

Nonlinear Dynamics and Chaos

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Part I

Nonlinear Dynamics and Chaos 1

Chapter 1

Introduction

First we shall introduce the most important characters in our following exploration. The ideas and definitions here will be recurring regularly as we examine them from different perspectives and using different tools. The content covered by this course can be found in the following books. For further details on some of the results, we recommend consulting these.

- J. Guckenheimer & P. Holmes, Nonlinear Oscillations, Dynamical Systems and Bifurcations of Vector Fields [[Guckenheimer and Holmes, 1990](#)];
- F. Verhulst, Nonlinear Differential Equations and Dynamical Systems [[Verhulst, 1989](#)];
- V. I. Arnold, Ordinary Differential Equations [[Arnold, 1992](#)];
- S. Strogatz, Nonlinear Dynamics and Chaos [[Strogatz, 2000](#)].

Definition 1.1 (Dynamical System (DS)). A triple (P, E, \mathcal{F}) , with

- P : the phase space for the dynamical variable $x \in P$,
- E : base space of the evolutionary variable (e.g. time) $t \in E$,
- \mathcal{F} : the evolution rule (deterministic) which defines the transition from one state to the next.

The two main types of evolutionary variable spaces are

- (i) Discrete dynamical systems (DDS) $t \in E = \mathbb{Z}$ with trajectory $\{x_0, x_1, \dots\}$,
- (ii) Continuous dynamical systems (CDS) $t \in E = \mathbb{R}$ with trajectory $\{x_t\}_{t \in \mathbb{R}}$.

Corresponding to these there are various types of evolution rules

(i) In a DDS we have iterated mappings

$$x_{n+1} = F(x_n, n). \quad (1.1)$$

If there is no explicit dependence on n , i.e. $\frac{\partial F}{\partial n} = 0$, then

$$x_{n+1} = F(x_n) = F(F(x_{n-1})) = \underbrace{F \circ \dots \circ F}_{n+1 \text{ times}}(x_0) = F^{n+1}(x_0). \quad (1.2)$$

Notebook 1.1. *Construction of the cobweb diagram*

Example 1.1 (Cobweb diagram of a one-dimensional DDS). In such cases and in one-dimensional problems, a simple way to analyze the behavior of the system is the so-called *cobweb* diagram. We may plot x_{n+1} as a function of x_n , as demonstrated in Fig. 1.1. The image of an initial condition x_0 lies on the graph at $x_{n+1} = F(x_0)$. We can also compute the next iterate by horizontally projecting the point $(x_0, F(x_0))$ to the diagonal line defined by $x_{n+1} = x_n$. Following the projection of this point to the horizontal axis (x_n) we find the intersection with the graph at the point $(x_1, F(x_1))$. It follows that fixed points on the cobweb diagram correspond to the intersection of the graph of F with the diagonal line $x_{n+1} = x_n$.

(ii) In a CDS we have a first order system of ordinary differential equations (ODE)

$$\dot{x} = f(x, t) \quad (1.3)$$

for $x \in P$ and $t \in E$. This yields the initial value problem (IVP):

$$\begin{cases} \dot{x} = f(x, t) \\ x(t_0) = x_0 \end{cases} \quad (1.4)$$

Assuming there exists a unique solution $\varphi(t; t_0, x_0)$ with $\dot{\varphi} = f(\varphi, t)$ and $\varphi(t_0) = x_0$, then the following flow map is well defined

$$F_{t_0}^t(x_0) := \varphi(t; t_0, x_0). \quad (1.5)$$

Geometrically, this solution can be viewed as a trajectory in phase space (cf. Fig. 1.2).

Such an $F_{t_0}^t$ has the properties

(a) $F_{t_0}^t$ is as smooth as $f(x, t)$,

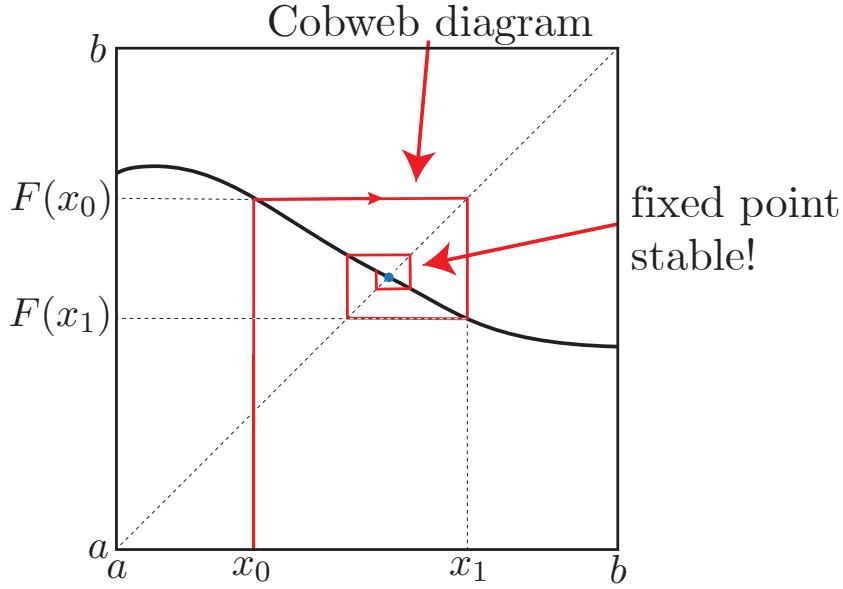


Figure 1.1: Analysis of a one-dimensional system defined on the interval $x \in [a, b]$ using the cobweb diagram

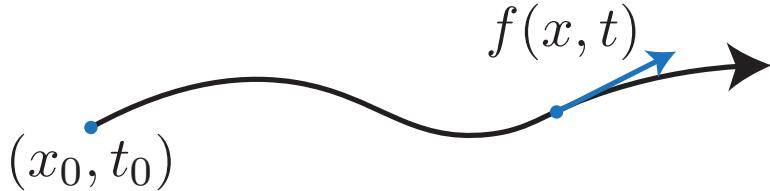


Figure 1.2: Trajectory of a continuous dynamical system. The RHS is given by $f(x, t)$, which is the tangent vector to this curve at the point x at time t .

- (b) $F_{t_0}^{t_0} = I$ and $F_{t_0}^{t_2} = F_{t_2}^{t_2} \circ F_{t_0}^{t_1}$,
- (c) $(F_{t_0}^t)^{-1} = F_t^{t_0}$ exists and is smooth.

Properties (b) and (c) together are called the group property. A special case of continuous dynamical systems is the autonomous system.

$$\boxed{\dot{x} = f(x).} \quad (1.6)$$

The autonomy of a system implies

$$x(s, t_0, x_0) = x(\underbrace{s - t_0}_t, 0, x_0) \stackrel{!}{=} x(t, x_0). \quad (1.7)$$

The induced flow map in this case is the one-parameter family of maps

$$F^t = F_0^t : x_0 \mapsto x(t, x_0). \quad (1.8)$$

Example 1.2 (Logistic Equation). For a resource-limited population, we have the following dynamical system for $a > 0$, $b > 0$, and the population $x \in \mathbb{R}_+ \cup \{0\}$

$$\dot{x} = ax(b - x). \quad (1.9)$$

In this case we have $E = \mathbb{R}$ and $\mathcal{F} = \{F^t\}_{t=-\infty}^{+\infty}$. This system has globally existing unique solutions (see later). We may analyze the behavior of this system by plotting \dot{x} as a function of x , analogously to the cobweb diagram. This is demonstrated in Fig. 1.3. At x values, where \dot{x} is positive $x(t)$ is growing, while at negative values it is decreasing. This means, that fixed points, at which $x(t) = \text{const.}$ correspond to intersections of the graphs with the horizontal axis.

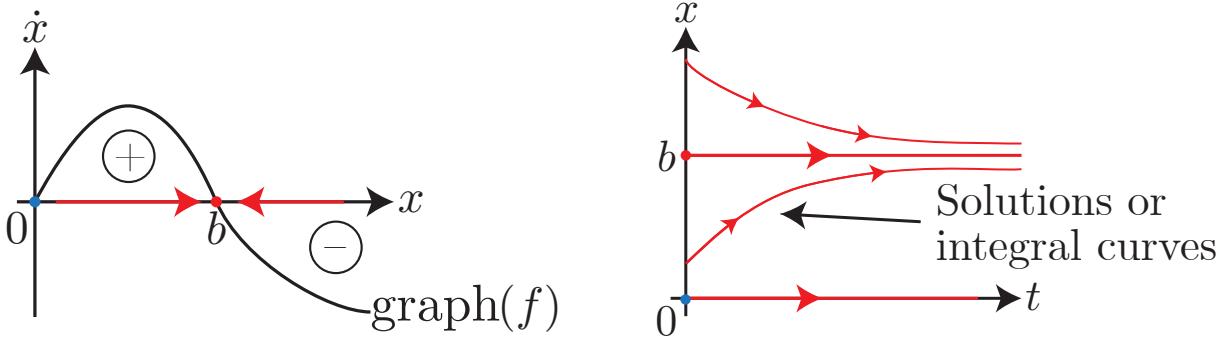


Figure 1.3: Left: Analysis of the right hand side. Right: Evolution in the extended phase space $P \times \mathbb{R}$.

Example 1.3 (Pendulum). Given the equation of motion

$$ml^2\ddot{\varphi} = -mgl \sin(\varphi). \quad (1.10)$$

We let $x_1 = \varphi$ and $x_2 = \dot{\varphi}$ to transform into the first-order ODE form

$$\begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -\frac{g}{l} \sin(x_1). \end{cases} \quad (1.11)$$

Thus we have

$$x = \begin{pmatrix} x_1 \\ x_2 \end{pmatrix}; \quad f(x) = \begin{pmatrix} x_2 \\ -\frac{g}{l} \sin(x_1) \end{pmatrix}. \quad (1.12)$$

Qualitative analysis gives the following facts

- $(x_1, x_2) = (0, 0)$ and $(x_1, x_2) = (\pi, 0)$ are zeros of f .
- Energy is conserved, hence both small and large amplitude oscillations are expected.
- The function $f(x)$ has symmetries: it is invariant under the transformation $(x_1, x_2, t) \mapsto (x_1, -x_2, -t)$ and $(x_1, x_2, t) \mapsto (-x_1, x_2, -t)$. See the left panel of Fig. 1.4.

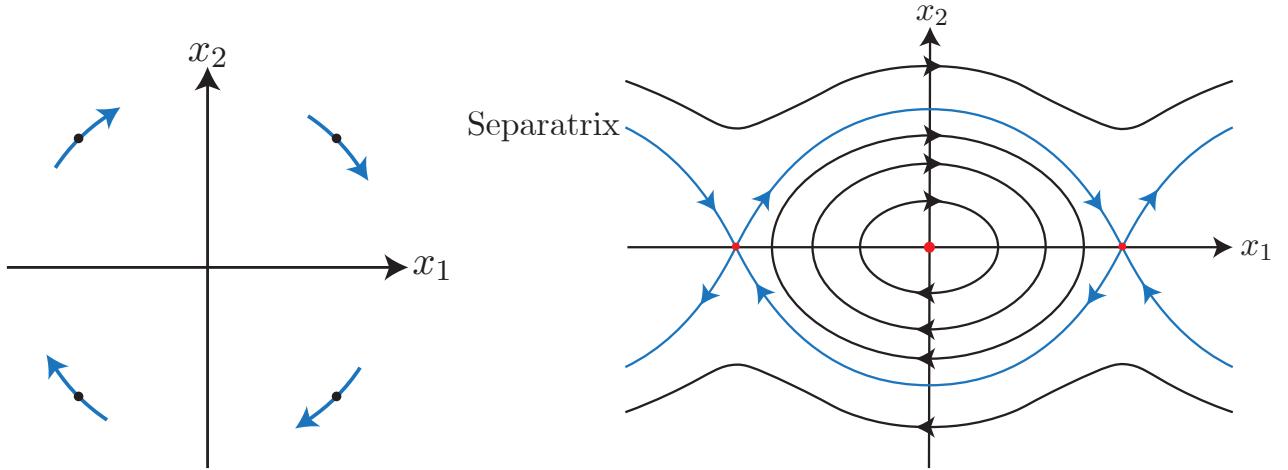


Figure 1.4: Left: The symmetries of the dynamical system. Right: Phase portrait of the pendulum. Red dots show the fixed points, while the blue trajectories make up the separatrix.

Definition 1.2. A separatrix is a boundary (i.e. a codimension-1 surface) in phase space which separates regions of qualitatively different behaviors. In practice, it is unobservable by itself and connects different fixed points. The separatrix of the pendulum is shown in the right panel of Fig. 4.

Example 1.4 (Exploit geometry of phase space for analysis). Consider two cities, A and B . The two cities are connected by two roads, denoted by the blue and green curves of the left panel of Fig. 1.5. We assume that travelling on the two roads, it is possible for two bikes to make it from A to B without ever being further away from each other than a distance $d < D$.

Assume two trucks are trying to make it between A and B , on different roads in the opposite direction, carrying a load of width D . Given this information, can the trucks make it without hitting each other? We can view this problem as a continuous dynamical system with two coordinates x_1 and x_2 that parameterize the two routes between A and B . This dynamical system is, in general, non autonomous.

The right panel of Fig 1.5 shows the trajectories of the two trucks and the two bikes in phase space. The two trajectories must intersect by continuity, thus at that point the trucks

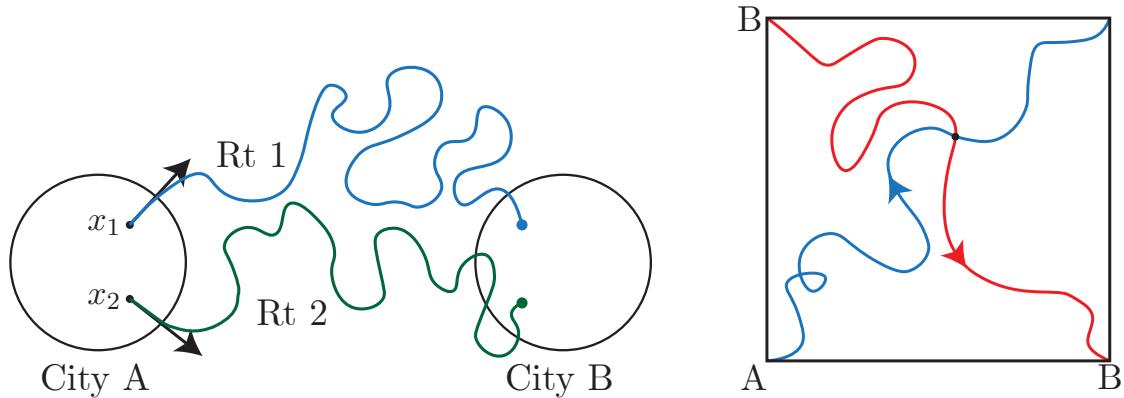


Figure 1.5: Left: An example of the two bike routes. Right: Blue represents the trajectory of the two bikes, red represents the trajectory of the two trucks.

must be at the same positions as the bikes, implying they are within distance D . Therefore the trucks must crash!

Exercise 1.1. Consider a pendulum that strikes an inclined wall repeatedly, as shown in Fig. 1.6 below. Using the phase portrait of the pendulum discussed in class, sketch the trajectories in the phase space of this impact dynamical system for positive and negative values of the angle α , when

- (i) there is no loss of energy at impact
- (ii) the coefficient of restitution is 0.5.

Identify the asymptotic behavior of the pendulum in each case.

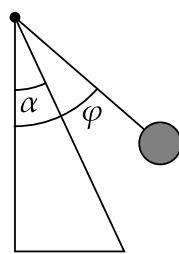


Figure 1.6: Pendulum attached to the inclined wall.

Hint: During impact, the pre-impact velocity of the pendulum is assumed to jump instantaneously to its post-impact value, while the pendulum's position remains the same.

Chapter 2

Fundamentals

In this chapter, we first review some fundamental properties of continuous dynamical systems that will be used heavily in later chapters. As we will see, these technical results are interesting in their own right. They can help in interpreting or cross-checking numerical results or physical models for self-consistency or accuracy.

2.1 Existence and uniqueness of solutions

Consider

$$\begin{cases} \dot{x} = f(x, t); & x \in \mathbb{R}^n \\ x(t_0) = x_0 \end{cases}. \quad (2.1)$$

Does this initial value problem have a unique solution? We have the following theorems to help us answer that question.

Theorem 2.1 (Peano). *If $f \in \mathcal{C}^0$ near (x_0, t_0) , then there exists a local solution $\varphi(t)$, i.e.,*

$$\dot{\varphi}(t) = f(\varphi(t), t), \varphi(t_0) = x_0; \quad \forall t \in (t_0 - \varepsilon, t_0 + \varepsilon); \quad 0 < \varepsilon \ll 1. \quad (2.2)$$

Example 2.1 (Free falling mass). Consider a point mass of mass m at position x . The acceleration due to gravity is denoted by g . Measuring the potential energy from the reference point $x = x_0$, we have the total energy is conserved.

$$\frac{1}{2}m\dot{x}^2 = mg(x - x_0). \quad (2.3)$$

This implies that

$$\begin{cases} \dot{x} = \sqrt{2g(x - x_0)} \\ x(0) = x_0 \end{cases} \quad (2.4)$$

on the set $P = \{x \in \mathbb{R} : x \geq x_0\}$. Therefore we have that $f \in \mathcal{C}^0$ in phase space, so by Peano's theorem (cf. Theorem 2.1), there exists a local solution. A schematic diagram is shown in Fig. 2.1. The solution is actually $x(t) = x_0 + \frac{g}{2}(t - t_0)^2$, however $x(t) = x_0$ is also a solution to

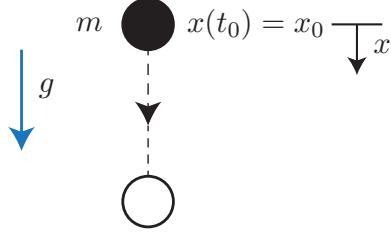


Figure 2.1: Schematic diagram of the point mass in free fall.

the IVP, therefore we do not have a unique solution. Physically there exists a solution, but this IVP was derived from a heuristic energy-principle, not from Newton's laws, which are not equivalent.

Definition 2.1. A function f is called locally Lipschitz around x_0 if there exists an open set U_{x_0} and $L > 0$ such that for all $x, y \in U_{x_0}$

$$\|f(y, t) - f(x, t)\| \leq L\|y - x\|. \quad (2.5)$$

Example 2.2 (Lipschitz functions). Fig. 2.2 shows an example of a Lipschitz and a non-Lipschitz function around x_0 .

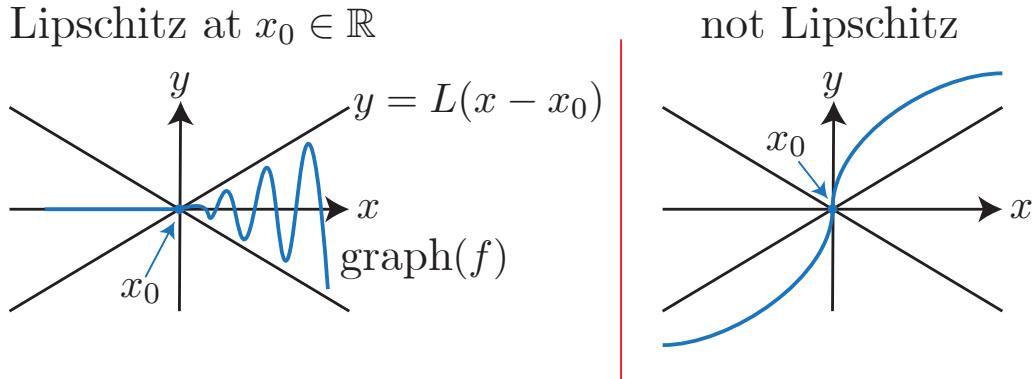


Figure 2.2: Interpretation of the Lipschitz property.

Theorem 2.2 (Picard). Assume

- (i) $f \in \mathcal{C}^0$ in t near (t_0, x_0) ,

(ii) f is locally Lipschitz in x near (t_0, x_0) .

Then there exists a unique local solution to the IVP. The proof can be found in [Arnold, 1992].

Note the following relations. If f is $\mathcal{C}^1 \implies f$ is Lipschitz $\implies f$ is \mathcal{C}^0 .

Example 2.3 (Free falling mass revisited). We check if f is Lipschitz.

$$\frac{|f(x) - f(x_0)|}{|x - x_0|} = \frac{\sqrt{2g}}{\sqrt{|x - x_0|}} \geq L. \quad (2.6)$$

Thus f is not Lipschitz near x_0 .

Exercise 2.1. Many important properties of nonlinear dynamical systems follow from Gronwall's inequality. Assume that two positive, continuous scalar functions $u(t)$ and $v(t)$ satisfy the condition

$$u(t) \leq C + \int_{t_0}^t u(\tau)v(\tau) d\tau \quad (2.7)$$

for some constant $C \geq 0$ and for all $t \geq t_0$. Then Gronwall's inequality asserts that

$$u(t) \leq Ce^{\int_{t_0}^t v(\tau) d\tau} \quad (2.8)$$

for all $t \geq t_0$. The significance of this result is that it gives a $u(t)$ -independent upper bound on the growth of $u(t)$. Using Gronwall's inequality, give an upper bound on how fast the solutions of a nonlinear ODE can separate from each other in time. In particular, show that for an ODE of the form

$$\dot{x} = f(x, t); \quad x \in \mathbb{R}^n,$$

and for two solutions starting from the initial conditions x_0 and \hat{x}_0 at time t_0 , we have

$$|x(t, x_0) - x(t, \hat{x}_0)| \leq |x_0 - \hat{x}_0|e^{L(t-t_0)},$$

where L is a Lipschitz constant for the function f over a domain containing the trajectories of the system over the time interval $[t_0, t]$.

Hint: Substitute both solutions into the ODE, integrate the resulting two equations from t_0 to t , and estimate their normed difference.

2.2 Geometric consequences of uniqueness

If the solution is unique, we have a few facts that can be derived from the geometric point of view.

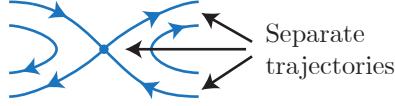


Figure 2.3: The phase portrait of the pendulum. Trajectories do not intersect since each arrow is pointing at separate trajectories.

- (i) The trajectories of autonomous systems cannot intersect. Note that fixed points do not violate this (e.g. pendulum equations). See Fig. 2.3 which shows the phase portrait of the pendulum.
- (ii) For non-autonomous systems, intersections in phase space are possible: a trajectory may occupy the same point x at a different time instants (see the left panel of Fig. 2.4. In this case we can extend the phase space in order to get an autonomous system where there cannot be any intersections.

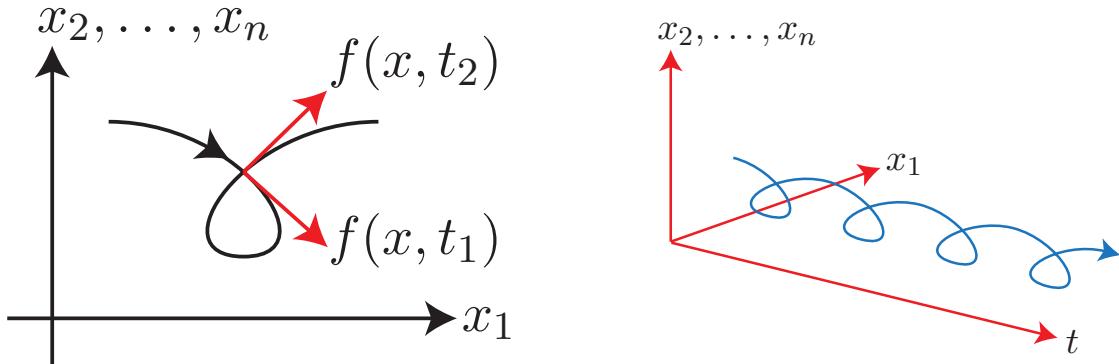


Figure 2.4: Left: Intersecting trajectories in phase space for a non-autonomous system. Right: The same trajectory in the extended phase space, without intersections.

$$X = \begin{pmatrix} x \\ t \end{pmatrix}, \quad F(X) = \begin{pmatrix} f(x, t) \\ 1 \end{pmatrix}; \quad \dot{X} = F(X). \quad (2.9)$$

2.3 Local vs global existence

Example 2.4 (Exploding solution).

$$\begin{cases} \dot{x} = x^2 \\ x(t_0) = 1. \end{cases} \quad (2.10)$$

Integrating yields the solution $x(t) = \frac{1}{1-(t-t_0)}$. This solution blows up at $t_\infty = t_0 + 1$, therefore the solution is only local. This is demonstrated in Fig. 2.5.

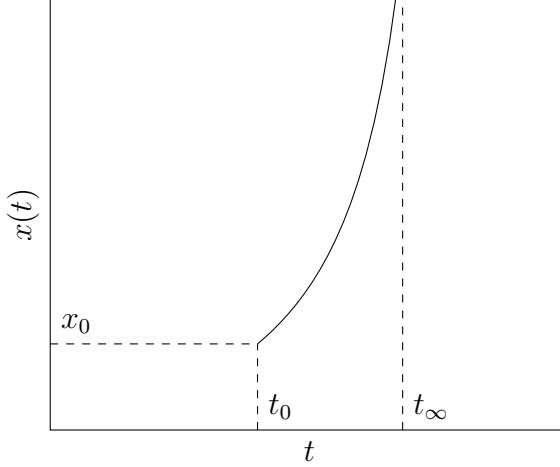


Figure 2.5: Solution to the ODE $\dot{x} = x^2$ started from $x(t_0) = 1$.

To address this problem of local solutions not being able to be continued into global solution, we have the following theorem.

Theorem 2.3 (Continuation of solution). *If a local solutions cannot be continued to a time $t = T$, then we must have*

$$\boxed{\lim_{t \rightarrow T} \|x(t)\| = \infty.} \quad (2.11)$$

The proof can be found in [Arnold, 1992].

Example 2.5 (Coupled Pendulum System). Consider two pendula of masses m_1 and m_2 . They both have length l . The angles of these pendula are denoted by φ_1 and φ_2 . Let us assume that they are coupled by a nonlinear spring, which can be described by a potential $V(\varphi_1, \varphi_2)$. This setup is illustrated in Fig. 2.6. We set $x_1 = \varphi_1$, $x_2 = \dot{\varphi}_1$, $x_3 = \varphi_2$, $x_4 = \dot{\varphi}_2$ and get the following equation of motion

$$\begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = \dots \\ \dot{x}_3 = x_4 \\ \dot{x}_4 = \dots \end{cases} \quad (2.12)$$

The RHS is smooth, therefore there exists a unique local solution to any IVP. The phase space

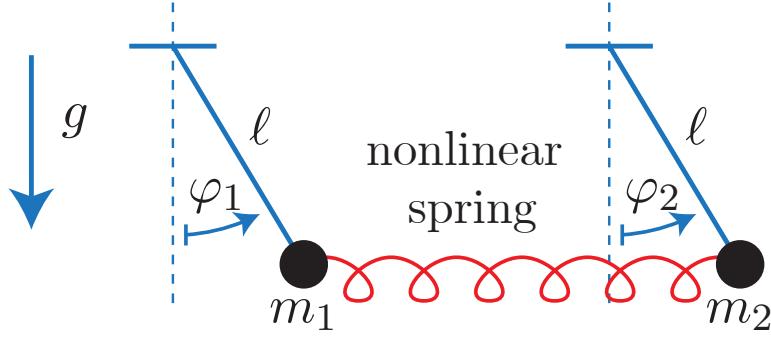


Figure 2.6: Physical setup of the coupled pendulum with a nonlinear spring.

is given by

$$P = \{x : x_1 \in S^1, x_2 \in \mathbb{R}, x_3 \in S^1, x_4 \in \mathbb{R}\} = S^1 \times \mathbb{R} \times S^1 \times \mathbb{R}. \quad (2.13)$$

Where S^1 is the 1 dimensional sphere (i.e. a circle). With this space we know that $\|x_1\|$ and $\|x_3\|$ are bounded. Due to energy being conserved we have

$$E = T + V = \frac{1}{2}m_1l_1x_2^2 + \frac{1}{2}m_2l_2x_4^2 + \underbrace{V(x_1, x_3)}_{\geq 0} \quad (2.14a)$$

$$E = E_0 = \text{constant} \geq 0. \quad (2.14b)$$

Hence $\|x_2\|$ and $\|x_4\|$ are also bounded, therefore all solutions exist globally.

Definition 2.2. A linear system is one such that for $x \in \mathbb{R}^n$, $A(t) \in \mathbb{R}^{n \times n}$ and $A \in \mathcal{C}^0$

$$\boxed{\dot{x} = A(t)x.} \quad (2.15)$$

Remark 2.4. Note that A can be written as $A = S + \Omega$ where $S = \frac{1}{2}(A + A^T)$ is symmetric (i.e. $S = S^T$) and $\Omega = \frac{1}{2}(A - A^T)$ is skew symmetric (i.e. $\Omega = -\Omega^T$). Furthermore the eigenvalues of S , λ_i , are all real and their respective eigenvectors, e_i , are orthogonal.

Example 2.6 (Global existence in linear systems).

$$\langle x, \dot{x} \rangle = \frac{1}{2} \frac{d}{dt} \|x(t)\|^2 = \langle x, A(t)x \rangle = \langle x, (S(t) + \Omega(t))x \rangle \quad (2.16a)$$

$$= \langle x, S(t)x \rangle + \underbrace{\langle x, \Omega(t)x \rangle}_{=0} \stackrel{(*)}{=} \sum_{i=1}^n \lambda_i(t)x_i^2 \quad (2.16b)$$

$$\leq \lambda_{\max}(t) \sum_{i=1}^n x_i^2 = \lambda_{\max}(t) \|x(t)\|^2. \quad (2.16c)$$

Where in $(*)$ we used that $x = \sum_{i=1}^n x_i e_i$ with $\|e_i\| = 1$ and $e_i \perp e_j$ for all $i \neq j$. Thus we get

$$\frac{\frac{1}{2} \frac{d}{dt} \|x(t)\|^2}{\|x(t)\|^2} \leq \lambda_{\max}(t) \implies \frac{1}{2} \log \left(\frac{\|x(t)\|^2}{\|x(t_0)\|^2} \right) \leq \int_{t_0}^t \lambda_{\max}(s) ds. \quad (2.17)$$

By exponentiating both sides, we obtain

$$\|x(t)\| \leq \|x(t_0)\| \exp \left(\int_{t_0}^t \lambda_{\max}(s) ds \right).$$

(2.18)

Therefore, by the continuation theorem, global solutions exist as long as $\int_{t_0}^t \lambda_{\max}(s) ds < \infty$.

2.4 Dependence on initial conditions

Given the IVP

$$\begin{cases} \dot{x} = f(x, t) \\ x(t_0) = x_0. \end{cases} \quad (2.19)$$

With $x \in \mathbb{R}^n$ and $f \in \mathcal{C}^r$ for some $r \geq 1$, we have the solution $x(t; t_0, x_0)$.

The dependence of the solution on initial data is of interest to us. This is due to us wanting the solution to be robust with respect to errors and uncertainties in the initial data. To address this, we have Theorem 2.5.

Theorem 2.5. *If $f \in \mathcal{C}^r$ for $r \geq 1$ then $x(t; t_0, x_0)$ is \mathcal{C}^r in (t_0, x_0) . Proof in [Arnold, 1992].*

The geometric meaning of this is that for $U \subset P \subset \mathbb{R}^n$ we have that $F_{t_0}^t(U)$ is a smooth deformation of U (cf. Fig. 2.7). It turns out $(F_{t_0}^t)^{-1} = F_t^{t_0}$ is also \mathcal{C}^r , hence we have that $F_{t_0}^t$ is a diffeomorphism.

Remark 2.6 (The total differential). We denote the total differential of a function $f : \mathbb{R}^n \rightarrow \mathbb{R}^m$ as Df . The total differential is a function which takes a location x as the argument and returns the derivative of f at the point x , i.e. the Jacobian. This implies evaluating the Jacobian at the point x . For a function $f(x, y) = f(x_1, \dots, x_n, y_1, \dots, y_m) : \mathbb{R}^{n+m} \rightarrow \mathbb{R}^k$ the total differential Df means the differential with respect to all of the variables and the total differential with respect to x , written $D_x f$ is the total differential only taken with respect to the x variables.

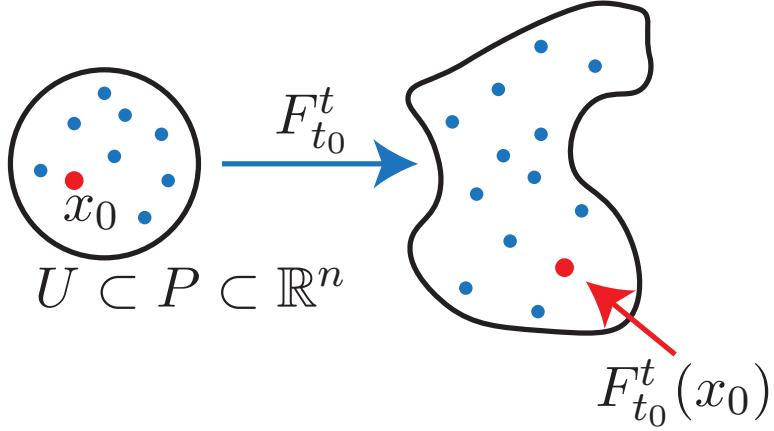


Figure 2.7: The smooth transformation of U . The red point on the right is $F_{t_0}^t(x_0)$, i.e. the image of x_0 under the evolution operator.

Thus for $f(x, y) : \mathbb{R}^{n+m} \rightarrow \mathbb{R}^k$ we have the total differential

$$Df = \begin{pmatrix} \frac{\partial f_1}{\partial x_1} & \dots & \frac{\partial f_1}{\partial x_n} & \frac{\partial f_1}{\partial y_1} & \dots & \frac{\partial f_1}{\partial y_m} \\ \vdots & & \vdots & \vdots & & \vdots \\ \frac{\partial f_k}{\partial x_1} & \dots & \frac{\partial f_k}{\partial x_n} & \frac{\partial f_k}{\partial y_1} & \dots & \frac{\partial f_k}{\partial y_m} \end{pmatrix}; \quad (2.20)$$

$$Df(x_0, y_0) = \begin{pmatrix} \frac{\partial f_1}{\partial x_1}(x_0, y_0) & \dots & \frac{\partial f_1}{\partial x_n}(x_0, y_0) & \frac{\partial f_1}{\partial y_1}(x_0, y_0) & \dots & \frac{\partial f_1}{\partial y_m}(x_0, y_0) \\ \vdots & & \vdots & \vdots & & \vdots \\ \frac{\partial f_k}{\partial x_1}(x_0, y_0) & \dots & \frac{\partial f_k}{\partial x_n}(x_0, y_0) & \frac{\partial f_k}{\partial y_1}(x_0, y_0) & \dots & \frac{\partial f_k}{\partial y_m}(x_0, y_0) \end{pmatrix}, \quad (2.21)$$

and the total differential with respect to x

$$D_x f = \begin{pmatrix} \frac{\partial f_1}{\partial x_1} & \dots & \frac{\partial f_1}{\partial x_n} \\ \vdots & & \vdots \\ \frac{\partial f_k}{\partial x_1} & \dots & \frac{\partial f_k}{\partial x_n} \end{pmatrix}; \quad D_x f(x_0, y_0) = \begin{pmatrix} \frac{\partial f_1}{\partial x_1}(x_0, y_0) & \dots & \frac{\partial f_1}{\partial x_n}(x_0, y_0) \\ \vdots & & \vdots \\ \frac{\partial f_k}{\partial x_1}(x_0, y_0) & \dots & \frac{\partial f_k}{\partial x_n}(x_0, y_0) \end{pmatrix}. \quad (2.22)$$

Now, how can we compute the Jacobian of the flow map $\frac{\partial x(t; t_0, x_0)}{\partial x_0} = DF_{t_0}^t(x_0)$? We start from the IVP and take the gradient (with respect to x_0) of both sides. On the left hand side we can exchange the order of the time derivative and the gradient and on the right hand side we use the chain rule. We end up with the equation

$$\frac{d}{dt} \frac{\partial x}{\partial x_0} = D_x f(x(t; t_0, x_0), t) \frac{\partial x}{\partial x_0}; \quad \frac{\partial x}{\partial x_0} \in \mathbb{R}^{n \times n}. \quad (2.23)$$

This means, that the flow map gradient satisfies the IVP

$$\frac{d}{dt} [DF_{t_0}^t(x_0)] = D_x f(F_{t_0}^t(x_0), t) DF_{t_0}^t(x_0) \quad (2.24a)$$

$$DF_{t_0}^{t_0}(x_0) = I. \quad (2.24b)$$

This is called the equation of variations, which is a linear, non-autonomous ODE for the matrix $M = DF_{t_0}^t(x_0)$

$$\begin{cases} \dot{M} = D_x f(x(t; t_0, x_0)) M \\ M(t_0) = I. \end{cases} \quad (2.25)$$

Example 2.7 (Locations of extreme deformation in phase space). We define

$$\xi(t) := \tilde{x}(t) - x(t) = x(t; t_0, \tilde{x}_0) - x(t; t_0, x_0) \quad (2.26a)$$

$$= x(t; t_0, x_0) + \frac{\partial x}{\partial x_0}(t; t_0, x_0) \xi_0 + \mathcal{O}(\|\xi_0\|^2) - x(t; t_0, x_0) \quad (2.26b)$$

$$= DF_{t_0}^t(x_0) \xi_0 + \mathcal{O}(\|\xi_0\|^2). \quad (2.26c)$$

Where we used the Taylor expansion and assume the perturbation to x_0 is small, i.e. $\|\xi_0\| \ll 1$. Therefore we have

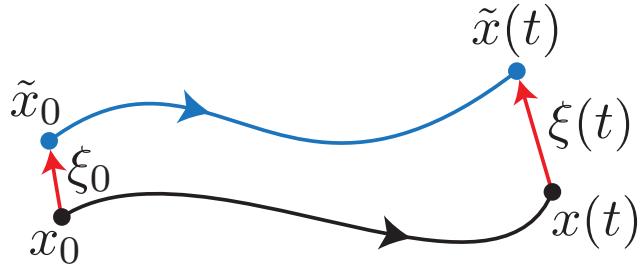


Figure 2.8: Illustration of the quantities ξ_0 and $\xi(t)$.

$$\|\xi(t)\|^2 = \langle DF_{t_0}^t(x_0) \xi_0, DF_{t_0}^t(x_0) \xi_0 \rangle + \mathcal{O}(\|\xi_0\|^3) \quad (2.27a)$$

$$= \langle \xi_0, \underbrace{[DF_{t_0}^t(x_0)]^T DF_{t_0}^t(x_0)}_{=: C_{t_0}^t(x_0)} \xi_0 \rangle + \mathcal{O}(\|\xi_0\|^3). \quad (2.27b)$$

$C_{t_0}^t(x_0)$ is known as the Cauchy-Green strain tensor. It is positive definite and symmetric and due to its dependence on the initial condition, $C_{t_0}^t(x_0)$ actually defines a *tensor field*. Therefore the largest possible deformation is

$$\max_{x_0, \xi_0} \frac{\|\xi(t)\|^2}{\|\xi_0\|^2} = \max_{x_0, \xi_0} \frac{\langle \xi_0, C_{t_0}^t(x_0) \xi_0 \rangle}{\|\xi_0\|^2} = \max_{x_0} \lambda_n(x_0). \quad (2.28)$$

Where we used that $C_{t_0}^t$ is positive definite in the last equality, and that $\lambda_n(x_0)$ is the largest eigenvalue of $C_{t_0}^t(x_0)$. Because we typically have exponential growth we introduce the following quantity.

Definition 2.3. The finite-time Lyapunov exponent is defined as

$$\text{FTLE}_{t_0}^t(x_0) := \frac{1}{2|t - t_0|} \log(\lambda_n(x_0)). \quad (2.29)$$

The FTLE is a diagnostic quantity for Lagrangian Coherent Structures (LCS), i.e. influential surfaces governing the evolution in P .

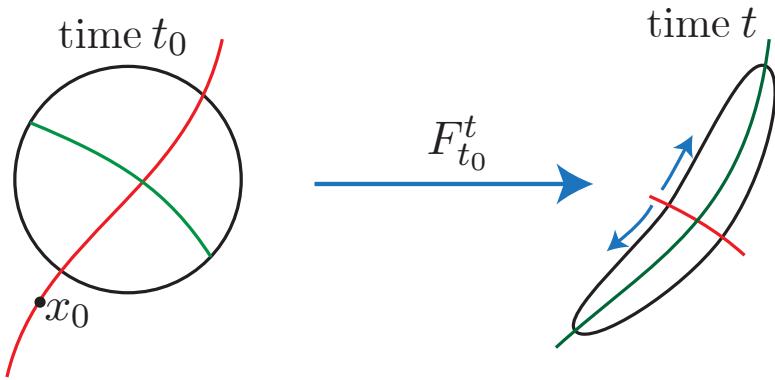


Figure 2.9: On the left the red ridge represents large values of $\text{FTLE}_{t_0}^t$, on the right the green ridge the high values of $\text{FTLE}_t^{t_0}$.

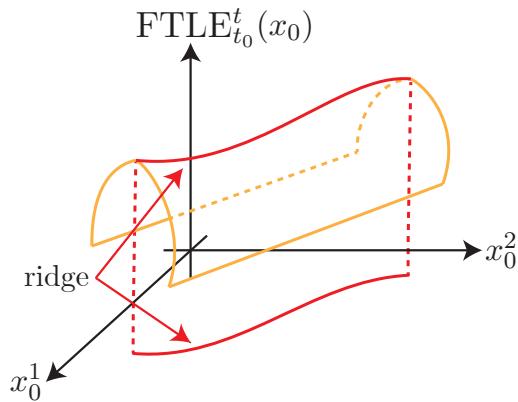


Figure 2.10: The projection of the FTLE ridge onto the initial value space.

The ridges of $\text{FTLE}_{t_0}^t$ are the repelling LCS, meanwhile the ridges of $\text{FTLE}_t^{t_0}$ are the attracting LCS as depicted in Fig. 2.9. Now we are left with the problem of computing $F_{t_0}^t(x_0)$.

Recall that analytically we start with $F_{t_0}^t(x_0)$ and use this to calculate $DF_{t_0}^t(x_0)$. From here we can find $C_{t_0}^t(x_0)$, giving us $\lambda_n(x_0)$ and thereby the FTLE. We now outline a process to compute the FTLE numerically.

- (i) Define an initial $M \times N$ grid of initial data $x_0(i, j) \in \mathbb{R}^2$.
- (ii) Launch trajectories numerically from grid points to obtain a discrete approximation of $F_{t_0}^t(x_0)$ as $F_{t_0}^t(x_0(i, j))$.
- (iii) Use finite differencing to approximate

$$DF_{t_0}^t(x_0(i, j)) \approx \begin{pmatrix} \frac{x(t; t_0, x_0(i, j) + \delta e_1)_1 - x(t; t_0, x_0(i, j) - \delta e_1)_1}{2\delta} & \dots & \frac{x(t; t_0, x_0(i, j) + \delta e_n)_1 - x(t; t_0, x_0(i, j) - \delta e_n)_1}{2\delta} \\ \vdots & & \vdots \\ \frac{x(t; t_0, x_0(i, j) + \delta e_1)_n - x(t; t_0, x_0(i, j) - \delta e_1)_n}{2\delta} & \dots & \frac{x(t; t_0, x_0(i, j) + \delta e_n)_n - x(t; t_0, x_0(i, j) - \delta e_n)_n}{2\delta} \end{pmatrix}. \quad (2.30)$$

This process then yields the surface we see in Fig. 2.10.

Example 2.8 (Calculating the FTLE for the double gyre). Due to incompressibility, we can define the two dimensional flow using a single scalar function called the stream function.

$$\Psi(x, y) = -\sin(\pi x) \sin(\pi y). \quad (2.31)$$

The components (u, v) of the fluid velocity ($v = (u, v)$) are obtained as partial derivatives of the stream function, according to the formulas

$$\begin{cases} u = \frac{\partial \Psi}{\partial y} \\ v = -\frac{\partial \Psi}{\partial x}. \end{cases} \quad (2.32)$$

The Lagrangian trajectories of fluid particles obey the differential equations (i.e. we have the fluid velocity field)

$$\begin{cases} \dot{x} = u = \frac{\partial \Psi}{\partial y} \\ \dot{y} = v = -\frac{\partial \Psi}{\partial x}. \end{cases} \quad (2.33)$$

Interestingly, in this case, the phase space coincides with the physical space spanned by the coordinates (x, y) .

Remark 2.7. This is an example of a Hamiltonian system, where Ψ is the Hamiltonian (usually denoted as H).

For any autonomous Hamiltonian system we have that the Hamiltonian is constant along trajectories. We can verify this as follows

$$\frac{d}{dt}\Psi(x(t), y(t)) = \frac{\partial\Psi}{\partial x}\dot{x} + \frac{\partial\Psi}{\partial y}\dot{y} = 0. \quad (2.34)$$

So we have that trajectories are level curves of $\Psi(x, y)$. We can then derive the phase portrait from the level curves of Ψ . Further, we have that $\dot{x} = \frac{\partial\Psi}{\partial y} = -\pi \sin(\pi x) \cos(\pi y)$ which yields that $\text{sign}(\dot{x}) = -\text{sign}(\sin(\pi x))\text{sign}(\cos(\pi y))$. Putting these together we can construct the contour plot with arrows. The contour plot, and FTLE approximation are shown in Fig. 2.11.

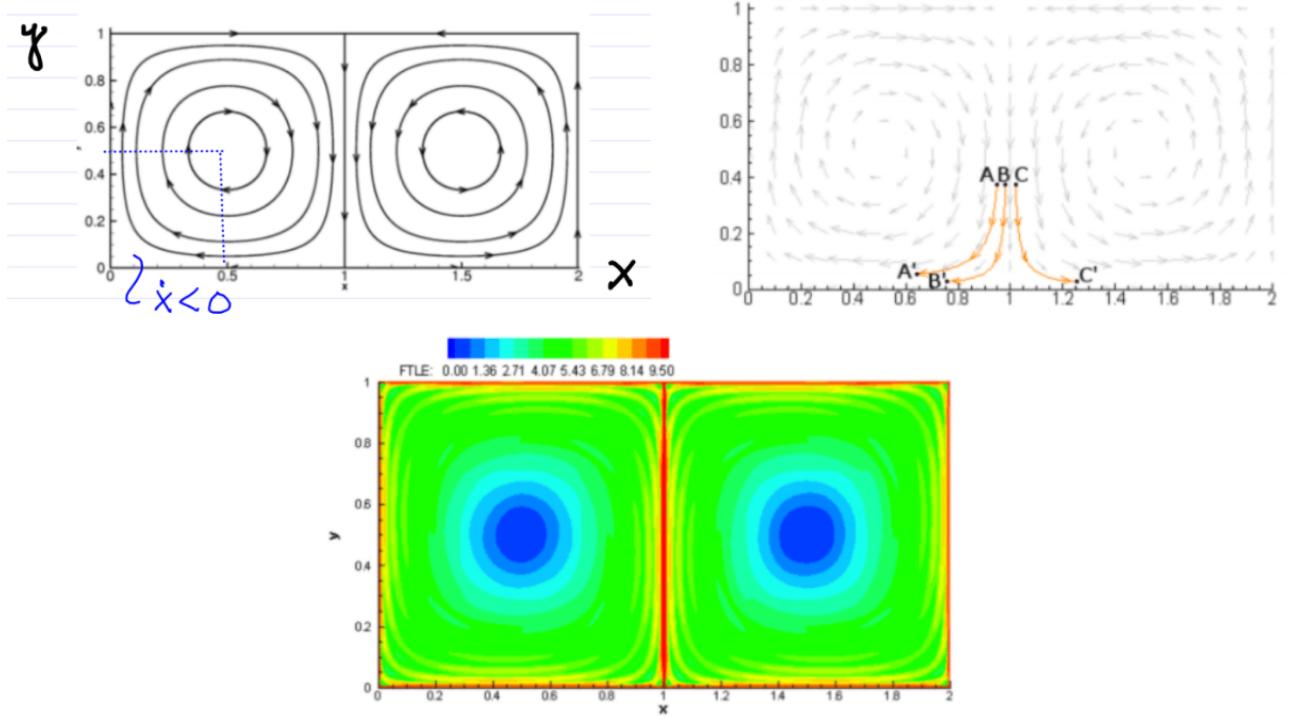


Figure 2.11: Top left: The analytic phase plot. Top right: The exploration done to calculate FTLE. Bottom: The FTLE plot. Figures here were taken from Shawn Shadden of UC Berkeley.

Notebook 2.1. [Calculation of the FTLE field and extraction of its ridges](#)

Example 2.9 (ABC flow). Let our dynamical system be defined as follows with $A, B, C \in \mathbb{R}$

$$\begin{cases} \dot{x} = A \sin(z) + C \cos(y) \\ \dot{y} = B \sin(x) + A \cos(z) \\ \dot{z} = C \sin(y) + B \cos(x). \end{cases} \quad (2.35)$$

This is an exact solution to Euler's equations. We have an autonomous velocity field. Depending on parameters it can even generate chaotic fluid trajectories. The numerical approximation of the FTLE for the ABC flow is depicted in Fig. 2.12.

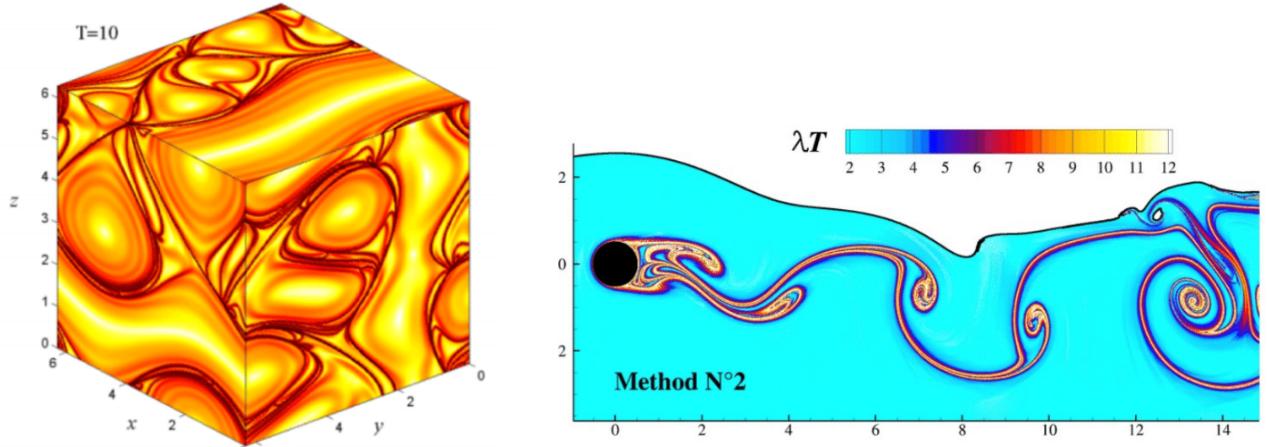


Figure 2.12: Left: numerically calculated FTLE field of the ABC flow. Darker colors signify higher FTLE values [Haller, 2001]. Right: Again the FTLE is plotted, for vortex shedding behind a cylinder under a free surface [Sun et al., 2016].

Exercise 2.2. Consider the non-dimensionalized, force-damped pendulum equation

$$\ddot{x} + k\dot{x} + \sin(x) = a \sin(t), \quad (2.36)$$

where $k \geq 0$ is the damping coefficient and $a \geq 0$ is the forcing amplitude.

- (a) For vanishing damping and forcing ($a = k = 0$), compute and plot numerically the FTLE field for this system over a 100×100 grid of initial conditions, covering the square $[-\pi, \pi] \times [-\pi, \pi]$ in the phase space of the pendulum. Perform the computation for long enough times so that the FTLE plot fully reveals the separatrices of the system, as we discussed in class for the undamped pendulum.

- (b) To explore the fate of these separatrices under mild damping and forcing, repeat the same FTLE computations for $a = 0.5$ and $k = 0.1$. Discuss domains of attractions and their boundaries based on the results.

Hint: Consider using MATLAB's ODE45 routine for trajectory integration. To this end, you need to assemble all initial grid points into a single long vector, and pass that vector on to ODE45 as an initial condition. You will then need to define an appropriate extended ODE, whose right-hand side will be called by ODE45 to produce an advected image of the extended initial condition vector. Subsequently, you will need to convert the elements of this advected vector back into advected grid positions. Alternatively, you could write your own 4-point Runge-Kutta solver to track trajectories simultaneously on the full grid of initial conditions (see the algorithm discussed in class).

2.5 Dependence on parameters

We now have the IVP

$$\begin{cases} \dot{x} = f(x, t, \mu) \\ x(t_0) = x_0. \end{cases} \quad (2.37)$$

With $x \in \mathbb{R}^n$, $f \in \mathcal{C}^r$, $r \geq 1$, therefore we have a solution $x(t; t_0, x_0, \mu) \in \mathcal{C}_{x_0}^r$.

We now examine how solutions depend μ . This is critical as solutions should be robust to changes or uncertainties in the model.

Example 2.10 (Perturbation Theory). Given a weakly nonlinear oscillator

$$m\ddot{x} + c\dot{x} + kx = \varepsilon f(x, \dot{x}, t), \quad 0 \leq \varepsilon \ll 1, \quad x \in \mathbb{R}. \quad (2.38)$$

The usual approach is to seek solutions by expanding from the known solution of the linear limit $\varepsilon = 0$, i.e.

$$x_\varepsilon(t) = \varphi_0(t) + \varepsilon\varphi_1(t) + \varepsilon^2\varphi_2(t) + \dots + \mathcal{O}(\varepsilon^r). \quad (2.39)$$

If $x_\varepsilon(t)$ is in $\mathcal{C}_\varepsilon^r$, we have $\varphi_1(t) = \left. \frac{\partial x_\varepsilon(t)}{\partial \varepsilon} \right|_{\varepsilon=0}$ and $\varphi_2(t) = \left. \frac{\partial^2 x_\varepsilon(t)}{\partial \varepsilon^2} \right|_{\varepsilon=0}$

Regularity with respect to the parameter μ actually follows from regularity with respect to the initial condition x_0 . We can use the following trick to extend the IVP with a dummy variable μ

$$\begin{cases} \dot{x} = f(x, t, \mu) \\ \dot{\mu} = 0 \\ x(t_0) = x_0 \\ \mu(t_0) = \mu_0. \end{cases} \quad (2.40)$$

Thus with $X = \begin{pmatrix} x \\ \mu \end{pmatrix} \in \mathbb{R}^{n+p}$ and $F(X_0) = \begin{pmatrix} f \\ 0 \end{pmatrix}$; $X_0 = \begin{pmatrix} x_0 \\ \mu_0 \end{pmatrix}$. We have the extended IVP

$$\begin{cases} \dot{X} = F(X) \\ X(t_0) = X_0. \end{cases} \quad (2.41)$$

Applying the previous result on regularity with respect to x_0 to (2.41), we have that $f \in \mathcal{C}_{x,\mu}^r$ implies that $X(t) \in \mathcal{C}_{X_0}^r$ in turn implying that $x(t; t_0, x_0, \cdot) \in \mathcal{C}_\mu^r$. The solution is as smooth in parameters as the RHS of the dynamical system.

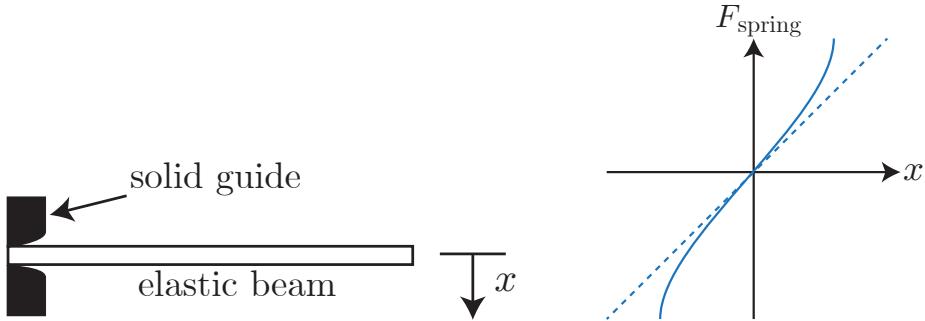


Figure 2.13: Setup for the nonlinear springboard.

Example 2.11 (Periodic Oscillations of a nonlinear springboard). Given an elastic beam extending from a solid guide, we measure the deflection of this beam with the variable x . This system is illustrated in the left panel of Fig. 2.13. By increasing x , the effective free length of the beam is shortened, thereby stiffening the spring nonlinearly. The effect of this nonlinearity on the force exerted on the spring is illustrated in the right panel of Fig. 2.13. This setup yields the following equations of motion

$$\begin{cases} \ddot{x} + x + \varepsilon x^3 = 0; & 0 \leq \varepsilon \ll 1 \\ x(0) = a_0; & \dot{x}(0) = 0. \end{cases} \quad (2.42)$$

So we have weak nonlinearity with no known explicit solution. Although weak, this nonlinearity is still significant, as can be seen in Fig. 2.14. Rewriting this as a first order ODE ($x_1 = x$; $x_2 = \dot{x}$), and note that the RHS is $\mathcal{C}_{x,\mu}^r$, therefore there exists a unique local solution that is also \mathcal{C}_μ^r . Thus the expansion is justified

$$x_\varepsilon(t) = \varphi_0(t) + \varepsilon \varphi_1(t) + \dots + \mathcal{O}(\varepsilon^r). \quad (2.43)$$

We can see, by substitution, that for $\varepsilon = 0$ we find that $\varphi_0(t) = a_0 \cos(t)$.

Now we look specifically for T -periodic solutions, as we would expect such a solution physically, therefore we have

$$\varphi_i(t) = \varphi_i(t + T). \quad (2.44)$$

The period T still has to be determined. Plugging this power series into (2.43) to get

$$\mathcal{O}(1) : \ddot{\varphi}_0 + \varphi_0 = 0 \quad (2.45a)$$

$$\mathcal{O}(\varepsilon) : \ddot{\varphi}_1 + \underbrace{\varphi_1}_{\omega=1} = -\varphi_0^3 = -a_0^3 \cos^3(t) = -a_0^3 \left[\frac{1}{4} \cos(3t) + \frac{3}{4} \underbrace{\cos(t)}_{\text{resonance}} \right]. \quad (2.45b)$$

We can see that (2.45b) is a linear oscillator with a forcing coming from the zeroth order solution. Since the zeroth order solution $\varphi_0 = a_0 \cos(t)$ already solves the IVP we have the following initial conditions

$$\varphi_1(0) = 0; \quad \dot{\varphi}_1(0) = 0. \quad (2.46)$$

This holds as $\varphi_0 = a_0 \cos(t)$ already solves the IVP. The general solution to this equation is the sum of two terms. We add the general solution of the homogeneous part and a particular solution to the inhomogeneous part. We can write this solution to (2.45b) as

$$\varphi_1(t) = \varphi_1^{\text{hom}}(t) + \varphi_1^{\text{part}}(t) \quad (2.47a)$$

$$= \underbrace{A \cos(t) + B \sin(t)}_{\text{TBD from initial conditions}} + \underbrace{C \cos(3t) + D t \cos(t) + E t \sin(t)}_{\text{TBD from (2.45b)}}. \quad (2.47b)$$

Observe that due to a resonance between the natural frequency of the oscillator and the forcing secular terms, $t \cos(t)$ and $t \sin(t)$ appear. Thus it cannot be periodic, so our Ansatz already fails for $i = 1$. We conclude that no solution of this type exists. Our Ansatz was too restrictive and T should depend on ε .

Lindstedt's idea We should seek a solution of the form

$$x_\varepsilon(t) = \varphi_0(t; \varepsilon) + \varepsilon \varphi_1(t; \varepsilon) + \varepsilon^2 \varphi_2(t; \varepsilon) + \mathcal{O}(\varepsilon^3). \quad (2.48)$$

Furthermore φ_i should be T_ε periodic, i.e. the period should depend on the strength of the nonlinearity ε .

$$\varphi_i(t + T_\varepsilon; \varepsilon) = \varphi_i(t; \varepsilon). \quad (2.49)$$

Rewriting the period as

$$T_\varepsilon = \frac{2\pi}{\omega(\varepsilon)}; \quad \omega(\varepsilon) = 1 + \varepsilon \omega_1 + \varepsilon^2 \omega_2 + \mathcal{O}(\varepsilon^3). \quad (2.50)$$

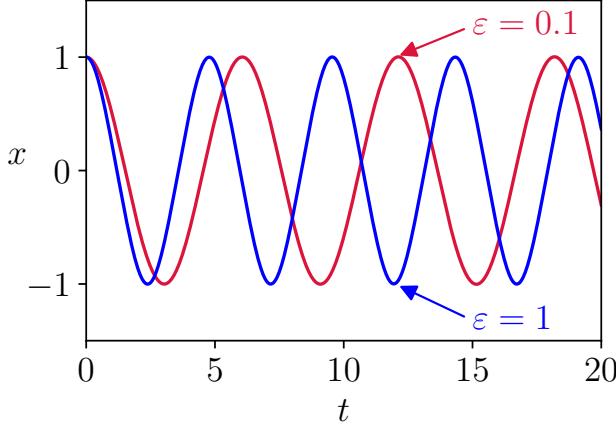


Figure 2.14: Numerical integration of x for $a_0 = 1$ and different values of ε .

We then rescale time according to $\tau = \omega(\varepsilon)t$ to find

$$\frac{d}{d\tau} = \frac{1}{\omega(\varepsilon)} \frac{d}{dt} \implies [\omega(\varepsilon)]^2 x'' + x + \varepsilon x^3 = 0. \quad (2.51)$$

Where we have taken x' to represent $\frac{dx}{dt}$. Plugging our expression into the rescaled ODE yields

$$(1 + 2\varepsilon\omega_1 + \mathcal{O}(\varepsilon^2)) [\varphi_0'' + \varepsilon\varphi_0''' + \mathcal{O}(\varepsilon^2)] + [\varphi_0 + \varepsilon\varphi_1 + \mathcal{O}(\varepsilon^2)] + \varepsilon [\varphi_0^3 + \mathcal{O}(\varepsilon)] = 0. \quad (2.52)$$

Matching equal powers of ε yields

$$\mathcal{O}(1) : \varphi_0'' + \varphi_0 = 0 \implies \varphi_0(\tau) = a_0 \cos(\tau); \quad \varphi_0(0) = a_0; \quad \dot{\varphi}_0(0) = 0 \quad (2.53a)$$

$$\mathcal{O}(\varepsilon) : \varphi_1'' + \varphi_1 = -\varphi_0^3 - 2\omega_1\varphi_0'' = \left(2\omega_1 a_0 - \frac{3}{4}a_0^3\right) \underbrace{\cos(\tau)}_{\text{resonance}} - \frac{a_0^3}{4} \cos(3\tau); \quad (2.53b)$$

$$\varphi_1(0) = 0; \quad \dot{\varphi}_1(0) = 0. \quad (2.53c)$$

From the first line, we can see the initial conditions are fulfilled. In this step we used that $\dot{\varphi}(t=0) = 0$ if and only if $\omega(\varepsilon)\varphi'(0) = 0$. We get the solution

$$\varphi_1(t) = A \cos(\tau) + B \sin(\tau) + C \cos(3\tau) + D\tau \cos(\tau) + E\tau \sin(\tau). \quad (2.54)$$

The presence of resonance again excludes periodic solutions, but now we can select ω_1 to eliminate these terms.

$$2\omega_1 a_0 - \frac{3}{4}a_0^3 = 0 \implies \omega_1 = \frac{3}{8}a_0^2. \quad (2.55)$$

This successfully eliminates the resonance and determines the missing frequency term at $\mathcal{O}(\varepsilon)$. Thus we find

$$x_\varepsilon(\tau) = a_0 \cos(\tau) - \frac{\varepsilon}{32} a_0^3 (\cos(\tau) - \cos(3\tau)) + \mathcal{O}(\varepsilon^2). \quad (2.56)$$

In the original time scaling this is

$$x_\varepsilon(t) = a_0 \cos(\omega t) - \frac{\varepsilon}{32} a_0^3 (\cos(\omega t) - \cos(3\omega t)) + \mathcal{O}(\varepsilon^2); \quad \omega = 1 + \frac{3}{8}\varepsilon a_0^2 + \mathcal{O}(\varepsilon^2). \quad (2.57)$$

This procedure can be continued to higher order terms, where we select ω_2 so that the $\mathcal{O}(\varepsilon^2)$ terms cancel.

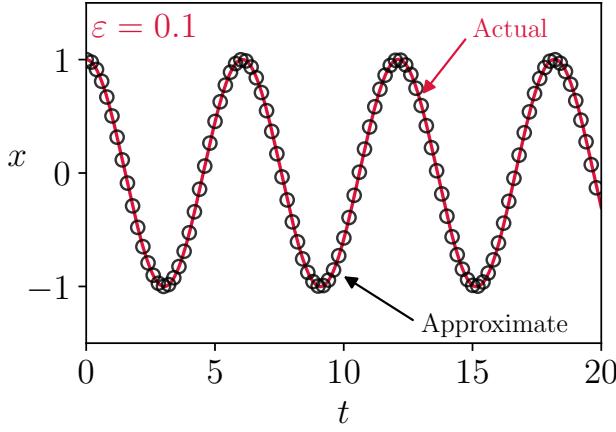


Figure 2.15: Approximation (dots) vs analytic solution (solid line) of x on the time interval $[0, 20]$.

Notebook 2.2. *Higher order computation of the perturbation expansion*

Exercise 2.3. Consider the nonlinear oscillator

$$\ddot{x} + \omega_0^2 x = \varepsilon M x^2, \quad (2.58)$$

where $\varepsilon M x^2$ represents a small nonlinear forcing term ($0 \leq \varepsilon \ll 1, M > 0$).

Using Lindstedt's method, find an $\mathcal{O}(\varepsilon)$ approximation for nonlinear motions as a function of their initial position, with zero initial velocity.

Exercise 2.4. Consider the forced van der Pol equation

$$\ddot{x} + \varepsilon(x^2 - 1)\dot{x} + x = F \cos(\omega t), \quad (2.59)$$

which arises in models of self-excited oscillation, such as those of a valve generator with a cubic valve characteristic. Here $F, \omega > 0$ are parameters, and $0 \leq \varepsilon \ll 1$.

- (i) For small values of ε , find an approximation for an **exactly** $2\pi/\omega$ -periodic solution of the equation. The error of your approximation should be $\mathcal{O}(\varepsilon)$.
- (ii) For $\varepsilon = 0.1$, $\omega = 2$, and $F = 1$, verify your prediction numerically by solving the equation numerically. Plot your numerical solution along with your analytic prediction computed in (i).

Note: For chaotic dynamics in the forced van der Pol equation, see Section 2.1 of [Guckenheimer and Holmes, 1990].

Chapter 3

Stability of fixed points

Now we would like to begin to explore the behaviour of dynamical systems around fixed points. This will allow us to find out if we should expect to observe a fixed state, and to understand what happens if we perturb the system away from this fixed state.

3.1 Basic definitions

Consider

$$\dot{x} = f(x, t), \quad x \in \mathbb{R}^n, \quad f \in \mathcal{C}^1. \quad (3.1)$$

Assume that $x = 0$ is a fixed point, i.e. $f(0, t) = 0$ for all $t \in \mathbb{R}$. If the fixed point is originally at $x_0 \neq 0$, shift it to zero by letting $\tilde{x} := x - x_0$, therefore

$$\dot{\tilde{x}} = \dot{x} = f(x_0 + \tilde{x}, t) = \tilde{f}(\tilde{x}, t). \quad (3.2)$$

We would like to understand how the dynamical system behaves near its equilibrium state. To this end we introduce the following definitions.

Definition 3.1 (Lyapunov Stability). The fixed point $x = 0$ is stable if for all t_0 , for all $\varepsilon > 0$ small enough, there exists a $\delta = \delta(t_0, \varepsilon)$, such that for all $x_0 \in \mathbb{R}^n$ with $\|x_0\| \leq \delta$, we have

$$\boxed{\|x(t; t_0, x_0)\| \leq \varepsilon \quad \forall t \geq t_0.} \quad (3.3)$$

Remark 3.1 (N-dimensional ball). When writing $\mathcal{B}(r)$ we refer to the ball of radius r in \mathbb{R}^n , i.e. the set $\{x : \|x\| < r\}$.

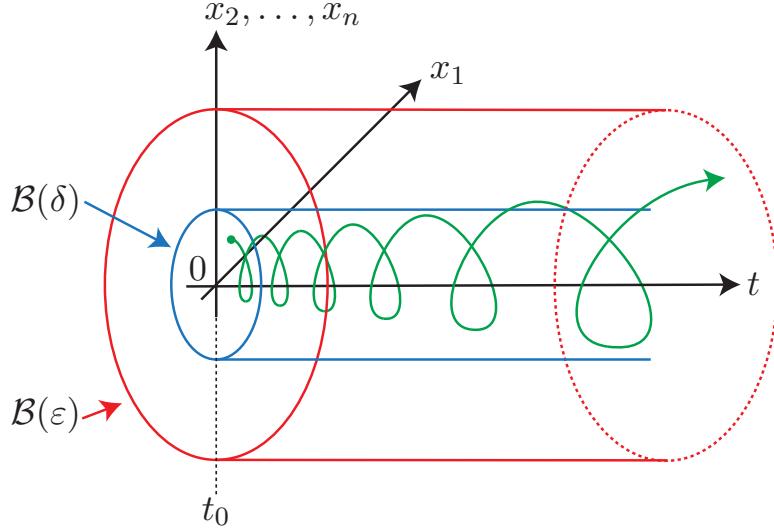


Figure 3.1: An example of δ for a given Lyapunov stable fixed point.

Example 3.1 (Stability of the lower equilibrium of the pendulum). Recall the equation of motion of the pendulum $\ddot{\varphi} + \sin(\varphi) = 0$, that we transform into a first order ODE by setting $x_1 = \varphi$ and $x_2 = \dot{\varphi}$ to obtain

$$\begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -\sin(x_1). \end{cases} \quad (3.4)$$

For small $\varepsilon > 0$, this geometric procedure gives a $\delta(\varepsilon) > 0$ such that the definition of stability is satisfied for $x = 0$. We can see in Fig. 3.2 that for any initial point chosen within the blue circle, its trajectory remains within the red circle for all time (cf. Fig. 3.1). Therefore $x = 0$ is (Lyapunov) stable.

