Astrostatistics: Tue 05 Feb 2017

https://github.com/CambridgeAstroStat/PartIII-Astrostatistics

- Examples Classes (sheets provided ~1 week prior)
 - Fri Feb 16, Fri Mar 2, Wed Mar 14 (1pm, RoomTBD)
 - One more + Revision Class in Easter Term
- Fitting Statistical Models to Astronomical Data
 - Generative / Latent Variable Modeling / Bayes
 - Hogg, Bovy & Lang. "Data analysis recipes: Fitting a model to data". https://arxiv.org/abs/1008.4686

Statistical Modelling Wisdom

- Have an objective function [e.g. Likelihood or posterior] that you
 optimise or sample to fit the data not just a procedure/recipe
- Objective function helps you evaluate relative fits of data with under different parameter values / models
- Derive your objective function from your modelling assumptions (physical or statistical)
- Write down your assumptions!
- First question: what is the likelihood $L(\theta)$? Derive it from the assumptions underlying your sampling distribution $P(D \mid \theta)$!
- Second question: what is your prior $P(\theta)$? (if Bayesian)
- Third question: How do I optimise/sample objective function to fit the data?

Fitting Models to Astro Data

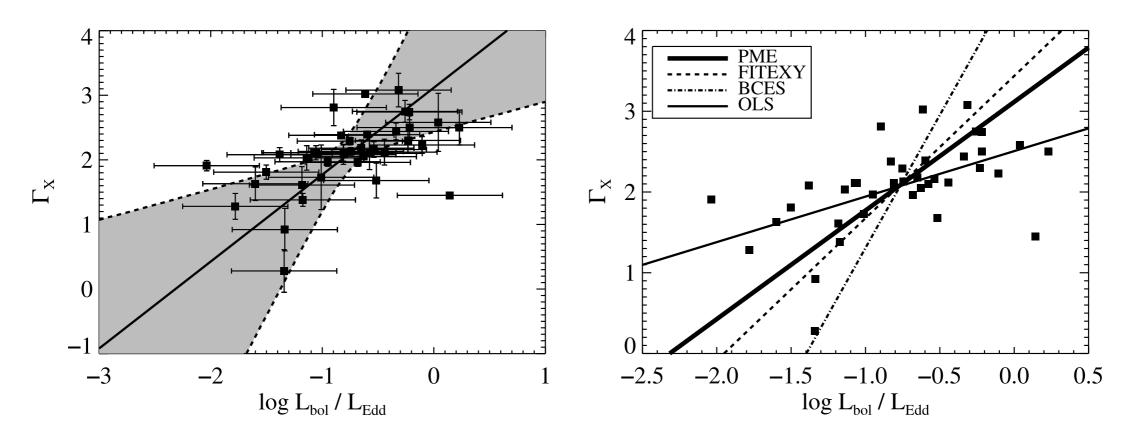


Fig. 10.—X-ray photon index Γ_X as a function of $\log L_{\rm bol}/L_{\rm Edd}$ for $39~z \lesssim 0.8$ radio-quiet quasars. In both plots, the thick solid line shows the posterior median estimate (PME) of the regression line. In the left panel, the shaded region denotes the 95% (2 σ) pointwise confidence intervals on the regression line. In the right panel, the thin solid line shows the OLS estimate, the dashed line shows the FITEXY estimate, and the dot-dashed line shows the BCES(Y|X) estimate; the error bars have been omitted for clarity. A significant positive trend is implied by the data.

