ARTS Developer Guide

edited by

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The content and usage of ARTS are not only described by this document. An overview of ARTS documentation and help features is given in *ARTS User Guide*, Section 1.2. For continuous reports on changes of the source code and this user guide, subscribe to the ARTS developers mailing list at http://www.sat.ltu.se/arts/support/.

We welcome gladly comments and reports on errors in the document. Send then an e-mail to: patrick.eriksson (at) chalmers.se or sbuehler (at) ltu.se.

If you use data generated by ARTS in a scientific publication, then please mention this and cite the most appropriate of the ARTS publications that are summarized on http://www.sat.ltu.se/arts/docs/.

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Chapter 1

The art of developing ARTS

The aim of this section is to describe how the program is organized and to give detailed instructions how to make extensions. That means, it is addressed to the ARTS developers, not the users. If you only want to use ARTS, you should not need to read it. But if you want to make changes or additions, you should definitely read this carefully, since it can safe you a lot of work to understand how things are organized.

1.1 Organization

ARTS is written in C++ and uses the cross-platform, open-source build system CMake (http://www.cmake.org/). It is organized in a similar manner as most GNU packages. The top-level ARTS directory is either called arts or arts-x.y, where x.y is the release number. It contains various sub-directories, notably doc for documentation, src for the C++ source code, and aii for the IDL interface. The document that you are reading right now, the ARTS Developer Guide, is located in doc/uguide.

There are two different versions of the ARTS package: The development version and the end-user version. Both contain the complete source code, the only difference is that the developers version also includes the SVN housekeeping data. If you want to join in the ARTS development (which we of course encourage you to do), you should write an email to the ARTS Development mailing list. Furthermore, for serious development work you need a computer running Unix, CMake, LaTeX, and the Doxygen program. All this is freely and easily available on the Internet, and, what is more, all these tools are included in the standard Linux distributions like Ubuntu or openSUSE.

The end-user version contains everything that you need in order to compile and install ARTS in a fairly automatic manner. The only thing you should need is an ANSI-C++ compiler, and the CMake utility. Please see files arts/README and arts/INSTALL for installation instructions. We are developing with the GNU C++ compiler, no other compilers have been tried so far.

| History | |
|---------|---|
| 020425 | Stefan Buehler: Put this part back in the AUG. Updated. |
| 000728 | Stefan Buehler: Added stuff about build system and howto cut a release. |
| 000615 | Created by Stefan Buehler. |

1.2 The ARTS build system

As mentioned above, CMake is used to construct the ARTS build system. A good introduction to the CMake system can be found in:

```
http://www.cmake.org/cmake/project/about.html
```

Using this tool makes a lot of things very easy, but also some things slightly more complicated.

The support for documentation other than info and man pages is not very good in the GNU system, so we had to use some tricks to make sure that the Doxygen automatic documentation and the User Guide work as they should.

1.2.1 Configure options

Here are some interesting options for CMake:

- -DCMAKE_BUILD_TYPE=Release: Removes '-g' from the compiler flags and includes #define NDEBUG 1 in config.h. The central switch to turn off all debugging features (index range checking for vectors, the trace facility, assertions,...). The default configuration RelWithDebInfo turns on some optimizations and provides debugging information.
- **-DCMAKE_BUILD_TYPE=Debug:** This switch turns off all optimizations. This should only be used if the default configuration makes debugging a certain problem difficult.
- **-DNO_OPENMP=1:** Disables the generation of multi-threaded code. CMake tries to detect if the compiler supports OpenMP and enables it by default.

1.3 Conventions

Here are some general rules for ARTS programming:

1.3.1 Numeric types

Never use float or double explicitly, use the type Numeric instead. This is set by CMake (to double by default). In the same way, use Index for all integers. It can take on positive or negative values and defaults to long. To change the default types, run cmake with the options -DINDEX=long or -DNUMERIC=double:

```
cmake -DINDEX=int --DNUMERIC=float ..
```

Note that changing the numeric type to a lower precision type than double might have unforseen impacts on the numerical precision and could lead to wrong results. In a similar way, reducing the index type can make it impossible to handle larger Vectors, Matrices or Tensors. The maximum range of the index type determines the maximum number of elements the container types can handle.

1.3 Conventions 3

1.3.2 Container types

Use Vector and Matrix for mathematical vectors and matrices (with elements of type Numeric). Use Array<something> to create an array of somethings. Commonly used Arrays have been predefined, they have names like ArrayOfString, ArrayOfMatrix, and so forth.

1.3.3 Terminology

Calculations are carried out in the so called workspace (WS), on workspace variables (WSVs). A WSV is for example the variable containing the absorption coefficients. The WSVs are manipulated by workspace methods (WSMs). The WSMs to use are specified in the controlfile in the same order in which they will be executed.

1.3.4 Global variables

Are not visible by default. To use them you have to declare them like this:

```
extern const Numeric PI;
```

which will make the global constant PI=3.14... available. Other important globals are:

full_name Full name of the program, including version.

parameters All command line parameters.

basename Used to construct output file names.

out_path Output path.

messages Controls the verbosity level.

wsv_data WSV lookup data.

wsv_group_names Lookup table for the names of types of WSVs.

WsvMap The map associated with wsv_data.

md_data WSM lookup data.

MdMap The map associated with md_data.

workspace The workspace itself.

species_data Lookup information for spectroscopic species.

SpeciesMap The map associated with species_data.

The only exception from this rule are the output streams out0 to out3, which are visible by default.

1.3.5 Files

Always use the open_output_file and open_input_file functions to open files. This switches on exceptions, so that any error occurring later on with this file will result in an exception. (Currently not really implemented in the GNU compiler, but please use it anyway.)

1.3.6 Version numbers

The package version number is set in the ChangeLog file in the top level ARTS directory. Always increase this when you do a SVN commit, even for small changes. In such cases increase the last digit by one. If you make a new distribution, increase the middle digit by one and omit the last digit. If you make a bug-fix distribution, you can add the last digit to indicate this.

1.3.7 Header files

The global header file arts. h *must* be included by every file. Apart from that you have to see yourself what header files you need. If you use functions from the C or C++ standard library, you have to also include the appropriate header file.

1.3.8 Documentation

Doxygen is used to generate automatic source code documentation. See

```
http://www.stack.nl/~dimitri/doxygen/
```

for information. There is a complete User manual there. At the moment we only generate the output as HTML, although latex, man-page, and rtf format is also possible. The HTML version is particularly useful for source code browsing, since it includes the complete source code! You should add Doxygen headers to the following:

- 1. Files
- 2. Classes (Including all private and public members)
- 3. Functions
- 4. Global Variables

The documentation headers are comment blocks that look like the examples below. They should be put above the *definition* of a function, i.e., in the .cc file. Some functions are defined in the .h file (e.g., inline member functions). In that case the comment can be put in the .h file.

There is an Emacs package (Doxymacs) that makes the insertion of documentation headers particularly easy. You can find documentation of this on the Doxymacs webpage: http://doxymacs.sourceforge.net/. To use it for ARTS (provided you have it), put the following in your Emacs initialization file:

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```
(doxymacs-mode))))

(add-hook 'font-lock-mode-hook 'my-doxymacs-font-lock-hook)

(setq doxymacs-doxygen-root "../doc/doxygen/html/")
(setq doxymacs-doxygen-tags "../doc/doxygen/arts.tag")
```

The only really important lines are the first two, where the second line is the one selecting the style of documentation. The next block just turns on syntax highlighting for the Doxygen headers, which looks nice. The last two lines are needed if you want to use the tag lookup features (see Doxymacs documentation if you want to find out what this is). The package allows you to automatically insert headers. The standard key-bindings are:

```
C-c d? look up documentation for the symbol under the point.
C-c d r rescan your Doxygen tags file.
C-c d f insert a Doxygen comment for the next function.
C-c d i insert a Doxygen comment for the current file.
C-c d ; insert a Doxygen comment for a member variable on the current line (like M-;).
C-c d m insert a blank multi-line Doxygen comment.
C-c d g insert a blank single-line Doxygen comment.
C-c d @ insert grouping comments around the current region.
```

You can call the macros also by name, e.g., doxymacs-insert-file-comment.

File comment

Generated by doxymacs-insert-file-comment.

```
/*!
\file dummy.cc
\author John Doe <john.doe (at) example.com>
\date 2011-03-02
\brief A dummy file.

This file has no purpose at all,
it just servers as an example...
*/
```

Function comment

Generated by doxymacs-insert-function-comment. If arguments are modified by the function you should add '[out]' after the \param command, just like for the parameter a in the example below. If a parameter is both input and output, you should say '[in,out]'. Parameters that are passed by value or not modified inside the function should carry an '[in]'. The documentation for each parameter should start with a capital letter and end with a period, like in the example below.

Author and date tags are not inserted by default, since they would be overkill if you have many small functions. However, you should include them for important functions.

Generic multi-line comment

Generated by doxymacs-insert-blank-multiline-comment.

```
//! A dummy comment.
/*!
Some more elaborate description about this variable,
class, or whatever.
*/
```

Generic single-line comment

Generated by doxymacs-insert-blank-singleline-comment.

```
//! Short comment here.
```

1.4 Extending ARTS

1.4.1 How to add a workspace variable

You should read Section 2.2 to understand what workspace variables are. Here is just the practical description how a new variable can be added.

1. Create a record entry in file workspace.cc. (Just add another one of the wsv_data.push_back blocks.) Take the already existing entries as templates. The ARTS concept works best if WSVs are only of a rather limited number of different types, so that generic WSMs can be used extensively, for example for IO.

The name must be *exactly* like you use it in the source code, because this is used to generate interface functions.

Make sure that the documentation string you give explains the variable and its purpose well. In particular, state the dimensions (in the case of matrices) and the units! This string is used for the online documentation. Please take some time to write it carefully. Use the template at the beginning of function define_wsv_data() in file workspace.cc as a guideline.

2. That's it!

1.4 Extending ARTS 7

1.4.2 How to add a workspace variable group

You should read Section 2.2 to understand what workspace variable groups are. Here is just the practical description how a new group can be added.

- 1. Add a wsv_group_names.push_back("your_type") function to the function define_wsv_group_names() in groups.cc. The name must be *exactly* like you use it in the source code, because this is used to generate interface functions.
- 2. XML reading/writing routines are mandatory for each workspace variable group. Two steps are necessary to add xml support for the new group:
 - (a) Implement an xml_read_from_stream and xml_write_to_stream function. Depending on the type of the group the implementation goes into one of the three files xml_io_basic_types.cc, xml_io_compound_types.cc, or xml_io_array_types.cc. Basic types are for example Index or Numeric. Compound types are structures and classes. And array types are arrays of basic or compound types. Also add the function declaration in the corresponding .h file.
 - (b) Add an explicit instantiation for xml_read_from_file<GROUP> and xml_write_to_file<GROUP> to xml_io_instantiation.h.
- 3. If your new group does not implement the output operator (operator<<), you have to add an explicit implementation of the Print function in m_general.h and m_general.cc.
- 4. That's it! (But as stated above, use this feature wisely)

1.4.3 How to add a workspace method

You should read Section 2.3 to understand what workspace methods are. Here is just the practical description how a new method can be added.

- 1. Create an entry in the function define_md_data in file methods.cc. (Make a copy of an existing entry (one of the md_data.push_back(...) blocks) and edit it to fit your new method.) Don't forget the documentation string! Please refer to the example at the beginning of the file to see how to format it.
- 2. Run: make.
- 3. Look in auto_md.h. There is a new function prototype

