

Astrostatistics: Tue 13 Feb 2017

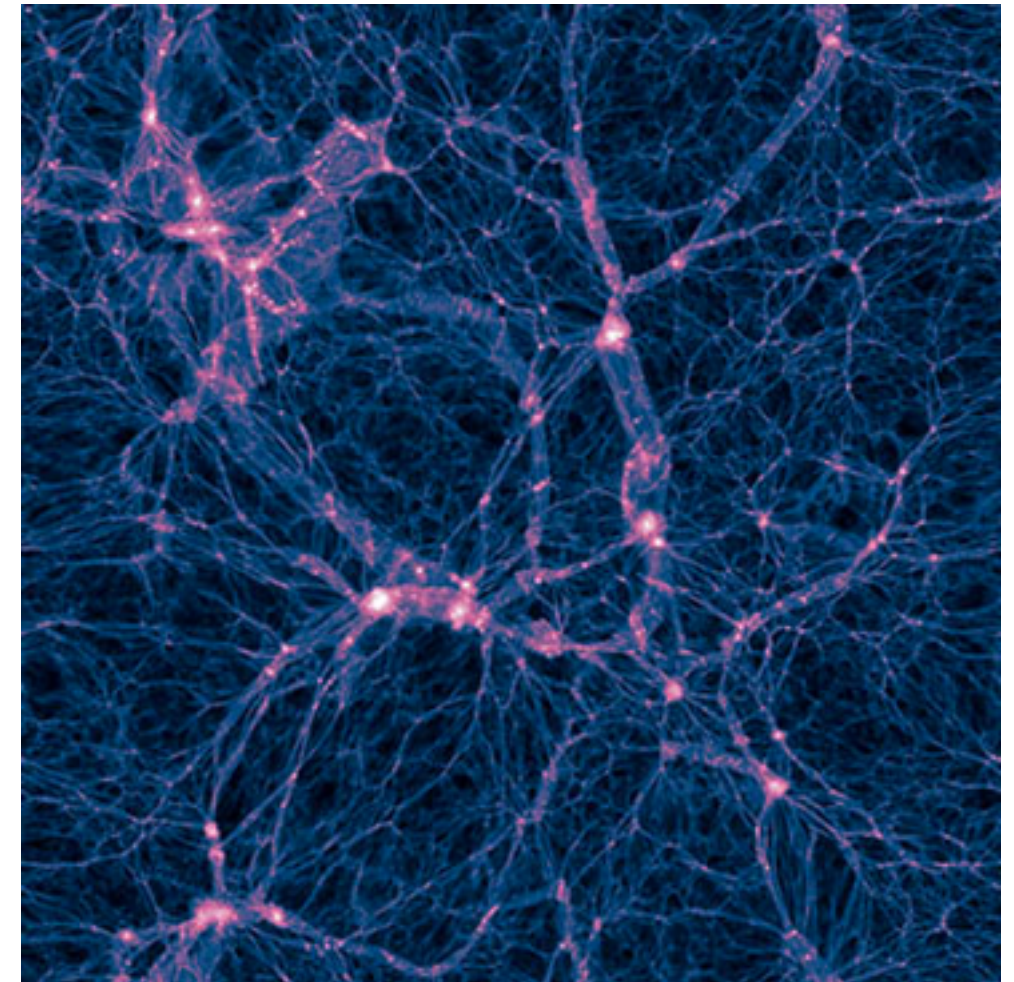
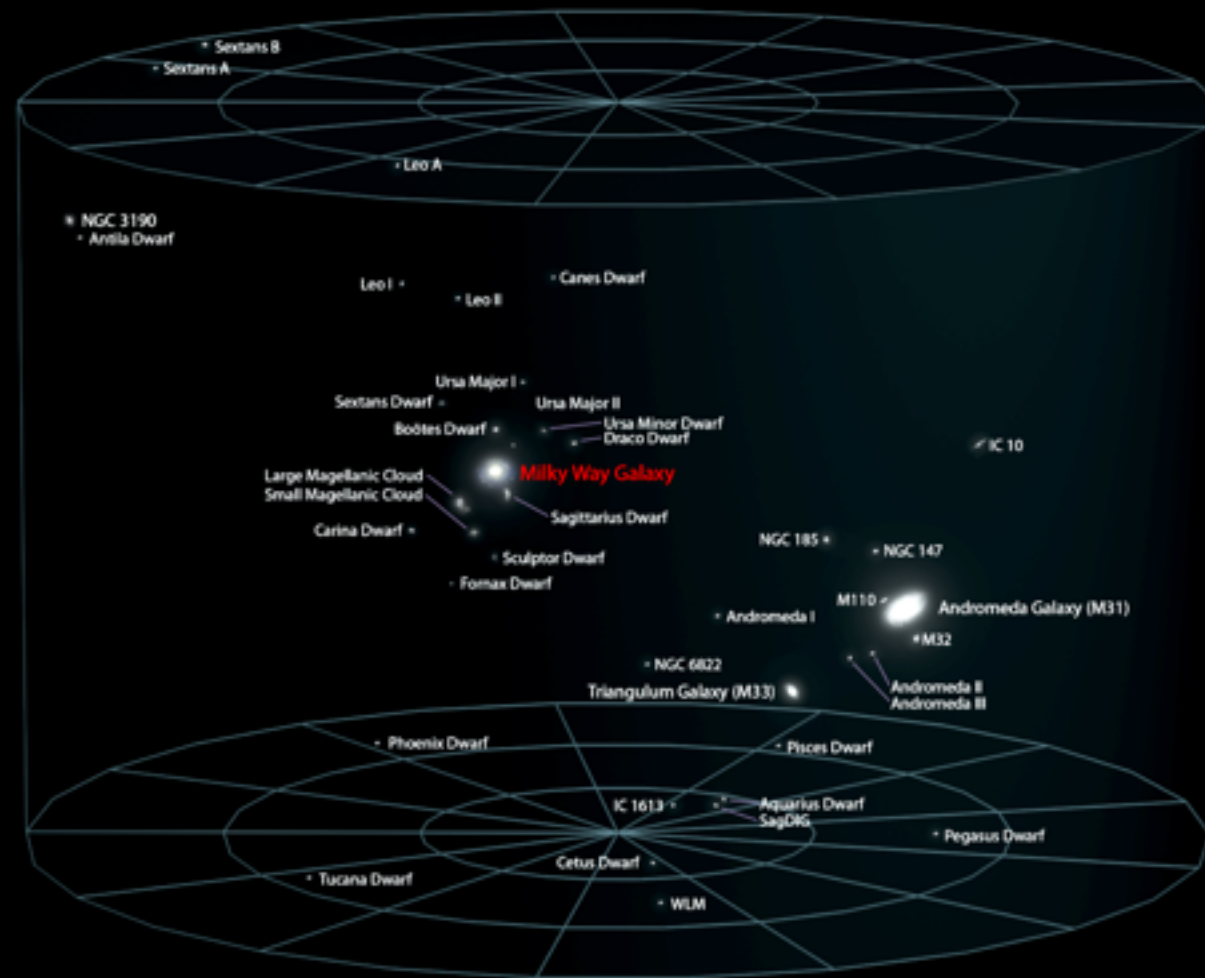
<https://github.com/CambridgeAstroStat/PartIII-Astrostatistics>

