

In this problem set, the questions marked with a (\star) are required for students enrolled in 583 and bonus for students enrolled in 483.

1. Lorentz Transformation

- (a) (2 points) The coordinate transformation for an observer moving at a constant velocity V_x relative to a stationary observer is given by

$$\begin{aligned} t' &= \gamma_x(t - V_x x) \\ x' &= \gamma_x(x - V_x t) \\ y' &= y \\ z' &= z. \end{aligned} \tag{1}$$

How can we write this as a 4×4 Lorentz transformation matrix?

- (b) (5 points) More generally, suppose we have

$$\begin{aligned} t' &= \mu t + \nu x \\ x' &= \sigma t + \gamma x \\ y' &= \rho y \\ z' &= \lambda z. \end{aligned} \tag{3}$$

How can we write this as a 4×4 Lorentz transformation matrix?

2. Coordinate Transformations

Let

$$x = r \cos \theta, \quad y = r \sin \theta.$$

- (a) (2 points) Compute $\partial x / \partial r$, $\partial x / \partial \theta$, $\partial y / \partial r$, and $\partial y / \partial \theta$.
 (b) (2 points) Use the chain rule to express $\partial / \partial r$ in terms of $\partial / \partial x$ and $\partial / \partial y$.
 (c) (2 points) Interpret $\partial / \partial r$ geometrically.

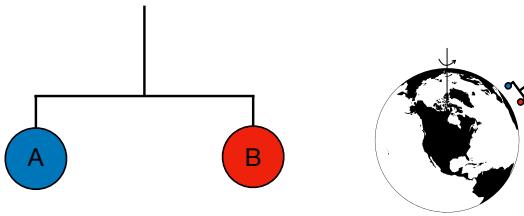
3. Differentials and Geometry

Given the scalar function $f(x, y) = x^2 y$:

- (a) (3 points) Compute the differential df .
 (b) (2 points) Interpret df geometrically.

4. (4 points) Change of Basis

Write the 2×2 rotation matrix $R(\theta)$ and show how vector components transform.



5. **Eötvos Experiment** Consider a torsion balance consisting of two equal inertial masses mounted at opposite ends of a light horizontal rod and suspended from a thin fiber. The apparatus is fixed to the surface of the rotating Earth at latitude θ . See figure for details.
- (3 points) Define a convenient coordinate system for analyzing the torsion balance. Using this coordinate system, state the condition that must be satisfied for the torsion balance to be in mechanical equilibrium.
 - (3 points) Determine the torque on the torsion balance arising from the Earth's rotation. Your answer should involve the Earth's angular velocity Ω , the latitude θ , and the relevant geometric parameters of the apparatus.
 - (2 points) What is the optimal latitude to perform this experiment? In other words, at what latitude would one expect to measure the largest potential torque?
 - (2 points) State the condition required for the weak equivalence principle to be satisfied for the two test masses in this experiment.
6. (10 points) **Thomas Precession** Suppose that students A and B start out in the lab frame. Student A remains in the lab frame. Student B gets in a rocket ship and changes their velocity by V_x in the x -direction (with $V_x \ll 1$) by firing the “+ X ” thrusters on their rocket. Then they change their velocity by V_y in the new y -direction (again with $V_y \ll 1$) by firing the “+ Y ” thrusters. Then they fire the “- X ” thrusters (changing their velocity by $-V_x$), and the “- Y ” thrusters (changing their velocity by $-V_y$). To lowest order in the V s, how does B 's reference frame now differ from A 's?
- Note: The problem is designed so that in Newtonian physics, B 's final reference frame is the same as A 's: they are at rest with respect to each other, with no rotation of the coordinate systems.*
7. (*) **The sky as viewed from a spaceship.** Let's suppose that observer \mathcal{O} remains on Earth in the lab frame. A second observer $\bar{\mathcal{O}}$ moves in a spaceship

at velocity $V = \tanh \alpha$ in the z -direction with respect to Earth. As you may recall from watching science-fiction movies, if V is large enough, $\bar{\mathcal{O}}$ sees the stars bunch up in front of them (the $+z$ direction). This problem works through the effect.

We suppose that the direction to the star makes an angle θ to the z -axis as seen from Earth, and $\bar{\theta}$ as seen from the spaceship. Without loss of generality, we will place the direction to the star at zero longitude (i.e., in the xz -plane).

- (a) (2 points) Show that in the Earth's frame, in time Δt , a photon from the star undergoes a displacement $\Delta x^\alpha = (\Delta t, -\Delta t \sin \theta, 0, -\Delta t \cos \theta)$.
- (b) (5 points) Apply a Lorentz transformation to find the photon's displacement in $\bar{\mathcal{O}}$'s frame. You may leave some results in terms of $\gamma = 1/\sqrt{1-V^2}$. Show that in the barred frame, the direction of the photon satisfies

$$\cos \bar{\theta} = \frac{V + \cos \theta}{1 + V \cos \theta}. \quad (8)$$

- (c) (3 points) Show that a star that appears on the “Equator” as seen from Earth ($\theta = \pi/2$) has an apparent position $\bar{\theta} = \cos^{-1} V$ as seen from the spaceship. How far from the North Pole does the star appear in the spaceship frame if $V = 0.9c$? What about $0.99c$?
- (d) (5 points) Now take the limit of small $\theta \ll 1$ (i.e., we will consider a constellation that contains the North Pole). Show that

$$\bar{\theta} \simeq \sqrt{\frac{1-V}{1+V}} \theta. \quad (10)$$

Hint: Take the Taylor expansion of your answer to (a) to 2nd order in θ . This means that the constellation containing the North Pole appears shrunk by a factor of $\sqrt{(1-V)/(1+V)}$ when seen from the spaceship.