

A geometric origin of the Bohm potential from Einstein–Cartan spin–torsion coupling

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Academic Editors: M. K. Tavassoly, Hichem Eleuch

Abstract

A deterministic quantum-gravity framework is proposed that embeds Bohmian mechanics within the spin–torsion geometry of Einstein–Cartan–Kibble–Sciama theory. Starting from the Palatini form of the action for a Dirac field, a Gordon decomposition followed by a Foldy–Wouthuysen expansion shows that the familiar Bohm quantum potential, namely the term proportional to the Laplacian of the wave-function amplitude divided by that amplitude, coincides exactly with the axial-torsion contribution to the Ricci scalar. A covariant, spin-augmented guidance law then follows, and the algebraically determined torsion field, sourced by the global spin density of an entangled state, mediates non-local correlations without superluminal signalling. The same spin–torsion couplings produce an effective dark stress–energy tensor, for which observational tests are outlined across a wide range of scales: neutron-star magnetospheres, black-hole jets, gravitational-wave inspirals and spin-polarized interferometry. This geometric mechanism extends to spinless particles via torsion’s universal back-reaction on the connection, yielding a deterministic reformulation of quantum mechanics without metric quantization. A simple one-dimensional toy model reproduces the Tsirelson bound for Bell-inequality violations, underscoring the empirical and conceptual viability of this geometric route to quantum gravity, which dispenses with the need to quantize the metric.

Keywords: *quantum gravity, bohmian mechanics, Einstein–Cartan theory, spin-torsion coupling, non-local correlations*

Citation: Northey JGB. A geometric origin of the Bohm potential from Einstein–Cartan spin–torsion coupling. *Academia Quantum* 2025;2. <https://doi.org/10.20935/AcadQuant7901>

1. Introduction

Why has gravity eluded quantization strategies that succeeded for other forces? General relativity (GR) paints gravity as a smooth spacetime curvature sourced by energy–momentum [1], while quantum mechanics (QM) insists on fundamentally probabilistic wavefunctions and measurement outcomes [2]. Decades of effort—from Weyl’s gauge idea [3] (for the original 1918 paper see [4]) and Einstein’s affine-connection programme [5] (historical background in [6]) to string theory’s vibrating loops [7] (classic monograph: [8]) and loop–quantum–gravity spin networks [9]—have yet to deliver a fully satisfactory bridge between these paradigms or a clear experimental signature [10]. Recent spinfoam overviews, however, underscore torsion’s central role in background-independent quantization schemes [9, 11].

An intriguing alternative is to geometrize quantum effects themselves. In the Einstein–Cartan–Kibble–Sciama (ECKS) theory, spacetime torsion, an antisymmetric part of the affine connection, is sourced algebraically by the intrinsic spin of matter [12, 13]. Torsion is invisible in classical solar-system tests but can soften singularities [14] and introduce novel spin–gravity couplings [15]. Meanwhile, Bohmian mechanics restores determinism via a non-local quantum potential,

$$Q = -\frac{\hbar^2}{2m} \frac{\nabla^2 R}{R}, \quad (1)$$

guiding particles along definite trajectories [16, 17]. To date, however, no framework has shown that this Bohm potential can

emerge directly from a spacetime geometry that respects Lorentz covariance and yields a fully covariant guidance law.

This work closes that gap. Building on earlier hints [18, 19], the analysis

(i) begins with the Palatini–ECKS action minimally coupled to a Dirac spinor, within the fully covariant Dirac framework validated by exact torsionful solutions [20],

(ii) performs a Gordon decomposition and Foldy–Wouthuysen expansion to prove

$$Q = -\frac{\hbar^2}{2m} \frac{\nabla^2 R}{R} \iff Q = -\frac{\hbar^2}{8m} (\nabla_\mu S^\mu + S_\mu S^\mu), \quad (2)$$

identifying the Bohm potential with the axial torsion vector S^μ ,

(iii) derives a manifestly covariant, spin-augmented guidance law that preserves covariance under foliation changes,

(iv) shows that algebraic ECKS torsion—sourced by the global spin density $|\psi(x_1, x_2)|^2$ —automatically mediates quantum non-locality and reproduces Bell-inequality violations without superluminal signalling,

(v) constructs an effective “dark” stress–energy tensor from spin–torsion couplings and outline concrete observational

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tests in neutron-star X-ray polarimetry [21], black-hole jet magnetometry, gravitational-wave inspirals [22], and spin-polarized interferometry [23].

This synthesis resolves the tension between quantum non-locality and spacetime geometry, offering a deterministic alternative to metric quantization, where the H-J equation (with torsion-sourced potential) + continuity equation equals the Schrödinger equation, extending even to spinless particles via geometric back-reaction. Further developments, including cosmological implications of torsion as dark energy and a teleparallel scalar–torsion extension [24] and a unimodular-teleparallel variant [25], are detailed in a companion paper [19].

This analysis yields a wealth of testable predictions. As a striking example, estimates suggest that the Imaging X-ray Polarimetry Explorer (IXPE) could constrain torsion couplings at the 10^{-2} level by measuring polarization anomalies in neutron-star magnetospheres.

The remainder of this paper is organized as follows. In Section 2.2 the spin–phase relation is derived from the Dirac–ECKS action, and in Section 2.3 the nonrelativistic Foldy–Wouthuysen expansion to recover the Bohm equations are performed. Section 2.5 presents the covariant guidance law and its foliation consistency, while Section 3.1 shows how algebraic torsion encodes non-local entanglement. In Section 3.2 the dark-sector stress–energy is discussed and in Section 4 observational and laboratory-scale signatures are explored. The conclusion in Section 5 outlines future extensions.

2. Materials and methods

2.1. Conventions

The metric signature is chosen as $(-, +, +, +)$, with the Levi-Civita symbol $\epsilon^{\mu\nu\rho\sigma}$ defined by $\epsilon^{0123} = +1$. The hypersurface normal n_μ aligns with the quantum fluid’s rest frame.

2.2. Dirac–ECKS action and Gordon decomposition

Starting with the Palatini–ECKS action for Dirac fields:

$$S = \int d^4x \sqrt{-g} \left[\frac{1}{16\pi G} (\mathcal{R} - 2\Lambda) + i\bar{\psi}\gamma^\mu D_\mu \psi - m\bar{\psi}\psi \right], \quad (3)$$

where $D_\mu = \nabla_\mu + \frac{i}{4} K_{\rho\sigma\mu} \gamma^{[\rho} \gamma^{\sigma]}$. The Gordon decomposition of the current $j^\mu = \bar{\psi}\gamma^\mu \psi$ yields:

$$j^\mu = \frac{\hbar}{2m} (\bar{\psi} \nabla^\mu \psi - (\nabla^\mu \bar{\psi}) \psi) + \frac{\hbar}{2m} \epsilon^{\mu\nu\rho\sigma} \nabla_\nu (\bar{\psi} \gamma_\rho \gamma_\sigma \psi). \quad (4)$$

Identifying $\nabla^\mu \ln R$ with the spin–torsion coupling term gives the fundamental relation:

$$\frac{\nabla^\mu R}{R} = \frac{m}{\hbar} \epsilon^{\mu\nu\rho\sigma} s_\nu T_{\rho\sigma}, \quad (5)$$

derived without ansatz. This identity rewrites the amplitude gradient entirely in terms of the axial torsion vector and will be the key step in eliminating ψ from the quantum potential below. Although the Palatini–ECKS action is introduced in metric–connection variables, it can be derived equally well by gauging the Lorentz

group $SO(1, 3)$, or more precisely its double cover $\text{Spin}(1, 3) \simeq \text{SL}(2, \mathbb{C})$ [12, 26, 27]. In this language, the co-frame $e^a{}_\mu$ plays the role of the gauge potential for spacetime translations, while the spin connection $\omega^{ab}{}_\mu$ is the Lorentz gauge field; their field strengths are torsion T^a and curvature R^{ab} , respectively. The two-to-one homomorphism $\text{Spin}(1, 3) \rightarrow \text{SO}(1, 3)$ mirrors the familiar $\text{SU}(2) \rightarrow \text{SO}(3)$ covering that underlies the electroweak sector of the standard model, so the present construction may be viewed as a classical geometric gauge theory of Weyl type. In the Bohmian interpretation, the dynamically selected orthonormal frames—or equivalently the foliation defined by $n_\mu = j_\mu / \sqrt{-j_\alpha j^\alpha}$ —supply the additional “hidden variables” envisaged by Bohm, completing the quantum description without contradicting standard predictions.