Modelling heteroskedastic, correlated measurement errors in both y and x, intrinsic scatter, nondetections, selection effects

B. Kelly et al. 2007, "Some Aspects of Measurement Error in Linear Regression of Astronomical Data." ApJ, 665, 1489

Ad-hoc " χ^2 " approaches vs. Likelihood formulation

FITEXY Estimator

• Press et al.(1992, *Numerical Recipes*) define an 'effective χ^2 ' statistic:

$$\chi_{EXY}^{2} = \sum_{i=1}^{n} \frac{(y_{i} - \alpha - \beta x_{i})^{2}}{\sigma_{y,i}^{2} + \beta^{2} \sigma_{x,i}^{2}}$$

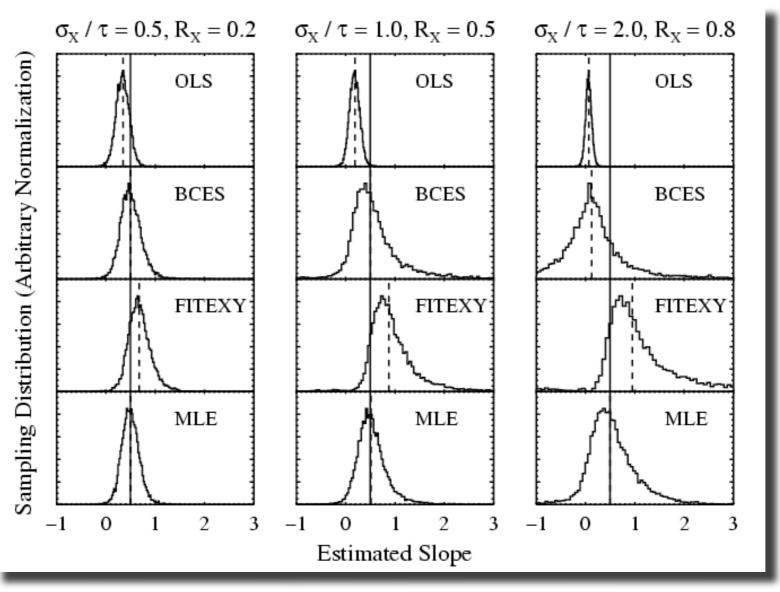
- Choose values of α and β that minimize χ^2_{EXY}
- Modified by Tremaine et al.(2002, ApJ, 574, 740), to account for intrinsic scatter:

$$\chi_{EXY}^{2} = \sum_{i=1}^{n} \frac{(y_{i} - \alpha - \beta x_{i})^{2}}{\sigma^{2} + \sigma_{y,i}^{2} + \beta^{2} \sigma_{x,i}^{2}}$$

http://astrostatistics.psu.edu/su07/kelley_measerr07.pdf

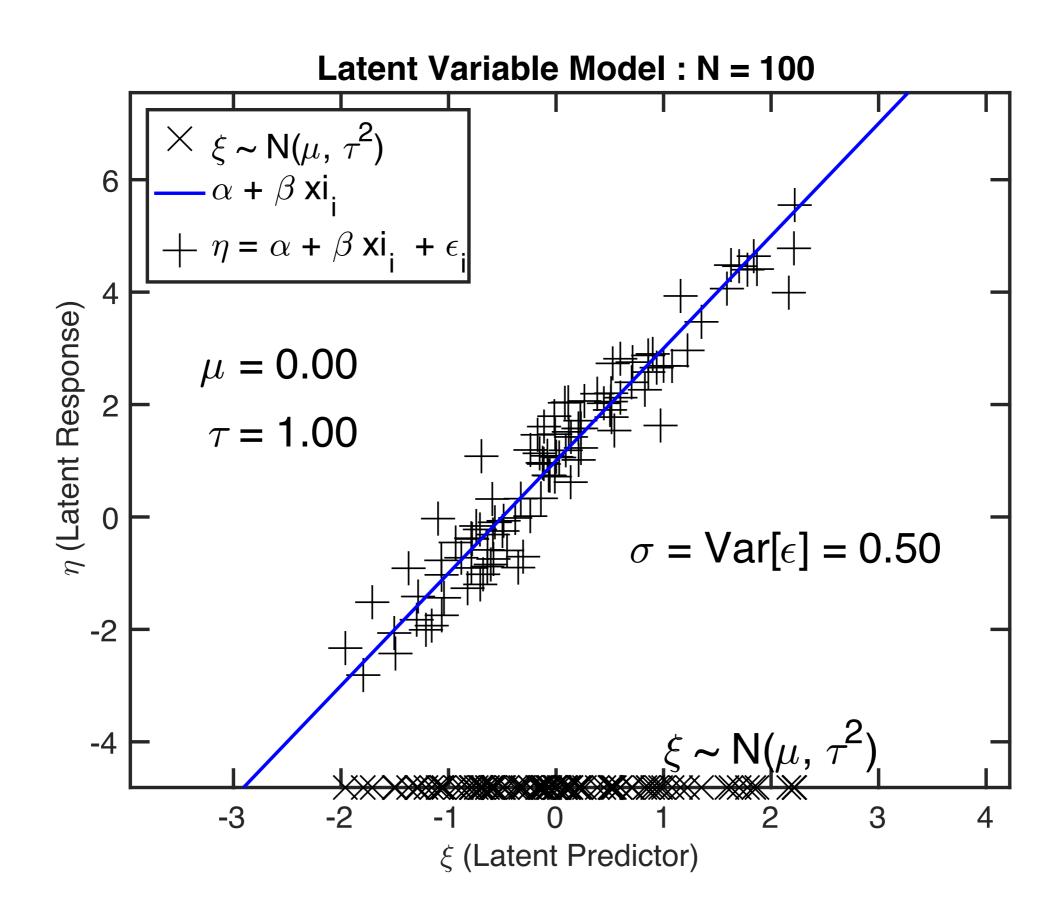
Kelly et al. 2017, Latent Variable Likelihood approach vs. Bad