Definition 3.2 (Asymptotic stability). The fixed point $x = 0$ is *asymptotically stable* if

- (i) it is stable,
- (ii) for all t_0 , there exists $\delta_0(t_0, \varepsilon)$ such that for every x_0 with $\|x_0\| \leq \delta_0$ we have

$$\boxed{\lim_{t \rightarrow \infty} x(t; t_0, x_0) = 0.} \quad (3.5)$$

Definition 3.3 (Domain of attraction). The *domain of attraction* of the fixed point $x = 0$ is the set of all x_0 's for which

$$\boxed{\lim_{t \rightarrow \infty} x(t; t_0, x_0) = 0.} \quad (3.6)$$

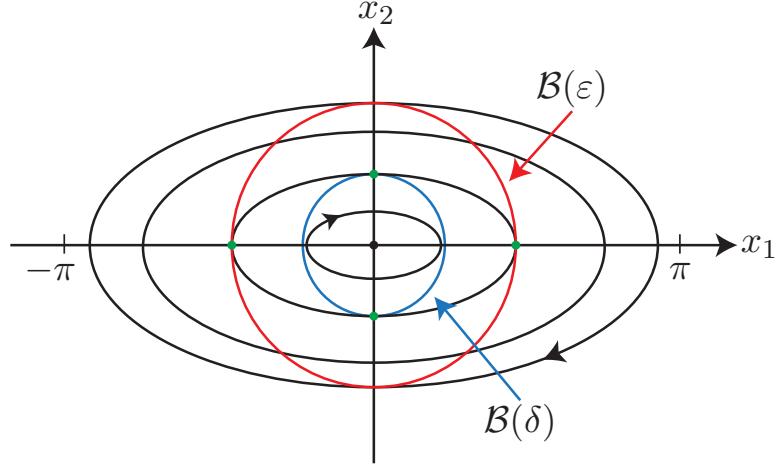


Figure 3.2: Stability of the lower equilibrium for the pendulum, here $0 < \varepsilon < \pi$.

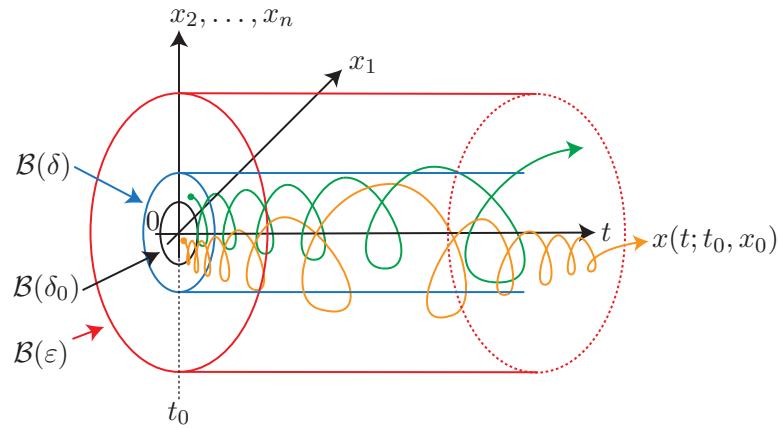


Figure 3.3: An example for an asymptotically stable fixed point (black trajectory).

Example 3.2 (Damped pendulum). We have the equation of motion with the linear damping coefficient c

$$\ddot{\varphi} + c\dot{\varphi} + \sin(\varphi) = 0, \quad c > 0. \quad (3.7)$$

Transforming into a first-order ODE with $x_1 = \varphi$ and $x_2 = \dot{\varphi}$ gives

$$\begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -cx_2 - \sin(x_1). \end{cases} \quad (3.8)$$

The total energy is given by

$$E = \frac{1}{2}x_2^2 + (1 - \cos(x_1)). \quad (3.9)$$

Further we have the rate of energy change

$$\frac{d}{dt} E(x_1(t), x_2(t)) = x_2 (\dot{x}_2 + \sin(x_1)) = -cx_2^2. \quad (3.10)$$

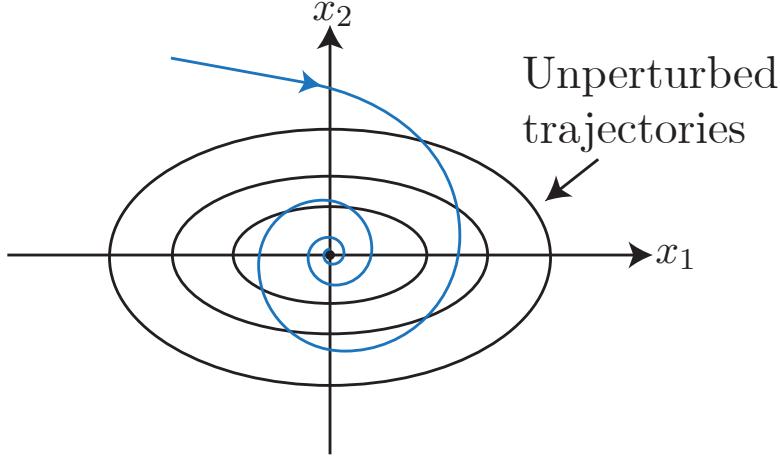


Figure 3.4: An example of a trajectory which loses energy, in this case due to damping.

Therefore, along trajectories energy decreases monotonically as shown in Fig 3.4. By the \mathcal{C}^0 dependence of the trajectory on initial conditions, the trajectories remain close to the undamped oscillations for small $c > 0$. We conclude that trajectories are inward spirals for a small dissipation $c > 0$. The fixed point $x = 0$ is still Lyapunov stable, but asymptotic stability does not yet follow (is the limit of $x(t)$ equal to 0?).

Remark 3.2 (LaSalle's invariance principle). This conclusion follows rigorously from LaSalle's invariance principle, namely if we assume that $\dot{x} = f(x)$, $f \in \mathcal{C}^1$, and that there exists a $V \in \mathcal{C}^1$ with

$$\dot{V} = \frac{dV(x(t))}{dt} \leq 0. \quad (3.11)$$

Then the set of accumulation points for any trajectory is contained in the set of trajectories that stay within the set $I = \{x \in \mathbb{R}^n : \dot{V}(x) = 0\}$.

Example 3.3. Consider the following dynamical system in polar coordinates, i.e. $r \cos(\theta) = x$ and $r \sin(\theta) = y$,

$$\begin{cases} \dot{r} = r(1 - r) \\ \dot{\theta} = \sin^2\left(\frac{\theta}{2}\right). \end{cases} \quad (3.12)$$

Note that $r = 0$ is a fixed point, the set $r = 1$ is an invariant circle, and the set $\theta = 0$ is an invariant set. An invariant set is a set such that if the dynamical system is started on the set, it remains in the set for all time. Examining the radial evolution reveals that the equation of motion decouples. We see that $\dot{\theta} \geq 0$, so rotation is either positive or null.

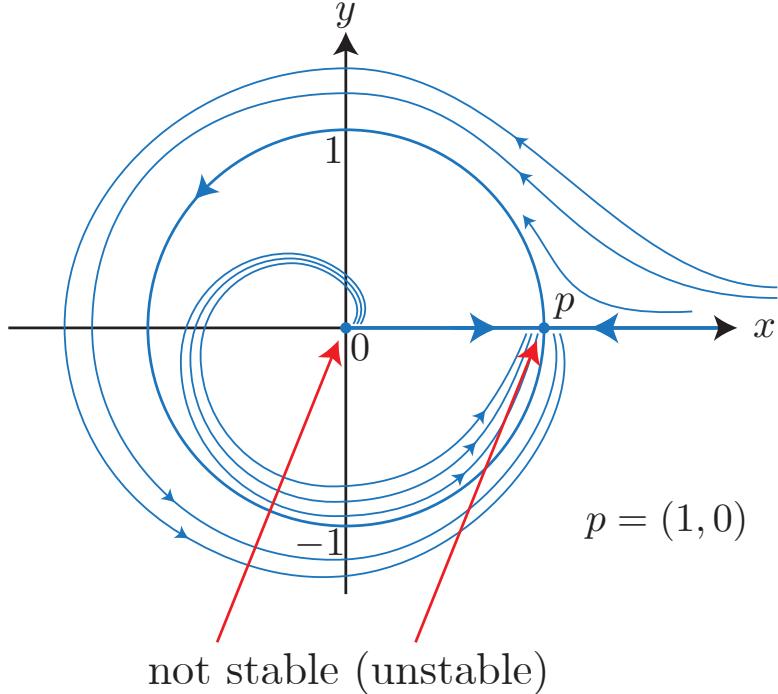


Figure 3.5: Phase portrait of the dynamical system in cartesian coordinates, with the red arrows pointing to the two unstable equilibria.

From Fig. 3.5 we can see that both of the fixed points, $(0, 0)$ and $(1, 0)$, are not stable. However, inspecting Fig. 3.5 we see that that $p = (1, 0)$ is an example of an attractor: a set with an open neighborhood of points that all approach the set as $t \rightarrow \infty$.

Definition 3.4 (Invariant set). The set $S \subset P$ is an *invariant set* for the flow map $F^t : P \rightarrow P$ if $F^t(S) = S$ for all $t \in \mathbb{R}$.

Definition 3.5 (Unstable point). A fixed point $x = 0$ is unstable if it is not stable.

Remark 3.3. We can negate a mathematical statement by using the reverse relational operators outside the statements involving these operators i.e. $\exists \rightarrow \forall$ and $\forall \rightarrow \exists$. For example we have for continuity $\forall \varepsilon \exists \delta : \|f(x) - f(y)\| < \varepsilon$ if $\|x - y\| < \delta$, meanwhile for discontinuity we have $\exists \varepsilon : \forall \delta : \|f(x) - f(y)\| \geq \varepsilon$ for $\|x - y\| < \delta$.

In our case for stability we have

$$\forall \varepsilon, t_0 : \exists \delta > 0 : \forall x_0 \text{ with } \|x_0\| < \delta : \|x(t)\| \leq \varepsilon \quad \forall t \geq t_0. \quad (3.13)$$

Meanwhile for instability

$$\exists \varepsilon, t_0 : \underbrace{\forall \delta > 0}_{\text{"for arbitrarily small"}} : \exists x_0 \text{ with } \|x_0\| < \delta : \|x(t)\| > \varepsilon \quad \underbrace{\exists t \geq t_0}_{\text{"for some}}. \quad (3.14)$$

This negation is demonstration in Fig. 3.6.

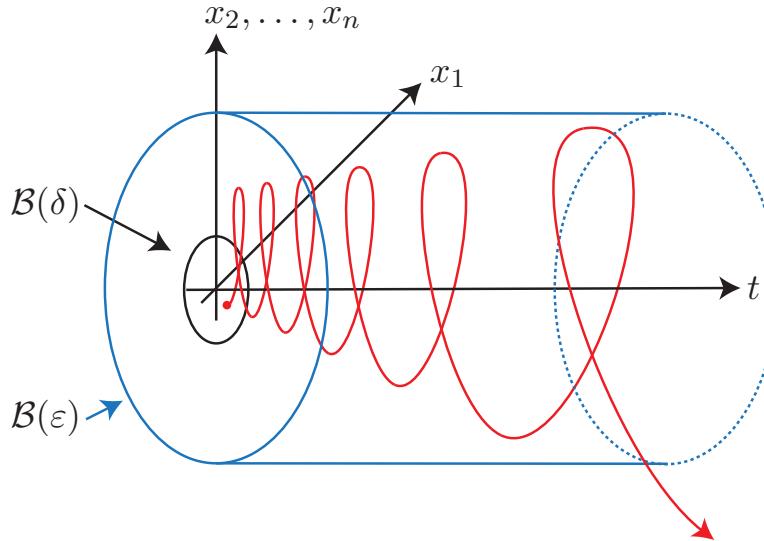


Figure 3.6: Example of an unstable fixed point, with the red trajectory representing a trajectory starting arbitrarily close to the fixed point, leaving a given ε -ball.

Remark 3.4. By C^0 dependence of trajectories on initial conditions, if $x(t; t_0, x_0)$ leaves $B(\varepsilon)$, then for \tilde{x}_0 close enough to x_0 , $x(t; t_0, \tilde{x}_0)$ also leaves $B(\varepsilon)$. Since this is true on an open set around x_0 , the measure of such trajectories is nonzero, the instability is observable!

Example 3.4 (Unstable fixed point of pendulum). In contrast, we can have that infinitely many trajectories converge to the fixed point, yet it is still unstable, as illustrated in Fig. 3.7. In fact, the converging trajectories form a measure-zero set, thus the stability near the unstable equilibrium is unobservable.

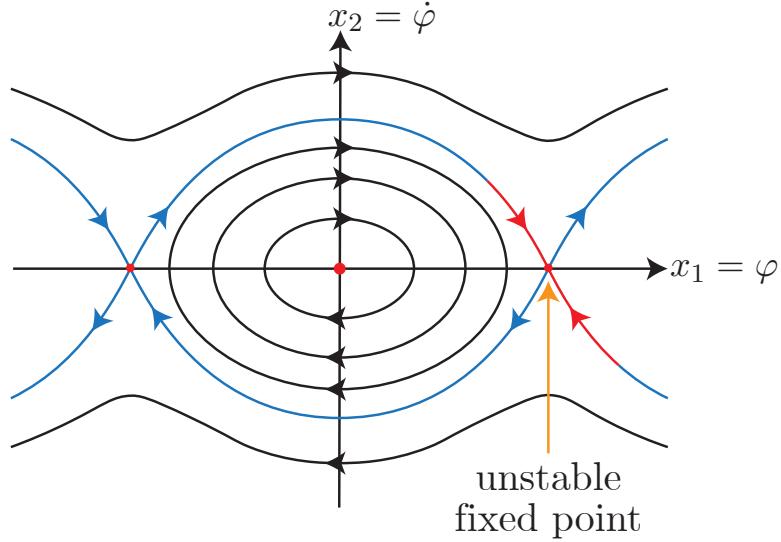


Figure 3.7: The phase portrait around the unstable fixed point of the pendulum, with the stable trajectories (red).

3.2 Stability based on linearization

We would like to derive a more general method to analyze the stability of fixed points, thus we try to simplify our system around the fixed point and discover what this can tell us about the full (unimplified) system. In the following section we shall always assume that our system is autonomous. We will have the following setup

$$\dot{x} = f(x), \quad f \in \mathcal{C}^1, \quad x = \begin{pmatrix} x_1 \\ \vdots \\ x_n \end{pmatrix} \in \mathbb{R}^n, \quad p = \begin{pmatrix} p_1 \\ \vdots \\ p_n \end{pmatrix} \in \mathbb{R}^n. \quad (3.15)$$

If \$f(p) = 0\$, then \$p\$ is a fixed point. By transforming using \$y = x - p\$, we have that in the transformed system \$y = 0\$ is a fixed point. Furthermore, we have that around \$y = 0\$ the ODE is

$$\dot{y} = f(p + y) = \underbrace{f(p)}_{=0} + Df(p)y + o(\|y\|) = Df(p)y + o(\|y\|). \quad (3.16)$$

Remark 3.5. Since we only assumed one continuous derivative for the function \$f\$ the remainder term in the Taylor approximation is \$o(\|y\|)\$. The little o notation means that

$$\lim_{\|y\| \rightarrow 0} \frac{o(\|y\|)}{\|y\|} = 0. \quad (3.17)$$

Definition 3.6 (Linearized ODE). We define the *linearization* of (3.15) at the fixed point p as

$$\boxed{\dot{y} = Ay; \quad y \in \mathbb{R}^n, \quad A := Df(p) \in \mathbb{R}^{n \times n}; \quad Df(p) = \left(\begin{array}{ccc} \frac{\partial f_1}{\partial x_1} & \dots & \frac{\partial f_1}{\partial x_n} \\ \vdots & & \vdots \\ \frac{\partial f_n}{\partial x_1} & \dots & \frac{\partial f_n}{\partial x_n} \end{array} \right) \Big|_{x=p}} \quad (3.18)$$

Now we would like to study the stability of the fixed point $y = 0$ in (3.18). From this analysis, we want to know the relevance of our results for the full nonlinear system (3.15).

3.3 Review of linear dynamical systems

Recall the setup

$$\dot{y} = A(t)y, \quad y \in \mathbb{R}^n, \quad A \in \mathbb{R}^{n \times n}, \quad A \in \mathcal{C}_t^0. \quad (3.19)$$

The following facts have already been established

- We know that the global existence and uniqueness of solutions is guaranteed.
- The superposition principle holds; namely the linear combination of solutions is also a solution.
- There exists a set of n linearly independent solutions: $\varphi_1(t), \dots, \varphi_n(t) \in \mathbb{R}^n$.
- The general solution is

$$y(t) = \sum_{i=1}^n c_i \varphi_i(t) = \underbrace{\begin{bmatrix} \varphi_1(t) & \dots & \varphi_n(t) \end{bmatrix}}_{\Psi(t): \text{ fundamental matrix solution}} \underbrace{\begin{bmatrix} c_1 \\ \vdots \\ c_n \end{bmatrix}}_c = \Psi(t)c; \quad \dot{\Psi} = A(t)\Psi. \quad (3.20)$$

- We have the initial value problem $y(t_0) = y_0$ which implies

$$\Psi(t_0)c = y_0 \implies y(t) = \underbrace{\Psi(t) [\Psi(t_0)]^{-1}}_{\phi(t) := F_{t_0}^t} y_0 \quad (3.21)$$

Where we used that the $\varphi_i(t)$ are linearly independent in the last equality. And we have the *normalized fundamental matrix* $\phi(t)$ equal to the flow map, with $\phi(t_0) = I$.

- In the autonomous case $\dot{x} = Ax$ solutions can be practically constructed.

(i) **Explicit Solution**

$$\phi(t) = e^{At} := \sum_{j=0}^{\infty} \frac{1}{j!} (At)^j \quad (3.22)$$

With $0! = 1$ and $0^0 = I$. We can verify that this is indeed a solution

$$\dot{\phi}(t) = \sum_{j=1}^{\infty} \frac{1}{(j-1)!} (At)^{j-1} A = A \sum_{j=0}^{\infty} \frac{1}{j!} (At)^j = Ae^{At} = A\phi(t). \quad (3.23)$$

Where we used that A commutes with its powers in the second equality. We now have that each column of $\phi(t)$ satisfies $\dot{y} = Ay$.

Remark 3.6. For a scalar ODE $\dot{y} = a(t)y$ for $y \in \mathbb{R}$, the solution is known $y(t) = e^{\int_{t_0}^t a(s)ds} y_0$. However, this does not extend to the higher dimensional $\dot{y} = A(t)y$. In fact, in general, $\phi(t) = e^{\int_{t_0}^t A(s)ds}$ is not a solution. We can check this

$$\dot{\phi} = \sum_{j=0}^{\infty} \frac{1}{j!} \frac{d}{dt} \left(\int_{t_0}^t A(s)ds \right)^j = \sum_{j=1}^{\infty} \frac{1}{(j-1)!} \left(\int_{t_0}^t A(s)ds \right)^{j-1} A(t) \neq A(t)\phi(t). \quad (3.24)$$

The nonequality holds as $A(t)$ does not generally commute with $\int A(s)ds$.

- (ii) **Solution from eigenfunctions** If we have an autonomous system, we can solve the ODE without an infinite series. We have

$$\dot{y} = Ay, \quad y \in \mathbb{R}^n, \quad y(0) = y_0. \quad (3.25)$$

Substituting $\varphi(t) = e^{\lambda t}s$ for $\lambda \in \mathbb{C}$ and $s \in \mathbb{C}^n$ into (3.25) yields

$$\lambda s = As \implies (A - \lambda I)s = 0 \iff \det(A - \lambda I) = 0. \quad (3.26)$$

Therefore λ must be an eigenvalue of A and s must be the corresponding eigenvector. We call $\det(A - \lambda I)$ the *characteristic equation* of A . Let $\lambda_1, \dots, \lambda_n$ be the eigenvalues and s_1, \dots, s_n be the corresponding eigenvectors. In the case that some eigenvalues are repeated, some of the s_i may be generalized eigenvectors. We then have two cases.

- (a) A is semisimple, i.e. the eigenvectors are linearly independent (which is always the case if the λ_i all have algebraic multiplicity of one). Then we have the solution

$$y(t) = \sum_{i=1}^n c_i e^{\lambda_i t} s_i = \sum_{j=1}^n c_j e^{(\operatorname{Re}\lambda_j)t} e^{i(\operatorname{Im}\lambda_j)t} s_j. \quad (3.27)$$

Where we used $\lambda_j = \operatorname{Re}\lambda_j + i\operatorname{Im}\lambda_j$.

- (b) A is not semisimple, i.e. has repeated eigenvalues (but not enough linearly independent eigenvectors). Then we assume that λ_k has algebraic multiplicity $a_k > g_k$, where a_k measures the multiplicity of λ_k as a root of $\det(A - \lambda I) = 0$, and g_k is the number of linearly independent eigenvectors for λ_k , also called the *geometric multiplicity* of λ_k . Even in this case, λ_k gives rise to a_k linearly independent solutions of the form

$$\underbrace{P_0}_{=s_k} e^{\lambda_k t}, P_1(t)e^{\lambda_k t}, P_2(t)e^{\lambda_k t}, \dots, P_{a_k-1}(t)e^{\lambda_k t} \quad (3.28)$$

where $P_j(t)$ is a vector polynomial of t of order j or less.

3.3.1 Evolution of observable functions

A common narrative originating from the operator-theoretic approach to dynamical systems, is that *all dynamical systems become linear when rewritten in terms of observables*. Let us consider the dynamical system given by

$$\dot{x} = f(x), \quad x \in \mathbb{R}^n, \quad (3.29)$$

and also the observable function defined on the phase space \mathbb{R}^n

$$g : \mathbb{R}^n \rightarrow \mathbb{R}^k. \quad (3.30)$$

To evaluate the function along trajectories of the dynamical system (3.29) we can take the composition of g and the flowmap F^t , which, for an initial condition $x_0 \in \mathbb{R}^n$ returns

$$g(x(t, x_0)) = g \circ F^t(x_0). \quad (3.31)$$

We can introduce the Koopman operator acting on observable functions as

$$\mathcal{K}^t : \mathcal{C}^1 \rightarrow \mathcal{C}^1, \quad (3.32)$$

realizing the composition with the flowmap, i.e.

$$\mathcal{K}^t : g(x_0) \mapsto [\mathcal{K}^t g](x_0) = g \circ F^t(x_0) = g(x(t, x_0)). \quad (3.33)$$

Note that the Koopman operator defined as (3.33) is linear. Indeed, its action on a linear combination of observable functions $\alpha_1 g_1 + \alpha_2 g_2$ is simply

$$[\mathcal{K}^t(\alpha_1 g_1 + \alpha_2 g_2)](x_0) = [\alpha_1 g_1 + \alpha_2 g_2] \circ F^t(x_0) \quad (3.34)$$

$$= \alpha_1 g_1(x(t, x_0)) + \alpha_2 g_2(x(t, x_0)) \quad (3.35)$$

$$= \alpha_1(g_1 \circ F^t) + \alpha_2(g_2 \circ F^t) \quad (3.36)$$

$$= \alpha_1 [\mathcal{K}^t g_1](x_0) + \alpha_2 [\mathcal{K}^t g_2](x_0). \quad (3.37)$$

The Koopman-operator thus is a linear operator acting on \mathcal{C}^1 , i.e. the space of functions, which is infinite-dimensional. A common narrative is, that the difficulty of solving the finite-dimensional nonlinear system can be sidestepped by investigating the dynamics of observables. The observables are advected by the linear Koopman operator, which gives hope that the dynamics of these observables may be also linear. This statement, however, is false.

Let us derive the evolution of equation of an observable. Differentiating the action of the Koopman operator, we get

$$\frac{d}{dt} [\mathcal{K}^t g] (x_0) = D_x g(x(t, x_0)) f(x(t, x_0)) = \mathcal{K}^t [D_x g] (x_0) [\mathcal{K}^t f] (x_0). \quad (3.38)$$

The operator on the right-hand side $\mathcal{K}^t [D_x g] (x_0)$, is not a linear operator on $\mathcal{K}^t g$. Therefore, the time evolution of $\mathcal{K}^t g$ is not governed by a linear operator, i.e. it is not of the form

$$\mathcal{K}^t g = \Psi(t) \mathcal{K}^0 g = \Psi(t) g, \quad (3.39)$$

for some $\Psi(t)$ operator.

The exception is, when g is chosen to be an eigenfunction of the Koopman operator. In this case, we may write

$$\mathcal{K}^t g_0 = e^{\lambda t} g_0, \quad \text{for some eigenvalue } \lambda \in \mathbb{C}. \quad (3.40)$$

In this case, the time evolution of g_0 is indeed linear, since

$$\frac{d}{dt} \mathcal{K}^t g_0 = \lambda \mathcal{K}^t g_0. \quad (3.41)$$

However, constructing such Koopman eigenfunctions is generally a very difficult task. In addition, g_0 may not even exist, or the spectrum of \mathcal{K}^t may be continuous.

3.4 Stability of fixed points in autonomous linear systems

First we note that we can bound our solution

$$\|y(t)\| = \|\phi(t)y_0\| \leq \underbrace{\|\phi(t)\|}_{\text{Operator norm}} \|y_0\| \leq C e^{\mu t} \|y_0\|. \quad (3.42)$$

Where $\mu = \max_j (\operatorname{Re} \lambda_j) + \nu$, with $\nu > 0$, as small as needed, provided we increase C appropriately. If A is semisimple, then $\nu = 0$ can be selected.

Theorem 3.7 (Stability of fixed points in linear systems). *Given $y = 0$ a fixed point of the linear system $\dot{y} = Ay$ with $A \in \mathbb{R}^{n \times n}$ the following statements hold:*

- (i) Assume that $\operatorname{Re}\lambda_j < 0$ for all j . Then $y = 0$ is asymptotically stable.
- (ii) Assume that $\operatorname{Re}\lambda_j \leq 0$ for all j , and for all λ_k with $\operatorname{Re}\lambda_k = 0$ we have $a_k = g_k$. Then $y = 0$ is stable.
- (iii) Assume there exists a k such that $\operatorname{Re}\lambda_k > 0$. Then $y = 0$ is unstable.

These scenarios are illustrated in Fig. 3.8.

Proof. (i) Pick $\varepsilon > 0$, and select $\nu > 0$ small, such that $\mu < 0$. Then pick $C > 0$ such that (3.42) holds, and let $\delta = \frac{\varepsilon}{C}$. This implies (since $\|y_0\| \leq \delta$) that

$$\|y(t)\| \leq \varepsilon e^{\mu t} \leq \varepsilon, \quad (3.43)$$

and

$$\|y(t)\| \leq \varepsilon e^{\mu t} \xrightarrow{t \rightarrow \infty} 0. \quad (3.44)$$

Where the limit holds as $\mu < 0$.

- (ii) Again choose $\delta = \frac{\varepsilon}{C}$ and note that $\mu = \max_j(\operatorname{Re}\lambda_j) + \nu = 0 + \nu = 0$ ($\nu = 0$ as $a_k = g_k$). Then stability follows by (3.42). However, asymptotic stability does not hold, as $\varphi(t) = Ce^{i(\operatorname{Im}\lambda_j)t}$ solutions exist.
- (iii) There exists a solution of the form

$$\varphi(t) = C_k e^{\lambda_k t} s_k = C_k e^{(\operatorname{Re}\lambda_k)t} e^{i(\operatorname{Im}\lambda_k)t} s_k. \quad (3.45)$$

In turn this implies

$$\|\varphi(t)\| = C_k e^{(\operatorname{Re}\lambda_k)t} \|s_k\| \xrightarrow{t \rightarrow \infty} \infty. \quad (3.46)$$

□

Example 3.5 (Stability analysis of 2 degrees of freedom coupled oscillators). Given a rectangular mass m with a spring of stiffness coefficient k attached to each side extending to fixed walls in each cardinal direction. We want to know the stability of the equilibrium where all of the springs are equally extended. This dynamical system is depicted in Fig. 3.9. First, note that this is a conservative system, i.e. $E = \text{const}$. Next we transform the coordinates so that the equations of motion can be brought into the form of an ODE for this dynamical system

$$x = \begin{pmatrix} x \\ \dot{x} \\ y \\ \dot{y} \end{pmatrix}. \quad (3.47)$$

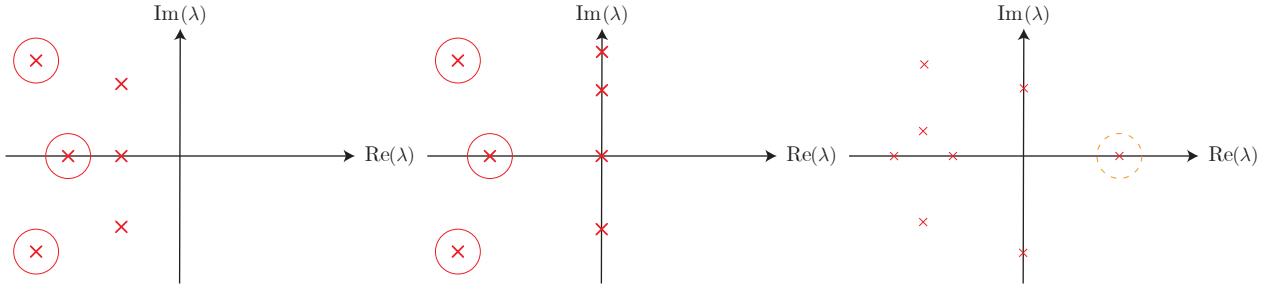


Figure 3.8: Eigenvalue arrangements for scenarios (i), (ii), and (iii) (from left to right) in Theorem 3.7

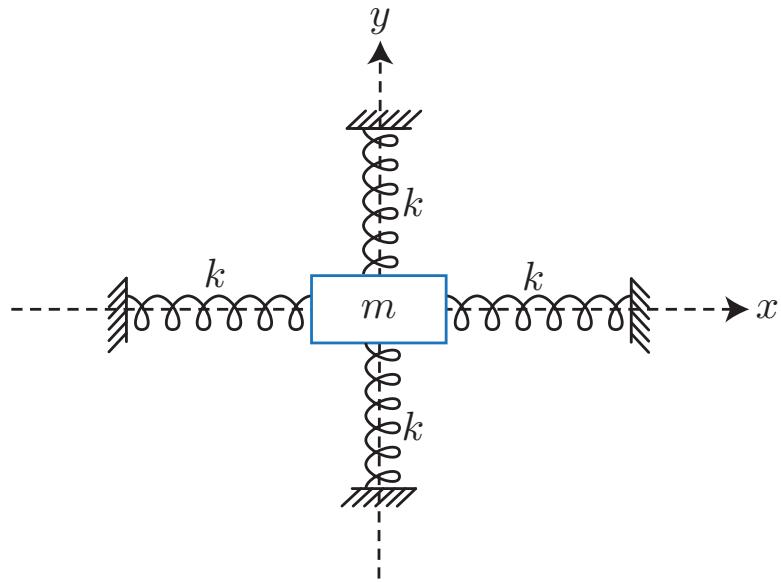


Figure 3.9: Arrangement of coupled oscillators with rectangular mass in the middle.

Thus we have a 4-dimensional, nonlinear, system of ODEs. We now linearize this at the fixed point $(x, y) = (0, 0)$, i.e. $\dot{x} = Ax$ with $x \in \mathbb{R}^n$.

The system exhibits full spatial symmetry in x and y , hence the eigenmodes will be the same in the x and y directions. This means we have repeated pairs of purely imaginary eigenvalues for A : $\lambda_{1,2} = \lambda_{3,4} = \pm i\omega$. It is clear that scenarios (i) and (iii) of Theorem 3.7 do not apply to the linearized ODE. So we need to check if (ii) applies.

We have that $\text{Re}\lambda_k = 0$ for $k = 1, 2, 3, 4$. Also $a_k = 2$ for $k = 1, 2, 3, 4$. Now assume $g_k < 2$. Then there would exist solutions of the form $te^{\pm i\omega t}s_k$, but this would contradict the conservation of energy, as either the (nonnegative) kinetic energy and/or the (nonnegative) potential energy would grow unbounded. Hence, the total energy could not be conserved.

Therefore we know that $g_k = a_k$ and we can apply (ii) to find $x = y = 0$ is Lyapunov stable for the linearized system. What does this imply for the nonlinear system?

Notebook 3.1. *Comparison of two-dimensional phase portraits for various eigenvalue configurations.*

Exercise 3.1. Consider a discrete dynamical system given by the iterated mapping

$$x_{k+1} = f(x_k); \quad f : \mathbb{R}^n \rightarrow \mathbb{R}^n, \quad x_k \in \mathbb{R}^n. \quad (3.48)$$

Assume that $x = p$ is a fixed point for the mapping, i.e., $p = f(p)$.

(a) Derive a linearized mapping of the form

$$y_{k+1} = Ay_k \quad (3.49)$$

to describe the discrete dynamics in the vicinity of the fixed point.

(b) Assume that A has eigenvalues $\lambda_1, \dots, \lambda_n \in \mathbb{C}$ with corresponding n linearly independent eigenvectors $s_1, \dots, s_n \in \mathbb{C}^n$. Show that the general solution of (3.49) is of the form

$$y_k = c_1\varphi_1(k) + \dots + c_n\varphi_n(k), \quad \varphi_i(k) = \lambda_i^k s_i. \quad (3.50)$$

(c) Formulate a definition of stability, asymptotic stability, and instability for the $y = 0$ fixed point of (3.49).

(d) Using (3.50), find a sufficient and necessary condition for the asymptotic stability you have defined in (c).

Exercise 3.2. Consider a ball of mass m that slides on a rotating hoop (see Fig. 3.10).

The angular velocity of the hoop is Ω , the viscous friction coefficient between the hoop and the ball is b , and the constant of gravity is g . The equation of motion for the sliding ball is given by

$$mR^2\ddot{\alpha} + bR^2\dot{\alpha} + mR^2(g/R - \Omega^2 \cos(\alpha)) \sin(\alpha) = 0. \quad (3.51)$$

(a) Plot the location of equilibria of the ball as a function of the non-dimensionalized rotation parameter $\nu = R\Omega^2/g$.

(b) Using linearization, determine the stability type of the different equilibrium branches on the plot. Identify the critical angular velocity at which a bifurcation of equilibria occurs.

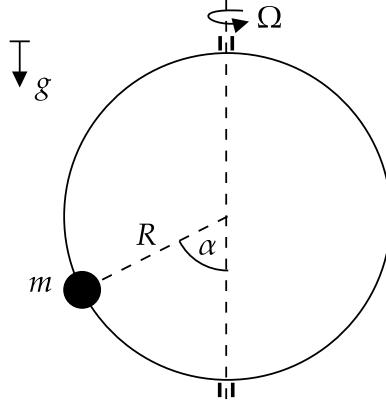


Figure 3.10: Mass on a hoop

3.5 Stability of fixed points in nonlinear systems

Following the previous example, we would like to know what information about the stability of fixed points of nonlinear systems we can derive from the linearized system. The full nonlinear system is

$$\dot{x} = f(x), \quad f(x_0) = 0, \quad x \in \mathbb{R}^n, \quad f \in \mathcal{C}^1. \quad (3.52)$$

And its linearization at the fixed point x_0

$$\dot{y} = Df(x_0)y, \quad y \in \mathbb{R}^n, \quad Df(x_0) \in \mathbb{R}^{n \times n}. \quad (3.53)$$

We would like to conclude that the linearized dynamics are qualitatively similar to the nonlinear dynamics. In order to study if this is the case, we have to formalize *similar* mathematically.

Definition 3.7 (\mathcal{C}^k equivalence of dynamical systems). Consider two autonomous dynamical systems:

(i)

$$\dot{x} = f(x), \quad x \in \mathbb{R}^n, \quad f \in \mathcal{C}^1; \quad F^t : x_0 \mapsto x(t; x_0). \quad (3.54)$$

(ii)

$$\dot{x} = g(x), \quad x \in \mathbb{R}^n, \quad g \in \mathcal{C}^1; \quad G^t : x_0 \mapsto x(t; x_0). \quad (3.55)$$

The two dynamical systems are \mathcal{C}^k *equivalent*, for $k \in \mathbb{N}$, on an open set $U \subset \mathbb{R}^n$, if there exists a \mathcal{C}^k diffeomorphism $h : U \rightarrow U$ that maps orbits of (i) into orbits of (ii), while preserving the orientation but not necessarily the exact parameterization of the orbit by time. Specifically for all $x \in U$, any $t_1 \in \mathbb{R}$ there exists a $t_2 \in \mathbb{R}$ such that

$$h(F^{t_1}(x)) = G^{t_2}(h(x)). \quad (3.56)$$

$h : U \rightarrow U$ does this for all $x \in U$ in a \mathcal{C}^k fashion. This equivalence through a function h is demonstrated in Fig 3.11.

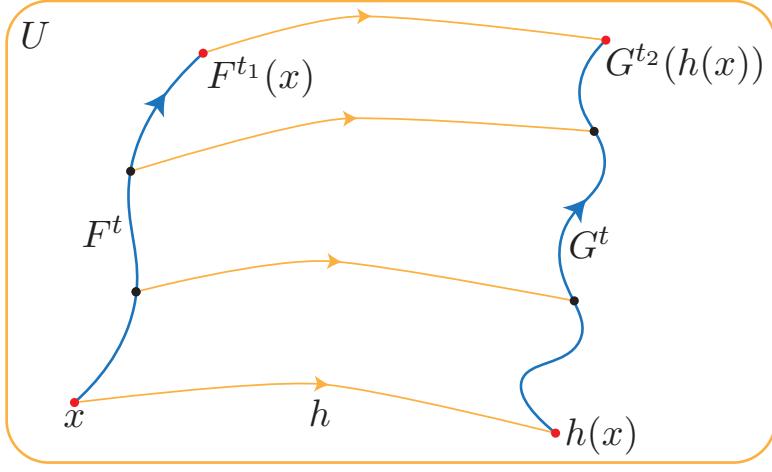


Figure 3.11: The function h mapping the orbits of the dynamical system describing F into the system describing G .

Proposition 3.8 (Restrictiveness of smooth equivalence). *Consider two \mathcal{C}^k -equivalent dynamical systems for $k > 0$ with the right hand sides f and g , as in the definitions above. Let x_0 be a fixed point of F and y_0 a fixed point of G such that these correspond to each other under the \mathcal{C}^k equivalence, i.e. $y_0 = h(x_0)$. The eigenvalues of $Df(x_0)$ must be constant, positive multiples of the eigenvalues of $Dg(y_0) = Dg(h(x_0))$.*

Proof. By the \mathcal{C}^k equivalence we have that there exists a $\tau(x, t)$ with $\frac{d\tau}{dt} > 0$ such that $h(F^t(x)) = G^{\tau(x,t)}(h(x))$. Further assume that τ is at least continuously differentiable in its arguments. Differentiating the previous equality with respect to x yields

$$DhDF^t = \dot{G}^\tau \frac{d\tau}{dx} + DG^\tau Dh. \quad (3.57)$$

At a fixed point x_0 we have that $\dot{G}^\tau(x_0) = g(G^\tau(x_0)) = 0$. This implies that the first term on the right hand side of the differentiated equality is equal to 0. Now denote the linearizations of

f and g at the fixed point by $A = Df(x_0)$ and $B = Dg(h(x_0))$ respectively. Then the matrices e^{At} and $e^{B\tau}$ must be similar matrices, in turn implying that At and $B\tau(x_0, t)$ must have the same spectrum, i.e.

$$\frac{\lambda_i(B)}{\lambda_i(A)} = \frac{t}{\tau(x_0, t)} = C = \text{const.} > 0. \quad (3.58)$$

Therefore the eigenvalues of the linearizations are constant, positive multiples of each other. \square

Definition 3.8 (\mathcal{C}^k conjugacy). Consider the same two dynamical systems as before. The two dynamical systems are \mathcal{C}^k *conjugate* if there exists a \mathcal{C}^k diffeomorphism $h : U \rightarrow U$ that maps orbits of (i) into (ii), while preserving orientation and the parameterization of the orbit by time. Specifically for all $x \in U$ and $t \in \mathbb{R}$ we have

$$h(F^t(x)) = G^t(h(x)). \quad (3.59)$$

Proposition 3.9 (Restrictiveness of smooth conjugacy). *Consider two \mathcal{C}^k conjugate systems, as in the definition above. In addition to the previous proposition, now the linearizations of the two systems will have the same spectra at each fixed point.*

Proof. The \mathcal{C}^k conjugacy requires preserving the parameterizations of the orbits, i.e. $\tau(x, t) = t$. Hence, this is just a specification of the previous proof. \square

Definition 3.9 (Topological equivalence). For $k = 0$, \mathcal{C}^k equivalence is also called *topological equivalence*. In this case, a continuous, invertible deformation takes orbits of one system into the orbits of the other. Under these conditions, $h : U \rightarrow U$ is called a *homeomorphism*.

Example 3.6 (Topologically equivalent linear systems for $n = 2$). To illustrate the meaning of topological equivalence, Fig. 3.12 shows three linear systems ($\dot{x} = Ax$ for $x \in \mathbb{R}^2$) which are topologically equivalent.

The stable spiral has the eigenvalues $\lambda_{1,2} = \alpha \pm i\beta$ for $\alpha < 0$ and $\beta \neq 0$. The sink has the eigenvalues $\lambda_1 = \lambda_2 < 0$. and The stable node has the eigenvalues $\lambda_1 < \lambda_2 < 0$. Note here that the number of eigenvalues λ_i with $\text{Re}\lambda_i < 0$, $\text{Re}\lambda_i = 0$, and $\text{Re}\lambda_i > 0$ is the same in all three of these cases, namely for each system the real part of both eigenvalues are less than 0.

Remark 3.10. The above systems are not \mathcal{C}^k equivalent or conjugate, as the spectra differ by more than a multiplicative constant.

Example 3.7 (Topologically inequivalent linear systems for $n = 2$). As a counter example, we now present three linear systems which are not topologically equivalent. The stable spiral (from before), the unstable spiral, and the saddle. The unstable spiral has the eigenvalues $\lambda_{1,2} = \alpha \pm i\beta$

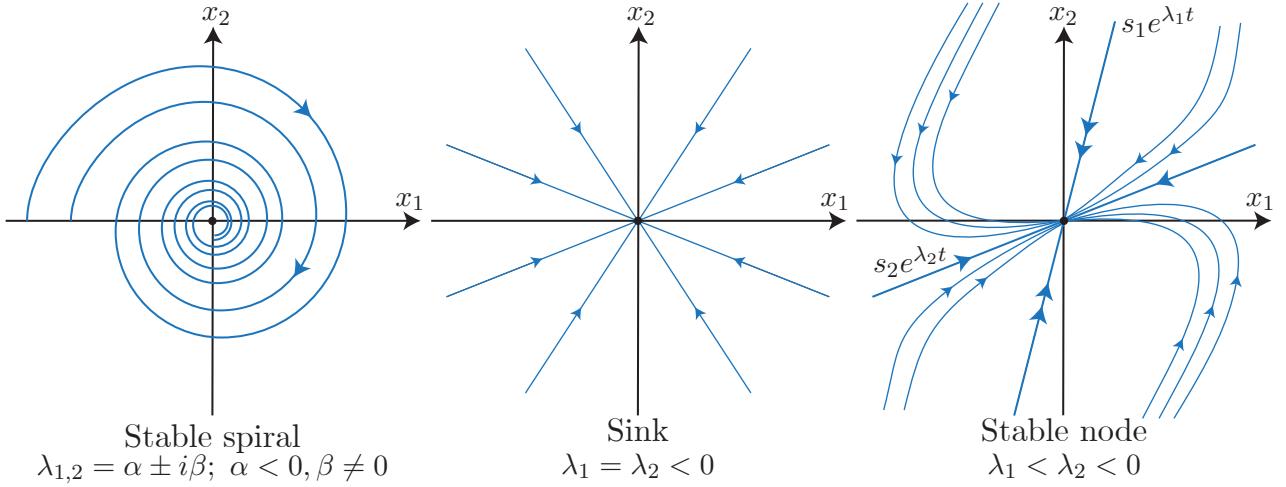


Figure 3.12: Three topologically equivalent 2-dimensional linear systems. Left: The stable spiral. Middle: The sink. Right: The stable node.

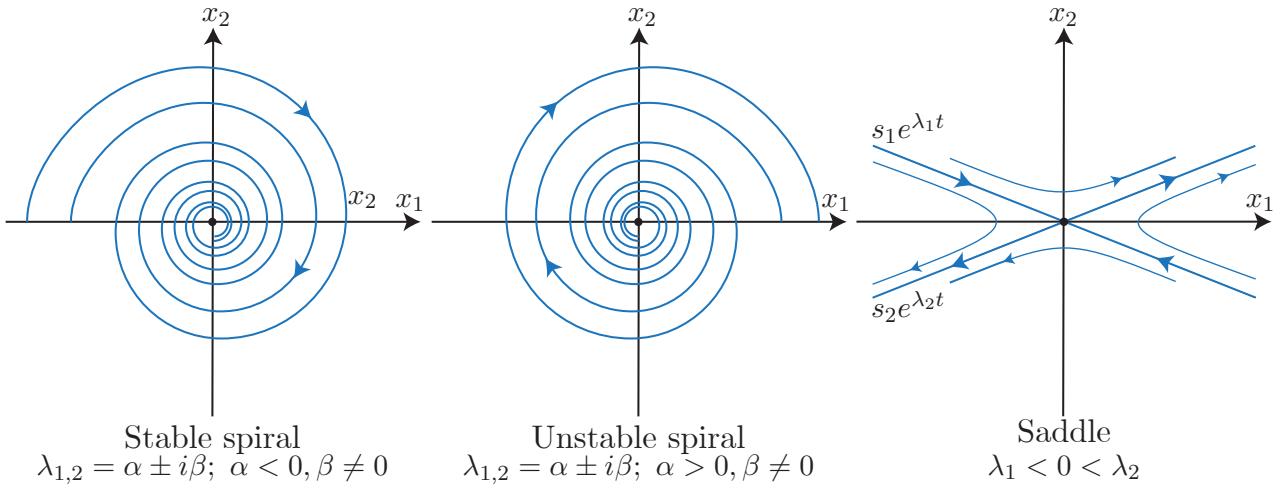


Figure 3.13: Three 2-dimensional linear systems which are not topologically equivalent. Left: The stable spiral. Middle: The unstable spiral. Right: The saddle.

for $\alpha > 0$ and $\beta \neq 0$ (note the different sign for α). The saddle has the eigenvalues $\lambda_1 < 0 < \lambda_2$. These systems are depicted in Fig 3.13.

Note here that the eigenvalue configurations in terms of the number of λ_i with real part less than 0 are different in each case. Building on the role of the eigenvalue configuration we noted in the previous examples, we introduce the concept of a hyperbolic fixed point.

Definition 3.10 (Hyperbolic fixed point). We call the fixed point $x = x_0$ a *hyperbolic fixed*

point of (3.52) if each of the eigenvalues λ_i of its linearization (3.53) satisfy

$$\boxed{\operatorname{Re}\lambda_i \neq 0.} \quad (3.60)$$

Geometrically the eigenvalue configuration of a hyperbolic fixed point is shown in Fig. 3.14.

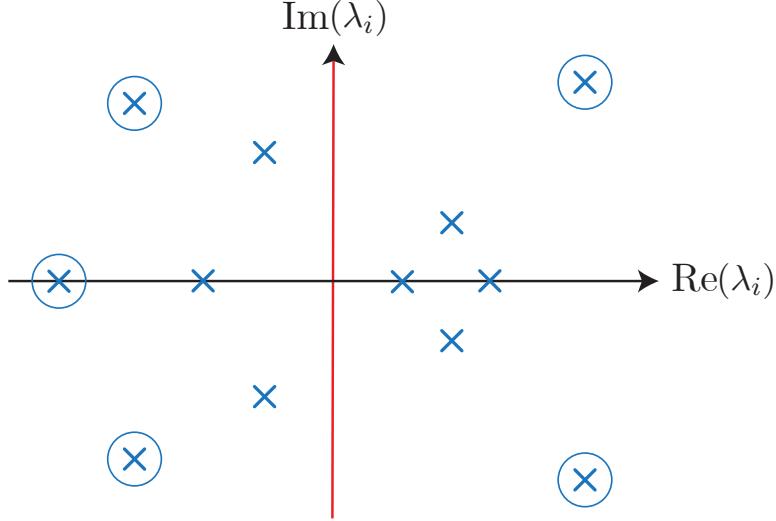


Figure 3.14: The eigenvalue configuration of a hyperbolic fixed point, i.e. no eigenvalues are on the imaginary axis (red).

Proposition 3.11. *Under small perturbations to the nonlinear system, the linearized stability type of the hyperbolic fixed point is preserved.*

Before proving this result, recall the Implicit Function Theorem (without proof).

Theorem 3.12 (Implicit Function Theorem ($n+1$ dimensional case)). *For a function $F : \mathbb{R}^{n+1} \rightarrow \mathbb{R}^n$ which is C^1 , if $F(x_0, y_0) = 0$ and the Jacobian $D_x F(x_0, y_0)$ is nonsingular (invertible), then there exists a nearby solution to $F(x, y) = 0$, for $x_y = x_0 + \mathcal{O}(|y - y_0|)$. Further x_y is as smooth in y as $F(x, y)$.*

Proof (Proposition). Add a small perturbation to (3.54) i.e.

$$\dot{x} = f(x) + \varepsilon g(x); \quad |\varepsilon| \ll 1, \quad f(x_0) = 0. \quad (3.61)$$

Now we ask if the perturbed system has a fixed point x_ε near x_0 . We frame this in terms of the implicit function theorem

$$F(x, \varepsilon) = f(x) + \varepsilon g(x) \stackrel{?}{=} 0; \quad F(x_0, 0) = 0; \quad x \in \mathbb{R}^n, \quad F : \mathbb{R}^{n+1} \rightarrow \mathbb{R}^1. \quad (3.62)$$

We check that $D_x F(x_0, 0)$ is nonsingular exactly when $Df(x_0)$ is; this is fulfilled as we have no zero eigenvalues. The linearization at the perturbed fixed point takes the form

$$\dot{y} = D [f(x) + \varepsilon g(x)]|_{x=x_\varepsilon} y = [Df(x_0 + \mathcal{O}(\varepsilon)) + \varepsilon Dg(x_0 + \mathcal{O}(\varepsilon))] y = \underbrace{[Df(x_0) + \mathcal{O}(\varepsilon)]}_{{=A_\varepsilon}} y. \quad (3.63)$$

In the last equality we used the Taylor expansion in ε . We have that the roots of $\det(A_\varepsilon - \lambda I) = 0$ depend continuously on the parameter ε . These roots correspond to the eigenvalues of A_ε . Therefore, the roots stay within an $\mathcal{O}(\varepsilon)$ neighborhood of the eigenvalues of $Df(x_0)$ (see Fig. 3.15). Hence we have that the eigenvalue configuration is unchanged for small enough ε . \square

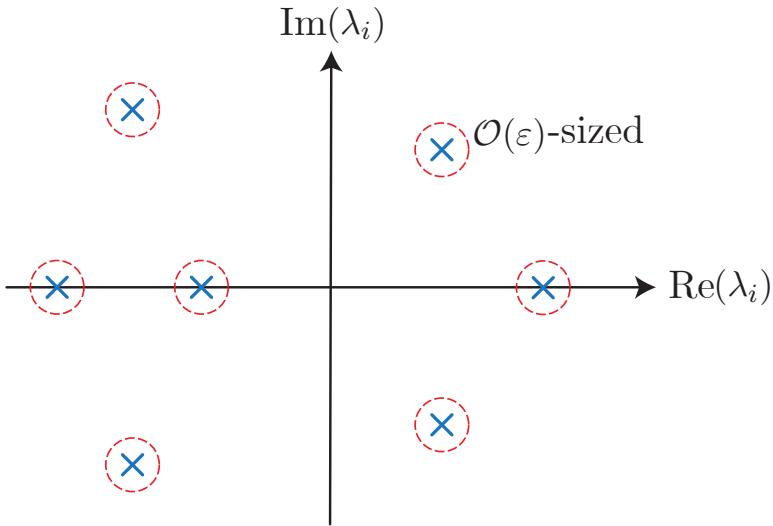


Figure 3.15: The eigenvalue configuration the $\mathcal{O}(\varepsilon)$ neighborhood (red) drawn around each eigenvalue (blue).

Remark 3.13. In the above proof, not only does the hyperbolicity of fixed points remain preserved, but also the stability type.

Meanwhile, for nonhyperbolic fixed points, this is not the case, and the smallest perturbation may change their stability type. This is due to the fact, that no matter how small the scale of the perturbation (ε) the $\mathcal{O}(\varepsilon)$ ball around eigenvalues on the imaginary axis will always intersect with $\mathbb{C} - \{\text{Im } \lambda_i = 0\}$ (i.e. points which are not on the imaginary axis).

Now we would like to connect the preservation of stability type under nonlinear perturbation to analyzing the stability type of fixed points of nonlinear dynamical systems based on their linearization.

Theorem 3.14 (Hartman-Grobman). *If the fixed point x_0 of the nonlinear system (3.52) is hyperbolic, then the linearization (3.53) is topologically equivalent to the nonlinear system in a neighborhood of x_0 .*

Consequence: *For hyperbolic fixed points, linearization predicts the correct stability type and orbit geometry near x_0 .*

Now we would like to apply this to the pendulum to systematically derive the stability type of its fixed points.

Example 3.8 (Stability analysis of the pendulum via Hartman-Grobman). Recall the transformed ODE for the pendulum

$$\dot{x} = f(x) = \begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -\sin(x_1) \end{cases}; \quad x = \begin{pmatrix} x_1 \\ x_2 \end{pmatrix}. \quad (3.64)$$

We have two fixed points $p = (\pi, 0)$ and $q = (0, 0)$. First we analyze the stability of the fixed point p . The differential of f at a point a is

$$Df(a) = \begin{pmatrix} 0 & 1 \\ -\cos(a_1) & 0 \end{pmatrix}. \quad (3.65)$$

Start by linearizing at p

$$\dot{y} = Ay; \quad A = Df(p) = \begin{pmatrix} 0 & 1 \\ -\cos(x_1) & 0 \end{pmatrix}_{x=p} = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}. \quad (3.66)$$

Now we have to check if p is hyperbolic

$$\det(A - \lambda I) = \lambda^2 - 1 = 0 \implies \lambda_{1,2} = \pm 1; \quad s_1 = \begin{pmatrix} -1 \\ 1 \end{pmatrix}, \quad s_2 = \begin{pmatrix} 1 \\ 1 \end{pmatrix}. \quad (3.67)$$

Neither of the eigenvalues lie on the imaginary axis, so p is hyperbolic. This allows us to move between the nonlinear and linearized system for the stability analysis without compromising our results. We find the linearized dynamics to be

$$y(t) = C_1 e^{-t} s_1 + C_2 e^t s_2. \quad (3.68)$$

Using the initial conditions y_0 the trajectory can be expressed as

$$y(t) = F^t y_0, \quad (3.69)$$

where F^t is the normalized fundamental matrix solution. We can now fully describe the phase portrait of the linearization. The *stable subspace* E^S is $\text{span}\{s_1\} = \{y_0 : F^t y_0 \xrightarrow{t \rightarrow \infty} 0\}$. The

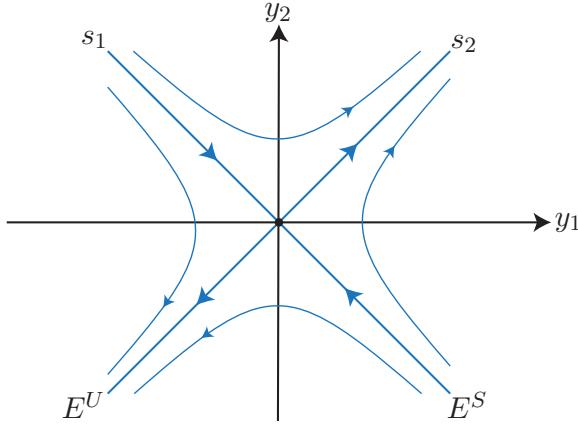


Figure 3.16: The phase portrait of the linearized pendulum in a neighborhood around p .

unstable subspace E^U is $\text{span}\{s_2\} = \{y_0 : F^t y_0 \xrightarrow{t \rightarrow -\infty} 0\}$. The phase portrait near the fixed point p is illustrated in Fig. 3.16

The nonlinear phase portrait is topologically equivalent to the linear one. Further we can define the stable and unstable manifolds of p for the nonlinear system. We designate $F^t(\cdot)$ to be the flow map for the nonlinear system after this point. The *stable manifold* of p is

$$W^S = \{x_0 : F^t(x_0) \xrightarrow{t \rightarrow \infty} p\}. \quad (3.70)$$

and the *unstable manifold* of p

$$W^U = \{x_0 : F^t(x_0) \xrightarrow{t \rightarrow -\infty} p\}. \quad (3.71)$$

Both of these are C^0 curves through p and their existence follows from the Hartman-Grobman theorem. These manifolds are shown in the nonlinear phase portrait around p in Fig. 3.17.

Next we analyze the stability of the fixed point q . Once again, our first step is to linearize

$$\dot{y} = Ay, \quad A = Df(q) = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}; \quad \det(A - \lambda I) = 0. \quad (3.72)$$

From here, we see that the determinant is equal $\lambda^2 + 1 = 0$, yielding the roots $\lambda_{1,2} = \pm i$, i.e. the fixed point is not hyperbolic. Thus the linearized dynamics is inconclusive for the nonlinear system. In this particular case, q turns out to be stable by the definition of Lyapunov stability. Later we will use another approach to show this directly.

In the last example, the importance of hyperbolicity was not accentuated, as the latter fixed point had the same stability type in the linearized system as in the full system. This leads us to question if there are cases where the stability type between the linear and nonlinear systems is not preserved.

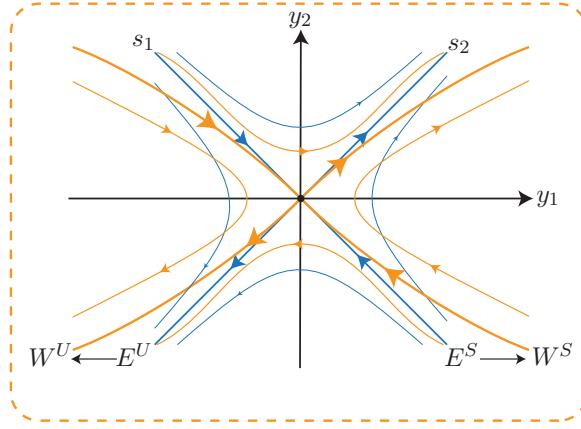


Figure 3.17: The phase portrait of the pendulum on a neighborhood around p with the stable and unstable manifolds of the nonlinear system as well as the stable and unstable spaces of the linearization.

Example 3.9 (Criticality of hyperbolicity in Hartman-Grobman). Let the dynamical system be

$$\dot{x} = ax^3, \quad x \in \mathbb{R}, \quad a \neq 0. \quad (3.73)$$

This system has a fixed point at $x = 0$, linearizing here gives

$$A = 3ax^2|_{x=0} = 0 \implies \dot{y} = 0y = 0. \quad (3.74)$$

We have a single root $\lambda_1 = 0$, hence $x = 0$ is a nonhyperbolic fixed point. Disregarding this fact, we may be inclined to conclude that $x = 0$ is a stable fixed point, since $y = 0$ is trivially a fixed point of the linearization $\dot{y} = 0$. This is not the case, as we can see by analyzing the full nonlinear dynamics for $a > 0$ as in Fig. 3.18, where we observe that $x = 0$ is an unstable fixed point.

Now that we have understood how to use the Hartman-Grobman theorem, we would like to be able to definitely conclude the stability type of a fixed point, once we have the linearization. To achieve this, we require a sufficient and necessary criterion for all of the eigenvalues of the linearized system to be left of the imaginary axis, i.e. $\text{Re}\lambda_i < 0$ for all i .

Theorem 3.15 (Routh-Hurwitz). *Consider the polynomial*

$$a_n\lambda^n + a_{n-1}\lambda^{n-1} + \dots + a_1\lambda + a_0 = 0. \quad (3.75)$$

Without loss of generality assume $a_0 > 0$, if $a_0 < 0$ then multiply by -1 and if $a_0 = 0$ then $\lambda = 0$ is a root and therefore we cannot have asymptotic stability. Next, define the following

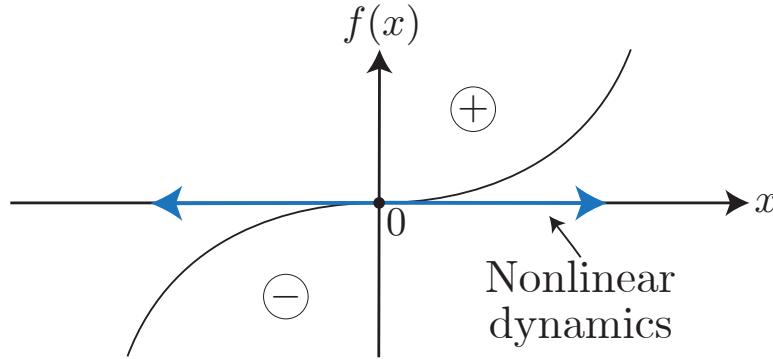


Figure 3.18: Nonlinear dynamics for the dynamical system $\dot{x} = ax^3$ with $a > 0$.

series of subdeterminants

$$D_0 = a_0, \quad D_1 = a_1, \quad D_2 = \begin{vmatrix} a_1 & a_0 \\ a_3 & a_2 \end{vmatrix}, \quad D_3 = \begin{vmatrix} a_1 & a_0 & 0 \\ a_3 & a_2 & a_1 \\ a_5 & a_4 & a_3 \end{vmatrix}, \dots, \quad D_n = \begin{vmatrix} a_1 & a_0 & 0 & \dots & 0 \\ a_3 & a_2 & a_1 & \dots & 0 \\ \vdots & & & & \vdots \\ 0 & \dots & a_n & a_{n-1} & a_{n-2} \\ 0 & \dots & & 0 & a_n \end{vmatrix} \quad (3.76)$$

Then we have that if and only if for all i $D_i > 0$ then $\operatorname{Re}\lambda_i < 0$ for all i .

A weaker necessary condition is that if for all i $\operatorname{Re}\lambda_i < 0$ then $a_i > 0$ for all i . Therefore if there exists an $a_k < 0$, we know immediately that the fixed point cannot be asymptotically stable as not all of the $\operatorname{Re}\lambda_i$ can be strictly negative.

Remark 3.16. For a given i we construct the matrix used for calculating D_i as follows: write the elements a_1, \dots, a_i along the diagonal, then in each row k write the a_j in descending index order such that a_k aligns with the placement inherited from us writing along the diagonal. The leftover spaces are filled with zeros.

Remark 3.17. Adolf Hurwitz discovered this criterion independently of Edward Routh in 1895 while holding a chair at the ETH.

Example 3.10 (Applying the Routh-Hurwitz criterion). Given the polynomial

$$a_3\lambda^3 + a_2\lambda^2 + a_1\lambda + a_0 = 0. \quad (3.77)$$

The Routh-Hurwitz criterion is

$$D_0 = a_0 > 0; \quad D_1 = a_1 > 0; \quad D_2 = a_1a_2 - a_0a_3 > 0; \quad D_3 = a_3D_2 > 0. \quad (3.78)$$

Therefore

$$\boxed{a_0 > 0, a_1 > 0, a_1 a_2 - a_0 a_3 > 0, a_3 > 0} \quad (3.79)$$

forms a sufficient and necessary condition for asymptotic stability ($a_i > 0$ follows from here) for $n = 3$.

Exercise 3.3. *The first three modes of a convecting fluid motion in a two-dimensional layer heated from below are described by the famous Lorenz equations*

$$\dot{x} = a(y - x), \quad (3.80)$$

$$\dot{y} = bx - y - xz, \quad (3.81)$$

$$\dot{z} = xy - cz. \quad (3.82)$$

Here $a > 0$ denotes the Prandtl number, $b > 0$ is the Rayleigh number, and $c > 0$ is the aspect ratio. Lorenz's original assumption is that $a > 1 + c$.

The above equations describe the evolution of the amplitudes of one velocity mode and two temperature modes. As a paradigm of chaotic dynamics, this system has inspired much of the development of the modern geometric theory of dynamical systems.

- (a) Complicated dynamics arise when all possible equilibria of the system become unstable, and hence solutions cannot settle down to any simple steady state. Show that this is the case when

$$b > \frac{a(3 + a + c)}{a - c - 1} \quad (3.83)$$

Note: Note: A negative sign for one of the Routh-Hurwitz determinants actually implies instability, not just the lack of asymptotic stability.

- (b) Solve the Lorenz equations numerically for $a = 10$, $b = 28$, and $c = 8/3$, choosing an initial condition close to $x = y = z = 0$. Plot the trajectory in three dimensions to show that it converges to a complicated surface, a chaotic attractor.

A more refined linearization result comes from Sternberg's Theorem [Chicone, 1999].

Theorem 3.18 (Sternberg). *Assume the following*

- (i) *The fixed point is asymptotically stable, i.e. $\operatorname{Re}\lambda_j < 0$ for all $j = 1, \dots, n$ (this also implies hyperbolicity);*
- (ii) *$f \in \mathcal{C}^r$ for $r \in \mathbb{N} \cup \{\infty\}$ and $r > \frac{\max_j |\operatorname{Re}\lambda_j|}{\min_j |\operatorname{Im}\lambda_j|}$;*

(iii) The system is nonresonant, i.e.

$$\sum_{j=1}^n m_j \lambda_j \neq \lambda_k, \quad \forall k \quad (3.84)$$

for any sequence $m_j \in \mathbb{N}$ with $\sum_{j=1}^n m_j \geq 2$.

Then the following holds: $\dot{x} = f(x)$ with $f(0) = 0$ is locally C^r conjugate to its linear part $\dot{y} = Ay$ for $A = Df(0)$.

Notebook 3.2. Calculation of the linearizing transformation as described by Sternberg's theorem.

Example 3.11 (Watt's centrifugal governor for steam engines). Now we put together everything built until now in the example of Watt's centrifugal governor for steam engines. Originally this system was used in mills in the 1700's, then it was adapted by Watt to the steam engine in 1788. This adaptation has been credited as a major factor in the industrial revolution, and is a first example of feedback control. The system is outlined in Fig. 3.19. The two masses (of mass

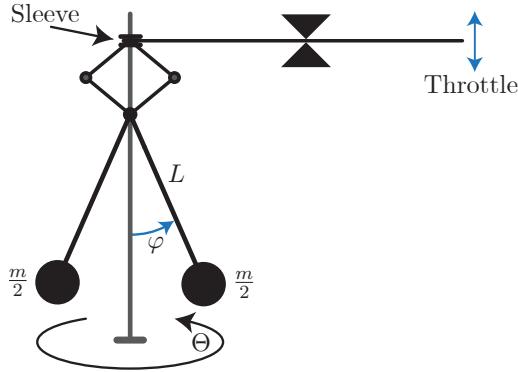


Figure 3.19: Schematic for Watt's centrifugal governor. The sleeve acts as a damper on the rotation about the spindle (drawn in grey).

$\frac{m}{2}$) rotate counter clockwise, and their position in radians is given by θ , with their rotational velocity $\dot{\theta}$. The masses are attached by a rod of length L and their deflection from the vertical position is measured by φ . Smaller $\dot{\theta}$ allowed for an increase in steam supply.

Following changes in the design, the systems suddenly became unstable. To address this Vyshnegradsky studied the root cause in 1877. We first derive the equation of motion. For the governor we use the equation of motion for a rotating hoop (with viscous damping coefficient b).

$$mL^2\ddot{\varphi} + bL^2\dot{\varphi} + mL^2 \left(\frac{g}{L} - \dot{\theta}^2 \cos(\varphi) \right) \sin(\varphi) = 0; \quad b > 0. \quad (3.85)$$

Next, let ω denote the angular velocity of the steam engine, i.e. with the gear ratio n we have $\dot{\theta} = n\omega$. Then we find

$$m\ddot{\varphi} = -b\dot{\varphi} - m \left(\frac{g}{L} - n^2 w^2 \cos(\varphi) \right) \sin(\varphi). \quad (3.86)$$

Now we derive the equation of motion for the steam engine. Denote the moment of inertia for the engine by J , the driving torque from the steam as P_1 and the constant load P , we obtain

$$J\dot{\omega} = P_1 - P. \quad (3.87)$$

In this case we have $P_1 = P^* + k (\cos(\varphi) - \cos(\varphi^*))$ for the desired operation angle φ^* , the gain k , and P^* the value of P at φ^* . Putting this together yields

$$J\dot{\omega} = k \cos(\varphi) - P_0; \quad P_0 = P - P^* + k \cos(\varphi^*). \quad (3.88)$$

Let $\dot{\varphi} = \Psi$ to transform into a three-dimensional set of equations (ODE)

$$\begin{cases} \dot{\varphi} = \Psi \\ \dot{\Psi} = -\frac{b}{m}\Psi - \left(\frac{g}{L} - n^2 \omega^2 \cos(\varphi) \right) \sin(\varphi); \\ \dot{\omega} = \frac{k}{J} \cos(\varphi) - \frac{P_0}{J}. \end{cases} \quad x = \begin{pmatrix} \varphi \\ \Psi \\ \omega \end{pmatrix}. \quad (3.89)$$

Then our operation point is the fixed point x_0 of this system

$$f(x_0) = 0 \implies \Psi_0 = 0; \quad \omega_0^2 = \frac{g}{Ln^2 \cos(\varphi_0)}; \quad \cos(\varphi_0) = \frac{P_0}{k}. \quad (3.90)$$

If $\sin(\varphi_0) = 0$, we have an unphysical state and ignore this case. For simplification set $L = 1$, this could be formally achieved by nondimensionalizing the length L . Now we linearize at the fixed point x_0

$$\dot{y} = Ay; \quad A = Df(x_0) = \begin{pmatrix} 0 & 1 & 0 \\ n^2 \omega_0^2 \cos(2\varphi_0) - g \cos(\varphi_0) & -\frac{b}{m} & n^2 \omega_0 \sin(2\varphi_0) \\ -\frac{k}{J} \sin(\varphi_0) & 0 & 0 \end{pmatrix}. \quad (3.91)$$

With this we obtain the characteristic equation $\det(A - \lambda I) = 0$

$$\underbrace{1}_{a_3} \lambda^3 + \underbrace{\frac{b}{m}}_{a_2} \lambda^2 + \underbrace{g \frac{\sin^2(\varphi_0)}{\cos(\varphi_0)}}_{a_1} \lambda + \underbrace{g \frac{2k \sin^2(\varphi_0)}{J \omega_0}}_{a_0} = 0. \quad (3.92)$$

Now check the Routh-Hurwitz criterion for asymptotic stability:

(i) The necessary condition for $\operatorname{Re}\lambda_k < 0$ for all k : $a_j > 0$ for all j is fulfilled.

