



# Implementing probabilistic photometric redshifts on DES SV data

Matías Carrasco Kind

Department of Astronomy
University of Illinois

#### DES SV data (fixed!)



VVDS Deep 02hr (3117)

CDFS (3721)

VVDS Wide 14hr (2970)

COSMOS (6148)

Secure redshifts 3 < ZFLAG < 4

 $0.01 \leq Z \leq 1.5$ 

15956 galaxies

7978 for training and 7978 for testing

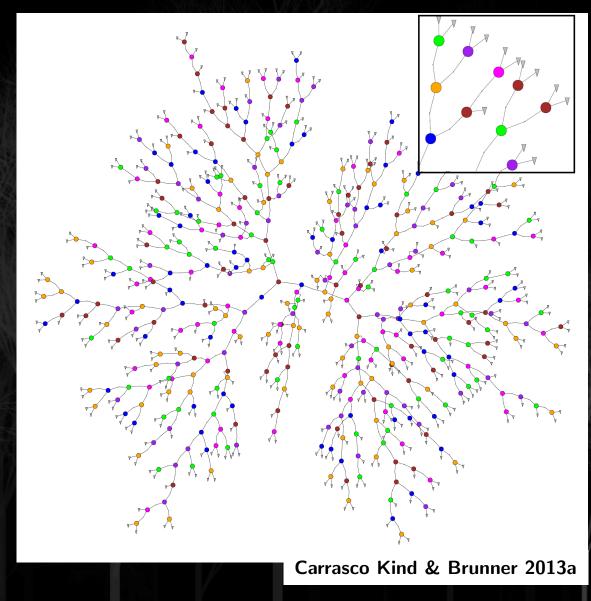
Use 5 bands from MAG\_AUTO and MAG\_DETMODEL

Use colors (8)

#### TPZ: Trees for Photo-Z



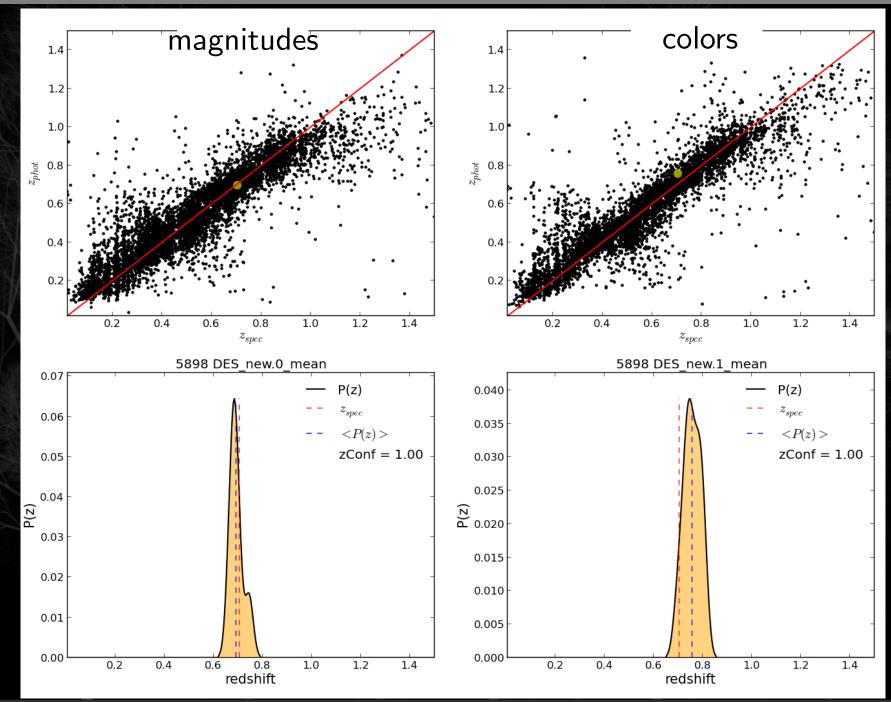
- ullet Provides photo-z PDF and confidence values
- Deals with missing data
- Includes measurements errors
- Provides useful ancillary information
- Out-of-Bag data for unbiased errors
- No need for validation set