Simulation Study: Slope



Dashed lines mark the median value of the estimator, solid lines mark the true value of the slope. Each simulated data set had 50 data points, and y-measurement errors of $\sigma_v \sim \sigma$.

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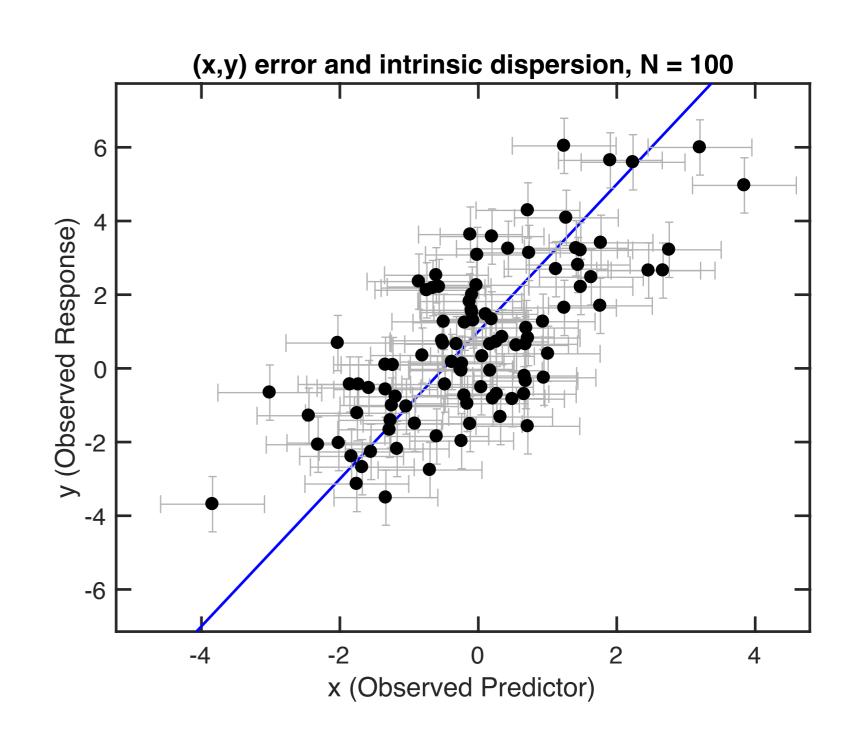