- Example Sheet 1 & datasets uploaded
 - Fri Feb 16 (2:30 pm, Room MR5)
- Fitting Statistical Models to Astronomical Data
 - Bayesian Inference —> Sampling, Computation, examples
 - References: Ivezić, Ch 5, F&B Ch 3, Gelman BDA
 - Hogg, D., 2012. “Data analysis recipes: Probability calculus for inference.” <https://arxiv.org/abs/1205.4446>
 - Hogg & DFM, 2017 “Data analysis recipes: Using Markov Chain Monte Carlo.” <https://arxiv.org/abs/1710.06068>

Next time: Astrostatistics Case Study 3:

Bayesian estimates of the Milky Way and Andromeda masses
using high-precision astrometry and cosmological simulations
(Patel et al. 2017, arXiv:1703.05767)

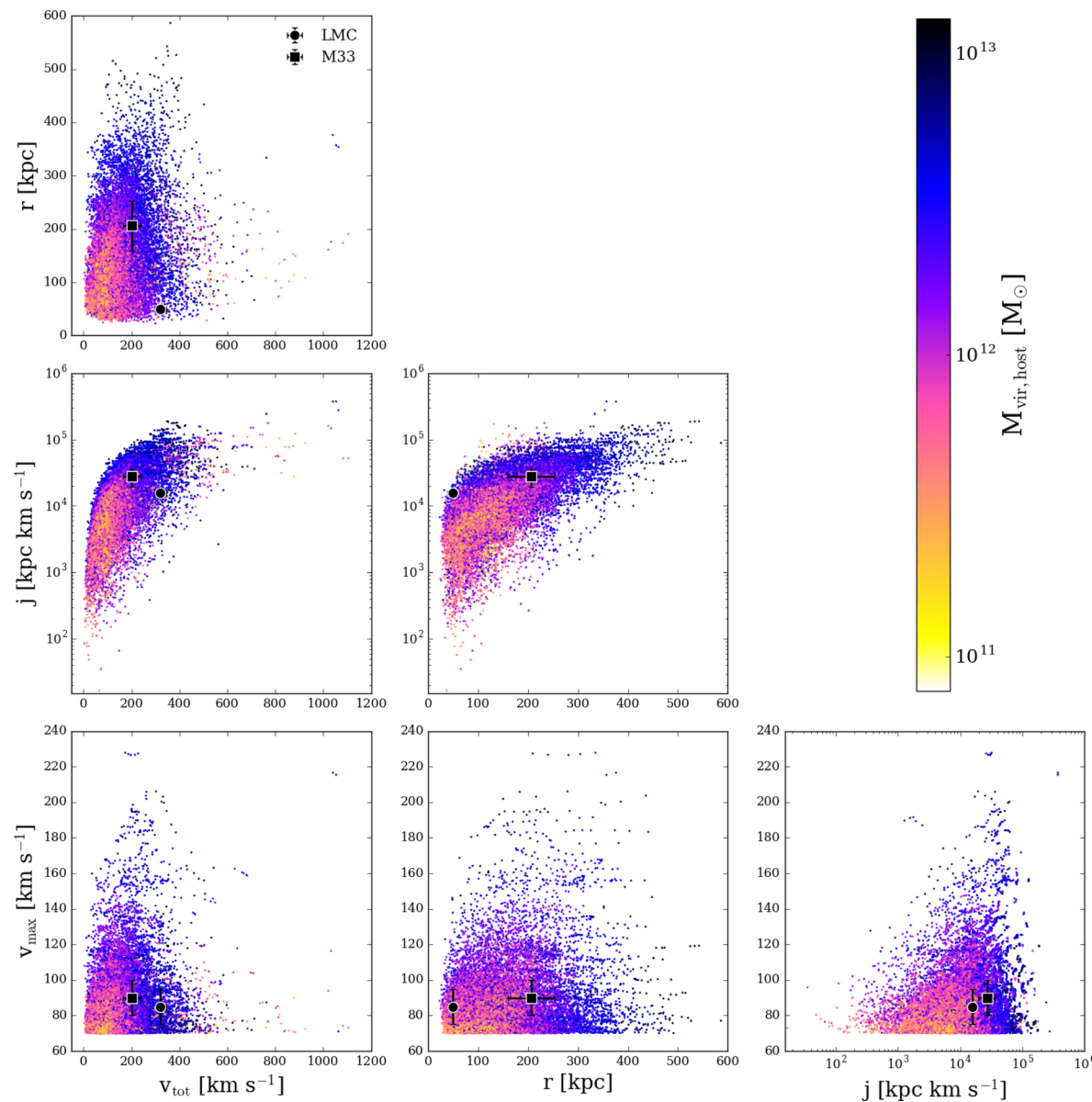
Local Galactic Group



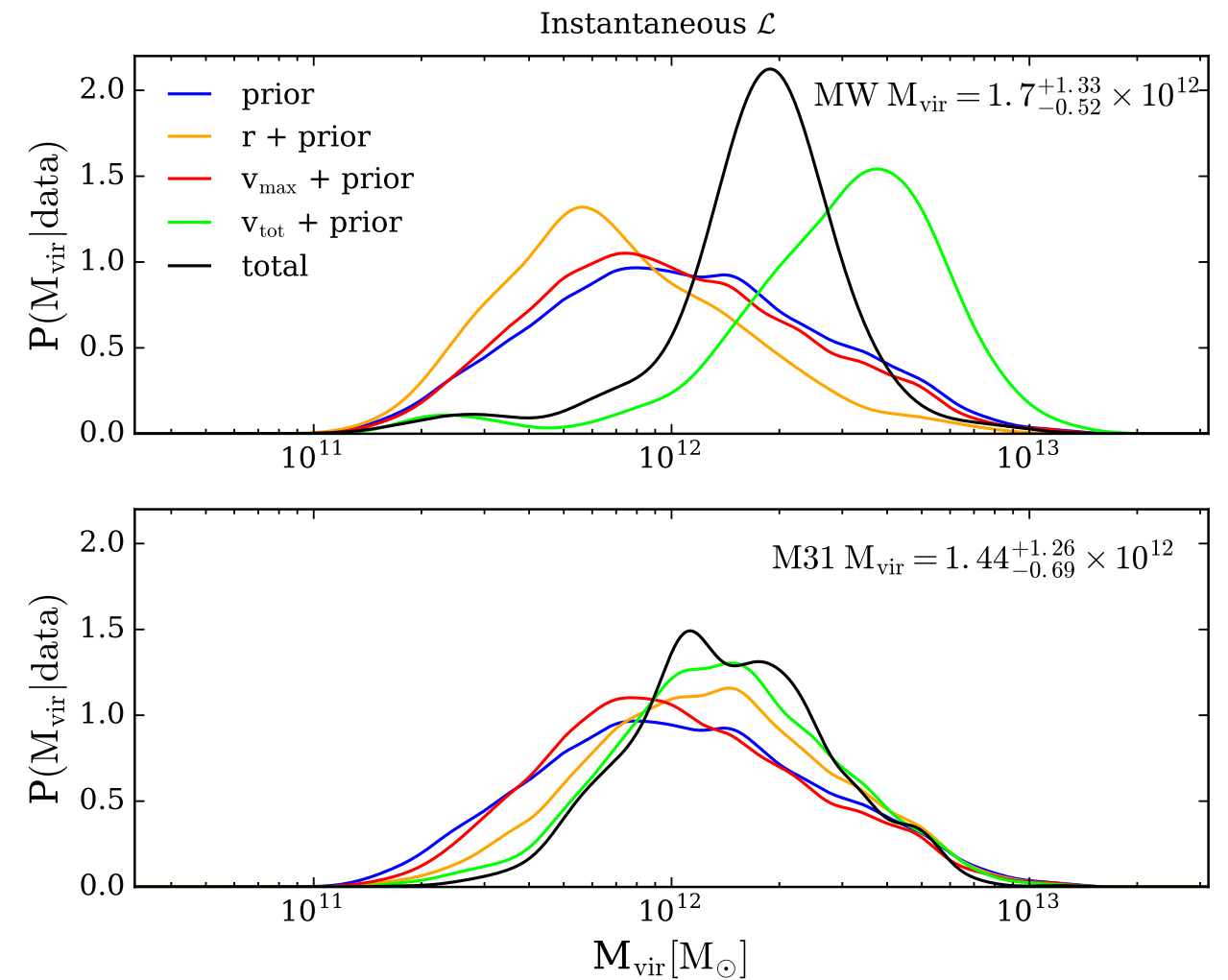
Illustris
Cosmological Simulation of
Galaxy Formation