(ii) Next we check the subdeterminants

$$D_0 = a_0 = g \frac{2k \sin^2(\varphi_0)}{J\omega_0} > 0; \quad (3.93a)$$

$$D_1 = a_1 = g \frac{\sin^2(\varphi_0)}{\cos(\varphi_0)} > 0; \quad (3.93b)$$

$$D_2 = \begin{vmatrix} a_1 & a_0 \\ a_3 & a_2 \end{vmatrix} = a_1 a_2 - a_0 a_3 = g \frac{b}{m} \frac{\sin^2(\varphi_0)}{\cos(\varphi_0)} - g \frac{2k \sin^2(\varphi_0)}{J\omega_0} > 0; \quad (3.93c)$$

$$D_3 > 0 \iff D_2 > 0 \text{ and } a_3 = 1 > 0. \quad (3.93d)$$

The only actual condition for x_0 to be asymptotically stable is $D_2 > 0$

$$\frac{bJ}{m} > \frac{2P_0}{\omega_0}. \quad (3.94)$$

From the equation for the fixed points we know

$$P_0 \omega_0^2 = \frac{gk}{Ln^2} = \text{const.} \quad (3.95a)$$

$$\omega_0^2 + 2P_0 \omega_0 \frac{d\omega_0}{dP_0} = 0. \quad (3.95b)$$

From the first equation, we realize that we must write $\omega_0 = \omega_0(P_0)$. To obtain the second equation, we derive the first equation (3.95a) with respect to P_0 . Therefore we find

$$\frac{d\omega_0}{dP_0} = -\frac{\omega_0}{2P_0}. \quad (3.96)$$

Next, define the *non-uniformity of performance*

$$\nu = \left| \frac{d\omega_0}{dP_0} \right| = \frac{\omega_0}{2P_0}. \quad (3.97)$$

Then our criterion for asymptotic stability becomes

$$\frac{bJ}{m} \nu > 1. \quad (3.98)$$

To conclude, the following have harmful effects of stability

- Bigger engines which increase m

- Better machining of surfaces decreasing b
- Increased operating speed decreasing J
- Versatility in operation decreasing ν .

These have the effect of pushing the eigenvalues to the right in the complex plane (towards the imaginary axis) as is illustrated in Fig 3.20.

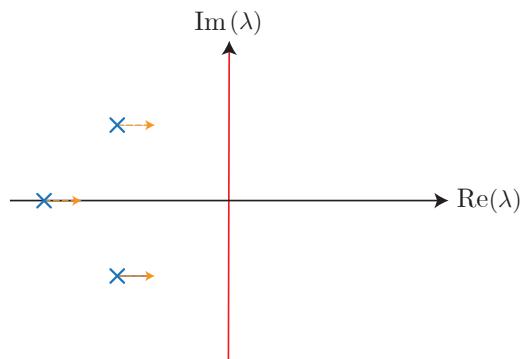


Figure 3.20: Change in the eigenvalue configuration as small modification are made to the Watt engine governor.

Notebook 3.3. *Analysis of the equations of motion for Watt's governor.*

3.6 Data-driven linear modeling of dynamical systems

Given are $m + 1$ vectors of n -dimensional observables $x_0, \dots, x_m \in \mathbb{R}^n$ observed at Δt intervals. An example of such data would be many different trajectories or snapshots of a single trajectory. Typically $m \gg 1$, i.e. there are a high number of observations.

Our objective is to find the best fitting linear dynamical system

$$\dot{x} = Ax; \quad x \in \mathbb{R}^n; \quad A \in \mathbb{R}^{n \times n}. \quad (3.99)$$

Since the data is given as discrete observations, this is equivalent to seeking the best $B = e^{A\Delta t}$ such that $x_{k+1} = Bx_k$ for $k = 1, \dots, m - 1$. To solve this we synthesize the data into matrices

$$X = [x_0 \dots x_{m-1}] \in \mathbb{R}^{n \times m}; \quad Y = [x_1 \dots x_m] \in \mathbb{R}^{n \times m}. \quad (3.100)$$

Notice that the matrix Y is shifted forward compared to the matrix X by Δt . Next, we solve for a matrix B that minimizes the distance between Y and BX . That is, we find

$$B_* = \underset{B \in \mathbb{R}^{n \times n}}{\operatorname{argmin}} g(B); \quad g(B) = \|Y - BX\|^2, \quad (3.101)$$

where the matrix norm is the Frobenius norm. We denote the components of B by b_{ij} , the components of Y by y_{ij} , and the components of X by x_{ij} . Note that

$$\frac{\partial g}{\partial b_{ij}} = \frac{\partial}{\partial b_{ij}} \sum_{l,m,p} (y_{lm} - b_{lp}x_{pm})^2 = -2 \sum_{l,m,p} (y_{lm} - b_{lp}x_{pm}) \delta_{il}\delta_{jp}x_{pm} \quad (3.102a)$$

$$= -2 \sum_{m,p} (y_{im} - b_{ip}x_{pm}) x_{jm} = 2(YX^T - BXX^T). \quad (3.102b)$$

In the third equality we have used the properties of the Kronecker-delta δ_{ij} . Therefore we must have that $D_B g = -2(YX^T - BXX^T) = 0$, so we must solve the linear system of equations

$$BXX^T = YX^T \in \mathbb{R}^{n \times n}. \quad (3.103)$$

In the case $n > m$, the matrix X^T has more columns than rows, thus X^T is singular. Hence XX^T has a nonempty kernel and is therefore not invertible. For $n < m$ this is not necessarily the case. For a solution, we can use the pseudo inverse $(XX^T)^\dagger$, the superscript here is called *dagger*.

Recall that if $A \in \mathbb{R}^{n \times n}$ is invertible then $A^{-1} = A^\dagger$. If A is not invertible, then calculate the singular value decomposition of $A = U\Sigma V^T$ for $U, V \in \text{SO}(n)$ and Σ a diagonal matrix.

The entries of Σ are the strictly positive singular values. The columns of U and V are the left and right singular vectors. We then define the pseudo inverse as

$$A^\dagger = V\Sigma^\dagger U^T; \quad \Sigma_{ij}^\dagger = \begin{cases} \frac{1}{\Sigma_{ij}} & \Sigma_{ij} \neq 0 \\ 0 & \Sigma_{ij} = 0. \end{cases} \quad (3.104)$$

This means that for $Ax = b$, the best solution in a least squares sense is given by $x = A^\dagger b$. In our case, the solution to 3.101 is

$$B = (YX^T)(XX^T)^\dagger. \quad (3.105)$$

This procedure is called *dynamic mode decomposition* (DMD) [Schmid, 2010, Kutz et al., 2015]. The eigenvectors of B are called *DMD modes* and the normalized exponents of those corresponding eigenvalues are called *DMD eigenvalues*.

The procedure has been extended in *extended DMD*, where DMD was performed on functions of the observables (typically polynomials) [Williams et al., 2015]. This effectively tries to construct the linearization from data. Note here that none of these procedures can describe intrinsically nonlinear phenomena (i.e. those with coexisting isolated compact invariant sets).

Remark 3.19. The numerical cost of (pseudo)inverting the $n \times n$ matrix, XX^T , can be reduced when $n \gg m$.

- (i) We can calculate the singular value decomposition of the data matrix X instead.
- (ii) By keeping only the leading r singular values in this decomposition, we obtain a reduced (rank- r) approximation of the matrix $B \in \mathbb{R}^{n \times n}$ as follows. Denote the singular value decomposition by $X = U\Sigma V^T$. This also means that $(XX^T) = U\Sigma^2 U^T$, i. e. we get the same U as before, but from an SVD of an $n \times m$ matrix, which is much less demanding numerically.

With the leading r singular vectors and singular values σ , X can be written as

$$X \approx U_r \Sigma_r V_r^T = \sum_{i=1}^r \sigma_i u_i v_i^T, \quad U = [u_1, \dots, u_n] \quad V = [v_1, \dots, v_m]. \quad (3.106)$$

- (iii) The matrix B can be approximately written as

$$B = YX^T(XX^T)^\dagger \quad (3.107a)$$

$$\approx YV_r \Sigma_r U_r^T U_r (\Sigma_r^2)^\dagger U_r^T \quad (3.107b)$$

$$= YV_r \Sigma_r^\dagger U_r^T = Y(X_r)^\dagger. \quad (3.107c)$$

We denote the truncated approximation of B as $B_r = Y(X_r)^\dagger$.

Notebook 3.4. *Finding approximate linear models with DMD.*

Exercise 3.4. Consider the damped pendulum equation

$$\ddot{x} + c\dot{x} + \sin(x) = 0. \quad (3.108)$$

Using the energy of the pendulum as a Lyapunov function, can we conclude the nonlinear asymptotic stability of the $x = 0$ equilibrium? Give detailed reasoning why.

3.7 Lyapunov's direct (second) method for stability

Stability analysis via linearization is not perfect. For nonhyperbolic fixed points, the results obtained are inconclusive, and this case turns out to be somewhat common for conservative systems. Furthermore, linearization does not give any insight into the size of the domain of stability. Hence, a method for determining stability type without relying on linearization would be desirable. This is given to us by Lyapunov's direct method.

Theorem 3.20. Consider

$$\dot{x} = f(x), \quad f \in \mathcal{C}^1, \quad x \in \mathbb{R}^n; \quad f(x_0) = 0. \quad (3.109)$$

Assume that there exists a function $V : U \rightarrow \mathbb{R}$ with $V \in \mathcal{C}^1(U)$ for an open set $U \subset \mathbb{R}^n$ and $x_0 \in U$ which fulfills the following

(i) V is positive definite in a neighborhood of x_0 , i.e.

$$V(x_0) = 0; \quad V(x) > 0, \quad x \in U - \{x_0\}. \quad (3.110)$$

(ii) \dot{V} is negative semidefinite in the same neighborhood, i.e.

$$\dot{V} = \frac{d}{dt}V(x(t)) = \langle DV(x(t)), \dot{x}(t) \rangle = \langle DV(x(t)), f(x(t)) \rangle \leq 0, \quad x \in U. \quad (3.111)$$

Then $x = x_0$ is Lyapunov stable (cf. Def. 3.1). V is called a Lyapunov function. The hypotheses are illustrated in Fig. 3.21.

Remark 3.21. To denote the boundary of a set A we write ∂A .

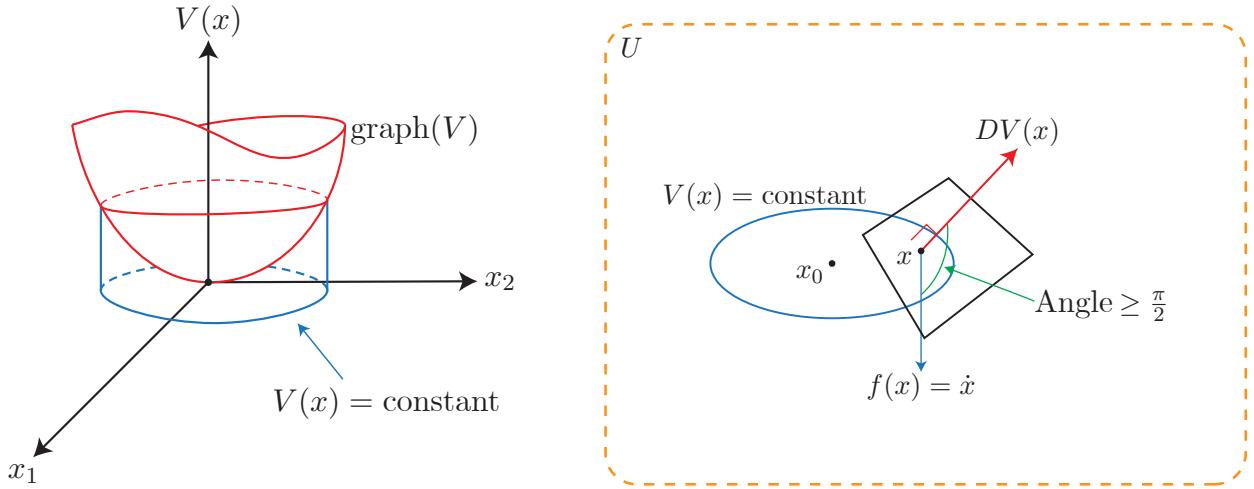


Figure 3.21: Geometric interpretation of hypotheses of Lyapunov's direct method. Left depicts hypothesis (i) and right hypothesis (ii). The angle between $DV(x)$ and $f(x)$ is at least $\frac{\pi}{2}$. In each image the blue $V(x) = \text{const.}$ is a level surface diffeomorphic to S^{n-1} .

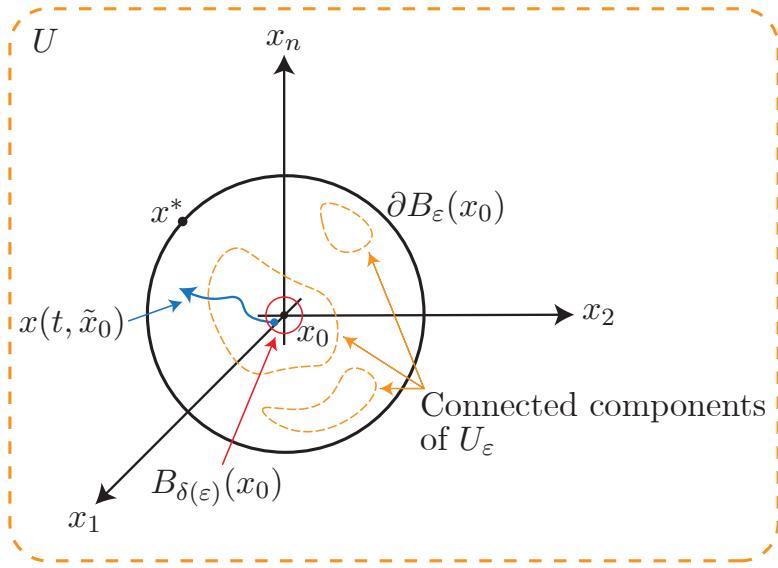


Figure 3.22: The constellation of U , U_ε , $\partial B_\varepsilon(x_0)$, and $B_{\delta(\varepsilon)}(x_0)$ from the proof of Lyapunov's direct method. The red arrow points at $B_{\delta(\varepsilon)}(x_0)$ and the yellow arrow at a connected component of U_ε .

Proof. First choose $\varepsilon > 0$ and define $\alpha(\varepsilon) = \min_{x \in \partial B_\varepsilon(x_0)} V(x) > 0$. Note that $\alpha(\varepsilon)$ is well defined as $V \in C^0(U)$ and $\partial B_\varepsilon(x)$ is compact and spherical for small enough ε . There exists an $x^* \in \partial B_\varepsilon(x)$ with $V(x^*) = \alpha(\varepsilon) \leq V(x)$ for all $x \in \partial B_\varepsilon(x_0)$. Next, define $U_\varepsilon = \{x \in B_\varepsilon(x_0) :$

$V(x) < \alpha(\varepsilon)$. Notice that $x_0 \in U_\varepsilon$ because $V(x_0) = 0$ and $V(x) \geq 0$ on U . Further U_ε is open due to the continuity of V . We have that $U_\varepsilon \cap \partial B_\varepsilon(x_0)$ is empty by definition, noting that for all $x \in \partial B_\varepsilon(x_0)$ we have $V(x) \geq \alpha(\varepsilon)$. Therefore there exists a ball $B_{\delta(\varepsilon)} \subset U_\varepsilon$ which contains x_0 . This can be seen in Fig. 3.22.

Now observe that for every $\tilde{x}_0 \in B_{\delta(\varepsilon)}(x_0)$ we have that along trajectories $V(x(t; \tilde{x}_0)) \leq V(\tilde{x}_0) < \alpha(\varepsilon)$. The first inequality comes from hypothesis (ii) and the second inequality from the definition of U_ε . This implies that for $x(t; \tilde{x}_0) \in U_\varepsilon$ we have that $x(t; \tilde{x}_0)$ is not in $\partial B_\varepsilon(x_0)$ for any $t > 0$. The trajectory $x(t; \tilde{x}_0)$ is continuous, in order for it to leave the ball $B_\varepsilon(x_0)$ it must intersect the boundary $\partial B_\varepsilon(x_0)$. At this point $V(x)$ will attain a value of at least $\alpha(\varepsilon)$ by definition, however this is in contradiction to the fact that $V(x(t; \tilde{x}_0))$ is strictly smaller than $\alpha(\varepsilon)$. Therefore must stay in the ball $B_\varepsilon(x_0)$ for all times. \square

Now we present some extensions to this theorem for various different stability types. These are necessary, as the previous theorem only provides a sufficient condition for stability (it is not “if and only if”).

Theorem 3.22 (Theorem 2). *Consider the same dynamical system. Assume*

- (i) $V(x)$ is positive definite,
- (ii) $\dot{V}(x)$ is negative definite, i.e.

$$\dot{V}(x) < 0, \quad x \in U - \{x_0\}. \quad (3.112)$$

Then $x = x_0$ is asymptotically stable. These hypotheses are illustrated in Fig. 3.23.

Theorem 3.23 (Theorem 3). *Consider the same dynamical system. Assume*

- (i) $V(x)$ is positive definite,
- (ii) $\dot{V}(x)$ is positive definite, i.e.

$$\dot{V}(x) > 0, \quad x \in U - \{x_0\}. \quad (3.113)$$

Then $x = x_0$ is unstable. The hypotheses are illustrated in Fig. 3.24.

Theorem 3.24 (Theorem 4). *Consider the same dynamical system. Assume*

- (i) $V(x)$ is indefinite, i.e. arbitrarily close to x_0 there exists $a, b \in U$ such that $V(x_1) \cdot V(x_2) < 0$ (they have opposite signs and are not equal to 0) and $V(x_0) = 0$. Furthermore, this is a critical point, i.e. $DV(x_0) = 0$.

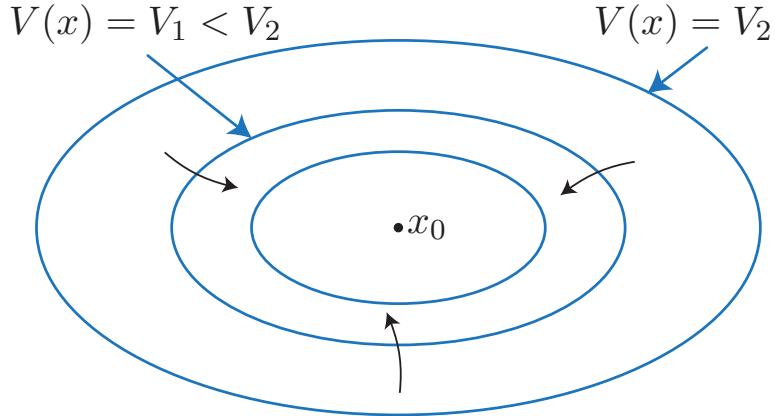


Figure 3.23: The hypotheses of Theorem 2 illustrated, the key difference being that the arrows (denoting the flow of the dynamical system) cross the level surfaces of V (blue rings) toward x_0 .

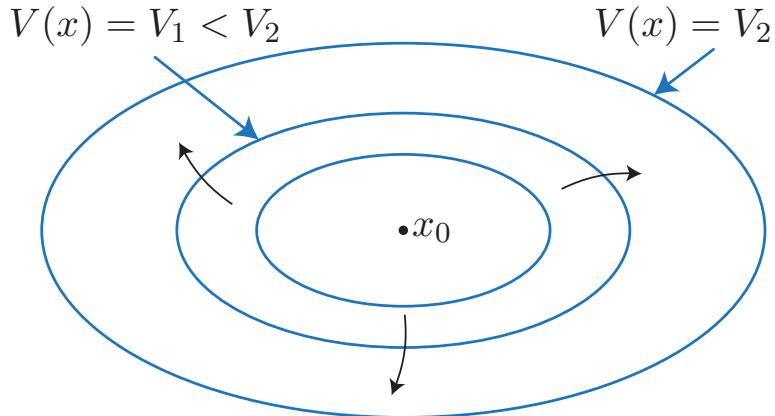


Figure 3.24: The hypotheses of Theorem 3 illustrated, the key difference being that the arrows (denoting the flow of the dynamical system) cross the level surfaces of V (blue rings) away from x_0 .

(ii) $\dot{V}(x)$ is definite near x_0 (either positive or negative).

Then $x = x_0$ is unstable. The geometry of the hypotheses are illustrated in Fig. 3.25.

Remark 3.25. In each of these theorems, the definiteness of \dot{V} can be replaced by semidefiniteness, if we add that the set $\{x \in U : \dot{V}(x) = 0\}$ does not contain full trajectories of the system. This is called Krasovskiy's condition.

Now we would like to put these theorems into practice with a few examples.

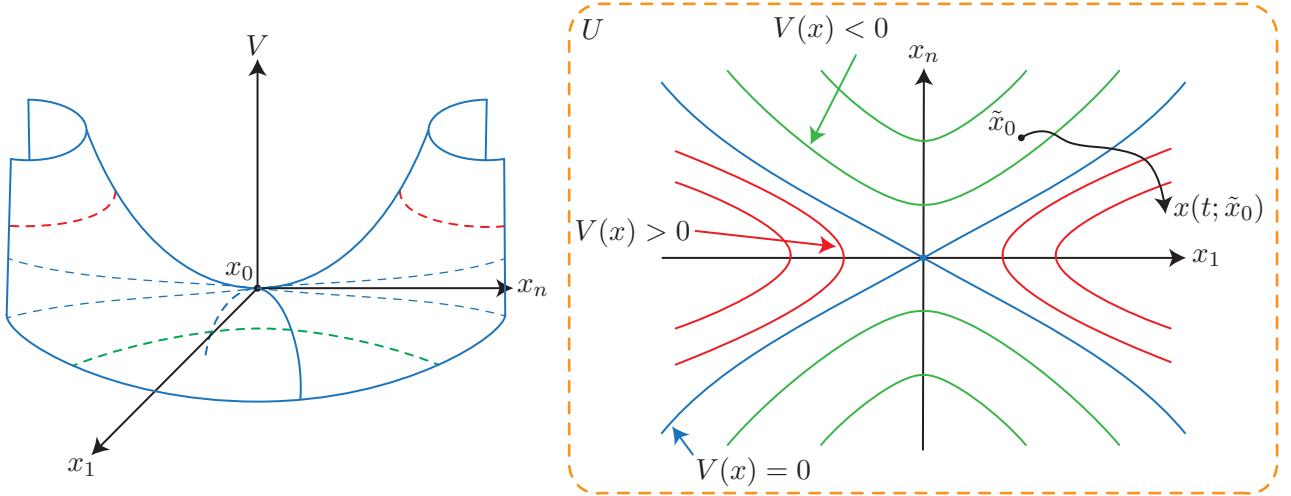


Figure 3.25: The geometry of hypotheses (i) (left) and (ii) (right) for \dot{V} positive definite of Theorem 4 illustrated. On the right level surfaces are designated by lines, blue corresponds to $V = 0$, red $V > 0$, and green $V < 0$, which can be seen as the dotted lines of the same colors on the left.

Example 3.12 (Stability analysis of the pendulum with Lyapunov's direct method). Recall that using linearization, we were only able to conclude the stability type of one of the fixed points for the dynamical system of the pendulum

$$ml^2\ddot{\varphi} + mgl \sin(\varphi) = 0. \quad (3.114)$$

Now we will use the energy as a Lyapunov function, which is often very useful. The energy is given by

$$E(x) = E(\varphi, \dot{\varphi}) = \frac{1}{2}ml^2\dot{\varphi}^2 + mgl(1 - \cos(\varphi)) = \frac{1}{2}ml^2\dot{x}_2^2 + mgl(1 - \cos(x_1)). \quad (3.115)$$

Transforming the dynamical system to be an system of first order ODEs we obtain

$$x = \begin{pmatrix} x_1 \\ x_2 \end{pmatrix} = \begin{pmatrix} \varphi \\ \dot{\varphi} \end{pmatrix}; \quad \dot{x} = \begin{pmatrix} \dot{x}_1 \\ \dot{x}_2 \end{pmatrix} = \begin{pmatrix} x_2 \\ -\frac{g}{l} \sin(x_1) \end{pmatrix} = f(x). \quad (3.116)$$

At the fixed point $x = (0, 0)$ we have

$$E(0, 0) = 0; \quad DE(0, 0) = 0 \in \mathbb{R}^{2 \times 2}; \quad D^2E(0, 0) = \begin{pmatrix} mgl & 0 \\ 0 & ml^2 \end{pmatrix}. \quad (3.117)$$

We have that the Hessian of E is positive definite. Therefore E is positive definite at $(0, 0)$. Further we have

$$\dot{E}(x) = \langle DE(x), f(x) \rangle = (mgl \sin(x_1) \quad ml^2 x_2) \begin{pmatrix} x_2 \\ -\frac{g}{l} \sin(x_1) \end{pmatrix} = 0. \quad (3.118)$$

Thus \dot{E} is negative semidefinite. Now the hypotheses of Theorem 3.20 are fulfilled, hence $x = (0, 0)$ is Lyapunov stable. Importantly, this is a nonlinear result and we did not need to refer to the linearization of the system!

At the fixed point $x = (\pi, 0)$. We check again using the energy. First we realize that $E(\pi, 0) = 2mgl$, so we subtract the constant and redefine to obtain $\tilde{E} = E - 2mgl$. Now $\tilde{E}(\pi, 0) = 0$, although this is not essential. Next we calculate

$$DE(\pi, 0) = 0 \in \mathbb{R}^{n \times n}; \quad D^2E(\pi, 0) = \begin{pmatrix} -mgl & 0 \\ 0 & ml^2 \end{pmatrix}. \quad (3.119)$$

Hence, E is indefinite at $(\pi, 0)$. From the previous calculation we already know that $\dot{E}(\pi, 0) = 0$, i.e. \dot{E} is semidefinite. We cannot apply Theorem 4, but linearization already concluded that $(\pi, 0)$ was unstable.

If we add damping the dynamics becomes

$$ml^2 \ddot{\varphi} + c\dot{\varphi}mgl \sin(\varphi) = 0. \quad (3.120)$$

Again we transform the system by introducing the coordinates $x_1 = \varphi$ and $x_2 = \dot{\varphi}$. Thus we have a first order ODE

$$\begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -\frac{c}{ml^2}x_2 - \frac{g}{l} \sin(x_1). \end{cases} \quad (3.121)$$

We still have the fixed points $(0, 0)$ and $(0, \pi)$ and the energy $E = \frac{1}{2}ml^2\dot{\varphi}^2 + mgl(1 - \cos(\varphi))$. Again we find

$$DE = \begin{pmatrix} mgl \sin(x_1) \\ ml^2 x_2 \end{pmatrix} \implies DE(0) = 0; \quad D^2E = \begin{pmatrix} mg\ell \cos(x_1) & 0 \\ 0 & ml^2 \end{pmatrix}. \quad (3.122)$$

Thus E is positive definite at $(0, 0)$ and indefinite at $(0, \pi)$. We have

$$\dot{E} = DE \cdot f = mgl \sin(x_1)x_2 + ml^2 x_2 \left(-\frac{c}{ml^2}x_2^2 - \frac{g}{l} \sin(x_1) \right) = -cx_2^2. \quad (3.123)$$

Note here that $\dot{E} = -cx_2^2$ is only negative semi-definite. Therefore the asymptotic stability of $x = (0, 0)$ and the instability of $x = (0, \pi)$ do not follow from theorems 1 or 4. However, the set $S = \{x : \dot{E}(x) = 0\} - \{(0, 0), (0, \pi)\}$ does not contain full trajectories. Indeed

$$\frac{d}{dt}x_2(t) \Big|_{x=0} = -\frac{g}{\ell} \sin(x_1) \neq 0, \quad x \in S. \quad (3.124)$$

Hence, the asymptotic stability of $(0, 0)$ and instability of $(0, \pi)$ follows by Krasovskiy's condition.

Example 3.13 (Stability analysis of the friction pendulum). We have a shaft which is constantly rotating with angular speed Ω , around this shaft is a sleeve which rubs against the shaft creating friction. To this sleeve is a mass m attached at distance l , the deflection of this mass from its standard position (directly below the shaft) is measured by φ . The gravity constant is given by g . This setup is depicted in Fig. 3.26. The torque driving the pendulum is given by

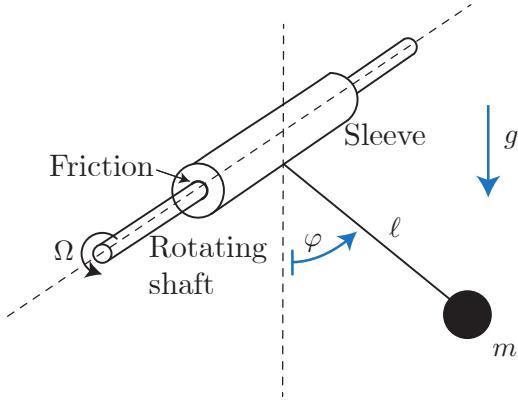


Figure 3.26: The setup of the friction pendulum.

$$T(\dot{\varphi}) = T_0 \text{sign}(\Omega - \dot{\varphi}). \quad (3.125)$$

From this we get the equation of motion

$$ml^2\ddot{\varphi} + mgl \sin(\varphi) = T(\dot{\varphi}) = T_0 \text{sign}(\Omega - \dot{\varphi}). \quad (3.126)$$

By assuming $\Omega \gg 1$, i.e. very fast rotation of the shaft, we find that $T(\dot{\varphi}) = T_0$. Next we transform the coordinates in order to get an ODE

$$\begin{pmatrix} x_1 \\ x_2 \end{pmatrix} = \begin{pmatrix} \varphi \\ \dot{\varphi} \end{pmatrix}; \quad \begin{pmatrix} \dot{x}_1 \\ \dot{x}_2 \end{pmatrix} = \begin{pmatrix} x_2 \\ -\frac{g}{l} \sin(x_1) + \frac{T_0}{ml^2} \end{pmatrix}. \quad (3.127)$$

Thus the fixed point x_0 is at (\bar{x}_1, \bar{x}_2) given by

$$\sin(\bar{x}_1) = \frac{T_0}{mgl}; \quad \bar{x}_2 = 0. \quad (3.128)$$

Since the system is forced, energy is not conserved and we cannot use it as a Lyapunov function. However, there may still exist a conserved quantity. Consider using

$$V(x) = [\text{total energy at time } t] - [\text{work put in between time } t \text{ and } t_0] = [\text{initial energy}] = \text{const}. \quad (3.129)$$

This is a conserved quantity in all mechanics problems, but generally cannot be calculated without already knowing the trajectories. Here we can, as

$$V(x(t)) = \frac{1}{2}mlx_2^2 + mgl(1 - \cos(x_1)) - T_0(\underbrace{x_1}_{\varphi(t)} - \underbrace{x_1(0)}_{\varphi(0)}). \quad (3.130)$$

We can drop the constant term $T_0x_1(0)$. Now verify that V is indeed constant

$$\dot{V}(x) = \frac{\partial V}{\partial x_1}\dot{x}_1 + \frac{\partial V}{\partial x_2}\dot{x}_2 = (mgl \sin(x_1) - T_0)x_2 + ml^2x_2\left(-\frac{g}{l} \sin(x_1) + \frac{T_0}{ml^2}\right) = 0. \quad (3.131)$$

Hence $\dot{V}(x)$ is semidefinite, and may be used to conclude stability or instability. We must check the Lyapunov conditions for $V(x)$

$$DV(x_0) = (mgl \sin(\bar{x}_1) - T_0 \quad ml^2\bar{x}_2) = (0 \quad 0) \quad (3.132a)$$

$$D^2V(x_0) = \begin{pmatrix} mgl \cos(\bar{x}_1) & 0 \\ 0 & ml^2 \end{pmatrix}. \quad (3.132b)$$

The Hessian is positive definite as long as $\bar{x}_1 \in (-\frac{\pi}{2}, \frac{\pi}{2})$, and all fixed points in this region are stable by Theorem 3.20. The Hessian is indefinite if $\bar{x}_1 \notin [-\frac{\pi}{2}, \frac{\pi}{2}]$, but in this case Theorem 4 is not applicable as \dot{V} is not definite. More detailed discussion and further results on Lyapunov's direct method can be found in [Rouche et al., 1977].

Exercise 3.5. Recall from Exercise 3.2 that a ball of mass m sliding on a hoop rotating with angular velocity Ω satisfies the differential equation

$$mR^2\ddot{\alpha} + mR^2(g/R - \Omega^2 \cos(\alpha)) \sin(\alpha) = 0 \quad (3.133)$$

if there is no friction between the hoop and the mass, i.e. we assume $b = 0$. Assume that the parameter values are such that the lower equilibrium position of the ball is stable in linear approximation.

(a) Show that in this case, the equilibrium is also nonlinearly stable.

Hint: Note that system (3.133) is not conservative: an external torque is required to maintain the constant rotation of the hoop. As a result, the total mechanical energy of the system is not expected to work as a Lyapunov function. To find another candidate for a Lyapunov function, find a quantity that is conserved along trajectories. Such a quantity can be found by multiplying equation (3.133) by $\dot{\alpha}$ and integrating in time.

(b) Prove that the equilibrium cannot be asymptotically stable for the nonlinear system.

Hint: use the Lyapunov function you have found in (a).

Exercise 3.6. Consider the damped pendulum equation

$$\ddot{x} + c\dot{x} + \sin(x) = 0. \quad (3.134)$$

(a) A theorem due to Krasovski states the following: Assume that $x = 0$ is a fixed point for the n -dimensional dynamical system $\dot{x} = f(x)$. Assume that there exists a smooth scalar function $V(x)$ such that

- (i) $V(x)$ is positive definite on an open neighborhood U of $x = 0$;
- (ii) \dot{V} is negative semi-definite on the same neighborhood;
- (iii) the only trajectory lying completely in the set $S = \{x \in U : \dot{V} = 0\}$ is the fixed point $x = 0$. Then $x = 0$ is asymptotically stable.

Use Krasovski's theorem to conclude the asymptotic stability of the origin for (3.134).

Exercise 3.7. Consider an n -degree-of-freedom holonomic mechanical system (i.e. one that has only position-dependent constraints) with generalized coordinates $q \in \mathbb{R}^n$ and generalized velocities $\dot{q} \in \mathbb{R}^n$. The total energy of the system is of the form

$$E(q, \dot{q}) = \frac{1}{2}\dot{q}^T M(q)\dot{q} + V(q), \quad (3.135)$$

where $M \in \mathbb{R}^{n \times n}$ is the mass matrix (symmetric and positive definite), and $V(q)$ is the potential energy. In the absence of external forces, the associated Lagrangian equations of motion are

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{q}} - \frac{\partial L}{\partial q} = 0, \quad (3.136)$$

where $L(q, \dot{q}) = \frac{1}{2}\dot{q}^T M(q)\dot{q} - V(q)$ is the Lagrangian of the mechanical system.

Show that if $V(q)$ admits a strict local minimum at a point q_0 , then q_0 is a (nonlinearly) stable equilibrium for the mechanical system. (This result is also known as Dirichlet's Theorem in classical mechanics).

Chapter 4

Bifurcations of fixed points

4.1 Local nonlinear dynamics near fixed points

We are interested in the local nonlinear dynamics around fixed points. Consider

$$\dot{x} = f(x); \quad f \in \mathcal{C}^r, r \geq 1; \quad f(p) = 0, \quad (4.1)$$

i.e. p is a fixed point of the dynamical system. The linearized system at p is

$$\dot{y} = Df(p)y, \quad y \in \mathbb{R}^n, \quad Df(p) \in \mathbb{R}^{n \times n}. \quad (4.2)$$

The linearization has the eigenvalues $\lambda_1, \dots, \lambda_n \in \mathbb{C}$ with multiplicities counted. Corresponding to these eigenvalues are the eigenvectors $e_1, \dots, e_n \in \mathbb{C}^n$, including generalized eigenvectors for when the algebraic multiplicity is greater than the geometric multiplicity. The eigenvector e_j is real when $\lambda_j \in \mathbb{R}$.

Definition 4.1. The following subspaces are invariant for the linearized dynamical system:

(i) The *stable subspace*

$$E^S = \text{span}_j \{ \text{Re}(e_j), \text{Im}(e_j) : \text{Re}(\lambda_j) < 0 \}, \quad (4.3)$$

(ii) The *unstable subspace*

$$E^U = \text{span}_j \{ \text{Re}(e_j), \text{Im}(e_j) : \text{Re}(\lambda_j) > 0 \}, \quad (4.4)$$

(iii) The *center subspace*

$$E^C = \text{span}_j \{ \text{Re}(e_j), \text{Im}(e_j) : \text{Re}(\lambda_j) = 0 \}. \quad (4.5)$$

Remark 4.1. Note here that the following facts hold

- (i) $E^C = \emptyset$ if and only if p is hyperbolic,
- (ii) $E^{U,S}$ and E^C are invariant subspaces of (4.2) by construction,
- (iii) Solutions of (4.2) in E^S (resp. E^U) decay to $y = 0$ as $t \rightarrow \infty$ (resp. $t \rightarrow -\infty$).

We now discuss what happens to these subspaces in the nonlinear system (4.1).

Theorem 4.2 (Center Manifold Theorem). *The following hold:*

- (i) *There exists a unique stable manifold $W^S(p)$ for (4.1), such that*

- $W^S(p)$ is a \mathcal{C}^r manifold (surface), tangent to E^S at p with $\dim W^S(p) = \dim E^S$,
- $W^S(p)$ is invariant for (4.1) and for $x \in W^S(p)$ we have

$$\|F^t(x) - p\| \leq K_S \exp \left[t \left(\max_{Re(\lambda_j) < 0} (Re(\lambda_j)) + \varepsilon_S \right) \right] \quad (4.6)$$

for $t \geq 0$, $0 < \varepsilon_S \ll 1$, and $\|x - p\|$ small enough.

- (ii) *There exists a unique unstable manifold $W^U(p)$ for (4.1), such that*

- $W^U(p)$ is a \mathcal{C}^r manifold (surface), tangent to E^U at p with $\dim W^U(p) = \dim E^U$,
- $W^U(p)$ is invariant for (4.1) and for $x \in W^U(p)$ we have

$$\|F^t(x) - p\| \leq K_U \exp \left[t \left(\min_{Re(\lambda_j) > 0} (Re(\lambda_j)) - \varepsilon_U \right) \right] \quad (4.7)$$

for $t \leq 0$, $0 < \varepsilon_U \ll 1$, and $\|x - p\|$ small enough.

- (iii) *There exists a (not necessarily unique) center manifold $W^C(p)$ for (4.1), such that*

- $W^C(p)$ is a \mathcal{C}^{r-1} manifold (surface), tangent to E^C at p with $\dim W^C(p) = \dim E^C$,

For a proof, we refer to Section 2 of [Marsden and McCracken, 1976]. The geometry of these manifolds and their corresponding subspaces are sketched in Fig. 4.1.

Notebook 4.1. *Approximation of stable and unstable manifolds by time-integration.*

Notice that Theorem 4.2 does not state anything about the asymptotic dynamics on the center manifold, in contrast to that of the unstable and stable manifolds. This means that the overall dynamics depends crucially on the center manifold. In the case when $W^u = \emptyset$, even the stability type of the fixed point is determined by $W^C(p)$. This is why it will be the subject of further investigations.

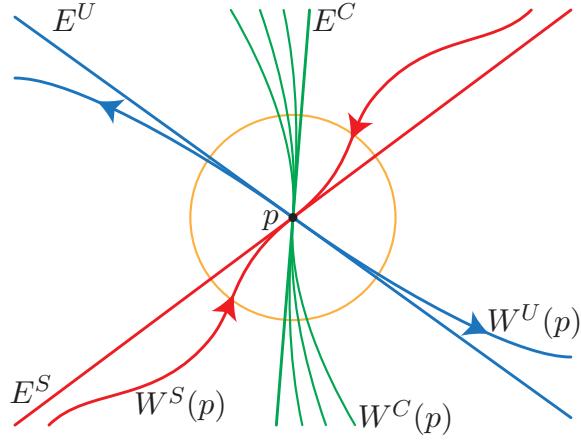


Figure 4.1: A sketch of the stable (red), unstable (blue), and center manifolds (green), along with their respective linear subspaces, that are invariant under the linearized dynamics. Note the existence of multiple center manifolds and the unique unstable/stable manifolds.

4.2 The center manifold

Example 4.1 (Uniqueness of the center manifold). Since Theorem 4.2 gave no guarantees on the uniqueness of the center manifold, now we explore this non-uniqueness in an example.

$$\begin{cases} \dot{x} = x^2 \\ \dot{y} = -y. \end{cases} \quad (4.8)$$

First, linearize at the origin to find the linearized dynamics

$$A = Df(0) = \begin{pmatrix} 0 & 0 \\ 0 & -1 \end{pmatrix}. \quad (4.9)$$

The linearized dynamics is illustrated in Fig. 4.2. We find the invariant subspaces

$$E^C = \text{span} \left\{ \begin{pmatrix} 1 \\ 0 \end{pmatrix} \right\}; \quad E^S = \text{span} \left\{ \begin{pmatrix} 0 \\ 1 \end{pmatrix} \right\}; \quad E^U = \emptyset. \quad (4.10)$$

The nonlinear manifolds are illustrated with the invariant subspaces from the linearization in Fig. 4.2. Observe that there exist infinitely many center manifolds which are all invariant and all tangent to \$E^C\$ at the origin. We also see that although the fixed point at the origin is stable in the linearized system, it is unstable in the nonlinear system.

Remark 4.3. All center manifolds \$W^C(0)\$ are \$\mathcal{C}^\infty\$, but not analytic at \$x = 0\$, for instance

$$y(x) = y_0 e^{-\frac{1}{x_0}} e^{\frac{1}{x}} \implies \forall n \in \mathbb{N}: \frac{d^n}{dx^n} y(0) = 0. \quad (4.11)$$

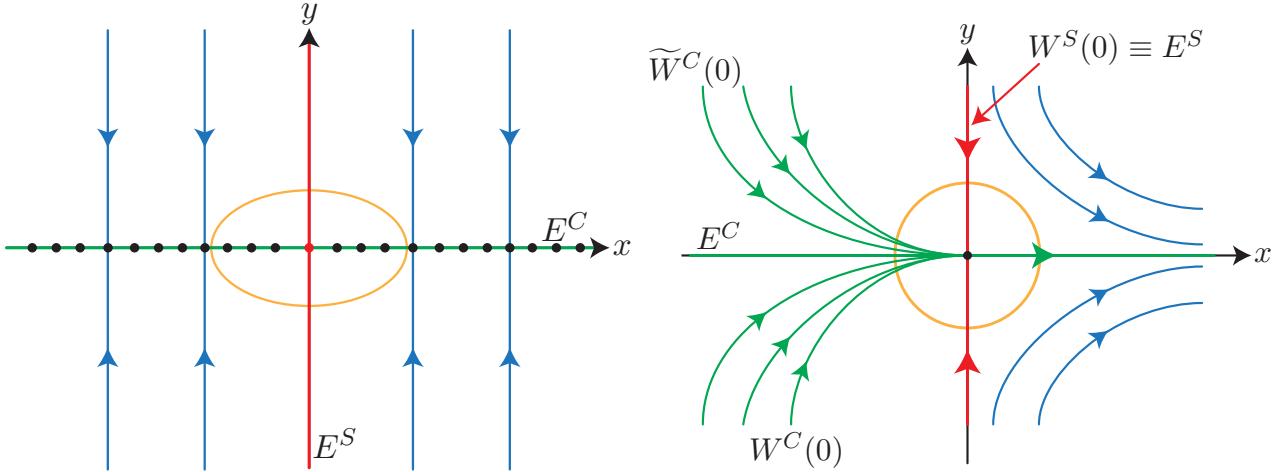


Figure 4.2: Left: The linearized dynamics around the origin. Right: The nonlinear phase portrait.

We now present a method to calculate $W^C(p)$ in general cases.

(i) Consider

$$\dot{z} = F(z); \quad F(0) = 0; \quad z \in \mathbb{R}^{c+d}; \quad F \in \mathcal{C}^r, \quad (4.12)$$

where c represents the number of center directions at the origin ($\dim E^C$) and d denotes the remaining directions ($\dim E^U + \dim E^S$).

(ii) Now block-diagonalize the linearization. This consists of four steps

- (a) First linearize the dynamics to find $\dot{z} = Mz + \mathcal{O}(\|z\|^2)$ with $M = DF(0) \in \mathbb{R}^{(c+d) \times (c+d)}$.
- (b) Define the transformation matrix

$$T = [a_1 \ \dots \ a_c \ \ b_1 \ \dots \ b_d] = [\text{basis in } E^C \ \ \text{basis in } E^U \oplus E^S]. \quad (4.13)$$

- (c) Pass to the basis from the transformation matrix by introducing a new set of coordinates as $z = T\xi$

$$\dot{\xi} = T^{-1}\dot{z} = T^{-1}MT\xi + T^{-1}\mathcal{O}(\|T\xi\|^2) = \begin{pmatrix} A & 0 \\ 0 & B \end{pmatrix} \xi + \mathcal{O}(\|\xi\|^2). \quad (4.14)$$

The matrices A and B are elements of $\mathbb{R}^{c \times c}$ and $\mathbb{R}^{d \times d}$ respectively.

(d) Let $\xi = \begin{pmatrix} x \\ y \end{pmatrix} \in \mathbb{R}^c \times \mathbb{R}^d$, the x -coordinates are aligned with E^C and the y -coordinates are perpendicular to them. We then find

$$\dot{x} = Ax + f(x, y); \quad \dot{y} = By + g(x, y), \quad (4.15)$$

for $f, g \in \mathcal{C}^r$ and $f, g = \mathcal{O}(\|x\|^2, \|y\|^2, \|x\|\|y\|)$. The geometry in these coordinates is depicted in Fig. 4.3. The center manifold is given by

$$W^C(0) = \{(x, y) \in U : y = h(x)\} \quad (4.16)$$

for $h : \mathbb{R}^c \rightarrow \mathbb{R}^d$ and $h \in \mathcal{C}^{r-1}$ as in the theorem.

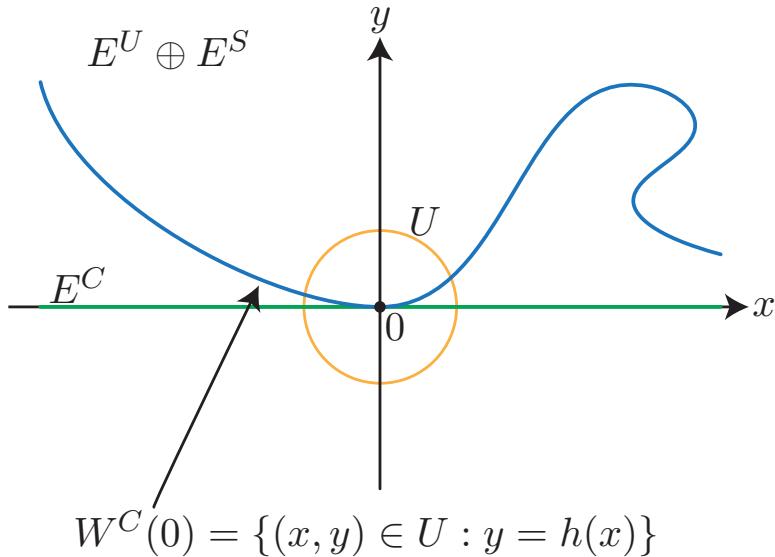


Figure 4.3: The geometry of the nonlinear system in the transformed coordinates aligned with the invariant subspaces of the linearization.

(iii) Now we use the invariance of $W^C(0)$, i.e. for all t , it holds that $y(t) = h(x(t))$. With this we find

$$\dot{y} = Dh(x(t))\dot{x}(t). \quad (4.17)$$

We can now substitute \dot{x} and \dot{y} from (4.15) to get the following nonlinear partial differential equation (PDE) for $h(x)$

$$\boxed{Bh(x) + g(x, h(x)) = Dh(x)[Ax + f(x, h(x))].} \quad (4.18)$$

In addition to (4.18) being nonlinear, the boundary conditions are also unknown. Therefore there is little hope for solving it analytically.

(iv) Instead take the Taylor expansion of (4.18) to approximate the solution

$$h(x) = \underbrace{h(0)}_{=0} + \underbrace{Dh(0)}_{=0} x + \frac{1}{2} \underbrace{D^2 h(0)}_{\text{3-tensor}} \otimes x \otimes x + \mathcal{O}(\|x\|^3), \quad (4.19)$$

where the first two terms are 0 due to the tangency to E^C at 0. We therefore have that $h = O(x^2)$ and we are justified in seeking $W^C(0)$ in this form. We can then restrict the system to the center manifold by substituting $y = h(x)$ to get the reduced dynamics as

$$\boxed{\dot{x} = Ax + f(x, h(x))}. \quad (4.20)$$

Exercise 4.1. Consider the discrete dynamical system

$$x_{n+1} = Ax_n + f(x_n, y_n), \quad (4.21)$$

$$y_{n+1} = By_n + g(x_n, y_n), \quad (4.22)$$

where $x_n \in \mathbb{R}^c$, $y_n \in \mathbb{R}^d$, $A \in \mathbb{R}^{c \times c}$, $B \in \mathbb{R}^{d \times d}$; f and g are \mathcal{C}^r functions with no linear terms. Assume that all eigenvalues of A have modulus one, and none of the eigenvalues of B have modulus one. Then the linearized system at the origin admits a center subspace E^c aligned with the x -coordinate plane.

- (a) Derive a general algebraic equation for the center manifold W^C , which is known to exist by a theorem analogous to the center manifold theorem for continuous dynamical systems.
- (b) Find a cubic order approximation for the center manifold of the discrete system

$$x_{n+1} = x_n + x_n y_n, \quad (4.23)$$

$$y_{n+1} = \lambda y_n - x_n^2, \quad (4.24)$$

where $\lambda \in (0, 1)$.

- (c) Reduce the dynamics to the center manifold and determine the stability of the origin. Verify your results by a numerical simulation of a few initial conditions near the origin.

Example 4.2 (Finding the center manifold). Consider the dynamical system

$$\begin{cases} \dot{x} = xy \\ \dot{y} = -y + \alpha x^2. \end{cases} \quad (4.25)$$

First we linearize at $(0, 0)$ to get

$$M = \begin{pmatrix} [0] & [0] \\ [0] & [-1] \end{pmatrix} \quad (4.26)$$

which is already in block-diagonal form. The dimensions of the stable, unstable, and center subspaces of the linearization are 1, 0, and 1 respectively. Hence the stability type depends on the dynamics on the center manifold $W^C(0)$. We now look for a parameterization of $W^C(0)$ in the form

$$h(x) = ax^2 + bx^3 + cx^4 + \mathcal{O}(x^5). \quad (4.27)$$

Since the right-hand side of the dynamical system was \mathcal{C}^∞ , this expansion could formally be continued to any order. However, the expansion will not converge as that would imply the uniqueness of the center manifold. In this example, we work with a finite, order 4 truncation. Now use the invariance (the PDE (4.18) we derived above) to find

$$\dot{y} = h'(x)\dot{x} = [2ax + 3bx^2 + 4cx^3 + \mathcal{O}(x^4)] x [ax^2 + bx^3 + cx^4 + \mathcal{O}(x^5)]. \quad (4.28)$$

On the other hand, from the dynamical system we know

$$\dot{y} = -h(x) + \alpha x^2 = (\alpha - a)x^2 - bx^3 - cx^4 + \mathcal{O}(x^5). \quad (4.29)$$

Equations (4.28) and (4.29) must hold for all x small enough, therefore the coefficients of equal powers of x in these two equations must match.

$$\mathcal{O}(x^2) : \alpha = a \quad (4.30a)$$

$$\mathcal{O}(x^3) : b = 0 \quad (4.30b)$$

$$\mathcal{O}(x^4) : 2a^2 = -c. \quad (4.30c)$$

Therefore we find

$$h(x) = \alpha x^2 - 2\alpha^2 x^4 + \mathcal{O}(x^5). \quad (4.31)$$

Then the dynamics on $W^C(0)$ becomes

$$\dot{x} = xh(x) = \alpha x^3(1 - 2\alpha x^2) + \mathcal{O}(x^6). \quad (4.32)$$

The dynamics is depicted in Fig. 4.4. For $\alpha > 0$ the origin is unstable, meanwhile for $\alpha < 0$ the origin is asymptotically stable.

The full local stable manifold for $\alpha < 0$ is $\overline{W}^S(0) = U$ and it is of dimension 2. The difference between $\overline{W}^S(0)$ and $W^S(0)$ is that in general the decay rate in $\overline{W}^S(0) - W^S(0)$ is generally weaker than the rate guaranteed in the Center Manifold Theorem.

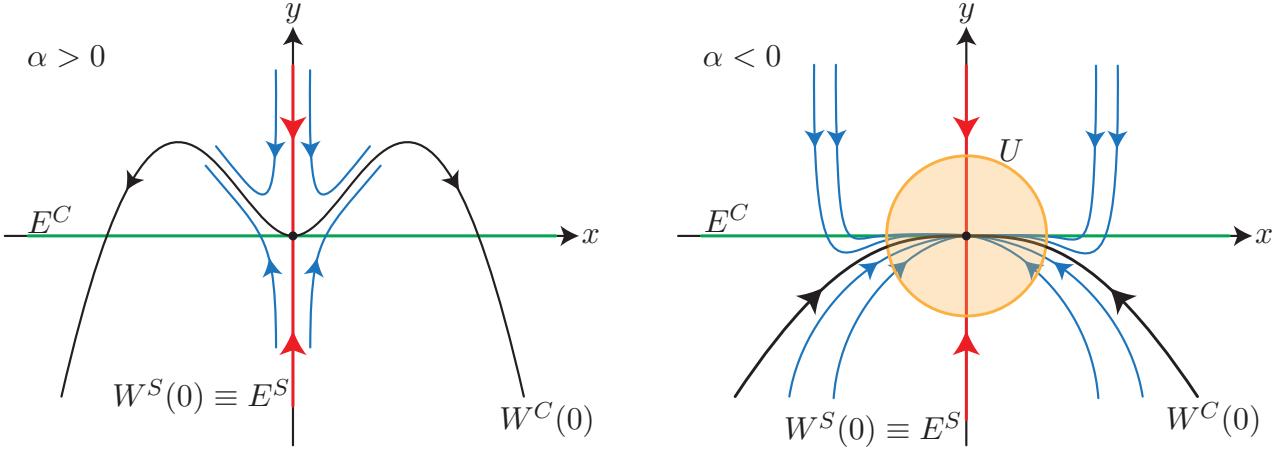


Figure 4.4: Left: The nonlinear dynamics on the center manifold for $\alpha > 0$. Right: The nonlinear dynamics on the center manifold for $\alpha < 0$.

Remark 4.4. The $\mathcal{O}(x^5)$ truncation has two hyperbolic fixed points at $x = \pm \frac{1}{\sqrt{2\alpha}}$, however the full system has no such fixed points. The reason for this is that away from the origin, the $\mathcal{O}(x^6)$ terms are no longer guaranteed to be small relative to the $\mathcal{O}(x^5)$ terms, and the truncation this far away from 0 is not justified.

Notebook 4.2. *Computing a higher-order approximation of the center manifold.*

After this example we would like to explore if the concept of the center manifold is robust, as the existence of eigenvalues with $\text{Re}(\lambda_i) = 0$ is not. We will explore this in an example.

Example 4.3 (Perturbing the previous example). Consider the following perturbed dynamical system

$$\begin{cases} \dot{x} = xy + \varepsilon x \\ \dot{y} = -y + \alpha x^2 \end{cases}; \quad |\varepsilon| \ll 1. \quad (4.33)$$

The linearization of this system yields

$$\begin{cases} \dot{x} = \varepsilon x \\ \dot{y} = -y. \end{cases} \quad (4.34)$$

Now the center manifold disappears as the center subspace E^C disappears for $\varepsilon \neq 0$!

Exercise 4.2. *Construct a cubic-order local approximation for the unstable manifold of the hyperbolic fixed point of the pendulum equation*

$$\ddot{x} + \sin(x) = 0. \quad (4.35)$$

4.3 Center manifolds depending on parameters

We begin with the setup

$$\begin{cases} \dot{x} = Ax + f(x, y, \varepsilon) \\ \dot{y} = By + g(x, y, \varepsilon) \end{cases}; \quad x \in \mathbb{R}^c, y \in \mathbb{R}^d, 0 \leq \|\varepsilon\| \ll 1; \varepsilon \in \mathbb{R}^p; \quad (4.36a)$$

$$f, g \in \mathcal{C}^r, f, g = \mathcal{O}(\|x\|^2, \|y\|^2, \|x\|\|y\|, \|\varepsilon\|\|x\|, \|\varepsilon\|\|y\|). \quad (4.36b)$$

The order $\|\varepsilon\|\|x\|$ and $\|\varepsilon\|\|y\|$ terms are due to the perturbation of the linear part. Now assume $\text{Re}(\lambda_j(A)) = 0$ for $j = 1, \dots, c$ (the center directions) and $\text{Re}(\lambda_j(B)) \neq 0$ for $j = 1, \dots, d$ (the hyperbolic directions). Next, rewrite $\tilde{x} = \begin{pmatrix} x \\ \varepsilon \end{pmatrix}$ and $\tilde{y} = y$ to obtain the system

$$\begin{cases} \dot{\tilde{x}} = \tilde{A}\tilde{x} + \tilde{f}(\tilde{x}, \tilde{y}) \\ \dot{\tilde{y}} = \tilde{B}\tilde{y} + \tilde{g}(\tilde{x}, \tilde{y}) \end{cases}; \quad \tilde{A} = \begin{pmatrix} A & 0 \\ 0 & 0 \end{pmatrix} \in \mathbb{R}^{(c+p) \times (c+p)}; \quad \tilde{f} = \begin{pmatrix} f \\ 0 \end{pmatrix}. \quad (4.37)$$

Here, $\tilde{g} = g$ and $\tilde{B} = B$. Further, note that $\text{span} \left\{ \begin{pmatrix} x \\ 0 \end{pmatrix} \right\}$ is an invariant subspace for \tilde{A} . The eigenvalues of \tilde{A} are the same as those of A and an additional zero with multiplicity p , thus there are $c+p$ center directions and d hyperbolic directions. Applying the center manifold theorem to the fixed point $0 \in \mathbb{R}^{c+p+d}$ of (4.37) we obtain that there exists a \mathcal{C}^{r-1} manifold $\tilde{W}^C(0)$ tangent to E^C at $\begin{pmatrix} \tilde{x} \\ \tilde{y} \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \end{pmatrix}$ which is invariant and is of dimension $c+p$. The geometry of this manifold is illustrated in Fig. 4.5. Note in the figure that there is no dynamics in the ε direction, as $\dot{\varepsilon} = 0$ and that $(x, y) = (0, 0) \in \mathbb{R}^{c+d}$ remains a fixed point for $\varepsilon \neq 0$.

Computing $\tilde{W}^C(0)$ is done in a similar fashion as before. We use the center manifold theorem to get

$$\tilde{y} = y = \tilde{h}(\tilde{x}) = \tilde{h}(x, \varepsilon) = \mathcal{O}(\|x\|^2, \|\varepsilon\|\|x\|, \|\varepsilon\|^2) = \mathcal{O}(\|x\|^2, \|\varepsilon\|\|x\|). \quad (4.38)$$

The order $\|\varepsilon\|^2$ term was dropped as $x = 0$ must remain a fixed point. The function h describes the graph of $W_\varepsilon^C(0)$. Then the reduced dynamics on $W_\varepsilon^C(0)$ is

$$\boxed{\dot{x} = Ax + f(x, \tilde{h}(x, \varepsilon), \varepsilon)}. \quad (4.39)$$

This can then be applied to the perturbed example from above.

Example 4.4 (Revisting the perturbation). Recall the dynamical system

$$\begin{cases} \dot{x} = xy + \varepsilon x \\ \dot{y} = -y + \alpha x^2. \end{cases} \quad (4.40)$$

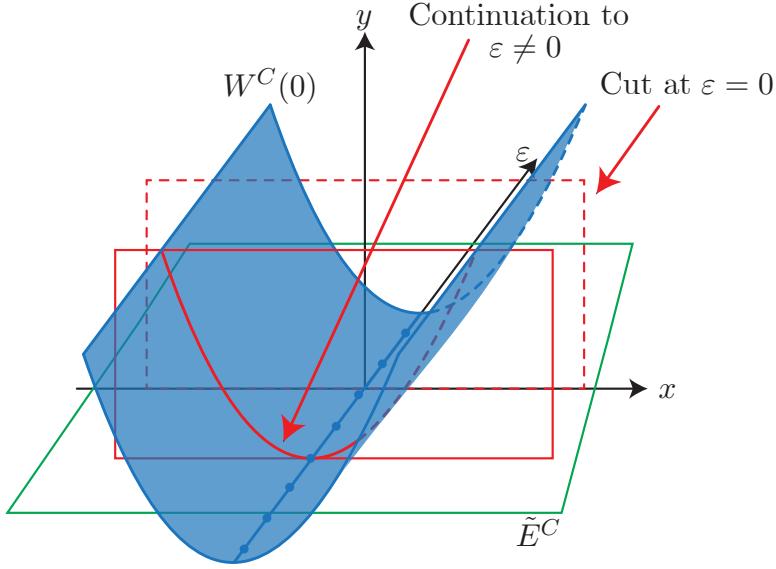


Figure 4.5: Geometry of the center manifold with the perturbation.

We have the persisting fixed point at $(x, y) = (0, 0) \in \mathbb{R}^{c+d}$ and the system is already in standard form with

$$A = 0; \quad B = -1; \quad f(x, y, \varepsilon) = xy + \varepsilon x; \quad g(x, y, \varepsilon) = \alpha x^2. \quad (4.41)$$

We apply the center manifold theorem and get the existence of $W_\varepsilon^C(0)$ for $|\varepsilon| \ll 1$. This manifold satisfies

$$y = \tilde{h}(x, \varepsilon) = ax^2 + bx\varepsilon + c\varepsilon^2 + dx^3 + ex^2\varepsilon + jx\varepsilon^2 + k\varepsilon^3 + lx^4. \quad (4.42)$$

The term $c\varepsilon^2$ must be equal to 0 for all ε such that the fixed point persists, therefore $c = 0$. Next the invariance $y(t) = \tilde{h}(x(t)), \varepsilon)$ is used, taking the time derivative on both sides yields

$$\dot{y} = [2ax + b\varepsilon + \mathcal{O}(2)] \underbrace{[\mathcal{O}(3) + \varepsilon x]}_{= \dot{x} \text{ from ODE and (4.42)}} = 2a\varepsilon x^2 + b\varepsilon^2 x + \mathcal{O}(4). \quad (4.43)$$

The $\mathcal{O}(n)$ designates terms of total degree n , for example x^n or $x^{n-k}\varepsilon^k$. From the ODE we find

$$\dot{y} = -y + \alpha x^2 = -ax^2 - bx\varepsilon - c\varepsilon^2 - dx^3 - ex^2\varepsilon - jx\varepsilon^2 - k\varepsilon^3 - \mathcal{O}(4) + \alpha x^2. \quad (4.44)$$

Comparing equal powers in these two equations we find

$$\mathcal{O}(x^2) : 0 = \alpha - a; \quad \mathcal{O}(\varepsilon^2) : 0 = -c; \quad \mathcal{O}(x\varepsilon) : 0 = -b; \quad (4.45a)$$

$$\mathcal{O}(\varepsilon^3) : 0 = -t; \quad \mathcal{O}(x^3) : 0 = -d; \quad \mathcal{O}(x^2\varepsilon) : 2a = -e; \quad (4.45b)$$

$$\mathcal{O}(x\varepsilon^2) : b = -j. \quad (4.45c)$$

Thus the shape of $W_\varepsilon^C(0)$ is given by

$$y = \tilde{h}(x, \varepsilon) = \alpha(1 - 2\varepsilon)x^2 + \mathcal{O}(4). \quad (4.46)$$

The dynamics on $W_\varepsilon^C(0)$ is

$$\dot{x} = \varepsilon x + \alpha(1 - 2\varepsilon)x^3 + \mathcal{O}(5). \quad (4.47)$$

We can see there is no substantial change in the shape of $W_\varepsilon^C(0)$ relative to the $\varepsilon = 0$ case. The stability type is determined by the sign of ε and a two time-scale dynamic persists.

From this example we may still wonder what effect the higher order terms have on the center manifold.

Notebook 4.3. *Computing the parameter-dependent center manifold.*

4.4 Normal forms

For a general treatment see [Guckenheimer and Holmes, 1990], here we will consider one example to illustrate the idea originally coming from Poincaré.

Example 4.5 (Reduced dynamics on 1-dimensional manifold). Consider the following 1-dimensional dynamical system

$$\dot{x} = x(\mu - x^2) + x^4; \quad 0 \leq |\mu| \ll 1. \quad (4.48)$$

The fixed points are at $x = 0$ and the roots of $g_\mu(x) = \mu - x^2 + x^3$. This function g_μ is illustrated in Fig. 4.6. By plotting x as a function of μ such that $g_\mu(x) = 0$ we get the *bifurcation diagram* as shown in Fig. 4.7.

The fold bifurcation (see caption of Fig. 4.7) is created by quartic (order 4) terms, which become more important away from the origin. The pitchfork bifurcation is already captured by the cubic truncation. We would like to know when the truncation captures the full dynamics correctly near the origin. Poincaré showed that, in fact, the truncated system is topologically equivalent to the full system near the origin by using a change of coordinates to remove $\mathcal{O}(4)$ terms.

- (i) Let $x = y + h_4(y) = y + ay^4 + \mathcal{O}(y^5)$, which is near the identity near the origin, hence it is also invertible near the origin (by the Implicit Function Theorem). Further, this preserves the ODE up to the $\mathcal{O}(3)$ terms.
- (ii) Plug in $x(t)$ and $y(t)$ and take the derivative with respect to time to get

$$\dot{x} = \dot{y}(1 + 4ay^3 + \mathcal{O}(y^4)). \quad (4.49)$$

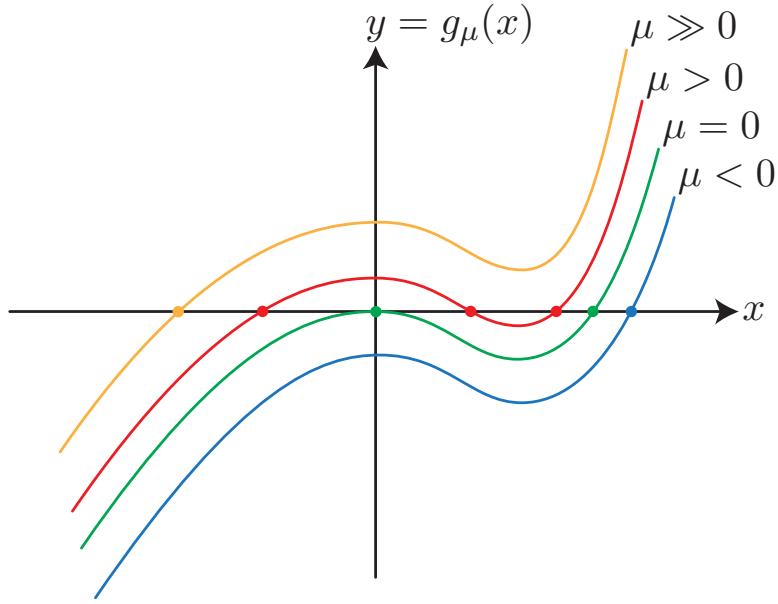


Figure 4.6: The functions g_μ for different values of μ .

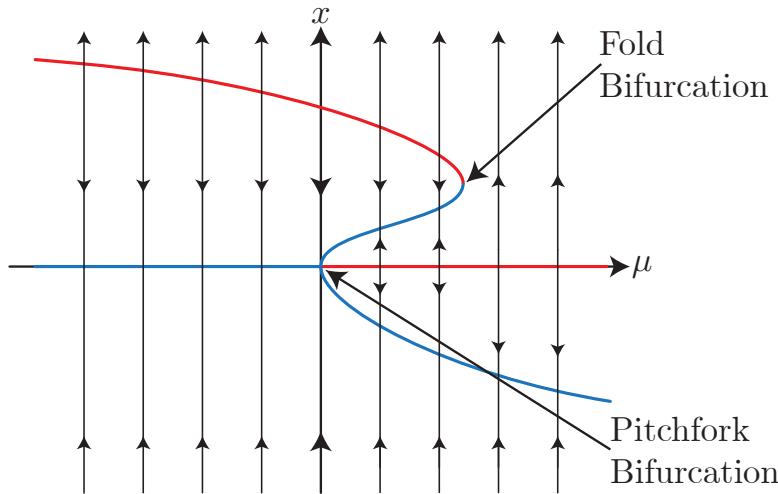


Figure 4.7: Bifurcation diagram for the 1-dimensional dynamical system. Red and blue demarcate if the fixed point at the given (x, μ) pair is stable (blue) or unstable (red). The curve is given by implicitly solving $g_\mu(x) = 0$.