Probabilistic Generative Modelling

- Forward Model comprises series of probabilistic steps describing conceptually how the observed data was generated from the parameters of interest
- Can introduce intermediate parameters / unobserved latent variables α (e.g. true values corresponding to the observed data).
- From Forward model, derive the sampling distribution, e.g. $P(D \mid \theta) = \int P(D \mid \alpha) P(\alpha \mid \theta) d\alpha$
- Using observed data D, draw inference from Likelihood function: $L(\theta) = P(D \mid \theta)$
- Or if Bayesian with prior P(θ): sample posterior:
 P(θ | D) = P(D | θ) P(θ)

Example: Structural Model for Linear Regression (B. Kelly et al. 2007, "Some Aspects of Measurement Error in Linear Regression of Astronomical Data." ApJ, 665, 1489)

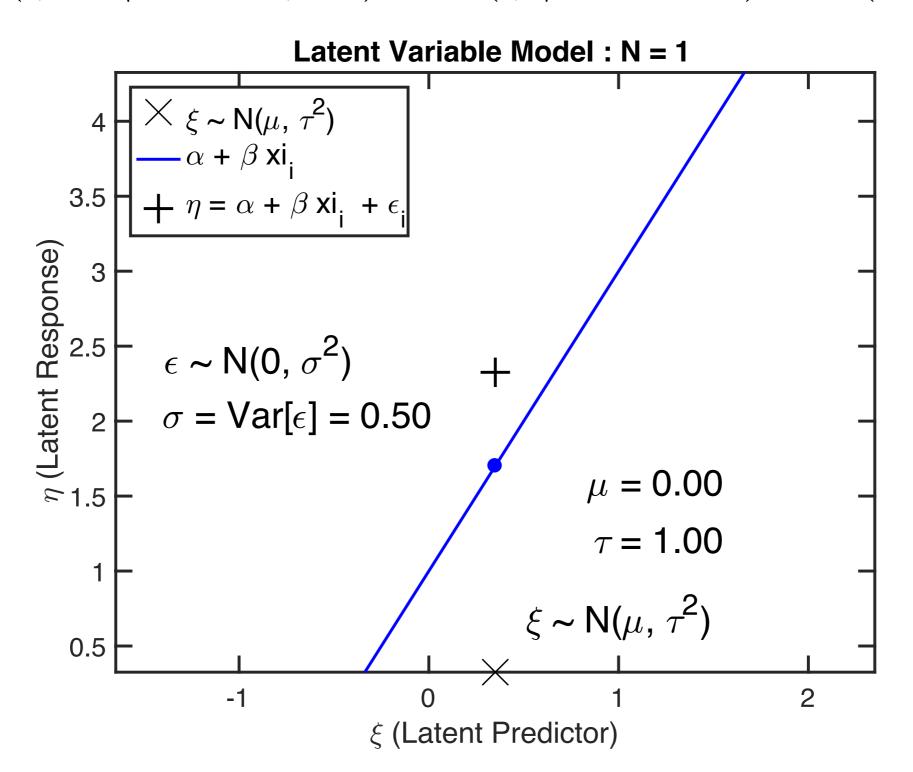
- Observed data has x and y meas. errors and intrinsic dispersion
- Estimate the true slope (and other parameters)



Step 1:

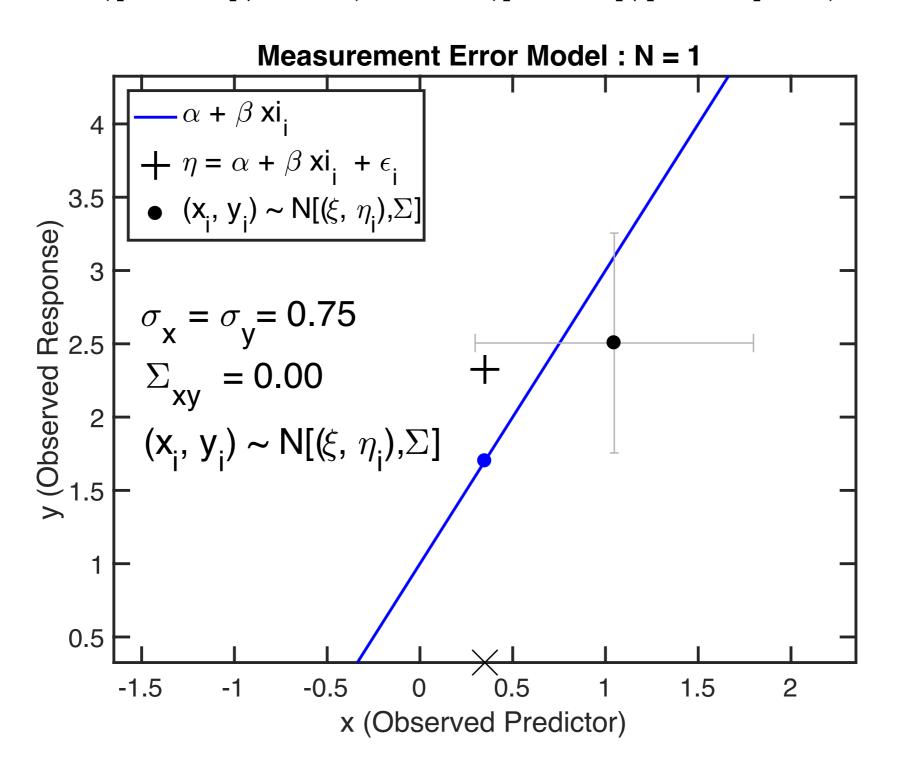
Generating Latent Variables from Parameters:

$$P(\eta_i, \xi_i | \alpha, \beta, \sigma, \mu, \tau) = P(\eta_i | \xi_i, \alpha, \beta, \sigma) \times P(\xi_i | \mu, \tau)$$