```
void <YourNewMethod>(...)
```

- 4. Add your function to one of the .cc files which contain method functions. Such files must have names starting with m_. (See separate HowTo if you want to create a new source file.) The header of your function must be compatible with the prototype in auto_md.h.
- 5. Check that everything looks nice by running

arts -d YourNewMethod

If necessary, change the documentation string.

6. Thats it!

1.4.4 How to add a source code file

- 1. Create your file. Names of files containing workspace methods should start with m_.
- 2. You have to register your file in the file src/CMakeLists.txt. This file states
 which source files are needed for arts. In the usual case, you just have to add your
 .cc file to the list of source files of the artscore library. Header files are not added to
 this list.
- 3. Go to src and run: svn add <my_file> to make your file known to SVN.

1.4.5 How to add a test case

- 1. Create a new subdirectory in tests/. If your test is closely related to another test case you can skip this step and instead add it to one of the existing subdirectories.
- 2. Create your own test controlfile. The filename should start with Test followed by the name of the subdirectory it is located in, e.g. tests/DOIT/TestDOIT.arts. If the subdirectory contains more than one test controlfile append a short descriptive text to the end of the filename like tests/MonteCarlo/TestMonteCarloGaussian.arts.
- 3. Copy all required input files into the subdirectory.
- 4. Add the necessary checks for your testcase to tests/testall.py.
- 5. In tests/CMakeLists.txt add an entry for your test case.

1.5 SVN issues

The arts project is controlled by Subversion. This section describes some basic SVN commands. For more information see the extensive SVN documentation:

http://svnbook.red-bean.com/

1.5.1 How to check out arts

- 1. Go to a temporary directory.
- 2. Run: svn co https://www.sat.ltu.se/svn/rt/arts/trunk arts

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1.5.2 How to update (if you already have a copy)

- 1. Go to the top ARTS directory (called simply arts).
- 2. Run: svn update

or C-x 4 a.

IMPORTANT! Always update, before you start to make changes to the program, especially after a longer pause. If you edit an outdated copy, it will be a lot more work to bring your changes into the current copy of the program.

1.5.3 How to commit your changes

- 1. You should make sure that the program compiles and runs without obvious errors before you commit.
- 2. If you have created a new source file, make it known to SVN by running the command svn add <my_file> in the directory where the file resides.

In general, when you run svn update, it will warn you about any files it doesn't know by marking them with a?. Files that are created during the compilation process, but should not be part of the package are ignored by SVN. You can view the list of ignored files for the current directory with svn propget svn:ignore.. The list can be edited with svn propedit svn:ignore..

- 3. Have you added the documentation for your new features?
- 4. Open the file ChangeLog in the top level ARTS directory with your favorite editor. With Emacs, you can very easily add an entry by typing either

```
M-x add-change-log-entry
```

Specify the new version number and describe your changes.

These keystrokes work also while you are editing some other file in Emacs. Thus it is best to write your ChangeLog entry already while you work on a file. Whenever you make a change to a file, there should be a ChangeLog Entry!

- 5. Make sure that you have saved all your files. Go to the top level ARTS directory and run: svn commit.
- 6. This will pop up an editor. Use the mouse to cut and paste the ChangeLog message also to this editor window. Safe the file and exit the editor. If you made changes in different directories, another editor will pop up, already containing your message. Save again and exit. Do this until no more editors come up. (Note: This works well if you set

```
export EDITOR=xedit
```

in your shell startup file.

With smart editors there can be problems, because they might refuse to safe your file if you haven't made changes to it. With xedit you just have to push the save button twice to override.

7. Tell the other developers about it. The best way to do this is to send an email to arts-dev@www.sat.ltu.se.

1.5.4 How to move your arts working directory

In general it is no problem to move your working directory. Only when you are using a build directory that is located outside the ARTS source tree, you have rerun cmake PATHTOARTS in that directory.

1.6 Debugging (use of assert)

The idea behind assert is simple. Suppose that at a certain point in your code, you expect two variables to be equal. If this expectation is a precondition that must be satisfied in order for the subsequent code to execute correctly, you must assert it with a statement like this:

```
assert (var1 == var2);
```

In general assert takes as argument a boolean expression. If the boolean expression is true, execution continues. Otherwise the abort system call is invoked and the program execution is stopped. If a bug prevents the precondition from being true, then you can trace the bug at the point where the precondition breaks down instead of further down in execution or not at all. The assert call is implemented as a C preprocessor macro, so it can be enabled or disabled at will.

In ARTS, you don't have to do this manually, as long as your source file includes arts.h either directly or indirectly. Instead, assertions are turned on and off with the global NDEBUG preprocessor macro, which is set or unset automatically by the cmake build configuration. Assertions are enabled in the default cmake build configuration (-DCMAKE_BUILD_TYPE=RelWithDebInfo). They are turned off in the release configuration (-DCMAKE_BUILD_TYPE=Release).

If your program is stopped by an assertion failure, then the first thing you should do is to find out where the error happens. To do this, run the program under the GDB debugger. First invoke the debugger:

```
qdb arts
```

You have to give the full path to the ARTS executable. Then set a breakpoint at the assertion failure:

```
(gdb) break __assert_fail
```

(Note the two leading underscores!) Now run the program:

```
(gdb) run
```

Instead of just exiting, under the debugger the program will be paused when the assertion fails, and you will get back the debugger prompt. Now type:

to see where the assertion failure happened. You can use the print command to look at the contents of variables and you can use the up and down commands to navigate the stack. For more information, see the GDB documentation or type help at the prompt of GDB.

For ARTS, the assertion failures mostly happen inside the Tensor / Matrix / Vector package (usually because you triggered a range check error, i.e., you tried to read or write beyond array bounds). In this case the up command of GDB is particularly useful. If you give this a couple of times you will finally end up in the part of your code that caused the error.

Recommendation: In Emacs there is a special GDB mode. With this you can very conveniently step through your code.

Chapter 2

The workspace

This chapter deals with the main components of ARTS: Workspace variables (WSVs) and workspace methods (WSMs). Furthermore, it explains the use of agendas, a special group of WSVs.

2.1 Implementation files

The most important files are:

- workspace.cc:
 Definition and documentation of WSVs.
- methods.cc:
 Definition and documentation of WSMs. The implementations of WSMs reside in files named m_something.cc.
- agendas.cc:
 Definition and documentation of agendas.

It is very likely that you will have to edit these. Less likely, but possibly, you also have to edit:

• groups.cc:
Definition of WSV groups.

When ARTS is built, a number of source code files are generated automatically. They are listed here in the order in which they are generated:

• auto_workspace.h:
Generated from groups.cc.

History

110622 Updated by Oliver Lemke.020605 Created by Stefan Buehler.

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auto_md.h, auto_md.cc:
 Generated from auto_workspace.h, agendas.cc, groups.cc, and methods.cc.

This is achieved by a set of simple C++ programs:

- make_auto_workspace_h.cc
- make_auto_md_h.cc
- make_auto_md_cc.cc

The meaning of the names should be self-explanatory. There is one program for each file to be generated. The generation of the auto_files happens automatically when you do a make. Therefore, never edit any of these files.

Next, there are some files that contain the internal implementation of WSVs and WSMs. These are:

- wsv_aux.h:
 Implementation of class WsvRecord, which stores the lookup information for one WSV, plus auxiliary stuff for the workspace.
- methods.h, methods_aux.cc:
 Implementation of class MdRecord, which stores the lookup information for one WSM.

Finally, there are some files that contain the internal implementation of agendas. These are:

- agenda_class.h, agenda_class.cc: Implementation of class MRecord, which stores runtime information for one WSM, and class Agenda, which stores an agenda.
- agenda_record.h, agenda_record.cc:
 Implementation of class AgRecord, which is used to store agenda lookup information.

As mentioned above, you will not have to modify any of the implementation files, they are listed here just for reference. Normally, you only have to modify workspace.cc, methods.cc, and agendas.cc.

2.2 Workspace Variables or WSVs

All important variables in ARTS are WSVs. This means that they can be manipulated by a list of WSMs, which is specified in the ARTS controlfile. There exists a predefined list of possible WSVs. This list defines the *workspace*. One can think of each WSV as a 'slot' in the workspace: The WSV can be either *set*, or *unset*. Set means that the WSV has a well-defined content, unset means that it has no well-defined content. At the start of an ARTS job all WSVs are unset.

WSVs are defined in the file workspace.cc. A typical definition looks like this:

All WSV definitions have the same three elements:

- 1. The *name*, exactly the same name has to be used in the code.
- 2. The *description*, which is normally much longer than in the example here. It must fully describe the WSV, its purpose, and its normal usage. See file workspace.cc for instructions how to write the documentation.
- 3. The *group* to which the WSV belongs. You can think of a group as something similar to a C++ data type. The WSV in the example belongs to the group Vector. The allowed groups are defined in file groups.cc.

See Section 1.4 for explicit instructions how to add a new WSV to ARTS.

2.3 Workspace Methods or WSMs

WSMs manipulate WSVs to produce other WSVs. There are three kinds of WSMs:

- 1. Specific WSMs.
- 2. Generic WSMs.
- 3. Agenda WSMs.

As in the case of WSVs, there is a central place in ARTS where information on the available WSMs is stored. This place is the file methods.cc. It contains a record for each WSM. Here is an example:

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```
AUTHORS( "Patrick Eriksson" ),

OUT( "r_geoid" ),

GOUT(),

GOUT_TYPE(),

GOUT_DESC(),

IN( "atmosphere_dim", "lat_grid", "lon_grid" ),

GIN( "r" ),

GIN_TYPE( "Numeric" ),

GIN_DEFAULT( NODEF ),

GIN_DESC( "Radius of the geoid sphere."),

));
```

All WSM definitions have the same elements:

- 1. The *NAME*, exactly as in the code.
- 2. The *DESCRIPTION*. This must fully describe the WSM, its purpose, and its normal usage. See file methods.cc for instructions how to write the documentation.
- 3. The OUT. This is a list of WSV names. All these WSVs are set by this WSM.
- 4. The GOUT. This is a list descriptive names for the generic outputs.
- 5. The *GOUT_TYPE*. This is a list of WSV group names. This defines the group to which the generic output arguments must belong (see below).
- 6. The GOUT_DESC, a list of short descriptions for the generic outputs.
- 7. The *IN*. This is a list of WSV names. All these WSVs are required as input by this WSM. This means they must have been set before.
- 8. The GIN, a list of descriptive names for the generic inputs.
- 9. The *GIN_TYPE*. This is a list of WSV group names. This defines the group to which the generic input arguments must belong.
- 10. The *GIN_DEFAULT*, a list of default values for the generic inputs. NODEF means that the generic input has no default and the user has to set it in the control file.
- 11. The GIN_DESC, a list of short descriptions for the generic inputs.