(iii) Now use the ODE and find

$$\dot{x} = \mu x - x^3 + x^4 = \mu(y + ay^4) - (y + ay^4)^3 + (y - ay^4)^5 + \dots \quad (4.50)$$

(iv) Combine the previous two steps and calculate

$$\dot{y} = [1 + 4ay^3 + \mathcal{O}(y^4)]^{-1} [\mu y + a\mu y^4 - y^3 + y^4 + \mathcal{O}(5)]. \quad (4.51)$$

At this point recall the alternating series (it could be verified with Taylor expansion)

$$\frac{1}{1+z} = 1 - z + \mathcal{O}(z^2); \quad 0 \leq |z| \ll 1. \quad (4.52)$$

We also note that a generalized result holds for an operator $A \in R^{m \times m}$, which could be used in higher dimensional calculations. This is referred to as the Neumann series. Applying this to the left term in the formula for \dot{y} yields

$$[1 + 4ay^3 + \mathcal{O}(y^4)]^{-1} = 1 - 4ay^3 + \mathcal{O}(y^4). \quad (4.53)$$

Therefore we find

$$\dot{y} = \mu y - y^3 + y^4 \underbrace{(-4a\mu + a\mu + 1)}_{\text{choose } a \text{ such that this } = 0} + \mathcal{O}(y^5). \quad (4.54)$$

We may now choose the parameter a to make the coefficient of the quartic term 0. The a that fulfills this is $a = \frac{1}{3\mu}$, using this we find

$$\boxed{\dot{y} = \mu y - y^3 + \mathcal{O}(y^5)}. \quad (4.55)$$

This transformation has removed the quartic terms from the equation.

(v) Now we remove the $\mathcal{O}(y^5)$ terms similarly. First set

$$y = \xi + h_5(\xi) = \xi + b\xi^5 + \mathcal{O}(\xi^6) \quad (4.56)$$

and then continue as before, but with y now playing the role of x and ξ playing the role of y .

(vi) The successive sequence of near identity coordinate changes turns out to converge usually. In general, it depends on the type of problem, sometimes resonant terms are not removable and must stay as they are crucial to the dynamics. These resonant terms depend only on the linear part of the RHS, for more information see Chapter 3 of [Guckenheimer and Holmes, 1990].

Thus

$$\boxed{\dot{x} = \mu x - x^3} \quad (4.57)$$

is the *normal form* for the ODE for the study of bifurcations at the origin for $0 \leq \mu \ll 1$. It is topologically equivalent to the full system near $x = 0$ and captures the pitchfork bifurcation.

4.5 Bifurcations

A *bifurcation* is a qualitative change in the dynamical system

$$\dot{x} = f(x, \mu); \quad x \in \mathbb{R}^n; \quad \mu \in \mathbb{R}^p. \quad (4.58)$$

Linear stability analysis led to reducing to the center manifold (depending on parameters). From there we moved to normal forms which enable the analysis of qualitative dynamics under varying parameters.

Definition 4.2 (Local bifurcation). A *local bifurcation* is a qualitative change in the behavior of a system near a non-hyperbolic equilibrium. More precisely, a *bifurcation* occurs in (4.58) at $\mu = \mu_0$ near the fixed point $x = 0$ if there exists no neighborhood of $x = 0$ in which $\dot{x} = f(x, \mu_0)$ is topologically equivalent to all systems $\dot{x} = f(x, \mu)$ for $\|\mu - \mu_0\|$ small enough.

This idea of a bifurcation can be illustrated by a bifurcation surface which separates the space of dynamical systems into components. Within each component the dynamical systems are topologically equivalent, however elements from separate components are not. This is sketched in Fig. 4.8.

Bifurcation point for
for a given family

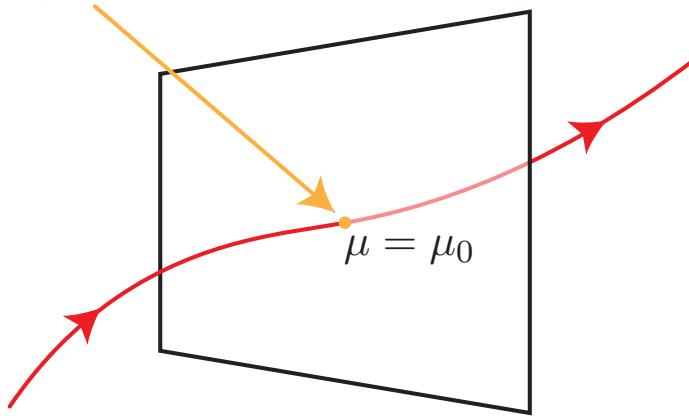


Figure 4.8: Illustration of a bifurcation surface. The near side of the surface is one component, and the far side the other. The red path is a family of dynamical systems $\{\dot{x} = f(x, \mu)\}_{\mu \in \mathbb{R}^p}$, going through the bifurcation point μ_0 .

We wish to understand what happens in the case that a given family of dynamical systems is nongeneric (atypical). For instance in the case that the family forms a tangency to the bifurcation surface, in which case a bifurcation does not take place, hence the family of dynamical systems is not general enough to capture all possible dynamics.

Example 4.6 (Nongeneric family of dynamical systems). In comparison to the family taken previously consider

$$\dot{x} = -a^2x - x^3; \quad \mu = -a^2 \leq 0. \quad (4.59)$$

For every μ which we consider, there is only one fixed point $x = 0$ and it is stable for all values of μ . However, we have unwittingly missed the full picture, as for $\mu > 0$ there are three fixed points, one of which are unstable. This situation is shown in Fig. 4.9.

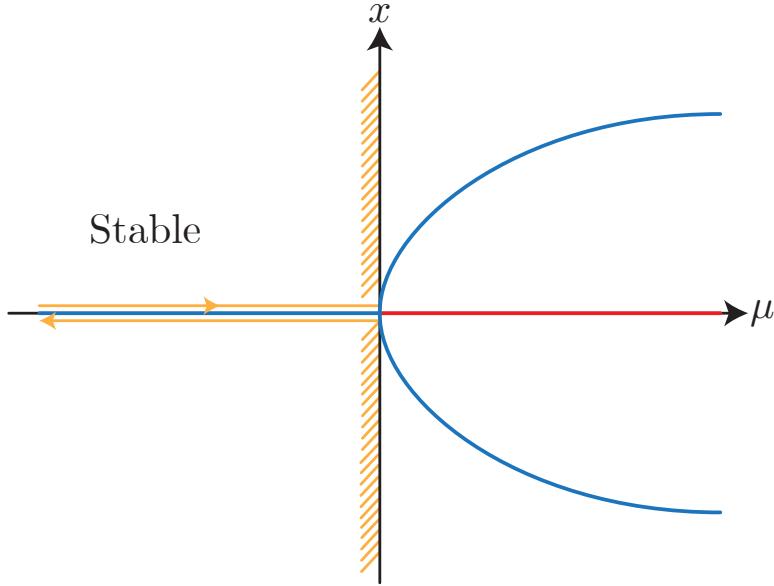


Figure 4.9: A nongeneric family of dynamical systems. We only see what happens to the left of the x -axis, and miss everything to the right, hence our family is tangent to the bifurcation surface.

This idea of nongeneric families motivates our next definition, which is also depicted in Fig. 4.10.

Definition 4.3 (Universal unfolding). A parameterized family of dynamical systems crossing all nearby topological equivalence classes as the parameters vary is called a *universal unfolding*.

Definition 4.4 (Codimension of a bifurcation). The *codimension of a bifurcation* is the minimum number of parameters required for a universal unfolding. Thus a more degenerate bifurcation requires a larger codimension.

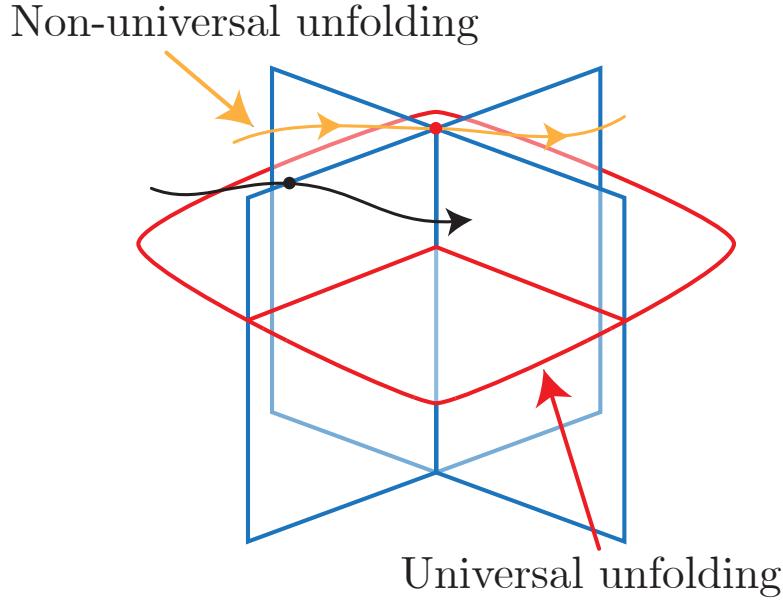


Figure 4.10: An example of universal unfolding (red) for the red bifurcation point which crosses the four topologically equivalent classes (components) created by the two bifurcation surfaces (blue). Furthermore, a nonuniversal unfolding is shown by the yellow 1-dimensional path at the top. Another universal unfolding in for the black bifurcation point, is shown by the 1-dimensional black family.

4.6 Codimension 1 bifurcations

We begin by considering different types of center manifolds. We start by classifying these by the number of zero eigenvalues which appear in the linearization.

Notebook 4.4. *Illustration of one- and two-dimensional normal forms.*

(i) **Single zero eigenvalue** Our system is as follows

$$\dot{y} = f(y, \mu), \quad y \in \mathbb{R}^n, \quad \mu \in \mathbb{R}^p; \quad f(0, 0) = 0; \quad \dim E^C = 1 \implies \dim W_\mu^C(0) = 1. \quad (4.60)$$

Thus we have a single zero eigenvalue, an example eigenvalue constellation which fulfills this is given in Fig. 4.11. The universal unfolding (generally parameterized by the normal form on $W_\mu^C(0)$) is given by

$$\dot{x} = \tilde{\mu} \pm x^2; \quad x \in \mathbb{R}; \quad \tilde{\mu} \in \mathbb{R}. \quad (4.61)$$

This is a codimension 1 bifurcation, i.e. only 1 parameter is needed for the universal unfolding. Hence we find a bifurcation diagram as shown in Fig. 4.12. This is called a

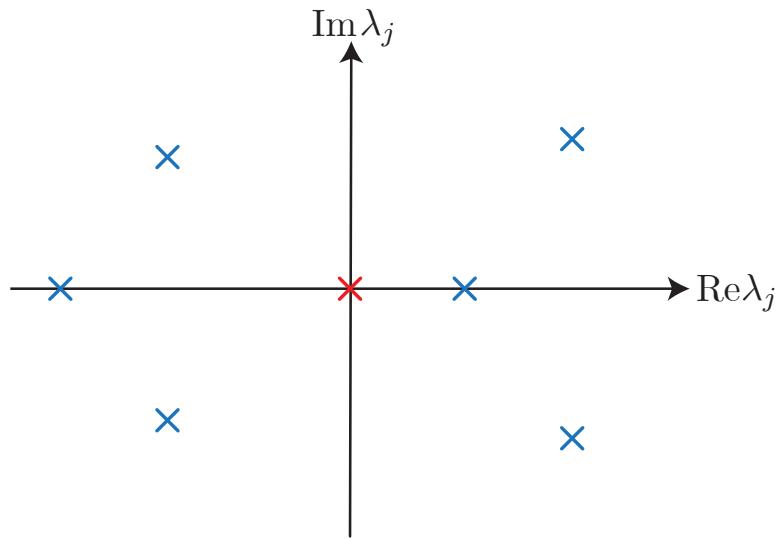


Figure 4.11: An example eigenvalue constellation with a single zero eigenvalue.

fold or *saddle node* bifurcation, in this example it is *supercritical* as we show the $-x^2$ case. *Subcriticality* occurs when the diagram is mirrored with respect to the x -axis, i.e. when the fixed points disappear for $\mu > 0$. The sign which is actually used is problem dependent.

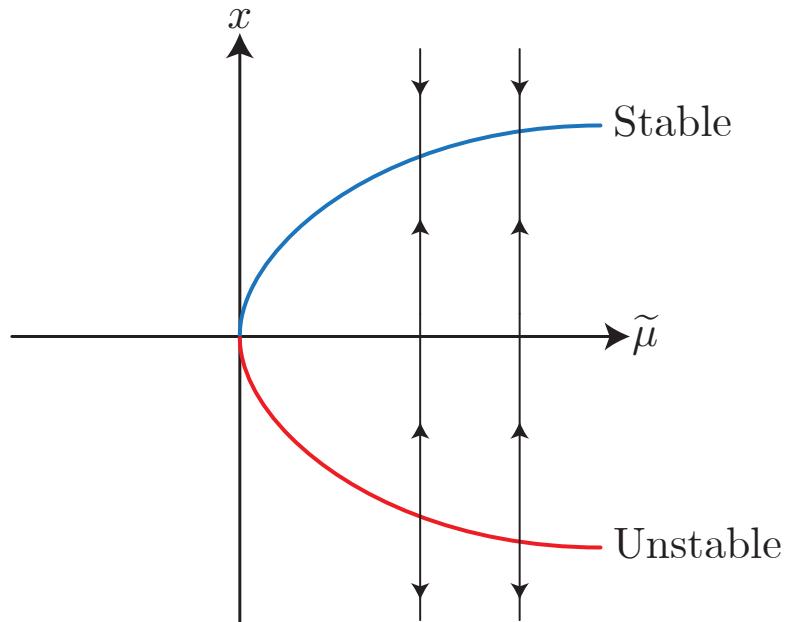


Figure 4.12: A codimension 1 fold bifurcation, which is supercritical.

Remark 4.5. *Hysteresis* is the dependence of the output on past input and is often related to interplay between two such fold bifurcations as shown in Fig. 4.13.

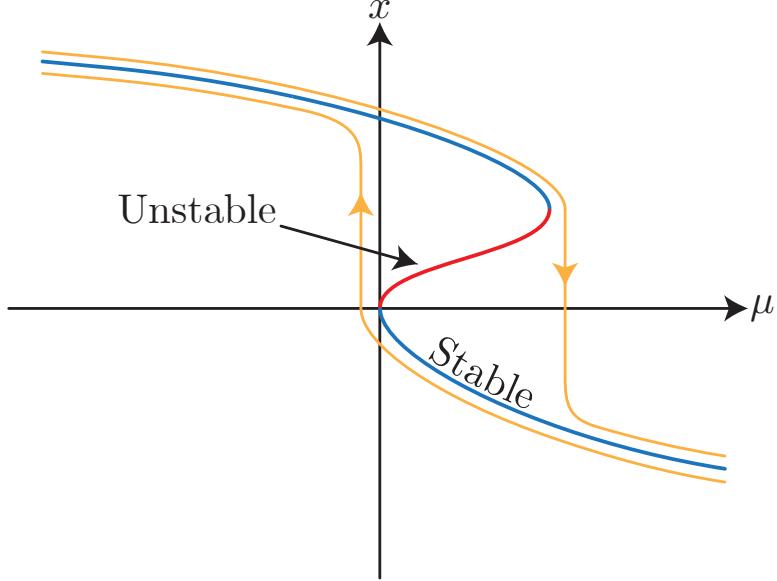


Figure 4.13: An example of hysteresis caused by two interacting fold bifurcations. In this case, the hysteresis is a result of purely deterministic dynamics, no additional memory (or delay) terms were needed.

(ii) **Single zero eigenvalue & origin remains a fixed point** Now we have more conditions and our setup is as follows

$$\dot{y} = f(y, \mu), \quad y \in \mathbb{R}^n, \quad \mu \in \mathbb{R}^p; \quad f(0, \mu) = 0, \quad \forall \mu; \quad \dim E^C = 1 \implies \dim W_\mu^C(0) = 1. \quad (4.62)$$

Thus we have the additional degeneracy $D_\mu f(0, 0) = 0$. We get the universal unfolding

$$\dot{x} = \tilde{\mu}x \pm x^2 = x(\tilde{\mu} \pm x); \quad x \in \mathbb{R}; \quad \tilde{\mu} \in \mathbb{R}. \quad (4.63)$$

This so called *transcritical bifurcation* is depicted in Fig. 4.14. In such bifurcations the observed stable state does not vary until the bifurcation, after which it begins to vary linearly.

Example 4.7 (A physical example of a transcritical bifurcation). The laser threshold is a physical example of a transcritical bifurcation [Strogatz, 2000, Haken, 1977]. The simplest model is of an active material held between two partially reflecting mirrors, with an energy

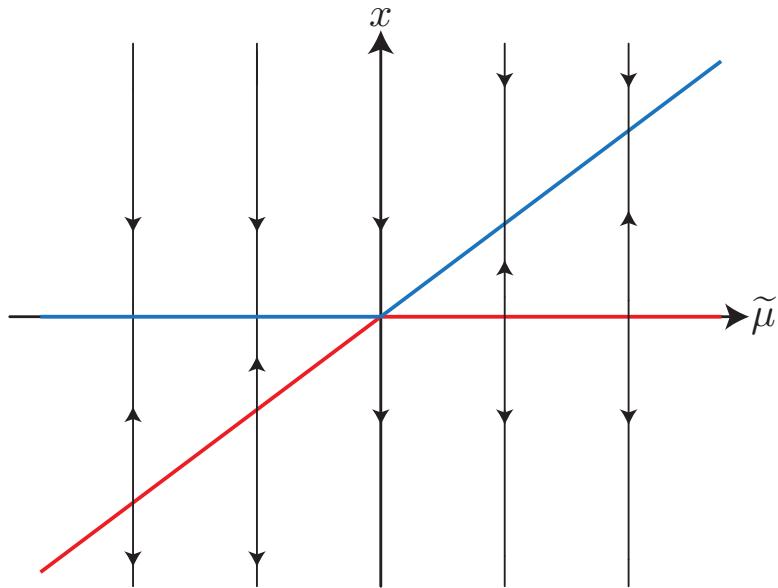


Figure 4.14: A transcritical bifurcation.

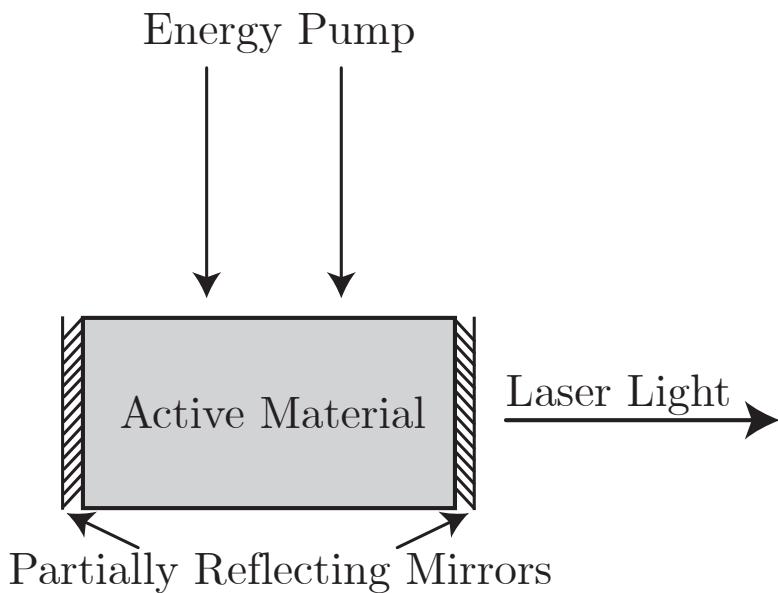


Figure 4.15: The simplest model of for the laser threshold example.

pump into the active material, from which a laser is emitted. This is depicted in Fig. 4.15.

The phenomenon we can observe in this system is that a coherent laser light is only gener-

ated over a certain pumping strength. Let $n(t)$ denote the number of photons emitted. Therefore \dot{n} is equal to the gain minus the loss, i.e. if N is the number of excited atoms and G is the gain from stimulated emission, then

$$\dot{n} = GnN - kn. \quad (4.64)$$

The number of excited atoms N decreases as an excited atom emits a photon. Therefore if α is the rate at which atoms drop back to the ground state and N_0 is the number of excited atoms (fixed without laser action) then we have

$$N(t) = N_0 - \alpha n. \quad (4.65)$$

Combining these two equations we find

$$\dot{n} = Gn(N_0 - \alpha n) - kn = \underbrace{(GN_0 - k)}_{=\mu} n - \alpha Gn^2. \quad (4.66)$$

Hence, we have a transcritical bifurcation, and the laser threshold is at $\frac{k}{G}$, as shown in Fig. 4.16.

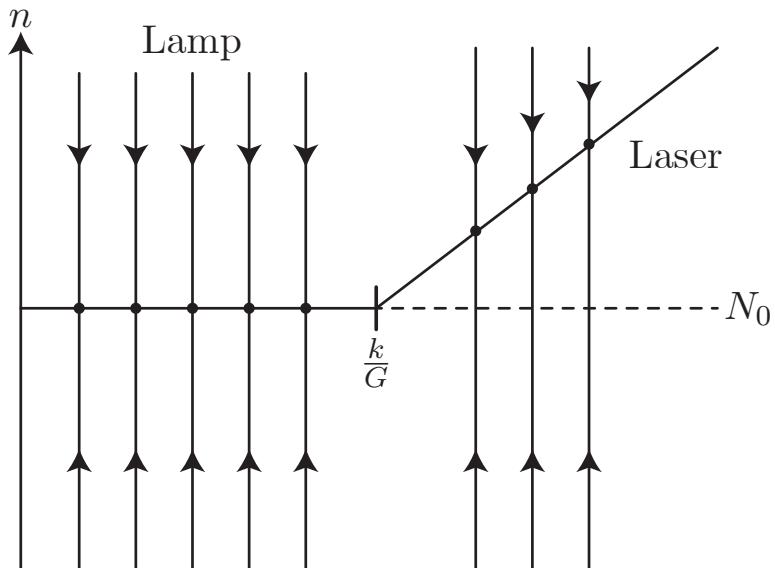


Figure 4.16: The laser threshold bifurcation at $\frac{k}{G}$.

- (iii) **Single zero eigenvalue, origin remains a fixed point, & RHS is an odd function**
We have the same setup as previously, with the additional condition that

$$f(y, 0) = -f(-y, 0) \implies D_y^2 f(y, 0) = -D_y^2 f(-y, 0) \implies D_y^2 f(0, 0) = 0. \quad (4.67)$$

This leads to the universal unfolding

$$\dot{x} = \tilde{\mu}x \pm x^3; \quad x \in \mathbb{R}; \quad \tilde{\mu} \in \mathbb{R}. \quad (4.68)$$

This type of bifurcation is called a *pitchfork bifurcation*. It is *supercritical* for the $-x^3$ case (cf. Fig. 4.17), and subcritical for the $+x^3$ case. The actual sign is problem dependent.

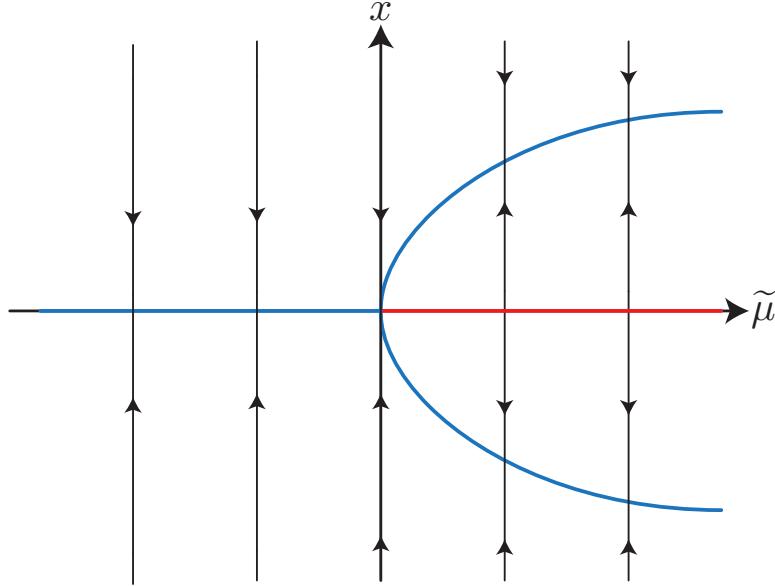


Figure 4.17: A supercritical (i.e with a $-$ sign) pitchfork bifurcation.

Example 4.8 (Buckling beam bifurcation). A beam fixed at the bottom with a downward force applied towards the top is an example of such a pitchfork bifurcation. As the force exceeds a critical threshold, the middle position loses stability and two new symmetric, stable fixed points arise. This is illustrated in Fig. 4.18.

Exercise 4.3. Consider the quadratic Duffing equation

$$\dot{u} = v, \quad (4.69)$$

$$\dot{v} = \beta u - u^2 - \delta v, \quad (4.70)$$

where $\delta > 0$, and $0 \leq |\beta| \ll 1$.

- (a) Construct a β -dependent center manifold up to quadratic order near the origin for small β values.
- (b) Construct a stability diagram for the reduced system on the center manifold using β as a bifurcation parameter.

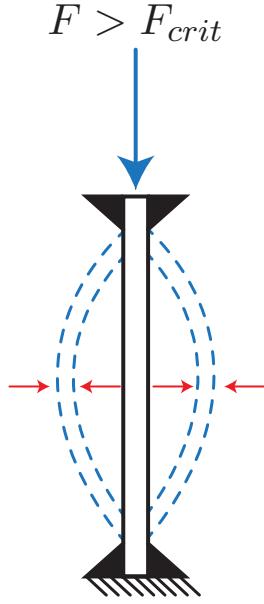


Figure 4.18: The buckling beam pitchfork bifurcation.

- (iv) **Purely imaginary pair of eigenvalues** Now we have the following setup, with the eigenvalue constellation shown in Fig. 4.19.

$$f(0, 0) = 0; \quad \dim E^C = 2; \quad \lambda_n(0) = \bar{\lambda}_{n-1}(0) = \pm i\omega \neq 0. \quad (4.71)$$

Such constellations are common in oscillatory systems at the loss of stability. This loss

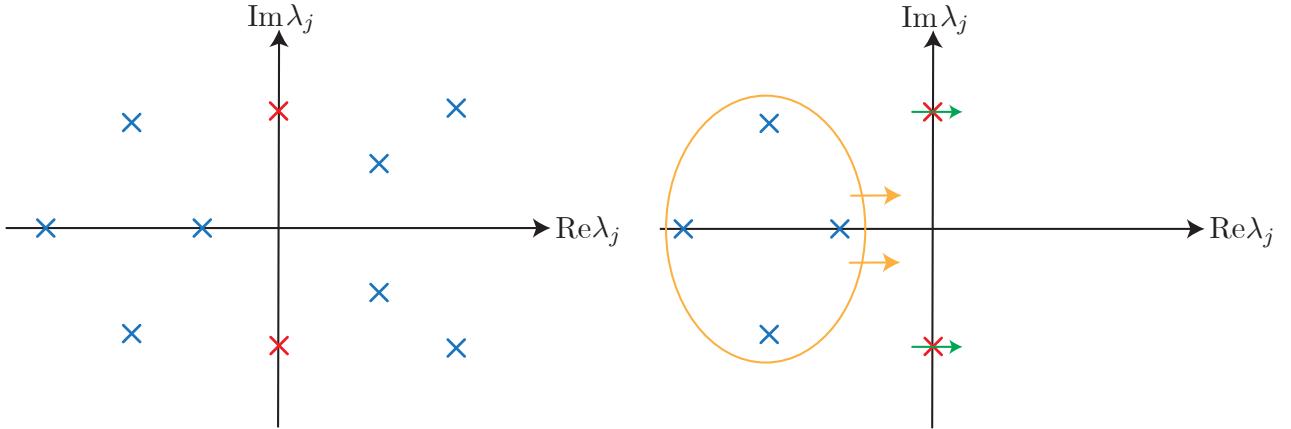


Figure 4.19: Left: Eigenvalue constellation for purely imaginary pair of eigenvalues. Right: Loss of stability in oscillating system as the pair of eigenvalues cross the imaginary axis.

of stability is sketched in the right pane of Fig. 4.19. The reduced system dynamics on

the 2-dimensional $W^C(0)$ for $\mu = 0$ is

$$\begin{pmatrix} \dot{u} \\ \dot{v} \end{pmatrix} = \begin{pmatrix} 0 & -\omega \\ \omega & 0 \end{pmatrix} \begin{pmatrix} u \\ v \end{pmatrix} + \begin{pmatrix} \tilde{f}(u, v) \\ \tilde{g}(u, v) \end{pmatrix}. \quad (4.72)$$

The coordinates (u, v) correspond to the basis given by $[\text{Re}(e_n), \text{Im}(e_n)]$. The normal form on $W_\mu^C(0)$ is

$$\dot{r} = r (d(\mu) + a(\mu)r^2) + \mathcal{O}(r^5); \quad d(\mu) = \text{Re}(\lambda_n(\mu)) \quad (4.73a)$$

$$\dot{\theta} = \omega(\mu) + e(\mu)r^2 + \mathcal{O}(r^4); \quad \omega(\mu) = \text{Im}(\lambda_n(\mu)). \quad (4.73b)$$

The functions a and e depend on the given \tilde{f} and \tilde{g} (thereby also on f). We then define

$$d_0 = \frac{d}{d\mu} \text{Re} [\lambda_n(\mu)]|_{\mu=0}; \quad \omega_0 = \text{Im}(\lambda_n(0)) \quad (4.74a)$$

$$a_0 = a(0) = \frac{1}{16} \left[\tilde{f}_{uuu} + \tilde{f}_{uvv} + \tilde{g}_{uuv} + \tilde{g}_{vvv} \right] \Big|_{(u,v)=0} \quad (4.74b)$$

$$+ \frac{1}{16\omega} \left[\tilde{f}_{uv}(\tilde{f}_{uu} + \tilde{f}_{vv}) - \tilde{g}_{uv}(\tilde{g}_{uu} + \tilde{g}_{vv}) - \tilde{f}_{uu}\tilde{g}_{uu} + \tilde{f}_{vv}\tilde{g}_{vv} \right] \Big|_{(u,v)=0}. \quad (4.74c)$$

Note that we must start from the standard form (4.72).

Building on this last case, we explore the extended center manifold in the next theorem.

Theorem 4.6 (Hopf-Bogdanov). *Assume the following*

- (i) *The eigenvalues λ_n and $\bar{\lambda}_n$ do cross the imaginary axis for $\mu \neq 0$, i.e. $d_0 \neq 0$.*
- (ii) *The leading order nonlinearity in \dot{r} is nonzero, i.e. $a_0 \neq 0$. Then there exists a unique extended center manifold $W_\mu^C(0)$ for $0 \leq |\mu| \ll 1$, on which the dynamical system is locally topologically equivalent to the universal unfolding*

$$\boxed{\dot{r} = r (d_0\mu + a_0r^2); \quad \dot{\theta} = \omega_0 + e_0r^2 + b_0\mu,} \quad (4.75)$$

for $b_0, e_0 \in \mathbb{R}$.

From here we explore two cases for the signs of d_0 and a_0 .

- (i) First, when each are strictly positive $d_0, a_0 > 0$. This yields a subcritical pitchfork bifurcation in r and a continued rotation in θ . The bifurcation of r and the dynamics of u and v for $\mu < 0$ are illustrated in Fig. 4.20, for $|\mu|$ and $|r|$ small enough. There is an increased sensitivity to perturbation as μ increases (the domain of attraction shrinks). This is shown in Fig. 4.21. Such behavior can be observed in the transition to turbulence in pipe flows. In that context, the equation is referred to as the Stuart-Landau equation.

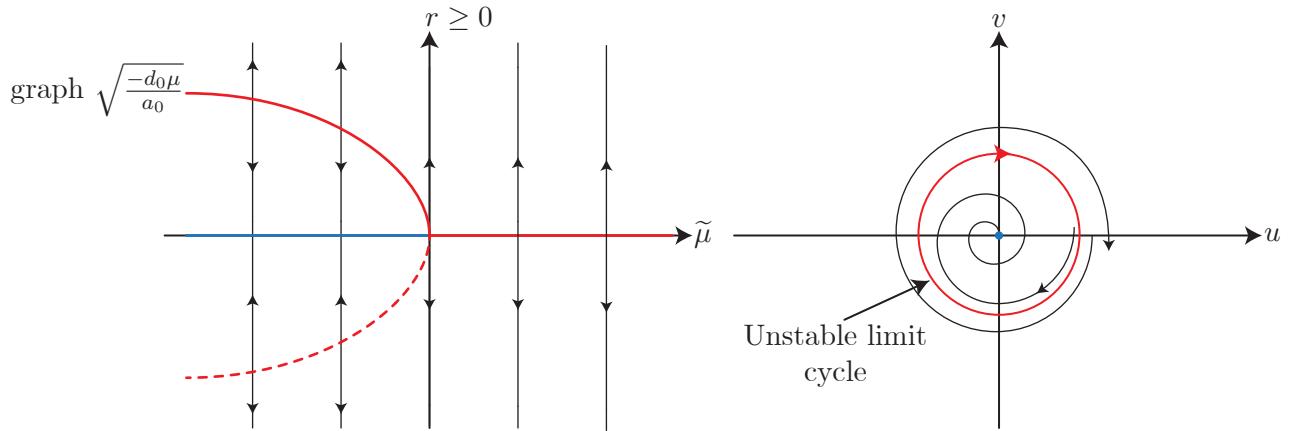


Figure 4.20: Left: The subcritical pitchfork bifurcation of r . The red arrow points towards the $r = \sqrt{\frac{-d_0 \mu}{a_0}}$, the symmetric negative part of this curve is dashed, as negative values of r are not relevant for this problem (the radius cannot be negative). Right: The effect of $\mu < 0$ on the system dynamics leading to an unstable limit cycle (red) in the $u - v$ plane.

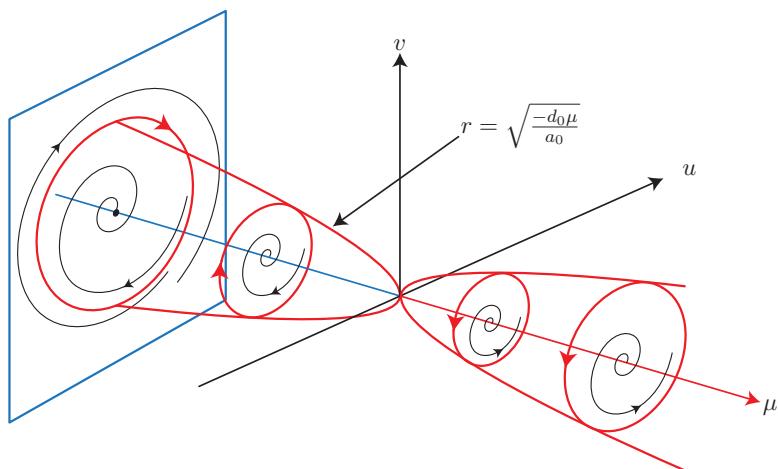


Figure 4.21: The full bifurcation diagram for the dynamical system in case (i).

- (ii) Next, the two coefficients have opposite signs $d_0 > 0$ and $a_0 < 0$. This leads to the supercritical Hopf bifurcation and a stable limit cycle as shown in Fig. 4.22

There is a sudden development of stable periodic oscillation as μ is increased. Often the parameter μ is a steady speed in the system which is changed gradually. Examples of this are

- increasing wind speed leading to bridge oscillations and collapse,

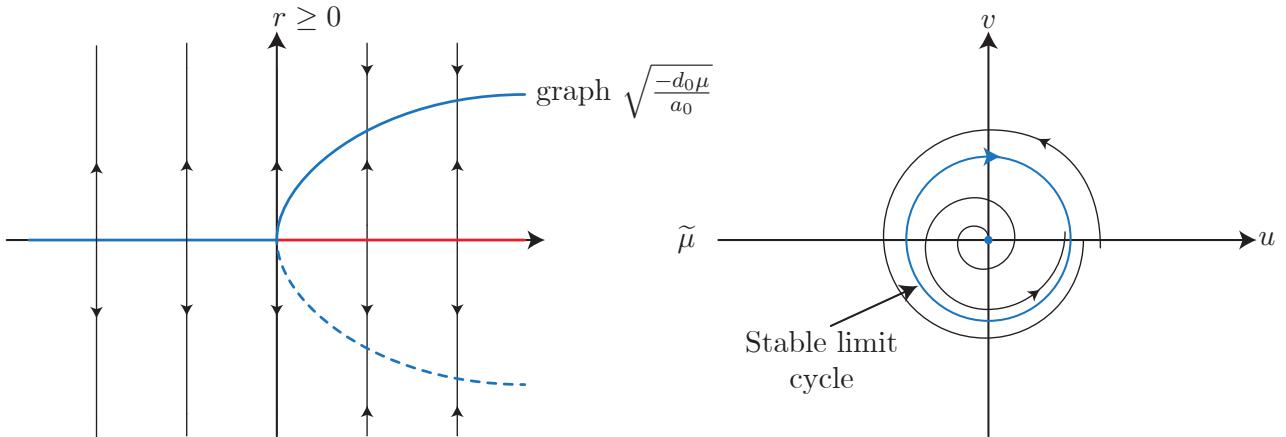


Figure 4.22: Left: The bifurcation in r , the blue arrow designates the curve given by $r = \sqrt{\frac{-d_0\mu}{a_0}}$, again the negative part is not relevant as the radius is always non-negative. Right: The stable limit cycle that arises for this system.

- increasing flight speed leading to wing flutter,
- increasing driving speed leading to death wobble on motorcycles.

Notebook 4.5. *Illustration of codimension-1 bifurcations.*

Exercise 4.4. Consider a dynamical system that has a pair of purely imaginary eigenvalues at its fixed point for the parameter value $\mu = 0$. As we have seen, a linear change of coordinates and a center manifold reduction gives the two-dimensional reduced dynamical system

$$\dot{x} = \delta(\mu)x - \omega(\mu)y + f(x, y, \mu), \quad (4.76)$$

$$\dot{y} = \delta(\mu)y + \omega(\mu)x + g(x, y, \mu), \quad (4.77)$$

where $\delta(\mu) = \operatorname{Re} \lambda(\mu)$, $\omega(\mu) = \operatorname{Im} \lambda(\mu)$. (Here $\lambda(\mu)$ and $\bar{\lambda}(\mu)$ is the pair of complex eigenvalues that crosses the imaginary axis at $\mu = 0$.)

Recall that in polar coordinates, the truncated normal form of (4.76) can be written as

$$\dot{r} = r(d_0\mu + a_0r^2), \quad (4.78)$$

$$\dot{\theta} = \omega_0 + b_0\mu + c_0r^2, \quad (4.79)$$

where

$$d_0 = \delta'(0), \quad \omega_0 = \omega(0) \quad (4.80)$$

$$a_0 = \frac{1}{16} [f_{xxx} + f_{xyy} + g_{xxy} + g_{yyy}]_{x=y=0, \mu=0} \quad (4.81)$$

$$+ \frac{1}{16\omega_0} [f_{xy}(f_{xx} + f_{yy}) - g_{xy}(g_{xx} + g_{yy}) - f_{xx}g_{xx} + f_{yy}g_{yy}]_{x=y=0, \mu=0}. \quad (4.82)$$

These classic formulae are used in all applications where Hopf bifurcations are analyzed. As an application of these results, consider now the stick-slip oscillator

$$m\ddot{x} + c\dot{x} + kx = F_f, \quad F_f = mg\mu_0 (1 + e^{-\beta|v_0 - \dot{x}|}) \operatorname{sign}(v_0 - \dot{x}), \quad (4.83)$$

where m is the mass of the oscillator, g is the constant of gravity, $\beta > 0$ is a constant, μ_0 is the Coulomb (static) friction coefficient, v_0 is the speed of the belt, x is the position of the mass on the belt, c is the coefficient of viscous damping, and k is the spring coefficient (see Fig. 4.23).

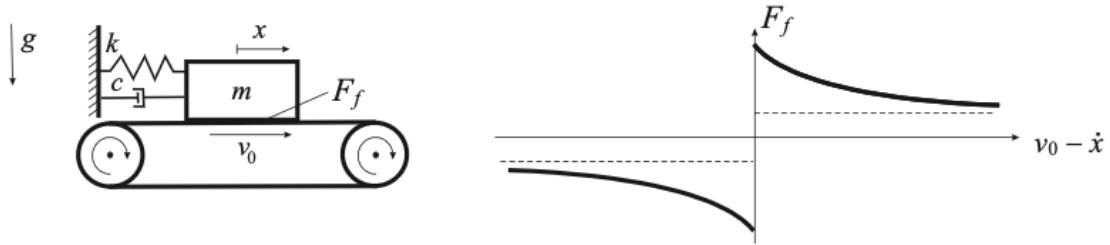


Figure 4.23: Stick-slip oscillator and its dry-friction force as a function of the relative velocity between the mass and the belt.

- (i) Find a condition under which the system has an asymptotically stable fixed point.
- (ii) Show that a subcritical Hopf bifurcation takes place when the above condition is violated. (Use v_0 as a bifurcation parameter.)
- (iii) Calculate the approximate amplitude of the bifurcating limit cycle.

Notebook 4.6. *Hopf bifurcation in the stick-slip oscillator.*

Chapter 5

Nonlinear dynamical systems on the plane

For this chapter the general setup will be

$$\dot{x} = f(x); \quad x \in \mathbb{R}^2; \quad f \in \mathcal{C}^1. \quad (5.1)$$

5.1 One degree of freedom conservative mechanical systems

Write $x = \begin{pmatrix} x_1 \\ x_2 \end{pmatrix}$, with x_1 denoting the position and $x_2 = \dot{x}_1$ denoting the speed. The total mechanical energy is given by $E(x) = \frac{1}{2}mx_2^2 + V(x_1)$, with the mass m and the potential function V . Newton's law then gives the equations of motion

$$m\ddot{x}_1 = -\frac{dV(x_1)}{dx_1} \implies \begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -\frac{1}{m}\frac{dV(x_1)}{dx_1}. \end{cases} \quad (5.2)$$

All of the forces derive from a time-independent potential. The energy is a first integral (conserved quantity of such systems), i.e.

$$\frac{d}{dt}E(x(t)) = \frac{\partial E}{\partial x_1}\dot{x}_1 + \frac{\partial E}{\partial x_2}\dot{x}_2 = 0. \quad (5.3)$$

Therefore on any given trajectory $x(t)$ the energy $E(x(t)) = E_0$ is constant. Hence we can derive an explicit equation for the velocity along a trajectory which suppresses the time dependence

$$x_2 = \pm \sqrt{\frac{2}{m}(E_0 - V(x_1))}. \quad (5.4)$$

This leads to multiple consequences for these systems

- (i) Trajectories form symmetric pairs (w.r.t. the x_1 -axis) of the same energy.
- (ii) There is a clockwise orientation for trajectories due to

$$\dot{x}_1 = x_2 \implies \begin{cases} x_2 > 0 \implies x_1 \text{ increases} \\ x_2 < 0 \implies x_1 \text{ decreases.} \end{cases} \quad (5.5)$$

This is shown graphically in Fig. 5.1.

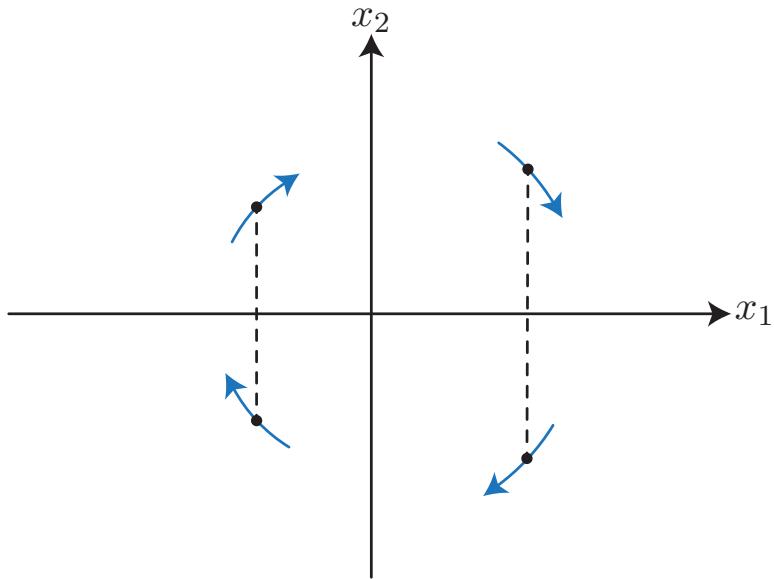


Figure 5.1: The clockwise orientation of trajectories in 1 degree of freedom conservative mechanical systems.

- (iii) Local minima of the potential give rise to a center fixed point surrounded by closed orbits. The minimum x^* is a fixed point if $\dot{x}_2 = 0$ (which occurs if $V(x_1^*) = E_0$) as $\dot{x}_2 \propto V'(x_1^*) = 0$. The trajectories of the closed orbits can be found using $x_2 = \pm \sqrt{\frac{2}{m}(E_0 - V(x_1))}$. Such a fixed point at a local minimum of the potential surrounded by closed orbits is shown in Fig. 5.2.
- (iv) At local maxima of the potential saddle type fixed points arise. The maximum x^* is a fixed point if $\dot{x}_2 = 0$ (which occurs if $V(x_1^*) = E_0$) as $V'(x_2^*) = 0$. The velocity can be

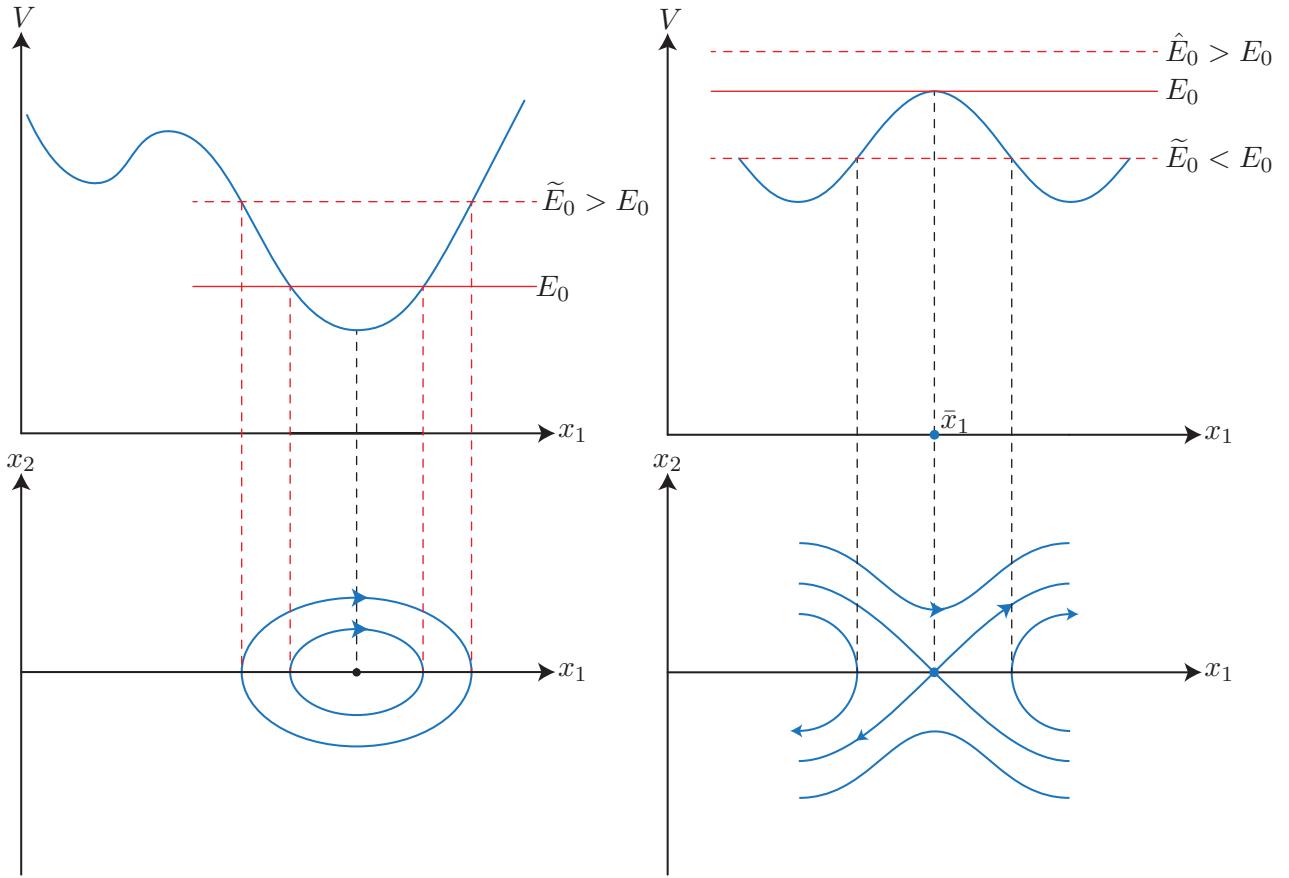


Figure 5.2: Left: Closed orbits around the fixed point at the local minimum of the potential function. Right: The saddle type fixed point at the local maxima of the potential. The clockwise orientations comes from consequence (ii).

approximated as follows

$$x_2 = \pm \sqrt{\frac{2}{m}(E_0 - V(x_1))} \quad (5.6a)$$

$$= \pm \sqrt{\frac{2}{m} \underbrace{(E_0 - V(x_1^*))}_{=0} - \underbrace{V'(x_1^*)(x_1 - x_1^*)}_{=0} - \frac{1}{2} \underbrace{V''(x_1^*)(x_1 - x_1^*)^2}_{<0} + \mathcal{O}((x_1 - x_1^*)^3)} \quad (5.6b)$$

$$= \pm \sqrt{\frac{2}{m}} \sqrt{-\frac{V''(x_1^*)}{2}} |x_1 - x_1^*| + \text{higher order terms.} \quad (5.6c)$$

The behavior is sketched in Fig. 5.2

- (v) A local minimum has local maxima to the left and right, unless V is monotonously nondecreasing to infinity. Then we have two cases to differentiate. One where the local maxima are at the same level, and the other where one local maximum has a smaller potential than the other (or in one direction V monotonously increases towards infinity). *Homoclinic orbits* are special periodic orbits that are formed on a level set that contains a single saddle point. *Heteroclinic orbits* are formed when the level set contains two saddle points, i.e. when the two maxima have the same potential. These two cases are illustrated in Fig. 5.3.

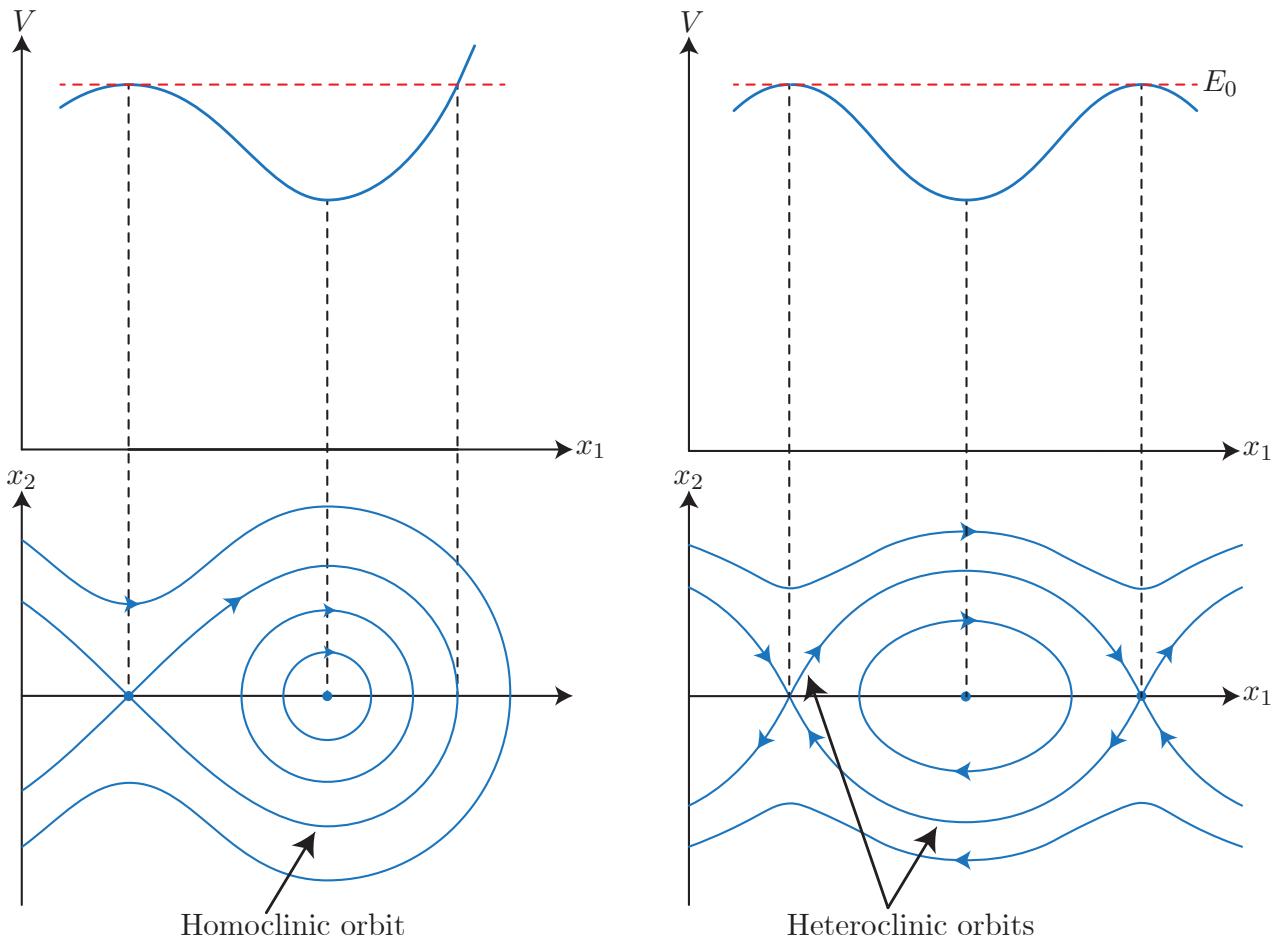


Figure 5.3: Left: One maximum has a lower potential than the other (or V is nondecreasing towards infinity on the right). The black arrow denotes a homoclinic orbit. Right: All maxima have the same potential. The black arrows denote heteroclinic orbits.

These consequences are illuminated in a basic example.

Example 5.1 (The Duffing oscillator). Let the equation of motion be given by

$$\ddot{x} - x + x^3 = 0; \quad x_1 = x; \quad x_2 = \dot{x}. \quad (5.7)$$

Therefore we get the first order ODE

$$\begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = x_1 - x_1^3 = -V'(x_1) \end{cases} \implies E(x) = \frac{1}{2}x_2^2 + V(x_1) = \frac{1}{2}x_2^2 + \left(-\frac{1}{2}x_1^2 + \frac{1}{4}x_1^4 \right). \quad (5.8)$$

The full phase portrait with two homoclinic orbits is shown in Fig. 5.4. We find two stable fixed points at the potential minima and they are separated by a potential maximum in the middle. Periodic orbits surround the stable fixed points. These orbits are separated from the rest of the phase space by the two homoclinic orbits of the saddle. Outside, we have periodic orbits that correspond to energy level-sets higher than the potential of the saddle.

Notebook 5.1. Computing the phase-portrait of a 1-DOF mechanical system.

5.2 Global behavior in two dimensional autonomous dynamical systems

Consider the following dynamical system

$$\dot{x} = f(x); \quad x \in \mathbb{R}^2; \quad f \in \mathcal{C}^1. \quad (5.9)$$

Further assume that solutions exist for all times, therefore there exists a unique solution $x(t; x_0)$ for all $t \in \mathbb{R}$ such that $x(0; x_0) = x_0$. Next, a few definitions are introduced enabling a deeper exploration of such systems.

Definition 5.1. A point $p \in \mathbb{R}^2$ is an ω -limit point of x_0 if there exists a monotone increasing unbounded sequence of times $\{t_i\}_{i=1}^\infty$ with $t_1 \geq 0$ such that

$$\lim_{i \rightarrow \infty} x(t_i; x_0) = p. \quad (5.10)$$

Definition 5.2. A point $q \in \mathbb{R}^2$ is an α -limit point of x_0 if it is an ω -limit point in backward time.

Definition 5.3. The ω -limit set of x_0 , denoted $\omega(x_0)$ is the set of all ω -limit points of x_0 . The α -limit set of x_0 , denoted $\alpha(x_0)$ is the set of all α -limit points of x_0 .

Remark 5.1. Note that $\omega(x_0)$ and $\alpha(x_0)$ is the same for all x_0 along a given trajectory, thus limit sets can be associated with full trajectories.

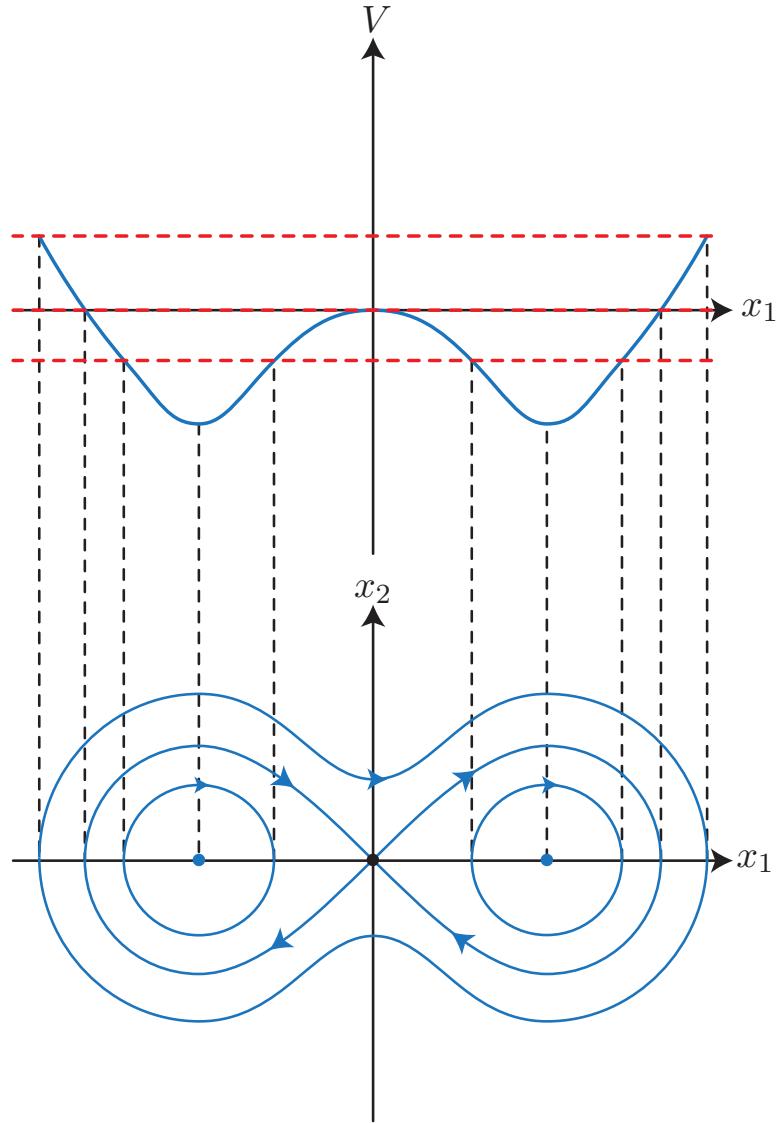


Figure 5.4: Phase portrait of the Duffing oscillator using the potential V .

Example 5.2 (Examples of limit sets). Below (Fig. 5.5) three different limit sets are depicted. For the leftmost dynamical system we have $\omega(p) = \alpha(p) = \{p\}$ and

$$\omega(x_0) = \emptyset, \quad \omega(\tilde{x}_0) = \{p\}, \quad \omega(\hat{x}_0) = \emptyset, \quad (5.11a)$$

$$\alpha(x_0) = \emptyset, \quad \alpha(\tilde{x}_0) = \emptyset, \quad \alpha(\hat{x}_0) = \{p\}. \quad (5.11b)$$

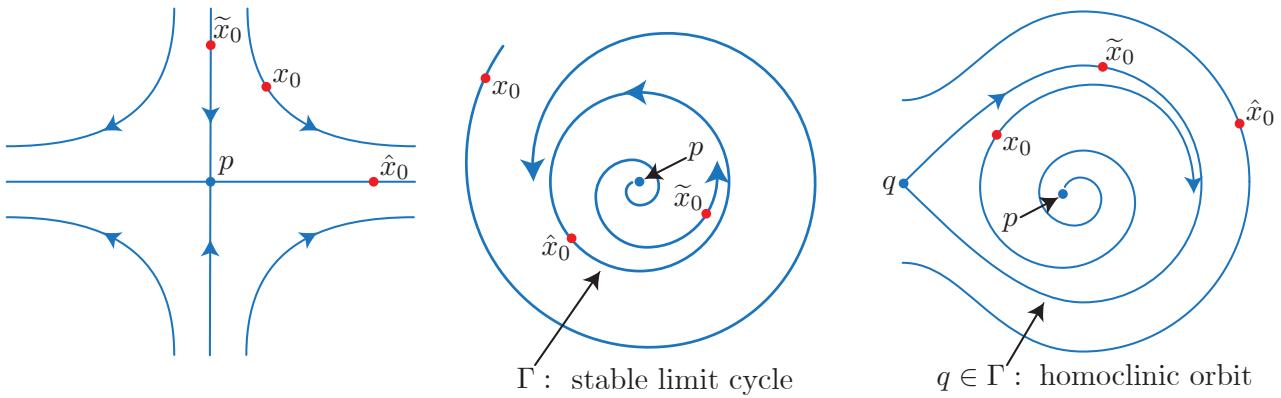


Figure 5.5: Three example systems for which we examine the respective limit sets. The arrows labelled Γ denote a stable limit cycle (middle) and a homoclinic orbit (right).

For the middle dynamical system we have

$$\omega(x_0) = \Gamma, \quad \omega(\tilde{x}_0) = \Gamma, \quad \omega(\hat{x}_0) = \Gamma, \quad (5.12a)$$

$$\alpha(x_0) = \text{unclear}, \quad \alpha(\tilde{x}_0) = \{p\}, \quad \alpha(\hat{x}_0) = \Gamma. \quad (5.12b)$$

For the rightmost dynamical system we have

$$\omega(x_0) = \Gamma, \quad \omega(\tilde{x}_0) = \{q\}, \quad \omega(\hat{x}_0) = \text{unclear}, \quad (5.13a)$$

$$\alpha(x_0) = \{p\}, \quad \alpha(\tilde{x}_0) = \{q\}, \quad \alpha(\hat{x}_0) = \text{unclear}. \quad (5.13b)$$

We now introduce two theorems which give us more insight into these limit sets.

Theorem 5.2. *If $x(t; x_0)$ is bounded, then $\omega(x_0)$ and $\alpha(x_0)$ are nonempty, closed, and connected. Further, the omega and alpha limit sets are invariant sets, i.e. they consist of full trajectories.*

Theorem 5.3 (Poincaré-Bendixson). *If $x(t; x_0)$ is bounded, then $\omega(x_0)$ and $\alpha(x_0)$ must be one of the following*

- (i) *A connected set of fixed points,*
- (ii) *A limit cycle,*
- (iii) *A set of fixed points and their homo-/heteroclinic orbits (illustrated in Fig. 5.6).*

There is no complex (chaotic) limiting behavior in 2-dimensional autonomous nonlinear systems.

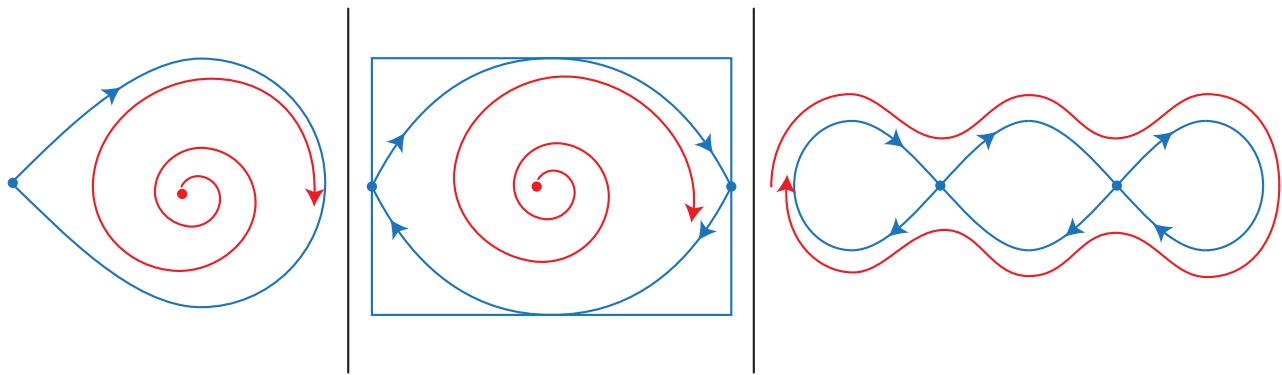


Figure 5.6: Examples of a set of fixed points connected by homo-/heteroclinic orbits.

Remark 5.4. The homo-/heteroclinic orbits are generally not robust, thus we should not expect to see them in a typical dynamical system. We can imagine that under small perturbations the continuous line of the orbit breaks up into separate unstable and stable manifolds as shown in Fig. 5.7. However, these orbits are robust within the class of conservative systems, and

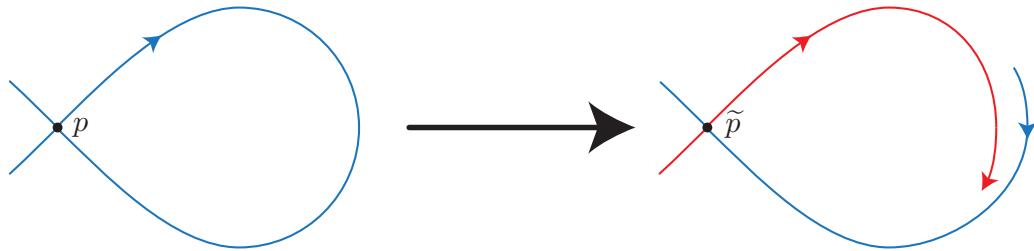


Figure 5.7: The loss of the homoclinic orbit under a small perturbation. Before the perturbation $\Gamma = W^S(p) = W^U(p)$ (p is the hyperbolic fixed point), meanwhile after the perturbation $W^S(p) \neq W^U(p)$. The unstable manifold is given in red, and the stable manifold in blue.

conservative perturbations do not lead to global bifurcations as seen in Fig. 5.8.

Remark 5.5. A consequence of the Poincaré-Bendixon Theorem is that a forward-invariant, bounded open set without fixed points must contain a limit cycle. This is due to the elimination of possibilities (i) and (iii) in the theorem. An example of such a set is shown in Fig. 5.9.

Exercise 5.1. Consider a planar dynamical system with the following phase portrait (Fig. 5.1).

Which of the following statement is true?

- (i) The ω -limit set of γ is empty.

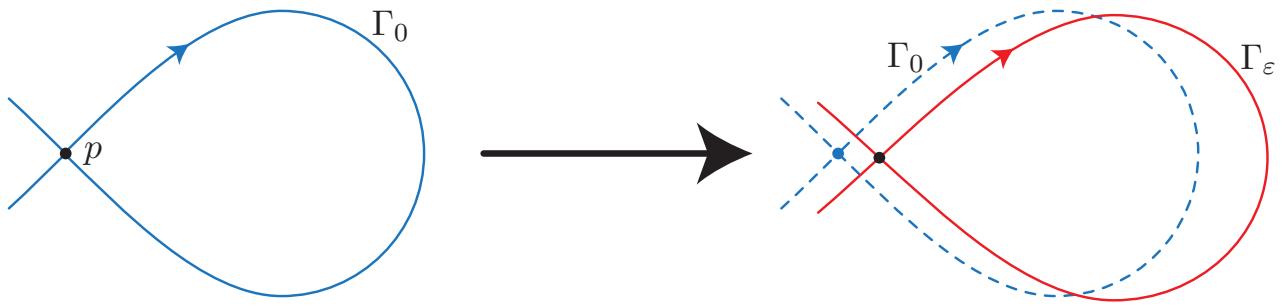


Figure 5.8: Small conservative perturbation does not lead to a global bifurcation in conservative systems. The energy level containing Γ_0 is robust, and perturbs to a new Γ_ε .

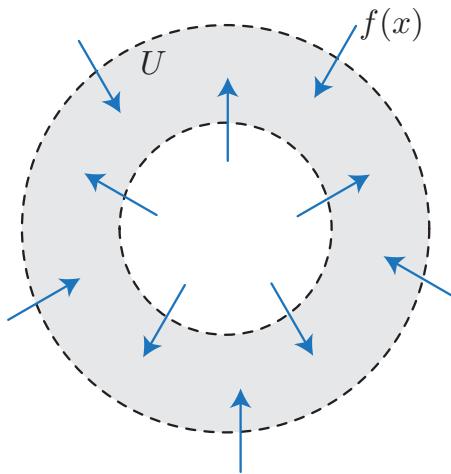


Figure 5.9: A forward-invariant, bounded open set without fixed points as in the remark. Such a set must contain a limit cycle.

- (ii) By the Poincaré-Bendixson theorem, the ω -limit set of γ is composed of the lines $y = 1$ and $y = -1$.
- (iii) The Poincaré-Bendixson theorem does not apply to γ .
- (iv) None of the above

Proposition 5.6 (Bendixson criterion). *Consider the following*

$$\dot{x} = f(x); \quad x \in \mathbb{R}^2; f \in \mathcal{C}^1; \quad U \text{ simply connected.} \quad (5.14)$$

And further assume

$$x \in U \implies \nabla \cdot f(x) \neq 0. \quad (5.15)$$

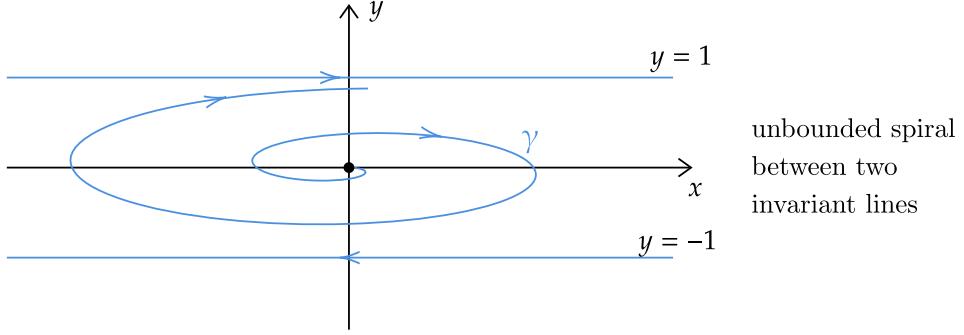


Figure 5.10: Phase portrait of the planar dynamical system

This in turn implies $\nabla \cdot f(x)$ has constant sign on U . Then there cannot exist a limit cycle in U .

