1 SPH w/ nuclear reaction

Suppose we know acceleration, time derviative of energy, coefficient viscosity, and the next timestep at the initial time. We integrate each particle as follows:

1. Calculate the position, and predict the velocity and energy:

$$\boldsymbol{r}_i = \boldsymbol{r}_i^{(0)} + \boldsymbol{v}_i^{(0)}\Delta t + \frac{1}{2}\boldsymbol{a}_i^{(0)}\Delta t^2, \tag{1}$$

$$\boldsymbol{v}_i^{(1/2)} = \boldsymbol{v}_i^{(0)} + \frac{1}{2}\boldsymbol{a}^{(0)}\Delta t, \tag{2}$$

$$u_i^{(1/2)} = u_i^{(0)} + \frac{1}{2}\dot{u}^{(0)}\Delta t + \epsilon_{\mathrm{nuc},i}(\Delta t/2), \tag{3}$$

$$\alpha_i^{(1/2)} = \alpha_i^{(0)} + \frac{1}{2}\dot{\alpha}^{(0)}\Delta t, \tag{4}$$

$$\alpha_i^{\mathrm{u},(1/2)} = \alpha_i^{\mathrm{u},(0)} + \frac{1}{2}\dot{\alpha}^{\mathrm{u},(0)}\Delta t, \tag{5}$$

$$\tilde{\boldsymbol{v}}_i = \boldsymbol{v}_i^{(0)} + \boldsymbol{a}_i^{(0)}\Delta t, \tag{6}$$

$$\tilde{\boldsymbol{u}}_i = \boldsymbol{u}_i^{(0)} + \dot{\boldsymbol{u}}_i^{(0)}\Delta t + \epsilon_{\mathrm{nuc},i}(\Delta t), \tag{7}$$

$$\tilde{\alpha}_i = \alpha_i^{(0)} + \dot{\alpha}_i^{(0)}\Delta t \tag{8}$$

$$\tilde{\alpha}_i^{\mathrm{u}} = \alpha_i^{\mathrm{u},(0)} + \dot{\alpha}_i^{\mathrm{u},(0)}\Delta t, \tag{9}$$

   where ($\epsilon_{\mathrm{nuc},i}$) is energy generated through nuclear reaction. When the reaction is exothermic, the density and temperature are fixed at this time. On the other hand, when the reaction is endothermic, only the density is fixed at this time.

2. Calculate the density, kernel-support length and grad-h term of each particle by an iterative method:

   (a) Estimate the density by the following expression:

   $$\rho_i = \sum_j m_j W_i \tag{10}$$

   where $W_i = W(r_{ij}, H_i)$, $r_{ij} = |\boldsymbol{r}_{ij}|$, and $\boldsymbol{r}_{ij} = \boldsymbol{r}_i - \boldsymbol{r}_j$. The kernel function is expressed as

   $$W(r, H) = H^{-D} w(r/H). \tag{11}$$

   The formula of $w(q)$ is described in Appendix A.

   (b) Calculate the kernel length as follow:

   $$H_i = \max\left[C_{H/h} \times \eta \left(\frac{m_i}{\rho_i}\right)^{1/D}, H_{\max}\right], \tag{12}$$

   where h is the kernel length. Here, $H_{\max}$ is the initial distance of the binary separation, or the maximum `double` in the case of the single WD. The values of η and $C_{H/h}$ are described in Appendix A.

(c) Return to step (2a) unless this is the 3rd time.

(d) Calculate divergence and rotation of $\boldsymbol{v}$, grad-h term, and gravity correction term:

$$\nabla \cdot \tilde{\boldsymbol{v}}_i = -\frac{1}{\Omega_i \rho_i} \sum \tilde{\boldsymbol{v}}_{ij} \cdot \left[m_j \frac{\partial W_i}{\partial r_{ij}} \frac{\boldsymbol{r}_{ij}}{r_{ij}} \right], \tag{13}$$

$$\nabla \times \tilde{\boldsymbol{v}}_i = \frac{1}{\Omega_i \rho_i} \sum \tilde{\boldsymbol{v}}_{ij} \times \left[m_j \frac{\partial W_i}{\partial r_{ij}} \frac{\boldsymbol{r}_{ij}}{r_{ij}} \right], \tag{14}$$

$$\Omega_i = 1 + \frac{1}{D} \frac{H_i}{\rho_i} \sum_j m_j \frac{\partial W_i}{\partial H_i}, \tag{15}$$

The derivatives of the kernel function are as follows:

$$\frac{\partial W(r, H)}{\partial r} = H^{-(D+1)} \frac{\partial w(r/H)}{\partial (r/H)} \tag{16}$$

$$\frac{\partial W(r, H)}{\partial H} = -H^{-(D+1)} \left[D w(q) + q \frac{\partial w}{\partial q} \right]. \tag{17}$$

The formula of $\partial w / \partial q$ is described in Appendix A.

3. Calculate the pressure ($\tilde{P}_i$) and sound speed ($\tilde{c}_{\mathrm{s},i}$) as follows:

$$P_i = (\gamma - 1) \rho_i \tilde{u}_i \tag{18}$$

$$c_{\mathrm{s},i} = \left(\gamma \frac{P_i}{\rho_i} \right)^{1/2}, \tag{19}$$

or Helmholtz EOS, in which case the temperature ($\tilde{T}_i$) is also obtained.

4. Calculate the Balsara switch:

$$f_i = \frac{|\nabla \cdot \tilde{\boldsymbol{v}}_i|}{|\nabla \cdot \tilde{\boldsymbol{v}}_i| + |\nabla \times \tilde{\boldsymbol{v}}_i| + 0.0001 \tilde{c}_{\mathrm{s},i} / h_i}. \tag{20}$$