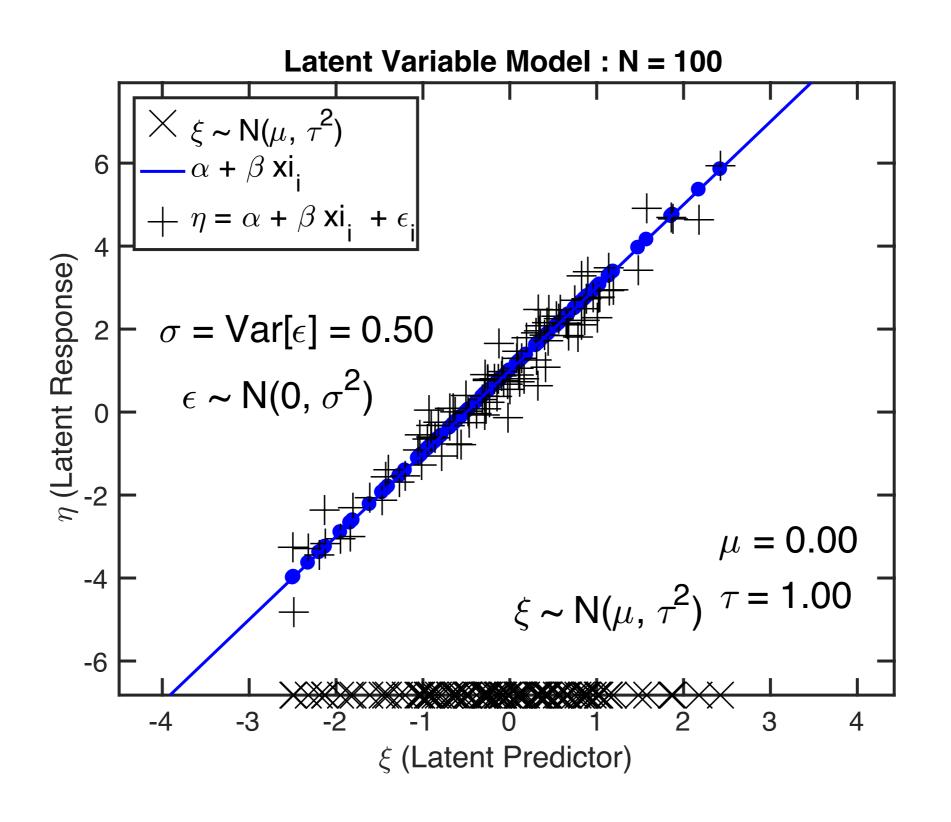


Step 2: Generating Observed Data from Latent Variables

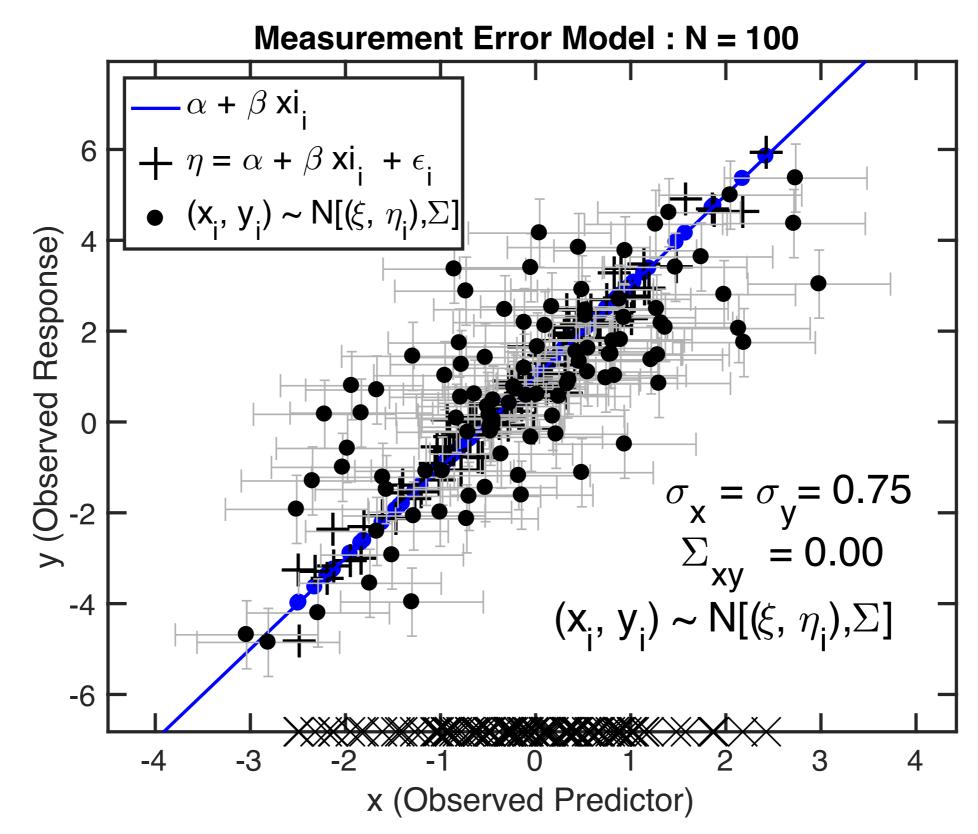
$$P([x_i, y_i]|\eta_i, \xi_i) = N([x_i, y_i]|[\eta_i, \xi_i], \Sigma)$$



Now repeat for N=100 objects



Now repeat for N=100 objects



Knowns and Unknowns

Regression Parameters

Independent Variable Population Distribution "Hyperparameters"