Proof. Assume, towards a contradiction, that there exists a closed orbit Γ of period τ . Calculate

$$f^\perp(x) = \begin{pmatrix} f_2(x) \\ -f_1(x) \end{pmatrix} \implies \oint_{\Gamma} f^\perp \cdot \underbrace{dr}_{=\dot{r}dt} = \int_0^\tau f^\perp(r(t)) \cdot f(r(t)) dt = 0. \quad (5.16)$$

On the other hand we have by Green's Theorem

$$\oint_{\Gamma} f^\perp \cdot dr = \iint_{\Gamma^\circ} (\nabla \times f^\perp)_z dA = - \iint_{\Gamma^\circ} \underbrace{\nabla \cdot f}_{\neq 0} dA \neq 0. \quad (5.17)$$

Here, Γ° denotes the interior of the area enclosed by Γ . These two equations stand in contradiction to each other, therefore no limit cycle can exist in U . \square

Remark 5.7. Another consequence of the Poincaré-Bendixson Theorem is that purely damped or purely forced perturbations of a conservative system cannot have limit cycles.

$$f(x) = \begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -V'(x_1) + g(x_1, x_2) \end{cases}; \quad \nabla \cdot f = \nabla \cdot \begin{pmatrix} 0 \\ g(x) \end{pmatrix} = \frac{\partial g}{\partial x_2}. \quad (5.18)$$

If we assume pure damping or forcing, the divergence of f cannot be 0 on any open set. If it is positive, we have pure forcing, alternatively if it is negative we have pure damping. In either case no limit cycle can exist.

Exercise 5.2. Consider a planar Hamiltonian system

$$\begin{cases} \dot{x} = \frac{\partial H(x, y)}{\partial y} + f_1(x, y) \\ \dot{y} = -\frac{\partial H(x, y)}{\partial x} + f_2(x, y), \end{cases} \quad (5.19)$$

where the twice continuously differentiable function $H(x, y)$ is the Hamiltonian associated with the system (say, the energy in classical mechanics, or the stream function in fluid mechanics), and the continuously differentiable $f = (f_1, f_2)$ is a dissipative term (say, damping in classical mechanics, or compressible terms in fluid mechanics). Assume that $\nabla \cdot f \neq 0$ for all $(x, y) \in \mathbb{R}^2$ (linear damping, for instance has this property.)

Show that the above system can have no limit cycles.

Chapter 6

Time-dependent dynamical systems

Our setup for this chapter will be of nonautonomous n -dimensional systems, i.e. the following

$$\dot{x} = f(x, t); \quad x \in \mathbb{R}^n; \quad t \in \mathbb{R}. \quad (6.1)$$

Note that by adding a dimension to our phase space, we can make the system autonomous

$$X = \begin{pmatrix} x \\ t \end{pmatrix} \in \mathbb{R}^{n+1}; \quad F(X) = \begin{pmatrix} f \\ 1 \end{pmatrix} \implies \dot{X} = F(X). \quad (6.2)$$

6.1 Nonautonomous linear systems

For this section, consider

$$\dot{x} = A(t)x + b(t); \quad x \in \mathbb{R}^n; \quad A(t) \in \mathbb{R}^{n \times n}; \quad b(t) \in \mathbb{R}^n. \quad (6.3)$$

We separate (6.3) into two parts. The homogeneous part is

$$\dot{x} = A(t)x. \quad (6.4)$$

For the homogeneous part we have the solution $x(t) = \Phi_{t_0}^t x_0$, with the fundamental matrix solution Φ with the property $\Phi_{t_0}^{t_0} = I \in \mathbb{R}^{n \times n}$. The general solution of (6.3) is the sum of the general solution of (6.4) and a particular solution of (6.3). We can find the particular solution by, for example, the variation of constants formula. We refer to [Arnold, 1992] for the derivation. Thus we find the general solution to (6.3) to be

$$x(t) = \Phi_{t_0}^t x_0 + \Phi_{t_0}^t \int_{t_0}^t (\Phi_{t_0}^s)^{-1} b(s) ds.$$

(6.5)

Note that due to the group property of the flow map, the solution can be equivalently written as

$$x(t) = \Phi_{t_0}^t x_0 + \int_{t_0}^t \Phi_s^t b(s) ds. \quad (6.6)$$

From here we can see that it is enough to understand the homogeneous part. The stability of $x = 0$ in (6.4) is equivalent to the stability of $x_p(t) = \Phi_{t_0}^t \int_{t_0}^t (\Phi_{t_0}^s)^{-1} b(s) ds$ in (6.3). Unfortunately the general solution to (6.4) is unknown.

6.2 Time-periodic nonlinear systems

We now specify our system to be T -periodic, thus the evolution rule defined by $\dot{x} = A(t)x$ repeats itself with period T , i.e.

$$A(t+T) = A(t). \quad (6.7)$$

This implies

$$\Phi_{t_0}^{t_0+T} = \Phi_{t_0+T}^{t_0+2T} = \Phi_{t_0+2T}^{t_0+3T} = \dots =: P_{t_0}, \quad (6.8)$$

The operator P_{t_0} is called the *Poincaré map* (or time T map) based at t_0 . Thus we have

$$x_0(t_0 + nT) = \underbrace{P_{t_0} \dots P_{t_0}}_{n \text{ times}} = P_{t_0}^n x_0. \quad (6.9)$$

The n -th iterate of the Poincaré map is denoted by $P_{t_0}^n$.

Time being a periodic variable allows us to study the geometry in the extended phase space $\mathbb{R}^n \times S^1$ as depicted in Fig. 6.1. Importantly, the stability of $x = 0$ can be studied as the stability of the $x = 0$ fixed point of the map P_{t_0} . Label the eigenvalues of P_{t_0} $\rho_1, \dots, \rho_n \in \mathbb{C}$ these are called *Floquet multipliers*. Since this is now a discrete dynamical system, we know how to assess the stability type of its fixed point. We have two cases

- (i) If there exists an eigenvalue with $|\rho_i| > 1$ for P_{t_0} or there exists a repeated eigenvalue with $|\rho_j| = 1$ for which the algebraic multiplicity is greater than the geometric multiplicity ($a > g$), then $x = 0$ is unstable for (6.4);
- (ii) Otherwise, if each eigenvalue of P_{t_0} has magnitude strictly less than 1, i.e. $|\rho_j| < 1$, then $x = 0$ is asymptotically stable for (6.4).

Floquet theory tells us that the following holds in general

$\Phi_{t_0}^t = B(t)e^{\Lambda(t-t_0)}; \quad B(t+T) = B(t) \in \mathbb{R}^{n \times n}; \quad \Lambda \in \mathbb{R}^{n \times n} \text{ is constant.}$

(6.10)

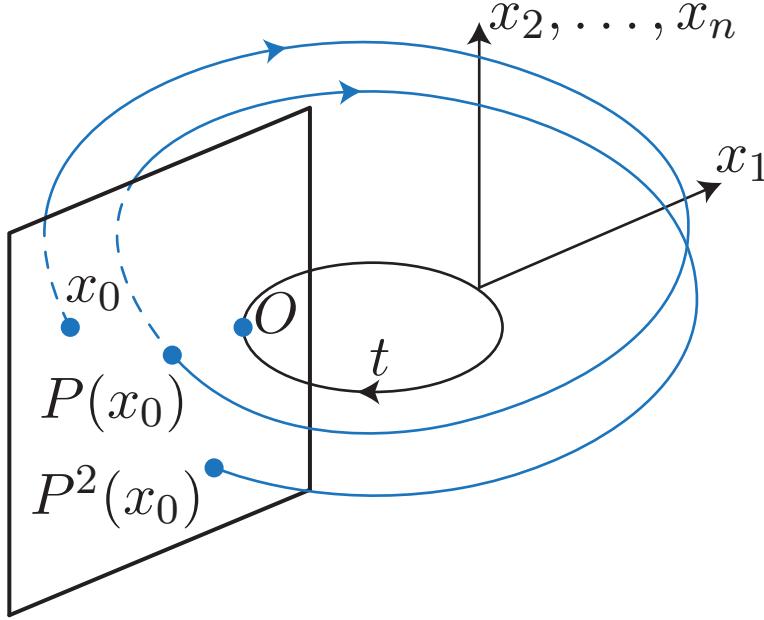


Figure 6.1: The map P records snapshots of the trajectory $x(t)$ taken at times nT . These snapshots are depicted along a single plane in the extended phase space. Due to periodicity, the extended phase space is made up as the product of the plane and the circle S^1 . The snapshots shown are $Px_0 = x_1$ and $P^2x_0 = x_2$.

Definition 6.1 (Floquet exponents). The eigenvalues of Λ are $\lambda_1, \dots, \lambda_n \in \mathbb{C}$ and they are called *Floquet exponents*. The previously defined Floquet multipliers are related to λ_j as

$$\rho_j = e^{\lambda_j T} = e^{(\alpha_j + i\beta_j)T}. \quad (6.11)$$

Hence, λ_j is well-defined only up to addition of $i2k\pi/T$ for $k \in \mathbb{Z}$. Therefore we have

$$|\rho_j| \stackrel{<}{\leq} 1 \Leftrightarrow \operatorname{Re} \lambda_j \stackrel{<}{\leq} 0. \quad (6.12)$$

This is generally only numerically computable.

We now explore a sufficient criterion for instability of $x = 0$. For this, recall Liouville's Theorem (Abel's Theorem)

$$\det(\Phi_{t_0}^t) = \det(\Phi_{t_0}^{t_0}) e^{\int_{t_0}^t \operatorname{Tr}[A(s)]ds}. \quad (6.13)$$

This holds true for any dynamical system $\dot{x} = A(t)x$ with aperiodic time-dependence.

Proposition 6.1 (Sufficient criterion for instability). *If we have $\int_0^T \operatorname{Tr}[A(s)]ds > 0$ then $x = 0$ is unstable for $\dot{x} = A(t)x$ with $A(t) = A(t + T)$*

Proof. Apply Liouville's Theorem to our T -periodic system to find

$$\prod_{j=1}^n \rho_j = e^{\sum_{j=1}^n \lambda_j T} = \det(P_{t_0}) = \det(\Phi_{t_0}^{t_0+T}) = e^{\int_{t_0}^{t_0+T} \text{Tr}[A(s)] ds}. \quad (6.14)$$

Now we can compare the arguments of the exponential functions to derive

$$\boxed{\sum_{j=1}^n \lambda_j = \frac{1}{T} \int_0^T \text{Tr}[A(s)] ds.} \quad (6.15)$$

Therefore our claim holds due to the case differentiation for stability above. Thus we can verify the stability without the numerical solution of (6.4). \square

Remark 6.2. The change of variables $x = P(t)y$, for some matrix $P(t)$ transforms the system (6.4) into the form $\dot{y} = By$, an autonomous linear ODE.

Carrying out the change of coordinates, we obtain that

$$\dot{y} = P^{-1}(A(t)P - \dot{P})y. \quad (6.16)$$

But

$$\frac{d}{dt} (P(t)e^{Bt}) = APe^{Bt} \implies \dot{P} + PB = AP. \quad (6.17)$$

Together this implies $\dot{y} = By$.

6.3 Averaging

We will now attempt to study the mean evolution of oscillatory systems

$$\dot{x} = f(x, t); \quad x \in \mathbb{R}^n; \quad f(x, t+T) = f(x, t). \quad (6.18)$$

In general the averaged system

$$\dot{y} = \frac{1}{T} \int_0^T f(y, t) dt \quad (6.19)$$

does not capture the overall (mean) behavior of the system. This can be demonstrated in an example.

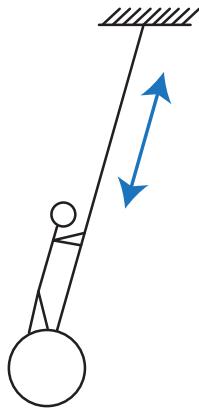


Figure 6.2: A swing (pendulum) with varying length and a mass m .

Example 6.1 (Averaging for a swing). Mechanically, we examine the stability of a pendulum with a periodically varying length as shown in Fig. 6.2. The first order model for this system is

$$\begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -\omega_0^2(1 + a(t))x_1 \end{cases}; \quad a(t) = a(t + T), \quad 0 < |a| \ll 1. \quad (6.20)$$

Averaging these equations yields

$$\begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -\omega_0^2(1 + \bar{a})x_1. \end{cases} \quad (6.21)$$

Where \bar{a} is the average of $a(t)$, a constant. The phase portrait of this is known, as it is just a standard pendulum, i.e. $x = 0$ is stable. This means that the small forcing, when averaged, shouldn't move us from the stable position at $x = 0$. But we know that by periodically swinging, we don't just stay stuck at the equilibrium point of the swing, but instead are able to reach larger amplitude oscillations. Therefore, averaging clearly does not work in this situation.

However, there are situations in which averaging does work.

Example 6.2 (Averaging in slowly varying periodic systems). Slowly varying periodic systems are given by

$$\dot{x} = \varepsilon f(x, t, \varepsilon); \quad x \in \mathbb{R}^n; \quad 0 \leq \varepsilon \ll 1; \quad f(x, t, \varepsilon) = f(x, t + T, \varepsilon). \quad (6.22)$$

Such systems are also called *adiabatic* systems. The averaged system in this case is

$$\dot{y} = \varepsilon \frac{1}{T} \int_0^T f(y, t, 0) dt = \varepsilon \bar{f}_0(y). \quad (6.23)$$

Here only the leading order terms with respect to ε in the integrand were taken, as $\varepsilon \ll 0$. Now we transform our original equation into the averaged equation via a near-identity, periodic change of variables

$$x = y + \varepsilon w(y, t); \quad w(y, t) = w(y, t + T). \quad (6.24)$$

Plugging this transformation into our differential equation, we find

$$\dot{x} = [I + \varepsilon D_y w] \dot{y} + \varepsilon \frac{\partial w}{\partial t} \quad (6.25a)$$

$$\dot{y} = \underbrace{[I + \varepsilon D_y w]^{-1}}_{I - \varepsilon D_y w + \mathcal{O}(\varepsilon^2)} \left[\underbrace{\varepsilon f(y + \varepsilon w(y, t), t, \varepsilon)}_{= \dot{x}} - \varepsilon \frac{\partial w}{\partial t} \right] \quad (6.25b)$$

$$= \varepsilon \left[f(y, t, 0) - \frac{\partial w}{\partial t} \right] + \mathcal{O}(\varepsilon)^2. \quad (6.25c)$$

In the last equation we used the Taylor expansion. We cannot set $w(y, t) = \int_0^t f(y, s, 0) ds$ as that would make ω aperiodic. Instead we separate f into two parts, the average and the deviation, $f(y, t, 0) = \bar{f}_0(y) + \tilde{f}(y, t)$. Note that the deviation has mean 0, therefore we set $w(y, t) = \int_0^t \tilde{f}(y, s) ds$ which is periodic. Therefore we find $\dot{y} = \varepsilon \bar{f}_0(y) + \mathcal{O}(\varepsilon^2)$, noting here that the order 2 term is still T periodic.

This has given us that the original system is a T -periodic, $\mathcal{O}(\varepsilon^2)$ perturbation of the averaged system in the y coordinates.

Averaging Principle Solutions of the original equation starting $\mathcal{O}(\varepsilon)$ close to those of the averaged system, stay $\mathcal{O}(\varepsilon)$ close for times of $\mathcal{O}(\frac{1}{\varepsilon})$ (very long).

The averaging principle is depicted in Fig. 6.3.

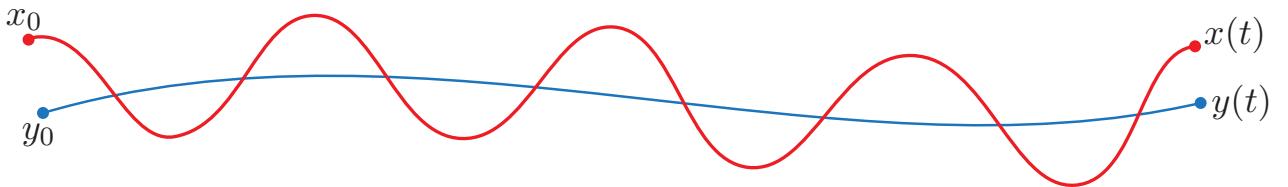


Figure 6.3: For the starting point x_0 which is $\mathcal{O}(\varepsilon)$ close to y_0 , the trajectories remain order ε close, i.e. $x(t)$ is $\mathcal{O}(\varepsilon)$ close to $y(t)$, for $t \ll \frac{K}{\varepsilon}$.

Exercise 6.1. Show that on time scales of $\mathcal{O}(1/\varepsilon)$, a solution $x(t)$ with $x(0) = x_0$ of the dynamical system

$$\dot{x} = \varepsilon f(x, t, \varepsilon); \quad x \in \mathbb{R}^n, \quad (6.26)$$

(ε is a small parameter and f is a smooth function that is T -periodic in time) remains $\mathcal{O}(\varepsilon)$ -close to any solution $y(t)$ with $y(0) = x_0 + \mathcal{O}(\varepsilon)$ of the averaged system

$$\dot{y} = \varepsilon \bar{f}_0(y); \quad y \in \mathbb{R}^n, \quad (6.27)$$

where

$$\bar{f}_0(y) = \frac{1}{T} \int_0^T f(y, t, 0) dt. \quad (6.28)$$

Hint: Subtract (6.27) from (6.26) and integrate to obtain an expression for $|x(t) - y(t)|$. Estimate $|x(t) - y(t)|$ from above using the facts that \bar{f} is Lipschitz and $|\hat{f} - \bar{f}|/\varepsilon$ is uniformly bounded, where \hat{f} is the right-hand-side of the system into which (6.26) is transformed by the averaging transformation $x = y + \varepsilon w(y, t)$. Then use the following generalized Gronwall inequality:

If $u(t), v(t), c(t)$ are non-negative functions, $c(t)$ is differentiable, and

$$v(t) \leq c(t) + \int_0^t u(s)v(s) ds,$$

then

$$v(t) \leq c(0)e^{\int_0^t u(s) ds} + \int_0^t c'(s)e^{\int_s^t u(\tau) d\tau} ds.$$

This has implications for the Poincaré map of the original equation

$$P_{t_0}^\varepsilon : x_0 \mapsto x(t_0 + T, x_0; \varepsilon); \quad P_{t_0}^\varepsilon = P_{t_0}^0 + \mathcal{O}(\varepsilon) \implies (P_{t_0}^\varepsilon)^n = (P_{t_0}^0)^n + \mathcal{O}(\varepsilon). \quad (6.29)$$

This only holds if $n \leq \frac{1}{\varepsilon}$. The time T map of the averaged system was denoted as $P_{t_0}^0$.

Example 6.3 (Averaged system with a saddle-type fixed point). Consider a system whose Poincaré map P_0 has a hyperbolic fixed point p_0 . By the persistence of hyperbolic fixed points, the perturbed Poincaré map P_ε also has a hyperbolic fixed point p_ε which is $\mathcal{O}(\varepsilon)$ close to p_0 . Furthermore, the \mathcal{C}^1 manifolds $W_{\text{loc}}^U(p_\varepsilon)$ and $W_{\text{loc}}^S(p_\varepsilon)$ are $\mathcal{O}(\varepsilon)$ close to their unperturbed counterparts. The unperturbed system and geometry of the perturbed system are shown in Fig. 6.4.

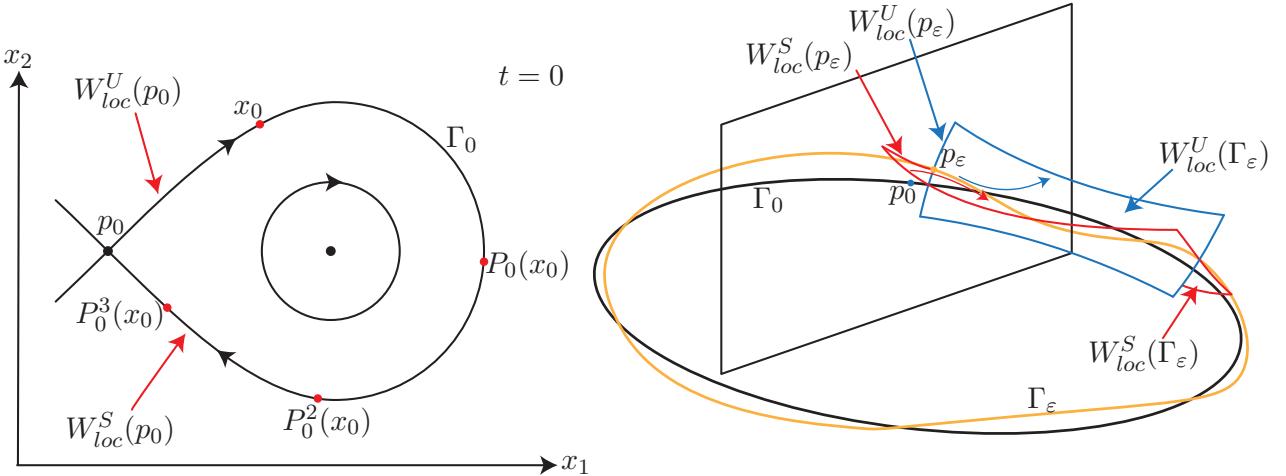


Figure 6.4: Left: The averaged system with the saddle type fixed point p_0 . The upper arrow designates the local unstable manifold, and the lower arrow the local stable manifold. Right: The same in the extended phase space, time is represented by S_1 and runs clockwise. The hyperbolic limit cycle Γ_ε is given by the yellow trajectory. Furthermore, the non-averaged local stable and unstable manifolds are drawn in red and blue respectively.

Remark 6.3. We can show that fixed points persist for P_ε as follows. First we take the flow map for the system $F_{t_0}^t(y_0; \varepsilon)$, then we use the Taylor expansion to find

$$F_{t_0}^{t_0+T}(y_0; \varepsilon) = F_{t_0}^{t_0+T}(y_0; 0) + \varepsilon \frac{\partial}{\partial \varepsilon} F_{t_0}^{t_0+T}(y_0; 0) + \mathcal{O}(\varepsilon^2) \quad (6.30a)$$

$$= y_0 + \varepsilon \left. \frac{\partial y(t_0 + T; t_0; \varepsilon)}{\partial \varepsilon} \right|_{\varepsilon=0} + \mathcal{O}(\varepsilon^2) \quad (6.30b)$$

$$= y_0 + \varepsilon \int_{t_0}^{t_0+T} \bar{f}(y_0) dt + \mathcal{O}(\varepsilon^2). \quad (6.30c)$$

In the last equality we used that

$$\left. \frac{\partial y(t_0 + T; t_0, \varepsilon)}{\partial \varepsilon} \right|_{\varepsilon=0} = \left. \frac{\partial}{\partial \varepsilon} \int_{t_0}^{t_0+T} \dot{y}(s; t_0, \varepsilon) ds \right|_{\varepsilon=0} \quad (6.31a)$$

$$= \int_{t_0}^{t_0+T} \left. \frac{\partial}{\partial \varepsilon} (\varepsilon \bar{f}(y(s)) + \mathcal{O}(\varepsilon^2)) \right|_{\varepsilon=0} ds \quad (6.31b)$$

$$= \int_{t_0}^{t_0+T} \bar{f}(y_0) ds. \quad (6.31c)$$

From this we have

$$P_\varepsilon(y_0) - y_0 \Leftrightarrow y_0 + \varepsilon T \bar{f}(y_0) + \mathcal{O}(\varepsilon^2) - y_0 = 0 \Leftrightarrow T \bar{f}(y_0) + \mathcal{O}(\varepsilon) = 0. \quad (6.32)$$

And now we can use the Implicit Function Theorem to conclude.

Example 6.4 (Averaged system with a stable focus). Consider the averaged system with a stable focus around a fixed point p_0 . The transition to the perturbed system is shown in Fig. 6.5.

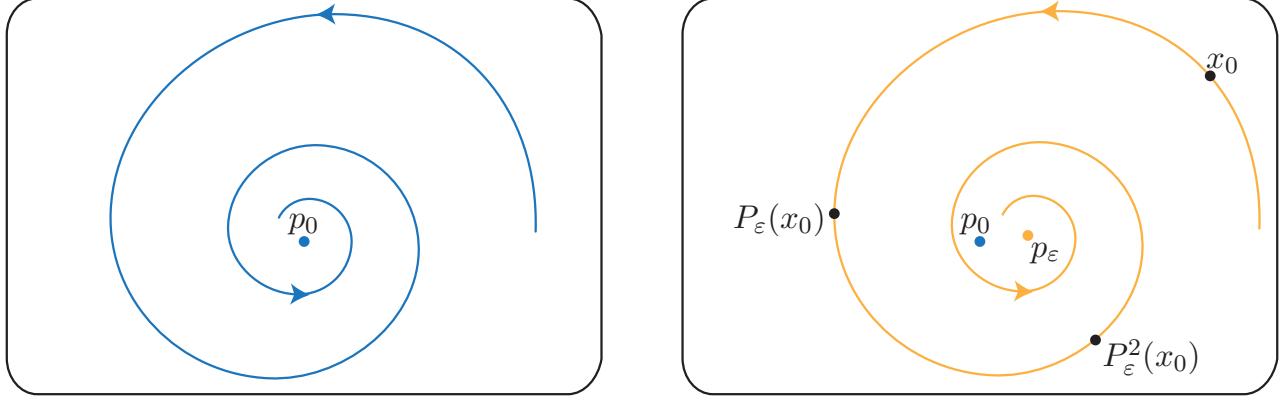


Figure 6.5: The stable focus around p_0 in the averaged system is preserved in the non-averaged system.

Example 6.5 (Averaged system with a limit cycle). Consider the averaged system with a limit cycle Γ_0 . The non-averaged system also has a stable limit cycle Γ_ε with $\dim W^S(\Gamma_\varepsilon) = \mathbb{R}^{n+1}$ in extended phase space. Thus in the full system we have an attracting 2-dimensional invariant torus. The averaged and non-averaged system, along with this torus are shown in Fig. 6.6.

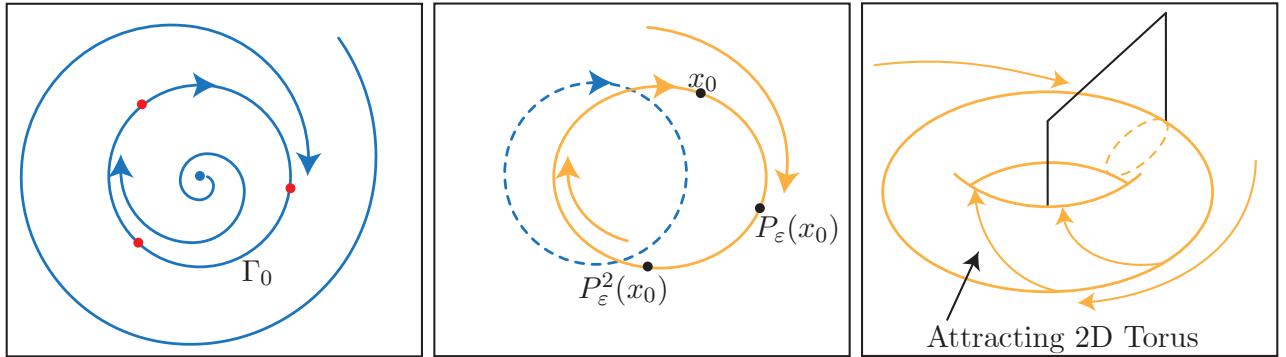


Figure 6.6: Left: The averaged system with the stable limit cycle Γ_0 . Middle: The non-averaged system with the limit cycle Γ_ε . Right: The attracting 2-dimensional invariant torus (black arrow), trajectories can be seen as the thin yellow lines.

Example 6.6 (Weakly nonlinear oscillations). An important application of averaging is for weakly nonlinear oscillations. This is covered in-depth in [Guckenheimer and Holmes, 1990].

The system is given by

$$\ddot{x} + \omega_0^2 x = \varepsilon f(x, \dot{x}, t) = \varepsilon f(x, \dot{x}, t + T); \quad T = 2\pi\omega. \quad (6.33)$$

Averaging applies after a change of coordinates moving with the solutions of the $\varepsilon = 0$ limit.

First transform the system to be a first-order ODE with $X = \begin{pmatrix} x \\ \dot{x} \end{pmatrix}$.

$$\dot{X} = AX + \varepsilon F(X, t); \quad A = \begin{pmatrix} 0 & 1 \\ -\omega_0^2 & 0 \end{pmatrix}; \quad F(X, t) = \begin{pmatrix} 0 \\ f(X_1, X_2, t) \end{pmatrix}. \quad (6.34)$$

We cannot directly apply averaging to (6.34). Instead we introduce a coordinate change along the trajectories of the system. Let $\Phi(t; \omega_0)$ be the fundamental matrix of solutions. In fact, this is explicitly computable for the harmonic oscillator as

$$\Phi(t; \omega_0) = \begin{pmatrix} \cos(\omega_0 t) & -\sin(\omega_0 t) \\ -\omega_0 \sin(\omega_0 t) & -\omega_0 \cos(\omega_0 t) \end{pmatrix}. \quad (6.35)$$

It can be checked that $\dot{\Phi} = A\Phi$. Thus we change coordinates with $X(t) = \Phi(t, \omega_0)Y(t)$. As Φ is invertible, we have $Y(t) = \Phi^{-1}(t, \omega_0)X(t)$ with the inverse

$$\Phi^{-1}(t, \omega_0) = \begin{pmatrix} \cos(\omega_0 t) & -\frac{1}{\omega_0} \sin(\omega_0 t) \\ -\sin(\omega_0 t) & -\frac{1}{\omega_0} \cos(\omega_0 t) \end{pmatrix}. \quad (6.36)$$

Now we substitute this transformation into the differential equation (6.34) and use the fact that $\dot{\Phi} = A\Phi$ to find

$$\dot{\Phi}Y + \Phi\dot{Y} = A\Phi Y + \varepsilon F(\Phi^{-1}Y, t) \implies \dot{Y} = \varepsilon\Phi^{-1}F(\Phi Y, t). \quad (6.37)$$

This appears to be in the required form for averaging, except the right hand side is not generally periodic, because Φ has frequency ω_0 , while the driving has frequency ω . It is only periodic if $\omega = k\omega_0$ for $k \in \mathbb{Z}$, i.e. resonance of order k . Often it is important to analyze near-resonant oscillations, i.e. $\omega \approx k\omega_0$. Thus our expectation is to find an almost sinusoidal response of frequency $\frac{\omega}{k}$. Therefore we instead use the transformation $\Phi(t; \frac{\omega}{k})$ and find

$$\dot{Y} = \Phi^{-1} \left[A\Phi - \dot{\Phi} \right] Y + \varepsilon\Phi^{-1}F(\Phi Y, t). \quad (6.38)$$

Now write $Y = \begin{pmatrix} u \\ v \end{pmatrix}$, and we find

$$\begin{cases} \dot{u} = -\frac{k}{\omega} \left[\left(\frac{\omega^2 - k^2\omega_0^2}{k^2} \right) x + \varepsilon f(x, \dot{x}, t) \right] \sin \left(\frac{\omega t}{k} \right) \\ \dot{v} = -\frac{k}{\omega} \left[\left(\frac{\omega^2 - k^2\omega_0^2}{k^2} \right) x + \varepsilon f(x, \dot{x}, t) \right] \cos \left(\frac{\omega t}{k} \right) \end{cases} \quad (6.39)$$

Now we have that if $\omega^2 - k^2\omega_0^2 = \mathcal{O}(\varepsilon)$, averaging applies to (6.39).

Next we consider a specific example, the Duffing oscillator

$$f(x, \dot{x}, t) = \beta \cos(\omega t) - \delta \dot{x} - \alpha x^3, \quad (6.40)$$

with $\omega_0^2 - \omega^2 = \varepsilon\Omega$, i.e. a nearly order one resonance. The transformed system then takes the form

$$\begin{cases} \dot{u} = \frac{\varepsilon}{\omega} [\Omega(u \cos(\omega t) - v \sin(\omega t)) - \omega\delta(u \sin(\omega t) + v \cos(\omega t)) \\ \quad + \alpha(u \cos(\omega t) - v \sin(\omega t))^3 - \beta \cos(\omega t)] \sin(\omega t) \\ \dot{v} = \frac{\varepsilon}{\omega} [\Omega(u \cos(\omega t) - v \sin(\omega t)) - \omega\delta(u \sin(\omega t) + v \cos(\omega t)) \\ \quad + \alpha(u \cos(\omega t) - v \sin(\omega t))^3 - \beta \cos(\omega t)] \cos(\omega t). \end{cases} \quad (6.41)$$

Averaging over one period $T = \frac{2\pi}{\omega}$ we find

$$\frac{1}{T} \int_0^T \sin(\omega t) \cos(\omega t) dt = \frac{1}{2T} \int_0^T \sin(2\omega t) dt = \left[-\frac{\cos(2\omega t)}{4\omega T} \right]_0^T = 0, \quad (6.42)$$

and

$$\frac{1}{T} \int_0^T \sin^2(\omega t) dt = \frac{1}{T} \int_0^T \frac{1 - \cos(2\omega t)}{2} dt = \frac{1}{T} \left[\frac{1}{2} - \frac{\sin(2\omega t)}{2\omega} \right]_0^T = \frac{1}{2}. \quad (6.43)$$

Using trigonometric identities then leads to the equation

$$\begin{cases} \dot{u} = \frac{\varepsilon}{2\omega} \left[-\omega\delta u - \Omega v - \frac{3\alpha}{4}(u^2 + v^2)v \right] \\ \dot{v} = \frac{\varepsilon}{2\omega} \left[\Omega u - \omega\delta v + \frac{3\alpha}{4}(u^2 + v^2)u - \beta \right]. \end{cases} \quad (6.44)$$

This can be further simplified by using the polar coordinates (r, ϕ) , i.e.

$$\begin{pmatrix} u \\ v \end{pmatrix} = \begin{pmatrix} r \cos(\phi) \\ r \sin(\phi) \end{pmatrix} \implies \begin{cases} \dot{r} = \frac{\varepsilon}{2\omega} [-\omega\delta r - \beta \sin(\phi)] \\ r\dot{\phi} = \frac{\varepsilon}{2\omega} \left[\Omega r + \frac{3\alpha}{4}r^3 - \beta \cos(\phi) \right]. \end{cases} \quad (6.45)$$

Next, recall that $x(t) = u(t) \cos(\omega t) - v(t) \sin(\omega t) = r(t) \cos(\omega t - \phi(t))$, i.e. we have slowly varying amplitude $r(t)$ and phase $\phi(t)$. The hyperbolic fixed points of (6.45) correspond to steady, almost sinusoidal response of the original system. We obtain the *forced response curves* by fixing α , β , and δ and plotting the fixed point $(\bar{r}, \bar{\phi})$ of (6.45) as a function of the frequency parameter Ω (or $\frac{\omega}{\omega_0}$) as in Fig. 6.7.

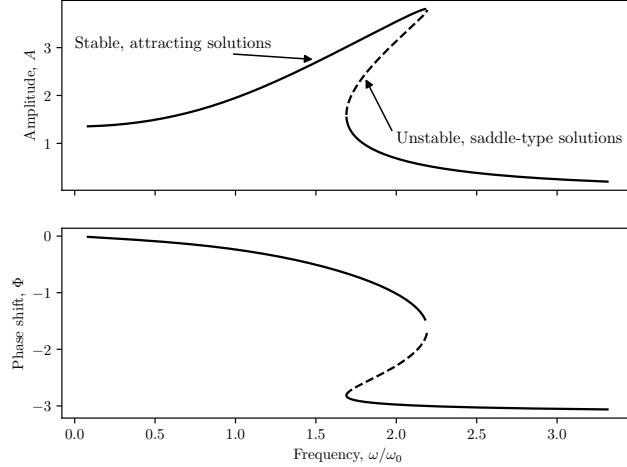


Figure 6.7: The forced response curve for the Duffing equation. The parameters are selected as $F_0 = 2, \omega_0 = 1, \alpha = 0.35, \zeta = 0.12$.

Exercise 6.2. *Fluid trajectories $x(t) = (x(t), y(t))$ in a two-dimensional time-periodic flow satisfy the differential equations*

$$\begin{cases} \dot{x} = u(x, y, t); & u(x, y, t) = u(x, y, t + T) \\ \dot{y} = v(x, y, t); & v(x, y, t) = v(x, y, t + T), \end{cases} \quad (6.46)$$

where $T > 0$ is the period, u and v are smooth velocity components satisfying the incompressibility condition $u_x + v_y \equiv 0$. Assume that the fluid is bounded by a wall at $y = 0$, on which the velocity field satisfies the no-slip boundary conditions $u(x, 0, t) = v(x, 0, t) = 0$. As a result, all boundary points are nonhyperbolic fixed points for (6.46).

We say that a boundary point $p_0 = (x_0, 0)$ is a separation point for the flow (6.46) if p_0 admits an unstable manifold $W^u(p_0)$. Physically, $W^u(p_0)$ is a time-dependent curve of fluid particles that shrinks to p_0 in backward time. In forward time, $W^u(p_0)$ attracts fluid particles, then ejects them from the vicinity of the wall (see Fig. 6.2). Show that separation points satisfy the criteria

$$\int_0^T u_y(x_0, 0, t) dt = 0, \quad \int_0^T v_{yy}(x_0, 0, t) dt > 0, \quad (6.47)$$

i.e., separation takes place where the average of the wall-shear is zero and the average of v_{yy} is positive.

Hint: Use incompressibility and the boundary conditions to show that (6.46) can be rewritten

as

$$\dot{x} = yU(x, y, t), \quad (6.48)$$

$$\dot{y} = y^2V(x, y, t). \quad (6.49)$$

To focus on the vicinity of the boundary, introduce the scaled variable $y = \varepsilon\eta$, where $0 \leq \varepsilon \ll 1$. Show that the resulting $(\dot{x}, \dot{\eta})$ equations are slowly varying and find the corresponding averaged equations. For the averaged equations, rescale time on trajectories by letting $\frac{d\tau}{dt} = \eta(t)$ in order to remove the common η factor from the right-hand-side, and look for a hyperbolic fixed point with an unstable manifold off the wall. Use the averaging theorem to relate your results to the original system (6.46).

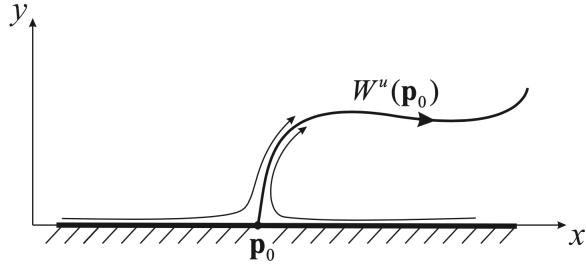


Figure 6.8: Unsteady separation from a no-slip wall.

6.4 The Harmonic Balance Method

Averaging and other perturbation techniques are often restrictive due to their small-parameter assumptions and result in cumbersome expressions limiting their application in high-dimensional systems. To avoid these constraints, the *Harmonic Balance* method computes the Fourier series approximation of the periodic response and is formally applicable without the previous restrictions. Consider the setup

$$\dot{x} = f(x, t) = f(x, t + T); \quad x \in \mathbb{R}^n; \quad f \in \mathcal{C}_x^r, \quad r \geq 1, \quad f \in \mathcal{C}_t^0. \quad (6.50)$$

Now we seek a T -periodic response via the Fourier series.

Remark 6.4. It is not immediately clear that a periodic response exists in a periodically forced multi-degree of freedom mechanical system. See for instance [Breunung and Haller, 2019] for more detail.

Assume (6.50) exhibits a T -periodic response, say $x_p(t) = x_p(t + T)$. Then it can be expressed as a convergent Fourier series

$$x_p(t) = \sum_{k \in \mathbb{Z}} \hat{x}_k e^{ik\omega t} = x_0 + \sum_{j \in \mathbb{N}} c_j \cos(j\omega t) + s_j \sin(j\omega t). \quad (6.51)$$

Substituting this into (6.50) and using that f is T -periodic (in both variables) and therefore admits a convergent Fourier series we find

$$\frac{d}{dt} x_p(t) = f(x_p(t), t) \quad (6.52a)$$

$$\frac{d}{dt} \sum_{k \in \mathbb{Z}} \hat{x}_k e^{ik\omega t} = \sum_{k \in \mathbb{Z}} \hat{f}_k(x_p(t), t) e^{ik\omega t} \quad (6.52b)$$

$$\sum_{k \in \mathbb{Z}} ik\omega \hat{x}_k e^{ik\omega t} = \sum_{k \in \mathbb{Z}} \hat{f}_k(x_p(t), t) e^{ik\omega t}. \quad (6.52c)$$

Now we can compare coefficients at different harmonics to find

$$ik\omega \hat{x}_k = \hat{f}_k(x_p(t), t) = \hat{f}_k \left(\sum_{j \in \mathbb{Z}} \hat{x}_j e^{ij\omega t}, t \right) \quad \forall k. \quad (6.53)$$

These are algebraic equations which give a solution for each coefficient. In general they are coupled, with a noteworthy exception if f is linear. However, there are infinitely many equations to solve, so in practice we only use a finite truncation of the Fourier series to approximate the response.

$$x_H = \sum_{k=-H}^H \hat{x}_k e^{ik\omega t}. \quad (6.54)$$

Now we instead solve $\dot{x}_H = f_H(x_H(t), t)$, leading to only $n(2H + 1)$ equations. Recall a few basic results from Fourier Analysis.

Theorem 6.5. *For any $F \in \mathcal{C}^0$ which is T -periodic, the truncation $F_H(t) = \sum_{k=-H}^H \hat{F}_k e^{ik\omega t}$ converges uniformly to $F(t)$ as $H \rightarrow \infty$.*

Theorem 6.6 (Decay of Fourier coefficients). *For $F \in \mathcal{C}^r$ which is T -periodic, the Fourier coefficients \hat{F}_k decay at least as fast as $\frac{1}{|k|^{r+1}}$. For an analytic function $F \in \mathcal{C}^a$, the coefficients decay faster than any power of k , i. e. exponentially.*

Remark 6.7. Note that the decay estimate is asymptotic, hence it is generally unclear exactly where to truncate. However, the method seems to work well for many practical applications and produces fast results.

We will now apply this method in a few examples.

Example 6.7 (Linear spring-mass-damper oscillator). Consider the following damped linear spring

$$\ddot{x} + 2\zeta\omega_0\dot{x} + \omega_0^2x = F(t) = F_0 \cos(\omega t). \quad (6.55)$$

As only a single harmonic is sufficient to describe F , the harmonic coefficients are decoupled. By taking the Fourier series of x and $F = \frac{1}{2}(F_0e^{i\omega t} + F_0e^{-i\omega t})$ we compare coefficients

$$k = 1 : -\omega^2\hat{x}_1e^{i\omega t} + i2\zeta\omega_0\omega\hat{x}_1e^{i\omega t} + \omega_0^2\hat{x}_1e^{i\omega t} = F_0e^{i\omega t}; \quad (6.56a)$$

$$k = -1 : -\omega^2\hat{x}_{-1}e^{-i\omega t} - i2\zeta\omega_0\omega\hat{x}_{-1}e^{-i\omega t} + \omega_0^2\hat{x}_{-1}e^{-i\omega t} = F_0e^{-i\omega t}. \quad (6.56b)$$

We therefore calculate

$$\hat{x}_1 = \frac{1}{2} \left(\frac{F_0}{\omega_0^2 - \omega^2 + i2\zeta\omega_0\omega} \right); \quad \hat{x}_{-1} = \frac{1}{2} \left(\frac{F_0}{\omega_0^2 - \omega^2 - i2\zeta\omega_0\omega} \right). \quad (6.57)$$

For all other values of k \hat{x}_k must equal zero. Now we can calculate the famous formula from linear vibration text books for the magnitude of the response

$$|x| = \frac{F_0}{\omega_0^2 \sqrt{\left(1 - \left(\frac{\omega}{\omega_0}\right)^2\right)^2 + \left(2\zeta\frac{\omega}{\omega_0}\right)^2}}. \quad (6.58)$$

Example 6.8 (Duffing Oscillator). Recall the Duffing oscillator

$$\ddot{x} + 2\zeta\omega_0\dot{x} + \omega_0^2x + \alpha x^3 = F_0 \cos(\omega t). \quad (6.59)$$

We consider the Ansatz

$$q_1(t) = c_1 \cos(\omega t) + s_1 \sin(\omega t) \implies \ddot{q}_1 + 2\zeta\omega_0\dot{q}_1 + \omega_0^2 q_1 + \alpha q_1^3 = F_0 \cos(\omega t). \quad (6.60)$$

Now calculate q_1^3 using multi-angle trigonometric identities

$$q_1^3 = (c_1 \cos(\omega t) + s_1 \sin(\omega t))^3 \quad (6.61a)$$

$$= c_1^3 \cos^3(\omega t) + 3c_1^2 s_1 \cos^2(\omega t) \sin(\omega t) + 3c_1 s_1^2 \cos(\omega t) \sin^2(\omega t) + s_1^3 \sin^3(\omega t) \quad (6.61b)$$

$$= \frac{1}{4} (3c_1^3 + 3c_1 s_1^2) \cos(\omega t) + \frac{1}{4} (c_1^3 - 3c_1 s_1^2) \cos(3\omega t) + \frac{1}{4} (3c_1^2 s_1 + 3s_1^3) \sin(\omega t) \quad (6.61c)$$

$$+ \frac{1}{4} (3c_1^2 s_1 - s_1^3) \sin(3\omega t). \quad (6.61d)$$

Now we neglect the harmonics higher than 1 and balance the remaining

$$0 = \left[(\omega_0^2 - \omega^2) c_1 + 2\zeta\omega_0\omega s_1 + \frac{3}{4}\alpha(c_1^3 + c_1 s_1^2) - F_0 \right] \cos(\omega t) \quad (6.62a)$$

$$+ \left[(\omega_0^2 - \omega^2) s_1 - 2\zeta\omega_0\omega c_1 + \frac{3}{4}\alpha(s_1^3 + c_1^2 s_1) \right] \sin(\omega t) + \text{higher order harmonics.} \quad (6.62b)$$

Therefore we have two algebraic equations in c_1 and s_1 . Using the polar coordinates $c_1 = A \cos(\phi)$ and $s_1 = A \sin(\phi)$ we find

$$(\omega_0^2 - \omega^2) A \cos(\phi) + 2\zeta\omega_0\omega A \sin(\phi) + \frac{3}{4}\alpha(A^3 \cos^3(\phi) + A^3 \cos(\phi) \sin^2(\phi)) = F_0 \quad (6.63a)$$

$$(\omega_0^2 - \omega^2) A \sin(\phi) - 2\zeta\omega_0\omega A \cos(\phi) + \frac{3}{4}\alpha(A^3 \sin^3(\phi) + A^3 \sin(\phi) \cos^2(\phi)) = 0. \quad (6.63b)$$

Adding $\cos(\phi)$ times the first equation to $\sin(\phi)$ times the second we obtain

$$(\omega_0^2 - \omega^2) A + \frac{3}{4}\alpha A^3 = F_0 \cos(\phi). \quad (6.64)$$

Similarly, adding $\sin(\phi)$ times the first equation to $-\cos(\phi)$ times the second we obtain

$$2\zeta\omega_0\omega A = F_0 \sin(\phi). \quad (6.65)$$

We now take the sum of the squares of these two new equations, which yields

$$A^2 \left[\left(\omega_0^2 - \omega^2 + \frac{3}{4}\alpha A^2 \right)^2 + 4\zeta^2\omega_0^2\omega^2 \right] = F_0^2. \quad (6.66)$$

This equation now allows us plot the amplitude A versus the frequency ω as in Fig. 6.9.

Unfortunately, in contrast to averaging, harmonic balance by itself does not return any stability information. We can, however, compute the Floquet multipliers of the periodic orbit by directly solving the equation of variations.

Example 6.9 (Unforced and undamped Duffing oscillator). For unforced systems harmonic balance can also be used to obtain internally parameterized periodic orbits. However, the expected period is a priori unknown. Consider the following

$$\ddot{x} + \omega_0^2 x + \alpha x^3 = 0. \quad (6.67)$$

Since the system is autonomous, the phase of the response is irrelevant. Thus the simplest Ansatz for a single harmonic response is

$$x_1(t) = A \cos(\omega t), \quad (6.68)$$

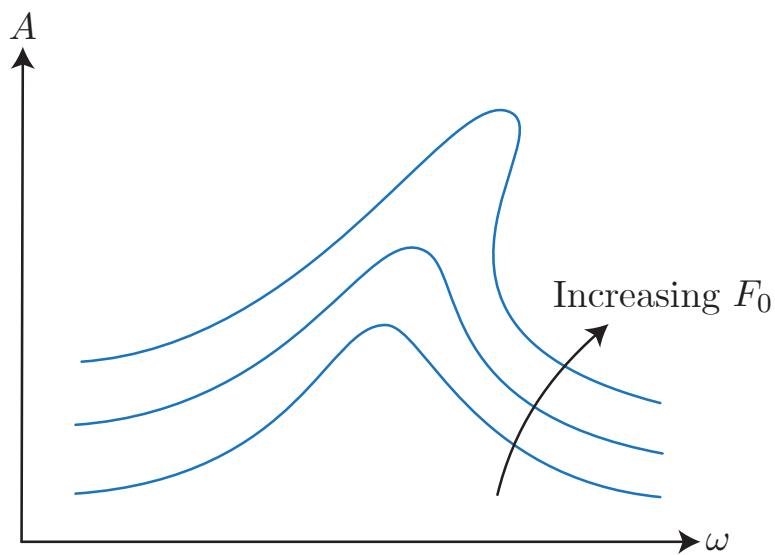


Figure 6.9: Forced response curves derived via harmonic balance for the Duffing oscillator.

for an unknown ω . Applying the procedure as before and neglecting the higher harmonics we find

$$\omega^2 = \omega_0 + \frac{3}{4}\alpha A^2. \quad (6.69)$$

For unforced, conservative, systems the frequency-amplitude relation gives the so-called *backbone curve*, illustrated in Fig. 6.10.

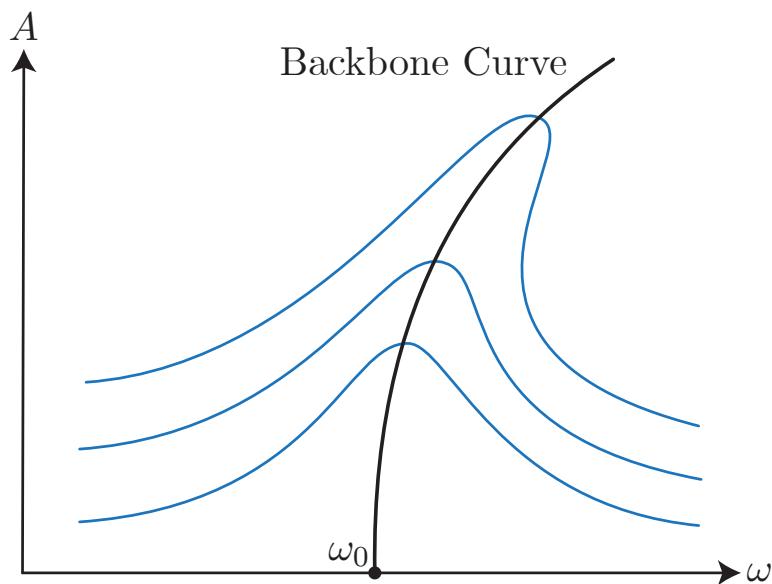


Figure 6.10: The backbone curve for the unforced and undamped Duffing oscillator. Under small forcing and damping, the backbone curve has been observed to connect the forced response curves' peaks.

Part II

Nonlinear Dynamics and Chaos 2

Chapter 7

Introduction to chaotic dynamics

We begin with a few motivating examples.

Example 7.1 (Periodically forced slender beam). A beam is hanging on the inside of a rectangular frame, attached to the upper edge. Two permanent magnets are attached to the lower edge. Furthermore, there is a $T = 2\pi$ -periodic forcing in the horizontal direction to the frame. The deflection of the beam is measured by the variable x . The setup is illustrated in Fig. 7.1. This leads to the equation of motion

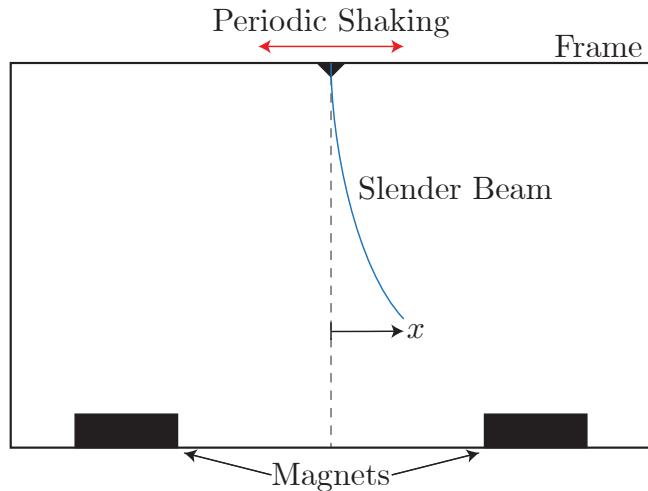


Figure 7.1: Depiction of the periodically forced slender beam. The deflection is given by x and the period is 2π . The two permanent magnets are represented by the black boxes.

$$\ddot{x} + \dot{x} - x + x^3 = \varepsilon \cos(t); \quad 0 \leq \varepsilon \ll 1. \quad (7.1)$$

Therefore we have a perturbed Duffing oscillator. We transform the system into a first order ODE

$$\begin{cases} \dot{x} = y \\ \dot{y} = -y + x - x^3 + \varepsilon \cos(t). \end{cases} \quad (7.2)$$

For $\varepsilon = 0$ driving and no damping, two homoclinic orbits arise for the three fixed points. However, for nonzero driving, we get seemingly chaotic behavior. Both of these regimes are depicted in Fig. 7.2.

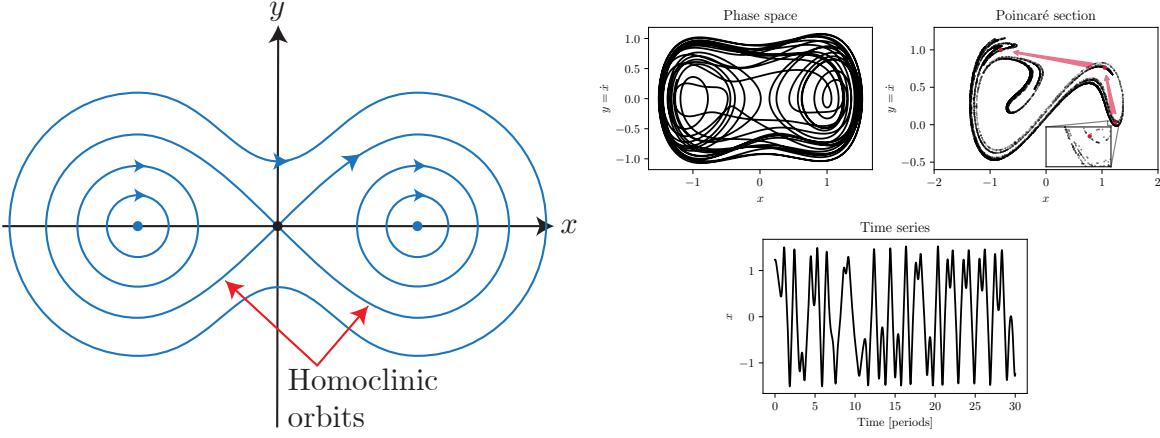


Figure 7.2: Left: Duffing oscillator with $\varepsilon = 0$ driving and no damping, the red arrows designate the homoclinic orbits. Right: (Clockwise from top left) The phase space of the driven Duffing oscillator; The Poincaré section of the driven Duffing oscillator, with the Poincaré map illustrated by the black arrows; The value of $x(t)$ over time, with no apparent pattern.

Example 7.2 (2-dimensional Rayleigh-Bénard convection). A fluid is held between two plates. The upper plate is cold and the lower plate is hot. This causes convective currents to start as hotter fluid rises and colder fluid falls. So called convection cells then form for low Rayleigh numbers. This process is illustrated in Fig. 7.3.

If the Rayleigh number R_a exceeds a critical value $R_{a_{\text{crit}}}$, we have a time periodic perturbation to the velocity field. The fluid trajectories have the following equations of motion

$$\begin{cases} \dot{x} = u(x, y) + \varepsilon u_1(x, y, t) \\ \dot{y} = v(x, y) + \varepsilon v_1(x, y, t) \end{cases}; \quad \varepsilon > 0. \quad (7.3)$$

Here, the functions u_1 and v_1 are T -periodic. The chaotic nature of this system is shown in Fig. 7.4.

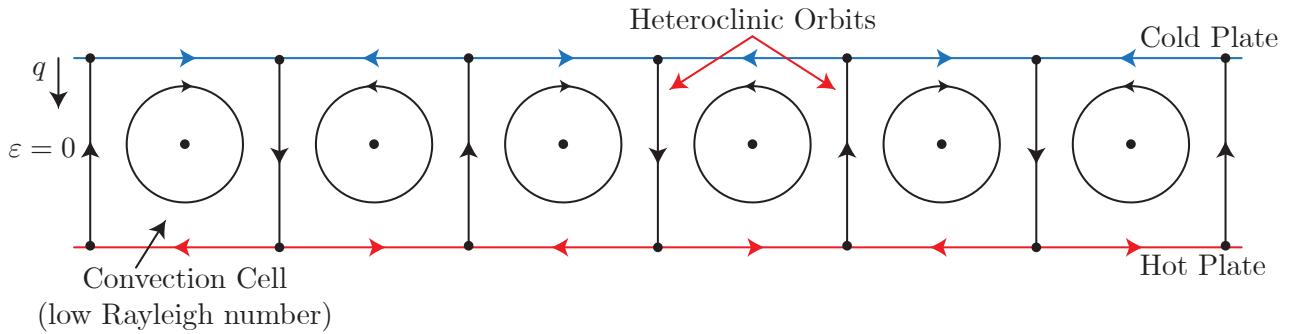


Figure 7.3: Convection cells forming between two plates of different temperature. The cold plate (blue) is vertically above the hot plate (red). Each black dot represents a fixed point of the system, with heteroclinic orbits connecting them (red arrows). This is in an unperturbed setting ($\varepsilon = 0$).

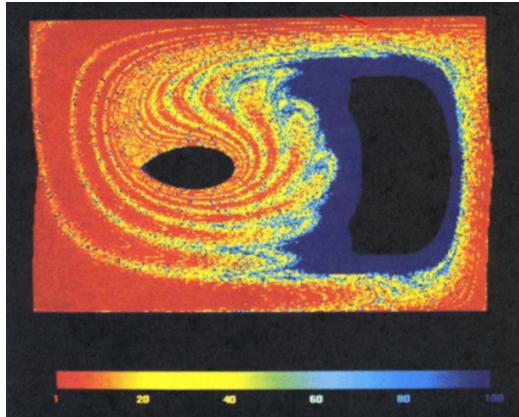


Figure 7.4: Escape time from one convection cell. The number of iterations of the Poincaré map needed for escape is given by the color scale. Here $0 < \varepsilon \leq 1$.

Moving forward we will have a common mathematical setting

$$\dot{x} = f(x) + \varepsilon g(x, t); \quad x \in \mathbb{R}^2, \quad g(x, t) = g(x, t + T); \quad 0 \leq \varepsilon \ll 1; \quad f, g \in \mathcal{C}^1. \quad (7.4)$$

We have a small T -periodic perturbation of a planar ODE which can be studied via a Poincaré map $P_\varepsilon^{t_0} : x_0 \mapsto x(t_0 + T; t_0, x_0)$. Assume that for $\varepsilon = 0$ the system (7.4) has a saddle type fixed point with a homoclinic (or heteroclinic) orbit $x^0(t - t_0)$. Such a system is depicted in Fig. 7.5.

Remark 7.1. The fixed point of $P_0^{t_0}$ given by p_0 is hyperbolic

$$P_0^{t_0}(p_0) = p_0; \quad DP_0^{t_0}(p_0) \text{ has eigenvalues } \lambda_1, \lambda_2 : |\lambda_1| < 1, |\lambda_2| > 1. \quad (7.5)$$

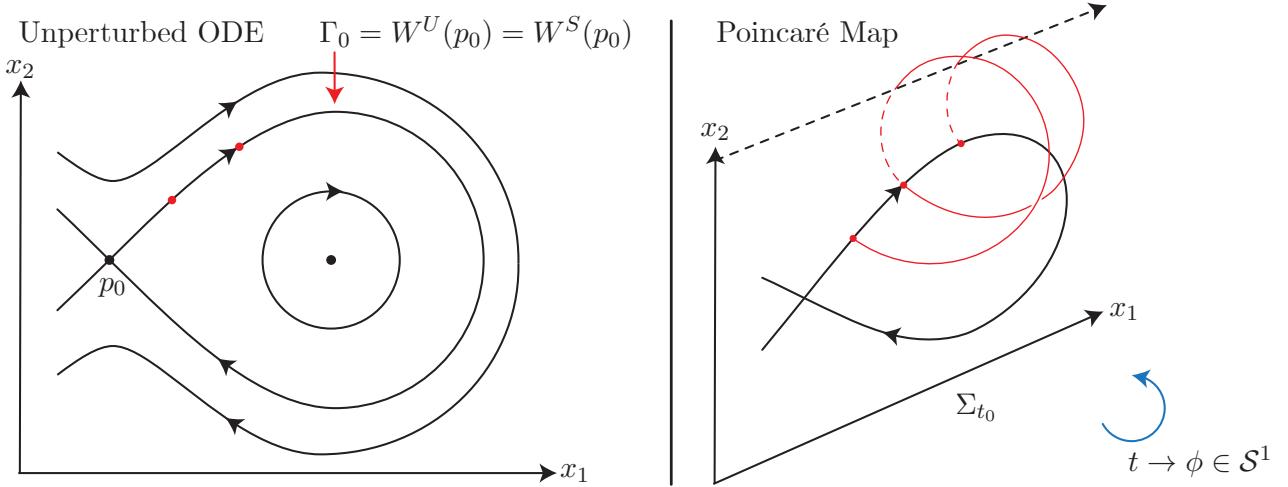


Figure 7.5: Left: An example of a system following the assumptions. The homoclinic orbit Γ_0 is equal to the local unstable and local stable manifolds. Right: The Poincaré map for $\varepsilon = 0$, with time running counterclockwise about the x_1 axis with period T .

More generally we define a hyperbolic fixed point for a map.

Definition 7.1. For a map $F : \mathbb{R}^n \rightarrow \mathbb{R}^n$ and a dynamical system with $x_{k+1} = F(x_k)$, the fixed point p_0 (i.e. $F(p_0) = p_0$) is *hyperbolic* if the linearization's $DF(p_0) \in \mathbb{R}^{n \times n}$ eigenvalues $\lambda_1, \dots, \lambda_n$ never have unitary length: $|\lambda_i| \neq 1$ for $i = 1, \dots, n$.

For each eigenvalue λ_i of the linearization the corresponding eigenvector is given by s_i . Assume that $DF(p_0)$ is semisimple, i.e. all of the eigenvectors are linearly independent. The linearized dynamics at p_0 for $y = x - p_0$ small is

$$y_{k+1} = DF(p_0)y_k \implies y_k = \lambda_1^k c_1 s_1 + \dots + \lambda_n^k c_n s_n. \quad (7.6)$$

If all of the eigenvalues have less than unitary magnitude, $|\lambda_i| < 1$ for $i = 1, \dots, n$, then p_0 is asymptotically stable. Otherwise if there exists an eigenvalue with modulus strictly larger than 1, p_0 is unstable. The relationship of the nonlinear and linearized dynamics are shown in Fig. 7.6.

The fixed point p_0 along with the stable and unstable manifolds ($W^S(p_0)$ and $W^U(p_0)$) persist smoothly under small smooth perturbations to F near p_0 .

7.1 Consequences of hyperbolicity

Hyperbolicity has a few important consequences in our setting. Foremost, the perturbed hyperbolic fixed point $p_\varepsilon^{t_0}$ translates to a hyperbolic periodic orbit for the ODE. Furthermore, the

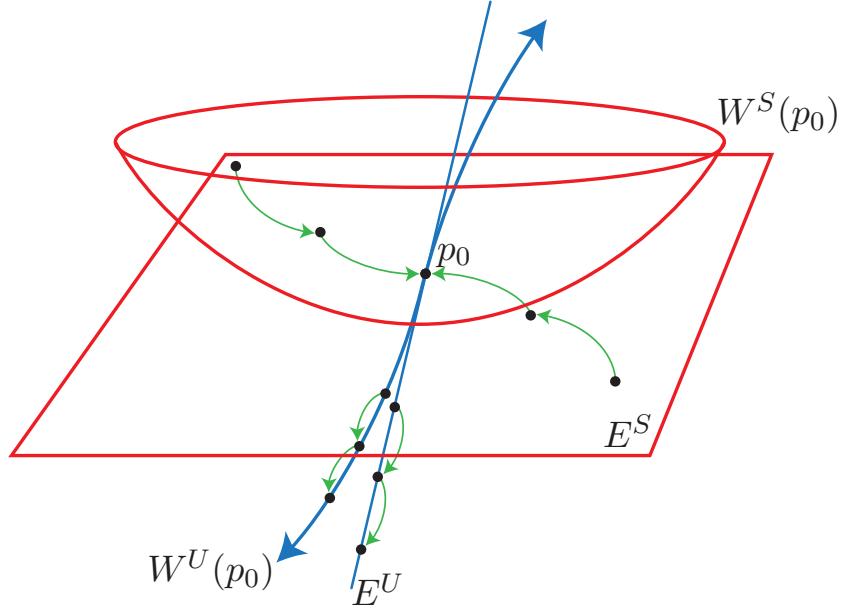


Figure 7.6: The stable ($W^S(p_0)$) and unstable ($W^U(p_0)$) manifolds drawn in relation to the stable (E^S) and unstable (E^U) subspaces. The subspaces are given as the span of the stable (resp. unstable) eigenvectors. The green arrows signify steps of the linearized (resp. original) system.

solutions of the ODE are smooth in ε , thus the solutions within $W^U(p_\varepsilon^{t_0})$ and $W^S(p_\varepsilon^{t_0})$ remain $\mathcal{O}(\varepsilon)$ close to Γ_0 , i.e.

$$x_\varepsilon^S(t; t_0) = x^0(t - t_0) + \varepsilon a^S(t) + \mathcal{O}(\varepsilon^2); \quad t \in [t_0, \infty) \quad (7.7a)$$

$$x_\varepsilon^U(t; t_0) = x^0(t - t_0) + \varepsilon a^U(t) + \mathcal{O}(\varepsilon^2); \quad t \in (-\infty, t_0]. \quad (7.7b)$$

Now we examine what the global shapes of these manifolds are for $\varepsilon > 0$ and if they interact. To do this we will follow an idea from Poincaré, Arnold, and Melnikov. To this end we define the perpendicular vector to f

$$f^\perp(x^0(0)) = \begin{pmatrix} -f_2(x^0(0)) \\ f_1(x^0(0)) \end{pmatrix}. \quad (7.8)$$

We would like to use this to measure the distance between $x_\varepsilon^U(t_0; t_0)$ and $x_\varepsilon^S(t_0; t_0)$. The outlook for this is shown in Fig. 7.7. Thus we have a signed distance function

$$d_\varepsilon(t_0) = \frac{\langle f^\perp(x^0(0)), x_\varepsilon^U(t_0; t_0) - x_\varepsilon^S(t_0; t_0) \rangle}{|f^\perp(x^0(0))|} = \varepsilon \frac{\langle f^\perp(x^0(0)), a^U(t_0) - a^S(t_0) \rangle}{|f^\perp(x^0(0))|} + \mathcal{O}(\varepsilon^2). \quad (7.9)$$

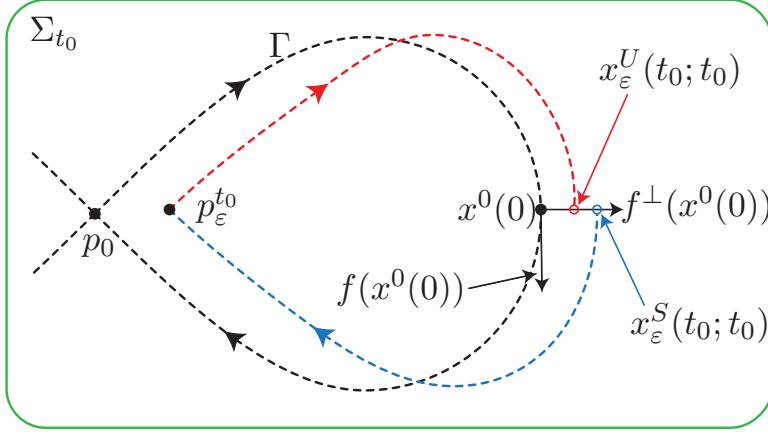


Figure 7.7: The Poincaré-Arnold-Melnikov idea. The dotted black line signifies Γ , the red dot $x_\varepsilon^U(t_0; t_0)$, the blue dot $x_\varepsilon^S(t_0; t_0)$, and the arrow pointing to the right f^\perp . The plane is Σ_{t_0} .

The numerator in the second equation is called the *Melnikov function*

$$M(t_0) = \langle f^\perp(x^0(0)), a^U(t_0) - a^S(t_0) \rangle. \quad (7.10)$$

In fact Melnikov proved an identity for this function.

Theorem 7.2 (Melnikov).

$$M(t_0) = \int_{-\infty}^{\infty} \langle f^\perp(x^0(t - t_0)), g(x^0(t - t_0), t) \rangle dt. \quad (7.11)$$

For the proof, see Theorem 4.5.3 in [Guckenheimer and Holmes, 1990].

Remark 7.3. The integral in Melnikov's theorem converges as $|g(x^0(t - t_0), t)|$ is globally bounded as $t \rightarrow \pm\infty$ and $|f^\perp| = |f|$ and

$$\lim_{t \rightarrow \pm\infty} |f(x^0(t - t_0))| = 0, \quad (7.12)$$

exponentially as p_0 is a hyperbolic fixed point.