5. Calculate the gravity and its correction term:

$$\boldsymbol{g}_{2,i} = -\frac{G}{2} \sum_j m_j \frac{\boldsymbol{r}_{ij}}{r_{ij}} \left[\frac{\partial \phi(r_{ij}, H_i)}{\partial r_{ij}} + \frac{\partial \phi(r_{ij}, H_j)}{\partial r_{ij}} \right], \tag{21}$$

$$\phi_i = \frac{G}{2} \sum_j m_j \left[\phi(r_{ij}, H_i) + \phi(r_{ij}, H_j) \right] \tag{22}$$

$$\eta_i = \frac{H_i}{D \rho_i} \frac{1}{\Omega_i} \sum_j m_j \frac{\partial \phi(r_{ij}, H_i)}{\partial H_i}, \tag{23}$$

where the potential is expressed as:

$$\phi(r,H) = -\frac{1}{\left(r^2 + H^2 C_{H/h}^{-2}\right)^{1/2}} \tag{24}$$

$$\frac{\partial \phi(r,H)}{\partial r} = \frac{r}{\left(r^2 + H^2 C_{H/h}^{-2}\right)^{3/2}}, \tag{25}$$

$$\frac{\partial \phi(r,H)}{\partial H} = \frac{H C_{H/h}^{-2}}{\left(r^2 + H^2 C_{H/h}^{-2}\right)^{3/2}}. \tag{26}$$

6. Calculate the hydro acceleration, the time derivative of the energy, and one term of the gravity:

$$\boldsymbol{a}_i = -\frac{1}{2}\sum_j \left(\frac{\tilde{P}_i}{\Omega_i \rho_i^2} + \frac{\tilde{P}_j}{\Omega_j \rho_j^2} + f_{ij}\Pi_{ij}\right)\left[m_j\left(\frac{\partial W_i}{\partial r_{ij}} + \frac{\partial W_j}{\partial r_{ij}}\right)\frac{\boldsymbol{r}_{ij}}{r_{ij}}\right], \tag{27}$$

$$\dot{u}_i = \frac{1}{2}\sum_j \left(\frac{\tilde{P}_i}{\Omega_i \rho_i^2} + \frac{f_{ij}\Pi_{ij}}{2}\right)\left[m_j\left(\frac{\partial W_i}{\partial r_{ij}} + \frac{\partial W_j}{\partial r_{ij}}\right)\frac{\boldsymbol{r}_{ij}}{r_{ij}}\right]\cdot \tilde{\boldsymbol{v}}_{ij}, \tag{28}$$

$$\boldsymbol{g}_{1,i} = \frac{G}{2}\sum_j m_j \left[\left(\eta_i \frac{\partial W_i}{\partial r_{ij}} + \eta_j \frac{\partial W_j}{\partial r_{ij}}\right)\frac{\boldsymbol{r}_{ij}}{r_{ij}}\right], \tag{29}$$

$$\left(\nabla^2 u\right)_i = \sum_j \frac{u_i - u_j}{\rho_{ij}}\left[m_j\left(\frac{\partial W_i}{\partial r_{ij}} + \frac{\partial W_j}{\partial r_{ij}}\right)\frac{1}{r_{ij}}\right], \tag{30}$$

where $f_{ij} = (f_i + f_j)/2$, and Π_{ij} is an artificial viscosity. The artificial viscosity is expressed as follows:

$$\Pi_{ij} = -\frac{\tilde{\alpha}_{ij}}{2}\frac{v_{ij}^{\rm sig}\min\left(w_{ij},0\right)}{\rho_{ij}} \tag{31}$$

$$v_{ij}^{\rm sig} = c_{{\rm s},i} + c_{{\rm s},j} - 3\min(w_{ij},0) \tag{32}$$

$$w_{ij} = \frac{\boldsymbol{r}_{ij}\cdot\tilde{\boldsymbol{v}}_{ij}}{|\boldsymbol{r}_{ij}|} \tag{33}$$

where $\tilde{\alpha}_{ij} = (\tilde{\alpha}_i + \tilde{\alpha}_j)/2$, and $\rho_{ij} = (\rho_i + \rho_j)/2$. If we introduce Price's thermal conductivity, equation (28) can be rewritten as:

$$\begin{aligned}\dot{u}_i =& \frac{1}{2}\sum_j m_j \left(\frac{\partial W_i}{\partial r_{ij}} + \frac{\partial W_j}{\partial r_{ij}}\right) \\ &\times \left\{ \left(\frac{\tilde{P}_i w_{ij}}{\Omega_i \rho_i^2} + \frac{f_{ij}\Pi_{ij}}{2}\right) w_{ij} \right. \\ &\left. + \frac{1}{\rho_{ij}}\left[\frac{1}{4}\tilde{\alpha}_{ij} v_{ij}^{\mathrm{sig}}\left(\hat{v}_i^2 - \hat{v}_j^2\right) + \tilde{\alpha}_{ij}^{\mathrm{u}} v_{ij}^{\mathrm{u,sig}}\left(u_i - u_j\right)\right]\right\} \end{aligned} \tag{34}$$

where $\tilde{\alpha}_{ij}^{\mathrm{u}} = 1$ temporarily, and

$$\hat{v}_i = \frac{\boldsymbol{v}_i \cdot \boldsymbol{r}_{ij}}{r_{ij}} \tag{35}$$

$$\hat{v}_j = \frac{\boldsymbol{v}_j \cdot \boldsymbol{r}_{ij}}{r_{ij}} \tag{36}$$

$$v_{ij}^{\mathrm{u,sig}} = \left(\frac{|\tilde{P}_i - \tilde{P}_j|}{\rho_{ij}}\right)^{1/2} \tag{37}$$

7. Calculate $\dot{\alpha}$ and $\dot{\alpha}^{\mathrm{u}}$:

$$\dot{\alpha}_i = -\frac{\tilde{\alpha}_i - \alpha_{\min}}{h_i/(0.25\tilde{c}_{\mathrm{s},i})} + \max\left[-(\nabla\cdot\tilde{\boldsymbol{v}}_i)(\alpha_{\max} - \tilde{\alpha}_i), 0\right] \tag{38}$$

$$\dot{\alpha}_i^{\mathrm{u}} = -\frac{\tilde{\alpha}_i^{\mathrm{u}} - \alpha_{\min}^{\mathrm{u}}}{h_i/(0.25\tilde{c}_{\mathrm{s},i})} + \frac{h_i|\nabla^2 u|_i}{(u_i + \epsilon^{\mathrm{u}})^{1/2}}\left(\alpha_{\max} - \tilde{\alpha}_i^{\mathrm{u}}\right), \tag{39}$$

where we set $\epsilon^{\mathrm{u}} = 0.0001 u_{0,\min}$.