http://lcdm.astro.illinois.edu/research/TPZ.html

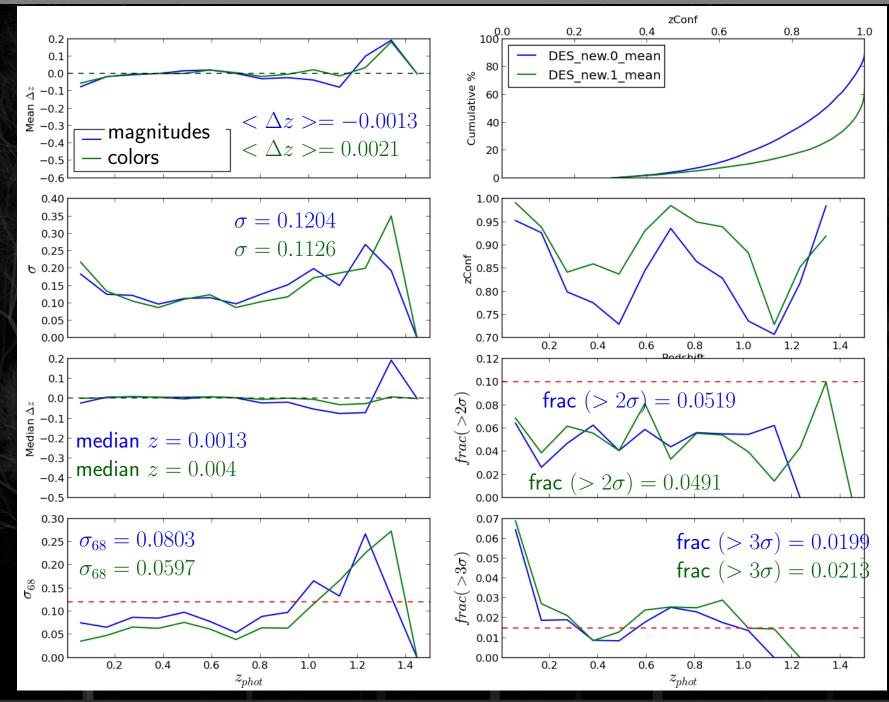
#### Preliminary results: True plots and PDF





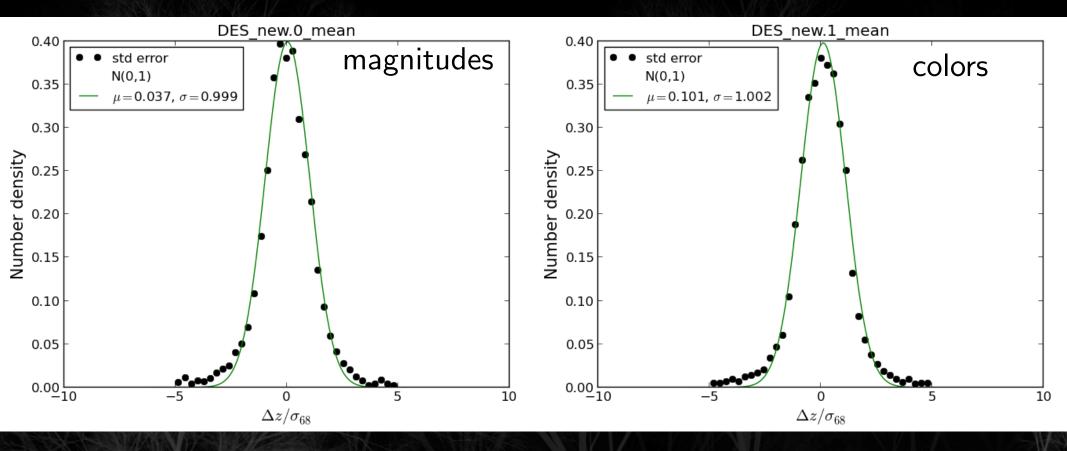
#### Preliminary results: Metrics





#### Preliminary results: Error distribution

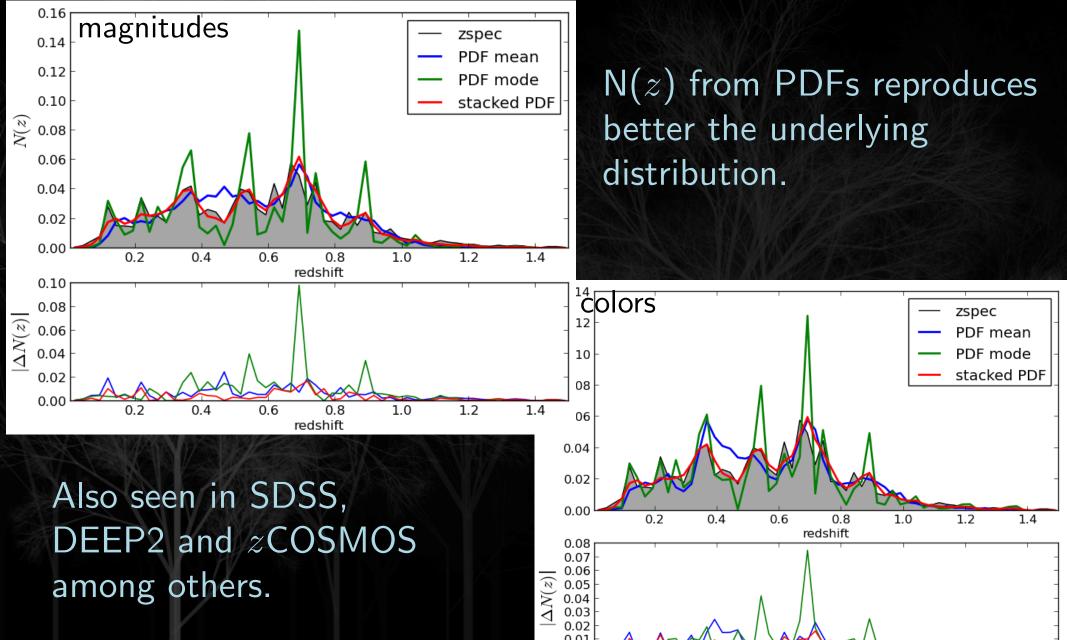




All these metrics were calculated using the mean of the probability density function