Remark 7.4. In order to evaluate $M(t_0)$ we do not need to solve the perturbed ODE $\dot{x} = f(x) + \varepsilon g(x, t)$, instead we only need $x^0(t - t_0)$.

Now observe that

$$d_\varepsilon(t_0) = 0 \iff \varepsilon \frac{M(t_0)}{|f^\perp(x^0(0))|} + \mathcal{O}(\varepsilon^2) = 0 \stackrel{\varepsilon \neq 0}{\iff} \underbrace{\frac{M(t_0)}{|f^\perp(x^0(0))|}}_{F(t_0, \varepsilon) = 0} + \mathcal{O}(\varepsilon) = 0. \quad (7.13)$$

Now we would like to know when we can find a solution $t_0(\varepsilon)$ such that $d_\varepsilon(t_0(\varepsilon)) = 0$ for $\varepsilon > 0$. To do this we use the Implicit Function Theorem. First assume that $F(\bar{t}_0, 0) = 0$, i.e. $M(\bar{t}_0) = 0$, and $\frac{\partial F}{\partial t_0}(\bar{t}_0, 0) \neq 0$, i.e. $\frac{\partial}{\partial t_0} M(\bar{t}_0) \neq 0$. If this condition is fulfilled the root is called *transverse*. Then there exists a unique $t_0(\varepsilon) = \bar{t}_0 + \mathcal{O}(\varepsilon)$ which solves $F(t_0(\varepsilon), \varepsilon) = 0$ for $\varepsilon \neq 0$ small enough. Also $t_0(\varepsilon)$ is smooth if $F(t_0, \cdot)$ is smooth.

A transverse zero for $M(t_0)$ implies that the Melnikov distance $d_\varepsilon(t_0(\varepsilon)) = 0$, in turn implying that the stable and unstable manifolds intersect, $W^S(p_\varepsilon^{t_0}) \cap W^U(p_\varepsilon^{t_0}) \neq \emptyset$. Therefore we have an element $q \in W^S(p_\varepsilon^{t_0}) \cap W^U(p_\varepsilon^{t_0})$, for this q we also have that for every $n \in \mathbb{Z}$

$$P_\varepsilon^n(q) \in W^S(p_\varepsilon^{t_0}) \cap W^U(p_\varepsilon^{t_0}). \quad (7.14)$$

Therefore we have that for each iterate of the Poincaré map there is a unique point in both the stable and unstable manifolds, so these must intersect each other infinitely many times. This behavior is shown in Fig. 7.8. In the homoclinic tangle, we can see the accumulation

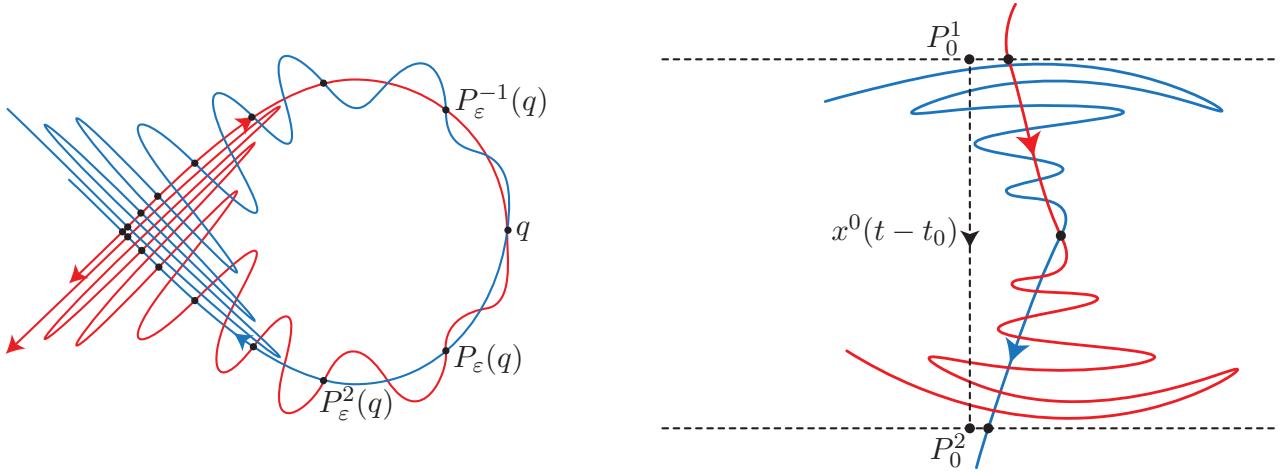


Figure 7.8: The infinite intersections of the stable and unstable manifolds for a transverse zero. Left: The homoclinic case, this is aptly called a *homoclinic tangle*. Right: The heteroclinic case.

of trajectories that are caused by the accumulation of the iterates of the Poincaré map. The accumulation can be formally seen as a consequence of the Λ -lemma [Palis, 1967, Palis, 1969].

Remark 7.5. It is impossible for stable and unstable manifolds to intersect themselves. This is usually argued by stating that q cannot have two distinctive preimages under a diffeomorphism. However, this is not sufficient as a self intersection does not imply that two distinct preimages exist, for instance a looping manifold structure and suitable Poincaré map as in Fig. 7.9. A priori, this is possible to occur. In this case the existence of one loop actually implies the

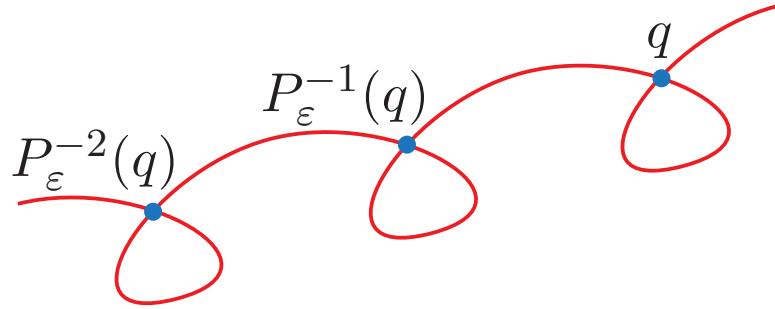


Figure 7.9: An example for a self-intersecting manifold without q having multiple preimages.

existence of infinitely many converging to p_ε . Thereby there must exist a loop in every arbitrarily small neighborhood of p_ε , this then contradicts the Hartman-Grobman Theorem as our system must be topologically equivalent to the linear saddle in a small enough neighborhood.

Remark 7.6. For every pair of intersections $P_\varepsilon^k(q)$ and $P_\varepsilon^{k+1}(q)$, there exists at least another intersection between the stable and unstable manifolds. This is due to the fact that P_ε is orientation preserving (i.e. $DP_\varepsilon(q)$ is an orientation preserving linear map). The implication here is illustrated in Fig. 7.10.

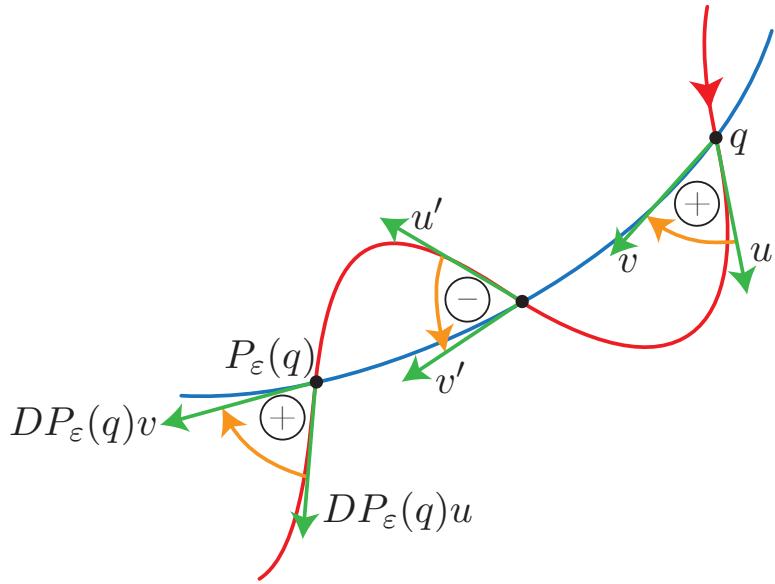


Figure 7.10: Another intersection of the stable and unstable manifolds exists between two sequential iterates of the Poincaré map due to the preservation of orientation under this map.

The orientation preserving nature of P_ε can be seen by using Liouville's Theorem. It states that for a general dynamical system $\dot{x} = f(x, t)$, the linearization of its flow-map $F_{t_0}^t : x_0 \mapsto$

$x(t; t_0, x_0)$ satisfies

$$\det(DF_{t_0}^t(p)) = \exp\left(\int_{t_0}^t \nabla \cdot (f(x(s; t_0, p), s)) ds\right). \quad (7.15)$$

In our case we find

$$\det(DP_\varepsilon(q)) = \exp\left(\int_{t_0}^{t_0+T} \nabla \cdot (f + \varepsilon g)|_{x_\varepsilon(t); x_\varepsilon(t_0)=q} ds\right) > 0 \quad (7.16)$$

Therefore $DP_\varepsilon(q)$ is orientation preserving, in fact Poincaré maps in general are orientation preserving.

Example 7.3 (The forced-damped Duffing equation). Recall the dynamical system

$$\begin{cases} \dot{x} = y \\ \dot{y} = x - x^3 + \varepsilon(\gamma \cos(\omega t) - \delta y) \end{cases}; \quad |\varepsilon| \ll 1. \quad (7.17)$$

Here the forcing amplitude is given by γ and the linear damping coefficient by δ . We separate the right hand side into two parts

$$f(x, y) = \begin{pmatrix} y \\ x - x^3 \end{pmatrix}, \quad g(x, y) = \begin{pmatrix} 0 \\ \gamma \cos(\omega t) - \delta y \end{pmatrix}. \quad (7.18)$$

The Hamiltonian for the unperturbed system is given by

$$H = \frac{1}{2}y^2 - \frac{1}{2}x^2 + \frac{1}{4}x^4. \quad (7.19)$$

Since the Hamiltonian is conserved, the trajectories of the unperturbed system are contained in the level sets of H . This allows us to draw the phase portrait that is shown in Fig. 7.11. The homoclinic orbit Γ_0 is contained in the $H = 0$ level set of the Hamiltonian, which means that along this orbit we have

$$\frac{1}{2}y^2 - \frac{1}{2}x^2 + \frac{1}{4}x^4 = 0. \quad (7.20a)$$

$$\frac{dx}{dt} = y = \pm \sqrt{x^2 - \frac{1}{2}x^2} = F(x). \quad (7.20b)$$

Locally we can obtain a parametrization of the heteroclinic orbit by solving this separable differential equation as

$$\frac{dx}{dt} = F(x) \implies \int_0^x \frac{d\tilde{x}}{F(\tilde{x})} = \int_0^t dt. \quad (7.21)$$

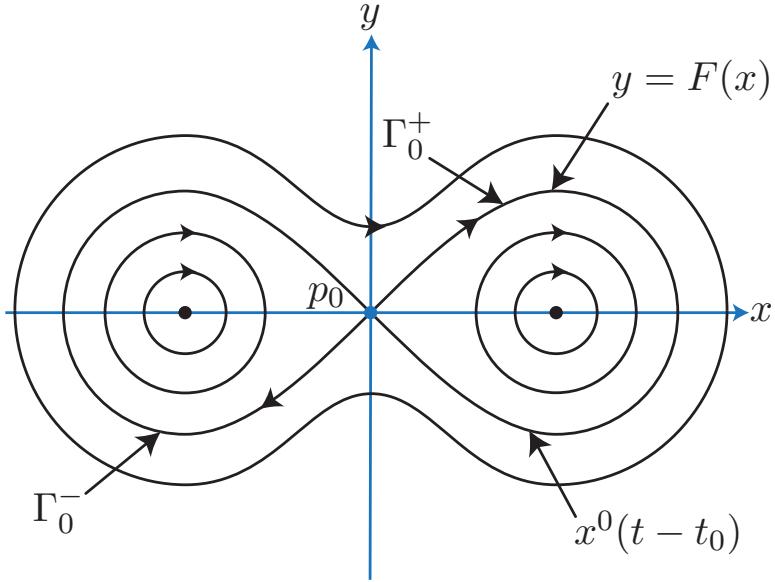


Figure 7.11: The phase portrait for the forced-damped Duffing oscillator for $\varepsilon = 0$.

Hence, on Γ_0^+ , we have

$$x^0(t) = \begin{pmatrix} \sqrt{2}\operatorname{sech}(t) \\ -\sqrt{2}\operatorname{sech}(t)\tanh(t) \end{pmatrix}. \quad (7.22)$$

As discussed above, for $\varepsilon > 0$, the homoclinic orbit splits into a stable and an unstable manifold. We can now calculate the Melnikov function to study whether these manifolds intersect.

$$M^+(t_0) = \int_{-\infty}^{\infty} \langle f^\perp(x^0(t-t_0), g(x^0(t-t_0), t)) \rangle dt \quad (7.23a)$$

$$= -\frac{4\delta}{3} + \sqrt{2}\gamma\pi\omega\operatorname{sech}\left(\frac{\pi\omega}{2}\right) \sin(\omega t_0). \quad (7.23b)$$

Next, we ask if the Melnikov function is equal to zero. For an accumulation of intersection points, a homoclinic tangle, we also need a transverse zero. To this end we define $R^0(\omega) = \frac{4\cosh(\frac{\pi\omega}{2})}{3\sqrt{2}\pi\omega}$. Thus the zeros of the Melnikov function are given by

$$R^0(\omega) = \frac{\gamma}{\delta} \sin(\omega t_0). \quad (7.24)$$

For a zero to exist, $R^0(\omega)$ cannot exceed the value $\frac{\gamma}{\delta}$, as $\sin(\omega t_0)$ is bounded by 1. Therefore when we have $R^0(\omega) \leq \frac{\gamma}{\delta}$, a root of the Melnikov function exists. This root is non-transverse only when it is quadratic, i.e. when only a peak of the sine function intersects with $R^0(\omega)$. This

occurs exactly when $R^0(\omega) = \frac{\gamma}{\delta}$. In all other cases, the sine wave intersects the curve of $R^0(\omega)$ transversely, i.e. the zero is transverse. This is illustrated in Fig. 7.12, where equality occurs along the graph of R^0 , that is the set on which non-transverse zeros exist. The area above this curve contains all of the point for which the zero is transverse.

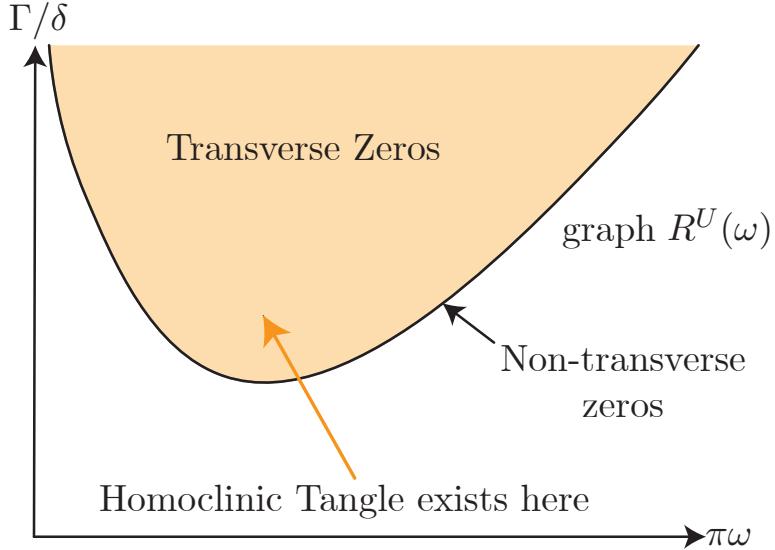


Figure 7.12: The yellow line designating the area above the curve where homoclinic tangles exist for the forced-damped Duffing oscillator.

7.2 Dynamics near the homoclinic tangle & Smale's horseshoe map

We now continue and examine the dynamics near the homoclinic tangle. Consider the dynamical system

$$\dot{x} = f(x) + \varepsilon g(x, t); \quad x \in \mathbb{R}^2; \quad f, g \in \mathcal{C}^1; \quad g(x, t) = g(x, t + T). \quad (7.25)$$

With a linear change of coordinates we can align our coordinate axes with the stable and unstable subspaces of the fixed point. Near the fixed point this leads to *Smale's construction*, which is as shown as in Fig. 7.13.

A model for the right panel of Fig. 7.13 is given by *Smale's horseshoe map*. This is defined by first considering the unit square $S = [0, 1]^2 \subset \mathbb{R}^2$, next the map itself is given by $f : S \rightarrow \mathbb{R}^2$. The map f will not be defined explicitly, and instead geometrically. The unit square is divided

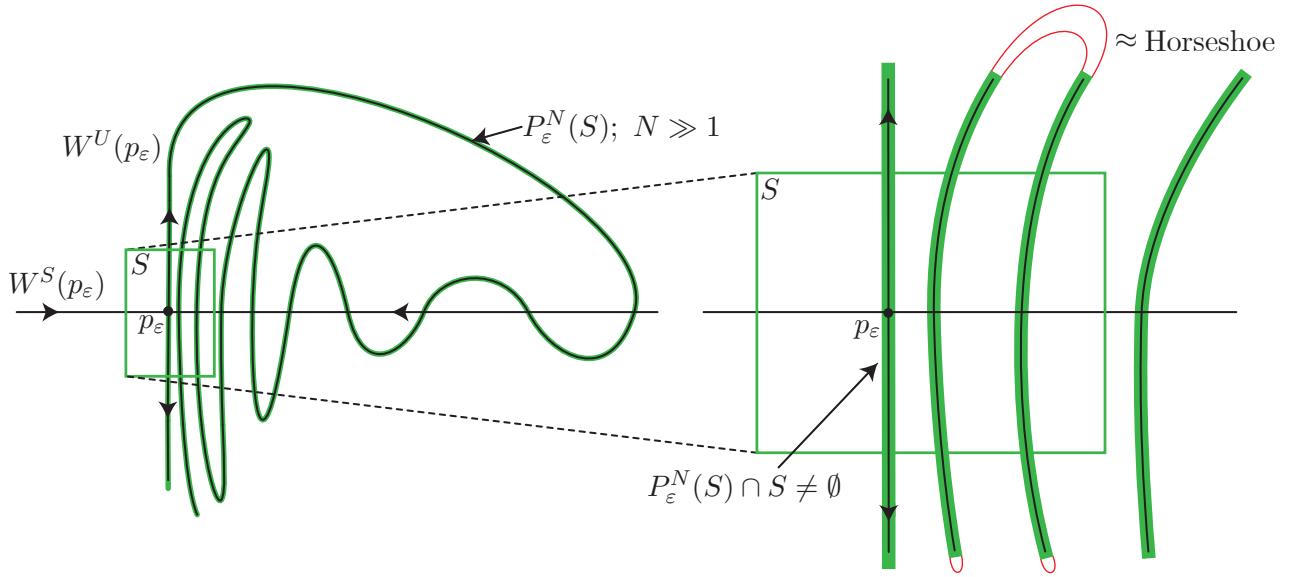


Figure 7.13: Smale's construction is found through an appropriate change of coordinates. Left: The full construction, the N is large enough such that $P_\varepsilon^N(S) \cap S = \emptyset$, where S is the green rectangle around p_ε . Right: Focus on the set S , the red lines designate a horseshoe-like structure.

into three horizontal strips, the upper one being labelled as H_2 and the lower H_1 . The square is then squeezed horizontally and stretched vertically, and then smoothly bent to a horseshoe such that the previously upper horizontal strip now forms a vertical strip on the right side of the box V_2 . Similarly, the lower horizontal strip now forms a vertical strip on the left V_1 . Now the unit square can also be divided into three vertical strips, on the left V_1 , an area in the middle, and on the right V_2 . What was previously part of the middle horizontal strip is now mapped outside the unit square. Because the map is continuous, these points connect V_1 and V_2 , but note that precisely defining what happens to these points is not essential in Smale's construction. The horseshoe map is depicted in Fig. 7.14. This is a good model for the dynamics near the homoclinic tangle, if we consider the horseshoe map f to be many iterations of the Poincaré map P_ε^N (for $N \gg 1$).

Using this construction, we would like to know if any initial conditions stay in S for long times (after many iterations of the horseshoe map), despite the overall instability coming from P_ε . To further study this we introduce a couple of definitions.

Definition 7.2. The set of all points that stay in S under k backwards iterations is given by V^k .

Definition 7.3. For a map $f : X \rightarrow Y$, we denote the *preimage* of a set $U \subset Y$ (or point

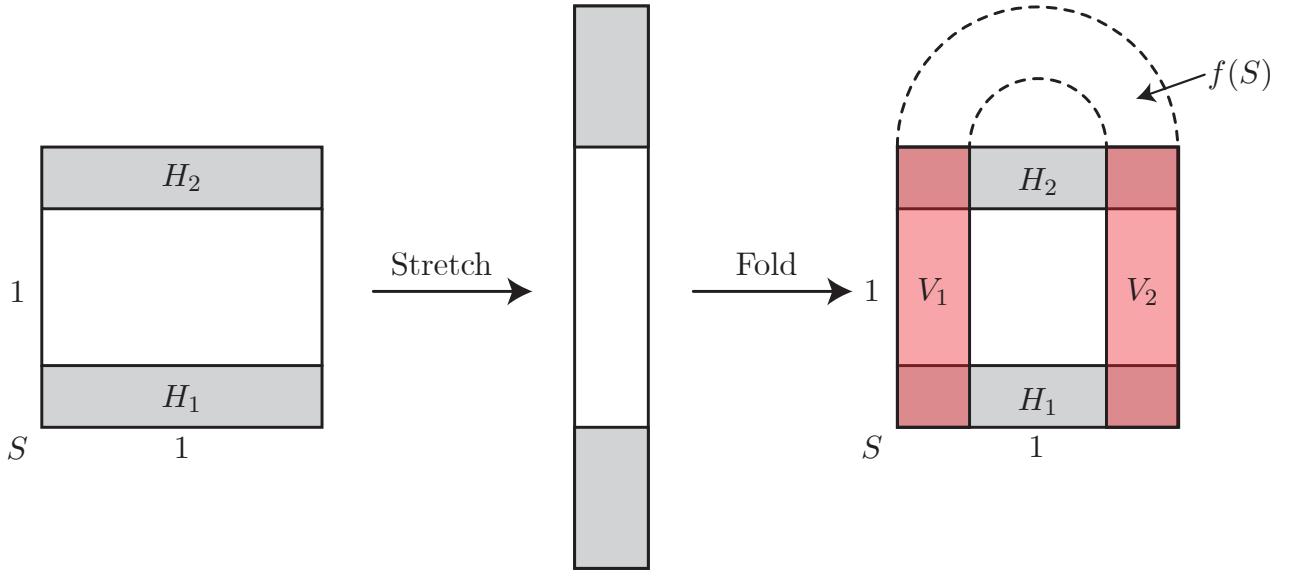


Figure 7.14: The two geometric steps of the Smale map, first a vertical stretch and then a fold into the horseshoe.

$p \in Y$) as

$$f^{-1}(U) = \{x \in X : f(x) \in U\}. \quad (7.26)$$

The set of points that will be mapped into S in one iteration is given by $S \cap f^{-1}(S)$. This means, that the set of points that stay in S for one backward iteration, V^1 , is

$$V^1 = f(S \cap f^{-1}(S)) = f(S) \cap f(f^{-1}(S)) = f(S) \cap S. \quad (7.27)$$

However, from the definition of the horseshoe map, $V^1 = f(S) \cap S$ is $V_1 \cup V_2$, which is a closed, nonempty set. Similarly, for 2 backward iterations we have

$$V^2 = f^2(S \cap f^{-1}(S) \cap f^{-2}(S)) = f^2(S) \cap f(S) \cap S = V_{11} \cup V_{12} \cup V_{22} \cup V_{21} = \bigcup_{i_k \in \{1,2\}} V_{i_1 i_2}. \quad (7.28)$$

The sets V_{ij} are defined in Fig. 7.15, the indices for each component are given by their orbit: The first index denotes which vertical strip V_i ($i \in \{1, 2\}$) the preimage of a point is in. The second index denotes which vertical strip it is mapped into. That is, V_{11} is the set of points that are mapped from V_1 to V_1 under one iteration, V_{12} is the set of points mapped from V_1 to V_2 and so on. Also the set V^2 is closed, nonempty, contained in V^1 , and has 2^2 components.

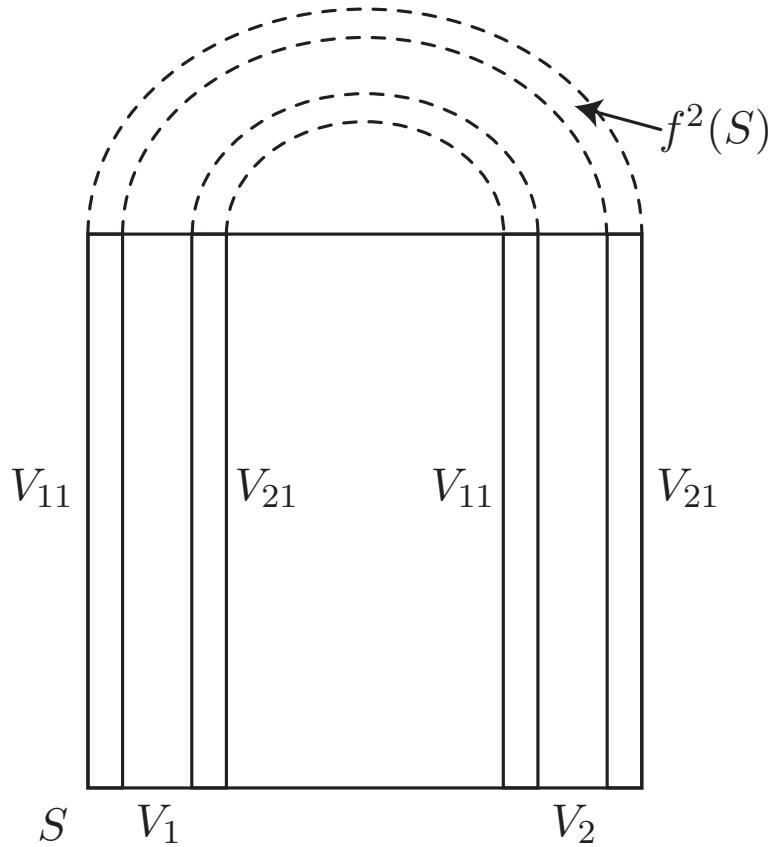


Figure 7.15: The four components of V^2 with their index labels.

This continues for any number of backward iterates k

$$V^k = f^k (S \cap f^{-1}(S) \cap \dots \cap f^k(S)) = f^k(S) \cap \dots \cap f^1(S) \cap S = \bigcup_{i_j \in \{1,2\}} V_{i_1 \dots i_k}. \quad (7.29)$$

This set has 2^k disjoint components (vertical strips) and is closed, nonempty, and contained in V^{k-1} . Therefore we have an infinite, nested sequence of nonempty and closed sets, therefore by Cantor's theorem we have

$$V^\infty = \bigcap_{i=1}^{\infty} V^i \neq \emptyset.$$

(7.30)

Remark 7.7. Cantor's theorem comes from the famous Cantor dust which arises by first removing the open middle third of the unit interval in \mathbb{R} , and then iteratively removing the open middle third of each remaining interval. The Cantor dust is a nonempty (in fact uncountably

infinite) and disconnected set of points. For more on Cantor's theorem and Cantor's dust, see any book on measure theory.

Definition 7.4. A set U is called *perfect* if it is closed and has no isolated points, i.e. given any point $x \in U$ and any distance ε we can find a distinct element $y \in U$ such that $\|x - y\| < \varepsilon$.

Remark 7.8. This definition uses that the topology on U was induced by a metric, which may not be the case in general, but is not relevant to the content here.

In our case V^∞ is in fact a *Cantor set of lines*, i.e. it is a closed, nonempty set which is totally disconnected (the largest connected component is a line), and perfect.

Remark 7.9. By a Cantor set of lines, we mean that there is a continuous, surjective map from this set of lines to a Cantor set of points. This map could be the one taking points along a vertical line to an intersection of the line with an arbitrarily selected horizontal line.

Definition 7.5. Analogous to before we define the set of all points that stay in S under k forwards iterations as H^k . For each k the set H^k is a nonempty set of 2^k closed, disjoint, horizontal strips, forming a nested sequence.

By similar reasoning as before we have

$$H^\infty = \bigcap_{k=1}^{\infty} H^k \neq \emptyset. \quad (7.31)$$

Definition 7.6. The set of all points that stay in S for all forward and backward iterates of the Smale horseshoe map is given by

$$\Lambda = H^\infty \cap V^\infty. \quad (7.32)$$

The geometry of this set is shown in Fig. 7.16.

Remark 7.10. The set Λ is a Cantor set of points.

We have then identified the invariant set Λ for the horseshoe map. Returning to our case study of a Poincaré map with a homoclinic tangle, the points in Λ keep coming back forever to the unstable periodic orbit as shown in Fig. 7.17.

7.3 Dynamics of Smale's horseshoe map on the invariant set Λ

We label any point $p \in \Lambda$ with an infinitely long binary sequence, which reflects the itinerary of p in Λ under iterations of f . For a given p , if we have

$$\dots, f^{-2}(p) \in V_{s_{-3}}, f^{-1}(p) \in V_{s_{-2}}, p \in V_{s_{-1}}, p \in H_{s_0}, f(p) \in H_{s_1}, f^2(p) \in H_{s_2}, \dots \quad (7.33)$$

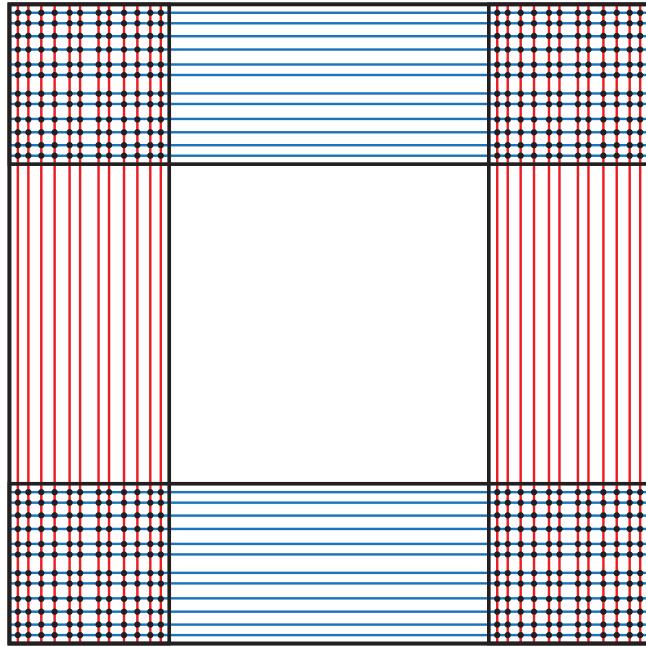


Figure 7.16: The geometry of the set Λ , with the four regions containing points circled in orange. The red vertical lines represent the set V^∞ and the blue horizontal lines the set H^∞ .

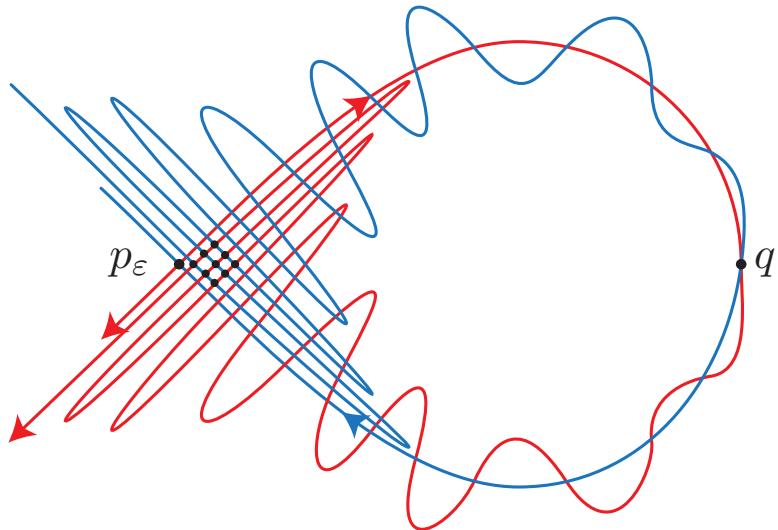


Figure 7.17: The homoclinic tangle with the elements of Λ being denoted by the black dots within S , the neighborhood around p_ε .

for $s_i \in \{1, 2\}$ then we would have the itinerary

$$s = \dots s_{-3} s_{-2} s_{-1} \bullet s_0 s_1 s_2 \dots \quad (7.34)$$

This is a unique symbolic encoding for all points in Λ . Therefore there is a one-to-one correspondence between points p in the invariant set Λ and sequences s in the space Σ of doubly infinite sequences of two symbols Σ . Such a space is just the space of outcomes for an infinite coin-tossing experiment. This equivalence of Λ to the doubly infinite sequence of two symbols leads us to *symbolic dynamics*: the analogue of the horseshoe map on the symbol space Σ .

In order to study these symbolic dynamics we need to know what the symbolic encoding of $f(p)$ is. For the analogue $\dots s_{-2}s_{-1}.s_0s_1s_2\dots = s \in \Sigma$ of $p \in \Lambda$, we can read off that $f(p) \in H_{s_1}$. Continuing in this fashion, we can see that $f(f(p)) = f^2(p) \in H_{s_2}$ and so on, therefore $f(p) \in H^\infty$. By recalling that $f(H_{s_0}) = V_{s_0}$ we can see that $f(p) \in V_{s_0}$. Applying the definition of f we find that $f^{-1}(f(p)) = p \in V_{s_{-1}}$. Therefore $f(p) \in V^\infty$ and thereby $f(p) \in \Lambda$.

From our examination of the forward iterates of $f(p)$ we know that the symbolic encoding of $f(p)$ to the right of the dot is

$$\dots s_1s_2s_3\dots \quad (7.35)$$

Similarly, from our examination of the backwards iterates of $f(p)$ we know that the symbolic encoded of $f(p)$ to the left of the dot is

$$\dots s_{-2}s_{-1}s_0\dots \quad (7.36)$$

Putting these together we get the symbolic encoding for $f(p)$ as

$$\dots s_{-2}s_{-1}s_0.s_1s_2s_3\dots \in \Sigma. \quad (7.37)$$

So the encoding of $f(p)$ in the symbolic space is just a shift to the left on the encoding of p .

Definition 7.7. We define the following three maps:

- (i) The *symbolic encoding map*

$\phi : \Lambda \rightarrow \Sigma; \quad p \mapsto s = \phi(p);$

(7.38)

This map can be shown to be a homeomorphism, i.e. it is continuous and has a continuous inverse. For a proof we refer to Theorem 5.1.1 in [Guckenheimer and Holmes, 1990].

- (ii) The *Smale horseshoe map restricted to Λ*

$\tilde{f} : \Lambda \rightarrow \Lambda; \quad p \mapsto f(p);$

(7.39)

- (iii) The *Bernoulli shift map*

$\sigma : \Sigma \rightarrow \Sigma; \quad \dots s_{-2}s_{-1}.s_0s_1s_2\dots \mapsto \dots s_{-2}s_{-1}s_0.s_1s_2s_3\dots$

(7.40)

We have that the following diagram commutes

$$\begin{array}{ccc} \Sigma & \xrightarrow{\sigma} & \Sigma \\ \phi \uparrow & & \uparrow \phi \\ \Lambda & \xrightarrow{\tilde{f}} & \Lambda \end{array}. \quad (7.41)$$

Therefore $\tilde{f} = f|_{\Lambda}$ is topologically conjugate to a Bernoulli shift on Σ

$$\tilde{f} = \phi^{-1} \circ \sigma \circ \phi. \quad (7.42)$$

Hence, orbits of f are taken to those of σ , with their orientation preserved.

Remark 7.11. To denote an infinitely repeating string, the bar notation will be used i.e.

$$\bar{1} = 111\dots, \quad \bar{123} = 123123\dots. \quad (7.43)$$

It is possible to classify all of the orbits of σ in Σ , the classes and their cardinality are as follows:

- (i) **Fixed points** There are exactly two fixed points of σ , namely $\bar{1}\bar{1}$ and $\bar{2}\bar{2}$;
- (ii) **Period-two orbits** A single such orbit exists, that of $\bar{12}\bar{12}$, as $\sigma^2(\bar{12}\bar{12}) = \sigma(\bar{21}\bar{21}) = \bar{12}\bar{12}$;
- (iii) **Period-three orbits** There exists two such orbits, namely $\bar{112}\bar{112} \rightarrow \bar{121}\bar{121} \rightarrow \bar{211}\bar{211}$ and $\bar{221}\bar{221} \rightarrow \bar{212}\bar{212} \rightarrow \bar{122}\bar{122}$.
- (iv) **Period- k orbits** For every natural number k , there exists

$$N(k) = \frac{1}{k} \left(2^k - \sum_{\langle i,k \rangle} i N(i) \right)$$

(7.44)

orbits of period exactly k , the notation $\langle i, k \rangle$ signifies integers i which are divisors of k .

Exercise 7.1. Show that the number of k -periodic orbits for the Bernoulli shift map on two symbols is

$$N(k) = \frac{1}{k} \left(2^k - \sum_{\langle i,k \rangle} i N(i) \right), \quad (7.45)$$

where $\langle i, k \rangle$ means that the integer i divides the integer k .

There exists a countable infinity (i.e. the elements can be organized uniquely into an infinite countable sequence) of periodic orbits of arbitrarily high period. The countable sequence ordering can be done by listing each orbit in the row corresponding to its minimal period. An orbit of period k is made up of k points along that orbit. We assign a single number to the orbit by concatenating the decimal representations of the individual points along the orbit, starting from the one with the smallest value. Within a row, we can then order the periodic orbits based on the magnitudes of the concatenated numbers representing them. Thus $\overline{112.112}$ would go in the third row as it has the minimal period 3, as the fixed points have period length 0 and occupy the first row, and in the leftmost position as $\overline{112} \prec \overline{122}$. Each row has a finite number of elements, so ordering each row is not a problem. Then the complete order is done by starting with 0 for the first element in the first row (the fixed point $\bar{1}.\bar{1}$), then reading from left to right before continuing to the next row. This list is well defined and countably infinite, containing all of the orbits of σ .

Before further study, it is important to introduce a notion of distance on the space Σ .

Definition 7.8. The distance on the space Σ between two elements s and s' is given by

$$d(s, s') = \sum_{i=-\infty}^{\infty} \frac{|s_i - s'_i|}{2^{|i|}}. \quad (7.46)$$

Therefore two symbols are close if they agree on large central blocks.

Remark 7.12. The function d satisfies all the axioms of a distance and the tuple (Σ, d) forms a complete metric space.

Proposition 7.13 (The periodic orbits are dense). *The set of all periodic orbits is dense in Σ .*

Proof. Given an arbitrary element $s \in \Sigma$ and $k \in \mathbb{N}$, we would like to find a 2^{-k} periodic approximation of s . To do this define the two strings

$$S_- = s_{-(k+1)} \dots s_{-1}, \quad S_+ = s_0 \dots s_{k+1}. \quad (7.47)$$

For any two strings (of infinite length) L and R the element

$$s' = LS_- . S_+ R \quad (7.48)$$

is 2^{-k} close to s . This is because they can only disagree at indices with modulus strictly greater than $k + 1$, therefore

$$d(s', s) \leq \sum_{j=-(k+2)}^{-\infty} \frac{1}{2^{-|j|}} + \sum_{j=k+2}^{\infty} \frac{1}{2^{-|j|}} = 2 \cdot 2^{-(k+1)} = 2^{-k}. \quad (7.49)$$

Now it is clear that the periodic element

$$s^* = \overline{S_+ S_-} \cdot \overline{S_+ S_-} = L' S_- \cdot S_+ R' \quad (7.50)$$

is a 2^{-k} approximation for s . As s and k were arbitrary, this completes the proof. \square

Proposition 7.14. *Consider a k -periodic point $s = \bar{a} \cdot \bar{a}$, where a is a k -long binary sequence. Let $s' = \bar{a}a' \cdot \bar{a}$, with $a' \neq a$, but an otherwise arbitrary finite length string. Clearly $s \neq s'$, therefore $d(s, s') \neq 0$, when we take the limit we find*

$$\lim_{n \rightarrow \pm\infty} d(\sigma^n(s), \sigma^n(s')) = 0. \quad (7.51)$$

Thus the underlying periodic orbit is unstable. Furthermore, s has a countable infinity of homoclinic orbits. As the periodic orbit was arbitrary, these statements hold for all periodic orbits, hence all periodic orbits are unstable and have a countable infinity of homoclinic orbits.

Remark 7.15. Also note that any two periodic orbits (of arbitrary period) are connected by a countable infinity of heteroclinic orbits.

Together, there exists a countable infinity of periodic and asymptotically periodic orbits, a sketch of this behavior is shown in Fig. 7.18.

Definition 7.9. A point s is asymptotically periodic if there is a periodic orbit p such that

$$\boxed{\lim_{n \rightarrow \infty} d(\sigma^n(s), \sigma^n(p)) = 0.} \quad (7.52)$$

For example, heteroclinic orbits are asymptotically periodic.

In fact there also exists an uncountable infinity of non-periodic (not even asymptotically periodic) orbits in Σ . To show this, we will show that Σ is uncountable, i.e. that it cannot be arranged into a well-defined, countable sequence.

Exercise 7.2. *Show that any two periodic orbits of the Bernoulli shift map are connected by infinitely many heteroclinic orbits.*

Definition 7.10. The binary expansion of a number $x \in [0, 1]$ is the sequence $(b_i)_{i=1}^\infty$ of 1s and 0s such that

$$x = \sum_{i=1}^{\infty} \frac{b_i}{2^i}. \quad (7.53)$$

To calculate this sequence define the map $g(x) = 2x \bmod 1$. Then generate the sequence $(x_i)_{i=1}^\infty = (x, g(x), g^2(x), \dots)$. Finally, set each binary coefficient as

$$b_i = \begin{cases} 0 & 2x_i < 1 \\ 1 & 2x_i \geq 1 \end{cases}. \quad (7.54)$$

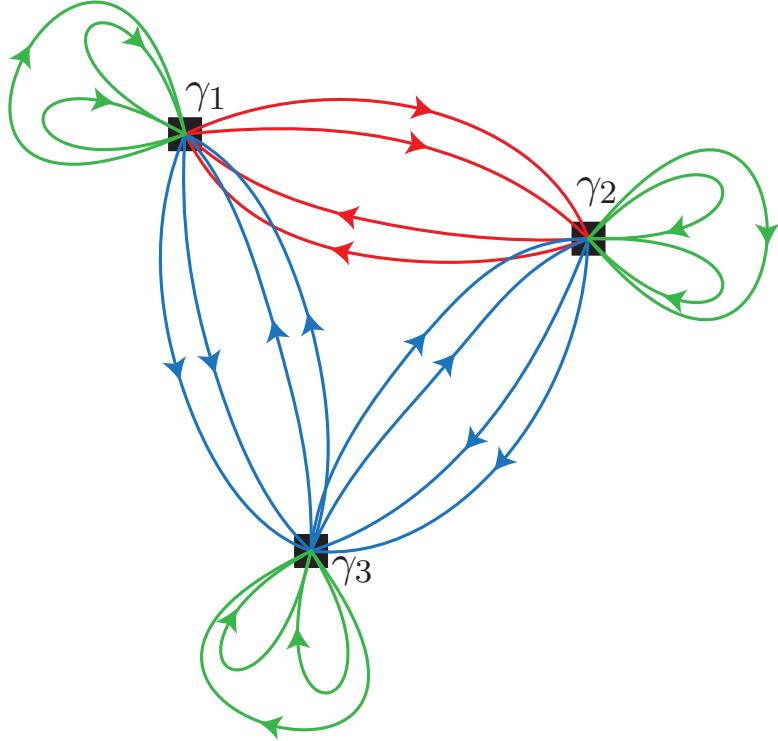


Figure 7.18: A depiction of the (infinitely many) homo- and heteroclinic orbits which connect periodic orbits of σ .

Proposition 7.16. *The set Σ is uncountable.*

Proof. It is enough to show that the set $\tilde{\Sigma}$ defined by

$$\tilde{\Sigma} = \{s_1 s_2 s_3 \dots : s_i \in \{0, 1\}\} \quad (7.55)$$

is uncountable. This suffices as we can injectively embed $\tilde{\Sigma}$ into Σ via the map $.s_1 s_2 \dots \mapsto \bar{1}.(s_1 + 1)(s_2 + 1) \dots$

Every binary expansion for all real numbers in the interval $[0, 1]$ is contained within $\tilde{\Sigma}$, therefore the cardinality of $\tilde{\Sigma}$ is at least as large as that of $[0, 1]$ which is uncountable. In fact, the binary expansion is a bijection, and these two sets have the exact same cardinality. \square

Example 7.4 (Binary expansion). Let $x = \frac{1}{5}$, we find the iterates of g

$$x = \frac{1}{5} \quad 2 \cdot \frac{1}{5} < 1 \implies x_1 = 0 \quad (7.56a)$$

$$g\left(\frac{1}{5}\right) = \frac{2}{5} \quad 2 \cdot \frac{2}{5} < 1 \implies x_2 = 0 \quad (7.56b)$$

$$g^2\left(\frac{1}{5}\right) = g\left(\frac{2}{5}\right) = \frac{4}{5} \quad 2 \cdot \frac{4}{5} > 1 \implies x_3 = 1 \quad (7.56c)$$

$$g^3\left(\frac{1}{5}\right) = g\left(\frac{4}{5}\right) = \frac{8}{5} \bmod 1 = \frac{3}{5} \quad 2 \cdot \frac{3}{5} > 1 \implies x_4 = 1 \quad (7.56d)$$

$$g^4\left(\frac{1}{5}\right) = g\left(\frac{3}{5}\right) = \frac{6}{5} \bmod 1 = \frac{1}{5} \quad 2 \cdot \frac{1}{5} < 1 \implies x_5 = 0. \quad (7.56e)$$

Since $g^4\left(\frac{1}{5}\right) = \frac{1}{5}$, we have reached a repeating loop (since our seed value was $\frac{1}{5}$), therefore we have that

$$\frac{1}{5} = 0 \cdot \frac{1}{2} + 0 \cdot \frac{1}{4} + 1 \cdot \frac{1}{8} + 1 \cdot \frac{1}{16} + \dots = .\overline{0011}. \quad (7.57)$$

Remark 7.17. Recall that for a metric space (X, d) when we write $\mathcal{B}(x, r)$ we refer to the open ball of radius r around x , i.e.

$$\mathcal{B}(x, r) = \{y \in X : d(x, y) < r\}. \quad (7.58)$$

Proposition 7.18. *The map σ has a dense orbit in Σ , i.e. an orbit which approaches any point in Σ arbitrarily closely.*

Proof. Without loss of generality, assume that $\delta < 1$. Next, note that the maximal distance between two elements occurs when they disagree at every entry, so they have the distance

$$\sum_{i=-\infty}^{\infty} \frac{1}{2^{|i|}} = 2. \quad (7.59)$$

Therefore, the entire space is included in the ball $\mathcal{B}(\bar{1}\bar{1}, 2)$. Now to show that for any given distance δ , we want to find an element s^* such that its orbit gets δ close to every element of Σ . Let

$$k = \min \{j \in \mathbb{N} : 2^{-j} \leq \delta\}, \quad (7.60)$$

i.e. the first index in the binary expansion of δ which is not 0. Now we would like to find a finite $2^{-(k+1)}$ -covering of Σ , i.e. a set of elements s_i^k such that $\bigcup_i \mathcal{B}(s_i^k, 2^{-(k+1)})$. We start out

with the covering $\bigcup_{j=1}^{\infty} \mathcal{B}(q_j, 2^{-(k+1)})$, with the q_j being an enumeration of the periodic orbits. As we previously noted, Σ is bounded as it can be contained in a finite ball, furthermore Σ is also closed (the intersection of two closed sets is always closed due to the de Morgan identity), thus Σ is compact and a finite subcovering exists. Hence we can deduce that only finitely many q_i are needed, call these $(s_i^k)_{i=1}^N$, and let the string representation of each s_i^k be given by $S_i^k \cdot \bar{S}_i^k$.

If we can construct s_k^* such that its orbit passes through each of the $\mathcal{B}(s_i^k, 2^{-(k+1)})$, then we know that the orbit of s_k^* gets within distance 2^{-k} of every element in Σ , by the triangle inequality.

For any given s_i^k the element

$$L \underbrace{S_i^k \dots S_i^k}_{(k+3) \text{ times}} \cdot \underbrace{S_i^k \dots S_i^k}_{(k+4) \text{ times}} R \quad (7.61)$$

is within distance $2^{-(k+1)}$ of s_i^k , for L and R arbitrary. This is because the string S_i^k is of at least length 1, so the two strings can disagree only at elements with index modulus strictly greater than $k+2$ (we need the one more copy of S_i^k on the right side as the 0th index is to the right of the \cdot), therefore the distance between them is upper bounded by

$$\sum_{j=-(k+3)}^{-\infty} \frac{1}{2^{-|j|}} + \sum_{j=k+3}^{\infty} \frac{1}{2^{-|j|}} = 2 \cdot 2^{-(k+2)} = 2^{-(k+1)}. \quad (7.62)$$

Therefore, using $2k+7$ copies of S_i^k we are able to approximate s_i^k within distance $2^{-(k+1)}$. We now construct the element

$$s_k^* = L \cdot \underbrace{S_1^k \dots S_1^k}_{2k+7 \text{ times}} \dots \underbrace{S_N^k \dots S_N^k}_{2k+7 \text{ times}} R \quad (7.63)$$

for L and R arbitrary. We now can see that for every $i \in \{1 \dots N\}$, the element

$$L' \underbrace{S_i^k \dots S_i^k}_{(k+3) \text{ times}} \cdot \underbrace{S_i^k \dots S_i^k}_{(k+4) \text{ times}} R', \quad (7.64)$$

is contained within the orbit of s_k^* for L' and R' chosen accordingly. Therefore we have now constructed the the desired s_k^* . Note here that this construction is well defined as each S_i^k is of finite length. This construction can be done for each k : start from the infinite cover using periodic orbits, then construct a finite subcover using the compactness of Σ , finally construct the element s_k^* by concatenating the $2^{-(k+1)}$ approximations of the elements which constitute the finite subcover.

To construct s^* , concatenate each of the strictly defined blocks (the block the the right of \cdot and left R) of s_k^* for each $k \in \mathbb{N}$. For any n , we know that we can get within distance 2^{-n}

distance of each element in Σ , therefore we have constructed a dense orbit. This construction is well defined as the strictly defined block of any given s_k^* is always of finite length. Thus we have constructed the element, whose orbit is dense in Σ . \square

Proposition 7.19. *The map σ is topologically transitive, i.e. for any two open sets A and B in Σ , there exists an $N \in \mathbb{N}$ such that*

$$\sigma^N(A) \cap B \neq \emptyset. \quad (7.65)$$

This is also called the *mixing property*. In the case of the Bernoulli shift map, it can be concluded from the existence of a dense orbit.

Exercise 7.3. *Prove Proposition 7.19, i.e. show that the Bernoulli shift map σ is topologically transitive on the symbol space Σ with respect to the metric $d(\cdot, \cdot)$ defined in class. Specifically, show that for any two open sets $A, B \subset \Sigma$, there exists an integer N such that $\sigma^N(A) \cap B \neq \emptyset$. (Hint: Use the existence of a dense orbit for σ in Σ : there exists a symbol sequence $s^* \in \Sigma$ with the following property: for any $s \in \Sigma$ and for any $\delta > 0$, there exists an integer $N(s, \delta)$ such that $d(\sigma^{N(s, \delta)}(s^*), s) < \delta$.)*

Proposition 7.20 (Sensitive dependence on initial conditions). *Regardless of how close two distinct initial conditions s and s' are chosen, they will not remain close. There exists a $\Delta > 0$ such that for every $s \in \Sigma$ and any $\delta > 0$, there exists an $s' \in \mathcal{B}(s, \delta)$ such that for an N large enough the following holds*

$$d(\sigma^N(s), \sigma^N(s')) > \Delta. \quad (7.66)$$

This is called the uniform instability of all orbits. This fact is depicted in Fig. 7.19.

Exercise 7.4. *Prove Proposition 7.20, i.e. show that the Bernoulli shift map σ has sensitive dependence on initial conditions on the symbol space Σ . Specifically, show that there exists a nonzero distance $\Delta > 0$, such that no matter how close two symbols s^* and \bar{s} are in Σ , we have*

$$d(\sigma^N(s^*), \sigma^N(\bar{s})) > \Delta \quad (7.67)$$

for some N .

Remark 7.21. This by itself implies no complexity: the linear saddle

$$\begin{cases} \dot{x} = \lambda x \\ \dot{y} = -\mu y, \end{cases} \quad (7.68)$$

also has this property.

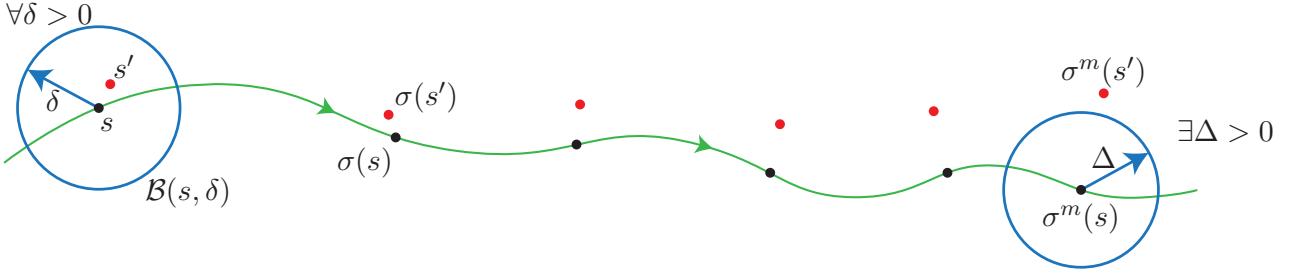


Figure 7.19: The uniform instability of orbits for the map σ .

Remark 7.22. In the case of this map, every $s' \in \mathcal{B}(s, \delta)$ has an N large enough such that the inequality holds.

These observations lead us to the following definition of a chaotic map.

Definition 7.11 (Chaotic map). Let $M : C \rightarrow C$ be a C^0 map on a compact metric space C . We call M *chaotic* if

- (i) M is topologically transitive (the mixing property);
- (ii) M has sensitive dependence on initial conditions (a lack of reproducibility);
- (iii) M has a dense set of periodic orbits in C .

The Bernoulli shift map σ is chaotic, therefore Smale's horseshoe map on the invariant set Λ is chaotic, therefore the horseshoe map on the unit square is chaotic. Since that was a model for the Poincaré map with a homoclinic tangle, we can rigorously conclude that the forced Duffing equation is chaotic. All this followed from a transverse zero of the Melnikov function.

Example 7.5 (Chaos in a 1-dimensional model). Consider the logistic map for resource-limited discrete population growth dynamics

$$x_{n+1} = x_n = ax_n(1 - x_n); \quad a \in \mathbb{R}^+. \quad (7.69)$$

We will now continue with $a = 4$. The points $x = 0$ and $x = \frac{3}{4}$ are fixed points. These are each unstable as $|f'(x)| > 1$ at both of these points. This is not an ODE, and we do not assume that the model is a small perturbation of a simpler system. That is, we are not assuming persistence of known structures. The cobweb diagram is illustrated in Fig. 7.20. We can see that for our choice of a , f maps $[0, 1]$ to $[0, 1]$ and is not invertible. In fact each preimage of $\frac{1}{2}$ reaches 0 in a finite number of iterations. Also note that because the map is quadratic, each point has two preimages. Following a point in backward time is not well defined in the usual sense, but instead, we may take preimages of preimages of $\frac{1}{2}$. This gives rise to two branches of backward

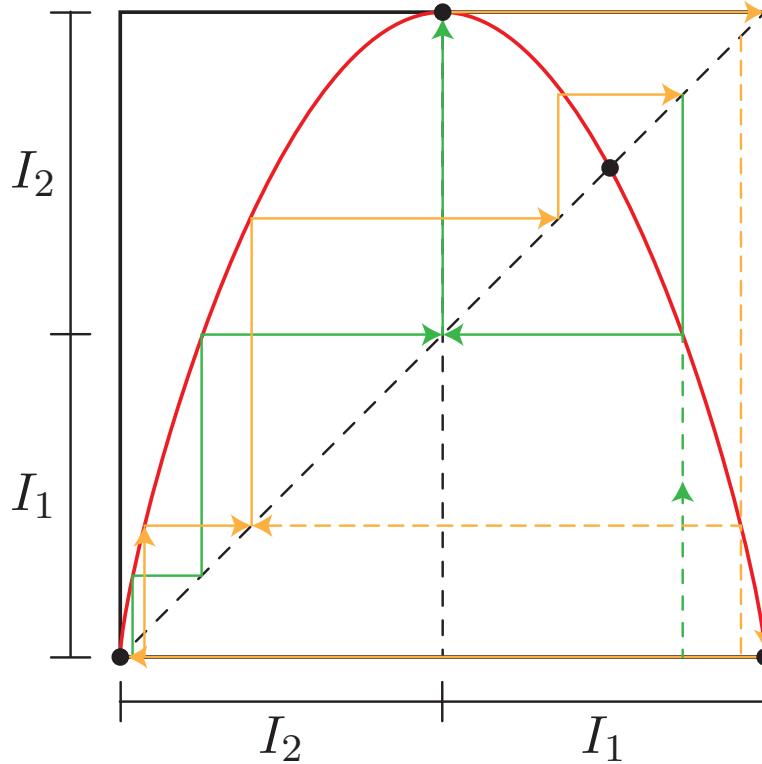


Figure 7.20: The cobweb diagram for the logistic map with $a = 4$. The sets I_1 and I_2 are defined as $[0, \frac{1}{2}]$ and $[\frac{1}{2}, 1]$.

iterates, i.e. two chains of preimages of $\frac{1}{2}$. One of these branches (branch A) converges to 0, and the other one (branch B) converges to $\frac{3}{4}$ in backward time. Points on branch A thus form a homoclinic orbit to zero, and points on branch B form a heteroclinic orbit between $\frac{3}{4}$ and 0. Since a point may follow branch A for a finite number of backward iterations before switching to branch B and vice versa, we actually have a countably infinite number of homo- and heteroclinic orbits. From Fig. 7.20, the dynamics appears complicated, with no stable fixed points as we have seen from the linearization.

To better understand the dynamics we can encode the itinerary of a point $x \in I = [0, 1]$ by if x is in I_1 or I_2 . We get a one-sided itinerary, because f is not invertible. For example if $x \in I_{s_0}$, $f(x) \in I_{s_1}$, $f^2(x) \in I_{s_2}$ then we would encode x as $.s_0 s_1 s_2 \dots$ for $s_i \in \{1, 2\}$.

Note here that this symbolic encoding is not well defined for every point in I , as all preimages of $x = \frac{1}{2}$ have multiple encodings, since their iterates eventually reach $I_1 \cap I_2 = \frac{1}{2}$. To address this, instead we hope to show that each encoding corresponds to a unique trajectory. To this

end define $\Psi : \Sigma \rightarrow I$ for Σ the set of semi-infinite binary symbols, by

$$\Psi : .s_0 s_1 s_2 \mapsto x = I_{s_0} \cap f^{-1}(I_{s_1}) \cap f^{-2}(I_{s_2}) \cap \dots \quad (7.70)$$

We now have to ask if this is well defined, i.e. if it yields a unique x value. Note here that for any subset $J \subset I$ which is closed and connected, we have that $I_{s_i} \cap f^{-1}(J)$ is closed, nonempty, and connected. This is due to the symmetry around $x = \frac{1}{2}$ and is shown visually in Fig. 7.21.

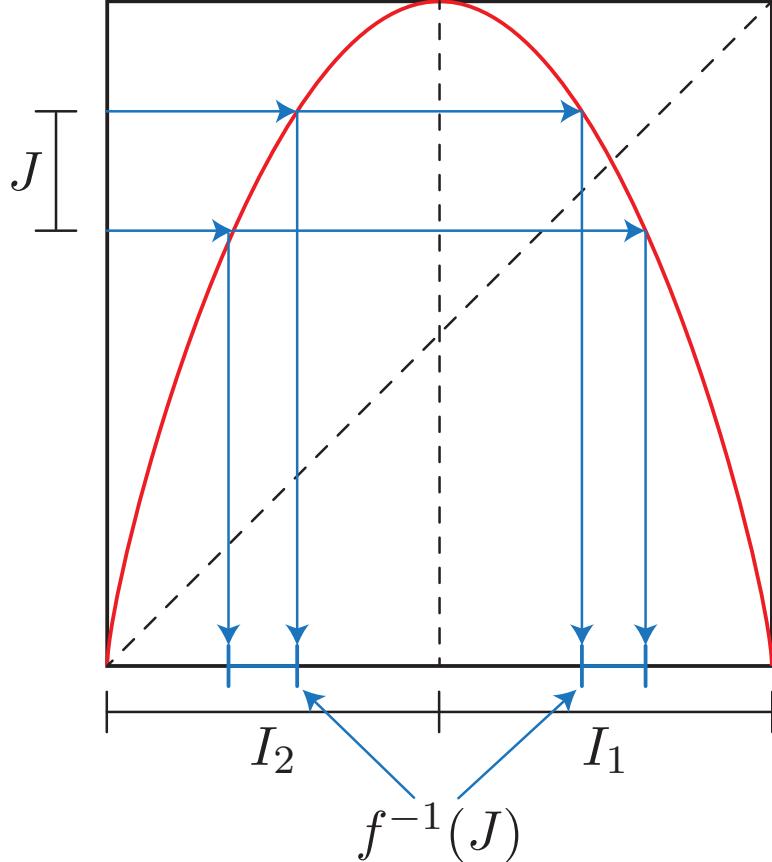


Figure 7.21: The preimage of a subset $J \subset I$, with its closed, nonempty, and connected intersections with I_1 and I_2 .

Remark 7.23. For any two sets A and B we have that

$$f^{-1}(A \cap B) = f^{-1}(A) \cap f^{-1}(B). \quad (7.71)$$

Further observe that for $I_{s_0 s_1} = I_{s_0} \cap f^{-1}(I_{s_1})$ is closed, nonempty, and connected for all s_0 and s_1 by the previous note. Continuing this we find that

$$I_{s_0 s_1 s_2} = I_{s_0} \cap f^{-1}(I_{s_1}) \cap f^{-2}(I_{s_2}) = I_{s_0} \cap f^{-1}(I_{s_1} \cap f^{-1}(I_{s_2})) \quad (7.72)$$

is closed, nonempty, connected, and contained in $I_{s_0 s_1}$. This line of argument can be continued to show for any $n \in \mathbb{N}$ the set $I_{s_0 \dots s_n}$ is closed, nonempty, connected, and contained in $I_{s_0 \dots s_{n-1}}$. Thus we have a nested sequence of closed, nonempty, and connected sets. By Cantor's theorem a nonempty limit exists, which is now not a Cantor set in this case, because it is connected. It is also possible to show that $\lim_{n \rightarrow \infty} |I_{s_0 \dots s_n}| = 0$, where $|\cdot|$ denotes the length of a interval. Now we can conclude that the limit set contains a single point x , therefore as $n \rightarrow \infty$ the function $\Psi(s) = x$ is well-defined. It is also true that every element $x \in I$ can be written in this form for some s , i.e. Ψ is a surjection.

Additionally with the metric

$$d(s, s') = \sum_{i=0}^{\infty} \frac{|s_i - s'_i|}{2^i} \quad (7.73)$$

the function Ψ is continuous.

By definition, a single iteration of the map f is equivalent to a shift to the left on the corresponding symbols in Σ , i.e.

$$f : x = .s_0 s_1 s_2 \dots \mapsto .s_1 s_2 \dots = f(x). \quad (7.74)$$

Therefore we have the following commuting diagram

$$\begin{array}{ccc} \Sigma & \xrightarrow{\sigma} & \Sigma \\ \downarrow \Psi & & \downarrow \Psi \\ I & \xrightarrow{f} & I \end{array} \quad (7.75)$$

Definition 7.12 (Topological semi-conjugacy). A map f is *topologically semi-conjugate* to σ if there exists a surjection Ψ such that

$$f \circ \Psi = \Psi \circ \sigma. \quad (7.76)$$

In comparison to the previous example, Ψ is not invertible (although it is well-defined and continuous). The invertibility of Ψ only fails on a countable set of points $\{f^{-n}(\frac{1}{2})\}_{n=1}^{\infty}$. Removing this countable set of points, a homeomorphism between Σ and $\tilde{I} = I \setminus f^{-n}(1/2)_{n=1}^{\infty}$ can be constructed. Therefore we have that the logistic map is chaotic on \tilde{I} .

7.4 Generalization of chaotic dynamics

We begin by defining a general symbol space.

Definition 7.13. Denote by Σ^N the symbol space defined as the space of doubly infinite sequences of N symbols (for $N \in \mathbb{N}$), i.e.

$$\boxed{\Sigma^N = \{s = \dots s_{-2}s_{-1} \cdot s_0s_1s_2 \dots : s_i \in \{0, 1, \dots, N-1\}\}.} \quad (7.77)$$

We may constrain this symbol space by constraining which sequences are admissible. This can be done by using a transition matrix $A \in \mathbb{R}^{N \times N}$ with entries $A_{ij} \in \{0, 1\}$. Then the constrained symbol space is given by

$$\Sigma_A^N = \{s \in \Sigma^N : A_{s_i s_{i+1}} \neq 0, \forall i\}. \quad (7.78)$$

In the constrained space, consecutive symbols s_i and s_{i+1} must correspond to a 1 in the transition matrix A . In other words if the symbol $s_i = k$ then the symbol s_{i+1} must be equal to j such that $A_{kj} = 1$.

Example 7.6 (Constrained symbol space). To demonstrate how a transition matrix can constrain the symbol space, lets start by examining Σ_A^2 for

$$A = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}. \quad (7.79)$$

By examining the entries of the transition matrix A we find which strings are admissible

$$\begin{aligned} A_{11} = 1 &\implies \dots 00\dots \text{ admissible} \\ A_{12} = 0 &\implies \dots 01\dots \text{ inadmissible} \\ A_{21} = 0 &\implies \dots 10\dots \text{ inadmissible} \\ A_{22} = 1 &\implies \dots 11\dots \text{ admissible.} \end{aligned}$$

Therefore the set Σ_A^2 is comprised only of two elements $\{\bar{0}\bar{0}, \bar{1}\bar{1}\}$. As this space is very simple and only contains fixed points of σ , no chaos can occur.

Definition 7.14. The map

$$\sigma : \Sigma_A^N \rightarrow \Sigma_A^N; \quad \dots s_{-1} \cdot s_0 \dots \mapsto \dots s_{-1} s_0 \cdot s_1 s_2 \dots, \quad (7.80)$$

is called a *Bernoulli subshift of finite type* (i.e. $N < \infty$) with transition matrix A .

Before more examination of chaos on such constrained spaces, we need to recall a definition from linear algebra.

Definition 7.15. A matrix A is called *irreducible* if there exists a $k \in \mathbb{N}$ such that $(A^k)_{ij} \neq 0$ for every $i, j \in \{1, \dots, N\}$.

Exercise 7.5. Let A denote the transition matrix for a sub-shift $\sigma: \Sigma_A^N \mapsto \Sigma_A^N$ of finite type on N symbols.

- (i) Show that the number of fixed points of σ is equal to $\text{tr}(A)$.
- (ii) Show that the total number of admissible k -periodic points (i.e., k -periodic points whose minimal period may be less than k) is equal to $\text{tr}(A^k)$.

Theorem 7.24. Assume that A is irreducible, then

- (i) Σ_A^N is a Cantor set; and
- (ii) $\sigma: \Sigma_A^N \rightarrow \Sigma_A^N$ is a chaotic map.

Note that in the previous example with $A = I_2$, the transition is not irreducible, which is why chaos is not guaranteed.

The role of irreducibility is critical for the theorem to work. Say for instance that the required k such that each entry of A^k is nonzero is $k = 3$. Then we know that $(A^3)_{ij} = \sum_{l,m=1}^N a_{il}a_{lm}a_{mj} \neq 0$ for any pair i, j . Therefore there exist l, m such that a_{il} , a_{lm} , and a_{mj} which are not equal to 0. Hence there exists an admissible ultimate transition from state i to state j for all pairs i, j , if the symbol sequence is viewed as an itinerary. Furthermore, different irreducible transition matrices generate different types of chaos.

Example 7.7 (Chaos in the Newton-Raphson iteration). Newton-Raphson iteration is a process to find the roots (zeros) of a nonlinear function f . The iteration is as follows

- (i) Start with the seed point x_0 ;
- (ii) Define $x_1 = x_0 - \frac{f(x_0)}{f'(x_0)} = N(x_0)$;
- (iii) Repeat (ii) as $x_{n+1} = x_n - \frac{f(x_n)}{f'(x_n)}$, until convergence, i. e. until $x_{n+1} = x_n$.

In practice this approach is often truncated when $x_{n+1} \approx x_n$. The roots of f coincide with the fixed points of the 1-dimensional dynamical system $x_{n+1} = N(x_n)$. If such a fixed point \bar{x} is stable and x_0 is in its domain of attraction, then $\lim_{i \rightarrow \infty} N^i(x_0) = \bar{x}$. The iteration process is sketched in Fig. 7.22.