Latent (true) Variables

Observed Data

$$\boldsymbol{\theta} = (\alpha, \beta, \sigma^2)$$

$$\psi = (\mu, \tau)$$

$$(\xi_i,\eta_i)$$

$$(x_i,y_i)$$

Generative Model

Population Distribution

$$\xi \sim N(\mu | \tau^2)$$

Regression

$$|\eta_i| \, \xi_i \sim N(\alpha + \beta x_i, \sigma^2)$$

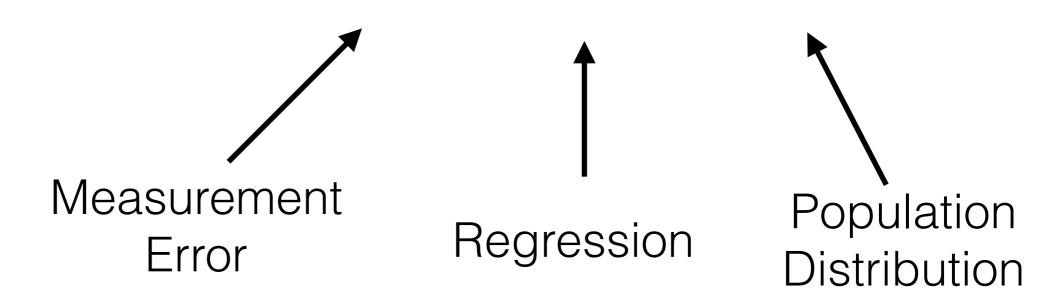
Measurement Error

$$[x_i, y_i] | \xi_i, \eta_i \sim N([\xi_i, \eta_i], \Sigma])$$

Formulating Likelihood Function: Marginalising (integrating out) latent variables

$$P(x_i, y_i | \boldsymbol{\theta}, \boldsymbol{\psi}) = \int \int P(x_i, y_i, \xi_i, \eta_i | \boldsymbol{\theta}, \boldsymbol{\psi}) d\xi_i d\eta$$

$$P(x_i, y_i | \boldsymbol{\theta}, \boldsymbol{\psi}) = \int \int P(x_i, y_i | \xi_i, \eta_i) P(\eta_i | \xi_i, \boldsymbol{\theta}) P(\xi_i | \boldsymbol{\psi}) d\xi_i d\eta$$



Solution: (Kelly 2007, Eqs. 16-23)

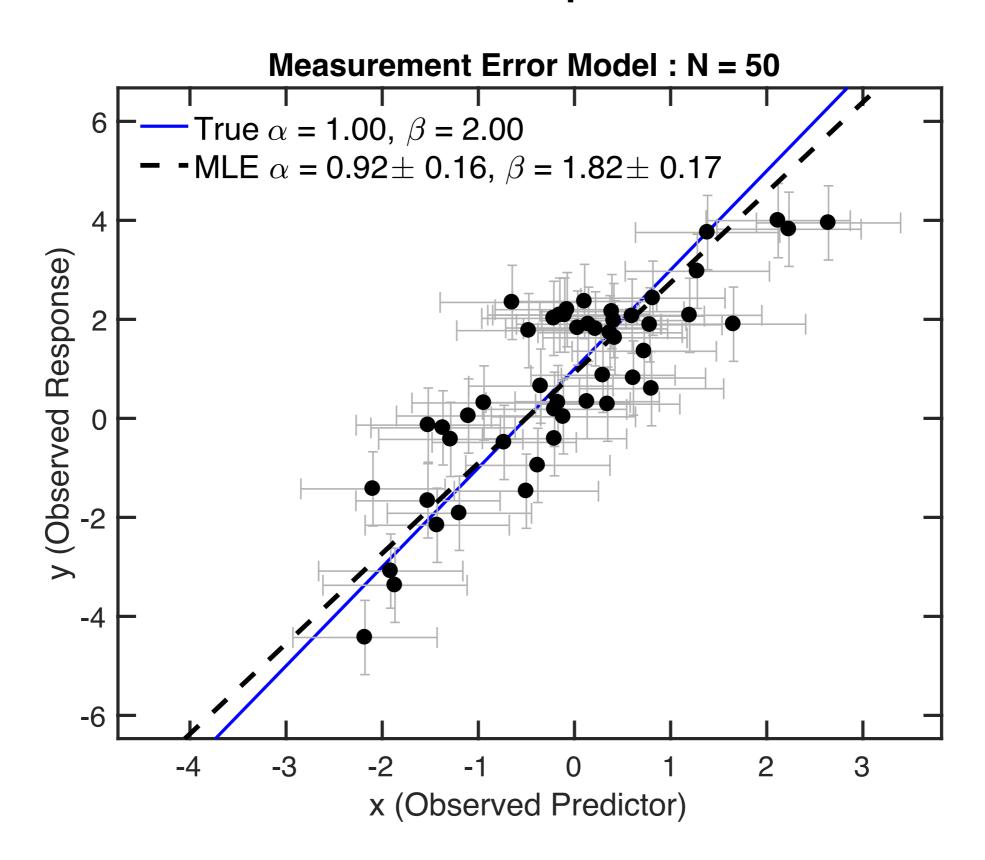
Gave More General Solution when $P(\xi|\Psi)$ is a Mixture of Gaussians (set K=1, π_1 = 1 for us)