2.3.1 Specific WSMs

```
GOUT_TYPE(),
GOUT_DESC(),
IN( "abs_lookup" ),
GIN(),
GIN_TYPE(),
GIN_DEFAULT(),
GIN_DESC()
));
```

For this type of WSM the output and input is fixed. Fields GIN and GOUT are empty. The example above belongs in this category. It sets the WSV p_grid, using the WSV abs_lookup as input.

To call this method in the controlfile, you just have to write p_gridFromGasAbsLookup.

2.3.2 Generic WSMs

This class of WSMs is more powerful, because it can be applied to any WSV that belongs to the right group. A good example is:

```
md_data_raw.push_back
  ( MdRecord
    ( NAME ( "VectorSetConstant" ),
      DESCRIPTION
       "Creates a vector and sets all elements to the specified value.\n"
       "The vector length is determined by *nelem*.\n"
      AUTHORS ( "Patrick Eriksson" ),
      OUT(),
                " V "
      GOUT (
                         ),
      GOUT_TYPE( "Vector" ),
      GOUT_DESC( "Variable to initialize." ),
      IN( "nelem" ),
      GIN(
                   "value" ),
      GIN_TYPE( "Numeric"),
      GIN_DEFAULT ( NODEF ),
                "Vector value." )
      GIN_DESC(
      ));
```

As you probably have guessed, this WSM resizes the output vector to have nelem elements and sets all elements to the given value. You would use it as follows:

```
IndexSet(nelem, 10)
VectorCreate(myvector)
VectorSetConstant(myvector, nelem, 0)
```

This would create the WSV myvector and then fill it with 10 elements set to 1. Note that output arguments always come first, input arguments last. Try arts -d VectorSetConstant to get more information on this method. (See Section 1.2.3 in ARTS User Guide for information on the built-in documentation.)

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For basic types it is allowed to pass values instead of variables directly to the WSM. In that case, the above example would look like this:

```
VectorCreate(myvector)
VectorSetConstant(myvector, 10, 0)
```

2.3.3 Agenda WSMs

2.4 Agendas

2.4.1 Introduction

Agendas are a special incarnation of a WSM. At runtime an arbitrary number of WSMs can be added to an agenda. On invocation, the agenda will execute its methods one after the other. The inputs and outputs defined for the agenda must be satisfied by the invoked WSMs. E.g., if an agenda has f_grid in its list of output WSVs, a WSM which generates f_grid must be added to the agenda in the control file.

Agendas run their methods in a separate scope. Although WSMs invoked by an agenda have full access to all workspace variables, only the WSVs defined as output of the agenda will keep their values after the agenda execution. All other WSVs retain the values from before the agenda run.

Even though it is possible to execute agendas directly from the control file with the AgendaExecute method, the more common and intended use case is the internal invocation by other WSMs. This adds a considerable amount of flexibility to arts. The iyEmission-StandardClearsky method for example calculates (besides other components) the emission term. Without the means of an agenda, it would only be possible to use always the same method for the emission calculation. By the use of an agenda the user can choose between different methods to calculate the emission and plug them into the emission agenda in the control file:

```
AgendaSet( emission_agenda ) {
  emissionPlanck
}
```

Chapter 3

Vectors, matrices, tensors, and arrays

This section describes how vectors and matrices are implemented in ARTS and how they are used. Furthermore it describes how arrays of arbitrary type can be constructed and used.

3.1 Implementation files

The Matrix and Vector classes described below reside in the files:

- matpackI.h
- make_vector.h
- matpackI.cc
- make_vector.cc

Tensors of order 3 to 7 reside in the files:

- matpackIII.h
- matpackIV.h
- matpackV.h
- matpackVI.h
- matpackVII.h

The template class Array (also described below) is implemented in the files:

• array.h

| History | |
|---------|---|
| 030807 | Sparse added by Mattias Ekström. |
| 030109 | Documentation for using jokers without Range added by Stefan Buehler. |
| 020516 | Tensors added by Stefan Buehler. |
| 011018 | Created and written by Stefan Buehler. |

• make_array.h

The Sparse class is described in the file:

• matpackII.h

The file test_matpack.cc contains test cases and usage examples. For Sparse there is a separate test file, test_sparse.cc.

3.2 Vectors

The class Vector implements the mathematical concept of a vector. (Surprise, surprise.) This means that:

- A Vector contains a list of floating point values of type Numeric.
- A Vector can be multiplied with another Vector (scalar product), or with a Matrix.
- Sub-ranges of a Vector can easily be accessed, and used as if they were Vectors.
- Resizing a Vector is expensive and should be avoided.

3.2.1 Constructing a Vector

You can construct an object of class Vector in any of these ways:

The last three examples all use the same constructor, which takes the three arguments 'start', 'extent', and 'stride'. It will create a Vector containing 'extent' elements, starting with 'start', with a step of 'stride'.

There also exists a special sub-class of Vector that can be initialized explicitly. This must be a special class in order to avoid ambiguities with the standard constructors. Usage:

```
MakeVector a(1.0,2.0,3.0); // Creates a vector of length 3 // containing the values // 1.0, 2.0, and 3.0.
```

You can use MakeVectors just like Vectors, except that the constructors are different. Otherwise you can mix them freely with Vectors.

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3.2.2 VectorViews

An object of class VectorView is, like the name says, just another view on an existing Vector. It does not have its own data. This has the important consequence that it cannot be resized, since that would mess up the original Vector that the view is referring to. You can create VectorViews from Vectors using the index operator '[]', the class Range, and the special joker object. Examples:

As you can see, most useful ways to create VectorViews involve the Range class. The general constructor to this class takes three arguments, 'start', 'extent', and 'stride'. This means that you will select 'extent' elements from the Vector, starting with index 'start', with a step-width of stride. Note that indices are 0-based, so 0 refers to the first element. The last argument, 'stride', can be omitted, in that case the default of 1 is assumed. As a special case, 'extent'==joker means 'to the end', and calling Range with only one argument joker means 'all elements'.

Usually, you will not have to use VectorView explicitly, because you can use expressions like:

However, VectorView and the related class ConstVectorView are extremely useful as the argument types of functions operating on Vectors. You should define your functions like this:

Note that there must not be any '&' after VectorView or ConstVectorView. In other words they have to be passed by value, not by reference. This is ok, since they do not contain the actual data, so that passing by value is efficient. Passing VectorViews by reference is forbidden.

You should use these kind of arguments for all input Vectors, and also for the output if you have a function that does not resize the output Vector. This has the great advantage that you can call the function with Vector sub-ranges, e.g.,

An exception to this rule are workspace methods, which use conventional argument types const Vector& for input and Vector& for output.

3.2.3 What you can do with a Vector (or VectorView)

All examples below (except for the first) assume that a is a Vector, MakeVector, or VectorView.

Resize (only for Vector, not for VectorView!):

```
a.resize(5);
```

This makes a a 5 element vector. The new Vector is not initialized (i.e., the contents will be unpredictable). Also, note that the previous content will be completely lost. Appending to a Vector is not possible.

Get the number of elements:

```
cout << a.nelem();</pre>
```

Sum up all elements:

```
cout << a.sum();</pre>
```

Element access:

```
cout << a[3];  // Print 4th element.
a[0] = 3.5;  // Assign 3.5 to first element.</pre>
```

Note that we use 0-based indexing! Furthermore note that the operator '[]' can be also used with Range, as explained above.

Copying Vectors:

```
Vector b;
b = a;
```

In this case the size of b will be adjusted to that of a automatically. Maybe you have noticed that there is a way to formulate the example above in even shorter fashion:

```
Vector b = a;
```

The result is exactly the same. Note, though, that in this case b is *constructed* from a, not copied (see section about constructing Vectors above).

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Copying in connection with views:

This one is a bit tricky. Obviously, the size of views can not be adjusted, because a view is just some selection of the underlying object. The '=' operator in this case copies the *contents*, so the sizes of the left-hand and right-hand argument must match. VectorView internally uses assertions to make sure of this. So, if you get an assertion failure one reason could be that you forgot to make the target the correct size. Here is an example:

Great, isn't it?

Assigning a scalar:

```
a = 1.0; // Assign 1 to all elements.
```

Mathematical operators:

All these operate element-wise. Note, that there are no return versions of these operators (i.e., expressions like b = a+1 are not possible). This is again for efficiency reasons. It is currently an active area of research in programming techniques how to make this kind of expression efficient. None of the available solutions works, so ARTS has to live without it.

Maximum and minimum:

```
cout << max(a);
cout << min(a);</pre>
```

Scalar product:

```
cout << a*a;
```

This is an exception to the rule not to have return versions of operators. The reason is quite obvious: The return value is only a scalar.

Arbitrary single-argument math functions:

The transform function operates on each element of a with the function you specify and puts the result in b. Note that the order of the arguments is swapped compared to the old function trans that we had in the pre-Matpack era.

3.3 Matrices

The class Matrix implements the mathematical concept of a matrix. (Who would have guessed this?) This means that:

- A Matrix contains floating point values of type Numeric.
- The values are arranged in rows and columns and can be accessed by indices. The first index is the row, the second the column. In other words, we use *row-major* order, similar to C, Matlab, and most math textbooks. Note, however, that some languages like FORTRAN and IDL use *column-major* order.
- A Matrix can be multiplied with a Vector, or with another Matrix.
- A sub-range of a Matrix in both dimensions (submatrix) can easily be accessed, and used as if it was just a normal matrix.
- Resizing a Matrix is expensive and should be avoided.

3.3.1 Constructing a Matrix

You can construct an object of class Matrix in any of these ways:

That is all. More fancy constructors, like for Vector, do not exist for Matrix. There is also no equivalent to the MakeVector class.

3.3.2 MatrixViews

A MatrixView is a view on an existing Matrix, in the same way as a VectorView is a view on an existing Vector. Like a VectorView, a MatrixView cannot be resized and does not contain the actual data. A view is generated by using Ranges:

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You probably get the idea. Note that the second argument of Range gives the number of elements to take, not the index of the last element. See the section about Vectors for more examples how to use Range. You can use <code>joker</code>, and also the third argument of Range to select only every nth row, or column, or reverse the order of the rows or columns.

In analogy to the Vector case, you should use the two classes MatrixView and ConstMatrixView as function arguments. Please refer to the discussion in the Vector section for details. As in the case of VectorViews, all arguments of these types should be passed by value, not by reference. Also, similar to the Vector case, workspace methods are the exception, because they have to use the conventional const Matrix& or Matrix& as input/output arguments.