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Simulation —> Prior

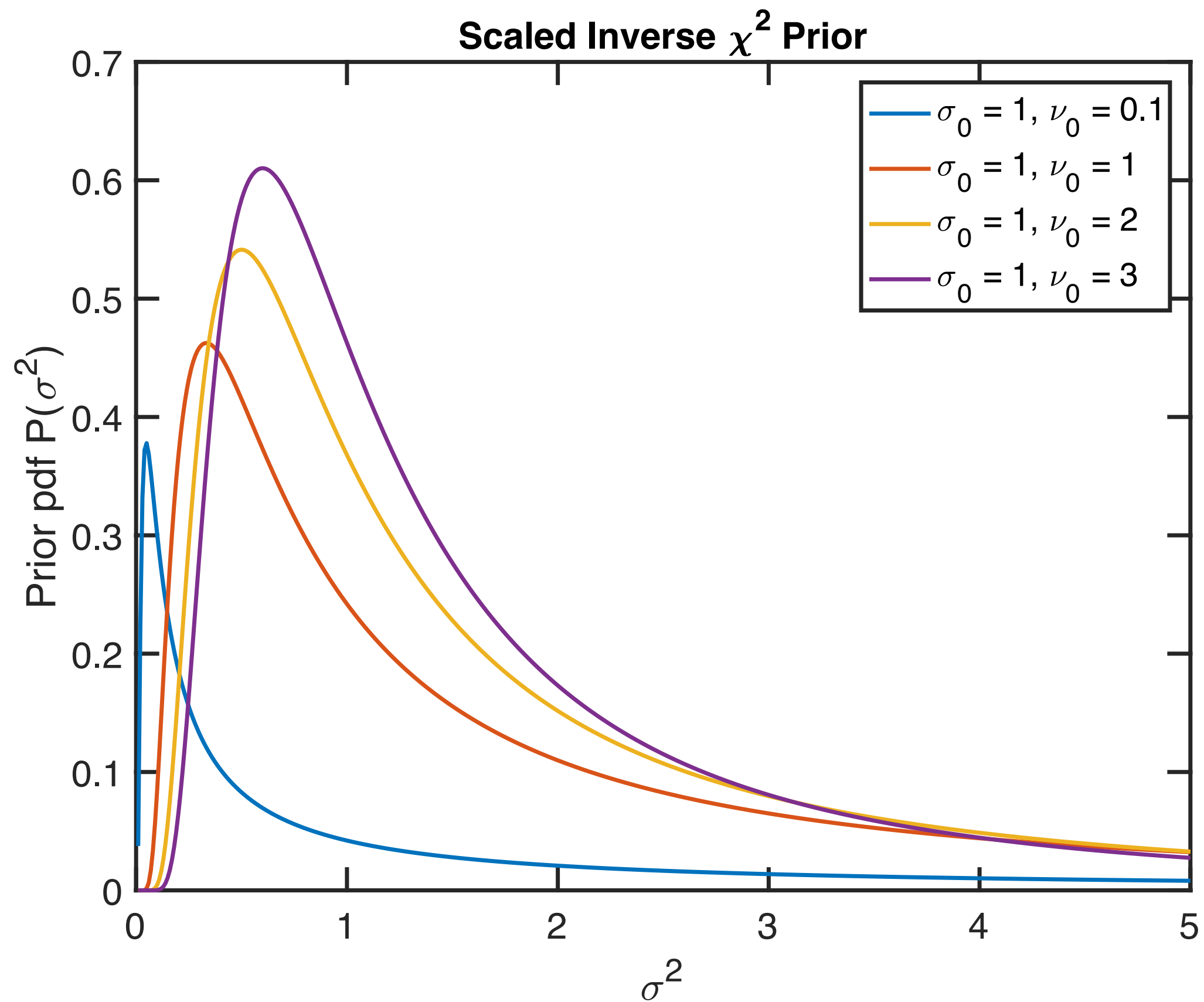


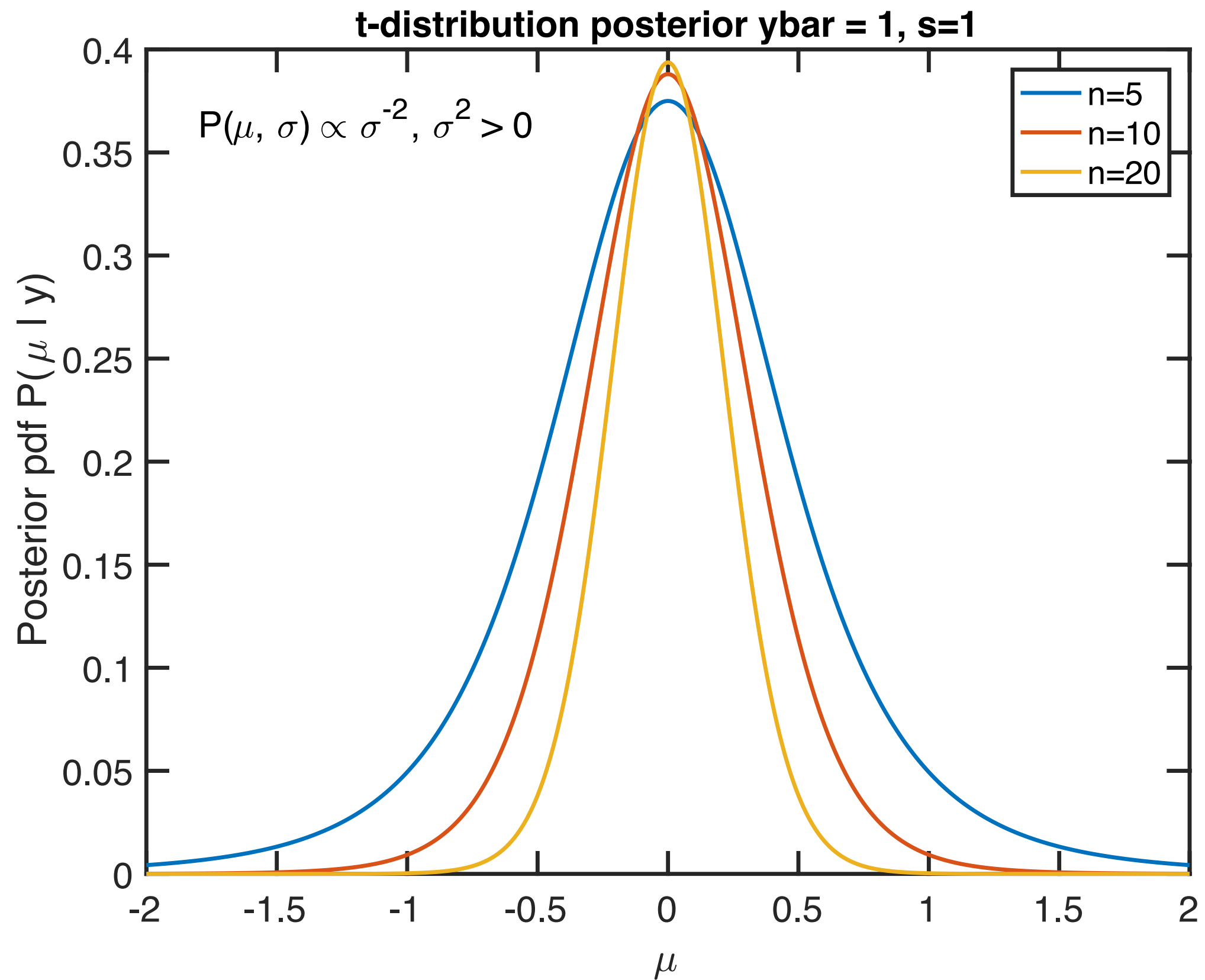
- Bayesian Inference
- Importance Sampling
- Kernel Density Estimation

