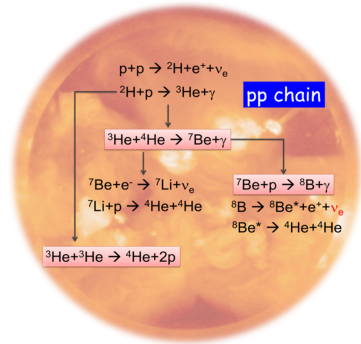
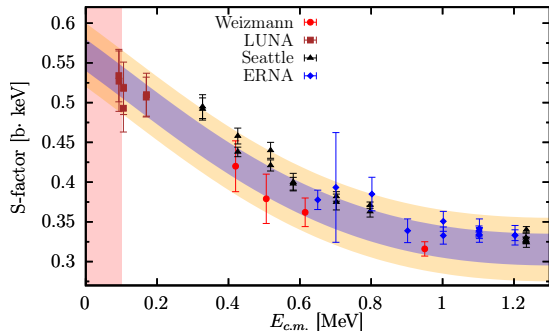
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# ${}^3\text{He}(\alpha, \gamma){}^7\text{Be}$ important for solar-model predictions

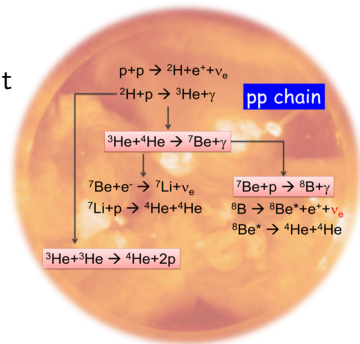
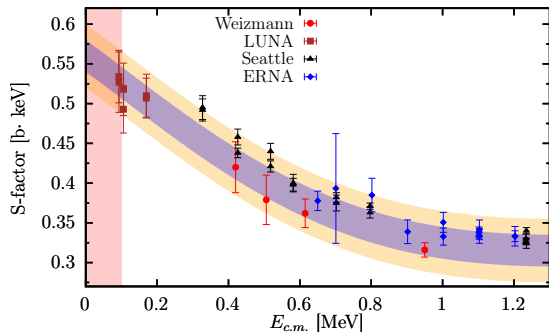
- Reaction rates too low at solar energies in the lab



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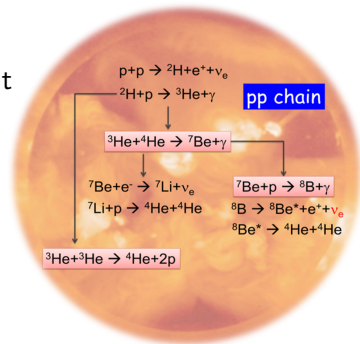
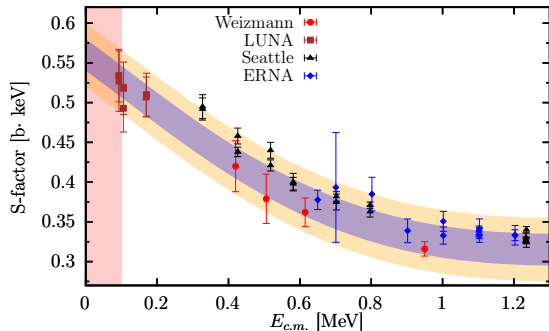
- Reaction rates too low at solar energies in the lab
- Current evaluations depend on both theory and experiment



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# $^3\text{He}(\alpha, \gamma)^7\text{Be}$ important for solar-model predictions

- Reaction rates too low at solar energies in the lab
- Current evaluations depend on both theory and experiment
- Ideally, theory will accurately predict  $S_{34}(0)$



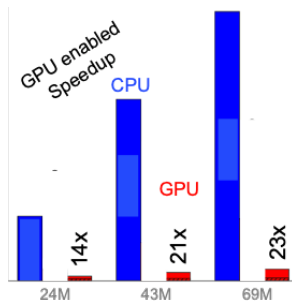
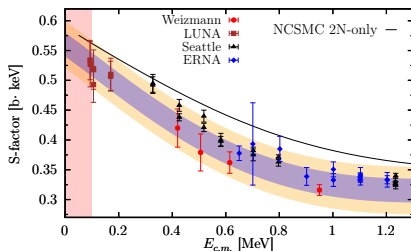
$$\sigma(E) = \frac{S_{34}(E)}{E} \exp \left\{ -\frac{2\pi Z_1 Z_2 e^2}{\hbar \sqrt{2E/m}} \right\}$$