3.3.3 What you can do with a Matrix (or MatrixView)

All examples below (except for the first) assume that a is a Matrix or MatrixView.

Resize (only for Matrix, not for MatrixView!):

```
a.resize(5,10);
```

This makes a 5x10 Matrix (5 rows, 10 columns). The new Matrix is not initialized (i.e., the contents will be unpredictable). Also, note that the previous content will be completely lost.

Get the number of rows or columns:

```
cout << a.nrows();
cout << a.ncols();</pre>
```

Refer to a row or column:

```
Vector x = a(0,Range(joker)); // First row.
Vector y = a(Range(joker),a.ncols()-1); // Last column.
```

Of course, you can use more complicated Range expressions to refer to only parts of a row or column. However, the case that you want all elements of a given dimension is so much more common than the more sophisticated uses of the Range class, that it is worth to introduce a simplified notation for this case. Therefore, you are allowed to omit the Range and just write:

```
Vector x = a(0, joker); // First row.
Vector y = a(joker, a.ncols()-1); // Last column.
```

Technically, expressions of this kind return the type VectorView. This means, they can be used in all cases where an object of that type is expected, for example with the function defined in Section 3.2.2:

Element access:

```
cout << a(3,4); // Print that element.
a(0,0) = 3.5; // Assign 3.5 to the top-left element
```

Note that we use 0-based indexing! Furthermore note that the operator '()' can be also used with one or two Range arguments, as explained above. To summarize:

- (Index,Index) returns Numeric (element access).
- (Index,Range) or (Range,Index) returns VectorView (row or column access).
- (Range, Range) returns Matrix View (sub-matrix access).

You may find it unlogical, that Matrix uses '()' for indexing, whereas Vector uses '[]'. However, using '[]' for Matrix is not possible, since it can have only one argument. On the other hand, using '()' for Vector element access seemed not a good idea, since that would break with the established use of '[]' for element access in C and C++.

Copying Matrices:

```
Matrix b;
b = a;
```

As in the case of Vectors, the '=' operator adjusts the size of the target automatically.

Copying in connection with views:

As in the case of Vectors, the '=' operator copies only the *contents* for views, so the dimensions must match. An attempt to justify this behavior has been made above in the Section about Vector. As for Vector, you can use '=' with complicated expressions. Here is a more elaborate example:

Note that in this case the dimensions must match exactly, as explained in the Section about Vector.

If you do not understand the use of Range here, refer to Section 3.2.2.

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Assigning a scalar:

```
a = 1.0; // Assign 1 to all elements.
```

Mathematical operators:

You can use the operators '+=', '-=', '*=', and '/=', which operate element-vise, just as for Vector.

Maximum and minimum:

```
cout << max(a);
cout << min(a);</pre>
```

Arbitrary single-argument math functions:

The function transform works just like for Vector.

Transpose:

```
Matrix b = transpose(a); // Make b the transpose of a.
```

The function transpose creates a MatrixView, for which rows and columns are interchanged. Note, that only the way the data is accessed is changed, not the data itself. So Matrix a in the example above is not changed. For this reason, transposing is very efficient. You can use transpose (a) instead of a in any matrix expression practically without additional cost. (This is not strictly true, after all, the view has to be generated and passed. But that cost should be negligible except for very small matrices.)

Matrix multiplication:

Note, that the result is put in the first argument, consistent with the general ARTS policy, but different from the old MTL based multiplication function. Furthermore note, that as you can see from the first example, a Vector is always considered to be a 1-column Matrix.

Important: The matrices or vectors that you give for the three arguments must not overlap, or you will get garbage. In particular, this means that

```
mult(x,y,x); // x = y * x FORBIDDEN!!!
```

does not work. No, even worse: It works, but it gives the wrong result. The reason for this behavior is that the result is constructed in the first argument variable. If that is also an input variable it will change while it is multiplied, which will lead to a different result. There is

no efficient way to detect overlap, so the only way to allow input and output arguments to be identical would be to use another internal dummy variable to store the result. However, this would be much less efficient.

Another thing: You can use transpose, of course. These two examples should obviously give the same result:

3.4 Tensors

ARTS has tensors with rank 3 to 7. They are called Tensor3, Tensor4, Tensor5, Tensor6, Tensor7, and work very much like matrices, just with more dimensions. Some properties:

- A Tensor contains floating point values of type Numeric.
- The *rank* of a tensor means the number of dimensions, so a Tensor4 has 4 dimensions. Tensors of different rank are different classes. That means, the rank is fixed at compile time and cannot be changed at runtime. We will use rank and dimension as synonyms.
- The different dimensions are named:
 - Library
 - Vitrine
 - Shelf
 - Book
 - Page
 - Row
 - Column

For example, Tensor3 b(2,4,3) defines a third order tensor with 2 pages, 4 rows, and 3 columns. Note that the column dimension is always last. (Incidentally, a Matrix behaves exactly like a second order tensor, except that it has some additional features.)

- A sub-range of a tensor in all dimensions (sub tensor) can easily be accessed, and used as if it was just a normal tensor.
- More importantly, you can easily access lower dimensional 'slices' of a tensor.
- Resizing a tensor is expensive and should be avoided.

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3.4.1 Constructing a tensor

You can construct an object of a tensor class like this:

3.4.2 Tensor views

Tensor views work exactly like matrix and vector views. Example:

```
Tensor4 a(10,20,5,4);
Tensor3View b = a(3,Range(1,3),joker,joker);
```

If you have read the previous sections carefully, it should be clear what this expression does.

This is what is meant by slicing: You can easily create a view of a tensor that picks out an object of lower dimension. Note that you can use either an Index or a Range argument¹ for any of the dimensions. The dimensionality of the result will adjust accordingly, as in the example above.

Everything that was said about matrix and vector views holds also here. In particular, please always use views as function arguments.

3.4.3 What you can do with a tensor (or tensor view)

All examples below (except for the first) assume that a is a Tensor7 or Tensor7View.

Resize (only for tensors, not for views):

```
a.resize(5, 10, 4, 5, 3, 6, 8);
```

This makes a the requested size. The new tensor is not initialized (i.e., the contents will be unpredictable). Also, note that the previous content will be completely lost.

Get the extent of the various dimensions:

```
Index nl = a.nlibraries();
Index nv = a.nvitrines();
Index ns = a.nshelves();
Index nb = a.nbooks();
Index np = a.npages();
Index nr = a.nrows();
Index nc = a.ncols();
```

Which of these functions are available depends on the dimension of your tensor. For example, nlibraries() is only available for Tensor7. Note, that I took care that the first letters of the dimension names are unique, which is very convenient if you prefer short names for your variables that refer so some dimension of a tensor.

¹Using just joker is equivalent to using Range (joker), as explained in Section 3.3.

Slicing:

```
Vector x = a(0,2,1,8,3,4,joker);
// Select row 4
// on page 3
// in book 8
// on shelf 1
// in vitrine 2
// in library 0
// and copy it to the Vector x.
```

Any Range or Index expression is allowed in any of the arguments, of course.

Element access:

```
cout << a(3,4,0,0,0,0); // Print that element. a(0,0,0,0,0,0) = 3.5; // Assign 3.5 to this element.
```

Copying tensors:

Works exactly like copying matrices. Size of output argument is adjusted for Tensors, but must already have the correct size for TensorViews.

Assigning a scalar:

```
a = 1.0; // Assign 1 to all elements.
```

Mathematical operators:

You can use the operators '+=', '-=', '*/', and '/=', which operate element-vise, just as for Vector.

Maximum and minimum:

```
cout << max(a);
cout << min(a);</pre>
```

Arbitrary single-argument math functions:

The function transform works just like for Vector.

3.4.4 Making things appear larger than they are

Assume that you have written a function that performs some calculation for a Tensor5:

```
void my_function(Tensor5View x);
```

Can you call this function with a Tensor4? Yes, you can:

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In general, you can always create a view that is one dimension bigger than what you have. The leading dimension then has extent 1. There is one important exception: If you interpret a Vector as a Matrix, the trailing dimension will be 1, not the leading dimension. This is necessary, because the vector has to act like a column vector, so that matrix-vector products work in the normal way. Of course you can use telescoping to blow up anything to Tensor7:

This kind of conversion works also implicitly. So, instead of the first example, you could have simply written:

```
my_function(a);
```

The purpose of this feature is to avoid having to make special versions of auxiliary functions for all different tensor dimensions. Use it wisely! There is very little runtime overhead for this, since the data itself is not copied.

3.4.5 Summary

This is all. In particular, we have no tensor products, since they are not needed. So, tensors are mostly used to store things, notably atmospheric field and so on. There is a set of sophisticated interpolation routines, though, which are described separately in Chapter 5.

3.5 Arrays

The template class Array can be used to make arrays out of anything. I do not know a good definition for 'array', but I guess anybody who has written a computer program in any programming language is familiar with the concept. Of course, it is rather similar to the concept of a Vector, just missing all the mathematical functionality like Matrix-Vector multiplication and sub-range access.

The implementation of our Array class is based on the STL class std::vector, whereas the implementation of our Vector class is done from scratch. So the two implementations are completely independent. Nevertheless, I tried to make Array behave consistently with Vector, as much as possible. There are a number of important differences, though, hopefully sufficiently explained in this part. A short summary of important differences:

- An Array can contain elements of any type, whereas a Vector always contains elements of type Numeric.
- No mathematical functionality for Array (no sub-ranges (nothing like VectorView); no +=, -=, *=, /=; no scalar product; no transform function; no mult function; no transpose function).
- On the other hand, resizing (for example adding to the end) of an Array is ok. (See the push_back method below.) It is still rather expensive, though, at least for large Arrays.

3.5.1 Constructing an Array

You can construct an object of an Array class like this:

There are already a lot of predefined Array classes. The naming convention for them is: ArrayOfIndex, ArrayOfString, etc.. Normally you should use these predefined classes. But if you want to define an Array of some uncommon type, you can do it with '<>', as in the above examples.

As for Vector, there is a special sub-class of Array that can be initialized explicitly. Usage:

3.5.2 What you can do with an Array

All examples below assume that a is an ArrayOfString.

Resize:

```
a.resize(5);
```

This adjusts the size of a to 5. Resizing is more efficiently implemented than for Vector, but still expensive.

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Get the number of elements:

```
cout << a.nelem(); // Just as for Vector.</pre>
```

In particular, note that the return type of this method is Index, just as for Vector. This is an extension compared to std::vector, which just has a method size() that returns the positive integer type size_t.

Element access:

```
cout << a[3];  // Print 4th element.
a[0] = "Hello"; // Assign string "Hello" to first element.</pre>
```

In other words, this works just like for Vector.

Copying Arrays:

This works also the same as for Vector. The size of the target must match! In this respect, I have modified the behavior with respect to the underlying std::vector, which has different copy semantics.

