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Please use	this identifier to cite or link to this item: http://hdl.handle.net/123456789/3474
Title:	AstroSat observations of the first galactic ULX pulsar swift J0243.6+6124
Authors:	Beri, A. (/jspui/browse?type=author&value=Beri%2C+A.) Singh, Kulinder Pal (/jspui/browse?type=author&value=Singh%2C+Kulinder+Pal)
Keywords:	Pulsars: individual: (Swift J0243.6+6124) X-rays: binaries
Issue Date:	2020
Publisher:	Oxford University Press
Citation:	Monthly Notices of the Royal Astronomical Society, 500(1), pp. 565-575.
Abstract:	Swift J0243.6+6124, the first Galactic ultraluminous X-ray pulsar, was observed during its 2017–2018 outburst with AstroSat at both sub- and super-Eddington levels of accretion with X-ray luminosities of LX \sim 7 \times 1037 and 6 \times 1038 erg s-1, respectively. Our broad-band timing and spectral observations show that X-ray pulsations at \sim 9.85 s have been detected up to 150 keV when the source was accreting at the super-Eddington level. The pulse profiles are a strong function of both energy and source luminosity, showing a double-peaked profile with pulse fraction increasing from \sim 10 per cent at 1.65 keV to 40–80 per cent at 70 keV. The continuum X-ray spectra are well modelled with a high-energy cut-off power law (Γ \sim 0.6–0.7) and one or two blackbody components with temperatures of \sim 0.35 and 1.2 keV, depending on the accretion level No iron line emission is observed at sub-Eddington level, while a broad emission feature at around 6.9 keV is observed at the super-Eddington level, along with a blackbody radius (121–142 km) that indicates the presence of optically thick outflows.
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