#### Preliminary results: Using PDFs for N(z)





0.00

1.0

1.2

1.4

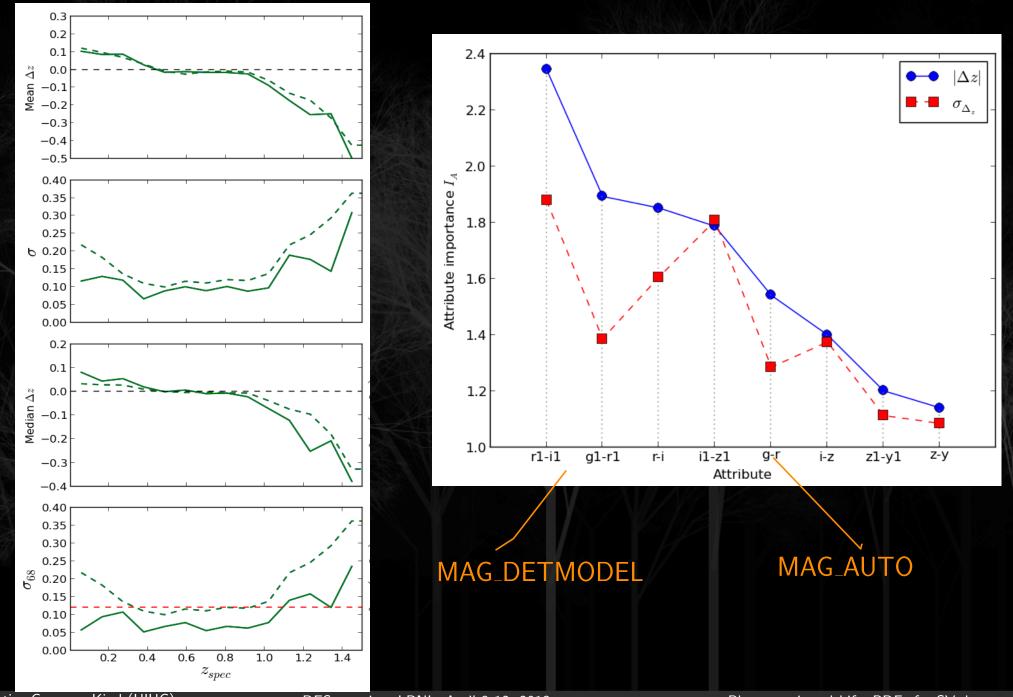
0.8

redshift

0.6

#### Ancillary examples: OOB error and importance





#### Combining techniques



Multiple "independent" techniques can be combined into a powerful one (Carrasco Kind & Brunner, 2013b, c in prep.)

We use a modified and parallel version of BPZ (Benitez, 2000) with a prior from Random Naïve Bayes Classifier method

Extend it using a 3rd technique from Self-Organized-Maps

Our approach

Template fitting

+

Prior information

+

Empirical method

+

Weigthing scheme



 $\overline{\mathsf{photo-}z}$   $\overline{\mathsf{PDF}}$ 

#### Template fitting photo-z



Suppose a set of templates T and n magnitudes  $m_1, m_2, ..., m_n$ , the probability is:

$$P(z|\vec{m}) = \sum_{T} P(z, T|\vec{m}) \propto \sum_{T} P(z, T|\vec{m}) P(\vec{m}|z, T)$$

where  $ec{m}=(m_1,m_2,...,m_n)$ 

#### Template fitting photo-z



Suppose a set of templates T and n magnitudes  $m_1, m_2, ..., m_n$ , the probability is:

$$P(z|\vec{m}) = \sum_T P(z,T|\vec{m}) \propto \sum_T P(z,T|\vec{m})P(\vec{m}|z,T)$$
 where  $\vec{m}=(m_1,m_2,...,m_n)$  Prior Likelihood

#### Template fitting photo-z

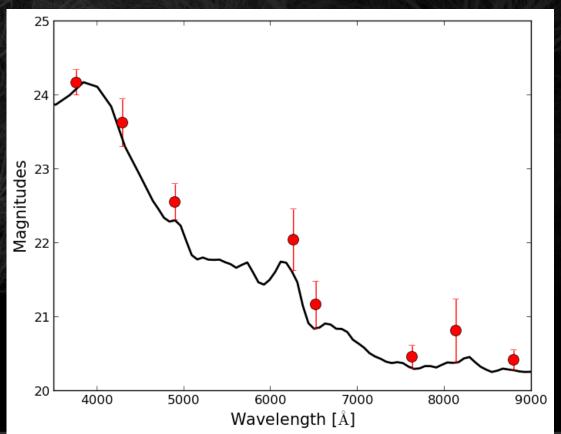


Likelihood

Suppose a set of templates T and n magnitudes  $m_1, m_2, ..., m_n$ , the probability is:

$$P(z|\vec{m}) = \sum_{T} P(z, T|\vec{m}) \propto \sum_{T} P(z, T|\vec{m}) P(\vec{m}|z, T)$$

where  $ec{m}=(m_1,m_2,...,m_n)$ 





DES meeting, LBNL, April 8-12, 2013

Photometric redshifts PDFs for SV data



$$P(z|\vec{m}) = \sum_{T} P(z,T|\vec{m}) \propto \sum_{T} P(z,T|\vec{m}) P(\vec{m}|z,T)$$
 Prior (Likelihood)