Note that if $N(\bar{x}) = \bar{x}$ then we have

$$N'(\bar{x}) = 1 - \frac{f'(\bar{x})^2 - f(\bar{x})f''(\bar{x})}{f'(\bar{x})^2} = 0. \quad (7.81)$$

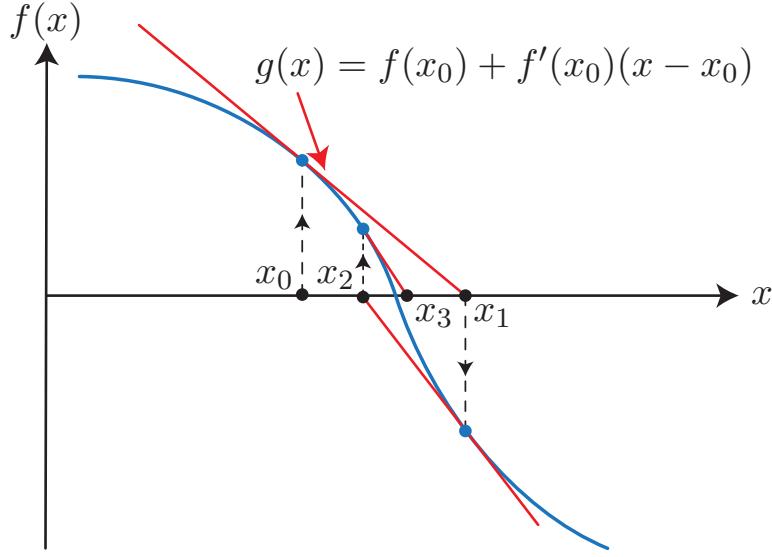


Figure 7.22: An illustration of the iteration of the Newton-Raphson process on the function f (blue). The red designates the function g , when given the point x_0 the next iterate x_1 fulfills the property $g(x_1) = 0$.

The latter equality is due to $f(\bar{x}) = 0$. Stability of \bar{x}_0 clearly follows, because $|N'(\bar{x}_0)| < 1$. Additionally, since $N'(\bar{x}_0) = 0$, we call the fixed point *super-stable*.

Specifically, once we are near the optimum \bar{x} , we have

$$N(x_n) = \underbrace{N(\bar{x})}_{=\bar{x}} + \underbrace{N'(\bar{x})(x_n - \bar{x})}_{=0} + \frac{1}{2}N''(\bar{x})(x_n - \bar{x})^2 + \mathcal{O}(|x_n - \bar{x}|^3). \quad (7.82)$$

This implies we have fast quadratic convergence, as

$$|N(x_n) - N(\bar{x})| \approx \frac{1}{2}|N''(\bar{x})||x_n - \bar{x}|^2. \quad (7.83)$$

Despite the fast convergence, erratic iterations may result if the initial point is not close enough to the root [Saari and Urenko, 1984]. In particular, consider a 4th order polynomial

$$f(x) = (x - x_0)(x - x_1)(x - x_2)(x - x_3); \quad x_0 < x_1 < x_2 < x_3. \quad (7.84)$$

This can be observed to lead to chaotic Newton-Raphson iterations. The proof will be sketched for this chaotic behaviour.

Retain the 4th order polynomial f from above. We then have $N(x_i) = x_i$ for $i = 0, 1, 2, 3$ as these are roots of f . Note then that we already know what the shapes of f and f' are, and these are illustrated in Fig. 7.23.

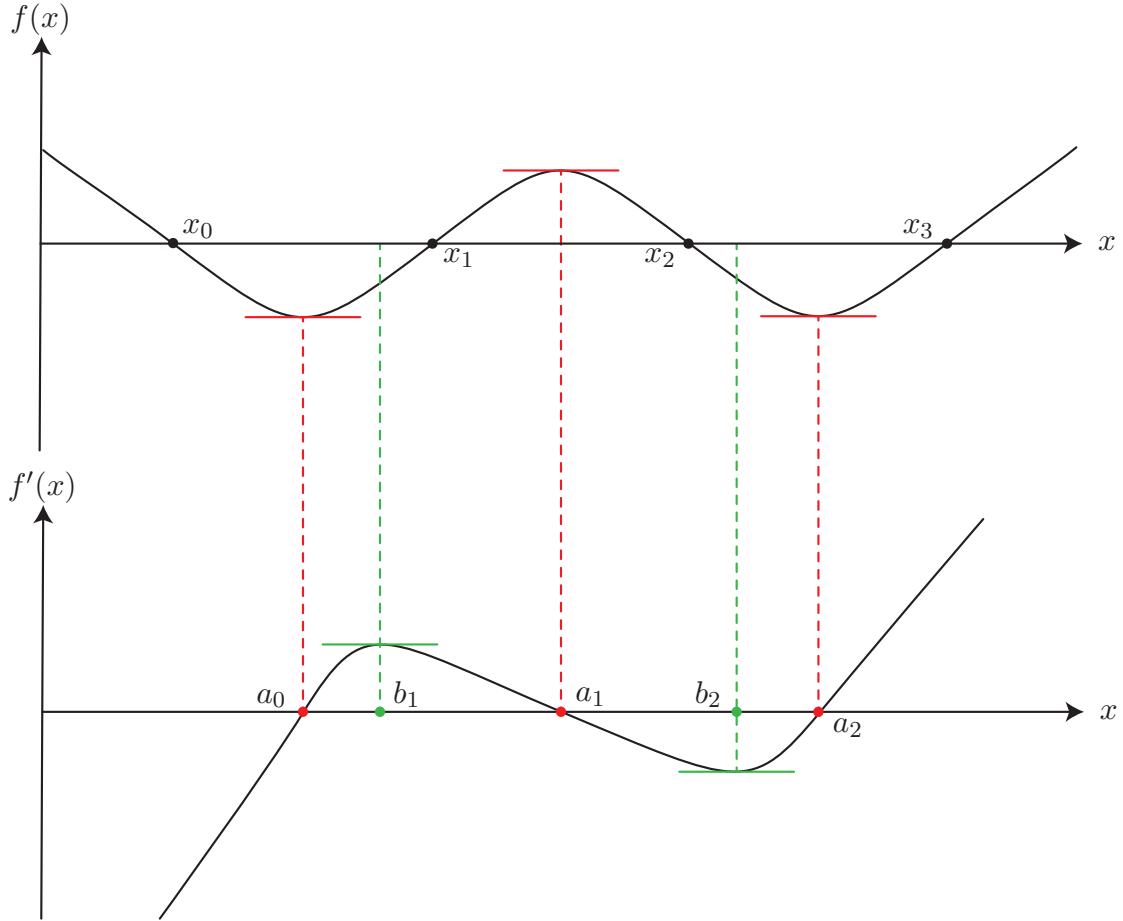


Figure 7.23: The shape of f (above) and f' (below). The points a_i are roots of f' and the points b_i are roots of f'' .

The points a_i denote the minima and maxima of $f(x)$ and b_i denote the minima and maxima of $f'(x)$ (See Fig. 7.23). As $f'(a_i) = 0$ we also have that $|N(a_i)| = \infty$, also $f''(b_i) = 0$ implies that $N'(b_i) = 0$. Thus via our knowledge of the shape of f , we have been able to derive information about N . If we assume that $b_1 \in (a_0, x_1)$ and $b_2 \in (x_2, a_2)$, then we get the graph of N as shown in Fig. 7.24.

Next, we seek non-convergent iterations in $[a_0, a_1] \cup [a_1, a_2]$. We begin by finding the preimages of a_0 (called $y_{0,i}$) and a_2 (called $y_{2,i}$) this can be done geometrically by starting from the point a_k , going down to the diagonal, and then searching horizontally for where the graph of N crosses the level a_k ; going down from this point to the x -axis yields a preimage. This is shown in Fig. 7.25. Using these points define the following sets

$$I_1 = [y_{2,1}, z_1^+]; \quad I_2 = [z_1^-, y_{0,1}]; \quad I_3 = [y_{2,2}, z_2^+]; \quad I_4 = [z_2^-, y_{0,2}]. \quad (7.85)$$

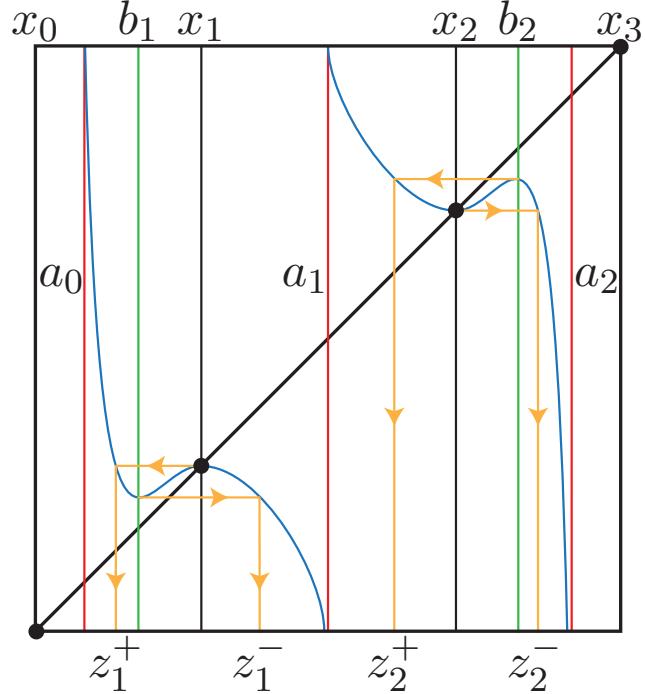


Figure 7.24: The graph of N , the points z_i^\pm are lower bounds on the domain of attraction as they surely converge to the fixed points. The are chosen by starting at a local extrema and working backwards.

These sets are illustrated in Fig. 7.25 as well.

Noting that $b_2 > N(b_2)$ and $b_1 < N(b_1)$ observe that the following hold

$$N(I_1) = [x_1, a_2] \supset I_2 \cup I_3 \cup I_4; \quad N(I_3) = [N(b_2), a_2] \supset [b_2, a_2] \supset I_4; \quad (7.86a)$$

$$N(I_2) = [a_0, N(b_1)] \supset [a_0, b_1] \supset I_1; \quad N(I_4) = [a_0, x_2] \supset I_1 \cup I_2 \cup I_3. \quad (7.86b)$$

Therefore we have that the possible transitions from I_i to I_j are described by the transition matrix

$$A = \begin{pmatrix} 0 & 1 & 1 & 1 \\ 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 1 & 1 & 1 & 0 \end{pmatrix}. \quad (7.87)$$

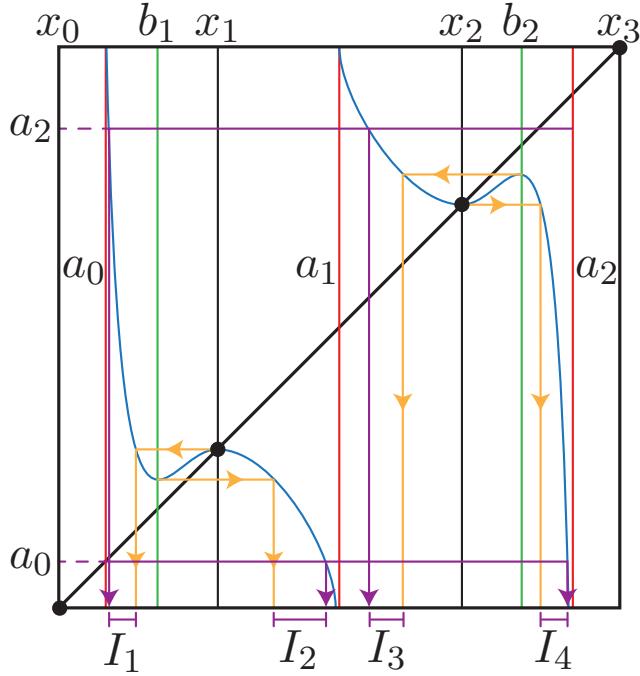


Figure 7.25: The preimages of the values a_0 and a_2 , along with the sets I_j for $j = 1, 2, 3, 4$.

This transition matrix is in fact irreducible with $k = 3$

$$A^2 = \begin{pmatrix} 2 & 1 & 1 & 1 \\ 0 & 1 & 1 & 1 \\ 1 & 1 & 1 & 0 \\ 1 & 1 & 1 & 2 \end{pmatrix}; \quad A^3 = \begin{pmatrix} 2 & 3 & 3 & 3 \\ 2 & 1 & 1 & 1 \\ 1 & 1 & 1 & 2 \\ 3 & 3 & 3 & 2 \end{pmatrix}. \quad (7.88)$$

We are not done yet, as we still need to establish a unique encoding. Note that if $a_{ij}^3 \neq 0$ then the set $I_i \cap f^{-1}(I_j) \neq \emptyset$ is closed for $i = 1, 2, 3, 4$. In light of this, with $\tilde{N} = N^3$, define

$$I_{s_0 s_1 \dots s_n} = I_{s_0} \cap \tilde{N}^{-1}(I_{s_1}) \cap \tilde{N}^{-2}(I_{s_2}) \cap \dots \cap \tilde{N}^{-n}(I_{s_n}) \quad (7.89a)$$

$$= I_{s_0} \cap \tilde{N}^{-1} \left(I_{s_1} \cap \tilde{N}^{-1} \left(I_{s_2} \cap \dots \underbrace{\tilde{N}^{-1} \left(I_{s_{n-1}} \cap \tilde{N}^{-1}(I_{s_n}) \right)}_{\neq \emptyset, \text{closed}} \right) \right). \quad (7.89b)$$

The condition $I_{s_{n-1}} \cap \tilde{N}^{-1}(I_{s_n}) \neq \emptyset$ and is closed is true as each entry $a_{ij}^3 \neq 0$. Therefore we have a nested sequence of nonempty, closed sets and by Cantor's Theorem their intersection yields a nonempty, closed set. By our previous result we have that $\Lambda = \bigcup_{s \in \Sigma_A^N} I_s$ is a Cantor

set. In this case $N = 4$ and A is as above. We then have the encoder function $\phi : \Lambda \rightarrow \Sigma_A^4$ and the chaotic function $\sigma : \Sigma_A^4 \rightarrow \Sigma_A^4$. These yield the commuting diagram

$$\begin{array}{ccc} \Sigma_A^4 & \xrightarrow{\sigma} & \Sigma_A^4 \\ \phi \uparrow & & \uparrow \phi \\ \Lambda & \xrightarrow{N} & \Lambda \end{array} . \quad (7.90)$$

Thus we have that $N = \phi^{-1} \circ \sigma \circ \phi$, which implies that N is a chaotic map. Hence, the Newton-Raphson iteration has an invariant Cantor set, on which the iteration is equivalent to a Bernoulli subshift of finite type on 4 symbols.

Notebook 7.1. *Chaos in the Newton-Raphson iteration.*

Chapter 8

Introduction to Hamiltonian Systems

A canonical Hamiltonian system is given by two state variables $p, q \in \mathbb{R}^n$ (i.e. the entire system is even dimensional), a *Hamiltonian* function $H(p, q, t) \in \mathcal{C}^1(\mathbb{R}^{2n} \times \mathbb{R})$ which maps $\mathbb{R}^n \times \mathbb{R}^n \times \mathbb{R} \rightarrow \mathbb{R}$. In general q denotes the generalized coordinates (position) and p the generalized momenta. The dynamics on the system is given by

$$\begin{cases} \dot{q} = \frac{\partial H(q, p, t)}{\partial p} \\ \dot{p} = -\frac{\partial H(q, p, t)}{\partial q}. \end{cases} \quad (8.1)$$

This is represented in shorthand notation by using the identity matrix, denoted by $I_n \in \mathbb{R}^{n \times n}$ and the symplectic matrix

$$J = \begin{pmatrix} 0 & I_n \\ -I_n & 0 \end{pmatrix}. \quad (8.2)$$

Denote the differential with respect to the state variables $x = \begin{pmatrix} p \\ q \end{pmatrix} \in \mathbb{R}^{2n}$ by

$$D_x = D = \begin{pmatrix} \frac{\partial}{\partial q} \\ \frac{\partial}{\partial p} \end{pmatrix}. \quad (8.3)$$

Then the shorthand notation for the system is

$$\boxed{\dot{x} = JD_x H(x, t).} \quad (8.4)$$

Example 8.1 (Holonomic mechanical systems are Hamiltonian). Assume that all external forces are derived from a potential, i.e. $F(q, t) = -\frac{\partial V(q, t)}{\partial q}$ where q denotes the position. Further

assume that all constraints are geometric, i.e. $f(q, t) = 0$. We have n degrees of freedom as $q \in \mathbb{R}^n$ and $\dot{q} \in \mathbb{R}^n$. The kinetic energy is given by $T(q, \dot{q})$ and the potential energy by $V(q, t)$. We then use the principle of least action (also referred to as Hamilton's principle) to derive the equations of motion. This can be shown to be equivalent to Newton's second law. This principle states that for the *Lagrangian* $L(q, \dot{q}, t) = T(q, \dot{q}) - V(q, t)$ we have that the variation of the action S is 0

$$\delta S = \delta \int_{t_0}^{t_1} L(q(s), \dot{q}(s), s) ds = 0. \quad (8.5)$$

This implies the *Euler-Lagrange equations*

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{q}} - \frac{\partial L}{\partial q} = 0. \quad (8.6)$$

Thus we have n , second-order ODEs.

Remark 8.1. In (8.5) the variation δS being equal to 0 stems from the principle of least action, as the action $\int L(q(s), \dot{q}(s), s) ds$ is then a minimum. The variational of a function δF (here $F[\gamma] = \int \gamma ds$) of a path (in our case $\gamma = (q(s), \dot{q}(s), s)$) corresponds to seeing how infinitesimal changes to the path change the output of the functional. These different, admissible paths are depicted in Fig. 8.1. Hamilton's principle states that among these different paths the one that will physically take place is the one along which the action S is minimal.

Now assume that L is a convex function of \dot{q} . Using this the *Legendre transformation* of L can be performed, which is essentially just a different way to look at the graph of $L(q, \cdot, t)$. This transformation gives us

$$H\left(q, \frac{\partial L}{\partial \dot{q}}, t\right) = \frac{\partial L}{\partial \dot{q}} \dot{q} - L(q, \dot{q}, t). \quad (8.7)$$

The geometric interpretation of this is shown in Fig. 8.2. At each velocity \dot{q} the tangent line of $L(q, \cdot, t)$ is taken and shifted such that the y -intercept is at 0. Then H is equal to the difference between this shifted line and the function $L(q, \dot{q}, t)$.

As L is convex, if we define the generalized momenta $p = \frac{\partial L}{\partial \dot{q}}$. There then exists a unique \dot{q} such that $\dot{q} = F(q, p, t)$, in fact the relationship between \dot{q} and p here is bijective. We then have

$$H(q, p, t) = pF(q, p, t) - L(q, F(q, p, t), t) \quad (8.8)$$

the Hamiltonian associated with L (or the Legendre transform of L).

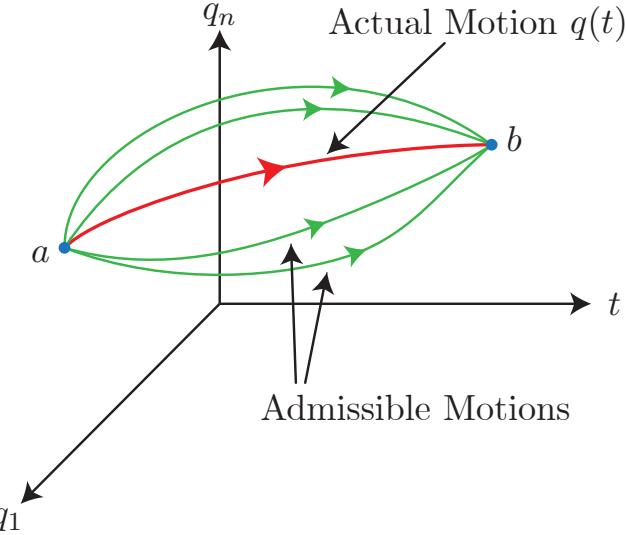


Figure 8.1: An illustration of various different paths (denoted here by $q(t)$). The red demarcates the actual motion which occurs in the system, whereas the green paths are each admissible motions which are consistent with the constraints.

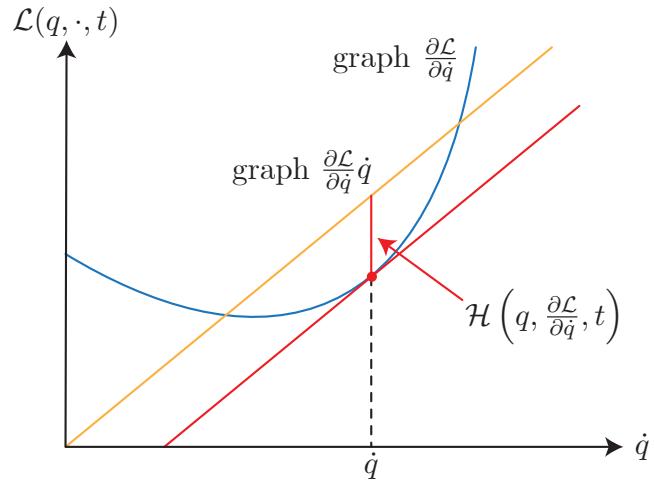


Figure 8.2: The Legendre transformation. The thin red linear function denotes the tangent of $L(q, \cdot, t)$ taken at \dot{q} , the yellow linear function is then the shifted version with a 0 y -intercept, and the thick red vertical line is the Legendre transformation H measuring the difference between this shifted line and $L(q, \cdot, t)$.

At this point note that the total differential of $H(q, p, t)$ is

$$dH = \frac{\partial H}{\partial q} dq + \frac{\partial H}{\partial p} dp + \frac{\partial H}{\partial t} dt \quad (8.9a)$$

$$= d(pF) - \underbrace{\frac{\partial L}{\partial q}}_{=\dot{p}} dq - \underbrace{\frac{\partial L}{\partial \dot{q}}}_{=p} - \frac{\partial L}{\partial t} dt \quad (8.9b)$$

$$= \underbrace{F dp}_{=\dot{q}} + pdF - \dot{p}dq - pdF - \frac{\partial L}{\partial t} dt \quad (8.9c)$$

$$= \dot{q}dp - \dot{p}dq - \frac{\partial L}{\partial t} dt. \quad (8.9d)$$

Comparing coefficients in the first and last equations then yields

$$\begin{cases} \dot{q} = \frac{\partial H}{\partial p} \\ \dot{p} = -\frac{\partial H}{\partial q} \end{cases}. \quad (8.10)$$

This first-order ODE is called the *Hamiltonian equations of motion*. The physical meaning of the Hamiltonian can be seen by substituting $L = T(q, \dot{q}) - V(q, t)$ into the definition 8.7. The kinetic energy only depends on \dot{q} through the kinetic energy,

$$\dot{q} \frac{\partial L}{\partial \dot{q}} = \dot{q} \frac{\partial T}{\partial \dot{q}}. \quad (8.11)$$

Since the kinetic energy is quadratic in the velocity, this simplifies to

$$\dot{q} \frac{\partial T}{\partial \dot{q}} = 2T(q, \dot{q}). \quad (8.12)$$

We then have

$$H = 2T - (T - V) = T(q, p) + V(q, t), \quad (8.13)$$

which is the full mechanical energy. Also note that we have $\frac{\partial H}{\partial t} = -\frac{\partial L}{\partial t}$.

We will now apply this result to a well known system.

Example 8.2 (The Hamiltonian of the pendulum). Recall that for the pendulum the position is given by $q = \varphi \in S^1$, with a string of length ℓ the vertical deflection of the pendulum is $\ell(1 - \cos(\varphi))$. The string is assumed massless and the mass at the end of the string is given by m .

We have the kinetic and potential energies

$$T = T(\dot{q}) = \frac{1}{2}m(\ell\dot{\varphi})^2 = \frac{1}{2}m\ell^2\dot{\varphi}^2; \quad V = V(q) = mgl(1 - \cos(\varphi)). \quad (8.14)$$

Therefore the Lagrangian is given by

$$L = T - V = \frac{1}{2}m\ell^2\dot{\varphi}^2 - mgl(1 - \cos(\varphi)). \quad (8.15)$$

To find the generalized momentum, we calculate the partial derivative of the Lagrangian with respect to \dot{q} , i.e.

$$p = \frac{\partial L}{\partial \dot{q}} = \frac{\partial L}{\partial \dot{\varphi}} = m\ell^2\dot{\varphi} \implies \dot{\varphi} = \frac{p}{m\ell^2}. \quad (8.16)$$

Now it is possible to calculate the Hamiltonian

$$H = p \frac{p}{m\ell^2} - \frac{1}{2} m\ell^2 \frac{p^2}{(m\ell^2)^2} + mg\ell(1 - \cos(\varphi)) = \underbrace{\frac{1}{2} \frac{p^2}{m\ell^2}}_T + \underbrace{mg\ell(1 - \cos(\varphi))}_V. \quad (8.17)$$

Hence for this Hamiltonian system we have the Hamiltonian equations of motion

$$\begin{cases} \dot{\varphi} = \frac{\partial H}{\partial p} = \frac{p}{m\ell^2} \\ \dot{p} = -\frac{\partial H}{\partial q} = -mg\ell \sin(\varphi). \end{cases} \quad (8.18)$$

Exercise 8.1. Derive the Hamiltonian equations of motion for a the coupled pendulums shown in Fig. 8.1. (The two point masses m are placed at the tips of two massless rods of length L . Both joints are frictionless; the constant of gravity is g .)

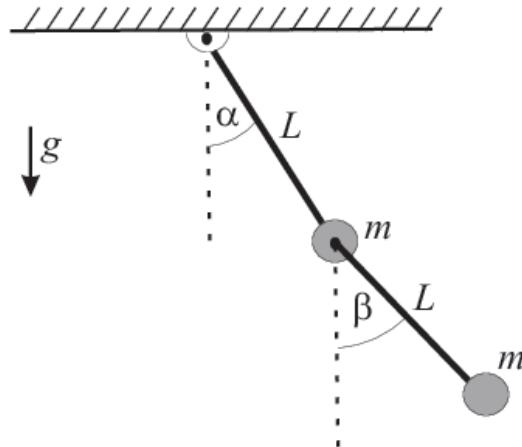


Figure 8.3: Coupled system of two pendulums.

Example 8.3 (2-dimensional incompressible fluids are Hamiltonian). We begin on a domain \mathcal{D} contained on the plane which is simply connected, i.e. it has no holes. For each position $x \in \mathcal{D}$ we have the velocity field

$$V(x, t) = \begin{pmatrix} u(x, y, t) \\ v(x, y, t) \end{pmatrix}, \quad (8.19)$$

i.e. for each position we have

$$\begin{pmatrix} \dot{x} \\ \dot{y} \end{pmatrix} = \begin{pmatrix} u \\ v \end{pmatrix}. \quad (8.20)$$

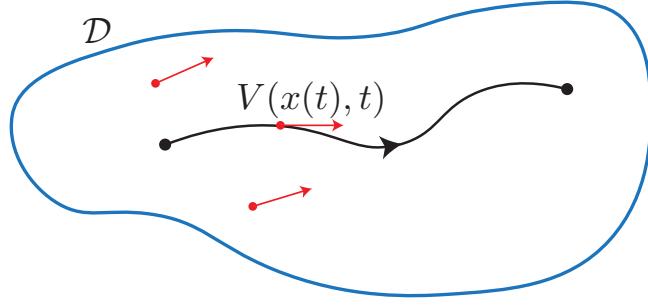


Figure 8.4: The simply connected domain \mathcal{D} with the velocity field V .

This is depicted in 8.4.

The incompressibility condition is given by

$$\nabla \cdot V = \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0 \quad \forall x \in \mathcal{D} \subset \mathbb{R}^2. \quad (8.21)$$

Now rewrite (8.21) in terms of a three dimensional vector field as

$$\nabla \times \begin{pmatrix} -v \\ u \\ 0 \end{pmatrix} = 0 \Leftrightarrow \nabla \cdot V = 0. \quad (8.22)$$

The third component is calculated as $\frac{\partial u}{\partial x} - \frac{\partial(-v)}{\partial y} = \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0$. On $\mathcal{D} \times \mathbb{R}$ (which is simply connected in \mathbb{R}^3) the vector field $\begin{pmatrix} -v \\ u \\ 0 \end{pmatrix}$ has zero curl. Therefore there exists a potential of this vector field $\Psi : \mathcal{D} \times \mathbb{R} \rightarrow \mathbb{R}$ such that

$$\begin{pmatrix} -v \\ u \\ 0 \end{pmatrix} = \nabla \Psi = \begin{pmatrix} \partial_x \Psi \\ \partial_y \Psi \\ \partial_z \Psi \end{pmatrix}. \quad (8.23)$$

Because of the third equation Ψ has no z dependence and we can write our potential as $\Psi = \Psi(x, y, t)$. Therefore we find

$$\begin{cases} -\dot{y} = \frac{\partial \Psi}{\partial x} \\ \dot{x} = \frac{\partial \Psi}{\partial y} \end{cases} \Rightarrow \begin{cases} \dot{x} = \frac{\partial \Psi(x, y, t)}{\partial y} \\ \dot{y} = -\frac{\partial \Psi(x, y, t)}{\partial x} \end{cases}. \quad (8.24)$$

Thus we have shown that particle motion is Hamiltonian.

Exercise 8.2. Consider a two-dimensional steady compressible fluid flow with velocity field $v(x) = (u(x, y), v(x, y))$, where $x = (x, y)$. Assume that the flow conserves mass, i.e., its density function $\rho(x) > 0$ satisfies the equation of continuity. The latter equation, in its general form for unsteady flows, reads

$$\rho_t + \nabla \cdot (\rho v) = 0, \quad (8.25)$$

valid or general, unsteady flow. Show that the equation of fluid particle motion becomes a canonical Hamiltonian system after a rescaling of time.

Example 8.4 (Planar, autonomous dynamical systems with a conserved quantity are Hamiltonian). We begin with the dynamical system which has a first integral U (which is not constant), i.e.

$$\dot{x} = f(x); \quad x \in \mathbb{R}^2; \quad \exists U : \mathbb{R}^2 \rightarrow \mathbb{R} : \frac{dU(x(t))}{dt} = 0. \quad (8.26)$$

Note that this first integral property has further implications

$$\frac{dU(x(t))}{dt} = DU(x(t)) \cdot \dot{x}(t) = DU \cdot f|_{x(t)} = 0 \implies DU \perp f. \quad (8.27)$$

Therefore, we can take the orthogonal complement of DU and scale this to be equal to f , i.e.

$$f(x) = p(x) JDU(x) \implies \dot{x} = p(x) JDU(x). \quad (8.28)$$

Such a system is called a *generalized Hamiltonian system* with the scalar function $p(x)$. Here the symplectic matrix in 2 dimensions is given by the 90° rotation

$$J = \begin{pmatrix} 0 & I_1 \\ -I_1 & 0 \end{pmatrix} = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}. \quad (8.29)$$

If we assume $p(x) \neq 0$ for all $x \in \mathcal{D} \subset \mathbb{R}^2$ then we may define the trajectory-dependent rescaling of time

$$\tau = \int_0^t p(x(s)) ds. \quad (8.30)$$

Then, define the notation for the derivative with respect to τ as follows

$$\frac{d(\)}{dt} = \frac{d(\)}{d\tau} \underbrace{\frac{d\tau}{dt}}_{p(x(t))} = (\)' p. \quad (8.31)$$

Hence, using this time-transformation we have

$$\dot{x} = p JDU \implies x' p = p JDU \implies x' = JDU(x). \quad (8.32)$$

Therefore, we have a canonical Hamiltonian system. In fact, all planar, autonomous systems with a nontrivial first integral are Hamiltonian systems after a possible reparameterization of time.

Remark 8.2. The fact that the first integral was not trivial (i.e. non-constant) is critical, as otherwise DU would be constantly 0, as we must be able to phrase the right hand side f in terms of DU . Only in the case $f = 0$ would this then be possible.

Example 8.5 (Lotka-Volterra model (1925) for predator-prey interactions). For more information about this model see [Lotka, 1925, Volterra, 1926]. Let the size of the prey population be represented by $h(t)$ and the size of the predator population be denoted by $p(t)$. Each of these populations must be at least 0. Now for strictly positive constants a_1 , a_2 , b , and c we have the following dynamics

$$\begin{cases} \dot{h} = a_1 h(1 - bp) \\ \dot{p} = -a_2 p(1 - ch). \end{cases} \quad (8.33)$$

The phase portrait of this system is depicted in Fig. 8.5.

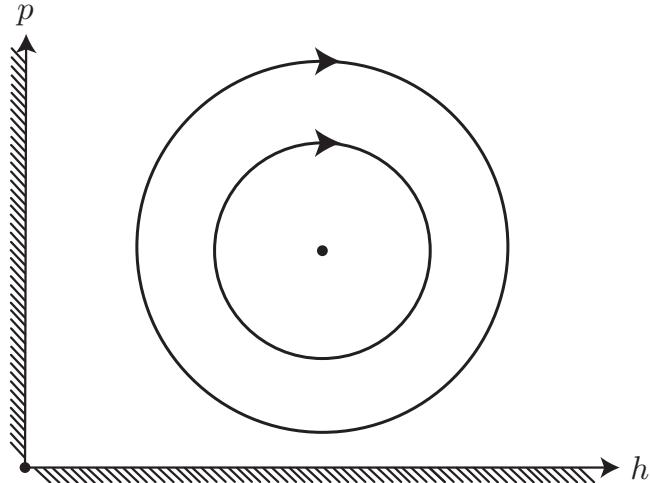


Figure 8.5: The phase portrait of the Lotka-Volterra dynamics, which has been used to explain the observed cycles in fish populations.

This system does in fact have a first integral, and thus it is Hamiltonian.

Now that we have seen multiple examples of Hamiltonian systems, we would like to better understand what properties they all share.

Exercise 8.3. Consider a dynamical system defined on the two-dimensional torus $\mathbb{T}^2 = \mathcal{S}^1 \times \mathcal{S}^1$. Such systems admit the general form

$$\dot{\phi}_1 = a(\phi_1, \phi_2) \quad (8.34)$$

$$\dot{\phi}_2 = b(\phi_1, \phi_2), \quad (8.35)$$

where $\phi_i \in \mathcal{S}^1$.

- (i) Show that a physical example of system (8.35) is found in the motion of two uncoupled linear undamped oscillators. Specifically, show that orbits of

$$\ddot{x} + x = 0 \quad (8.36)$$

$$\ddot{y} + y = 0, \quad (8.37)$$

are confined to two-dimensional invariant tori of the phase space.

- (ii) Assume that system (8.35) has no fixed point (which is the case in the oscillator example). Argue that (8.35) then cannot be Hamiltonian, even after a rescaling of time. (Hint: Use the fact that a continuous function defined on a compact set must have a minimum and a maximum).

Exercise 8.4. Show that for any dynamical system $\dot{q} = f(q, t)$, $q \in \mathbb{R}^n$, one can select a canonically conjugate variable $p \in \mathbb{R}^n$, such that the evolution of $(q(t), p(t))$ is governed by a Hamiltonian system. (Thus any type of dynamics can be viewed as a projection from a higher-dimensional Hamiltonian dynamical system.)

8.1 Properties of Hamiltonian systems

The setup going forward will be as follows for $n \geq 1$

$$\dot{x} = JDH(x, t); \quad x = \begin{pmatrix} q \\ p \end{pmatrix} \in \mathbb{R}^{2n}; \quad J = \begin{pmatrix} 0 & I_n \\ -I_n & 0 \end{pmatrix}. \quad (8.38)$$

Proposition 8.3 (Conservation of the Hamiltonian (energy)). If $\frac{\partial H}{\partial t} = 0$ (i.e. H is only a function of q and p) then $H(x(t))$ is constant on all trajectories. Therefore $H(x)$ is a first integral.

Proof. Use $J = -J^T$ to calculate

$$\frac{dH(x(t))}{dt} = DH(x(t)) \cdot JDH(x(t)) = \langle DH, JDH \rangle|_{x=x(t)} = 0. \quad (8.39)$$

□

The *energy surface* for some constant $E_0 = \{x \in \mathbb{R}^{2n} : H(x) = H_0\}$, for some constant H_0 , is an invariant surface for the system. By nature of this JDH must be tangent to E_0 . Assuming that E_0 contains no fixed points of (8.38) is equivalent to $DH(x_0) \neq 0$ for all x_0 in E_0 . Hence, by the Implicit Function Theorem, the solution of $H(x) - H_0 = 0$ can be smoothly continued from x_0 . Indeed, we have

$$DH(x_0) \neq 0 \in \mathbb{R}^{2n} \implies \exists i : \left. \frac{\partial H}{\partial x^i} \right|_{x=x_0} \neq 0. \quad (8.40)$$

Then we have $x^i = F(x^1, \dots, x^{i-1}, x^{i+1}, \dots, x^{2n})$ for $F \in \mathcal{C}^1$, because H is smooth. This implies that E_0 is a smooth $2n - 1$ -dimensional graph (i.e. a smooth hypersurface or codimension-one surface).

Example 8.6 (Two degree of freedom nonlinear oscillator). Between two fixed sides, three nonlinear springs hold two masses. Each spring has a potential given by V_i for $i = 1, 2, 3$ and the masses are given by m_1 and m_2 . The displacement of each mass is given by q_1 and q_2 respectively. The potentials have global minima at $x = (q_1, q_2)^T = 0$, and are otherwise strictly monotone. The system and example potential are illustrated in Fig. 8.6.

For this system the specific conditions on the potentials are

$$V_i(0) = 0; \quad V'_i(0) = 0; \quad V''_i(0) > 0; \quad \forall x \neq 0 : V''_i(x) \neq 0. \quad (8.41)$$

The Lagrangian is then

$$L = T - V = \frac{1}{2}m_1\dot{q}_1^2 + \frac{1}{2}m_2\dot{q}_2^2 - [V_1(q_1) + V_2(q_2 - q_1) + V_3(-q_2)]. \quad (8.42)$$

From the Lagrangian the generalized momenta can be derived

$$p_i = \frac{\partial L}{\partial \dot{q}_i} = m_i \dot{q}_i \implies \dot{q}_i = \frac{p_i}{m_i} \quad i = 1, 2. \quad (8.43)$$

And with this the Hamiltonian can be calculated as

$$H = T + V = \frac{1}{2} \frac{p_1^2}{m_1} + \frac{1}{2} \frac{p_2^2}{m_2} + V_1(q_1) + V_2(q_2 - q_1) + V_3(-q_2). \quad (8.44)$$

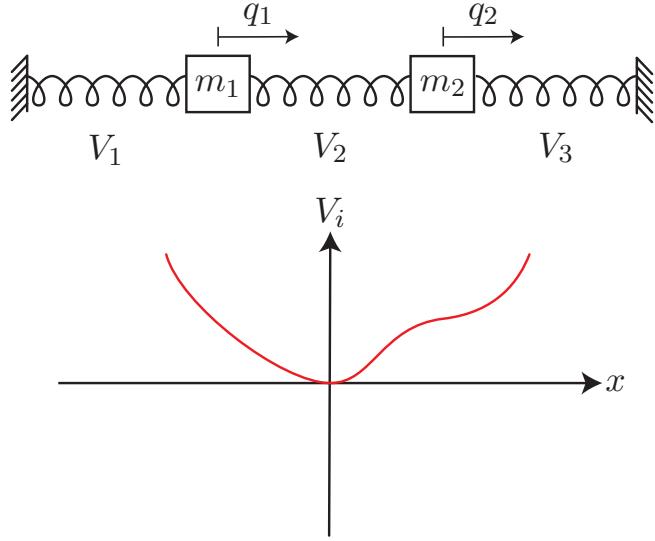


Figure 8.6: The two degree of freedom nonlinear oscillator and an example of a potential V_i .

Now recall that $\dot{x} = JDH(x)$, therefore the fixed point $x_0 = (q^0, p^0)^T$ can be found by identifying the roots of DH

$$DH(x_0) = 0 \Leftrightarrow \begin{pmatrix} V'_1(q_1^0) - V'_2(q_2^0 - q_1^0) \\ V'_2(q_2^0 - q_1^0) - V'_3(-q_2^0) \\ \frac{p_1^0}{m_1} \\ \frac{p_2^0}{m_2} \end{pmatrix}. \quad (8.45)$$

For instance the origin is a fixed point, i.e.

$$x_0 = \begin{pmatrix} q^0 \\ p^0 \end{pmatrix} = 0 \in \mathbb{R}^4. \quad (8.46)$$

Hence we have

$$p_1^0 = p_2^0 = 0; \quad V'_1(q_1^0) = V'_2(q_2^0 - q_1^0); \quad V'_2(q_2^0 - q_1^0) = V'_3(-q_2^0). \quad (8.47)$$

Assume that $q_1^0 > 0$ solves (8.47). Then we would have the following

$$V'_1(q_1^0) = V'_2(q_2^0 - q_1^0) > 0 \implies q_2^0 > q_1^0 > 0, \quad (8.48)$$

due to the conditions set out in (8.41). However we also have that

$$V'_3(-q_2^0) = V'_2(q_2^0 - q_1^0) > 0 \implies -q_2^0 > 0. \quad (8.49)$$

These two inequalities are clearly in contradiction. A similar argument shows that $q_1^0 < 0$ also leads to a contradiction. Therefore we find that at all equilibria $q_1^0 = 0$ and by symmetry $q_2^0 = 0$ as well. Thus the only fixed point is given by $q_1^0 = q_2^0 = 0$ and $p_1^0 = p_2^0 = 0$. We can conclude that all nonzero energy surfaces are codimension-one (3-dimensional) invariant manifolds (i.e. smooth surfaces).

Now it is possible to analyze the stability of this fixed point. To use Lyapunov's direct method, the Hamiltonian is a good candidate for a Lyapunov function, because $H(0) = 0$, $DH(0) = 0 \in \mathbb{R}^4$, and $\frac{dH(x(t))}{dt} = 0$. We still need to find out if H is cup shaped. First calculate the Hessian

$$D^2H = \left(\begin{array}{c|cc} D^2V & 0_{2 \times 2} \\ \hline 0_{2 \times 2} & \begin{matrix} \frac{1}{m_1} & 0 \\ 0 & \frac{1}{m_2} \end{matrix} \end{array} \right). \quad (8.50)$$

Next, we evaluate if each of the diagonal matrices are positive definite. The lower right matrix is clearly positive definite as it is a diagonal matrix with strictly positive entries. For the upper left, a calculation is required

$$D^2V|_0 = \left(\begin{matrix} V_1'' + V_2'' & -V_2'' \\ -V_2'' & V_2'' + V_3'' \end{matrix} \right) \Big|_{x=0 \in \mathbb{R}^4}. \quad (8.51)$$

Due to the fact that $V_1'' + V_2''|_0 > 0$ and $\det(D^2V|_0) > 0$ we have that $D^2V|_0$ is positive definite by Sylvester's theorem. Hence D^2H is positive definite, i.e. cup shaped. Now we have shown that $H(x)$ is a Lyapunov function, and therefore guarantees that the origin is nonlinearly stable irrespective of the actual form of the potentials.

From this we wonder what else can be said about the global phase space geometry, which leads us to the Morse lemma.

Lemma 8.4 (Morse). *Let x_0 be a nondegenerate critical point for $f : \mathbb{R}^n \rightarrow \mathbb{R}$, i.e.*

$$Df(x_0) = 0, \quad \det(D^2f(x_0)) \neq 0. \quad (8.52)$$

Let k be the number of negative eigenvalues of $D^2f(x_0)$. Then there exists a local change of coordinates (diffeomorphism) $x \mapsto y$ near x_0 such that in the new coordinates

$$f(y) = f(x_0) - y_1^2 - \dots - y_k^2 + y_{k+1}^2 + \dots + y_n^2. \quad (8.53)$$

By the Morse lemma, there exists a change of coordinates $x \mapsto y$ such that $H(y) = y_1^2 + y_2^2 + y_3^2 + y_4^2$ that is suitable near the origin. Therefore the energy surface $E_{h_0} = \{y \in \mathbb{R}^4 : H(y) = h_0\}$ is equal to $\mathcal{S}^3 \subset \mathbb{R}^4$, where \mathcal{S}^3 is the 3-dimensional hyper-sphere. This means in the original x -coordinates, all low-energy level sets of $H(x)$ are diffeomorphic to 3-dimensional spheres.

It is possible to draw on other results from differential geometry, namely the foliation theorem [Milnor, 1965]. This tells us that level surfaces of a scalar function can only change their topology (i.e. become no longer diffeomorphic to each other) through level surfaces that contain a critical point of the function (Df must vanish somewhere on the level set).

We can see this in the example of the pendulum equation, whose Hamiltonian we previously derived. The stable fixed point is a nondegenerate critical point of the Hamiltonian (as all other points around it have more energy) and by the Morse lemma, after a change of coordinates, level surfaces can be parametrized as a sum of quadratic terms. The loops representing periodic oscillation are also the level sets for the respective energy level and are all diffeomorphic to each other (they are all diffeomorphic to S^1) until the level set which intersects the unstable fixed points. This transition is depicted in Fig. 8.7.

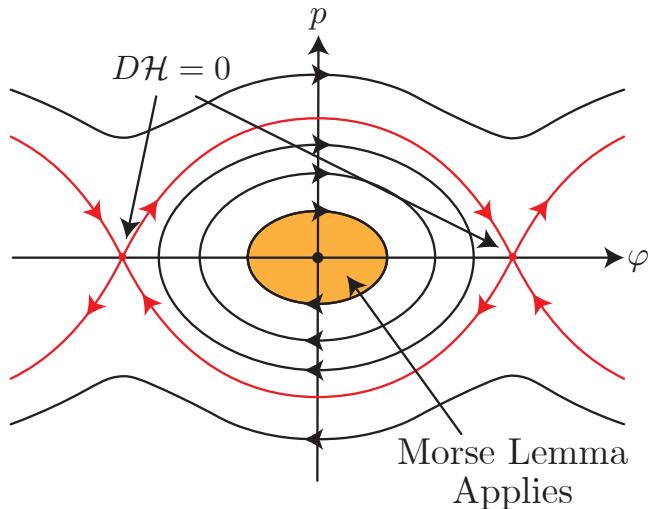


Figure 8.7: The effect of the foliation theorem. Around the stable fixed point where the Morse lemma applies is shaded orange, and the level sets containing critical points of H .

Regarding the two degree of freedom oscillator example previously handled, all surfaces are diffeomorphic to S^3 as $x = 0$ is the only critical point of H .

8.2 Volumes under Hamiltonian flow maps

We would like to understand what happens with a volume of the phase space under the flow map of a Hamiltonian system. When we talk about the transformation of a volume $V(t_0)$ under the flow map, we mean to study the set $V(t)$ which results from applying the flow map to $V(t_0)$, i.e. $V(t) = F_{t_0}^t(V(t_0))$. The evolution of such a volume is shown in Fig. 8.8. However, before

continuing to study this, we must prove Liouville's theorem.

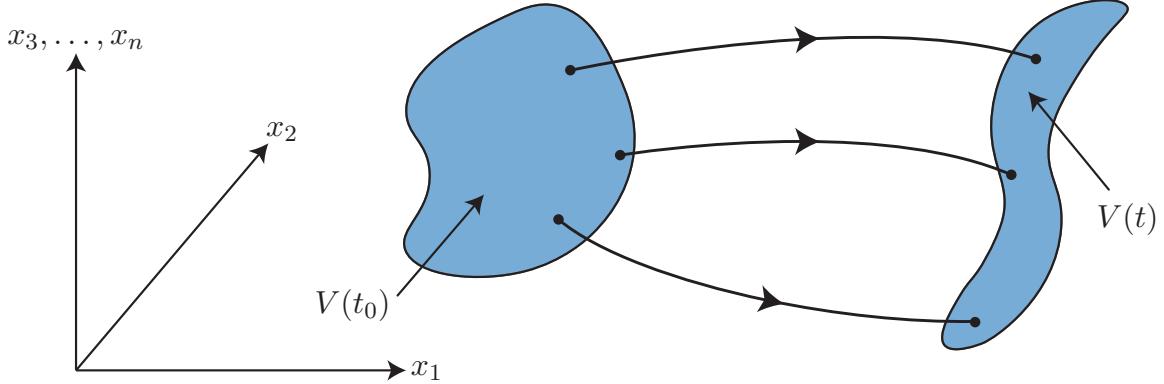


Figure 8.8: A transformed volume of the phase space under the flow map.

Theorem 8.5 (Liouville). *For a dynamical system*

$$\dot{x} = f(x, t); \quad x \in \mathbb{R}^n; \quad f \in \mathcal{C}^1, \quad (8.54)$$

we have that

$$\boxed{\frac{d}{dt} \text{vol}(V(t)) = \int_{V(t)} \nabla \cdot f(x, t) dV.} \quad (8.55)$$

Proof. Calculate the equation of variations for this system with $x(t) = x(t; t_0, x_0)$

$$\frac{d}{dt} \frac{\partial x}{\partial x_0} = Df(x(t), t) \frac{\partial x}{\partial x_0} \implies \frac{d}{dt} \left[\frac{\partial x}{\partial x_0} \right] = \underbrace{Df(x(t), t)}_{A(t)} \frac{\partial x}{\partial x_0}. \quad (8.56)$$

This measures how sensitive the resulting $x(t)$ is on the initial condition x_0 . We can generalize this to

$$\dot{X} = A(t)X; \quad A(t) = Df(x(t), t); \quad X, A \in \mathbb{R}^{n \times n}. \quad (8.57)$$

The only consistent initial condition for this is $\frac{\partial x}{\partial x_0} \Big|_{t=t_0} = X(t_0) = I_n \in \mathbb{R}^{n \times n}$. Recalling Abel's theorem, which applies to any linear system of ODE's we have

$$\det(\underbrace{X(t)}_{DF_{t_0}^t}) = \det(\underbrace{X(t_0)}_{I_n}) \cdot \exp \left(\int_{t_0}^t \text{Tr}[A(s)] ds \right). \quad (8.58)$$

Note that in present case $\frac{\partial x}{\partial x_0} = DF_{t_0}^t$. Finally, for a general dynamical system we have

$$\det(DF_{t_0}^t) = \exp\left(\int_{t_0}^t \nabla \cdot f(x(s), s) ds\right). \quad (8.59)$$

Using this we calculate

$$\frac{d}{dt} \text{vol}[V(t)] = \frac{d}{dt} \int_{V(t)} dV = \frac{d}{dt} \int_{V(t_0)} \det(DF_{t_0}^t(x_0)) dV_0 \quad (8.60a)$$

$$= \int_{V(t_0)} \frac{d}{dt} \det(DF_{t_0}^t(x_0)) dV_0 = \int_{V(t_0)} \nabla \cdot f(x(t), t) \underbrace{\det(DF_{t_0}^t)}_{=dV} dV_0 \quad (8.60b)$$

$$= \int_{V(t)} \nabla \cdot f(x(t), t) dV. \quad (8.60c)$$

Where on the last equation of the first line we used the change of coordinates $y = F_{t_0}^t(x_0)$, then on the last equation of the second line we used our previous result, and then finally switched back to the original coordinates for the final equality. \square

Proposition 8.6 (Conservation of phase space volume). *For any Hamiltonian system, the flow map $F_{t_0}^t : x_0 \mapsto x(t; t_0, x_0)$ is volume preserving for any $t_0, t \in \mathbb{R}$.*

Proof. For Hamiltonian systems we have that

$$\nabla \cdot (\underbrace{JDH(x, t)}_{f(x, t)}) = \text{Tr}[JD^2H(x, t)] = \text{Tr}\left[\begin{pmatrix} 0 & I_n \\ -I_n & 0 \end{pmatrix} \begin{pmatrix} D_{qq}^2 H & D_{qp}^2 H \\ D_{qp}^2 H & D_{pp}^2 H \end{pmatrix}\right] \quad (8.61a)$$

$$= \text{Tr}\left[\begin{pmatrix} D_{qp}^2 H & * \\ * & -D_{qp}^2 H \end{pmatrix}\right] = 0. \quad (8.61b)$$

Therefore by Liouville's theorem we have that $\text{vol}[V(t)]$ is constant which is equivalent to $\det(DF_{t_0}^t(x_0)) = 1$. This is called this *infitesimal (local) form*. \square

8.3 Consequences of area preservation

We now constrain ourselves to two dimensional, time-periodic Hamiltonian systems

$$\dot{x} = JDH(x, t); \quad x \in \mathbb{R}^2; \quad H(x, t) = H(x, t + T); \quad T > 0. \quad (8.62)$$

The Poincaré map is

$$P_{t_0} : x_0 \mapsto x(t_0 + T; t_0, x_0) = F_{t_0}^{t_0+T}(x_0). \quad (8.63)$$

$$H(x, t) = H(x_0) + \varepsilon H_1(x, t); \quad x = \begin{pmatrix} \varphi \\ p_\varphi \end{pmatrix}; \quad \dot{x} = JDH(x, t), \quad (8.64)$$

where H_1 and H are T -periodic in t .

It is possible to classify the different types of fixed points for P_{t_0} . Let x_0 be a fixed point of P_{t_0} , i.e. $P_{t_0}(x_0) = x_0$, then we have by area preservation that

$$\det(DP_{t_0}|_{x_0}) = \det(DF_{t_0}^{t_0+T}(x_0)) = 1. \quad (8.65)$$

Hence, the eigenvalues $\lambda_1, \lambda_2 \in \mathbb{C}$ of the linearized Poincaré map have the property

$$1 = \det(DP_{t_0}(x_0)) = \lambda_1 \lambda_2. \quad (8.66)$$

Furthermore, as the matrix $DP_{t_0}(x_0)$ is real and has even dimensions, the eigenvalues must occur in complex conjugate pairs. This implies that there are only a few restricted eigenvalue configurations which are possible.

- (i) **Robust (structurally stable) configurations** Saddles with $\lambda_1 \neq \lambda_2$ and $\lambda_1, \lambda_2 \in \mathbb{R}$ and centers with $\lambda_1 \neq \lambda_2$ with $\lambda_1 = e^{i\alpha} = \bar{\lambda}_2$ are the only robust eigenvalue configurations which are possible.
- (ii) **Non-robust configurations** The only other possibilities are $\lambda_1 = \lambda_2 = 1$ or $\lambda_1 = \lambda_2 = -1$.

The different eigenvalue configurations are shown in Fig. 8.9. As we can see the only robust fixed points of 2-dimensional Poincaré maps in Hamiltonian systems are saddles or centers (under linearization). The same conclusion holds, as a consequence, for any planar, autonomous Hamiltonian system (flow) $\dot{x} = JDH(x)$.

Next we examine what happens when we perturb homoclinic orbits. We recall the setting for homoclinic chaos in the Hamiltonian case with H_1 being T -periodic in t

$$\dot{x} = JD[H(x) + \varepsilon H_1(x, t)]. \quad (8.67)$$

We assumed that the unperturbed system had a homoclinic orbit and then asked if the unstable and stable manifolds of the perturbed system intersected using the Melnikov method, as sketched in Fig. 8.10.

Proposition 8.7. *The unstable manifold $W^U(x_\varepsilon)$ and stable manifold $W^S(x_\varepsilon)$ must intersect for $\varepsilon > 0$.*

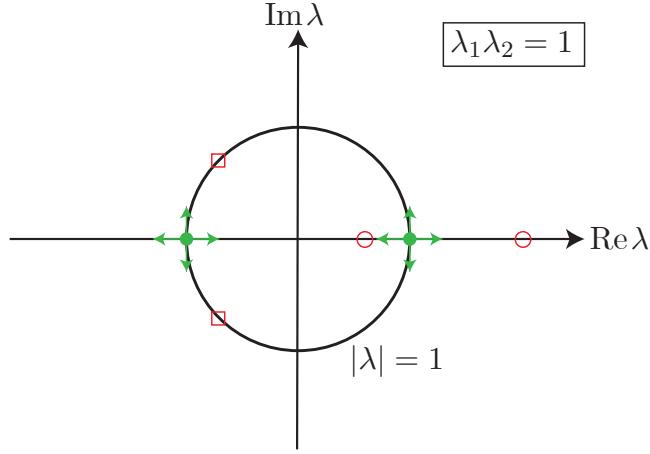


Figure 8.9: All possible eigenvalue configurations of the linearized Poincaré map.

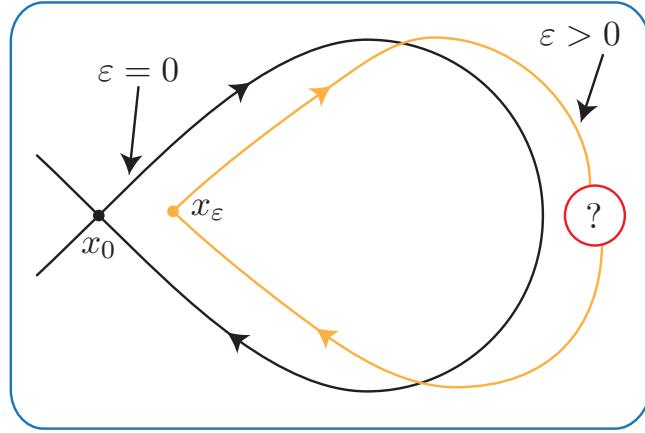


Figure 8.10: The unperturbed $\varepsilon = 0$ system with a homoclinic orbit in black and the perturbed $\varepsilon > 0$ system in yellow, the Melnikov method analyzes if the unstable and stable manifold intersect.

Proof. Assume contrarily that for $\varepsilon > 0$ the manifolds do not intersect. This leaves us with the image shown in Fig. 8.11.

As in Fig. 8.11, denote by A a point in $W^S(x_\varepsilon)$ and by B a point in $W^U(x_\varepsilon)$. Then call the region S the area enclosed by the curves of each manifold and the line AB . The region S under the Poincaré map now has a larger area than S , as can be seen by calculating:

$$\text{vol}[P_{t_0}(S) - S] = \text{vol}[\overline{ABP_{t_0}(A)P_{t_0}(B)}] \neq 0. \quad (8.68)$$

This violates the area preservation property of the Poincaré map, and therefore leads to a contradiction. \square

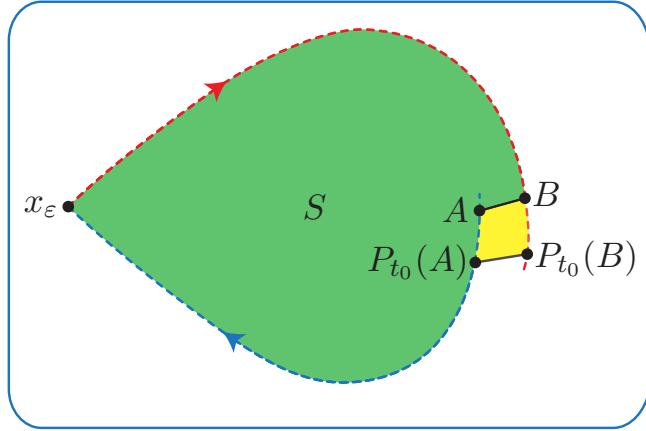


Figure 8.11: The image which we are left with if the two manifolds do not intersect.

There are four possible types of intersection geometries: transverse (chaotic), topologically transverse (chaotic), tangency, and equality (no chaos). These are all illustrated in Fig. 8.12. The tangency intersection lacks area preservation, and is also not possible.

An equality intersection can occur when the Hamiltonian is conserved even after perturbation. This implies that all trajectories continue to remain confined to the level sets of H .

As tangent intersections are not possible within the area preserving context, if $M(t_0) \neq 0$ then we must have homoclinic chaos. Thus in periodically forced Hamiltonian systems, chaos is generic! However, when we only have autonomous perturbations of the form $H(x) = H_0(x) + \varepsilon H_1(x)$, the equal case, in which the homoclinic orbit persists is generic.

Proposition 8.8 (Recurrence). *For an autonomous Hamiltonian system*

$$\dot{x} = JDH(x); \quad x \in \mathbb{R}^{2n}, \quad (8.69)$$

suppose there exists a subset S_0 of \mathbb{R}^{2n} which is compact, invariant, and has nonzero volume. Then for any open set $U \subset S_0$ and any time $T > 0$, there exists an $x_0 \in U$ and $t > T$ such that $x(t; t_0, x_0) \in U$.

This property is illustrated in Fig. 8.13.

Proof. The recurrence property follows from the more general Poincaré Recurrence Theorem, which can be applied here as the flow map $F_{t_0}^t : x_0 \mapsto x(t; t_0, x_0)$ is volume preserving and invertible and S_0 is compact, invariant, and has nonzero volume. \square

Proposition 8.9 (Poincaré Recurrence theorem). *For a compact domain B with nonzero volume equipped with an invertible, volume preserving map $P : B \rightarrow B$, for any open subset $U \subset B$*

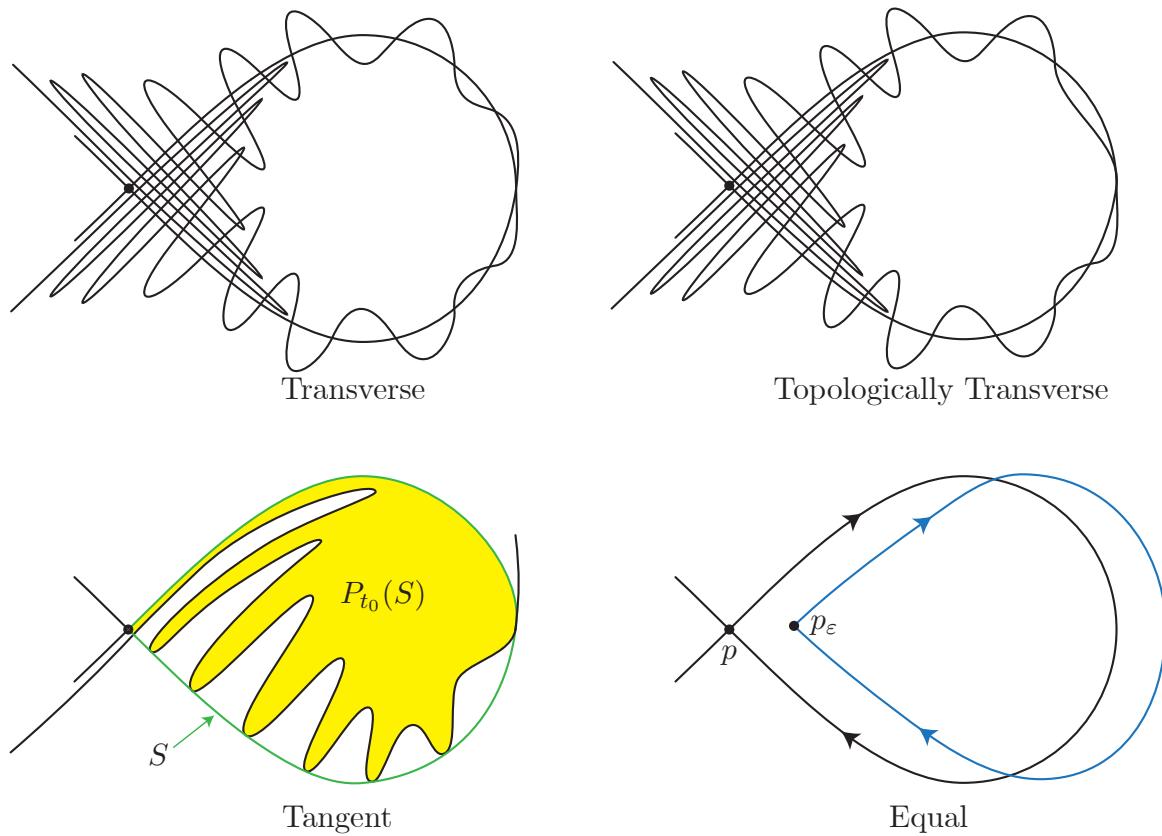


Figure 8.12: The four types of intersection geometries. Top row: transverse and topologically transverse, bottom row: tangent and equal.

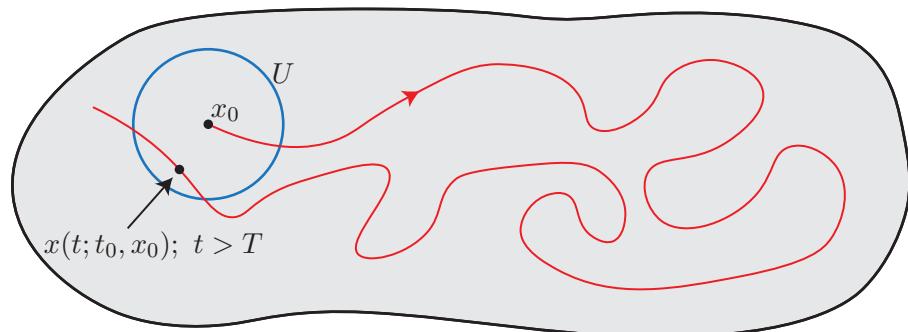


Figure 8.13: A depiction of the recurrence property of autonomous, area-preserving, Hamiltonian systems.

there exists a $N \in \mathbb{N}^+$ such that

$$P^N(U) \cap U \neq \emptyset. \quad (8.70)$$

Proof. Consider the sequence $U, P(U), P^2(U), \dots$, we claim that at least two of these sets intersect. Towards a contradiction, assume that each set is disjoint, therefore we have

$$\text{vol}[B] \geq \text{vol} \left[\bigcup_{i=0}^{\infty} P^i(U) \right] = \sum_{i=0}^{\infty} \text{vol}[P^i(U)] = \sum_{i=1}^{\infty} \text{vol}[U] = \infty. \quad (8.71)$$

The first inequality stems from the invariance of B , the second equality from the disjointness assumption, and the third from the volume preservation. This is clearly a contradiction as B is compact and must have finite volume. Therefore there exist m, n such that the intersection $P^m(U) \cap P^n(U)$ is non-empty. Call this intersection A and assume without loss of generality that $n > m$.

For an $x \in A$ we know that $x \in P^m(U)$, thus $P^{-m}(x) \in U$, similarly we know $x \in P^n(U)$ so $P^{-n}(x) \in P^{n-m}(U)$. Putting these two facts together tells us that $P^{-m}(x)$ is an element of U and $P^{n-m}(U)$, hence the intersection $U \cap P^{n-m}(U)$ cannot be empty. By calling $N = n - m$, we have shown the proposition. \square

Remark 8.10. There is a stronger version of Poincaré's Recurrence theorem, which states that almost all initial conditions return arbitrarily close to themselves. Here almost all is with respect to the assumed volume.

Example 8.7 (Recurrence of the pendulum). For the pendulum, the area enclosed by an energy level set can be used as S_0 . For large enough energy levels, this set may not be compact. This can be resolved by transforming the phase space $\mathbb{R}^2 \rightarrow \mathcal{S}^1 \times \mathbb{R}$ by taking advantage of the periodicity. This is shown in Fig. 8.14 where the importance of compactness and invariability are demonstrated. The figure also shows that one cannot simply apply the recurrence theorem to a Hamiltonian system restricted to an energy level, because this restricted system does not conserve volume any more.

Example 8.8 (Recurrence in non-smooth conservative systems). For a linear spring with a mass attached, which can bounce off an elastic wall as the spring is expanded, we can also observe recurrence. Denoting the deviation of the mass by q and the momentum by $m\dot{q}$ we still have recurrence under the map $P : x_0 \mapsto x(T; 0, x_0)$, which is volume preserving. The dynamical system and phase space are illustrated in Fig. 8.15.

Remark 8.11. These recurrent results extend to time-periodic Hamiltonian systems as well by applying them to the period- T map. Furthermore, they can be extended to quasi-periodic Hamiltonians.

Example 8.9 (Ideal gas in a partitioned container). An ideal gas in a partitioned three dimensional container, which we denote by $B_0 \subset \mathbb{R}^3$. One partition is full of $n \gg 1$ gas molecules, and the other partition is a vacuum. After removing the partition, statistical

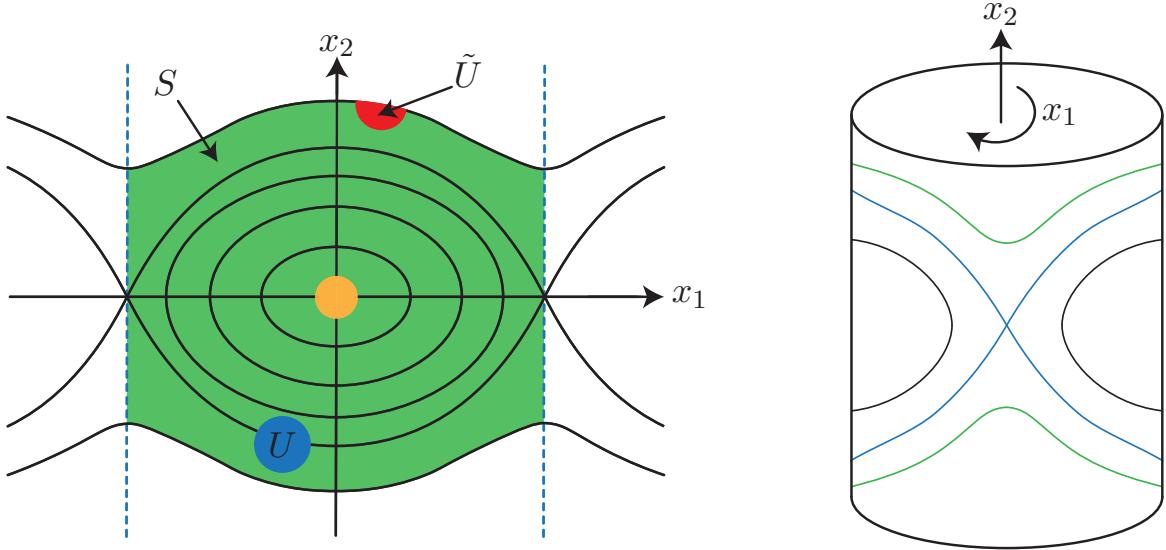


Figure 8.14: The recurrence of the pendulum is depicted along with the phase space transformation. The recurrence in the standard phase space does not hold for all points in the set U (blue), because initial conditions below the separatrix never return. This is due to the fact that in the phase space S_0 is not compact. If instead we identify the two dotted blue lines to transform the phase space based on periodicity, we find that recurrence does hold. Let us now restrict the dynamics to the separatrix, resulting in a one-dimensional dynamical system. The set in the intersection of the separatrix and the set U . This is a relatively open and compact set which is also invariant for the restricted dynamics, yet these points never return. However, in this case, the dynamics is not volume preserving any more, which means we cannot apply the recurrence theorem.

reasoning leads to the conclusion that the molecules will spread evenly. The layout of the partition is shown in Fig. 8.16. The position of the particles is described by the variable $q = (x_1, y_1, z_1, \dots, x_n, y_n, z_n)^T$ with $q \in B_0^n \subset \mathbb{R}^{3n}$, their respective velocities are described by $\dot{q} \in \mathbb{R}^{3n}$. The Lagrangian is made up of the kinetic energy $T(\dot{q}) = \frac{1}{2} \sum_{i=1}^n m_i \|v_i\|^2$ with $v_i = (\dot{x}_i, \dot{y}_i, \dot{z}_i)^T$ and the potential of the elastic interactions $V(q)$. With the generalized momentum given by $p = \frac{\partial L}{\partial \dot{q}}$, in this case the linear momentum, the Hamiltonian can be calculated as $H = H(q, p) = H(x)$ for $x = (q, p) \in \mathbb{R}^{6n}$. At each collision there is a discontinuity of the velocity, however the flow map $F_{t_0}^t$ is still invertible and volume preserving.