2.2.1. Bohm potential from axial torsion

Separating real and imaginary parts of the Dirac–ECKS equation and using **Equation (5)** one finds that only the axial combination $\nabla_\mu S^\mu + S_\mu S^\mu$ enters the non-relativistic Hamiltonian. We therefore define

$$\Delta \mathcal{R}_{\text{torsion}} \equiv -2(\nabla_\mu S^\mu + S_\mu S^\mu) [L^{-2}], \quad (6)$$

the torsion-only contribution to the Ricci scalar. We shall refer to $\Delta \mathcal{R}_{\text{torsion}}$ as “the curvature built purely from torsion”; it carries the correct inverse-length-squared dimension. With this shorthand the Foldy–Wouthuysen expansion gives the central identification

$$Q_{\text{tors}} = -\frac{\hbar^2}{8m} \Delta \mathcal{R}_{\text{torsion}} [ML^2 T^{-2}], \quad (7)$$

exactly matching Bohm’s $Q = -\hbar^2 \nabla^2 R / 2mR$ in both magnitude and units. **Equation (9)** is the central result: the Bohm quantum potential is geometrized as the axial-torsion contribution to curvature, scaled by $\hbar^2/8m$.

This identity arises because of the following:

- (i) Torsion’s irreducible decomposition (Section 2.4) isolates the axial vector S^μ as the sole low-energy contributor;
- (ii) The FW expansion maps the Dirac–ECKS Hamiltonian to a Schrödinger form where $\nabla_\mu S^\mu + S_\mu S^\mu$ captures amplitude-driven curvature;
- (iii) Gordon decomposition **Equation (5)** directly links $\nabla^\mu \ln R$ to S^μ , embedding the quantum potential in spacetime geometry.

It serves as the master relation for all subsequent sections.

2.3. Nonrelativistic limit of the Dirac–ECKS system

The Dirac equation with the contorted connection, expanded to leading order in $1/c$, becomes

$$i\hbar \partial_t \psi = \left[\frac{((\sigma \cdot \pi)^2)}{2m} + Q_{\text{tors}} + V \right] \psi, \quad (8)$$

where $\pi = (p - \frac{\hbar}{2} S)$, with S denoting the spatial components of the axial torsion 4-vector S^μ . The covariant divergence $\nabla_\mu S^\mu$ reduces to $\nabla \cdot S$ in the nonrelativistic limit, while $S_\mu S^\mu \approx |S|^2$. The torsion-induced quantum potential is given by the central result (**Equation (9)**):

$$Q_{\text{tors}} = -\frac{\hbar^2}{8m} (\nabla_\mu S^\mu + S_\mu S^\mu). \quad (9)$$

In the nonrelativistic regime, this simplifies to

$$Q_{\text{tors}} \approx -\frac{\hbar^2}{8m} (\nabla \cdot S + |S|^2),$$

recovering the standard Bohm potential form when combined with the spin-phase relation (**Equation (11)**). Here $V(x)$ is a classical external potential (e.g., electromagnetic or gravitational) that commutes with torsion operators. All intrinsic spin–torsion results assume $V = 0$; external potentials can be added without modifying Q_{tors} .

Matching to Bohm’s decomposition $\psi = R e^{iS/\hbar}$ yields:

$$\partial_t R^2 + \nabla \cdot \left(R^2 \frac{\nabla S}{m} \right) = 0, \quad (\text{continuity}) \quad (10)$$

$$\partial_t S + \frac{|\nabla S|^2}{2m} + Q_{\text{tors}} + V = 0, \quad (\text{Hamilton–Jacobi}) \quad (11)$$

establishing consistency with Bohmian mechanics at $\mathcal{O}(1/c^2)$.

2.4. Torsion irreducibles and spin coupling

The torsion tensor $T_{\mu\nu}^\rho$ decomposes into three irreducible parts under the Lorentz group:

$$T_{\mu\nu}^\rho = \frac{1}{3} (\delta_\mu^\rho T_\nu - \delta_\nu^\rho T_\mu) + \frac{1}{6} \epsilon_{\mu\nu\sigma}^\rho S^\sigma + q_{\mu\nu}^\rho, \quad (12)$$

where:

- $T_\mu \equiv T_{\mu\nu}^\nu$ is the trace torsion vector,
- $S^\mu \equiv \epsilon^{\mu\nu\rho\sigma} T_{\nu\rho\sigma}$ is the axial torsion vector, sourced by the spin density $\sigma^{\rho\mu\nu} = \bar{\psi} \gamma^{[\rho} \gamma^\mu \gamma^{\nu]} \psi$ of the Dirac field [3, 5, 28],
- $q_{\mu\nu}^\rho$ is the traceless, parity-even pure tensor torsion.

Only the axial torsion S^μ couples to the quantum potential. This coupling arises via the spin-phase gradient:

$$S^\mu = \frac{\hbar}{m} \nabla^\mu \ln R, \quad (13)$$

where $R = |\psi|$ is the wavefunction amplitude. **Equation (13)** directly links spacetime torsion to the Bohmian guiding equation. Hence the Bohm guidance equation can be viewed as a geodesic equation in a spacetime whose torsion is fixed by $\ln R$.

2.4.1. Tensor torsion decoupling

The pure tensor component $q_{\mu\nu}^\rho$ contributes to the Ricci scalar at $\mathcal{O}(T^2)$:

$$\Delta \mathcal{R} \supset \frac{1}{4} q_{\rho\mu\nu} q^{\rho\mu\nu} - \frac{1}{2} q_{\rho\mu\nu} q^{\mu\nu\rho}. \quad (14)$$

However, for Dirac fermions, Fierz identities—which enforce antisymmetry in spinor bilinears—constrain the spin density $\sigma^{\rho\mu\nu}$ to source only S^μ and T^μ . Consequently, $q_{\mu\nu}^\rho$ vanishes identically in the nonrelativistic limit. Parity-odd terms cancel due to the \mathcal{P} -symmetry of the axial coupling $S^\mu \nabla_\mu \ln R$, ensuring consistency with observed parity conservation in low-energy quantum systems.

2.4.2. Trace torsion decoupling

For Dirac fermions, Fierz identities (See, e.g., [29] for a derivation of spinor bilinear constraints.) enforce antisymmetry in $\sigma^{\rho\mu\nu}$,

suppressing the trace torsion T_μ . Specifically, the spin density satisfies $\sigma^{[\rho\mu\nu]} = 0$, eliminating T_μ in the nonrelativistic limit. Recent studies on torsional effects in gravity–spinor couplings further support this via Yukawa-like interactions [30].

2.4.3. Dimensional consistency of the quantum potential

The quantum potential $Q_{\text{tors}} = -\frac{\hbar^2}{8m} (\nabla \cdot S + S^2)$ maintains correct energy dimensions ($[ML^2T^{-2}]$). This is verified as follows:

- The axial torsion vector S has dimensions of inverse length, $[L^{-1}]$, as it is derived from the torsion tensor $T_{\nu\rho\sigma}$ (which has $[L^{-1}]$ in geometrized units).
- Therefore, $\nabla \cdot S$ (divergence) has dimensions $[L^{-2}]$, and $S^2 = S_\mu S^\mu$ also has $[L^{-2}]$.
- The constant $\frac{\hbar^2}{m}$ has dimensions $[ML^4T^{-2}]$ (since \hbar has $[ML^2T^{-1}]$ and m has $[M]$).
- Thus, Q_{tors} has dimensions $[ML^4T^{-2}] \times [L^{-2}] = [ML^2T^{-2}]$, matching the Bohm quantum potential $Q = -\frac{\hbar^2}{2m} \frac{\nabla^2 R}{R}$ (which has $[ML^2T^{-2}]$ because $\nabla^2 R/R$ has $[L^{-2}]$).

This resolves the apparent discrepancy in the Ricci scalar formulation and confirms the consistency of the torsion-derived quantum potential. The earlier notation $\Delta \mathcal{R}_{\text{torsion}}$ corresponds to $-2(\nabla \cdot S + S^2)$, preserving the relation $Q = -\frac{\hbar^2}{8m} \Delta \mathcal{R}_{\text{torsion}}$ with correct dimensions.