$$p(\boldsymbol{x}, \boldsymbol{y}|\boldsymbol{\theta}, \boldsymbol{\psi}) = \prod_{i=1}^{n} \sum_{k=1}^{K} \frac{\pi_{k}}{2\pi |\mathbf{V}_{k,i}|^{1/2}} \times \exp\left[-\frac{1}{2}(z_{i} - \boldsymbol{\zeta}_{k})^{T} \mathbf{V}_{k,i}^{-1}(z_{i} - \boldsymbol{\zeta}_{k})\right], \quad (16)$$

$$\boldsymbol{\zeta}_{k} = (\alpha + \beta \mu_{k}, \mu_{k}), \quad (17)$$

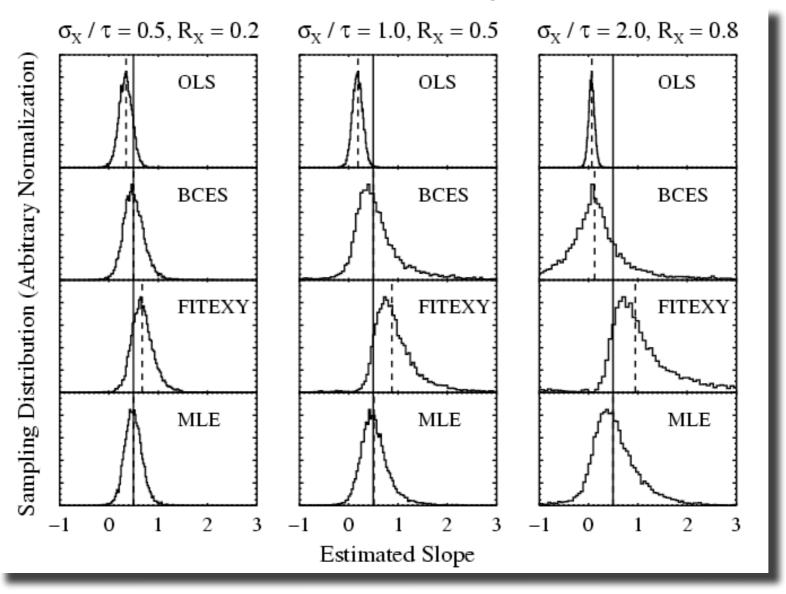
$$\mathbf{V}_{k,i} = \begin{pmatrix} \beta^{2} \tau_{k}^{2} + \sigma^{2} + \sigma_{y,i}^{2} & \beta \tau_{k}^{2} + \sigma_{xy,i} \\ \beta \tau_{k}^{2} + \sigma_{xy,i} & \tau_{k}^{2} + \sigma_{x,i}^{2} \end{pmatrix}, \quad (18)$$

Example



Kelly et al. 2017, Latent Variable Likelihood approach vs. Bad

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