

Comparing Random Forest and Convolutional Neural Networks on ASTRI dedicated MC production

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Outline

This (ongoing) work aims at:

- comparing performances of a C++ Shark implementation of machine learning tools to recognize gammas from hadrons, with a Python scikit-learn one;
- comparing performances of RandomForest model implemented in scikit-learn with a CNN implementation in keras.



ASTRI MC Data

I used an ASTRI dedicated MC simulated data (CTA Prod3b) for this work:

- 33 ASTRI telescopes in square layout
- Events from each telescope are used as independent, as if coming from a single average telescope
- A data challenge has been extracted from this dataset, but not used in this work



Workflow

Predict class probabilities, energy, direction with:

- → Scikit-learn libraries
- → Shark libraries

The inputs for the above set of tools are Hillas parameters files (lv1b) filtered like this:

COLUMNS = ['SIZE', 'WIDTH', 'LENGTH', 'CONC', 'DIST', 'M3LONG', 'COSDELTAALPHA'] FILTERS = "SIZE > 50 && WIDTH > 0 && LEAKAGE < 0.1 && NUMISLAND < 2 && NUMCORE > 2 && LENGTH > 0"

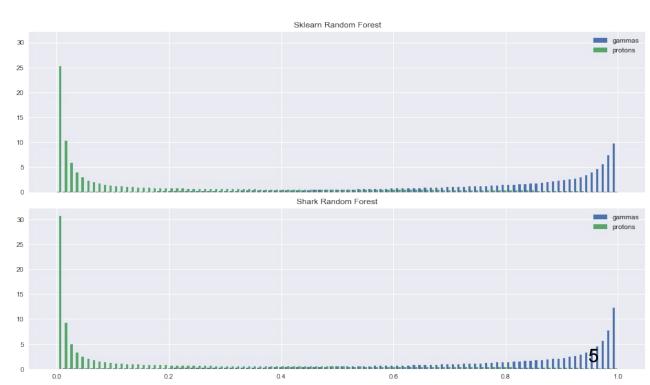


Gamma - Hadron separation

Gamma Hadron separation

Comparison between Shark and scikit-learn libraries

Normed histograms showing the density of particles falling in bins of gammaness.



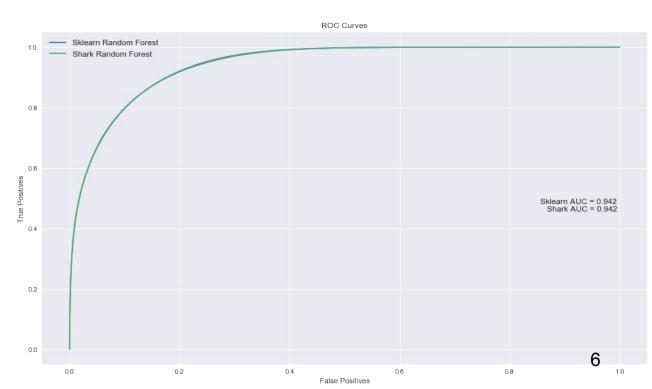


ROC curve and area

ROC curves typically feature true positive rate on the Y axis, and false positive rate on the X axis. This means that the **top left corner of the plot is the "ideal" point**.

Larger area under the curve (AUC) is usually better.

It can be seen that the two methods are somewhat equivalent, a direct measure being the score, namely the area under the curves.



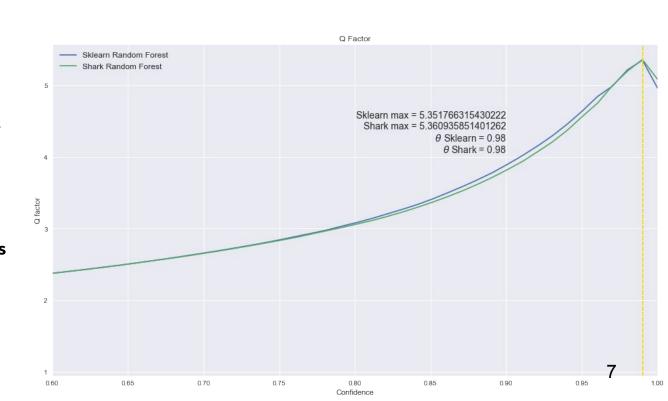


Q-factor

Q factor is widely used in astrophysics literature to measure the goodness of classification: it is measured as

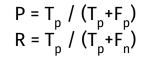
Q =
$$\epsilon \gamma / \sqrt{\epsilon p}$$

where $\epsilon \gamma$ is the fraction of well classified gammas, and ϵp is the fraction of badly classified protons (false positives). This is usually plotted against the acceptance (or threshold) for the probabilities.