8. Correct the velocity and energy:

$$\boldsymbol{v}_i = \boldsymbol{v}_i^{(1/2)} + \frac{1}{2}\boldsymbol{a}_i \Delta t \tag{40}$$

$$u_i = u_i^{(1/2)} + \frac{1}{2}\dot{u}_i \Delta t + \epsilon_{\mathrm{nuc},i}(\Delta t/2) \tag{41}$$

$$\alpha_i = \alpha_i^{(1/2)} + \frac{1}{2}\dot{\alpha}_i \Delta t, \tag{42}$$

$$\alpha_i^{\mathrm{u}} = \alpha_i^{\mathrm{u},(1/2)} + \frac{1}{2}\dot{\alpha}_i^{\mathrm{u}} \Delta t. \tag{43}$$

When the reaction is exothermic, the density and temperature are fixed at this time. On the other hand, when the reaction is endothermic, only the density is fixed at this time.

9. Calculate the next timestep:

$$\Delta t_{\mathrm{next}} = C \min_i \left[\frac{2H_i}{\max_j\left(v_{ij}^{\mathrm{sig}}\right)}, \Delta t \left|\frac{u_i^{(0)}}{u_i - u_i^{(0)}}\right|\right]. \tag{44}$$

10. Return to step 1.

2 Helmholtz EOS

Coorporate Timmes's EOS. The compositions are 100 % carbon, 50 % carbon and 50 % oxygen, and will be 100 % helium.

A Kernels

A.1 Cubic spline kernel

We describe the cubic spline kernel:

$$w(r/H) = 2^D\sigma \times \begin{cases} \left(1 - 6q^2 + 6q^3\right) & (0 \le q < 1/2) \\ \left[2(1-q)^3\right] & (1/2 \le q < 1) \\ 0 & (q \ge 1) \end{cases} \tag{45}$$

$$\frac{\partial w(q)}{\partial q} = (-6)2^D\sigma \times \begin{cases} 1/3 & (0 \le q < 1/3) \\ \left(2q - 3q^2\right) & (1/3 \le q < 1/2) \\ (1-q)^2 & (1/2 \le q < 1) \\ 0 & (q \ge 1) \end{cases}, \tag{46}$$

where $\sigma = 2/3$ (1D), $10/7\pi$ (2D), and $1/\pi$ (3D).

For this kernel, $\eta = 1.2$, and $C_{H/h} = 2$.

The equations (45) and (46) can be rewritten as

$$w(r/H) = 2^D\sigma \times 2\left[(1-q)_+^3 - 4(1/2-q)_+^3\right], \tag{47}$$

$$\frac{\partial w(q)}{\partial q} = (-6)2^D\sigma \times \left[(1-q)_+^2 - 4(1/2-q)_+^2 + 3(1/3-q)_+^2)\right], \tag{48}$$

where $(x)_+ \equiv \max(0, x)$.

A.2 Wendland C^2 kernel

We describe Wendland C^2 kernel: In the case of $D = 1$,

$$w(q) = C_{\mathrm{W}}(1-q)_+^3(1+3q) \tag{49}$$

$$\frac{\partial w(q)}{\partial q} = 3C_{\mathrm{W}}\left[(1-q)_+^3 - (1-q)_+^2(1+3q)\right], \tag{50}$$

and in the case of $D = 2, 3$,

$$w(q) = C_{\mathrm{W}}(1-q)_+^4(1+4q) \tag{51}$$

$$\frac{\partial w(q)}{\partial q} = 4C_{\mathrm{W}}\left[(1-q)_+^4 - (1-q)_+^3(1+4q)\right], \tag{52}$$

where $C_{\mathrm{W}} = 5/4$ $(D = 1)$, $7/\pi$ $(D = 2)$, and $21/(2\pi)$ $(D = 3)$. For this kernel, $\eta = 1.6$, and $C_{H/h} = 1.620185(D = 1)$, $1.897367(D = 2)$, and $1.93492(D = 3)$.

A.3 Wendland C^4 kernel

We describe Wendland C^4 kernel: In the case of $D = 1$,

$$w(q) = C_{\rm W}(1-q)_+^5(1+5q+8q^2) \quad (53)$$

$$\frac{\partial w(q)}{\partial q} = C_{\rm W}(1-q)_+^4\left[-5(1+5q+8q^2)+(1-q)_+(5+16q)\right] \quad (54)$$

and in the case of $D = 2, 3$,

$$w(q) = C_{\rm W}(1-q)_+^6\left[1+6q+(35/3)q^2\right] \quad (55)$$

$$\frac{\partial w(q)}{\partial q} = C_{\rm W}(1-q)_+^5\left\{-6\left[1+6q+(35/3)q^2\right]+(1-q)_+\left[6+(70/3)q\right]\right\} \quad (56)$$

where $C_{\rm W} = 3/2$ ($D = 1$), $9/\pi$ ($D = 2$), and $495/(32\pi)$ ($D = 3$). For this kernel, $\eta = 1.6$, and $C_{H/h} = 1.936492(D = 1)$, $2.171239(D = 2)$, and $2.207940(D = 3)$.

B SPH tests

1. 1D shock tube (CUBICSPLINE, USE_AT1D)
2. 3D shock tube (CUBICSPLINE)
3. Strong shock (CUBICSPLINE, USE_AT1D)
4. Point like explosion (CUBICSPLINE)
5. Evrard test (CUBICSPLINE, USE_AT3D, GRAVITY)

C EOS tests

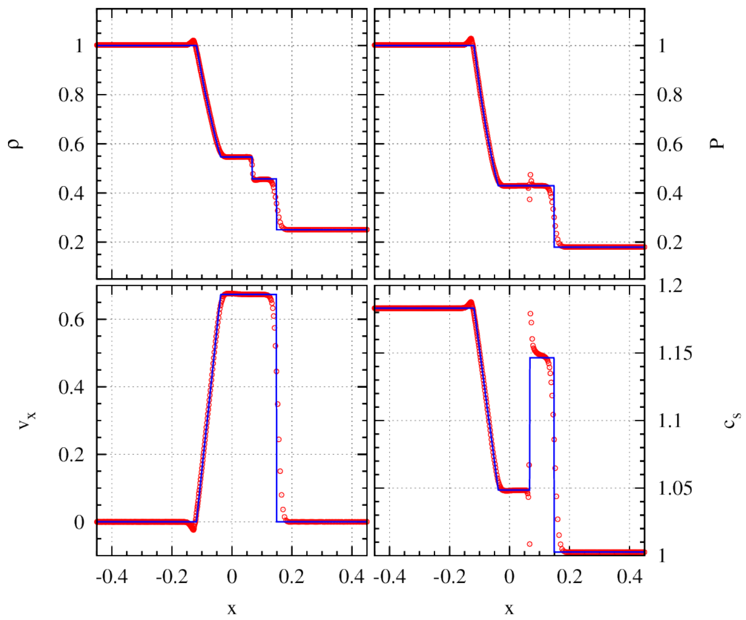

Figure 1: 1D shock tube (cubic spline, $\gamma = 1.4$, $\alpha = 1.0$).