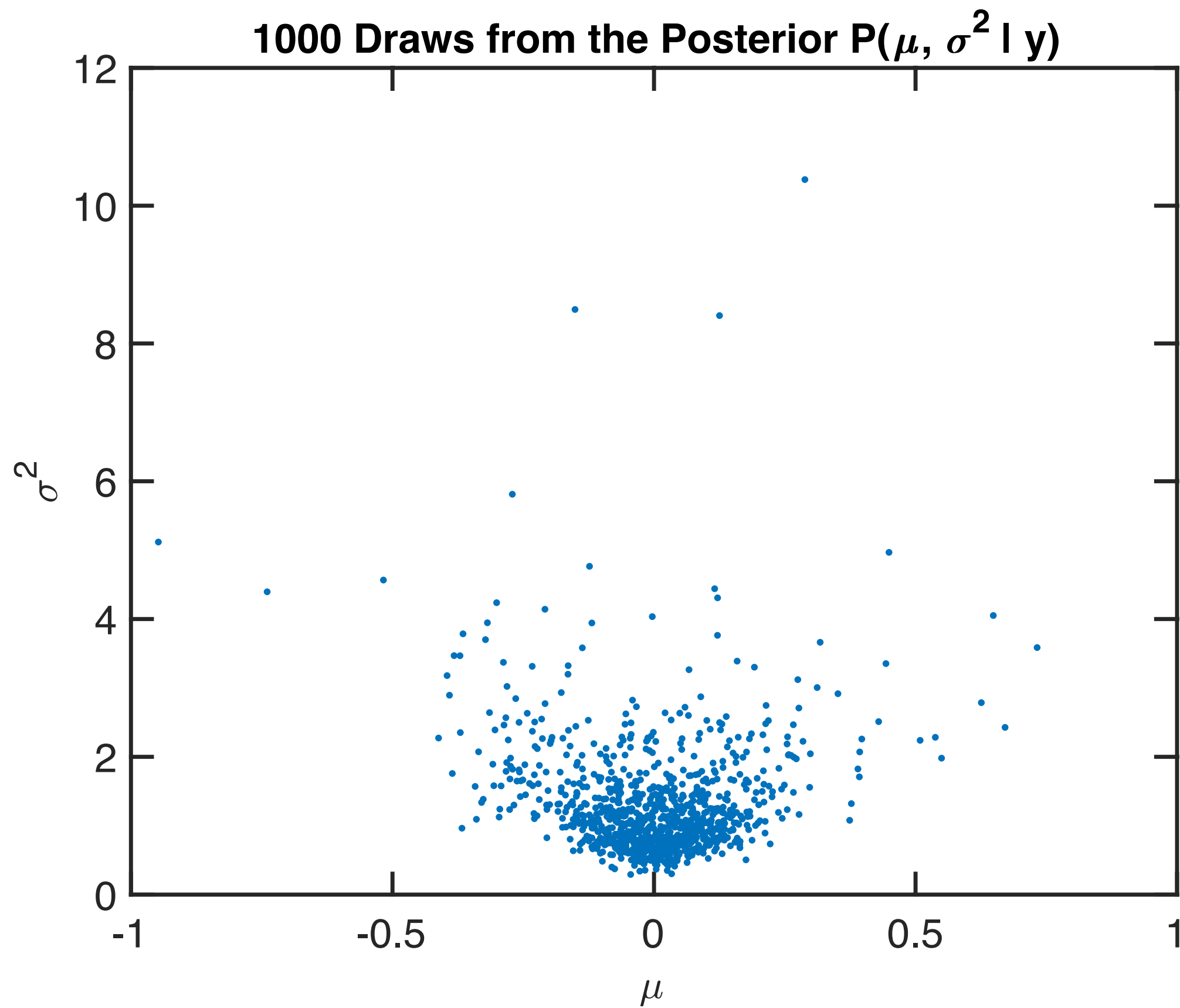
Multi-parameter Bayesian inference:

Gaussian example:

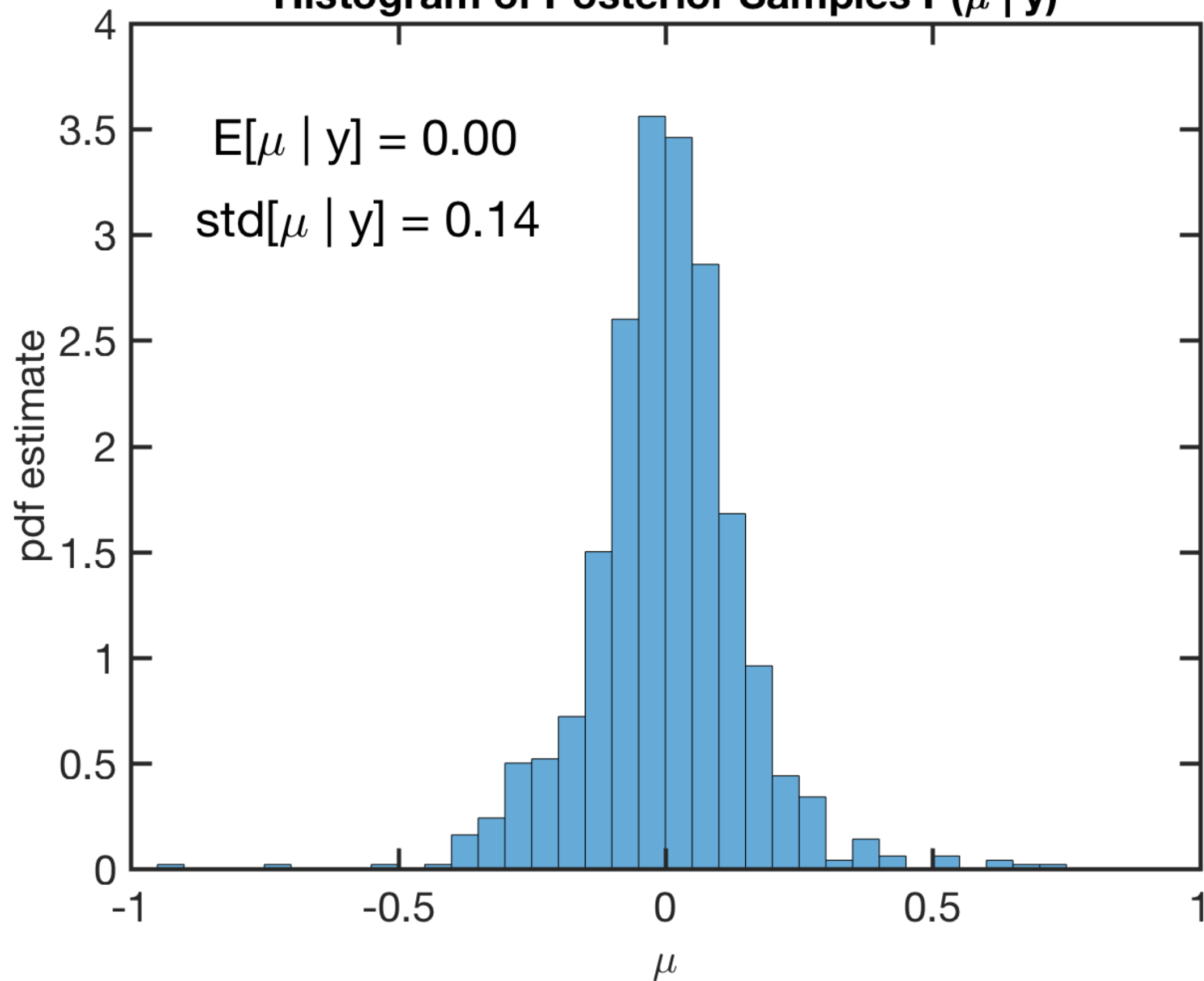
Gelman BDA Sec 3.2 - 3.3



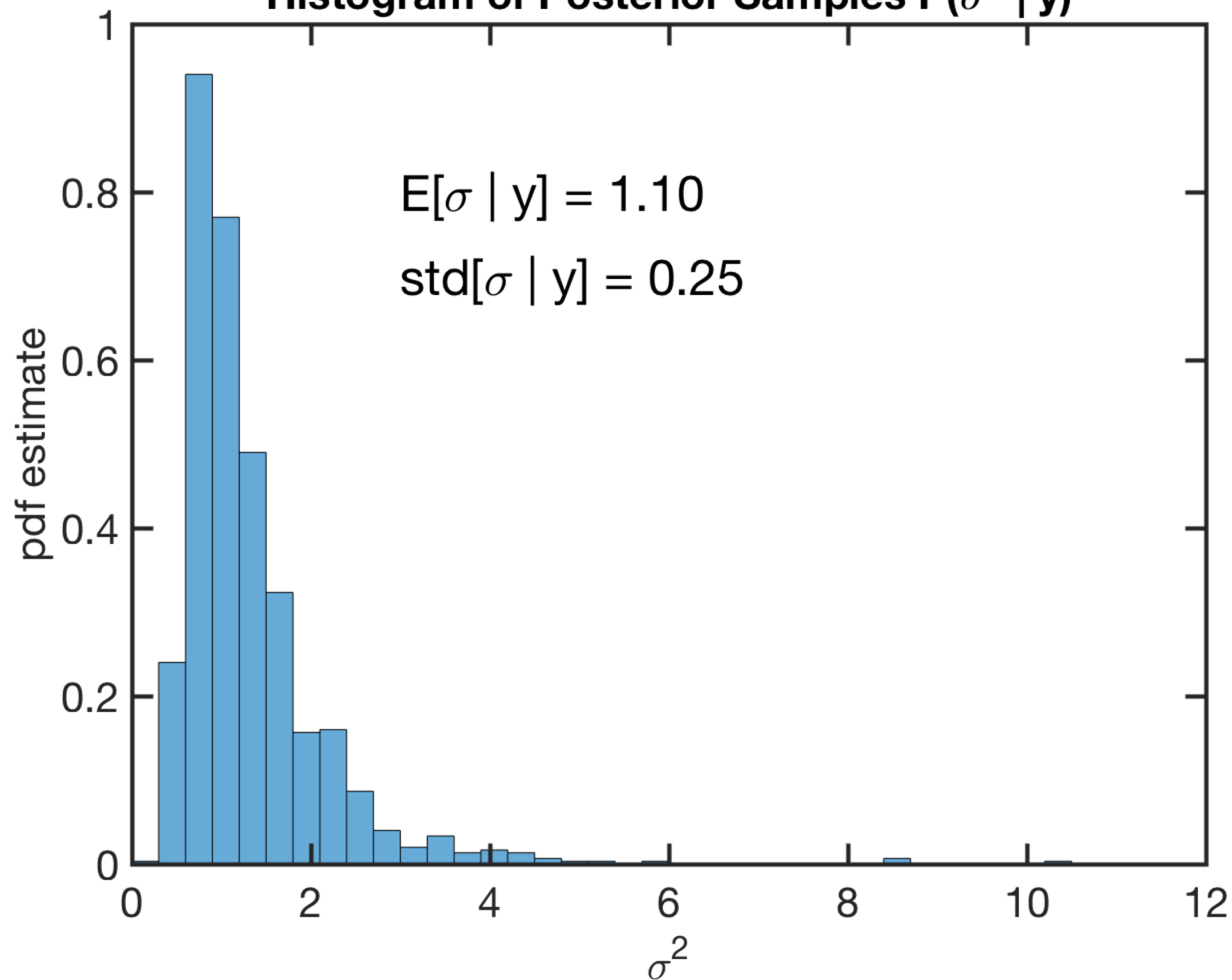




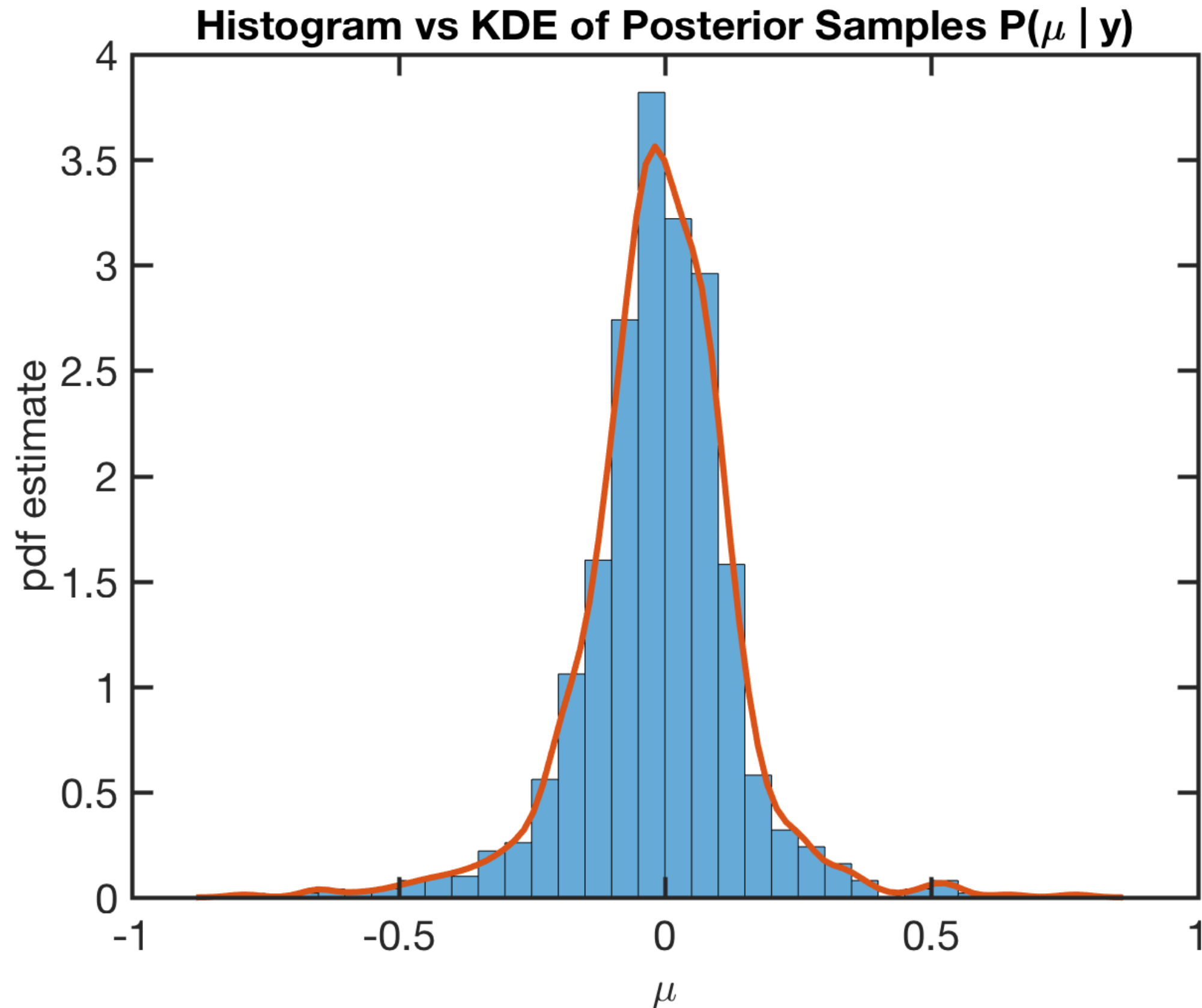
Histogram of Posterior Samples $P(\mu \mid y)$



Histogram of Posterior Samples $P(\sigma^2 | y)$



Kernel Density Estimate =
estimate a smooth density from samples



What if you can't directly sample
the posterior: $\theta_i \sim P(\theta | D)$?

$$\mathbb{E}[f(\boldsymbol{\theta}|D)] = \int f(\boldsymbol{\theta})P(\boldsymbol{\theta}|D)d\boldsymbol{\theta} \approx \sum_{i=1}^K f(\boldsymbol{\theta}_i)$$

- Posterior simulation - Markov Chain Monte Carlo, Nested Sampling, etc. generates draws
- Importance Sampling - draw from an easier (“tractable”) distribution $\theta_i \sim Q(\theta)$ and weight the samples by $w_i = P(\theta_i | D) / Q(\theta_i)$

$$\int f(\boldsymbol{\theta})P(\boldsymbol{\theta}|D)d\boldsymbol{\theta} = \int f(\boldsymbol{\theta})\frac{P(\boldsymbol{\theta}|D)}{Q(\boldsymbol{\theta})}Q(\boldsymbol{\theta})d\boldsymbol{\theta} \approx \sum_{i=1}^K f(\boldsymbol{\theta}_i)w_i$$