

Inverse equation of state construction from mass-radius-relations of compact stars

Bachelor thesis

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Contents

1	Introduction	2
2	Preparations	2
2.1	The Tolman-Oppenheimer-Volkoff equation	2
2.2	Numerical solution	3
2.2.1	Method	3
2.2.2	Implementation	3

1 Introduction

To this day, the nature of the equation of state [EOS] for neutron stars is highly debated. Previously discussed EOS have emerged from various topics of physics, such as nuclear matter, quark matter and many others. Every one of these could potentially be modified to resemble other physical aspects or different conditions.

For the past years, it has been common to derive these EOS from a theory and then use it to construct first one compact star, then a whole series, determining a relation between total mass and radius. This method obviously is constrained by the fact that one cannot put an arbitrary number of theories and therefore variables into one EOS, as that would result in a problem impossible to solve, even for computers. This is the main reason to write this thesis, as we here will, in a way, revert the previous method.

For compact stars up to a certain mass, we will assume that we know the EOS well enough. Above that mass, we will numerically reconstruct an EOS from a mass-radius relation, and further compare the results with results from EOS dictated and constrained by theories.

2 Preparations

2.1 The Tolman-Oppenheimer-Volkoff equation

Before getting into the actual thesis, one has to set up a theoretical basis for the topic.

This involves a derivation of the equation used for determining the structure of compact stars. Here, that will be the Tolman-Oppenheimer-Volkoff [TOV] equation.

To make analytical and numerical calculations easier, one can choose the unit system to be $c = G = 1$, so that every unit is a power of length.

The derivation is based off the assumption that the star matter can be described as a perfect/ideal fluid. Further, the system shall not evolve in time, therefore staying spherically symmetric. In terms of the energy-momentum tensor we are left with:

$$T_{\mu\nu} = (\epsilon + P) u_\nu u_\mu - P g_{\mu\nu} \quad (1)$$

Spherical symmetry leads to a certain form of the metric, which then gives the stress-energy tensor components:

$$T_{\mu\nu} = \text{diag}(\epsilon e^{\nu(r)}, P e^{\lambda(r)}, P r^2, P r^2 \sin^2(\theta)) \quad (2)$$

Imposing hydrostatic equilibrium,

$$\nabla_\nu T^{\mu\nu} = 0 \quad (3)$$

and some calculation, one obtains the expression

$$\text{content}... \quad (4)$$

To determine what $\nu(r)$ looks like, and using the previously defined metric, one can calculate the non-zero components of the Ricci tensor:

$$\text{content}... \quad (5)$$

The Einstein equations then yield more equations containing $\nu(r)$ and it's derivatives. After plugging some of the equations into one another, one obtains $\nu(r)$ and therefore, in conclusion, the full TOV equation:

$$\frac{dP}{dr} = \frac{(\epsilon + P)(m + 4\pi r^3 P)}{2mr - r^2} \quad (6)$$

Together with a second equation for the mass:

$$\frac{dm}{dr} = 4\pi r^2 \epsilon \quad (7)$$

2.2 Numerical solution

2.2.1 Method

In order to generate an initial mass-radius relation to test the "inverse" algorithm with, one has to solve the TOV and mass differential equation numerically. By looking at both equations, our ordinary differential equation [ODE] system is given in the form

$$\dot{y}(t) = f(y(t), t) \quad (8)$$

where $\dot{y}(t)$ is a two component "vector":

$$\dot{y}(t) = \begin{pmatrix} dP/dr \\ dm/dr \end{pmatrix} \quad (9)$$

$f(y(t), t)$ then contains the right hand side of equations 6 and 7.

To solve the ODE system we use a fourth order Runge-Kutta [RK4] code. This method works as follows:

$$y(t + \tau) = y(t) + \frac{1}{6}(k_1 + 2k_2 + 2k_3 + k_4) \quad (10)$$

with k_i :

$$\begin{aligned} k_1 &= f(y(t), t) \cdot \tau \\ k_2 &= f(y(t) + k_1/2, t + \tau/2) \cdot \tau \\ k_3 &= f(y(t) + k_2/2, t + \tau/2) \cdot \tau \\ k_4 &= f(y(t) + k_3, t + \tau) \cdot \tau \end{aligned}$$

All k_i are of course also two component "vectors".

2.2.2 Implementation

In the actual program (written in C), the left hand side of our ODE system is implemented as a two-dimensional array of form $y[N][x]$. N will be two, for the whole program, as we will not add other equations to the system. The k_i are all one-dimensional arrays; that way only x is reflecting the step count. x will therefore be in range zero up to the number of iteration steps.

Messung der Fadenlänge	l (m)
1	
2	
3	
4	
5	
Mittelwert \bar{l}	
Standardabweichung σ_l	

References

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