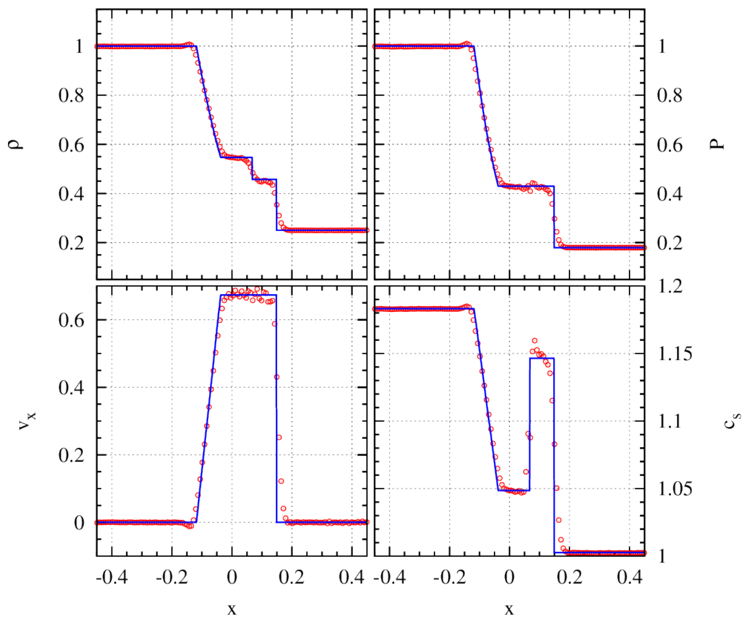

Figure 2: 3D shock tube (cubic spline, $\gamma = 1.4$, $\alpha = 1.0$).

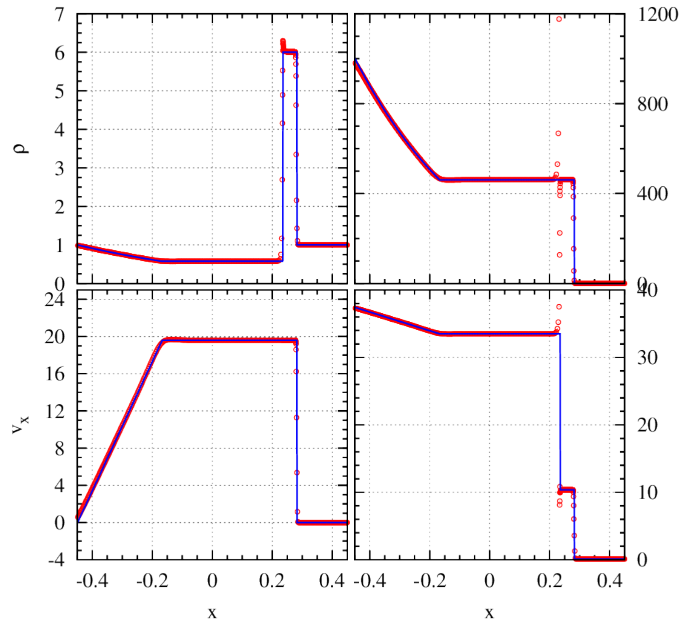

Figure 3: Strong shock (cubic spline, $\gamma = 1.4$, $\alpha = 1.0$).

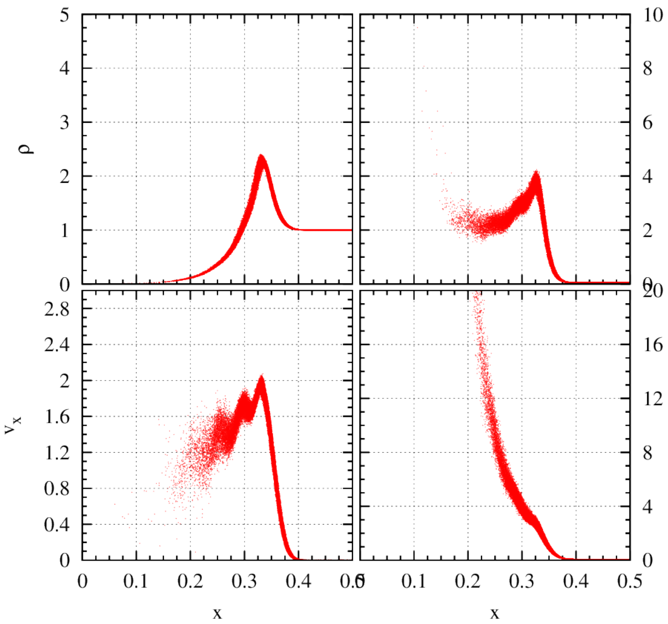

Figure 4: Point like explosion (cubic spline, $\gamma = 5/3$, $\alpha = 3.0$).