2.5. Covariant guidance law

To polar-decompose the Dirac spinor in the presence of torsion,

$$\psi = R e^{\frac{i}{\hbar} S_+} e^{\frac{i}{2} S_- \gamma_5} \chi,$$

where $R^2 = \bar{\psi} \psi$ and the real phases S_\pm are the symmetric (+) and antisymmetric (−) combinations that appear in the Gordon decomposition. The conserved Dirac current and the axial (“spin”) current are

$$j^\mu = \bar{\psi} \gamma^\mu \psi, \quad s^\mu = \bar{\psi} \gamma^\mu \gamma_5 \psi.$$

From the Gordon identity one finds

$$j^\mu = \frac{R^2}{m} (\cos S_- \nabla^\mu S_+ - \sin S_- \epsilon^{\mu\nu\rho\sigma} \hat{s}_\nu \nabla_\rho S_-), \quad \hat{s}_\mu \equiv \frac{s_\mu}{\sqrt{-j_\alpha j^\alpha}},$$

where \hat{s}_μ is the *unit* axial vector (dimensionless), ensuring proper mass–dimension balance. Identifying u^μ via $j^\mu = R^2 \cos S_- u^\mu$ and projecting onto an arbitrary spacetime slicing with unit normal n^μ yields the manifestly covariant Bohm–ECKS guidance equation

$$m \cos S_- u^\mu = -\frac{1}{2} (\nabla^\mu S_+ + \epsilon^{\mu\nu\rho\sigma} n_\nu \hat{s}_\sigma \nabla_\rho S_-) + \frac{\hbar}{2} \epsilon^{\mu\nu\rho\sigma} \hat{s}_\nu T_{\rho\sigma}{}^\alpha n_\alpha. \quad (15)$$

Notation (See Appendix A.2, **Equation (A8)**, for the Gordon/FW derivation that yields the normalized spin orientation \hat{s}_μ .) Throughout, $\hat{s}^\mu = s^\mu / \sqrt{-j_\alpha j^\alpha}$ denotes the unit spin orientation; it carries no density dimension. The last (torsion) term enters with a positive coefficient $+\hbar/2$, as dictated by the sign of the contorsion contribution to the covariant derivative and responsible for the repulsive quantum-potential correction discussed in Section 3.

Key features:

- Index balance: Every free index on the RHS is either μ or contracted; hence (15) is a bona-fide four-vector equation.
- Dimensions: $\nabla^\mu S_\pm$ and $T_{\rho\sigma}{}^\alpha$ carry mass dimension 1; all other factors are dimensionless, matching mu^μ on the LHS.
- Parity: \hat{s}_σ is axial (odd), the Levi-Civita tensor is pseudo (odd), so their product is even; no unintended parity violation is introduced.
- Limits: If $S_- = 0$ (no chiral phase) the middle term vanishes; if the spin lies in the hypersurface ($\hat{s} \cdot n = 0$) the torsion term vanishes, reproducing the expected special cases.

2.5.1. Dynamical foliation

The normal n^μ is not imposed by hand; it is locked to the conserved Dirac current:

$$n_\mu = \frac{j_\mu}{\sqrt{-j_\alpha j^\alpha}}, \quad j^\mu = \bar{\psi} \gamma^\mu \psi.$$

Because j^μ is a genuine four-vector variable, the identification $n^\mu = j^\mu / \sqrt{-j_\alpha j^\alpha}$ preserves diffeomorphism invariance: the hypersurface normal is built from the state itself and transforms covariantly under any coordinate change. This alignment ensures torsion-mediated effects remain frame-dependent yet respect relativistic causality. Kinematically the choice enforces $\mathcal{L}_n u^\mu = 0$, i.e., Bohmian trajectories do not drift when the spacetime slicing is deformed along n^μ . Dynamically it guarantees that information encoded in the algebraic torsion $T_{\rho\sigma\alpha}$ propagates only through the wavefunction ψ , preventing any superluminal signalling despite the manifest non-locality of the quantum potential, since ψ evolves causally under the Dirac-ECKS equation. Here $s^\mu = \bar{\psi} \gamma^\mu \gamma_5 \psi$ is the axial (spin) current, whereas the dimensionless spin orientation that appears in **Equation (15)** is $\hat{s}^\mu = s^\mu / \sqrt{-j_\alpha j^\alpha}$.

2.6. Spinless matter in a torsionful spacetime

In Einstein-Cartan theory the axial torsion vector $S^\mu = \epsilon^{\mu\nu\rho\sigma} T_{\nu\rho\sigma}$ is sourced algebraically by the fermionic spin density $\sigma^{\mu\nu\rho} \propto \bar{\psi} \gamma^{[\mu} \gamma^\nu \gamma^{\rho]} \psi$. Hence a purely scalar or electromagnetic sector ($\sigma^{\mu\nu\rho} \equiv 0$) does not generate torsion: $S^\mu = 0$ locally. Nevertheless spinless quanta do propagate through, and are guided by, any torsion field that is set up by other matter, e.g., nearby fermions. The mechanism works in three linked steps.

2.6.1. Gravitational coupling

The field equations with both scalar (ϕ) and fermion (ψ) content read

$$G_{\mu\nu} + \Lambda g_{\mu\nu} = 8\pi G \left(T_{\mu\nu}^{(\phi)} + T_{\mu\nu}^{(\psi)} + \kappa \sigma_{\alpha\beta\gamma} T^{\alpha\beta\gamma} g_{\mu\nu} \right),$$

so the contorted connection $\Gamma^\rho{}_{\mu\nu} = \{\rho{}_{\mu\nu}\} + K^\rho{}_{\mu\nu}$ governs all geodesics. A spinless particle's classical path therefore bends once torsion is present elsewhere in spacetime.

2.6.2. Quantum propagation of a scalar field

For a complex Klein-Gordon field in a Riemann-Cartan spacetime one has

$$(\square + \xi \mathcal{R} - m^2 c^2 / \hbar^2) \phi = 0, \quad (16)$$

where the d'Alembertian $\square \equiv g^{\mu\nu} \nabla_\mu \nabla_\nu$ contains the full (contorted) connection, and the curvature splits as $\mathcal{R} = \mathcal{R}_{\text{LC}} + \mathcal{R}_{\text{tors}}$. Thus, a scalar field feels any background torsion generated by nearby fermions. This establishes a unified geometric mechanism (as derived for Dirac fermions in Section 2.3): both fermionic and spinless quantum potentials arise from $\nabla_\mu S^\mu + S_\mu S^\mu$, with torsion sourced by fermionic matter but acting universally through the contorted ∇_μ . For minimal coupling ($\xi = 0$), the second term in **Equation (17)** survives purely from torsion in \square ; non-minimal ($\xi \neq 0$) rescales it via explicit $\xi \mathcal{R}_{\text{tors}}$, but the structure remains geometric. Writing $\phi = R e^{is/\hbar}$ and separating real and imaginary parts yields the continuity equation together with a Hamilton-Jacobi equation,

$$\partial_t S + \frac{(\nabla S)^2}{2m} + V + Q_\phi = 0,$$

whose quantum potential reads

$$Q_\phi = -\frac{\hbar^2}{2m} \frac{\nabla^2 R}{R} - \frac{\hbar^2}{8m} [\nabla_\mu S^\mu + S_\mu S^\mu] (1 - 6\xi). \quad (17)$$

Minimal coupling ($\xi = 0$). Even with $\xi = 0$ the second bracket survives because the contorted ∇_μ in **Equation (16)** already carries torsion through S^μ . Hence a scalar field feels any background torsion generated by nearby fermions. This establishes a unified geometric mechanism: both fermionic and spinless quantum potentials arise from $\nabla_\mu S^\mu + S_\mu S^\mu$, with torsion sourced by fermionic matter. Non-minimal coupling ($\xi \neq 0$). Choosing $\xi \neq 0$ adds the explicit $\xi \mathcal{R}_{\text{tors}}$ term familiar from curved-space scalar theories and merely rescales the same geometric combination $\nabla_\mu S^\mu + S_\mu S^\mu$. Therefore, if $S^\mu \neq 0$ in the environment, Q_ϕ inherits exactly the torsion structure that drives the fermionic Bohm potential, guaranteeing a unified guidance mechanism for spinful and spinless sectors alike.

2.6.3. Mixed (fermion + scalar) entangled systems

Consider a bipartite wavefunction $\Psi(x_f, x_s)$ describing a fermion “f” and a scalar “s”. The fermionic factor sources torsion, which in turn enters the scalar's guidance law

$$\frac{dx_s}{dt} = \frac{\nabla S_s}{m_s} - \frac{\hbar}{8m_s} \nabla(\nabla \cdot S + S^2),$$

so non-local Bohmian correlations, and thus Bell-type violations, are preserved even though the scalar carries no intrinsic spin.

In summary, the theory minimally extends to spinless matter: torsion originates only from fermionic spin but its geometric back-reaction acts universally on every field through the modified connection and curvature. Regions with $\langle \mathcal{R}_{\text{tors}} \rangle \neq 0$ therefore influence neutral particles or dark-matter candidates exactly as required for a fully consistent, deterministic quantum dynamics.

2.7. Dynamical back-reaction

The self-consistent iteration scheme:

1. Solve ECKS equations for $\Gamma^\rho{}_{\mu\nu}$ given initial ψ
2. Compute Q via revised potential:

$$Q = -\frac{\hbar^2}{8m} \Delta \mathcal{R} + \frac{\hbar^2}{4m} K^{\rho\sigma\mu} K_{\rho\sigma\mu} \quad (18)$$

3. Update trajectories using the covariant guidance law (**Equation (15)**)
4. Recalculate $\sigma^{\rho\mu\nu}$ and repeat

2.7.1. Algebraic non-locality

Although torsion is algebraically determined by $\sigma^{\rho\mu\nu}(x)$, the spin density itself depends on the *global* wavefunction $\psi(x_1, x_2)$. Hence $Q(x)$ becomes a non-local functional via:

$$Q[\psi] \sim \int d^3x' \frac{\delta Q}{\delta \psi(x')} \psi(x'), \quad (19)$$

where the variational derivative $\delta Q/\delta \psi$ includes contributions from all entangled particles.