RNBC needs training set

Assume *naïvely*  $m_i \perp m_j$ 

Bootstrap samples, random subset of magnitudes



$$P(z|\vec{m}) = \sum_{T} P(z,T|\vec{m}) \propto \sum_{T} P(z,T|\vec{m}) P(\vec{m}|z,T)$$
 Prior (Likelihood)

RNBC needs training set

Assume *naïvely*  $m_i \perp m_j$ 

Bootstrap samples, random subset of magnitudes

RNBC prior 
$$\Rightarrow P(z,T|\vec{m}) \propto P(z)P(T|z) \prod_{i=1}^{n} P(m_i|z,T)$$



$$P(z|\vec{m}) = \sum_{T} P(z,T|\vec{m}) \propto \sum_{T} P(z,T|\vec{m}) P(\vec{m}|z,T)$$

$$P(z|\vec{m}) = \sum_{T} P(z,T|\vec{m}) = \sum_{T} P(z,T|\vec{m}) P(\vec{m}|z,T)$$

$$P(z|\vec{m}) = \sum_{T} P(z,T|\vec{m}) = \sum_{T} P(z,T|\vec{m}) P(\vec{m}|z,T)$$

$$P(z|\vec{m}) = \sum_{T} P(z,T|\vec{m}) = \sum_{T} P(z,T|\vec{m}) P(\vec{m}|z,T)$$

RNBC needs training set

Assume *naïvely*  $m_i \perp m_j$ 

Bootstrap samples, random subset of magnitudes

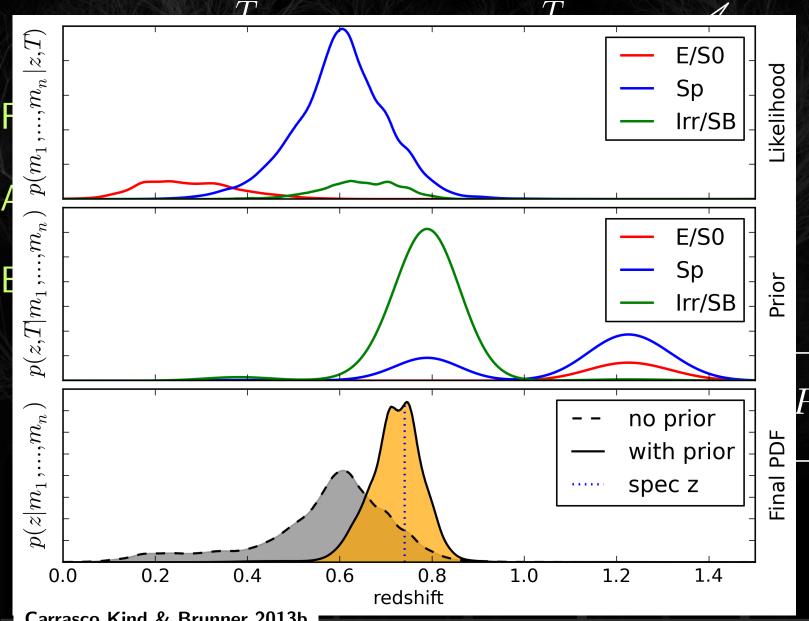
RNBC prior 
$$\Rightarrow P(z,T|\vec{m}) \propto P(z)P(T|z) \prod_{i=1}^{n} P(m_i|z,T)$$