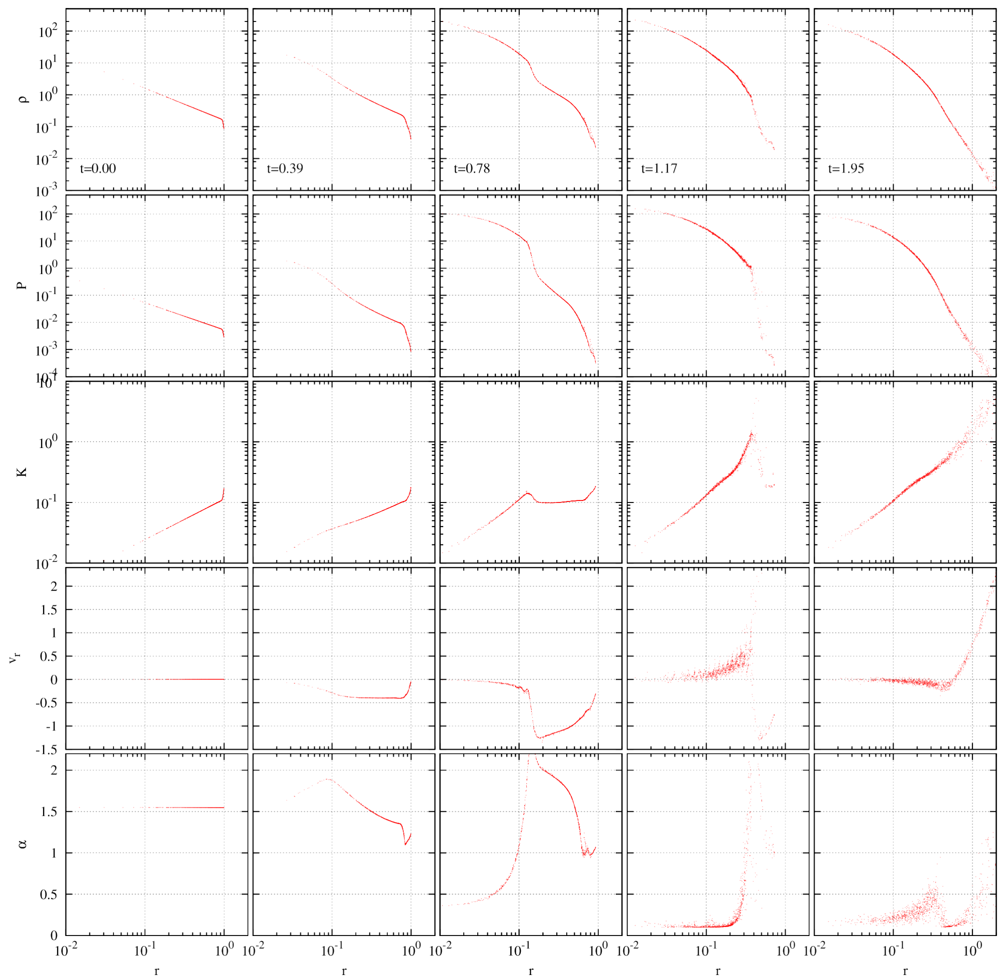

Figure 5: Evrard test (cubic spline, $\gamma = 5/3$, $\alpha_{\rm min} = 0.1$, $\alpha_{\rm max} = 3.0$).

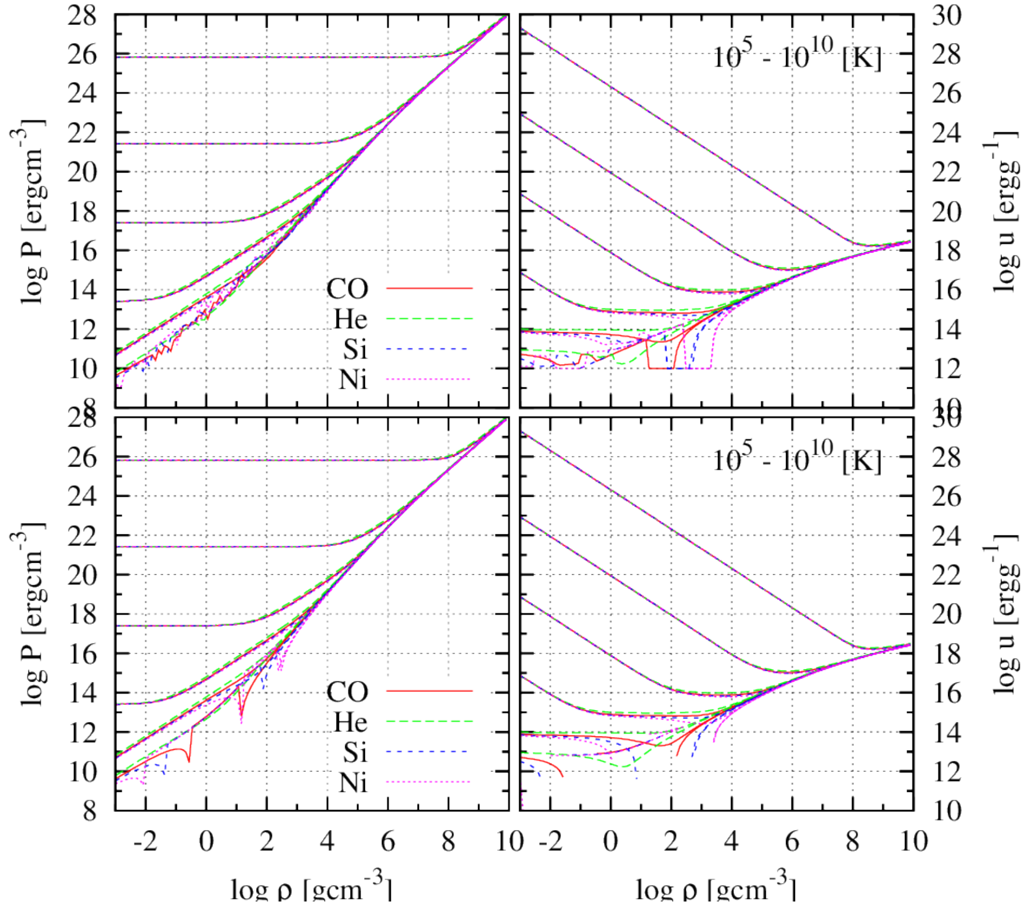

Figure 6: Helmholtz EOS w/ lookup table (top) and w/o (bottom).