Assigning a scalar of the base type:

```
a = "Hello"; // Assign string "Hello" to all elements.
```

Append to the end:

This can be an expensive operation, especially for large Arrays. Therefore, use it with care. Actually, the push_back method comes from the std::vector class that Array is based on. You can do a lot more with std::vector, all of which also works with Array. However, to explain the Standard Template Library is beyond the scope of this text. You can read about it in C++ or even dedicated STL textbooks.

3.6 Sparse matrices

The class Sparse implements the mathematical concept of a matrix, same as Matrix does, but the data is stored in a different manner. Sparse offers a memory saving storage when most of the matrix is filled with zeros. This means that:

- A Sparse contains floating point values of type Numeric.
- The values are arranged in rows and columns in the same ways as for ordinary matrices, in row-major order.
- A Sparse can be multiplied with a Vector, a Matrix or with another Sparse.
- There exist no views for Sparse.

• Resizing a Sparse is expensive and should be avoided.

To calculate the maximum number of non-zero elements for efficient storage, take the product of number of columns and number of rows, subtract the number of columns plus one and then divide by two, $(nnz \le 0.5 \times (ncols \times nrows - (ncols + 1))$.

3.6.1 Constructing a Sparse

You can construct an object of class Sparse in any of these ways:

```
Sparse a;  // Create empty Sparse.
Sparse b(3,4);  // Create Sparse with 3 rows
  // and 4 columns. When
  // created like this it will
  // contain only zeros, i.e.
  // be an empty Sparse.
Sparse d=c;  // Make d a copy of c.
```

That is all. As for Matrix there exist no more fancy constructors, like an equivalent to the MakeVector class.

3.6.2 What you can do with a Sparse

All examples below assume that a is a Sparse.

Identity matrix:

```
a.make_I(10,10);
```

This sets a to be the identity matrix of size 10x10 (10 rows and 10 columns). Using this function is much faster than setting the diagonal elements to one by yourself. The number of rows and columns doesn't have to match each other. In the case that they don't, the rule that each row and column only will have only one position occupied by a one is applied. That is, in the case there are more rows than columns, the last rows will be empty.

Resize:

```
a.resize(5,10);
```

This makes a a 5x10 Sparse (5 rows, 10 columns). Note that the previous content will be completely lost. The new Sparse will be empty.

Get the number of rows, columns or non-zero elements:

```
cout << a.nrows();
cout << a.ncols();
cout << a.nnz();</pre>
```

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Element access:

There are two different ways to access individual elements. One used for read only and one for read and write. The distinction is necessary since the read and write method creates elements if they don't already exist. Note that we use 0-based indexing. For reading only use:

For reading and writing, such as assigning values to elements, use:

Copying Matrices:

```
Sparse b;
b = a;
```

As in the case of Vectors, the '=' operator adjusts the size of the target automatically.

Transpose:

The function transpose works a bit differently for Sparse than for Vector and Matrix. This is due to the fact that we don't have any views for Sparse. Thus, transpose for a Sparse creates a new Sparse variable that contains the transpose of the original Sparse, whereas transpose for a Matrix just creates a transposed view of the original Matrix.

The target variable for the transposed Sparse has to have the right dimensions before the function is called.

Matrix multiplication:

The result is put in the first argument, consistent with the Matrix class. Note that for the Sparse – Matrix multiplication the output is a Matrix. **Important: As for Matrix, the matrices or vectors that you give for the three arguments must not overlap, or you will get garbage.**

Chapter 4

Gridded Fields

This section describes how gridded fields are implemented in ARTS and how they are used. Gridded fields consist of a data object like a Vector, Matrix, or Tensor and a grid for each dimension of its data. For example, a GriddedField1 consists of one grid and a Vector, whereas a GriddedField3 contains three grids and a Tensor3. Grids can be either numeric, like a pressure grid, or strings, like channel names.

4.1 Implementation files

The GriddedField1, GriddedField2, GriddedField3, GriddedField4 classes and their common base class GriddedField described below reside in the files:

- gridded_fields.h
- gridded_fields.cc

4.2 Design

4.2.1 The abstract base class GriddedField

The abstract base class GriddedField implements the properties all gridded fields have in common. These are mostly the methods to create, set, and access the grids. A GriddedField is never instantiated directly.

4.2.2 Inheritance

The GriddedFieldX classes use indirect inheritance to combine a data object with the grids, see Figure 4.1.

History

2010-09-28 Oliver Lemke: Updated for implementation changes.

2010-04-12 Created and written by Oliver Lemke.

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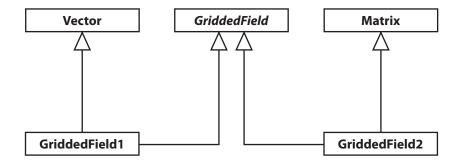


Figure 4.1: UML diagram of gridded field inheritance.

4.3 Constructing Gridded Fields

4.3.1 Creation

Each GriddedFieldX offers two constructors. One default constructor that creates an unnamed gridded field and a second constructor that takes a string with the name of the gridded field as an argument.

```
GriddedField1 gfone("I'm a GriddedField1");
GriddedField2 gftwo;
gftwo.set_name ("I'm a GriddedField2");
```

4.3.2 Initializing the grids

Once a gridded field has been created, we can start setting up the grids. There are two different types of grids, a numeric grid and a string grid. In the following example we set up two gridded fields: A GriddedField1 with a numeric grid and a GriddedField2 with a numeric grid for the rows and a string grid for the columns. Each grid can be assigned a name to describe its contents or unit.

4.3.3 Initializing the data

The data of a GriddedFieldX can be accessed through its data member. For a GriddedField1 data is a Vector, for a GriddedField2 a Matrix, for a GriddedField3 a Tensor3, and so on.

The following code shows how to fill the gridded fields from the previous example with data:

4.3.4 Consistency check

After initializing or changing either the grids or the data, it can happen that the size of the grids does not match the size of the data anymore. Each gridded field provides a convenience function which can be called to perform a consistency check.

The complete source code of the examples from this chapter can be found in src/test_gridded_fields.cc.

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Chapter 5

Interpolation

There are no general single-step interpolation functions in ARTS. Instead, there is a set of useful utility functions that can be used to achieve interpolation. Roughly, you can separate these into functions determining grid position arrays, functions determining interpolation weight tensors, and functions applying the interpolation. Doing an interpolation thus requires a chain of function calls:

- 1. gridpos (one for each interpolation dimension)
- 2. interpweights
- 3. interp

Currently implemented in ARTS is mulitlinear interpolation in up to 6 dimensions. (Is the 6D case called hexa-linear interpolation?) The necessary functions and their interaction will be explained in this chapter.

5.1 Implementation files

Variables and functions related to interpolation are defined in the files:

- interpolation.h
- interpolation.cc
- test_interpolation.cc

The first two files contain the declarations and implementation, the last file some usage examples.

| History | |
|---------|---|
| 100204 | Added documentation of grid checking functions by Stefan Buehler. |
| 020528 | Created by Stefan Buehler. |

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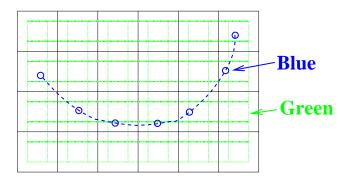


Figure 5.1: The two different types of interpolation. Green (dotted): Interpolation to a new grid, output has same dimension as input, in this case 2D. Blue (dashed): Interpolation to a sequence of points, output is always 1D.

5.2 Green and blue interpolation

There are two different types of interpolation in ARTS:

Green Interpolation: Interpolation of a gridded field to a new grid.

Blue Interpolation: Interpolation of a gridded field to a sequence of positions.

Figure 5.1 illustrates the different types for a 2D example.

The first step of an interpolation always consists in determining where your new points are, relative to the original grid. You can do this separately for each dimension. The positions have to be stored somehow, which is described in the next section.

5.3 Grid checking functions

Before you do an interpolation, you should check that the new grid is inside the old grid. (Or only slightly outside.) You can use the convenient function <code>chk_interpolation_grids</code> for this purpose, which resides in file <code>check_input.cc</code>. The function has the following parameters:

| const String& | which_interpolation | A string describing the interpolation for which the grids are intended. |
|--|-----------------------------------|--|
| ConstVectorView ConstVectorView const Numeric& | old_grid new_grid extpolfac | The original grid. The new grid. The extrapolation fraction. See gridpos function for details. Has a default |
| | | value, which is consistent with gridpos. |

There is also a special version for the case that the new grid is just a scalar. What the function does is check if old and new grid for an interpolation are ok. If not, it throws a detailed runtime error message.

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The parameter extpolfac determines how much extrapolation is tolerated. Its default value is 0.5, which means that we allow extrapolation as far out as half the spacing of the last two grid points on that edge of the grid.

The chk_interpolation_grids function is quite thorough. It checks not only the grid range, but also the proper sorting, whether there are duplicate values, etc.. It is not completely cheap computationally. Its intended use is at the beginning of workspace methods, when you check the input variables and issue runtime errors if there are any problems. The runtime error thrown also explains in quite a lot of detail what is actually wrong with the grids.

5.4 Grid positions

A grid position specifies where an interpolation point is, relative to the original grid. It consists of three parts, an Index giving the original grid index below the interpolation point, a Numeric giving the fractional distance to the next original grid point, and a Numeric giving 1 minus this number. Of course, the last element is redundant. However, it is efficient to store this, since it is used many times over. We store the two numerics in a plain C array of dimension 2. (No need to use a fancy Array or Vector for this, since the dimension is fixed.) So the structure GridPos looks like:

For example, idx=3 and fd=0.5 means that this interpolation point is half-way between index 3 and 4 of the original grid. Note, that 'below' in the first paragraph means 'with a lower index'. If the original grid is sorted in descending order, the value at the grid point below the interpolation point will be numerically higher than the interpolation point. In other words, grid positions and fractional distances are defined relative to the order of the original grid. Examples:

```
old grid = 2 3

new grid = 2.25

idx = 0

fd[0] = 0.25

old grid = 3 2

new grid = 2.25

idx = 0

fd[0] = 0.75
```

Note that fd[0] is different in the second case, because the old grid is sorted in descending order. Note also that idx is the same in both cases.

Grid positions for a whole new grid are stored in an Array<GridPos> (called ArrayOfGridPos).

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5.5 Setting up grid position arrays

There is only one function to set up grid position arrays, namely gridpos:

Some points to remember:

- As usual, the output qp has to have the right dimension.
- The old grid has to be strictly sorted. It can be in ascending or descending order. But there must not be any duplicate values. Furthermore, the old grid must contain at least two points.
- The new grid does not have to be sorted, but the function will be faster if it is sorted or mostly sorted. It is ok if the new grid contains only one point.
- The beauty is, that this is all it needs to do also interpolation in higher dimensions: You just have to call gridpos for all the dimensions that you want to interpolate.
- Note also, that for this step you do not need the field itself at all!
- If you want to use the returned gp object for something else than interpolation, you should know that gridpos guarantees the following:
 For the ascending old grid case:

```
old_grid[tgp.idx] <= tng || tgp.idx == 0</pre>
```

And for the descending old grid case:

```
old_grid[tgp.idx]>=tng || tgp.idx==0
```

• Finally, note that parameter extpolfac plays the same role as explained above in Section 5.3.