The contorsion $K_{\rho\sigma\mu}$ modifies the affine connection, while the torsion propagator $\mathcal{G}(r)$ (renamed from $K(r)$) governs non-local interactions:

$$T_{\rho\sigma}(x) = 8\pi G \int d^4x' \mathcal{G}(x - x') \sigma_{\rho\sigma}(x'). \quad (20)$$

3. Results

3.1. Non-locality and entanglement mediation

For two entangled particles with configuration–space wavefunction $\Psi(x_1, x_2)$ the axial torsion vector is sourced by the global spin density

$$S^\mu(x) = \frac{8\pi G \hbar}{m} |\Psi(x_1, x_2)|^2 [\delta^3(x - x_1) + \delta^3(x - x_2)].$$

The corresponding Bohm–torsion quantum potential is therefore

$$Q(x_1, x_2) = -\frac{\hbar^2}{8m} (\nabla_\mu S^\mu + S_\mu S^\mu),$$

a manifestly non-local functional that nevertheless carries the correct energy dimension $[ML^2T^{-2}]$. Because S^μ itself is a sum of two delta functions, Q couples the two particles instantaneously and cannot be decomposed into separate single-particle terms. Causality is preserved by the preferred foliation $n_\mu = j_\mu / \sqrt{-j_\alpha j^\alpha}$ aligned with the conserved Dirac current $j^\mu = \bar{\psi} \gamma^\mu \psi$, so no superluminal signalling can occur even though the guidance equation is non-local. The seeming “spooky action at a distance” dissolves once we recall that torsion in Einstein–Cartan theory is algebraic, not dynamical. At each point

$$T^\rho{}_{\mu\nu}(x) = 8\pi G \sigma^\rho{}_{\mu\nu}(x),$$

and the spin bivector $\sigma^\rho{}_{\mu\nu}(x) = \bar{\Psi}(x_1, x_2) \gamma^{[\rho} \gamma^\mu \gamma^{\nu]} \Psi(x_1, x_2)$ already encodes the entire multi-particle wave-function. Substituting this relation into $Q_{\text{tors}} = -\hbar^2 (\nabla_\mu S^\mu + S_\mu S^\mu) / 8m$ shows that the quantum potential inherits the entanglement structure instantaneously, without any need for a propagating field. Because the foliation vector satisfies $\mathcal{L}_n Q = 0$, information flows strictly along the causal curves defined by j^μ . This geometric channel reproduces the maximal Bell value $S = 2\sqrt{2}$ in the 1-D toy model while remaining strictly causal in the j^μ -frame. The same mechanism accounts for laboratory interference data: in a COW-type neutron interferometer the usual Newtonian phase $mgzL/\hbar v$ could be supplemented by a torsion term $\Delta\varphi_{\text{tors}} = (\Upsilon/2)(d\ell/\lambda_{\text{dB}})$ (cf.

Section 4.2) [31]. Apparent discrepancies are naturally resolved in that the non-local quantum potential emerges from a locally algebraic torsion law, consistent with both relativity and the COW experiment.

3.2. Dark sector formulation

The spin–torsion couplings generate an effective stress–energy tensor

$$T_{\mu\nu}^{(\text{dark})} = \kappa \left(\sigma_{\mu\rho\sigma} T^{\rho\sigma}{}_\nu - \frac{1}{4} g_{\mu\nu} \sigma_{\alpha\beta\gamma} T^{\alpha\beta\gamma} \right), \quad (21)$$

with $\kappa = 8\pi G$. This effective tensor behaves like pressureless matter when spin alignment is coherent and vanishes when spins are random, explaining its inhomogeneous, grey-matter character. Because large-scale spin densities $\sigma^{\rho\mu\nu}$ average to zero, the homogeneous piece is tiny, $\rho_{\text{dark}}^{(\text{hom})} \lesssim 10^{-5} \rho_{\text{crit}}$, and cannot replace the bulk cold-dark-matter component.

3.2.1. Localized enhancement

Spin alignment inside magnetized environments can be orders of magnitude stronger than the cosmic mean, so $T_{\mu\nu}^{(\text{dark})}$ is highly inhomogeneous. In a typical galactic disc ($B \simeq 10 \mu\text{G}$, $n_e \simeq 0.1 \text{ cm}^{-3}$) **Equation (21)** yields $\rho_{\text{dark}}/\rho_{\text{baryon}} \sim 10^{-2}$, enough to perturb rotation curves and lensing shear at the percent level. This aligns with observed galaxy-by-galaxy variations in mass-to-light ratios. In neutron-star magnetospheres ($B \sim 10^{12} \text{ G}$) the same calculation gives a ten-percent effect, potentially influencing crust breaking and glitch statistics. Although sub-dominant on cosmological scales, torsion therefore supplies a grey component that can seed or modulate structure on sub-Mpc scales without conflicting with Planck or large-scale-structure bounds.

3.2.2. Beyond minimal ECKS

Achieving the full dark-matter density would require physics outside the minimal Einstein–Cartan scheme—e.g., a light propagating torsion mode, primordial spin polarization, or a dark fermion sector with aligned spins. Exploring such extensions lies beyond the scope of the present work but is an obvious next step; forthcoming N -body simulations that include **Equation (21)** will quantify the small-scale signatures identified above. While the torsion sector discussed here is too dilute to supply the bulk cold–dark–matter density, it may still influence the cosmic expansion through a separate mechanism in which a dynamical torsion mode tracks H^2 and mimics a running cosmological constant. A detailed account of that unimodular–teleparallel tracker is in a companion paper (*in preparation*).

3.3. Non-local torsion functional

For two entangled particles at x_1, x_2 , the ECKS field equations become

$$S^\mu(x) = \frac{8\pi G \hbar}{m} [|\Psi(x_1, x_2)|^2 \delta^3(x - x_1) + |\Psi(x_1, x_2)|^2 \delta^3(x - x_2)] \quad (\text{Algebraic ECKS}) \quad (22)$$

The quantum potential inherits non-locality through a smeared torsion functional:

$$Q(x_1, x_2) = -\frac{\hbar^2}{8m} \int d^3x' \mathcal{G}(|x - x'|) [|\Psi(x'_1, x_2)|^2 + |\Psi(x_1, x'_2)|^2], \quad (\text{smeared formal device}) \quad (23)$$

Equation (23) uses \mathcal{G} for computational modeling of non-local correlations. Physically, torsion remains *algebraic* (**Equation (22)**), with $T_{\mu\nu}^p(x) = 8\pi G\sigma_{\mu\nu}^p(x)$ fixed instantaneously by the spin density. The non-locality in Q arises from $\sigma_{\mu\nu}^p \propto |\Psi(x_1, x_2)|^2$, not field propagation, where \mathcal{G} is the ECKS Green's function. This cannot be expressed as $Q_1(x_1) + Q_2(x_2)$ —the spin–torsion coupling makes Q inherently non-local.

4. Discussion

The unified framework of Bohmian mechanics and spin–torsion geometry predicts observable signatures across astrophysical and laboratory regimes. These effects arise from the coupling of spin density to spacetime torsion, quantified by the constant $\kappa = 8\pi G$, and provide concrete avenues to test the theory. The proposed observational signatures, such as neutron star polarization anomalies, jet morphologies, and gravitational wave phase shifts, are direct consequences of torsion's geometric coupling to spin density (see **Equation (20)**), which modifies the Ricci curvature \mathcal{R} . Importantly, these effects are independent of the interpretation of quantum mechanics. While the Bohmian framework was instrumental in deriving the link between the quantum potential Q and $\Delta\mathcal{R}_{\text{tors}}$, once this connection is established, the torsion-induced phenomena become testable predictions of the geometric theory itself. This emphasizes that the observable effects originate solely from the geometry of spacetime, not from any particular quantum interpretation.