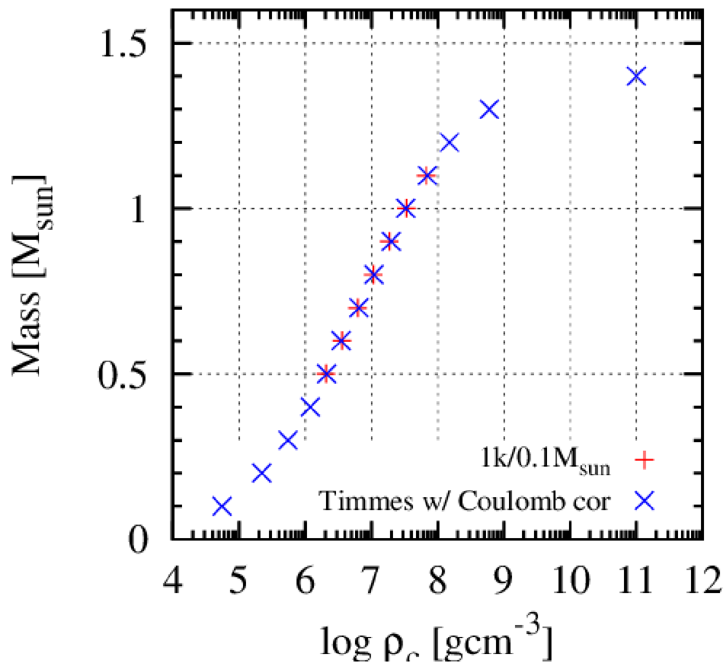

Figure 7: Single COWD equilibrium ($1k/0.1M_\odot$) w/ new damping. The new damping mode is the addtion of a sort of cooling during 100 s, such that $u_{\mathrm{new},i} = (u_{\mathrm{old},i} - u_{\min}(\rho_i)) \exp(-0.1\Delta t) + u_{\min}(\rho_i)$.

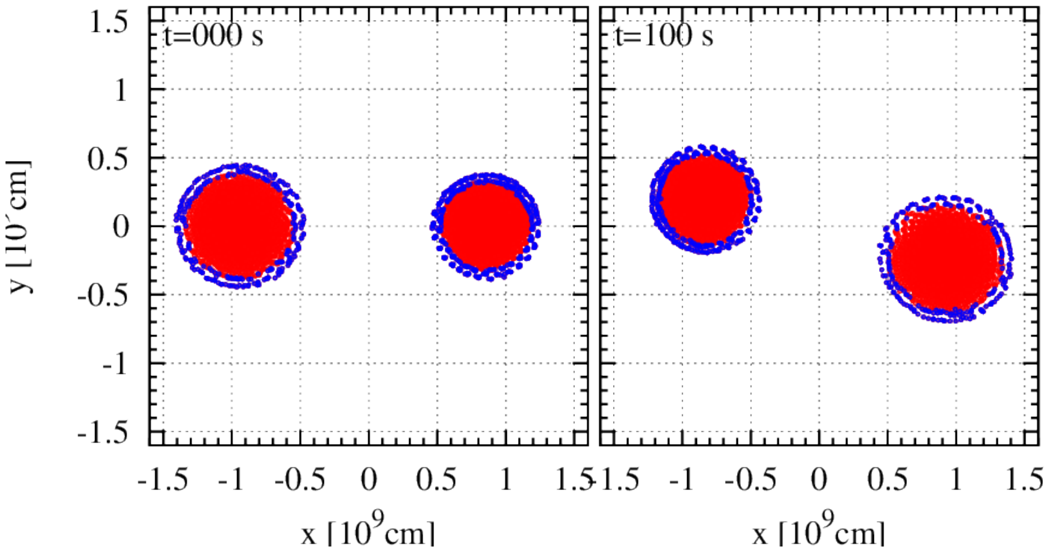

Figure 8: Density and temperature of $1.1M_{\odot}$ and $1.0M_{\odot}$ COWDs ($1k/0.1M_{\odot}$) from 1.8×10^9cm. Blue points indicate helium particles ($f_{\mathrm{He}} = 0.1$).

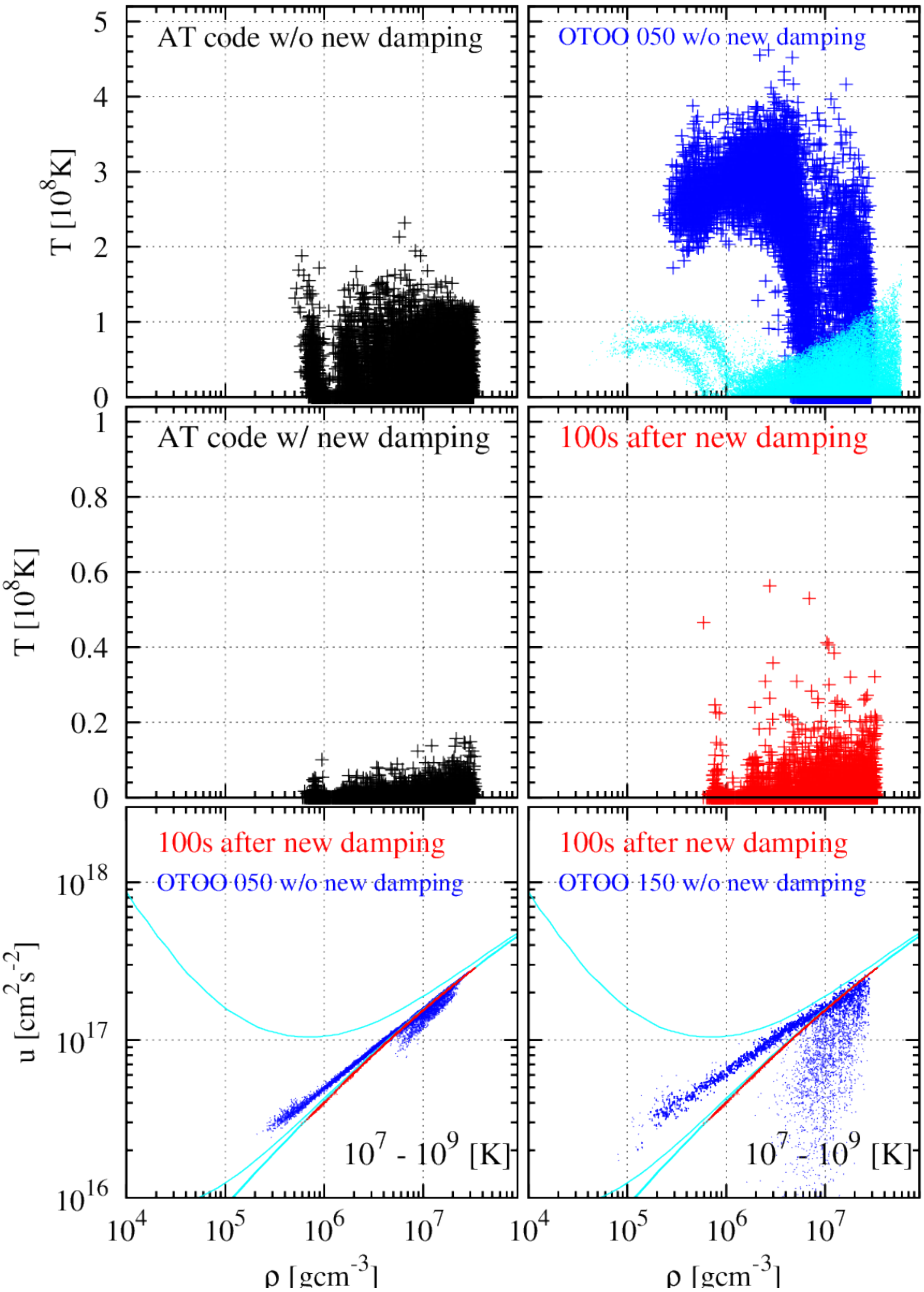

Figure 9: $1M_\odot$ COWD ($1k/0.1M_\odot$).