Precision - Recall



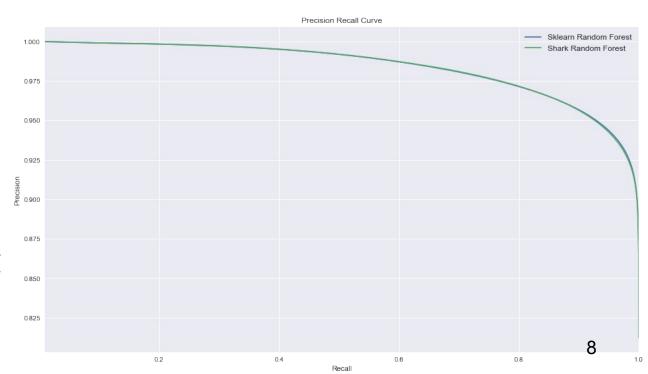
Where

T_n are gammas correctly labelled

 F_p are hadrons labelled as gammas

F_n gammas labelled as hadrons

Low Precision means many hadrons have been mis-labelled, while a low Recall means many gammas have been classified as protons (and then lost!)





Comments

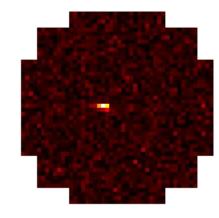
It can be said Shark and scikit-learn implementations of Random Forest seem to have the same performances on this dataset.

From the gamma-hadron separation plot, it seems Shark is able to better fill the edge bins, thus assigning probabilities close to 0 (hadrons) and 1 (gammas) to more individuals in the population.

Other thing worth noting, sklearn is much **faster** in execution: the same dataset used for training the classifier object is run in minutes from the Python library, hours from Shark.

Convolutional Neural Network

A five layers **CNN** has been trained with ASTRI calibrated data (*lv1a*), and then used to predict.



Network arch was taken from an example on keras blog to perform classification on dog and cats pictures: https://blog.keras.io/building-powerful-image-classification-models-using-very-little-data.html

The aim was to fast prototype, but very nice results anyway!

Layer (type)	Output Shape	Param #
conv2d_1 (Conv2D)	(None, 54, 54, 32)	320
batch_normalization_1 (Batc	h (None, 54, 54, 32)	128
activation_1 (Activation)	(None, 54, 54, 32)	0
max_pooling2d_1 (MaxPooling	2 (None, 27, 27, 32)	0
conv2d_2 (Conv2D)	(None, 25, 25, 32)	9248
batch_normalization_2 (Batc	h (None, 25, 25, 32)	128
activation_2 (Activation)	(None, 25, 25, 32)	Ø
max_pooling2d_2 (MaxPooling	2 (None, 12, 12, 32)	0
conv2d_3 (Conv2D)	(None, 10, 10, 64)	18496
batch_normalization_3 (Batc	h (None, 10, 10, 64)	256
activation_3 (Activation)	(None, 10, 10, 64)	0
max_pooling2d_3 (MaxPooling	2 (None, 5, 5, 64)	0
flatten_1 (Flatten)	(None, 1600)	0
dense_1 (Dense)	(None, 64)	102464
batch_normalization_4 (Batc	h (None, 64)	256
activation_4 (Activation)	(None, 64)	0
dropout_1 (Dropout)	(None, 64)	0
dense_2 (Dense)	(None, 1)	65
batch_normalization_5 (Batc	h (None, 1)	4
activation_6 (Activation)	(None, 1)	0
Total params: 131,365 Trainable params: 130,979 Non-trainable params: 386		



Convolutional Neural Network (keras + tensorflow)

Input

I extracted 16 bit grayscale png images from ASTRI lv1a (calibrated) data; this allowed to store full photo electron equivalent values.

Dataset was splitted this way:

- 140000 events for training (70000 each population, gammas and protons);
- 60000 events for testing (30000 each);
- 200000 events for prediction (200000 each).