# Goal: Reduce the theoretical uncertainty in the determination of $S_{34}(0)$

Current evaluation:

$$S_{34}(0) = 0.56 \pm 0.02(\text{expt.}) \pm \mathbf{0.02}(\text{theor.})$$

- **How?:** Perform an *ab initio* calculation of the  ${}^3\text{He}(\alpha, \gamma){}^7\text{Be}$  reaction
  - Previously only possible using 2N forces



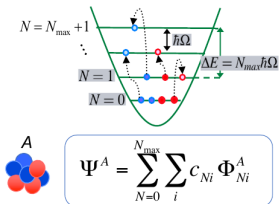
- I present the first *ab initio* calculation of  ${}^3\text{He}(\alpha, \gamma){}^7\text{Be}$  including 3N forces



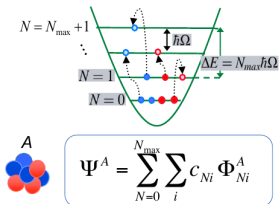
# No-Core shell model (NCSM)

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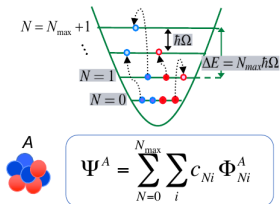


# No-Core shell model (NCSM)



- Construct the Hamiltonian from  $\chi_{EFT}$

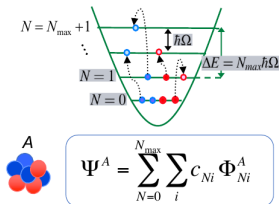
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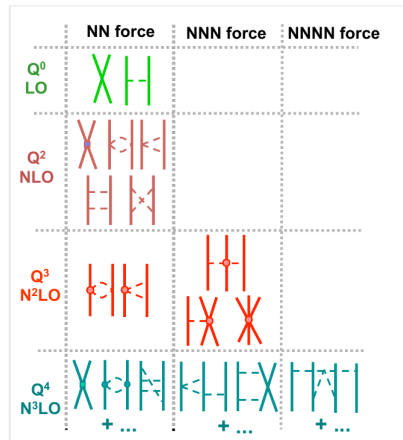
$$\hat{H} = \hat{T} + \hat{V}_{2N} + \hat{V}_{3N}$$

# No-Core shell model (NCSM)

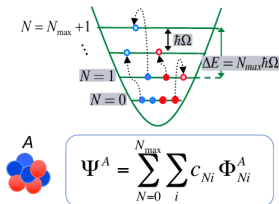


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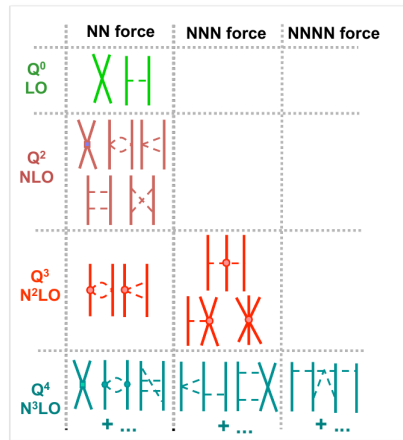
# No-Core shell model (NCSM)



- Construct the Hamiltonian from  $\chi_{EFT}$

$$\hat{H} = \hat{T} + \hat{V}_{2N} + \hat{V}_{3N}$$

$$\hat{H} |\Psi^A\rangle = E |\Psi^A\rangle$$



# No-core shell model with continuum (NCSMC)

$$\Psi^{(A)} = \sum_{\lambda} c_{\lambda} \left| \begin{array}{c} (A) \\ \text{NCSMC} \end{array} \right\rangle + \sum_{\nu} \int d\vec{r} \gamma_{\nu}(\vec{r}) \hat{A}_{\nu} \left| \begin{array}{c} \text{NCSMC} \\ (A-a) \end{array} \right\rangle, \nu \rangle$$



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$$\left| \begin{array}{c} (A) \\ \text{cluster} \end{array}, \lambda \right\rangle = |^7\text{Be}\rangle$$

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- $|^7\text{Be}\rangle$ ,  $|^3\text{He}\rangle$ , and  $|\alpha\rangle$  from NCSM