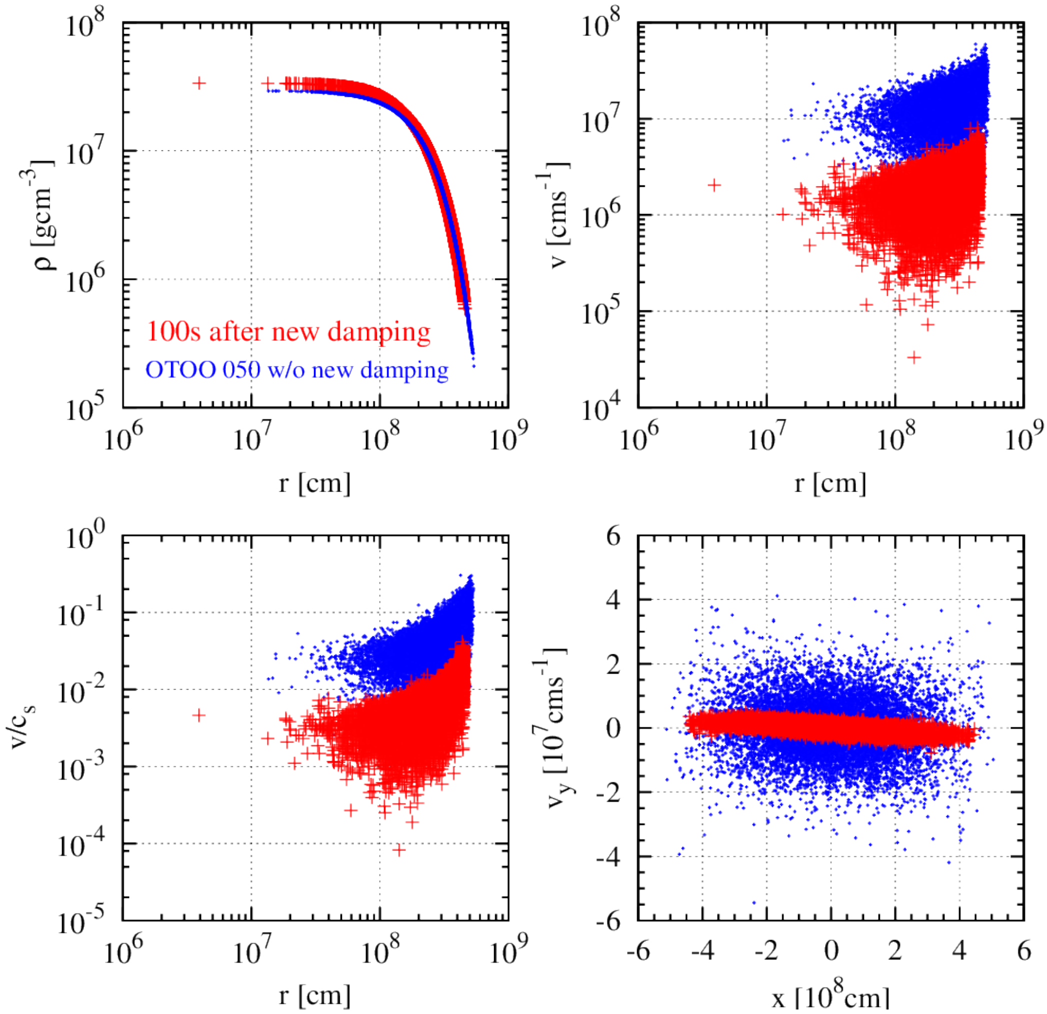

Figure 10: $1M_{\odot}$ COWD $(1k/0.1M_{\odot})$.

