

Cosmic muons model: Notes

A) Bethe Bloch for protons in water (from source code) (bethebloch.py)

Bethe Bloch formula:

$$\left[-\frac{dE}{dx} = \frac{4\pi}{m_e c^2} \cdot \frac{n z^2}{\beta^2} \cdot \left(\frac{e^2}{4\pi E_0} \right)^2 \left[\ln \left(\frac{2 m_e c^2 \beta^2}{I \cdot (1 - \beta^2)} \right) - \beta^2 \right] \right]$$

Where:

- $\beta = \frac{v}{c}$
- v : speed of particle
- E : energy of particle
- x : length of path
- n : numeric density of electrons in material
- I : mean excitation potential of material
- $z e$: charge of particle
- e : elementary charge
- c : speed of light
- m_e : mass of electron

Note:

- Block approx: $[I = 10 \text{ eV} \cdot Z]$, Z atomic number of the medium.
- Electron number density: $n = \frac{N_A \cdot Z \cdot \rho}{A \cdot M_u \approx 1}$

With

- ρ density
- A mass number
- N_A avogadro number

(TEST) Electron density of water

We know $n_e = 3,34 \cdot 10^{29} \text{ m}^{-3}$

- density of water: $997 \text{ kg/m}^3 = 997 \cdot 10^3 \text{ g/m}^3$

- avogadro number: $6,02 \cdot 10^{23} \text{ mol}^{-1}$

- $Z \approx 7,42$ (effective)

- $A = 18 \text{ g/mol}$

$$\Rightarrow n = \frac{N_A \cdot Z \cdot \rho}{A} = \frac{6,02 \cdot 10^{23} \text{ mol}^{-1} \cdot 7,42 \cdot 997 \cdot 10^3 \text{ g/m}^3}{18 \text{ g/mol}}$$

$$= 2,47 \cdot 10^{25}$$

*Not exact but
good magnitude*

✓ TESTED

Plain Idea of the code

Start with a particle at a given E_{kin} at $x=0$

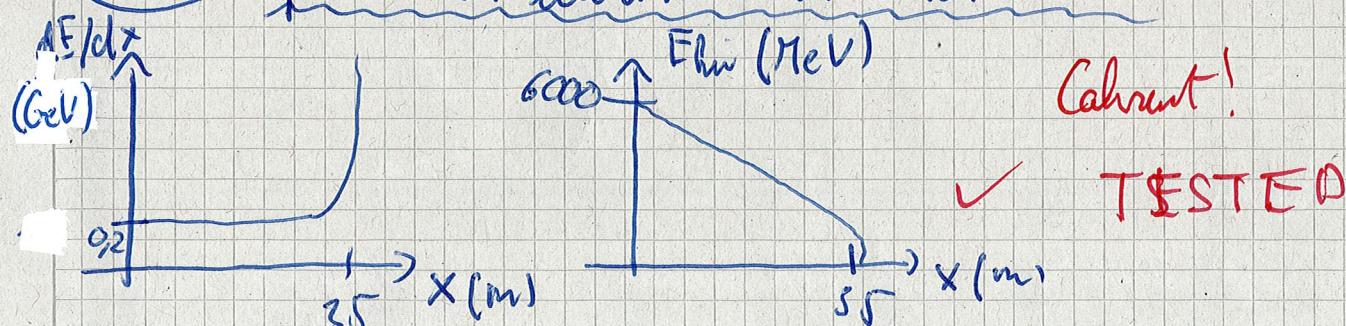
- Compute dE/dx

- $E_{\text{kin}} = E_{\text{kin}} - dE$, $x = x + dx$

loop until $E_{\text{kin}} < 0$

(TEST)

proton in water with $E_{\text{kin}} = 6 \text{ GeV}$



(bethebloch-muons.py)

B) Bethe-Bloch for muons in Air (adiabatic atmosphere)

TEST*

Convergence of the previous method for muon in air (sea level)

For air:

- effective $Z: 7,6$
- $A: 28,97 \text{ g/mol}$
- $\rho \text{ at sea level at } 0^\circ\text{C} \quad 1,060 \text{ kg/m}^3$

For muon:

- $Z_1 = 1$
- Mass $M_0 = 207 \cdot m_e$

Initial $E_{hi}: 2 \text{ GeV}$

Results: Distance for $E_{hi} \leq 0: \underline{\sim 17'000 \text{ m}}$

TESTED

TEST **

Muons with $E_{hi} 6 \text{ GeV}$ after 15'000 m

The energy of muons with initial $E_{hi} = 6 \text{ GeV}$ has a residual energy of 2 GeV after 15'000 m of interaction.

TESTED

Cohesive with cosmic. lab. gov

C) Insertion of the adiabatic atmosphere model

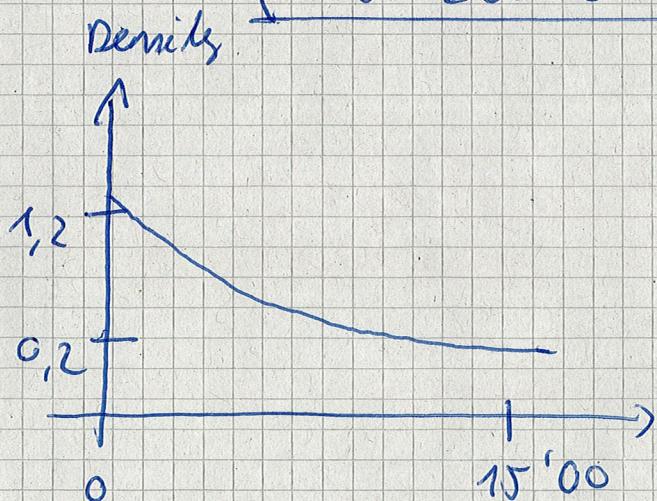
Variation of density of air with altitude:

$$\boxed{\rho = \frac{P_0 M}{R T_0} \left(1 - \frac{C h}{T_0}\right)^{g M / R L - 1}}$$

- M: molar mass of air : 0,029 kg/mol
- R: ideal gas constat 8,314 J/(mol·K)
- P₀: sea level pressure : 101325 Pa
- T₀: sea level temperature : 288,15 K
- C: temperature layer rate : 0,0065 K/m
- h: altitude
- g = 9,81 m/s²

(TEST)

Variation of density with altitude
(0 to 20000 m)



OK, coherent with
cosmic.ell.gau



TESTED

Intention in the code we already have in the following script: `bethebloch-muons-adia.py`

 here, we suppose that the creation altitude of muons is about 15 km. So the "0" distance of the previous code is supposed to be at this creation altitude

(TEST) Scheme as **, but we have an adia.
atmosphere

Energy after 15'000 m 4,789 GeV

OK, but seems very high, but the model is more accurate!

TESTED

(TEST) Scheme as **, but with an adia.
atmosphere

muon take appose 20'000 m to stop

Much bigger than the expectations...

D) Computation of the mass. path TESTED

length for energies between 0,2 GeV and 100 GeV

Range : 0,2 GeV (minimum on the graph) minimum
100 GeV maximum (after, flux $< 10^{-4}$)