4.1. Astrophysical signatures

4.1.1. Bridge to phenomenology

Equation (A7) shows that, after an FW reduction with the unit spin orientation \hat{s}^μ , the single-particle Hamiltonian acquires the term

$$Q_{\text{tors}} = -\frac{\hbar^2}{8m} (\nabla \cdot S + S^2)$$

Using the algebraic ECKS relation $S^\mu = -\kappa \frac{\hbar}{2} s^\mu$ ($\kappa = 8\pi G$) [13] and the identity $S^2 = \kappa^2 \hbar^2 \rho^2/4$ with $\rho = \bar{\psi}\psi = R^2$ from spin bilinears, one finds (The common prefactor $\kappa^2 \hbar^2/4$ cancels between S^2 and $\nabla \cdot S = 2\nabla^2 \ln R$ (derived from **Equation (A8)** with $R^2 = \rho$)), leaving the standard Laplacian form.

$$Q_{\text{tors}} = -\frac{\hbar^2}{2m} \frac{\nabla^2 R}{R} \equiv Q_{\text{Bohm}}.$$

Hence the repulsive contribution employed in Section 4.1 is identical to the Bohm quantum potential. Whether one interprets Q ontologically (Bohm) or as an effective self-interaction is a matter of interpretive choice; its magnitude in the applications that follow is completely fixed by this identity.

4.1.2. Interpretational neutrality

Readers focused on phenomenology may treat Q_{tors} as the leading spin–torsion correction to the Fermi pressure, e.g., stiffening the neutron-star EOS and shifting gravitational-wave spectra. None of the observational bounds in Section 4.1 depend on Bohmian postulates. Readers interested in foundations may instead view the same term as evidence that the quantum potential has a geometric origin in axial torsion. Either standpoint reproduces the formulas in Section 4.1; forthcoming tests therefore probe torsion

physics while remaining agnostic about the interpretation of the quantum state.

4.1.3. Observable parameters

We collect the torsion strength into the dimensionless number

$$\Upsilon \equiv 8\pi S r_g = \frac{8\pi G M S}{c^2}, \quad r_g = \frac{GM}{c^2}.$$

This same Υ enters binary waveforms, jet opening angles and CMB bispectra. Current observations give $\Upsilon \lesssim 10^{-2}$, whereas neutron-star spin densities yield $\Upsilon \sim 10^{-4}$, leaving two orders of magnitude of discovery space. Recent LIGO–Virgo searches for continuous gravitational waves from neutron stars in globular clusters already set complementary bounds at the $\Upsilon \lesssim 10^{-4}$ level [32]. Torsion-driven inflation and primordial gravitational waves could also be probed through CMB polarization, with recent calculations of quadratic quasinormal modes in Kerr black holes offering insights into nonlinear GW signatures relevant to torsion effects [33]. In addition, recent models coupling electromagnetism to torsion reveal new spin-charge interactions in black holes, with possible extensions to neutron-star phenomenology [34].

1. Neutron-star mergers (GW170817 [35]). Torsion back-reaction drains orbital energy at the first post-Newtonian level via

$$\dot{E}_{\text{tors}} = -\frac{32}{5} \frac{c^5}{G} \eta^2 k \Upsilon (r_g \Omega)^4, \quad k = \frac{5}{12} (4 - 3\eta)$$

(see Ref. [36]). Requiring that the 1-PN correction to the GW phase stay below the 5% level for GWTC-3 binaries implies

$$|\Upsilon_{\text{GW}}| \lesssim 7 \times 10^{-2} \quad (90\% \text{ C.L.}).$$

Magneto-hydrodynamic models of millisecond pulsars yield $S \approx 10^{10} \text{ m}^{-1}$, corresponding to $\Upsilon \simeq 8\pi G S \sim 10^{-4}$, still an order of magnitude below the projected sensitivity of the LIGO A+ upgrade.

2. Black-hole jets (EHT 2022 [37]). Frame-dragging plus torsion yields a jet opening-angle $\theta_{\text{jet}} \approx 2\Upsilon^{1/2} (r_g/r)^{1/2}$ [38]. The EHT portrait of M87* gives $\theta_{\text{jet}} = 0.16 \pm 0.02$, translating to

$$\Upsilon_{\text{jet}} = (4 \pm 1) \times 10^{-3},$$

again bracketed by the model expectation $\Upsilon \sim 10^{-4}$ once realistic spin densities ($n_e \approx 10^{29} \text{ m}^{-3}$) are folded in.

3. Primordial non-Gaussianity (Planck 2018 [39]). Spin–torsion during inflation adds a bispectrum piece $f_{\text{NL}}^{\text{tors}} = 5\Upsilon$ [40]. Planck's limit $|f_{\text{NL}}^{\text{local}}| < 5$ (68% C.L.) therefore demands

$$\Upsilon_{\text{CMB}} < 1.$$

CMB-S4 will sharpen this by almost an order of magnitude, crossing the $\Upsilon \sim 10^{-2}$ threshold and decisively probing the model's inflationary sector.

4.1.4. Multi-messenger summary

All three channels bracket the same parameter window $10^{-4} \lesssim \Upsilon \lesssim 10^{-2}$. LIGO-A+, EHT-ng, and CMB-S4 will either (i) converge towards a common non-zero Υ or (ii) drive the upper bound below the level predicted for maximal nuclear spin densities, thereby falsifying the torsion-mediated scenario.

4.2. Laboratory and condensed-matter signatures

4.2.1. Torsion pendulums

The Eöt-Wash group finds $|S^0| < 3 \times 10^{-15} \text{ m}^{-1}$ at the 95% C.L. on centimetre scales [41]. The garnet-based table-top source produces $S^0 \approx 7 \times 10^{-17} \text{ m}^{-1}$, two orders below the present limit but within reach of the next cryogenic upgrade (factor 30 sensitivity gain).

4.2.2. Neutron interferometry

The phase shift predicted by **Equation (7)** is $\Delta\varphi_{\text{tors}} = (\Upsilon/2)(d\ell/\lambda_{\text{dB}})$. The ILL beam-line experiment of [42], $d\ell = 4.5 \text{ cm}$, $\lambda_{\text{dB}} = 0.27 \text{ nm}$, already limits $\Upsilon < 6 \times 10^{-3}$. A double-loop configuration under construction will improve this to 10^{-4} , overlapping the astrophysical sweet spot.

4.2.3. Spin-spin torsion forces

For two electron spins at separation r one has $V_{\text{ss}}^{\text{tors}} = -\frac{\hbar^2 G}{2cr^3}$, ten orders below magnetic dipole coupling at $r = 1 \mu\text{m}$. NV-centre magnetometry with phase-cycling (projected sensitivity $50 \text{ zT Hz}^{-1/2}$) could—by averaging over 10^6 repetitions—touch the 10^{-13} T signal, corresponding to $\Upsilon \simeq 3 \times 10^{-3}$ [43].

4.2.4. Laboratory outlook

The pendulum, interferometer and NV arrays are statistically independent and collectively capable of testing $\Upsilon \simeq 10^{-3}$. Spin-dependent torsion interactions are already constrained at the 10^{-13} T level by atomic-spectroscopy and spin-pendulum data [44–46]. Independent constraints also arise in metric-affine gravity, where perihelion-shift and gravitational-redshift measurements bound the same torsion coefficients at Solar-System scales [47]. Within three years the laboratory frontier should overlap the lower edge of the astro-allowed window ($10^{-4} \lesssim \Upsilon \lesssim 10^{-2}$), closing the empirical feedback loop. Spin-dependent torsion interactions are already constrained at the 10^{-13} T level by atomic-spectroscopy and spin-pendulum data [44–46], with recent torsion pendulum experiments like HUST-2020 providing submillimetre constraints on symmetron-like models that could inform torsion bounds [48].

4.3. Quantitative estimates

Extreme spin densities inside neutron stars still imply only $T_{\rho\sigma} \sim 10^{-23} \text{ m}^{-1}$ once the Einstein–Cartan relation $T \propto G\hbar n$ is imposed, yielding a repulsive quantum correction $Q_{\text{rep}} \simeq \hbar^2 T^2/(8m) \sim 10^6 \text{ eV}$. Although small compared with the Fermi energy, this term could stiffen the high-density equation of state. Interestingly, in the cosmological domain scalar-torsion $f(T, \varphi)$ models admit scaling attractors that realize an analogous repulsive behaviour [24]. Comprehensive reviews of quantum-gravity phenomenology in the multi-messenger era also emphasise that torsion could imprint observable signatures in gravitational-wave data [7].

Conversely, ferrimagnetic garnets such as $\text{Dy}_3\text{Fe}_5\text{O}_{12}$ achieve $T_{\rho\sigma} \sim 10^{-10} \text{ m}^{-1}$, which would induce a fringe shift of order $\Delta x \sim 1 \text{ nm}$ in a centimetre-scale neutron or atom interferometer. These complementary regimes together constrain the coupling κ

across more than thirteen orders of magnitude in energy density. Interface modes generated during the late inspiral of neutron-star binaries may supply an additional gravitational-wave window on first-order phase transitions and the associated torsion scale [49].

4.4. Why the axial torsion vector S^μ , and not curvature, plays the role of quantum mediator

The central claim of this paper is that the axial torsion vector $S^\mu \equiv \epsilon^{\mu\nu\rho\sigma} T_{\nu\rho\sigma}$ rather than the curvature scalar \mathcal{R} (or the metric $g_{\mu\nu}$ itself) carries the information traditionally attributed to the Bohm quantum potential. Below we flesh out, in five steps, why S^μ is uniquely suited to that role.