Convolutional Neural Network (keras + tensorflow)

General results

- Training (and testing) took less than two hours on a Nvidia Tesla K20 GPU
- Prediction took 1 hour (30 epochs, stopped after 11 checking on val_acc, with patience=5 and min_delta=0.001), and gave this performance:
 - val_acc = 0.967
 - val_loss = 0.096
 - val_rmse = 0.045

As a first comparison, RF (optimized via GridsearchCV) best oob_scores was 0.86

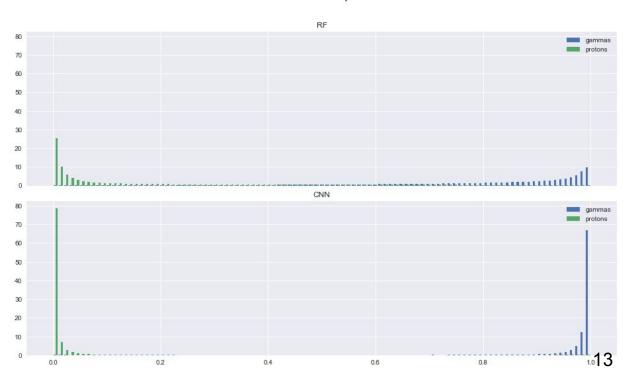


<u>Gamma - Hadron separation</u>

Gamma Hadron separation

Normed histograms showing the density of particles falling in bins of gammaness.

CNN better separates populations.



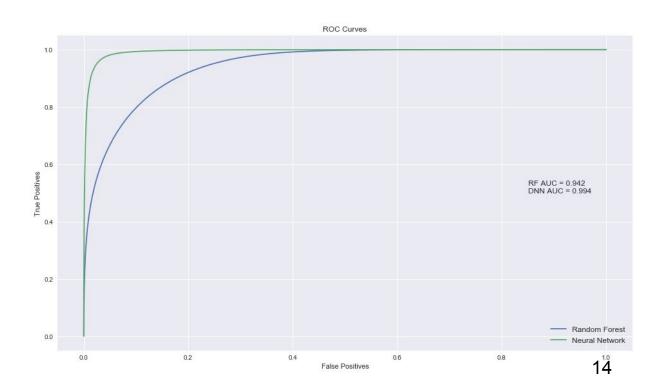


ROC curve and area

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CNN performs significantly better.



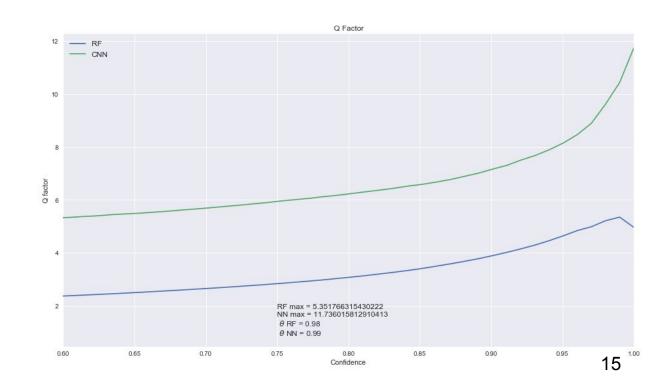


Q-factor

Q factor is widely used in astrophysics literature to measure the goodness of classification: it is measured as

$$Q = \epsilon_{v} / \sqrt{\epsilon_{p}}$$

where ε_{γ} is the fraction of well classified gammas, and ε_{p} is the fraction of badly classified protons (false positives). This is usually plotted against the acceptance (or threshold) for the probabilities.





<u>Precision - Recall</u>

$$P = T_p / (T_p + F_p)$$

$$R = T_p / (T_p + F_n)$$

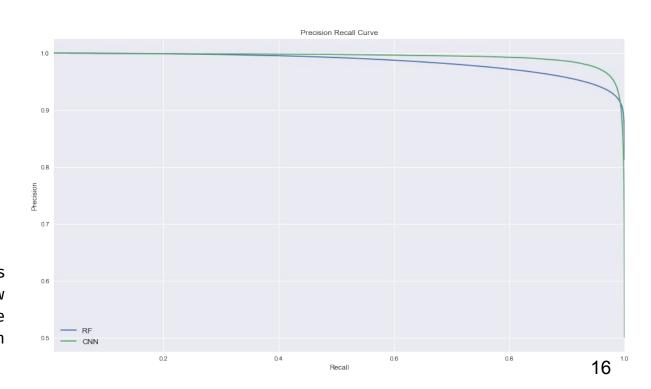
Where

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F_D are hadrons labelled as gammas

F_n gammas labelled as hadrons

Low Precision means many hadrons have been mis-labelled, while a low Recall means many gammas have been classified as protons (and then lost!)