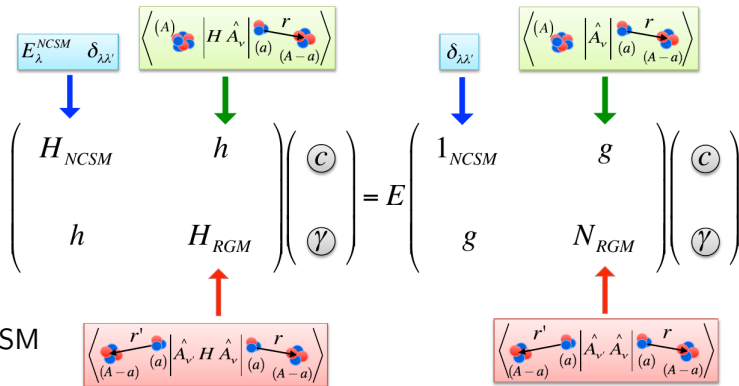
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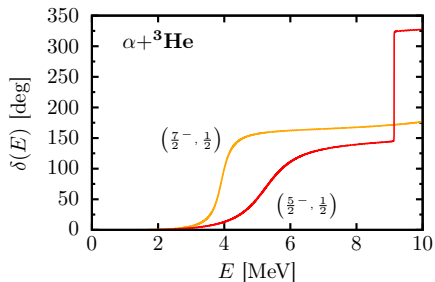
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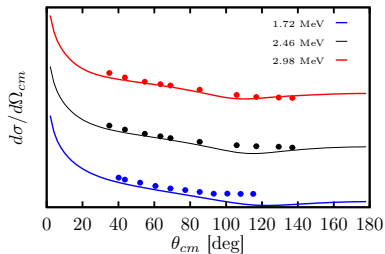
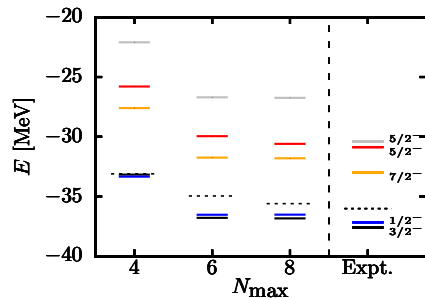
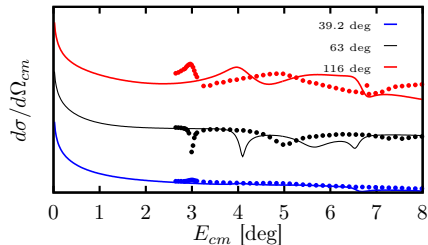
- $|^7\text{Be}\rangle$ ,  $|^3\text{He}\rangle$ , and  $|\alpha\rangle$  from NCSM



# NCSMC Calculation of $^3\text{He}+^4\text{He}$ shows reasonable agreement with data



NN-n3lo+3N1nl SRG1.9 hw20 he3+he4 Nmax8

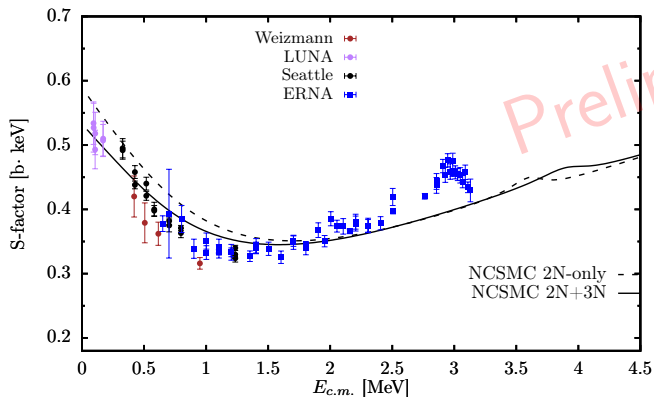


NN-N3LO+3N1nl

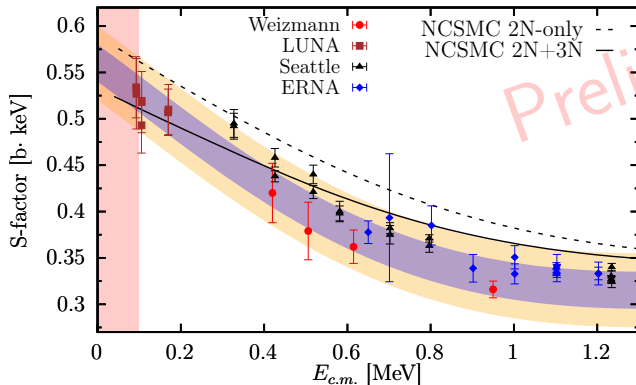
$\hbar\Omega = 20$  MeV

$\lambda_{\text{SRG}} = 2.0 \text{ fm}^{-1}$

Results are promising but convergence needs to be explored



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Preliminary

# Summary and Outlook

- Included 3N forces in the calculation of  $4\text{He}+3\text{He}$  using the boost formalism
- Results are promising but convergence needs to be explored



# Thanks

- Sofia Quaglioni
- Kostas Kravvaris
- Petr Navratil
- Guillaume Hupin