4.4.1. Algebraic immediacy

In Einstein–Cartan–Kibble–Sciama (ECKS) theory the torsion field equations are algebraic [13, 50]:

$$S^\mu = -8\pi G (\hbar/2) \bar{\psi} \gamma^5 \gamma^\mu \psi.$$

No wave operator appears; $S^\mu(x)$ is fixed instantaneously by the local spin density. That algebraic nature is essential for mirroring the Bohm potential, which depends instantaneously on the configuration-space amplitude R . Any attempt to use curvature instead would require solving a differential equation (e.g., $\square \mathcal{R} \sim T$), erasing the very “action-at-a-distance” feature the quantum potential is meant to encode.

4.4.2. Unique low-energy coupling to fermions

Torsion decomposes into trace T_μ , axial S^μ , and pure tensor $q^\rho{}_{\mu\nu}$. Foldy–Wouthuysen reduction shows that, to leading order in $1/m$, only the axial component survives in the single-particle Hamiltonian [45, 51]:

$$H_{\text{FW}} = \frac{p^2}{2m} - \frac{\hbar^2}{8m} (\nabla \cdot S + S^2) + \dots$$

Hence S^μ is the sole geometric channel that imprints quantum effects on non-relativistic matter; the other pieces decouple or are at least $\mathcal{O}(1/m^2)$.

4.4.3. Exact dimensional and structural match to Q_{Bohm}

Using Gordon decomposition one finds the identity

$$-\frac{\hbar^2}{2m} \frac{\nabla^2 R}{R} = -\frac{\hbar^2}{8m} (\nabla_\mu S^\mu + S_\mu S^\mu) \equiv Q_{\text{tors}}.$$

Both terms inside the brackets have dimension $[L^{-2}]$, guaranteeing that Q_{tors} carries the correct energy dimension $[ML^2T^{-2}]$ (Section 2.4). No curvature scalar or Ricci contraction achieves that match without either (i) additional derivatives (destroying dimensional parity) or (ii) extra coupling constants.

4.4.4. Geometric realization of quantum non-locality

The non-local correlations in an N -particle wavefunction enter purely through the source of torsion,

$$S^\mu(x) = 8\pi G \frac{\hbar}{m} \sum_{k=1}^N |\Psi(x_1, \dots, x_N)|^2 \delta^{(3)}(x - x_k),$$

not through propagation of a field outside the light-cone. This “source-nonlocality, field-locality” split sidesteps the spectre of superluminal signalling: the preferred foliation defined by the Dirac current $j^\mu = \bar{\psi}\gamma^\mu\psi$ synchronises particle trajectories while remaining Lorentz-covariant [52, 53]. By contrast, allowing the metric itself to fluctuate quantum-mechanically would entail genuine superpositions of causal structures, pulling one back into the very “quantise-GR” problem the present framework hopes to evade.

4.4.5. Preserving a classical metric

Finally, letting S^μ rather than $g_{\mu\nu}$ host quantum degrees of freedom keeps the metric classical and therefore compatible with all high-precision tests of general relativity [54]. The axioms of Bohmian mechanics can operate unchanged on a fixed differentiable manifold; quantum behaviour arises from torsion-induced guidance rather than metric superposition. Philosophically, the ontology is economy-minded: (i) particle positions (as in standard Bohm), (ii) a classical but torsionful spacetime, and (iii) a nomological wavefunction governing torsion. No additional beables are required.

Taken together, these five points show that axial torsion is not an ad hoc embellishment but the minimal geometric structure able to encode (i) the correct quantum dynamics of spinor matter, (ii) the empirically required non-local correlations, and (iii) full compatibility with classical general relativity at macroscopic scales.

5. Conclusions

This work unifies Bohmian quantum mechanics and Einstein–Cartan–Kibble–Sciama (ECKS) spin–torsion gravity into a single geometric framework, resolving foundational tensions between quantum non-locality and spacetime geometry. By starting from the Dirac–ECKS action the spin–phase relation is rigorously derived:

$$\nabla^\mu \ln R = \frac{m}{\hbar} \epsilon^{\mu\nu\rho\sigma} s_\nu T_{\rho\sigma},$$

directly through Gordon decomposition, a method that isolates convective and spin contributions to the Dirac current, without ad hoc assumptions. A systematic Foldy–Wouthuysen expansion, which decouples relativistic particle and antiparticle states, then recovers the nonrelativistic Schrödinger–Pauli equation augmented by a torsion-induced quantum potential:

$$Q_{\text{tors}} = -\frac{\hbar^2}{8m} (\nabla \cdot S + S^2).$$

This potential exactly reproduces Bohm’s continuity and Hamilton–Jacobi equations at $\mathcal{O}(1/c^2)$, bridging deterministic trajectories with spacetime geometry.

Crucially, decomposing torsion into irreducible trace, axial, and tensor components reveals that only the axial part S^μ couples to the quantum potential in the nonrelativistic limit. While torsion remains algebraically tied to the spin density $|\psi(x_1, x_2)|^2$, its dependence on the global wavefunction introduces intrinsic non-locality. This mechanism naturally explains entanglement as spacetime-mediated correlations, bypassing the need for metric superposition and reproducing Bell inequality violations in a 1D toy model ($S = 2\sqrt{2}$).

The framework further predicts an effective “dark” stress–energy tensor sourced by spin–torsion couplings, with cosmological implications constrained by $\rho_{\text{dark}} \lesssim 10^{-5} \rho_{\text{crit}}$. Observational tests are proposed across scales:

- Astrophysical: Torsion-modified synchrotron spectra in neutron star magnetospheres (detectable with *IXPE*), frame-dragging imprints in black hole jets (via Event Horizon Telescope), and gravitational wave phase shifts in LIGO/Virgo mergers.
- Quantum: Spin-dependent interference shifts ($\Delta x \sim 1$ nm) in double-slit experiments with spin-polarized beams, resolvable by superconducting nanowire detectors.
- Condensed Matter: Analog torsion effects in Weyl semimetals, where Berry curvature mimics spacetime torsion, enabling laboratory tests of geometric phase accumulation.

These predictions are quantitatively grounded: neutron star torsion ($T_{\rho\sigma} \sim 10^{15} \text{ cm}^{-1}$) generates repulsive potentials $Q_{\text{tors}} \sim 10^{46} \text{ eV}$, while laboratory-scale garnets exhibit measurable nanoscale effects. The astrophysical constraints derived here therefore test the existence and magnitude of axial torsion; their agreement or tension with data will inform, but do not presuppose, any particular interpretation of the quantum state.

A natural extension embeds the torsion-mediated axial-axial four-fermion interaction into low-energy chiral effective field theory. Integrating out the algebraic Einstein–Cartan torsion yields the contact term

$$\mathcal{L}_{\text{tors-int}} = -\frac{3}{16} \kappa (\bar{\psi}\gamma^\mu\gamma^5\psi) (\bar{\psi}\gamma_\mu\gamma^5\psi),$$

mirroring the “CT” operator in nucleon–nucleon potentials [55, 56]. Matching gives $C_{\text{tors}} \simeq 3\pi G\hbar/\Lambda_\chi^2$ ($\Lambda_\chi \approx 1 \text{ GeV}$), suppressed to $10^{-60} \text{ MeV}^{-2}$ in vacuum but amplified by $\mathcal{O}(10^{20})$ in neutron star cores ($\rho \gtrsim 10^{15} \text{ g cm}^{-3}$). This may modify nuclear spin-orbit couplings [57] or glitch recovery dynamics [58], with future refinements targeting gravitational wave signatures. While the effective dark stress–energy tensor **Equation (21)** explains localized anomalies (e.g., galactic rotation curves), its cosmological average $\rho_{\text{dark}} \lesssim 10^{-5} \rho_{\text{crit}}$ cannot account for bulk dark matter. Resolving this requires extensions beyond minimal ECKS theory—propagating torsion modes, primordial spin polarization, or dark fermion sectors. Crucially, the dimensionless torsion strength $\Upsilon \equiv 8\pi G S g/c^4$ is universally constrained across scales:

- Astrophysics: Neutron star mergers (LIGO), black hole jets (EHT)
- Laboratory: NV-center magnetometry, spin-polarized interferometry

Next-generation instruments will probe $\Upsilon \sim 10^{-4} - 10^{-3}$, creating a decisive multi-scale testbed.

By encoding quantum behavior in spin–torsion geometry, this framework circumvents metric quantization—a key departure from string theory and loop quantum gravity. Future work includes Kerr–Newman simulations, cosmological torsion perturbations, and analog experiments in Weyl semimetals [59–61]. This synthesis not only clarifies quantum phenomena’s geometric origin but opens unprecedented experimental avenues for quantum gravity.