Backup slides



ASTRI Data Challenge #1

Main facts on ASTRI DC1:

The events were properly filtered so to follow the experimental energy slope of -2.70, as measured by the BESS Coll. After this procedure, the available statistics resulted in ~4.2×10⁶ triggered events. Since the event rate was calculated to be ~100 Hz, these amount of events corresponded to ~11.6 hours of (single telescope) data taking.

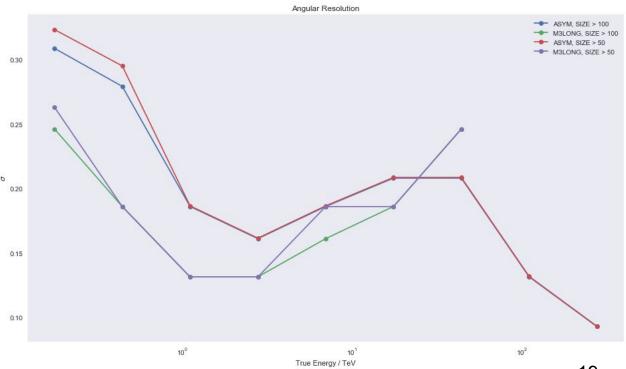
The overall data sample was then split in two subsamples, "ON" and "OFF". We combined the ON sample (~5.8 hours) with a number of randomly selected MC gamma-ray events calibrated to match the flux of the Crab Nebula as measured by the HEGRA Coll. The OFF sample, instead, was slightly reduced from ~5.8 to ~5.5 hours in order to keep a proper amount of independent proton events in their original MC format for the generation of gamma/hadron separation look-up-tables (LUTs)

Future plan: use ON and OFF subsamples for an end to end analysis with CNN.



Angular resolution

We explored performances of two head-tail parameters in angular resolution, finding M3LONG performs better than ASYM up to ~30 TeV





Confusion Matrix for Sklearn Random Forest Classifier

From wikipedia:

In the field of machine learning and specifically the problem of statistical classification, a confusion matrix, also known as an error matrix, is a specific table layout that allows visualization of the performance of an algorithm. The name stems from the fact that **it makes it easy to see if the system is confusing two classes** (i.e. commonly mislabelling one as another).

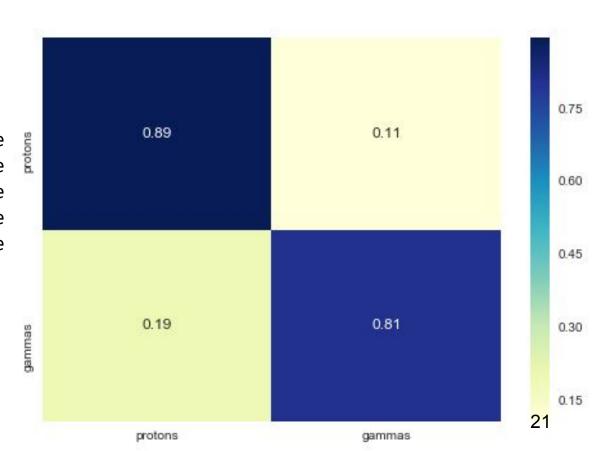
Sounds useful!



Confusion Matrix

The diagonal elements represent the number of points for which the predicted label is equal to the true label, while off-diagonal elements are those that are mislabeled by the classifier.

['True negatives', 'False positives']
['False negatives', 'True positives']

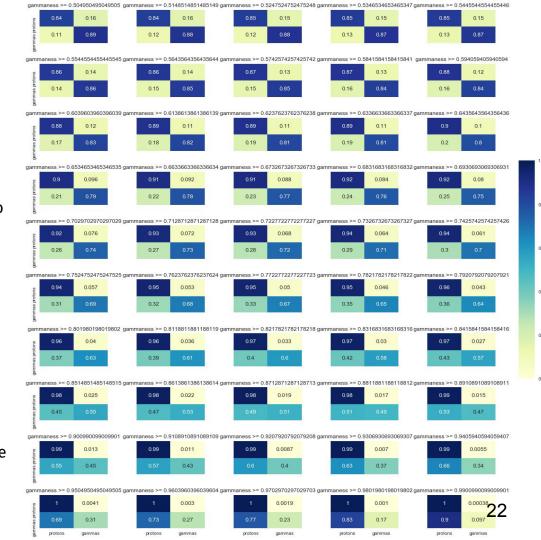


Confusion Matrix

Some interpretations:

from a gammaness value of ~64% on, you start to throw away more than 20% of true gammas: this can be useful to know for if you need a proper amount of gammas to perform some scientific analysis (DL4 data);

Q-factor is max at gammaness = 98%; this is because √fp (false positives, background, protons classified as gammas) are the denominator of the Q-factor and its value become very small as gammaness increase. The counter effect is for those values you are wasting ~80% of gammas.