PCA transformations

A model is assumed for this term



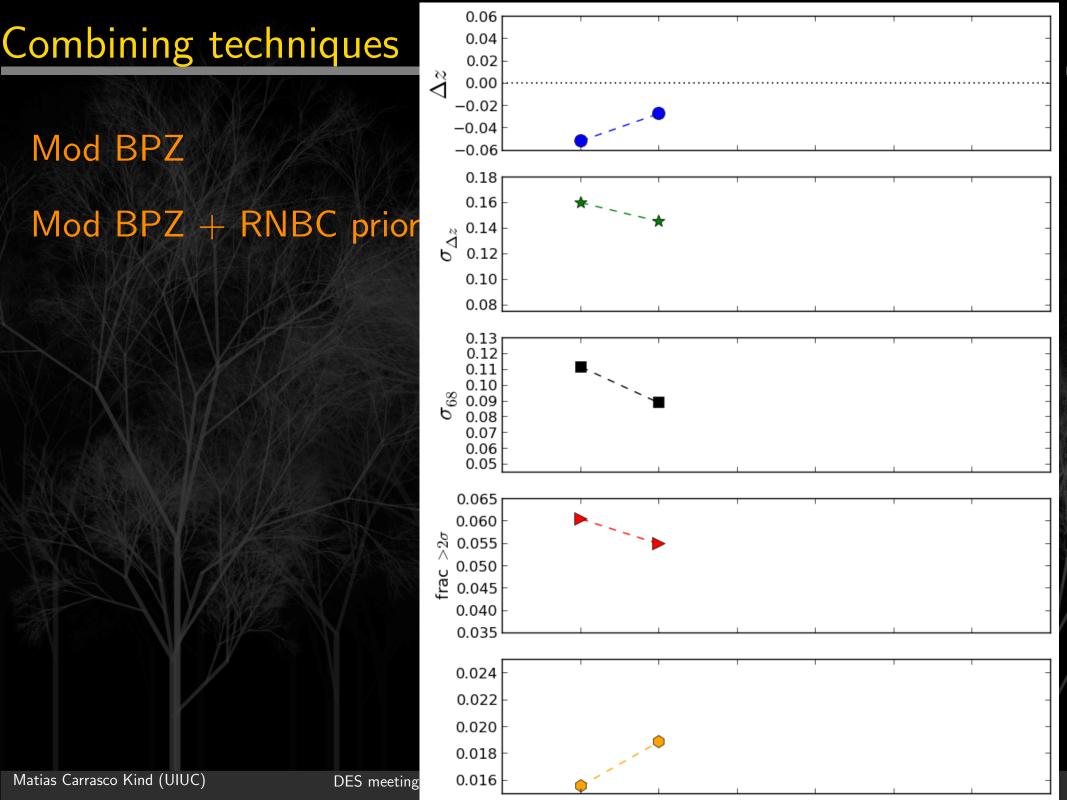


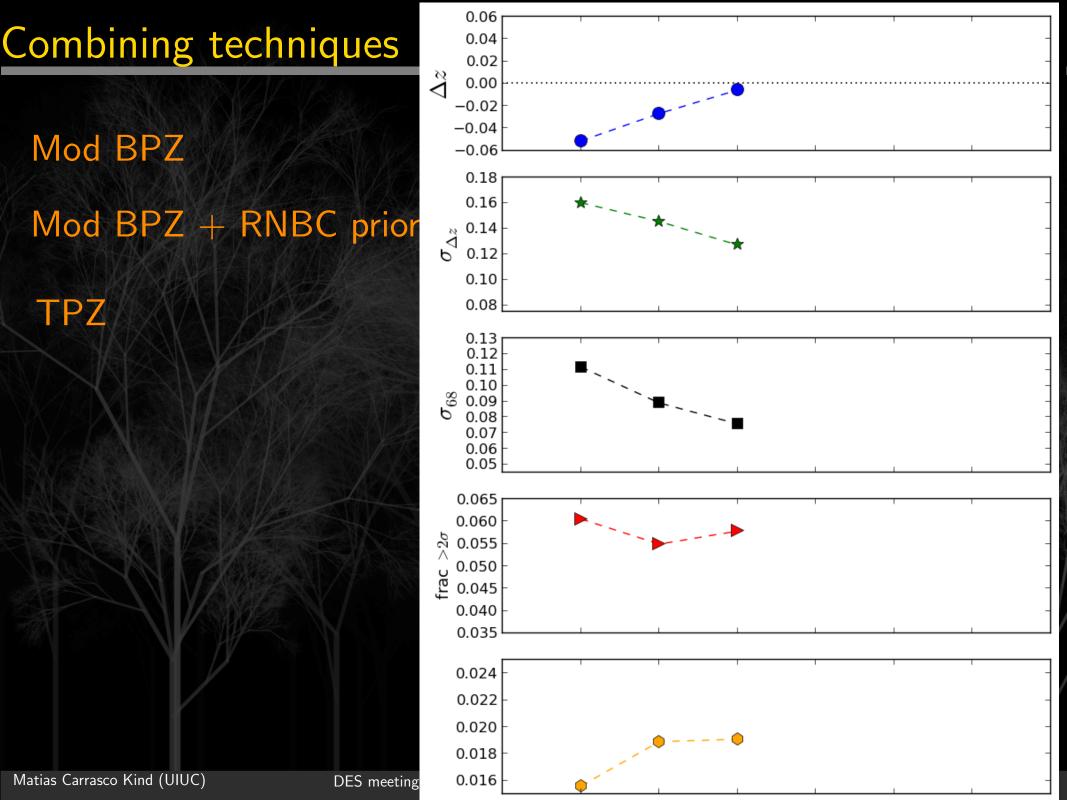


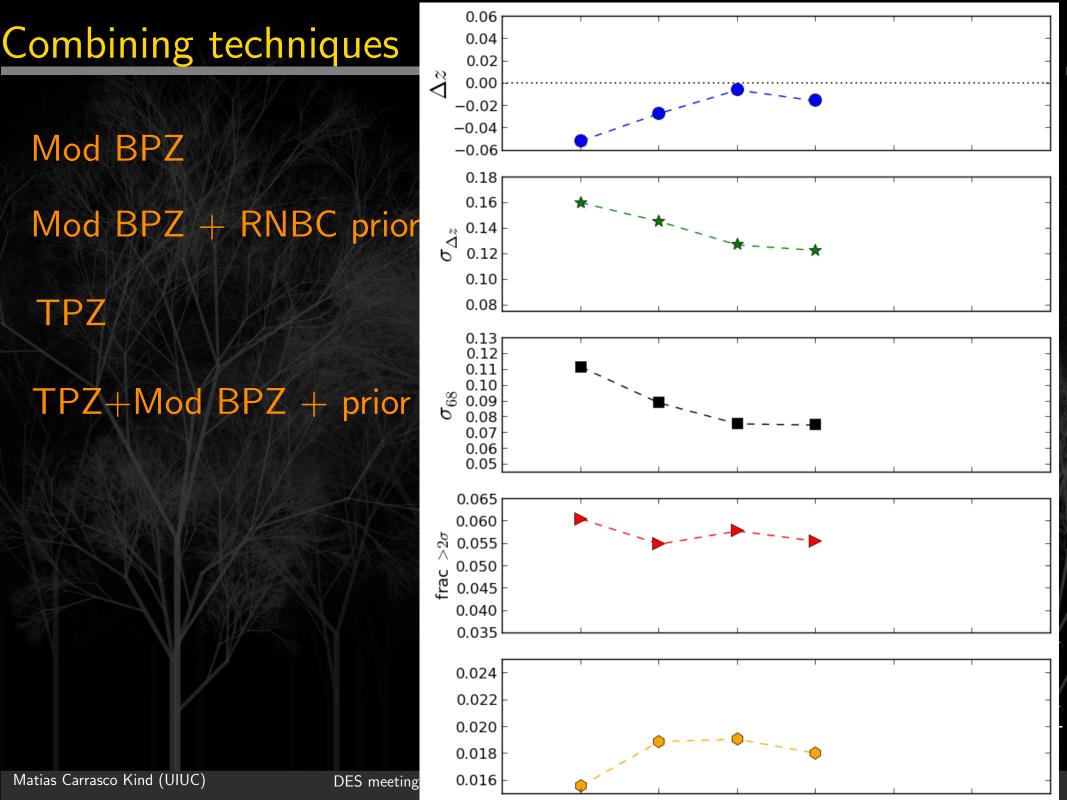
(Likelihood)

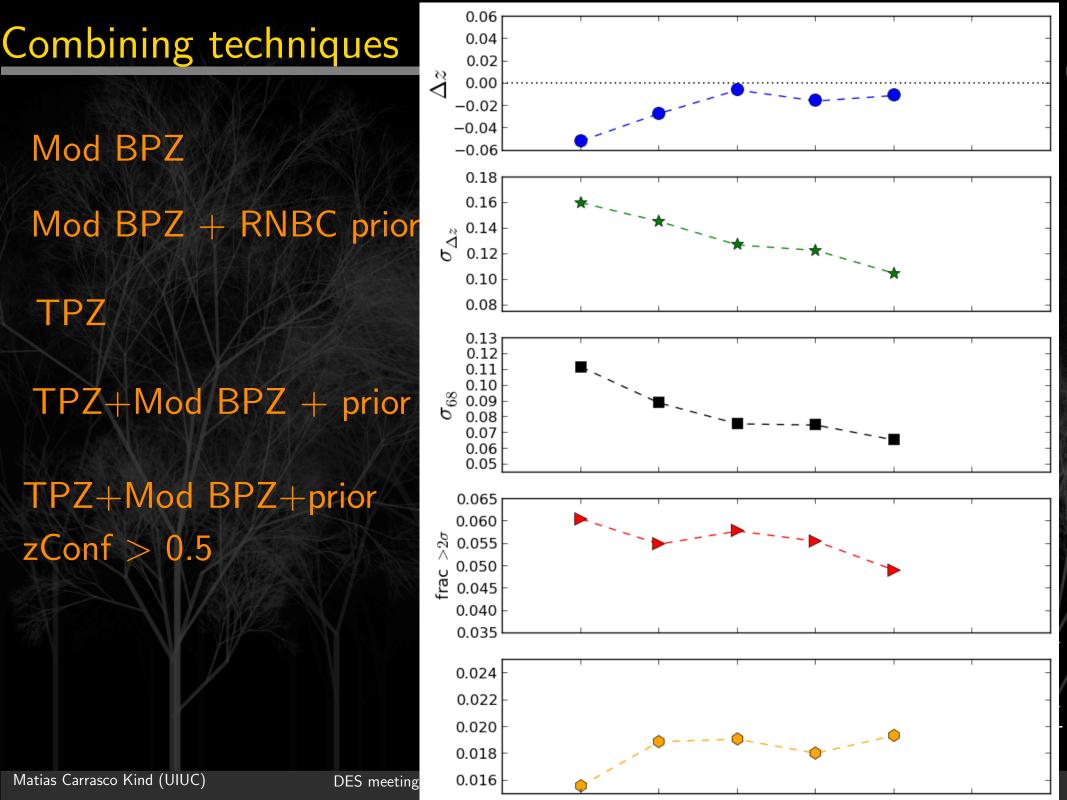
 $P(m_i|z,T)$ 

#### 0.06 Combining techniques 0.04 0.02 0.00 -0.02-0.04Mod BPZ -0.060.18 0.16 0.14 0.12 0.10 0.08 0.13 0.12 0.11 0.10 0.09 0.08 0.07 0.06 0.05 0.065 0.060 $\frac{6}{5}$ 0.055 0.050 0.050 0.045 يا 0.040 0.035 0.024 0.022 0.020 0.018 Matias Carrasco Kind (UIUC) 0.016 DES meeting