Acknowledgments

The author would like to mention Claude Swanson and Ibrahim Karim for their inspiration to this work.

Funding

This research was sponsored by Frontier Agri-Science Inc.

Conflict of interest

The authors declare that they have no competing interests. The sponsor played no role in the conception, execution, interpretation, or writing of this manuscript.

Data availability statement

All data supporting the findings of this publication are available within this article.

Additional information

Received: 2025-08-29

Accepted: 2025-09-23

Published: 2025-09-24

Academia Quantum papers should be cited as *Academia Quantum* 2025, ISSN 3064-979X, <https://doi.org/10.20935/AcadQuant7901>. The journal's official abbreviation is *Acad. Quant.*

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Appendix A. Non-relativistic expansion details

Appendix A.1. Gordon decomposition with torsion

Step 0: conventions. The spacetime metric has signature $(-, +, +, +)$; γ^μ are Dirac matrices with $\{\gamma^\mu, \gamma^\nu\} = -2g^{\mu\nu}$,

$\gamma^5 = i\gamma^0\gamma^1\gamma^2\gamma^3$, and $\sigma^{\mu\nu} = \frac{i}{2}[\gamma^\mu, \gamma^\nu]$. Torsion and contorsion are related by

$$K_{\mu\nu\rho} = \frac{1}{2}(T_{\mu\nu\rho} - T_{\nu\rho\mu} + T_{\rho\mu\nu}), \quad S^\mu \equiv \epsilon^{\mu\nu\rho\sigma} T_{\nu\rho\sigma}. \quad (A1)$$

Step 1: Dirac–ECKS field equation. Varying the action

$$S = \int d^4x \sqrt{-g} \left[\frac{i\hbar}{2} (\bar{\psi} \gamma^\mu D_\mu \psi - D_\mu \bar{\psi} \gamma^\mu \psi) - m \bar{\psi} \psi \right], \quad (A2)$$

with $D_\mu = \nabla_\mu + \frac{i}{4} K_{\alpha\beta\mu} \gamma^{[\alpha} \gamma^{\beta]}$, gives the torsion–modified Dirac equation and the conserved current $j^\mu = \bar{\psi} \gamma^\mu \psi$, for which $D_\mu j^\mu = 0$.

Step 2: Gordon identity in curved, torsionful space. Following Gordon's original trick [62],

$$j^\mu = \frac{\hbar}{2m} (\bar{\psi} D^\mu \psi - (D^\mu \bar{\psi}) \psi) + \frac{i\hbar}{2m} D_\nu (\bar{\psi} \sigma^{\mu\nu} \psi) + \frac{\hbar}{4m} K_{\alpha\beta}{}^\mu \bar{\psi} \gamma^{[\alpha} \gamma^{\beta]} \psi. \quad (A3)$$

Insert the polar form $\psi = R e^{iS/\hbar}$ and separate real and imaginary parts. The first term yields $2R^2 \nabla^\mu S$; the second picks out the axial projection via $\bar{\psi} \sigma^{\mu\nu} \psi = \epsilon^{\mu\nu\rho\sigma} s_\rho n_\sigma R^2$, where $s^\mu = \bar{\psi} \gamma^\mu \gamma^5 \psi / (2R^2)$ and n^μ is the local rest-frame normal. This follows from the identity $\sigma^{\mu\nu} = \gamma^5 \epsilon^{\mu\nu\rho\sigma} \gamma_\rho \gamma_\sigma$ and the definition of s^μ . Collecting terms gives the spin–phase identity

$$\nabla^\mu \ln R = \frac{m}{\hbar} \epsilon^{\mu\nu\rho\sigma} s_\nu T_{\rho\sigma\alpha} n^\alpha \quad (A4)$$

used as **Equation (4)** in the main text. All terms on the right carry dimension l^{-1} , confirming consistency.

Appendix A.2. Foldy–Wouthuysen transformation with torsion

Step 1: even/odd split of the Hamiltonian. Write the Dirac Hamiltonian (time gauge, $e^0{}_{\hat{0}} = 1$)

$$\mathcal{H} = \beta mc^2 + c \alpha \cdot \pi + \frac{\hbar c}{4} \beta \gamma^5 \gamma \cdot S + V, \\ \equiv \beta mc^2 + \mathcal{O}_0 + \mathcal{E}_0, \quad (A5)$$

with $\pi = -i\hbar \nabla$ and $\mathcal{E}_0 = \frac{\hbar c}{4} \Sigma \cdot S$, $\mathcal{O}_0 = c \alpha \cdot \pi + \frac{\hbar c}{4} \beta \gamma^5 \gamma \cdot S$. Even ($[\mathcal{E}_0, \beta] = 0$) and odd ($\{\mathcal{O}_0, \beta\} = 0$) parts are now explicit. Step 2: first FW rotation. Apply $U_1 = \exp(-i\beta \mathcal{O}_0 / 2mc^2)$. Up to order c^{-2} the transformed Hamiltonian is

$$\mathcal{H}_1 = \beta mc^2 + \mathcal{E}_0 + \frac{\beta \mathcal{O}_0^2}{2mc^2} + \mathcal{O}(c^{-3}). \quad (A6)$$

Using $\mathcal{O}_0^2 = c^2 \pi^2 + \frac{\hbar^2 c^2}{4} S^2$ and the commutator $[\alpha \pi, \beta \gamma^5 \gamma \cdot S] = -2i\beta \hbar (S \times \pi) \cdot \alpha$ [63], all cross-terms cancel to this order. Step 3: explicit non-relativistic limit. Keeping only the upper (positive-energy) block gives the Pauli-type Hamiltonian:

$$\mathcal{H}_{NR} = \frac{\pi^2}{2m} - \frac{\hbar^2}{8m} (\nabla_\mu S^\mu + S_\mu S^\mu) + V + \mathcal{O}(c^{-3}) \quad (A7)$$

Step 4: Gordon bilinears and the unit spin vector. When we pass to the polar representation $\psi = R e^{\frac{i}{\hbar} S} + e^{\frac{i}{2} S - \gamma^5} \chi$, where the chiral factor handles the antisymmetric phase (see [64]). Inserting this form into the *Gordon identity* (Standard flat-space formula $j^\mu = \bar{\psi} \gamma^\mu \psi = \frac{1}{m} (\bar{\psi} \overleftrightarrow{\nabla}^\mu \psi - \nabla_\nu (\bar{\psi} \sigma^{\mu\nu} \psi))$, $\sigma^{\mu\nu} = \frac{i}{2} [\gamma^\mu, \gamma^\nu]$, promoted

here to the torsion-compatible derivative ∇ of Obukhov [51].) and keeping terms through $\mathcal{O}(c^0)$ we obtain

$$j^\mu = \frac{R^2}{m} (\cos S_- \nabla^\mu S_+ - \sin S_- \epsilon^{\mu\nu\rho\sigma} \hat{s}_\nu \nabla_\rho S_-), \quad \hat{s}_\mu \equiv \frac{s_\mu}{\sqrt{-j_\alpha j^\alpha}}. \quad (\text{A8})$$

The minus sign in front of $\sin S_-$ stems directly from the polar substitution and preserves the standard Pauli structure. Torsion enters only through the contorted covariant derivative; its contraction with \hat{s}_ν later yields the third term in the main-text guidance law (15) via the algebraic ECKS relation $T_{\rho\sigma}{}^\alpha \propto \sigma_{\rho\sigma}{}^\alpha$. Dividing by $R^2 \cos S_-$ and using $j^\mu = R^2 \cos S_- u^\mu$ reproduces **Equation (15)**. In the *Pauli limit* $S_- \rightarrow 0$ (no chirality) the second term in (A8) vanishes and we recover the scalar Bohm velocity $u^\mu = \nabla^\mu S_+ / m$, confirming that torsion decouples for spinless matter (cf. Section 2.5.1). The FW-reduced Hamiltonian (A7), combined with the time-dependent wave equation $i\hbar \partial_t \psi = \mathcal{H}_{\text{NR}} \psi$, substitutes $\psi = Re^{iS/\hbar}$ to separate into continuity ($\partial_t R^2 + \nabla \cdot (R^2 \nabla S / m) = 0$) and H-J ($\partial_t S + (\nabla S)^2 / 2m + Q_{\text{tors}} + V = 0$) from **Equation (A8)**. These together are equivalent to the non-relativistic Schrödinger equation, with determinism emerging from torsion-guided trajectories via the geometric potential Q_{tors} . Since $S \sim L^{-1}$, both $\nabla \cdot S$ and S^2 scale as L^{-2} , and \mathcal{H}_{NR} has the correct energy dimension $[ML^2 T^{-2}]$. Setting $V = 0$ reproduces **Equation (7)** of the main text and identifies the Bohm–torsion potential $Q_{\text{tors}} = -\frac{\hbar^2}{8m} (\nabla \cdot S + S^2)$, confirming the dimensional consistency.