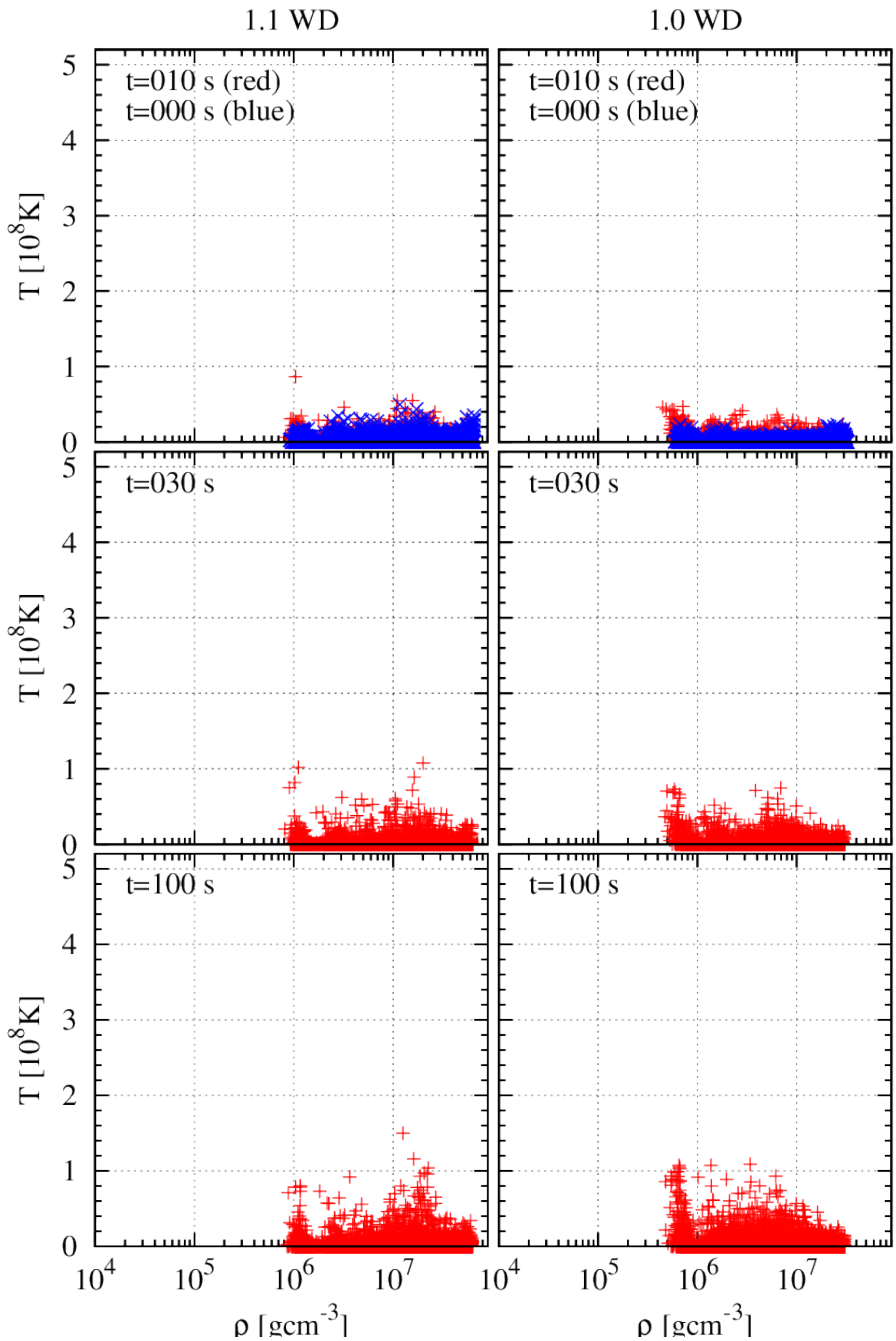

Figure 11: Density and temperature of $1.1M_{\odot}$ and $1.0M_{\odot}$ COWDs ($1k/0.1M_{\odot}$) from 1.8×10^9cm.

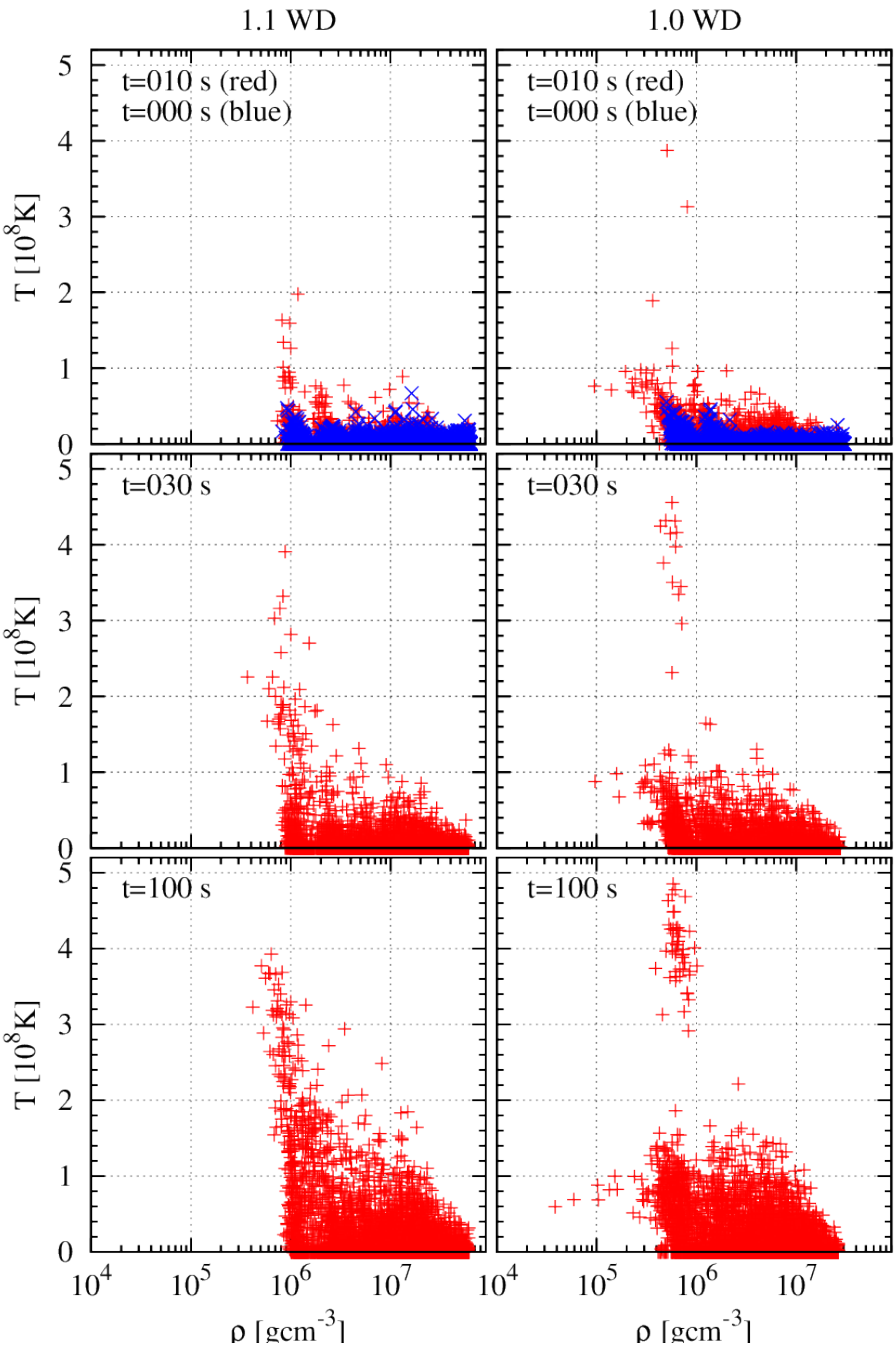

Figure 12: Density and temperature of $1.1M_\odot$ and $1.0M_\odot$ COWDs ($1k/0.1M_\odot$) from 1.5×10^9cm.