5.6 Interpolation weights

As explained in the 'Numerical Recipes' [*Press et al.*, 1997], 2D bi-linear interpolation means, that the interpolated value is a weighted average of the original field at the four corner points of the grid square in which the interpolation point is located. For simplicity, we label the four corner points counterclockwise, starting from the lower left point (Figure 5.2). Then the interpolated value is given by:

$$y(t,u) = (1-t)*(1-u)*y_1 + t*(1-u)*y_2 + t*u*y_3 + (1-t)*u*y_4 = w_1*y_1 + w_2*y_2 + w_3*y_3 + w_4*y_4$$
 (5.1)

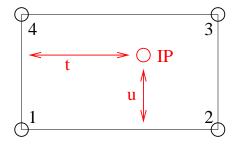


Figure 5.2: The grid square for 2D interpolation. The numbers 1...4 mark the corner points, IP is the interpolation point, t and u are the fractional distances in the two dimensions.

where t and u are the fractional distances between the corner points in the two dimensions, y_i are the field values at the corner points, and w_i are the interpolation weights.

(By the way, I have discovered that this is exactly the result that you get if you first interpolate linearly in one dimension, then in the other. I was playing around with this a bit, but it is the more efficient way to pre-calculate the w_i and do all dimensions at once.

How many interpolation weights one needs for a multilinear interpolation depends on the dimension of the interpolation: There are exactly 2^n interpolation weights for an n dimensional interpolation. These weights have have to be computed for each interpolation point (each grid point of the new grid, if we do a 'green' type interpolation. Or each point in the sequence, if we do a 'blue' type interpolation).

This means, calculating the interpolation weights is not exactly cheap, especially if one interpolates simultaneously in many dimensions. On the other hand, one can save a lot by re-using the weights. Therefore, interpolation weights in ARTS are stored in a tensor which has one more dimension than the output field. The last dimension is for the weight, so this last dimension has the extent 4 in the 2D case, 8 in the 3D case, and so on (always 2^n).

In the case of a 'blue' type interpolation, the weights are always stored in a matrix, since the output field is always 1D (a vector).

5.7 Setting up interpolation weight tensors

Interpolation weight tensors can be computed by a family of functions, which are all called interpweights. Which function is actually used depends on the dimension of the input and output quantities. For this step we still do not need the actual fields, just the grid positions.

5.7.1 Blue interpolation

In this case the functions are:

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```
const ArrayOfGridPos& pgp,
const ArrayOfGridPos& rgp,
const ArrayOfGridPos& cgp );
void interpweights( MatrixView itw,
const ArrayOfGridPos& vgp,
const ArrayOfGridPos& sgp,
const ArrayOfGridPos& bgp,
const ArrayOfGridPos& pgp,
const ArrayOfGridPos& rgp,
const ArrayOfGridPos& cgp );
```

In all cases, the dimension of itw must be consistent with the given grid position arrays and the dimension of the interpolation (last dimension 2^n). Because the grid position arrays are interpreted as defining a sequence of positions they must all have the same length.

5.7.2 Green interpolation

In this case the functions are:

```
void interpweights ( Tensor 3 View itw,
                    const ArrayOfGridPos& rgp,
                    const ArrayOfGridPos& cgp );
void interpweights ( Tensor4View itw,
                    const ArrayOfGridPos& pgp,
                    const ArrayOfGridPos& rgp,
                    const ArrayOfGridPos& cgp );
void interpweights ( Tensor5View itw,
                    const ArrayOfGridPos& bgp,
                    const ArrayOfGridPos& pgp,
                    const ArrayOfGridPos& rgp,
                    const ArrayOfGridPos& cgp );
void interpweights ( Tensor6View itw,
                    const ArrayOfGridPos& sqp,
                    const ArrayOfGridPos& bgp,
                    const ArrayOfGridPos& pgp,
                    const ArrayOfGridPos& rgp,
                    const ArrayOfGridPos& cgp );
void interpweights ( Tensor7View itw,
                    const ArrayOfGridPos& vgp,
                    const ArrayOfGridPos& sqp,
                    const ArrayOfGridPos& bgp,
                    const ArrayOfGridPos& pgp,
                    const ArrayOfGridPos& rgp,
                     const ArrayOfGridPos& cgp );
```

In this case the grid position arrays are interpreted as defining the grids for the interpolated field, therefore they can have different lengths. Of course, itw must be consistent with the length of all the grid position arrays, and with the dimension of the interpolation (last dimension 2^n).

5.8 The actual interpolation

For this final step we need the grid positions, the interpolation weights, and the actual fields. For each interpolated value, the weights are applied to the appropriate original field values and the sum is taken (see Equation 5.1). The interp family of functions performs this step.

5.8.1 Blue interpolation

```
void interp( VectorView
                                  ia,
             ConstMatrixView
                                  itw,
            ConstVectorView
                                  a,
            const ArrayOfGridPos& cgp);
void interp( VectorView
            ConstMatrixView
                                 itw,
            ConstMatrixView
            const ArrayOfGridPos& rgp,
            const ArrayOfGridPos& cgp);
void interp( VectorView
                         ia,
                                 itw,
            ConstMatrixView
            ConstTensor3View
             const ArrayOfGridPos& pgp,
             const ArrayOfGridPos& rgp,
             const ArrayOfGridPos& cgp);
void interp( VectorView
            ConstMatrixView
                                  itw,
                              1ı
a,
            ConstTensor4View
            const ArrayOfGridPos& bgp,
             const ArrayOfGridPos& pgp,
            const ArrayOfGridPos& rgp,
            const ArrayOfGridPos& cgp);
void interp( VectorView
                                  ia,
            ConstMatrixView
                                  itw,
            ConstTensor5View
             const ArrayOfGridPos& sgp,
             const ArrayOfGridPos& bgp,
             const ArrayOfGridPos& pgp,
             const ArrayOfGridPos& rgp,
             const ArrayOfGridPos& cgp);
void interp( VectorView
                                  ia,
             ConstMatrixView
                                  itw,
            ConstTensor6View
                                  a,
             const ArrayOfGridPos& vgp,
             const ArrayOfGridPos& sgp,
             const ArrayOfGridPos& bgp,
             const ArrayOfGridPos& pgp,
             const ArrayOfGridPos& rgp,
             const ArrayOfGridPos& cgp);
```

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5.8.2 Green interpolation

```
void interp( MatrixView
                                   ia,
             ConstTensor3View
                                   itw,
             ConstMatrixView
             const ArrayOfGridPos& rgp,
             const ArrayOfGridPos& cgp);
void interp( Tensor3View
             ConstTensor4View
                                   it.w.
             ConstTensor3View
             const ArrayOfGridPos& pgp,
             const ArrayOfGridPos& rgp,
             const ArrayOfGridPos& cgp);
void interp( Tensor4View
             ConstTensor5View
                                  itw,
             ConstTensor4View
             const ArrayOfGridPos& bgp,
             const ArrayOfGridPos& pgp,
             const ArrayOfGridPos& rgp,
             const ArrayOfGridPos& cgp);
void interp( Tensor5View
                                  ia.
             ConstTensor6View
             ConstTensor5View
             const ArrayOfGridPos& sgp,
             const ArrayOfGridPos& bgp,
             const ArrayOfGridPos& pgp,
             const ArrayOfGridPos& rgp,
             const ArrayOfGridPos& cgp);
void interp( Tensor6View
                                  ia,
             ConstTensor7View
                                  itw,
             ConstTensor6View
             const ArrayOfGridPos& vgp,
             const ArrayOfGridPos& sgp,
             const ArrayOfGridPos& bgp,
             const ArrayOfGridPos& pgp,
             const ArrayOfGridPos& rgp,
             const ArrayOfGridPos& cgp);
```

5.9 Examples

5.9.1 A simple example

This example is contained in file test_interpolation.cc.

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```
cout << "New grid:\n" << ng << "\n";
// To store the grid positions:
ArrayOfGridPos gp(ng.nelem());
gridpos(gp,og,ng);
cout << "Grid positions:\n" << gp;</pre>
// To store interpolation weights:
Matrix itw(gp.nelem(),2);
interpweights(itw,gp);
cout << "Interpolation weights:\n" << itw << "\n";</pre>
// Original field:
Vector of(og.nelem(),0);
                                // 0, 0, 10, 0, 0
of [2] = 10;
cout << "Original field:\n" << of << "\n";
// Interpolated field:
Vector nf(ng.nelem());
interp(nf, itw, of, gp);
cout << "New field:\n" << nf << "\n";</pre>
```

Ok, maybe you think this is not so simple, but a large part of the code is either setting up the example grids and fields, or output. And here is how the output looks like:

```
Very simple interpolation case
Original grid:
 1 2
         3 4
New grid:
  2 2.25 2.5 2.75
Grid positions:
  1 0
       1
  1 0.25 0.75
  1 0.5 0.5
  1 0.75 0.25
  1 1
         0
Interpolation weights:
 1 0
0.75 0.25
0.5 0.5
0.25 0.75
 0
    1
Original field:
 0 0 10 0
                 0
New field:
  0 2.5
         5 7.5 10
```

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5.9.2 A more elaborate example

What if you want to interpolate only some dimensions of a tensor, while retaining others? — You have to make a loop yourself, but it is very easy. Below is an explicit example for a more complicated interpolation case. (Green type interpolation of all pages of a Tensor3.) This example is also contained in file test_interpolation.cc.

```
void test04()
  cout << "Green type interpolation of all "</pre>
       << "pages of a Tensor3\n";
  // The original Tensor is called a, the new one n.
  // 10 pages, 20 rows, 30 columns, all grids are: 1,2,3
 Vector a_pgrid(1,3,1), a_rgrid(1,3,1), a_cgrid(1,3,1);
  Tensor3 a( a_pgrid.nelem(),
             a_rgrid.nelem(),
             a_cgrid.nelem() );
  a = 0;
  // Put some simple numbers in the middle of each page:
  a(0,1,1) = 10;
  a(1,1,1) = 20;
  a(2,1,1) = 30;
  // New row and column grids:
  // 1, 1.5, 2, 2.5, 3
 Vector n_{\text{rgrid}}(1,5,.5), n_{\text{cgrid}}(1,5,.5);
  Tensor3 n( a_pgrid.nelem(),
             n_rgrid.nelem(),
             n_cgrid.nelem() );
  // So, n has the same number of pages as a,
  // but more rows and columns.
  // Get the grid position arrays:
  ArrayOfGridPos n_rgp(n_rgrid.nelem()); // For rows.
  ArrayOfGridPos n_cgp(n_cgrid.nelem()); // For columns.
  gridpos( n_rgp, a_rgrid, n_rgrid );
  gridpos( n_cgp, a_cgrid, n_cgrid );
  // Get the interpolation weights:
  Tensor3 itw( n_rgrid.nelem(), n_cgrid.nelem(), 4 );
  interpweights( itw, n_rgp, n_cgp );
  // Do a "green" interpolation for all pages of a:
  for (Index i=0; i<a.npages(); ++i)
    {
      // Select the current page of both a and n:
      ConstMatrixView ap = a(i,
```