#### Combining techniques

Mod BPZ

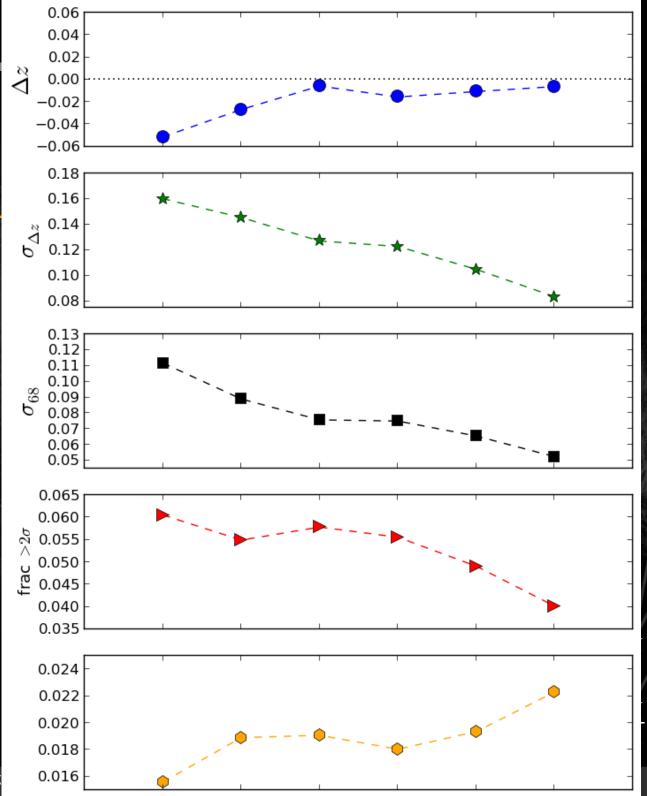
Mod BPZ + RNBC prior

TPZ

TPZ+Mod BPZ + prior

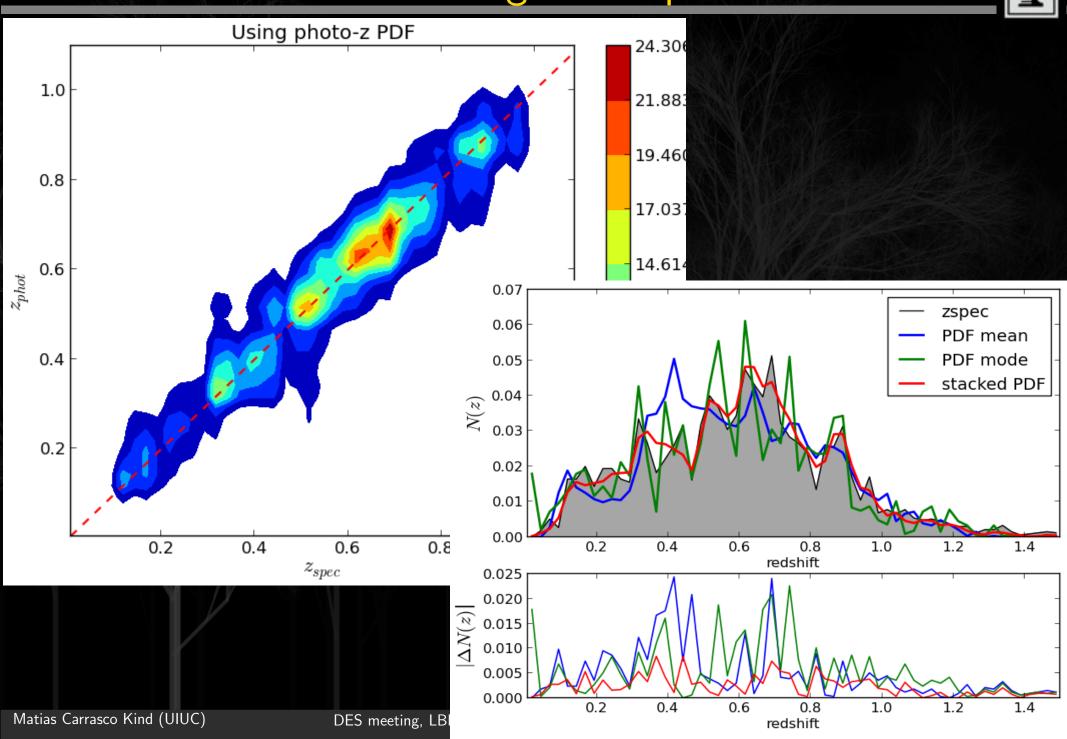
TPZ+Mod BPZ+prior zConf > 0.5

TPZ+Mod BPZ+prior zConf > 0.9



#### Combining techniques





#### Simulated DES



We use TPZ (Carrasco Kind & Brunner, 2013a) to generate photo-z for all galaxies.