Other Classifiers

Same input data have been passed through other machine learning methods, namely:

Multi Layer Perceptron (Neural Network)

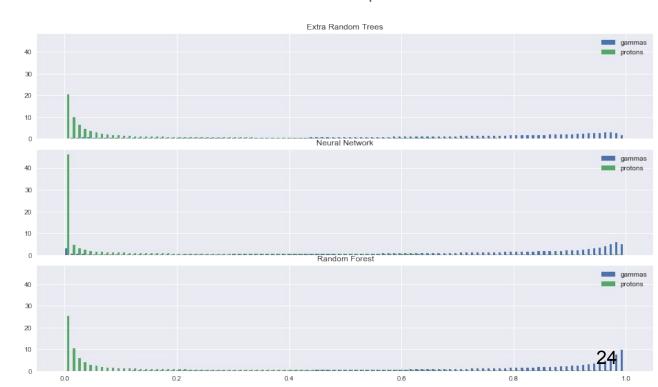
About Multi Layer Perceptron (Neural Network): it's been used a shallow network with 20 neurons in an only hidden layer, after a number of tests with two hidden layers and varying the number of neurons also: this shallow network gives the best result on this dataset, in a reasonable execution time.



Gamma - Hadron separation

Gamma Hadron separation

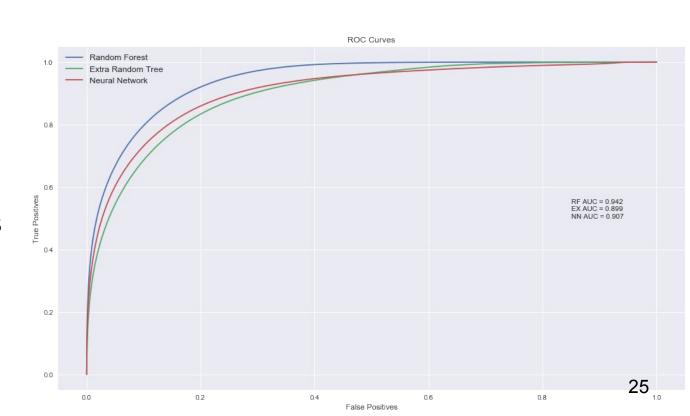






ROC curve

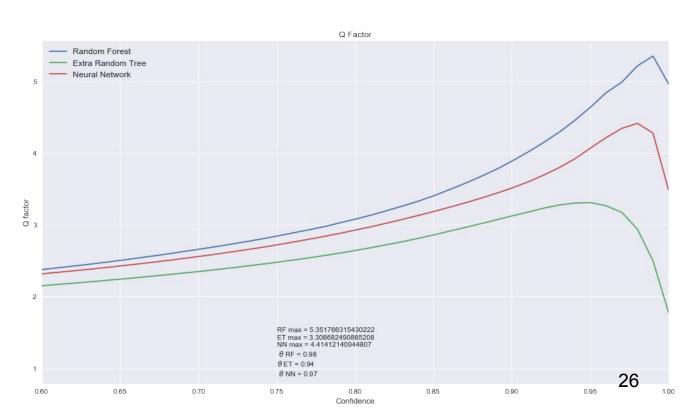
Random Forest leads





Q-factor

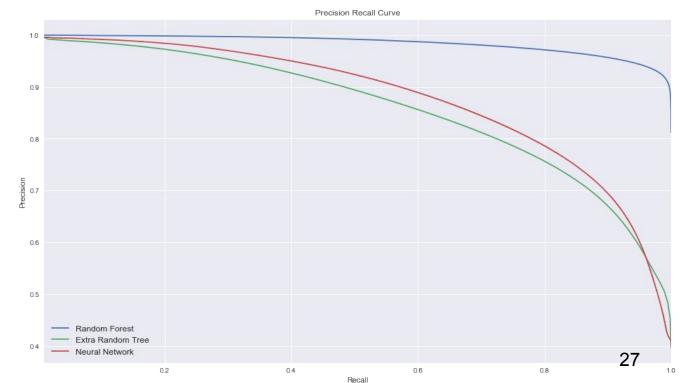






<u>Precision Recall curve</u>







Comments

Random Forest leads