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```
Range(joker), Range(joker) );
      MatrixView
                       np = n(i,
                                Range(joker), Range(joker) );
      // Do the interpolation:
      interp( np, itw, ap, n_rgp, n_cgp );
      // Note that this is efficient, because interpolation
      // weights and grid positions are re-used.
  cout << "Original field:\n";</pre>
  for ( Index i=0; i<a.npages(); ++i )</pre>
      cout << "page " << i << ":\n"
           << a(i,Range(joker),Range(joker)) << "\n";
  cout << "Interpolated field:\n";</pre>
  for ( Index i=0; i<n.npages(); ++i )</pre>
      cout << "page " << i << ":\n"
           << n(i,Range(joker),Range(joker)) << "\n";</pre>
}
The output is:
Green type interpolation of all pages of a Tensor3
Original field:
page 0:
  0
          0
      0
    10
  0
          0
  0
      0
          0
page 1:
          0
  0
      0
  0
     20
          0
  0
     0
page 2:
          0
  0
     0
  0
    30
          0
  0
     0
          0
Interpolated field:
page 0:
          0
               0
  0
     0
                   0
  0 2.5
          5 2.5
                   0
      5 10
  0
               5
                   0
  0 2.5
          5 2.5
                   0
  0
     0
          0
               0
                   0
page 1:
  0
      0
          0
              0
                   0
  0
     5
         10
               5
                   0
  0
     10
         20
              10
                   0
     5
         10
              5
  0
                   0
  0
      0
         0
              0
                   0
page 2:
```

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```
0 0 0 0 0
0 7.5 15 7.5 0
0 15 30 15 0
0 7.5 15 7.5 0
0 0 0 0
```

5.10 Higher order interpolation

Everything that was written so far in this chapter referred to linear interpolation, which uses 2 neighboring data points in the 1D case. But ARTS has also a framework for higher order polynomial interpolation. It is defined in the two files

```
• interpolation_poly.h
```

• interpolation_poly.cc

We define interpolation order O as the order of the polynomial that is used. Linear interpolation, the ARTS standard case, corresponds to O=1. O=2 is quadratic interpolation, O=3 cubic interpolation. The number of interpolation points (and weights) for a 1D interpolation is O+1 for each point in the new grid. So, linear interpolation uses 2 points, quadratic 3, and cubic 4.

Note, that if you use even interpolation orders, you will have an unequal number of interpolation points 'to the left' and 'to the right' of your new point. This is an argument for preferring O=3 as the basic higher order polynomial interpolation, instead of O=2.

Overall, higher order interpolation works rather similarly to the linear case. The main difference is that grid positions for higher order interpolation are stored in an object of type GridPosPoly, instead of GridPosPoly object contains grid indices and interpolation weights for all interpolation points. For each point in the new grid, there are O+1 indices and O+1 weights.

The reason why we store all interpolation point indices, and not only the index of the first point, is to allow correct handling of circular interpolation, for example in scattering phase function ϕ angle. If the angle goes from 0 to 360°, then points just below 360 should be used in interpolations to points just above 0, so the indices to use are not contiguous in memory. Functions to handle this are not yet implemented, but this should be a relatively simple matter.

In contrast to GridPos, GridPosPoly stores weights w rather than fractional distances fd. For the linear case:

```
w[0] = fd[1]
w[1] = fd[0]
```

So the two concepts are almost the same. Because the w are associated with each interpolation point, they work also for higher interpolation order, whereas the concept of fractional distance does not.

The weights are calculated according to section 3.1, eq. 3.1.1 of [*Press et al.*, 1997]. These are for the 1D case. For 2D and higher dimensional cases, the weights of the individual dimensions have to be multiplied, just as in the linear interpolation case.

Instead of gridpos, you have to use the function gridpos_poly for higher order interpolation. It works exactly like gridpos, but has an additional argument that gives the interpolation order O.

After setting up the <code>GridPosPoly</code> object with <code>gridpos_poly</code>, you have to call interpweights and interp, exactly as in the linear case. (The actual functions used are not the same, since the name is overloaded. The <code>interpweights</code> and <code>interp</code> functions for use with <code>GridPosPoly</code> are implemented in <code>interpolation_poly.cc.</code>) So, a complete interpolation chain involves:

```
gridpos_poly
interpweights
interp
```

For O=1 the result of the interpolation chain will be the same as for the linear interpolation routines. Below is a simple complete example, taken from the file test_interpolation.cc in the arts source directory:

```
void test08()
  cout << "Very simple interpolation case for the "</pre>
       << "new higher order polynomials.\n";
  Vector og(1,5,+1);
                                 // 1, 2, 3, 4, 5
  Vector ng(2,5,0.25);
                                  // 2.0, 2,25, 2.5, 2.75, 3.0
  cout << "Original grid:\n" << og << "\n";</pre>
  cout << "New grid:\n" << ng << "\n";
  // To store the grid positions:
  ArrayOfGridPosPoly gp(ng.nelem());
                                  // Interpolation order.
  Index order=2;
  gridpos_poly(gp,og,ng,order);
  cout << "Grid positions:\n" << gp;</pre>
  // To store interpolation weights:
  Matrix itw(gp.nelem(), order+1);
  interpweights(itw,gp);
  cout << "Interpolation weights:\n" << itw << "\n";</pre>
  // Original field:
  Vector of(og.nelem(),0);
                                  // 0, 0, 10, 0, 0
  of [2] = 10;
  cout << "Original field:\n" << of << "\n";</pre>
  // Interpolated field:
  Vector nf(ng.nelem());
```

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```
interp(nf, itw, of, gp);

cout << "New field (order=" << order << "):\n" << nf << "\n";

cout << "All orders systematically:\n";
  for (order=1; order<5; ++order)
    {
      gridpos_poly(gp,og,ng,order);
      itw.resize(gp.nelem(),order+1);
      interpweights(itw,gp);
      interp(nf, itw, of, gp);

      cout << "order " << order << ": ";
      for (Index i=0; i<nf.nelem(); ++i)
            cout << setw(8) << nf[i] << " ";
      cout << "\n";
      }
}</pre>
```

5.11 Summary

Now you probably understand better what was written at the very beginning of this chapter, namely that doing an interpolation always requires the chain of function calls:

- 1. gridpos or gridpos_poly (one for each interpolation dimension)
- 2. interpweights
- 3. interp

If you are interested in how the functions really work, look in file interpolation.cc or interpolation.poly.cc. The documentation there is quite detailed. When you are using interpolation, you should always give some thought to whether you can re-use grid positions or even interpolation weights. This can really save you a lot of computation time. For example, if you want to interpolate several fields — which are all on the same grids — to some position, you only have to compute the weights once.

Chapter 6

Integration functions

A radiative transfer model which takes into account the effect of scattering involves integration of certain quantities over the angles of observation. For example, from Section 10.1.2 in *ARTS User Guide* it is clear that computing scattering cross-section and scattering integral term requires integration over zenith and azimuth directions. There are a wide range of methods that can be used for numerical integration. They can be used depending on various factors starting from how accurate the result should be to the behaviour of the function. The one which is implemented in ARTS is the trapezoidal integration method.

6.1 Implementation files

The integration functions can be found in the files:

- math_funcs.h
- math_funcs.cc

The implementation function AngIntegrate_trapezoidis discussed in the second file.

6.2 Trapezoidal Integration

Trapezoidal Integration method comes under the Newton-Cotes formulas where integration of a function is approximated by the area under the curve described by the function. Trapezoidal integration assumes that the area under the curve is trapezoid.

Trapezoidal rule:

$$\int_{x_1}^{x_2} f(x)dx = \frac{1}{2}h(f_1 + f_2) + O(h^3 f'')$$
(6.1)

| History | | | | | | |
|---------|-----------|------------------|--------------|-----|-------------|-------------|
| 220802 | Created a | nd written by Sr | eerekha T.R. | | | |
| 220103 | Included | mathematical | description | for | implemented | integration |
| | method(C | E). | | | | |

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This is a two-point formula $(x_1 \text{ and } x_2)$. It is exact for polynomials upto and including degree 1, i.e., f(x) = x. $O(h^3 f'')$ signifies how far is the true answer from the estimate.

If we use eq. 6.1 N-1 times, to do the integration in the intervals (x_1, x_2) , (x_2, x_3) , ..., (x_{N-1}, x_N) , and then add the results, we obtain extended formula for the integral from x_1 to x_N .

Extended Trapezoidal rule:

$$\int_{x_1}^{x_N} f(x)dx = \frac{1}{2}h\left[f_1 + 2(f_2 + f_3 + \dots + f_{N-1}) + f_N\right] + O\left[\frac{(b-a)^3 f''}{N^2}\right]$$
(6.2)

The last term tells how much the error will be decreased by taking more number of steps.

6.3 Solid Angle Integration

In our scattering problem, we are often encountered with a double integration of functions over zenith and azimuth angles (see *ARTS User Guide*, Chapter 10). One way to achieve double integration is to use repeated one-dimensional trapezoidal integration. This is effective of course only if the boundary is simple and the function is very smooth. If the function is strongly peaked and if know where it occurs, integral should be broken into smaller regions so that the integrand is smooth in each. Another thing is to take into account the symmetry of the function as well as the boundary. For example in our case, if the radiation is symmetric about the azimuth, the integration in that direction returns constant value of 2π and we need to do only integration over zenith directions.

The general form of a solid angle integration is

$$S = \int_{4\pi} f(\omega) d\omega \tag{6.3}$$

In spherical coordinates we can write:

$$S = \int_0^{\pi} \int_0^{2\pi} f(\theta, \phi) \sin \theta \, d\theta d\phi \tag{6.4}$$

A double integration can be splitted into two single integrations:

$$S = \int_0^{\pi} \left(\int_0^{2\pi} f(\theta, \phi) \sin \theta d\phi \right) d\theta$$
 (6.5)

$$= \int_0^{\pi} g(\theta) d\theta \tag{6.6}$$

If we have to integrate a vector, we can apply this method componentwise.

To solve the integral numerically we discretize θ and ϕ and obtain two angular grids ($[\theta_0, \theta_1, \cdots, \theta_n]$ and $[\phi_0, \phi_1, \cdots, \phi_m]$). Then we can first calculate $g(\theta_j)$ for all θ_j unsing the trapezoidal method.

$$g(\theta_j) = \sum_{i=1}^{m} \sin \theta_j \frac{f(\theta_j, \phi_i) + f(\theta_j, \phi_{i+1})}{2} \cdot (\phi_{i+1} - \phi_i)$$
 (6.7)

The final step is to sum up all $g(\theta_i)$, again applying the trapezoidal method.

$$S = \sum_{j=1}^{n} \frac{g(\theta_j) + g(\theta_{j+1})}{2} \cdot (\theta_{j+1} - \theta_j)$$
(6.8)

If the radiation is symmetric about the azimuth we just calculate:

$$S_{sym} = 2\pi \int_0^{\pi} f(\theta) \sin(\theta) d\theta$$
 (6.9)

Unsing the trapezoidal method this can be written as:

$$S_{sym} = 2\pi \sum_{j=1}^{n} \frac{h(\theta_j) + h(\theta_{j+1})}{2} \cdot (\theta_{j+1} - \theta_j)$$
(6.10)

where $h(\theta) = \sin \theta \cdot f(\theta)$.

The function AngIntegrate_trapezoid takes as input the integrand and the angles over which the integration has to be done. For example in this case it can be the zenith and azimuth angle grid.