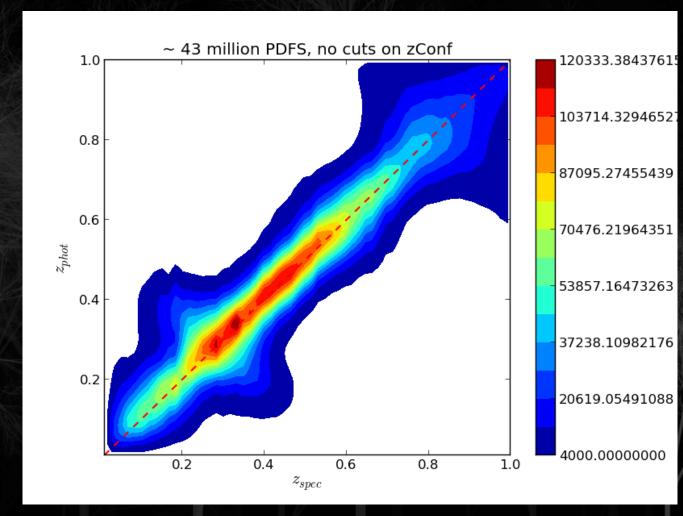
100,00 for training

5 magnitudes only

 $\sim 0.17$  sec per PDF

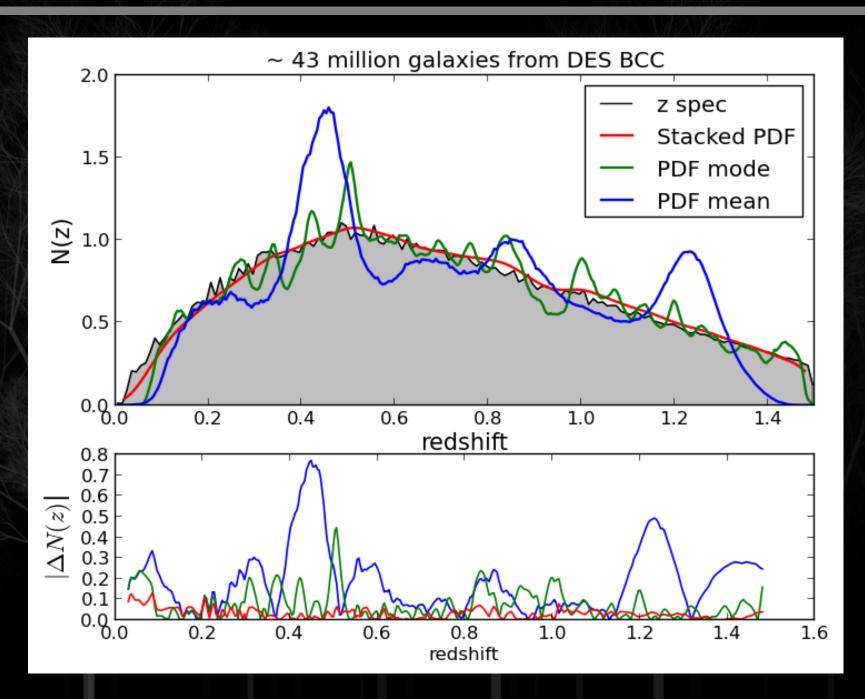
Store 43 million PDFs for analysis

No outlier removal



#### Simulated DES





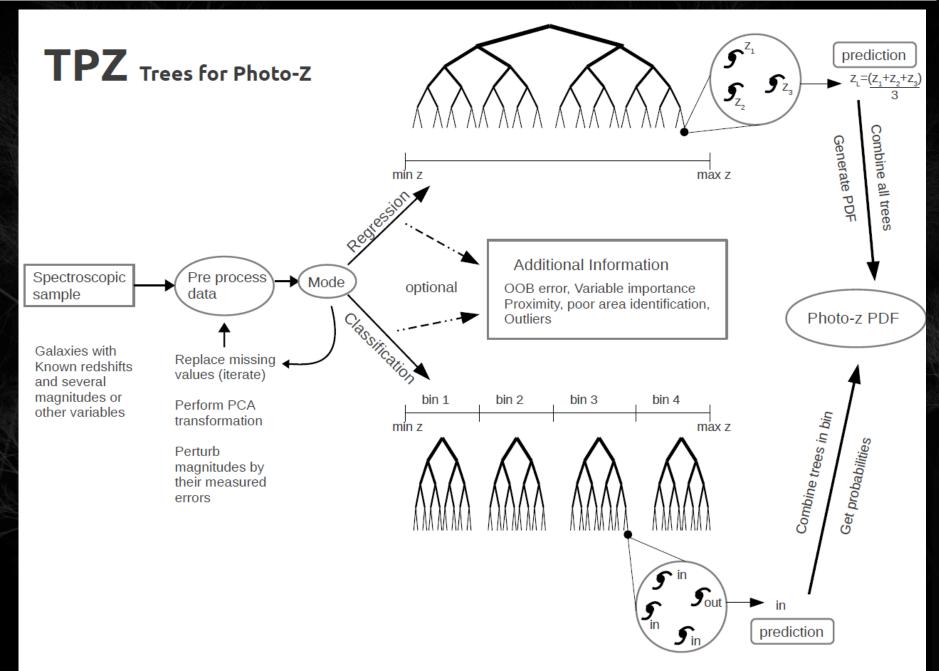
## Thanks!





#### TPZ scheme





#### Performance tests