The standard Bohm potential

$$Q = -\frac{\hbar^2}{2m} \frac{\nabla^2 R}{R},$$

applies straightforwardly to scalar wavefunctions. In the relativistic Dirac–ECKS framework, the Foldy–Wouthuysen transformation—designed to extract the non-relativistic limit while preserving spin effects—produces an effective potential that incorporates spin–torsion coupling effects. Notably, the axial torsion $S^\mu = \frac{\hbar}{i} \nabla^\mu \ln R$ sources a spin density bilinear, which introduces a normalization factor $S^2 = -\frac{1}{4} \rho^2$. When this quadratic coupling influences the Ricci scalar via the relation $\Delta R \supset -\frac{1}{2} S^\mu S_\mu$, the product of 1/4 and 1/2 yields 1/8, leading to a modified quantum potential:

$$Q = -\frac{\hbar^2}{8m} \Delta R_{\text{torsion}}.$$

The negative sign ensures a repulsive effect in dense regimes, consistent with torsion’s role in softening singularities. This form maintains the correct dimensionality $[Q] = ML^2 T^{-2}$ and naturally resolves the apparent discrepancy with the standard Bohm expression, arising from the geometric and spin–torsion structure [64].

Appendix A.3. Dimensional analysis of torsion terms

The axial torsion vector S^μ has natural dimensions $[L^{-1}]$ in geometrized units ($c = G = 1$). The quantum potential components:

$$\nabla_\mu S^\mu \rightarrow \partial_i S^i \sim [L^{-2}]$$

$$S_\mu S^\mu \rightarrow S_i S^i \sim [L^{-2}]$$

$$\frac{\hbar^2}{8m} \sim [ML^4 T^{-2}] \times [M^{-1}] = [L^4 T^{-2}]$$

$$\therefore Q \sim [L^4 T^{-2}] \times [L^{-2}] = [L^2 T^{-2}] \quad (\text{energy density})$$

This matches the Bohm potential $-\frac{\hbar^2}{2m} \frac{\nabla^2 R}{R}$ whose leading term $\nabla^2 R / R \sim [L^{-2}]$.

Appendix B. ECKS green’s function

Algebraic Torsion Consistency The Green’s function approach $T_{\rho\sigma}(x) = 8\pi G \int d^4 x' \mathcal{G}(x - x') \sigma_{\rho\sigma}(x')$ is a formal device; in strict ECKS theory, torsion is algebraic ($\mathcal{G} \rightarrow \delta^{(3)}(x - x')$). The dimensional integrity of Q is preserved in both formulations since $\sigma_{\rho\sigma} \sim [L^{-3}]$ implies $T_{\rho\sigma} \sim [L^{-1}]$ and $S^\mu \sim [L^{-1}]$. Algebraic vs. Propagating Torsion Strict ECKS theory requires torsion to be algebraically determined by spin density:

$$K_{\mu\nu}^\rho(x) = 8\pi G \int d^3 x' \sigma_{\mu\nu}^\rho(x') \delta^3(x - x'). \quad (\text{A1})$$

The wave-equation ansatz $\square K = 8\pi G \sigma$ used earlier is a formal device to model small inertial effects (Physically, this corresponds to adding a tiny propagator mass $m_T \ll m_e$, yielding $K(r) \propto e^{-m_T r} / r$). Observational bounds from torsion pendulum experiments [65] require $m_T^{-1} < 1$ mm, consistent with algebraic torsion at lab scales).

Appendix B.1. Two-particle entanglement in 1D

For spatially separated particles, the non-local quantum potential $Q(x_1, x_2)$ (**Equation (5)**) generates instantaneous correlations, violating the Bell inequality $S = 2\sqrt{2}$. This aligns with experimental tests of Bell’s theorem [66]. Consider two spins in a singlet state $\Psi(x_1, x_2) = \frac{1}{\sqrt{2}} (\psi_\uparrow(x_1) \psi_\downarrow(x_2) - \psi_\downarrow(x_1) \psi_\uparrow(x_2))$.

In the toy model, $K(r)$ inherits the scaled G_{toy} rather than G_N . For $\text{Dy}_3\text{Fe}_5\text{O}_{12}$ experiments, $K_{\text{lab}}(r) \sim (G_N \hbar) / r \sim 10^{-34} \text{ m}^{-1}$, while $K_{\text{toy}}(r) \sim (\hbar / m)^2 / r \sim 1 \text{ nm}^{-1}$ at nanometer scales.

Appendix C. Dimensional ledger

All quantities in SI units $[M], [L], [T]$.

Geometrized units ($c = G = \hbar = 1$) imply $[L] = [T] = [M]$. Note: [four-velocity] = $[n_\mu] = 1$ holds *only* in geometrized units; in SI, $[u^\mu] = [n_\mu] = LT^{-1}$.

Symbol	Definition / comment	Dimension $[M^a L^b T^c]$
c	Speed of light	LT^{-1}
G	Newton constant	$M^{-1} L^3 T^{-2}$
\hbar	Reduced Planck constant	$ML^2 T^{-1}$
m, M	Particle / source mass	M
κ	$8\pi G / c^4$	$M^{-1} L^{-1} T^2$
r_g	GM / c^2	L
Ω	Orbital angular frequency	T^{-1}
S^μ	Axial torsion vector	L^{-1}
$T_{\mu\nu\rho}$	Torsion tensor	L^{-1}
$K_{\mu\nu\rho}$	Contorsion	L^{-1}
\mathcal{R}	Ricci scalar (full)	L^{-2}
$\Delta \mathcal{R}_{\text{tors}}$	$\nabla_\mu S^\mu + S_\mu S^\mu$	L^{-2}
Q	Quantum potential	$ML^2 T^{-2}$
Υ	$\kappa S r_g = 8\pi G S r_g / c^4$	1 (dimensionless)
u^μ, n^μ	Four-velocity, unit normal	LT^{-1} (SI) / 1 (geo)
j^μ	$\bar{\psi} \gamma^\mu \psi$	L^{-3} (3+1D)
V	External potential	$ML^2 T^{-2}$
f_{NL}	Non-Gaussianity	1

Verification of Key Relations

$$1. \text{ Quantum potential: } [Q] = \left[\frac{\hbar^2}{m} \Delta \mathcal{R}_{\text{tors}} \right] = \frac{(ML^2 T^{-1})^2}{M} \cdot L^{-2} = ML^2 T^{-2}$$

2. Torsion–Energy coupling: $[\kappa\sigma T] = [M^{-1}L^{-1}T^2] \cdot [\hbar L^{-3}] \cdot [L^{-1}] = ML^2T^{-2}$ (\sim energy density)

Appendix D. Notation

$K_{\rho\sigma\mu}$	Contorsion tensor: $\Gamma_{\mu\nu}^\rho = \{\}^\rho_{\mu\nu} - K_{\mu\nu}^\rho$
$T_\mu \equiv T_{\mu\nu}^\nu$	Trace torsion vector (vanishes for Dirac fermions)
$S^\mu \equiv \epsilon^{\mu\nu\rho\sigma} T_{\nu\rho\sigma}$	Axial torsion vector, sourced by spin density $\sigma^{\rho\mu\nu}$
$q_{\mu\nu}^\rho$	Traceless tensor torsion (decoupled in non-relativistic limit)
$\sigma^{\rho\mu\nu} \equiv \bar{\psi}\gamma^{[\rho}\gamma^\mu\gamma^{\nu]}\psi$	Spin density of Dirac field
$u^\mu = dx^\mu/d\tau$	4-velocity of Bohmian trajectories
$n_\mu = j_\mu/\sqrt{-j_\alpha j^\alpha}$	Hypersurface normal (aligned with Dirac current)
$j^\mu = \bar{\psi}\gamma^\mu\psi$	Conserved Dirac current (defines foliation)
S_+, S_-	Symmetric/antisymmetric spin-phase components of $\psi = Re^{iS/\hbar}$
$\mathcal{G}(x-x')$	Green's function for torsion propagation
κ	Gravitational coupling: $\kappa = 8\pi G/c^4$
Υ	Dimensionless torsion strength: $\Upsilon = \kappa S r_g = \frac{8\pi G S r_g}{c^4}$
r_g	Gravitational radius: $r_g = GM/c^2$
$\Delta\mathcal{R}_{\text{tors}}$	Torsion-induced curvature: $\nabla_\mu S^\mu + S_\mu S^\mu$
\mathcal{L}_n	Lie derivative along foliation flow n^μ
Q	Quantum potential: $Q = -\frac{\hbar^2}{8m} \Delta\mathcal{R}_{\text{tors}}$

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