```
Numeric AngIntegrate_trapezoid(MatrixView Integrand,
ConstVectorView za_grid,
ConstVectorView aa_grid)
```

The integrand has the same number of rows as zenith angle grid and columns as azimuth angle grid. The inner loop does trapezoidal integration of the integrand over all azimuth angles and the result is stored in a Vector res1[i]. Note that the integrand at every point has to be multiplied with sin (za_grid[i] * DEG2RAD) since we are integrating over solid angles. The outer loop does an integration of res1[i] over all zentih angles. The result of this is returned back to the calling function.

Chapter 7

Linear algebra functions

Solving the vector radiative transfer equation requires the computation of linear equation systems and the matrix exponential. This section describes the functions which are implemented in ARTS and it gives instructions how these functions can be used, also for other purposes than the radiative transfer calculations.

7.1 Implementation files

All the functions described below can be found in the files:

- lin_alg.h
- lin_alg.cc

The template class Array and the classes Matrix and Vector are used, therefore the linear algebra functions require the files:

- matpackI.h
- make_vector.h
- array.h
- matpackI.cc
- make_vector.cc
- array.cc

Furthermore logical functions contained in

- logic.h
- logic.cc

are used to check the dimensions of input matrices for various functions.

History

020502 Created and written by Claudia Emde.

7.2 Linear Equation Systems

For solving a set of linear equations

$$\mathbf{A}\mathbf{x} = \mathbf{b} \tag{7.1}$$

the LU decomposition method is implemented. A slightly modified version of the algorithm described in [*Press et al.* [1997]] is used here. An alternative method is the Gauss-Jordan elimination, but this method is three times slower than the LU decomposition method [*Press et al.* [1997], p.36]. The LU decomposition method reqires two functions, ludemp and lubacksub, which will be decribed below.

The following example for a three dimensional equation system demonstrates how to solve a linear equation system of the type (7.1):

• Create matrix A, vector b:

```
A = Matrix(3,3);

A(1,1) = 4;

A(2,1) = 3;

...

b = Vector(3);

b(1) = 7;

...
```

• Initialize solution vector x and two other variables needed for storing intermediate results:

```
x = Vector(3);
LU = Matrix(3,3);
indx = ArrayOfIndex(3);
```

- Call LU decomposition function (see Section 7.2.1):
 - ludcmp(LU, indx, A);
- Call LU backsubstitution function (see Section 7.2.2): lubacksub(x, LU, b, indx);
- Print the solution vector:

```
cout << x;
```

7.2.1 LU Decomposition

A LU decomposition is a procedure for decomposing a square matrix A with dimension n into a product of a lower triangular matrix L (has elements only on the diagonal elements and below) and an upper triangular matrix U (has elements only on the diagonal and above):

$$\mathbf{L} \cdot \mathbf{U} = \mathbf{A} \tag{7.2}$$

For a 3 x 3 matrix equation 7.2 would look like this:

$$\begin{pmatrix} l_{11} & 0 & 0 \\ l_{21} & l_{22} & 0 \\ l_{31} & l_{32} & l_{33} \end{pmatrix} \cdot \begin{pmatrix} u11 & u_{12} & u_{13} \\ 0 & u_{22} & u_{23} \\ 0 & 0 & u_{33} \end{pmatrix} = \begin{pmatrix} a_{11} & a_{12} & a_{13} \\ a_{21} & a_{22} & a_{23} \\ a_{31} & a_{32} & a_{33} \end{pmatrix}$$

The decomposition can be used to rewrite the linear set of equations (7.1) in the following way:

$$\mathbf{A} \cdot \mathbf{x} = (\mathbf{L} \cdot \mathbf{U}) \cdot \mathbf{x} = \mathbf{L} \cdot (\mathbf{U} \cdot \mathbf{x}) = \mathbf{b} \tag{7.3}$$

First

$$\mathbf{L} \cdot \mathbf{y} = \mathbf{b} \tag{7.4}$$

is solved for the vector \mathbf{y} which can be done by forward substitution (see section 7.2.2). Then

$$\mathbf{U} \cdot \mathbf{x} = \mathbf{y} \tag{7.5}$$

is solved again by backsubstitution. The advantage in breaking up one linear set into two successive ones is that the solution of a triangular set of equations is quite trivial.

The function ludemp requires a square matrix of arbitrary dimension n as input and performs the LU decomposition. It returns one matrix which contains both matrices, \mathbf{L} and \mathbf{U} . For the lower triangular matrix \mathbf{L} the diagonal elements are chosen to be 1, then the other elements of \mathbf{L} and \mathbf{U} are determined. This is possible, as the LU decomposition is an under determined equation system with n^2 equations for $n^2 + n$ unknowns. The output matrix does not include the diagonal of \mathbf{L} , in the three-dimensional case it has the following elements:

$$\begin{pmatrix}
u_{11} & u_{12} & u_{13} \\
l_{21} & u_{22} & u_{23} \\
l_{31} & l_{32} & u_{33}
\end{pmatrix}$$

This special arrangement of the LU decomposition is named *Crout's algorithm* and a matrix arranged in this form is named *Crout matrix* in this context.

Another output variable of the function ludemp is an index vector which contains information about pivoting which is absolutely essential for the stability of Crout's algorithm. Here partial pivoting, i.e. interchange of rows is implemented. That means that not $\bf A$ is decomposed into LU-form but a rowwise permutation of $\bf A$. If the index vector contains for example the elements (2,1,0) the first and the last row of a three dimensional matrix would be exchanged.

7.2.2 Forward- and Backsubstitution

An equation system of the form

$$\begin{pmatrix} a11 & a_{12} & a_{13} \\ 0 & a_{22} & a_{23} \\ 0 & 0 & a_{33} \end{pmatrix} \cdot \begin{pmatrix} x_1 \\ x_2 \\ x_3 \end{pmatrix} = \begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix}$$

can be solved very easy. The last element, here x_3 , is already isolated, namely

$$x_3 = b_3/a_{33} (7.6)$$

As x_3 is known x_2 can be calculated using the second row of the equations. Then, finally, x_1 can be calculated as well using the first row. This procedure is called backsubtitution. The

same method applied for an equation system including a lower triangular matrix is named forward substitution.

The function lubacksub does forward and backward substitution to solve the equation system described in 7.2.1. As input it requires the output variables of ludemp which are the *Crout matrix* and the index vector. Output of the function is the solution vector \mathbf{x} to the equation system.

7.2.3 More Applications of the LU Decomposition

• Inverse of a matrix:

To compute $(\mathbf{K})^{-1} \cdot \mathbf{b}$, which is a part of the solution to the vector radiative transfer equation (Equation 10.10 in *ARTS User Guide*) the LU decomposition method can be used. The following equations show, that the problem is equivalent to solving a linear equation system of the type 7.1.

$$\mathbf{K}^{-1} \cdot \mathbf{b} = \mathbf{x} \tag{7.7}$$

$$\Leftrightarrow \quad \mathbf{K} \cdot \mathbf{x} = \mathbf{b} \tag{7.8}$$

• To solve the equation system

$$\mathbf{A} \cdot \mathbf{X} = \mathbf{B} \tag{7.9}$$

where A, B and X are matrices of dimension n, the LU decomposition functions can be applied as well. Assume that A and B are known and you want to solve for X. First you should do a LU decomposition of A and then backsubstitute with the columns of B and you get the columns of A as solution vectors.

7.3 Matrix Exponential Function

A very important function for solving differential equations is the matrix exponential:

$$e^{\mathbf{A}s} = \sum_{k=0}^{\infty} \frac{(\mathbf{A}s)^k}{k!} \tag{7.10}$$

In principle it could be computed using the Taylor power series but this method is not efficient. MOLER and VAN LOAN have shown for the simple example [*Moler and Loan* [1979]]

$$\mathbf{A} = \left(\begin{array}{cc} -49 & 24 \\ -64 & 31 \end{array} \right)$$

that convergence is obtained not until 59 terms. And if a relative accuracy of only 10^{-5} is taken, the method even leads to a wrong result due to rounding errors.

7.3.1 Padé Approximation

One of the better algorithms for computing the matrix exponential is the Padé approximation which is also shortly described in [*Moler and Loan* [1979]] and outlined in the book "Matrix Computations" by *Golub and Loan* [1991]. The method uses perturbation theorie as well as the so called Padé functions. It is possible to derive an algorithm which calculates

$$\mathbf{F} = e^{\mathbf{A} + \mathbf{E}} \tag{7.11}$$

where

$$\|\mathbf{E}\|_{\infty} \le \delta \|\mathbf{A}\|. \tag{7.12}$$

The accuracy of the computation given by δ can be chosen. The parameter q has to be the smallest non-negative integer such that $\epsilon(q,q) \leq \delta$ where

$$\epsilon(p,q) = 2^{3-(p+q)} \frac{p!q!}{(p+q)!(p+q+1)!}. (7.13)$$

The following table shows values of epsilon for different values of q.

| q | $\epsilon(q,q)$ |
|---|-------------------------|
| 1 | 0.1667 |
| 2 | $6.9444 \cdot 10^{-4}$ |
| 3 | $1.2401 \cdot 10^{-6}$ |
| 4 | $1.2302 \cdot 10^{-9}$ |
| 5 | $7.7667 \cdot 10^{-13}$ |
| 6 | $3.3945 \cdot 10^{-16}$ |

The algorithm is implemented in the function $matrix_exp$. Input to this function is the matrix **A** and the parameter q. As output it gives the matrix **F** which is defined above. The following example shows how to use the $matrix_exp$ function:

• Initialize A and assign values:

```
Matrix A(3,3);

A(1,1) = 45;

A(1,2) = 3;

...
```

• Initialize **F**:

```
Matrix F(3,3);
```

• Give a paramater for the accuracy: Index q=6;

 Call the matrix exponential function: matrix_exp(F, A, q);

• Print the result:
 cout << "exp(A) = " << F;</pre>

Part I Bibliography and Appendices

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