The initial configuration of the gas is given by $x_0 \in \mathbb{R}^{6n}$. The unique solution $x(t)$ describes the complete evolution of all of the molecules (their positions and momenta). This is just a single trajectory in this (large) phase space $B_0^n \times \mathbb{R}^{3n} = \{q \in \mathbb{R}^{3n} : q_i \in [0, K]\} \times \mathbb{R}^{3n}$. An example of such a path is depicted in Fig. 8.17. Observe that due to the conservation of energy

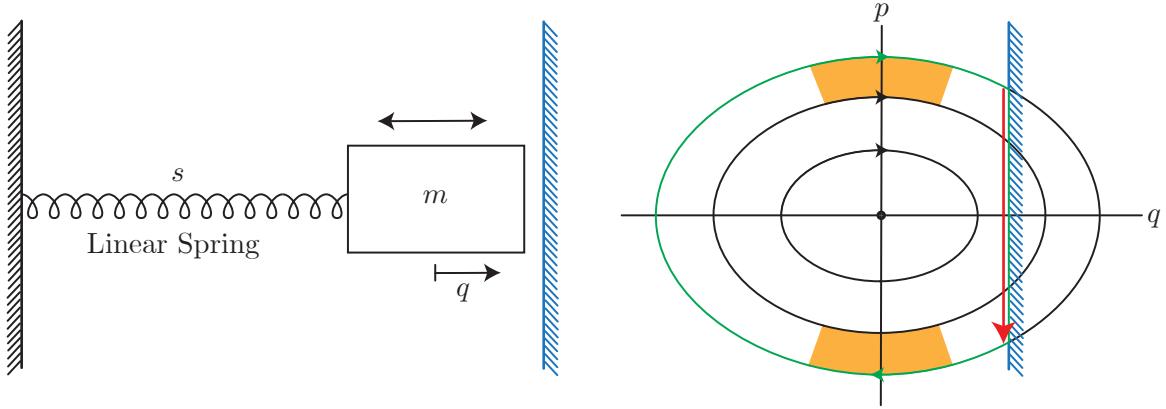


Figure 8.15: Left: A sketch of the dynamical system, with the elastic wall in blue. Right: The phase space of the dynamical system, with the red arrow designating that when the mass hits the wall (positive x_2) it immediately moves to the negative lower half (at the same absolute height) due to the elasticity. Furthermore, the volume preserving nature is shown by the two equal volume orange areas.

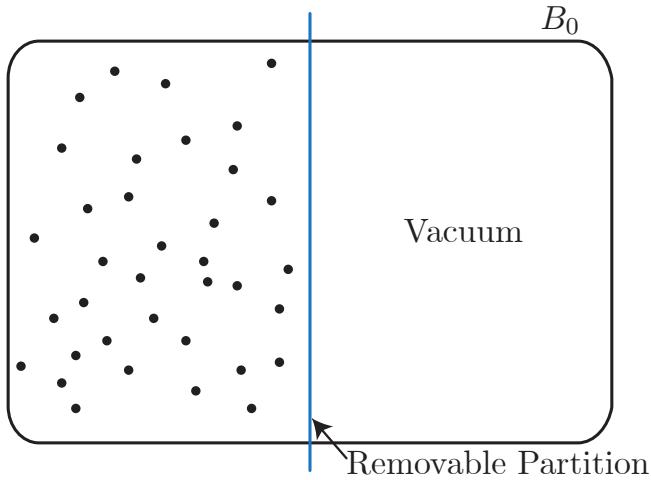


Figure 8.16: The gas contained in the container B_0 with the removable partition (blue) separating the gas and vacuum.

we have that the following is conserved for all times

$$H_0 = H(x_0) = H(x(t)) = T(\dot{q}(t)) + V(q(t)). \quad (8.72)$$

Furthermore, the potential is always positive, i.e. $V(q) \geq 0$. This implies that $T(\dot{q}(t)) \leq H_0$, thereby bounding $\|\dot{q}_i(t)\|$ for all times t . Thus the momenta $\|p(t)\|$ are also bounded for all times as $p = m\dot{q}$. The positions $q_i(t)$ are bounded by construction (they must remain within

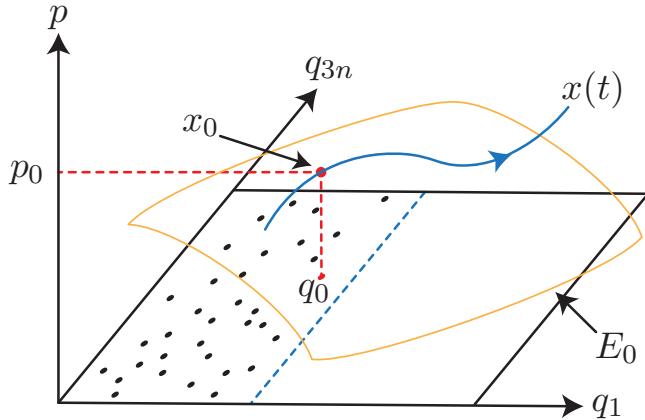


Figure 8.17: A single trajectory in the phase space for the partitioned container. Note that the trajectory is started from x_0 and remains within the energy level surface E_0 .

the container), therefore the region $S_0 = \{x \in \mathbb{R}^{6n} : H(x) \leq H_0\}$ is compact and invariant. By recurrence, almost all initial gas configurations return arbitrarily close to themselves over time. However, the recurrence time would be comparable to the life of the universe.

Remark 8.12. We cannot apply the recurrence to the energy surface as the restricted map may not be volume preserving.

The next example is of some historical value as this was the first ever numerical experiment to test a theory.

Example 8.10 (Fermi-Ulam-Pasta experiment (1955)). The experiment was conducted for large systems of linear, undamped, coupled oscillators. Since the system is linear, the energy content of each of the normal modes cannot change in time, and hence there is no equipartition of energy. According to the conclusions from statistical mechanics, a weak nonlinearity should change this completely and distribute the energy equally among the modes.. Initially, the experiments confirmed this hypothesis. However, after forgetting to turn off their computer for the night, they observed repeated returns close to the original energy configuration in their simulation, with the energy practically localized again in the same mode (ca. 3% error). More and more close returns were observed over longer time intervals.

The reason for this is that in Hamiltonian systems energy surfaces are compact and therefore recurrence applies for the phase space volume contained inside the energy surfaces.

8.4 Stability of fixed points in Hamiltonian systems

Recall the setup for autonomous Hamiltonian systems

$$\dot{x} = JDH(x), \quad x \in \mathbb{R}^{2n}. \quad (8.73)$$

The fixed points of such systems are critical points of the Hamiltonian, i.e. $DH(x_0) = 0$. Observe therefore that no asymptotically stable fixed points may exist in a Hamiltonian system. This is due to the volume preservation property. Assuming contrarily that the fixed point x_0 is asymptotically stable implies that there exists a ball B_0 within the open domain of attraction such that $F_{t_0}^t(B_0)$ is strictly contained within B_0 for t large enough. This would require that $\text{vol}[F_{t_0}^t(B_0)] < \text{vol}[B_0]$ for such a t , clearly violating the volume preservation property. Thus, at best, x_0 can be Lyapunov stable, and the nonlinear terms are relevant.

We first explore what we can find from the linear stability. To this end let $y = x - x_0$ and calculate the linearized system at x_0

$$\dot{y} = JD^2H(x_0)y. \quad (8.74)$$

Note that the linearized system is also Hamiltonian

$$\tilde{H}(y) = \frac{1}{2}\langle y, D^2H(x_0)y \rangle \implies \dot{y} = JD\tilde{H}(y). \quad (8.75)$$

Thus the phase space volume is also preserved for the linear system. This can be shown as we have before with

$$\nabla \cdot JD\tilde{H}(y) = \nabla \cdot JD^2H(x_0)y = \text{Tr}[JD^2H(x_0)] = 0. \quad (8.76)$$

This also tells us that the sum of the eigenvalues of the linearized Hamiltonian system is zero.

Further, observe that if λ is an eigenvalue of (8.74) (whether real or complex), then so is $-\lambda$. This leaves us with a few possible eigenvalue configurations for Hamiltonian systems, as seen in Fig. 8.18.

Exercise 8.5. *Show that if λ is an eigenvalue of a linearized Hamiltonian system, then so is $-\lambda$. (Hint: For any symmetric matrix A and any nonsingular matrix B , the eigenvalues of BA and $B^{-1}(BA)B$ coincide.)*

Suppose $\lambda = 0$ is an eigenvalue, then it is necessarily at least a double eigenvalue. Furthermore, under a variation of the parameters, purely imaginary eigenvalues can only leave the imaginary axis through a collision with another eigenvalue. Recalling that a linearized Hamiltonian can only be neutrally stable leads to the conclusion that linear stability analysis, is never conclusive to establish Lyapunov stability in Hamiltonian systems.

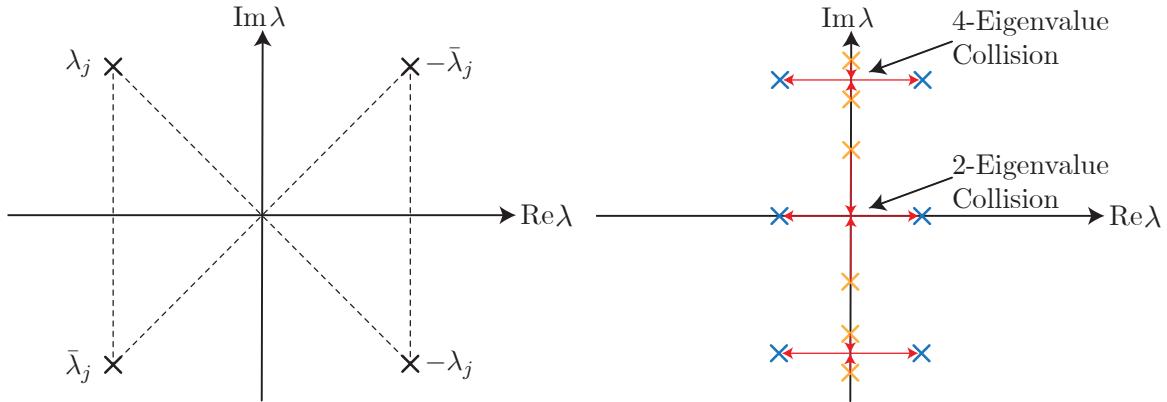


Figure 8.18: Left: An example eigenvalue configuration with the constraints found previously. Right: The possible losses of stability through eigenvalue collisions

The previous examination leads us to focusing on the nonlinear stability of fixed points of Hamiltonian. The setup is the same as before with an autonomous Hamiltonian system underlying the dynamics. Recall that for such systems we have that $\frac{dH(x(t))}{dt} = 0$, i.e. the Hamiltonian is conserved under the flow map, and as before a fixed point x_0 corresponds directly to a critical point of the Hamiltonian, i.e. $DH(x_0) = 0$.

These properties of H make it a candidate for a Lyapunov function. Observe that if $D^2H(x_0)$ is positive definite, then it is cup shaped and choosing H as a Lyapunov function implies that x_0 is Lyapunov stable. Meanwhile, if $D^2H(x_0)$ is negative definite, then choosing $-H$ as a Lyapunov function also implies that x_0 is stable. Hence, if $D^2H(x_0)$ is definite then x_0 is nonlinearly Lyapunov stable.

Exercise 8.6. A gradient system is a dynamical system of the form

$$\dot{x} = -DV(x); \quad x \in \mathbb{R}^n, \quad (8.77)$$

with some smooth function V . Apart from the absence of the symplectic matrix J , gradient systems appear similar to Hamiltonian systems. In fact, they are rather different, as the following exercises show.

- (i) Show that the eigenvalues of a linearized gradient system are always real.
- (ii) Find a condition for V under which a fixed point of a gradient system is asymptotically stable.
- (iii) Show that gradient systems cannot have periodic orbits.

- (iv) Given a smooth function $f(x)$, propose a numerical method for finding the local minima and maxima of f .

Example 8.11 (Multiple degree of freedom nonlinear mechanical system). Consider three pendulums, with masses m_1 , m_2 , and m_3 attached in order by nonlinear springs. The mass m_3 is attached to a mass m_4 which is constrained to move on a loop of radius R . The position of each mass is measured by its displacement from the unstretched spring position by the variable q , thus the springs are in their unstretched positions at the position $q = 0$. The system is depicted in Fig. 8.19.

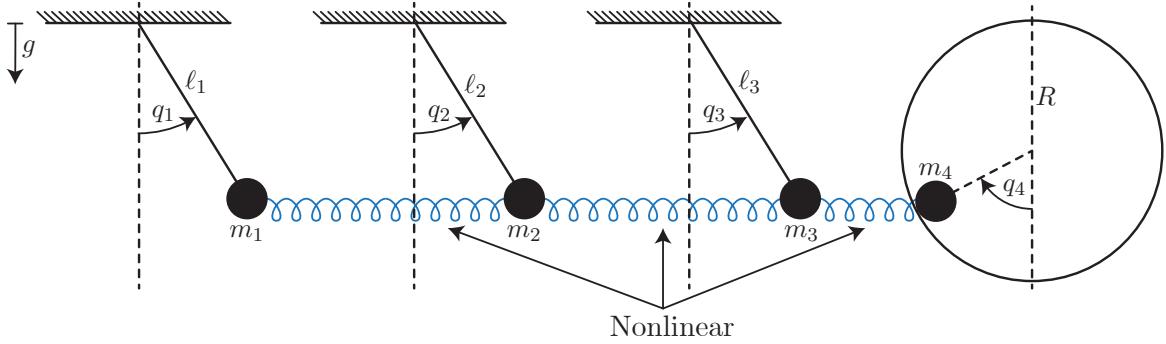


Figure 8.19: A depiction of the multiple degree of freedom nonlinear dynamical system consisting of three pendulums and a loop.

Concatenating the position and momentum into the variable x we have $x = (q, p)^T \in \mathbb{R}^8$. It is clear that $DH(0) = 0$ and due to the total mechanical energy having a strict minimum at $x = 0$ we have that $D^2H(0)$ is positive definite. Thus $x = 0$ is Lyapunov stable.

We may wonder if all linearly stable Hamiltonian fixed points are in fact always nonlinearly stable, i.e. if $D^2H(x_0)$ is always definite. This is not the case, as we can see by taking

$$H = \frac{1}{2} (q_1^2 + p_1^2) - \frac{1}{2} (q_2^2 + p_2^2); \quad x = \begin{pmatrix} q_1 \\ p_1 \\ q_2 \\ p_2 \end{pmatrix} \in \mathbb{R}^4. \quad (8.78)$$

This leads to the dynamics

$$\begin{cases} \dot{q}_1 = p_1 \\ \dot{q}_2 = -p_2 \\ \dot{p}_1 = -q_1 \\ \dot{p}_2 = q_2 \end{cases} \implies \dot{x} = \left(\begin{array}{cc|cc} 0 & 1 & 0 & 0 \\ -1 & 0 & 0 & 0 \\ \hline 0 & 0 & 0 & -1 \\ 0 & 0 & 1 & 0 \end{array} \right) x. \quad (8.79)$$

The eigenvalues of this system are $\pm i$, thus linear stability analysis and use of H as a Lyapunov function (see Fig. 8.20) are both inconclusive.

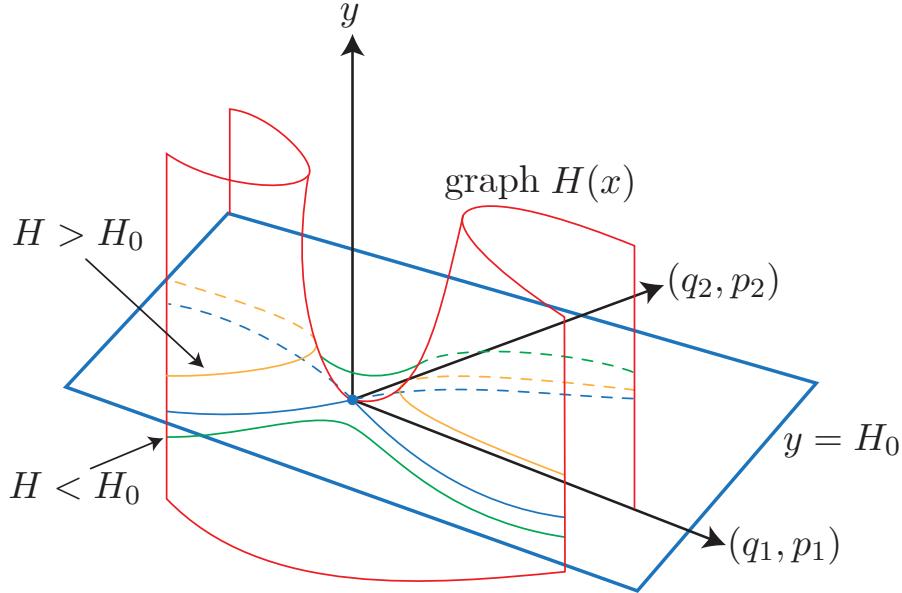


Figure 8.20: The graph of the Hamiltonian H showing the lack of definiteness which leads to an inconclusive Lyapunov stability analysis.

Example 8.12 (Gyroscopic systems). A classic example is that of a gyroscopic system such as the tippe-top toy (a spinning top like a dreidel). When spun there is a relative equilibrium fixed point for the Hamiltonian obtained by factoring out the cyclic coordinate, which is achieved via a Routh-reduction of cyclic coordinates. Fast spinning stabilizes this Hamiltonian equilibrium, but the Hamiltonian has no minimum there (nonlinear analysis is needed to prove this). Additional dissipation will then always destabilize these equilibria in a process called *dissipation-induced instability* which is shown in Fig. 8.21.

With these examples, it is clear that further study of nonlinear stability of Hamiltonian fixed points is necessary for the case that $D^2H(x_0)$ is not definite. The setting we will work in to study this is as follows

$$\dot{x} = JDH(x); \quad x \in \mathbb{R}^{2n}; \quad DH(0) = 0. \quad (8.80)$$

The final equality is always achievable by a coordinate shift. Once again we begin by linearizing to find $\dot{y} = Ay$. Assume for simplicity that A is semisimple, i.e. it has $2n$ linearly independent eigenvectors (even if some eigenvalues are repeated). Through a change of coordinates A can

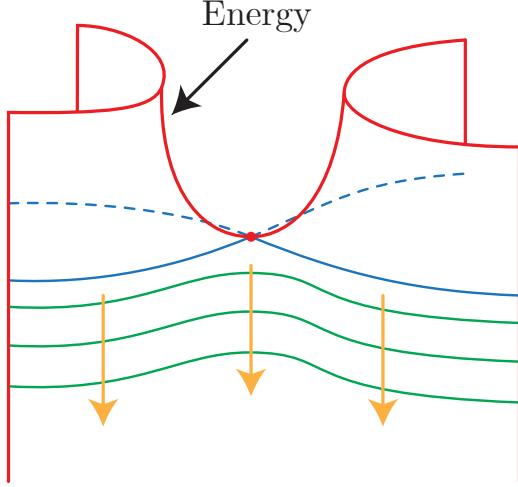


Figure 8.21: The energy is shown in red, the dissipation of energy takes the motion towards lower energy levels as shown by the yellow arrows.

be block-diagonalized

$$A = \begin{pmatrix} A_1 & & \\ & \ddots & \\ & & A_n \end{pmatrix}; \quad A_j = \begin{pmatrix} 0 & \omega_j \\ -\omega_j & 0 \end{pmatrix}. \quad (8.81)$$

Each block A_j corresponds to the eigenvalues $\pm i\omega_j$. There is no real part to the eigenvalue, otherwise $x = 0$ would be unstable, however $\omega_j < 0$ is possible. This change of coordinates allows us to frame the geometry of the linearization as the product of n different two dimensional dynamical systems, each with clockwise rotating cyclical orbits, i.e. $\prod_{i=1}^n S^1 = \mathbb{T}^n$ the n -dimensional torus. This product is shown in Fig. 8.22.

Exercise 8.7. Consider the Lotka-Volterra model

$$\begin{cases} \dot{h} = a_1 h (1 - bp) \\ \dot{p} = -a_2 p (1 - ch), \end{cases} \quad (8.82)$$

for the interaction of a predator and a prey population. Here $h(t)$ and $p(t)$ denote the predator and prey populations, respectively, as a function of time; a_1 , a_2 , b , and c are positive parameters.

- (i) Show that system (8.82) is Hamiltonian for $h, p > 0$ after an appropriate rescaling of time. Find the Hamiltonian. (Hint: Rewrite (8.82) as $\dot{h} = A(h, p)C(p)$, $\dot{p} = A(h, p)D(h)$ by defining the functions A, C and D appropriately.)

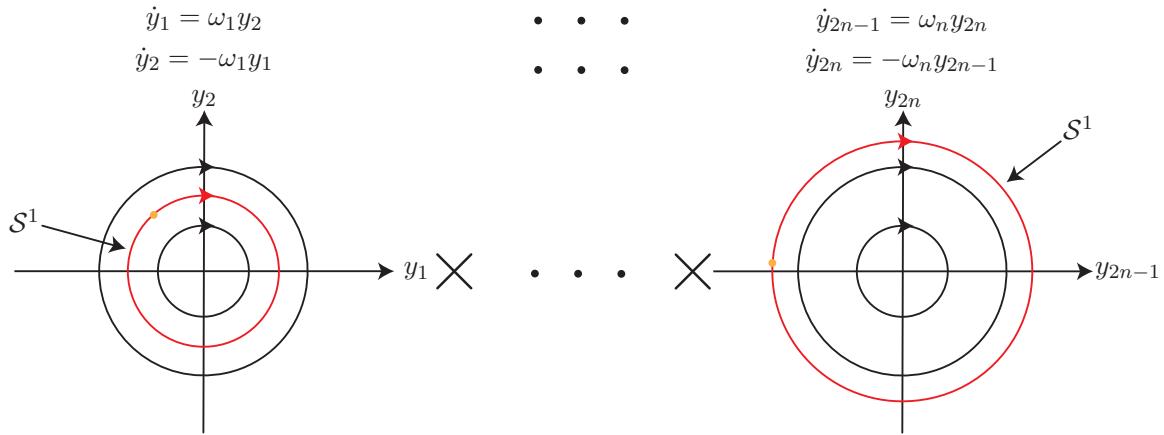


Figure 8.22: The product of the n dynamical systems of the linearization forming the n -dimensional torus.

- (ii) *Using the Hamiltonian, argue that the two species can exhibit stable coexistence, i.e., the system admits a stable fixed point. Hint: establish full nonlinear stability for the fixed point.*

Defining the vector of angular velocities $\omega = (\omega_1, \dots, \omega_n)$ we get three cases for the structure of the orbits

- (i) **Resonance** If there exists a $k \in \mathbb{Z}_+^n$ (i.e. $k_i \neq 0$ for all i) with $\langle k, \omega \rangle = 0$, trajectories close up. Therefore all orbits are closed and periodic on \mathbb{T}^n .
- (ii) **Non-resonance** If for each $k \in \mathbb{Z}^n$ with $\|k\| \neq 0$ we have that $\langle k, \omega \rangle \neq 0$, trajectories never close up. Therefore all orbits are quasi-periodic and yield a dense winding.
- (iii) **Partial resonance** If there exists a $k \in \mathbb{Z}^n$ with $\|k\| \neq 0$ and $\langle k, \omega \rangle = 0$ we get a combination of the previous two cases. Trajectories fill lower dimensional tori densely and are not closed.

8.5 Stability of Hamiltonian fixed points: normal forms

The setting of the previous section remains the same as we move to introduce the idea of the normal form of $H(x)$ near $x = 0$. This requires two additional assumptions

- (i) For each $k \in \mathbb{Z}^n$ with $\|k\|_1 = \sum_{i=1}^n |k_i| \leq 4$ we have $\langle k, \omega \rangle \neq 0$, i.e. there are no resonances below order 5.

The first assumption implies that a nonlinear, near-identity canonical change of coordinates puts $H(x)$ in the form

$$H(\varphi, r) = \sum_{i=1}^n \omega_i r_i + \frac{1}{2} \sum_{k,l=1}^n \omega_{k,l} r_k r_l + \mathcal{O}(|r|^4); \quad r \in \mathbb{R}_{\geq 0}^n; \quad \varphi \in \mathbb{T}^n. \quad (8.83)$$

The φ -dependence starts only in the higher order terms, i.e. $\mathcal{O}(|r|^n)$ in this case. By a near-identity change of coordinates, we mean that (r, φ) is close to the polar coordinates introduced for the linearized Hamiltonian.

Definition 8.1. A *canonical* (also *symplectic*) coordinate change is one which preserves the Hamiltonian structure. That is, in the new coordinates, Hamilton's equations still hold

$$\begin{cases} \dot{\varphi} = \frac{\partial H}{\partial r} \\ \dot{r} = -\frac{\partial H}{\partial \varphi}. \end{cases} \quad (8.84)$$

These reductions bring us to the next theorem. In order to state this, some terminology must be rigorously defined for this section.

Definition 8.2. *Persists* in this cases corresponds to surviving the transition back from the linearization to the nonlinear system through a smooth deformation by a small amount.

Definition 8.3. *Most* (in this case invariant tori) means

$$\lim_{\varepsilon \rightarrow 0} \frac{\mu(\text{persisting tori in } \varepsilon\text{-ball})}{\mu(\mathcal{B}(\varepsilon))} = 1, \quad (8.85)$$

where μ is the measure that is preserved, i.e. the phase space volume.

(ii) The *nondegeneracy condition* holds, i.e. $\det[\omega_{kl}] \neq 0$.

Theorem 8.13 (Kolmogorov-Arnold-Moser (KAM) theorem). *We have three items which hold:*

- (i) *Under the assumptions (i) and (ii), most invariant tori of the linearized Hamiltonian persist in the nonlinear system close to $x = 0$.*
- (ii) *Surviving tori are quasi-periodic, implying that trajectories are dense on these and obey ODEs of the form*

$$\begin{cases} \dot{r} = \mathcal{O}(|r|^4) \\ \dot{\varphi} = \tilde{\omega}(r, \varphi) = \frac{\partial H}{\partial r} = \omega + [\omega_{kl}]r + \mathcal{O}(|r|^3). \end{cases} \quad (8.86)$$

- (iii) Under the additional nondegeneracy condition known as the isoenergetic nondegeneracy condition

$$\det \begin{bmatrix} [\omega_{kl}] & \omega^T \\ \omega & 0 \end{bmatrix} \neq 0, \quad (8.87)$$

most tori persist within each energy surface close enough to the origin. This is shown in Fig. 8.23.

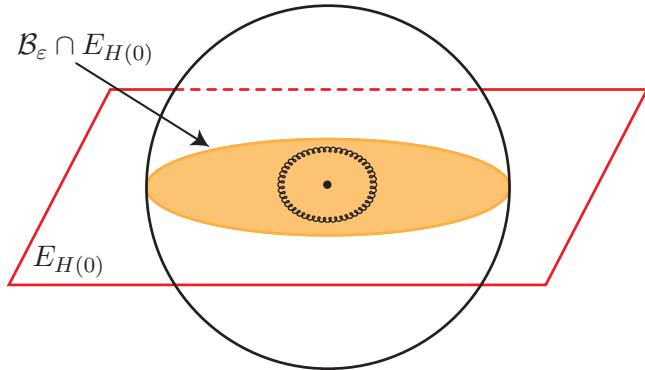


Figure 8.23: Most tori persist with respect to the measure defined inside the energy surface $E_{H(0)}$ (red). The ε -ball is given in black, and its intersection with $E_{H(0)}$, which is of dimension $2n - 1$.

This has different implications for the dynamics near $x = 0$ depending on the degrees of freedom based on the dimension of $E_{H(x)}$ ($= 2n - 1$). For $x \neq 0$ close to the origin the dimension of the surviving tori is $\dim \mathbb{T}^n = n$, thus the codimension of \mathbb{T}^n within $E_{H(x)}$ is $(2n - 1) - n = n - 1$.

- (i) **1 degree of freedom** (e.g. pendulum) The codimension of \mathbb{T}^1 is 0, therefore the torus is the energy surface itself and $x = 0$ is Lyapunov stable.
- (ii) **2 degrees of freedom** (e.g. two coupled pendulums) The codimension of \mathbb{T}^2 is 1, so the two-dimensional torus locally divides the three-dimensional energy surface into an interior and exterior. The interior is a compact invariant set that trajectories cannot escape from within the given energy surface. Hence $x = 0$ is Lyapunov stable, even if $D^2H(0)$ is indefinite.
- (iii) **≥ 3 degrees of freedom** (e.g. three coupled pendulums) The codimension of \mathbb{T}^n is at least 2, thus the tori no longer form barriers and trajectories can sneak out. Generically, trajectories do sneak out along a chaotic invariant set formed near the remnants of destroyed (resonant) tori such as in *Arnold diffusion*.

In Fig. 8.24 we show a nested family of surviving KAM-tori and the internal structure of the gaps between them. The gaps are due to resonances and are filled with lower dimensional hyperbolic and elliptic tori. A heteroclinic tangle forms as the intersection of the stable and

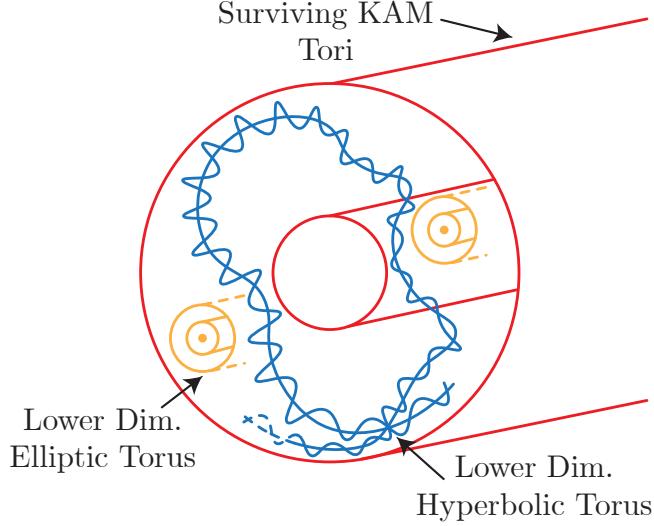


Figure 8.24: The geometry of Arnold diffusion. The red tori are two surviving KAM-tori of a large dimensional dynamical system. In the gaps between them we find hyperbolic and elliptic tori, which are lower dimensional.

unstable manifolds of the hyperbolic (saddle-type tori). This tangle works as a transition chain and can produce $\mathcal{O}(1)$ displacements in r for trajectories started near the surviving torus, even when $\varepsilon \rightarrow 0$.

Therefore we have a mechanism for instability when $D^2H(0)$ is not definite. When it is definite, the transition chains are still there but energy surfaces are small spheres near the origin. This implies that complicated dynamics does not lead to instability.

Theorem 8.14 (Nekhoroshev's theorem). *Under further nondegeneracy conditions, the speed of diffusion is less than $C \exp(-\frac{1}{\varepsilon^d})$, where d depends on $H(x)$. This means that the speed is exponentially small.*

Chapter 9

Normally hyperbolic invariant manifolds

The objective of this chapter is to extend the theory of low-dimensional stable and unstable manifolds to a general, higher-dimensional setting. The focus will lie on invariant manifolds rather than individual trajectories as invariant manifolds are more observable, more influential, and more robust. We begin with an example.

Example 9.1 (Dynamics near a saddle-type fixed point). The difference between the behavior of individual trajectories and of the invariant manifolds is depicted in Fig. 9.1. The idea being that under a diffeomorphic transformation of the invariant manifolds of the linear system into those of the nonlinear system, a trajectory starting from a point x_0 may be drastically different even if the dynamics remain topologically conjugate to each other. I.e. individual trajectories are sensitive with respect to initial conditions and changes in parameters, meanwhile robust invariant manifolds persist (perturb smoothly).

Thus our examination now turns to determining which invariant manifolds are persistent. In order to understand this, first an intermezzo in differentiable manifolds is necessary.

9.1 A crash course on differentiable manifolds

Definition 9.1. A set $M \subset \mathbb{R}^n$ is a *k-dimensional differentiable manifold*, if it is locally diffeomorphic to \mathbb{R}^k , i.e. for every $x \in M$ there exists an open (in M) neighborhood $V \subset M$, such that V is diffeomorphic to an open set $U \subset \mathbb{R}^k$. The diffeomorphism is called the *local parameterization* as is denoted $\Phi : U \rightarrow V$. Its smooth inverse is the *local coordinate system* $\Phi^{-1} : V \rightarrow U$. These are illustrated in Fig. 9.2.

The intuitive meaning of a *k*-dimensional manifold is a set which locally looks like \mathbb{R}^k .

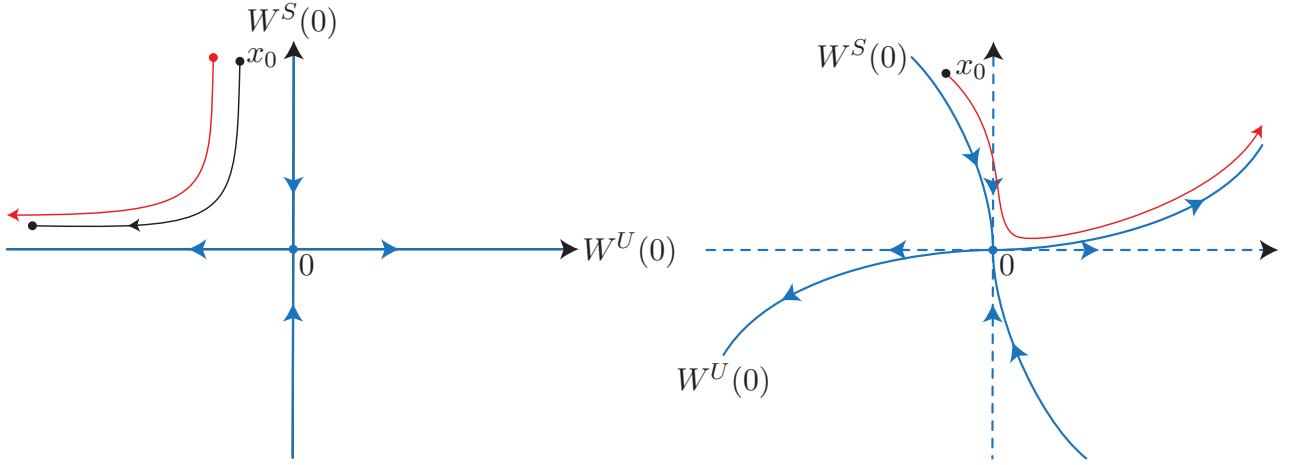


Figure 9.1: The dynamics near a saddle-type fixed point demonstrating the sensitive dependence of individual trajectories on initial conditions and changes in parameters, and the robust nature of invariant manifolds.

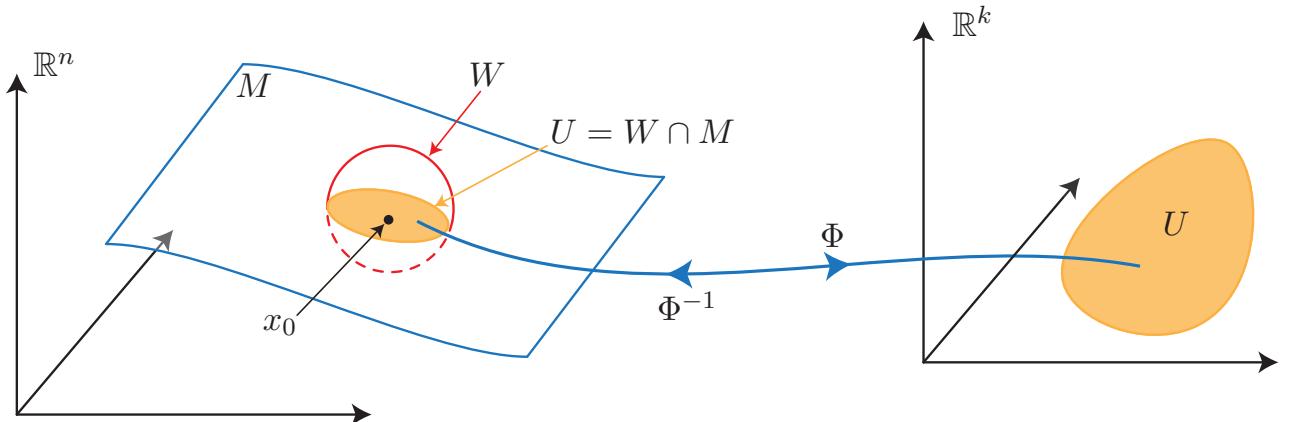


Figure 9.2: The differentiable manifold M is given in blue in the left. The red sphere W is an open ball in \mathbb{R}^n , intersecting this with M yields and open (in M) neighborhood $V = W \cap M$ of x . The smooth diffeomorphism Φ transforms U into the open set U of \mathbb{R}^k .

Definition 9.2. For a set $M \subset \mathbb{R}^n$ and an open set $U \subset \mathbb{R}^n$, the set $M \cap U$ is called *relatively open* in M .

Example 9.2 (The circle is a 1-dimensional manifold). The circle S^1 is a 1-dimensional differentiable manifold. To see define the four open subsets V_i of S^1 as open semicircles which are each shifted by $\frac{i\pi}{2}$ for $i = 1, 2, 3, 4$. Then $S^1 = \bigcup_{i=1}^4 V_i$. These are drawn in Fig. 9.3. Any $x \in S^1$ belongs to at least one of the V_i due to the previous equality. Furthermore each V_i is

diffeomorphic to $U = (-1, 1)$ which is open in \mathbb{R}^1 via the diffeomorphisms

$$\Phi_1 = \begin{pmatrix} t \\ \sqrt{1-t^2} \end{pmatrix}; \quad \Phi_2 = \begin{pmatrix} t \\ -\sqrt{1-t^2} \end{pmatrix}; \quad (9.1a)$$

$$\Phi_3 = \begin{pmatrix} -\sqrt{1-t^2} \\ t \end{pmatrix}; \quad \Phi_4 = \begin{pmatrix} \sqrt{1-t^2} \\ t \end{pmatrix}. \quad (9.1b)$$

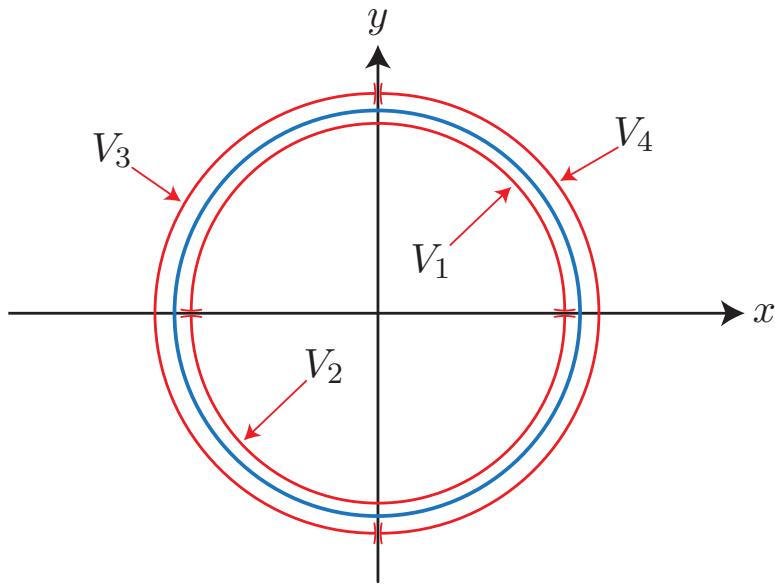


Figure 9.3: The subsets V_i of S^1 which are used to show that the circle is a 1-dimensional manifold.

There is no global parameterization of S^1 , but this is not needed to be a differentiable manifold, and these local parameterizations suffice.

Example 9.3 (Not all sets are manifolds). A few examples of sets which are not manifolds.

- (i) The figure-eight is not a manifold, as for any point except the central intersection, there exists a local neighborhood diffeomorphic to \mathbb{R}^1 , however at this central intersection, no open neighborhood exists which is diffeomorphic to \mathbb{R}^1 . This is depicted in Fig. 9.4.
- (ii) A dense orbit on a 2-dimensional torus is not a manifold, as around any point infinitely many adjacent trajectories exist, but do not form a plane (there is a distance between two instances of the trajectory). Hence any neighborhood around any point cannot be hyperbolic to \mathbb{R}^2 or \mathbb{R}^1 . Such a neighborhood is illustrated in Fig. 9.5.

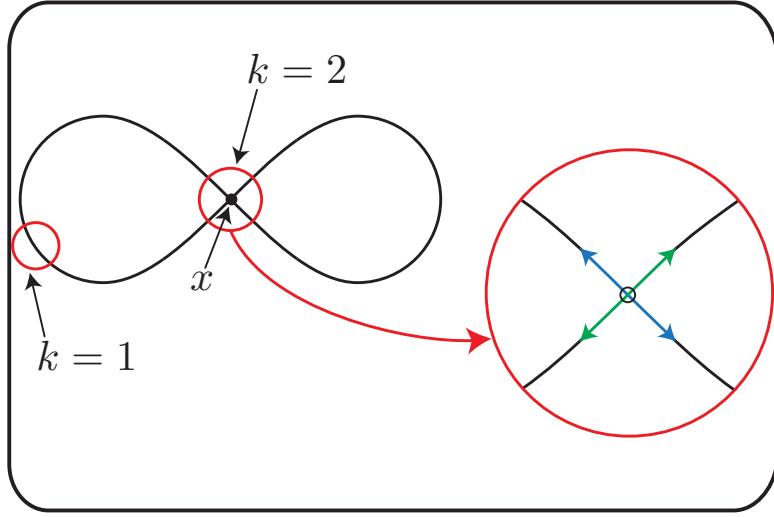


Figure 9.4: The figure-eight with the critical intersection which prevents it from being a manifold.

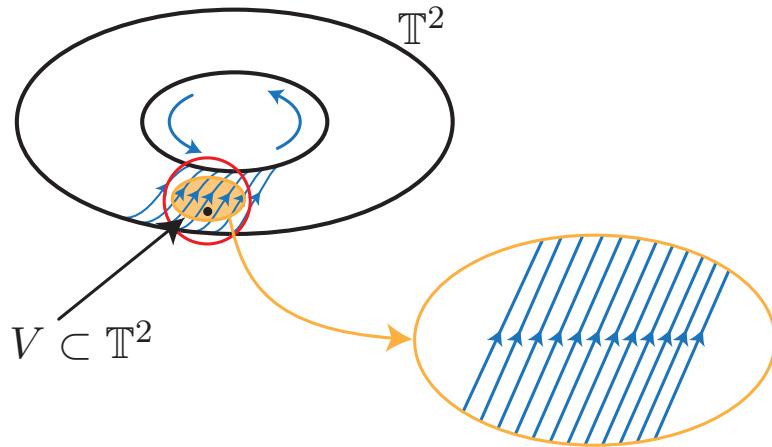


Figure 9.5: A depiction of a local neighborhood of a point in a dense orbit of the torus.

- (iii) The union of a semi-infinite spiral and semi-open (left closed, right open) interval do not form a differentiable manifold. For this set, at the intersections of the two sets, points with neighborhoods as in the figure eight appear.

We can construct manifolds very easily, for instance by taking any smooth function $f : X \rightarrow Y$, the graph of f namely the set $\text{graph}(f) = \{(x, y) : x \in X, y = f(x)\}$ is a differentiable manifold.

Definition 9.3. For two manifolds X and Z , each subsets of \mathbb{R}^n , with $X \subset Z$ we call X a

submanifold of Z .

An example of a submanifold would be $X = \mathcal{S}^1 \subset Z = \mathcal{S}^2 \subset \mathbb{R}^3$. We say \mathcal{S}^1 is a 1-dimensional submanifold of the 2-dimension manifold Z .

Definition 9.4. A manifold $M \subset \mathbb{R}^n$ is a *k-dimensional manifold with boundary*, if it is locally diffeomorphic to a relatively open set in the non-negative half-space $H^k = \{x \in \mathbb{R}^k : x_1 \geq 0\} \subset \mathbb{R}^k$. The *boundary of M* , denoted ∂M is always mapped into ∂H^k under any parameterization, and is a $k - 1$ -dimensional manifold. Such a manifold and its parameterization are illustrated in Fig. 9.6.

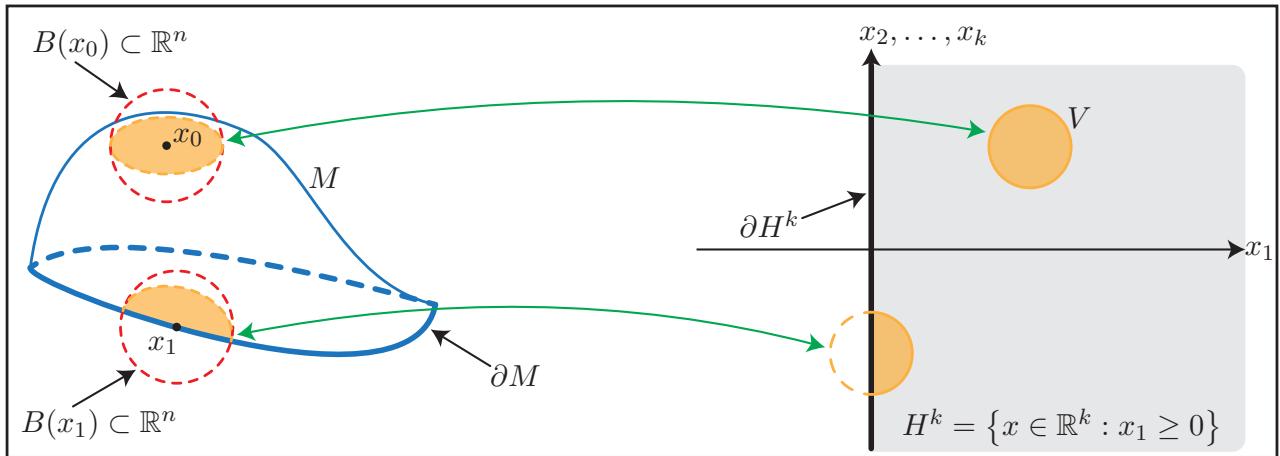


Figure 9.6: An illustration of the manifold with boundary M and a parameterization mapping its boundary to ∂H^k .

Closed intervals, closed cylinders, and closed hollow spheres are all manifolds with boundaries, however it is not always the case that simply taking the closure of a manifold yields a manifold with boundary. For instance the interior of a rectangle is a manifold (using the identity as a parameterization), however as the corners are not smooth the closure is not a manifold.

Exercise 9.1. Let $f : X \rightarrow Y$ be a smooth function, where $X \subset \mathbb{R}^n$ is a manifold. Prove that the graph of f ,

$$\text{graph}(f) = \{(x, y) \in X \times Y : y = f(x)\} \quad (9.2)$$

is always a manifold.

Definition 9.5. Given a k -dimensional manifold $M \subset \mathbb{R}^n$ with the local parameterization at x given by $\Phi : U \subset \mathbb{R}^k \rightarrow V \subset M \subset \mathbb{R}^n$ (assume for simplicity that $\Phi(0) = x \in V$), the *tangent space of the manifold* is then defined as

$$T_x M = \text{Range}[D\Phi(0)]. \quad (9.3)$$

The intuition as to why this is called the tangent space is shown in Fig. 9.7.

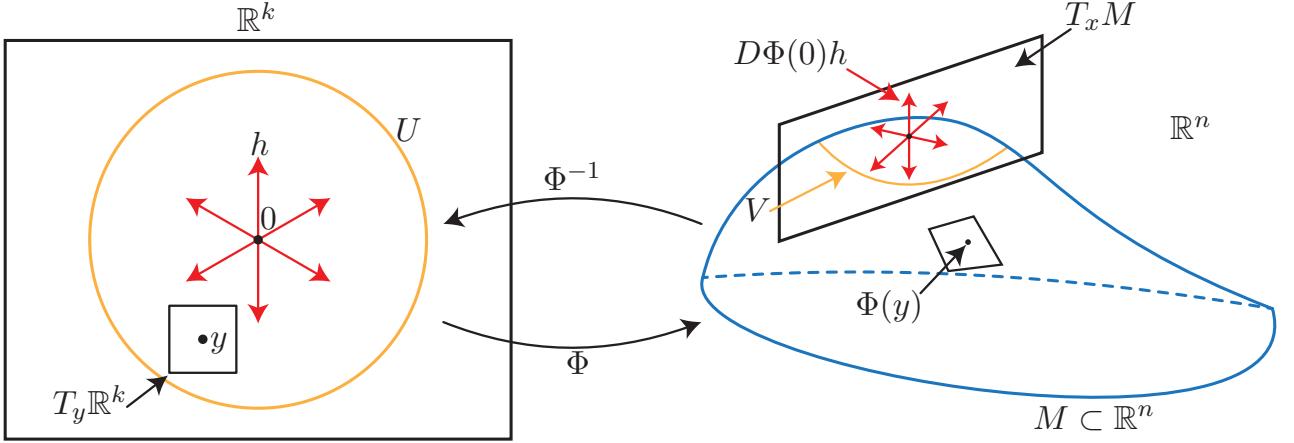


Figure 9.7: The tangent space of a manifold M . The map $D\Phi(0)$ maps the left image to the plane indicated in the right image.

Note here how the differential is defined

$$D\Phi(x)h = \lim_{s \rightarrow 0} \frac{\Phi(x + sh) - \Phi(x)}{s}. \quad (9.4)$$

Therefore the map $D\Phi(y)$ is a map from the tangent space of \mathbb{R}^k at y to the tangent space of M at $\Phi(y)$, i.e. $D\Phi(y) : T_y \mathbb{R}^k \rightarrow T_{\Phi(y)} M$.

Remark 9.1. Note a few facts about tangent spaces:

- (i) Although it appears that $T_x M$ depends on the parameterization Φ , it actually does not;
- (ii) The dimension of the tangent space, which is well defined as it is a linear subspace of \mathbb{R}^n , is equal to the dimension of M , namely k ;
- (iii) A map $f : M \rightarrow N$ between manifolds induces a linear mapping between the appropriate tangent spaces of the manifolds. Denoting by Φ and Ψ the respective local parameterizations, the induced map between parameter spaces is $F = \Psi^{-1} \circ f \circ \Phi$. This is a mapping

between linear spaces and DF is well-defined. In turn, this implies that Df is well-defined and by the chain rule is equal to $D\Psi \circ DF \circ D\Phi^{-1}$, therefore it is possible to differentiate mappings between manifolds. Then we have the sought-after mapping between tangent spaces

$$Df : T_x M \rightarrow T_{f(x)} N. \quad (9.5)$$

These maps are illustrated in Fig. 9.8.

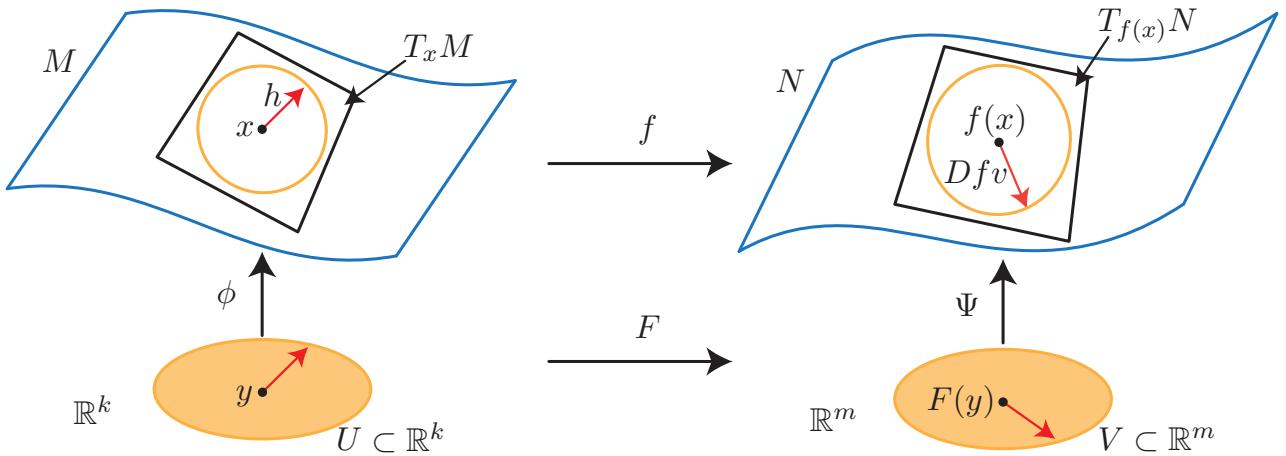


Figure 9.8: The commutative diagram along with the respective spaces illustrated for the induced map between tangent spaces.

Exercise 9.2. Show that the tangent space of a manifold at any point is independent of the local parametrization used in its construction (i.e., another local parametrization would give the same tangent space).

Definition 9.6. The set of all tangent spaces, labelled by their base points is called the *tangent bundle*

$$TM = \{(x, v) : x \in M, v \in T_x M\} = \bigcup_{x \in M} (x, T_x M). \quad (9.6)$$

The tangent bundle is a $2k$ -dimensional manifold.

Exercise 9.3. Prove that the tangent bundle of a manifold is a manifold by constructing an explicit local parametrization.

By labelling tangent vectors with their respective base points, different tangent spaces are separate *fibers* of the tangent bundle. In general $TM \neq M \times \mathbb{R}^k$, i.e. the tangent bundle is not a global direct product. In fact, locally, each tangent bundle is trivial (the direct product), this just does not extend globally. However, for $M = S^2$ the tangent bundle is diffeomorphic to $S^2 \times \mathbb{R}^2$.

Definition 9.7. The *normal space*, unlike the tangent space, depends on the ambient space. For a k -dimensional manifold M embedded in \mathbb{R}^n , the normal space at a point x , $N_x M$, is the orthogonal complement of the tangent space. For their direct sum it holds, that

$$T_x M \oplus N_x M = T_x \mathbb{R}^n. \quad (9.7)$$

Therefore, the dimensional of the normal space is $n - k$. The normal space is drawn in Fig. 9.9.

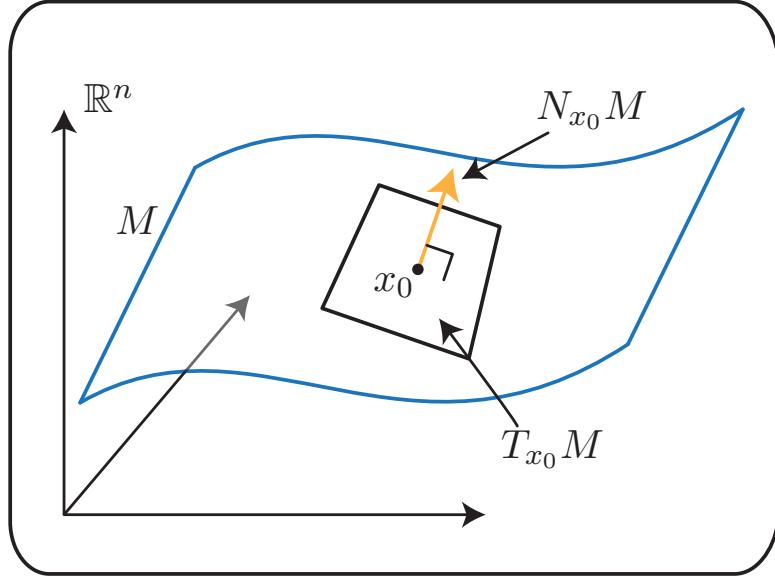


Figure 9.9: The normal space of a manifold M at the point x_0 along with the tangent space at the same point.

Definition 9.8. Analog to the tangent bundle, the *normal bundle* is the collection of normal spaces of the manifold M labelled by their base points

$$NM = \{(x, v) : x \in M, v \in N_x M\}. \quad (9.8)$$

By labelling normal vectors with their base points, intersections are avoided as shown in Fig. 9.10.

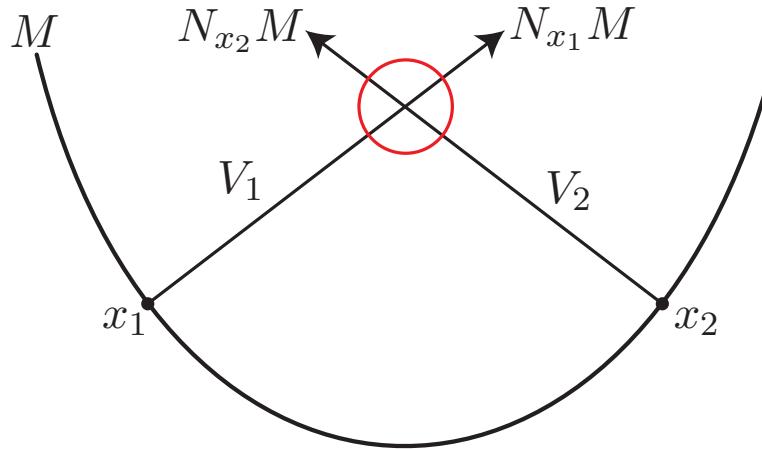


Figure 9.10: Intersecting normal vectors, these are avoided by labelling vectors with their base points.

Definition 9.9. A *subbundle* of the normal bundle is any x -dependent family of subspaces of $N_x M$, which is smooth in x .

An example of such a subbundle is given in Fig. 9.11, where the smooth family of subspaces is given by $S(x) \subset N_x M$.

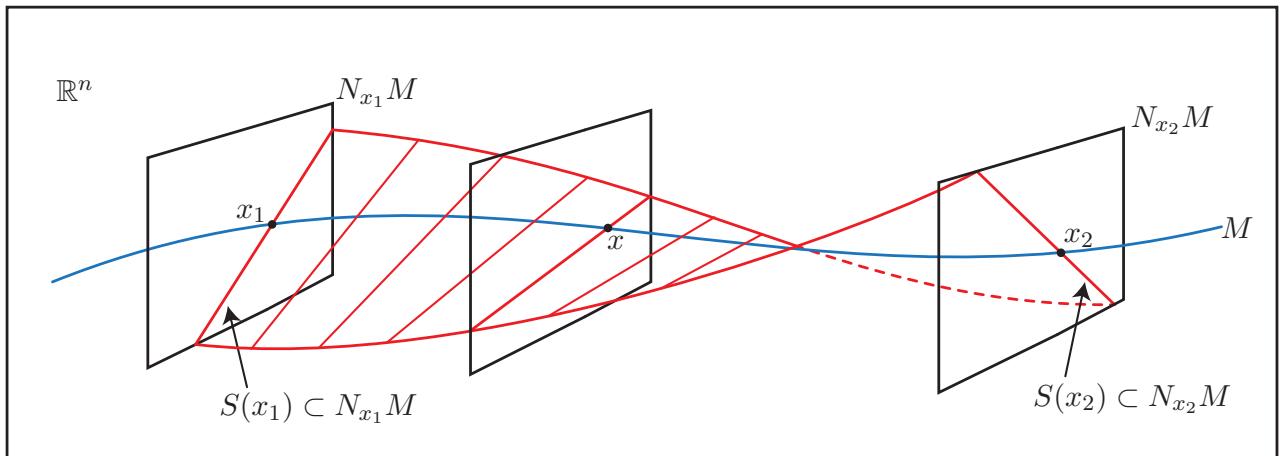


Figure 9.11: A subbundle of the manifold M , each base point $x \in M$ is a base point of $S(x)$, a fiber of the normal bundle.

9.2 Theory of normally hyperbolic invariant manifolds

The starting point for the following section will be of an autonomous dynamical system

$$\dot{x} = f(x); \quad x \in \mathbb{R}^n; \quad f \in \mathcal{C}^r \quad r \geq 1. \quad (9.9)$$

Furthermore, let $M_0 \subset \mathbb{R}^n$ be a k -dimensional, compact, \mathcal{C}^1 -manifold with boundary ∂M_0 .

Definition 9.10. The manifold M_0 is an *overflowing invariant manifold* if $f(x)$ points strictly outwards of ∂M_0 . A special case of this is when $\partial M_0 = \emptyset$.

There are three central questions for what happens to this manifold which we will address:

- (i) Does it persists (survive smoothly)?
- (ii) Is it stable (what happens to nearby trajectories)?
- (iii) What is the interior structure of the stable/unstable manifolds of M_0 ?

Example 9.4. Consider the system

$$\begin{cases} \dot{x} = -x(1-x)(1+x) \\ \dot{y} = \alpha y \end{cases}; \quad 0 < \alpha \leq 1. \quad (9.10)$$

Now define M_0 to be the closed interval $[-1 - \varepsilon, 1 + \varepsilon] \times \{0\}$. This manifold is of dimension 1, has nonempty boundary $\{-1 - \varepsilon, 1 + \varepsilon\}$, is compact, and smooth. At each point in the boundary, f is pointing away from M_0 . Therefore, M_0 is an overflowing invariant manifold. We also have that $W_{\text{loc}}^S(M_0)$ is nonempty, meanwhile $W_{\text{loc}}^U(M_0)$ is empty. Trajectories of this dynamical system along with M_0 and the stable local manifold are depicted in Fig. 9.12.

Now we may ask if M_0 is robust under perturbations (if it persists). For example under the perturbation

$$\begin{cases} \dot{x} = -x(1-x)(1+x) - \varepsilon y \\ \dot{y} = -\alpha y + \varepsilon x \end{cases}; \quad 0 \leq \varepsilon \ll 1. \quad (9.11)$$

The fixed points $A_0 = (-1, 0)$, $B_0 = (0, 0)$, and $C_0 = (1, 0)$ are hyperbolic, therefore they persist. In fact B_0 does not move at all. The other fixed points move to new perturbed counterparts, i.e. $A_0, C_0 \rightarrow A_\varepsilon, C_\varepsilon$. From the second equation, we see that the y coordinate of the perturbed fixed point becomes $y_\varepsilon = \varepsilon / \alpha x_\varepsilon$. That is, C_ε moves upwards and A_ε moves downwards.

At $B_0 = B_\varepsilon$ we can linearize to find

$$\begin{pmatrix} \dot{x} \\ \dot{y} \end{pmatrix} = \begin{pmatrix} -1 & -\varepsilon \\ \varepsilon & -\alpha \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix}. \quad (9.12)$$

Depending on α we differentiate two cases.

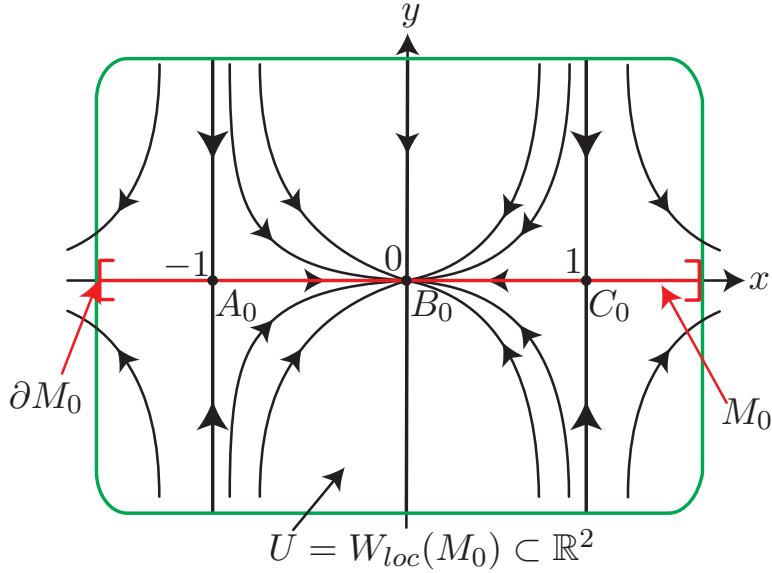


Figure 9.12: An illustration of the trajectories of the dynamical system, the overflowing invariant manifold M_0 , its boundary, and the local stable manifold drawn in green. The fixed points A_0 , B_0 , and C_0 are also shown.

- (i) **$\alpha = 1$** In this case the eigenvalues split in a vertical fashion, as shown in Fig. 9.13. The stable node turns into a stable spiral, which, when connected to the perturbed fixed points A_ε and C_ε cannot be diffeomorphic to M_0 .
- (ii) **$0 < \alpha < 1$** The eigenvalues remain real, but the manifold M_0 does not persist smoothly. To show this, consider a perturbation of the form

$$\begin{cases} \dot{x} = -x(1-x)(1+x) - \varepsilon y \\ \dot{y} = -\alpha y + \varepsilon x^2 \end{cases}; \quad 0 \leq \varepsilon \ll 1. \quad (9.13)$$

The perturbation in the \dot{y} -equation, εx^2 , causes the perturbed fixed points A_ε and C_ε to move upwards, since their y -coordinates are $y_\varepsilon = \frac{\varepsilon}{\alpha} x_\varepsilon^2$. The perturbed curve connecting A_ε , B_ε and C_ε has a cusp at B_ε and cannot be diffeomorphic to M_0 .

Thus in each case, no smooth transformation to M_0 exists and therefore there is no persistence. The perturbed geometries are shown in Fig. 9.14.

The reason for the destruction of the manifold here was the normal rate of attraction being less than or equal to the tangential rate of attraction at one point. This can be exemplified as follows.

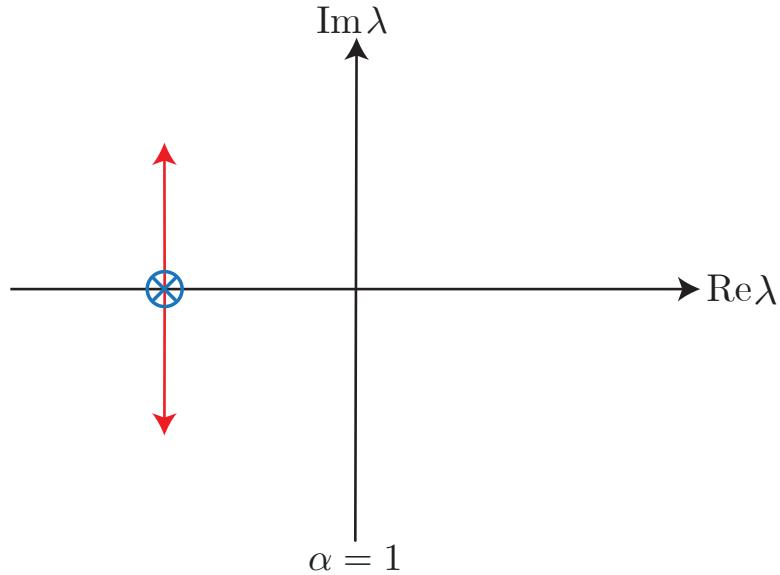


Figure 9.13: The eigenvalues splitting in a vertical fashion for $\alpha = 1$.

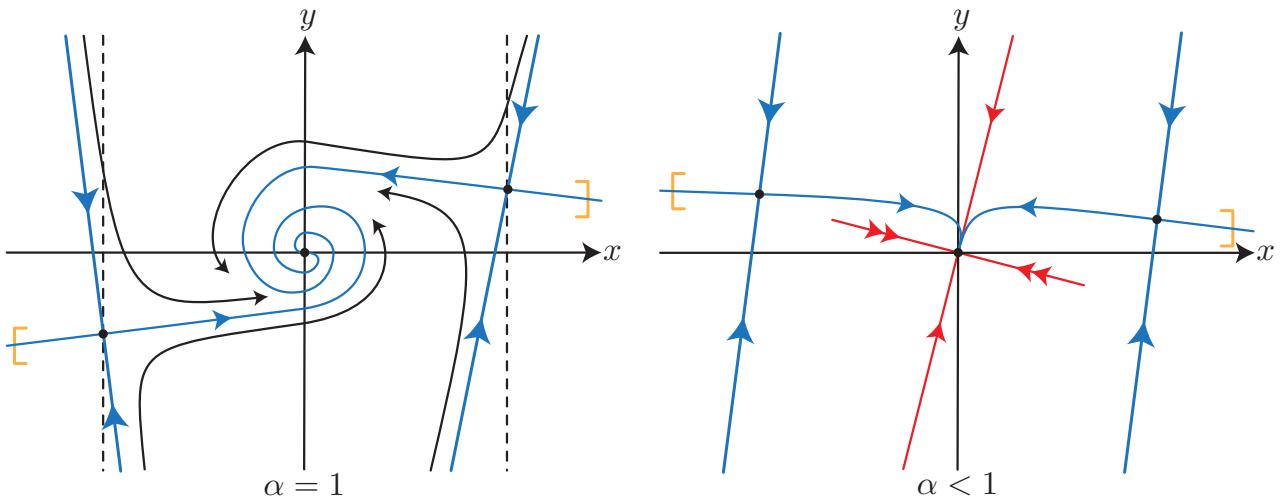


Figure 9.14: Left: The geometry for $\alpha = 1$. Right: The geometry for $\alpha < 1$.

Example 9.5. Consider the dynamical system

$$\begin{cases} \dot{x} = -x(1-x)(1+x) + \varepsilon y \\ \dot{y} = -\frac{3}{2}y - \varepsilon x \end{cases}. \quad (9.14)$$

This is the specification of the previous example to $\alpha = \frac{3}{2}$. If we linearize at the point B_0 for

$\varepsilon = 0$ we find

$$\begin{cases} \dot{x} = -x \\ \dot{y} = -\frac{3}{2}y \end{cases}. \quad (9.15)$$

The solution to the linearized system is

$$\begin{cases} x(t) = x_0 e^{-t} = x_0 (e^{-1})^t \\ y(t) = y_0 e^{-\frac{3}{2}t} = y_0 \left(e^{-\frac{3}{2}}\right)^t \end{cases} \implies y = y_0 \left(\frac{x}{x_0}\right)^{\frac{3}{2}}. \quad (9.16)$$

The eigenvalue constellation and linearized phase portrait, along with the implied nonlinear phase portrait for this system are show in Fig. 9.15. From (9.16) we see that the phase curves of the system are only C^1 because of the power of $\frac{3}{2}$. In the nonlinear system, this implies that the perturbed manifold M_ε exists, but is also only C^1 -smooth. Thus, the ratio between normal and tangential contraction determines the smoothness of M_ε .

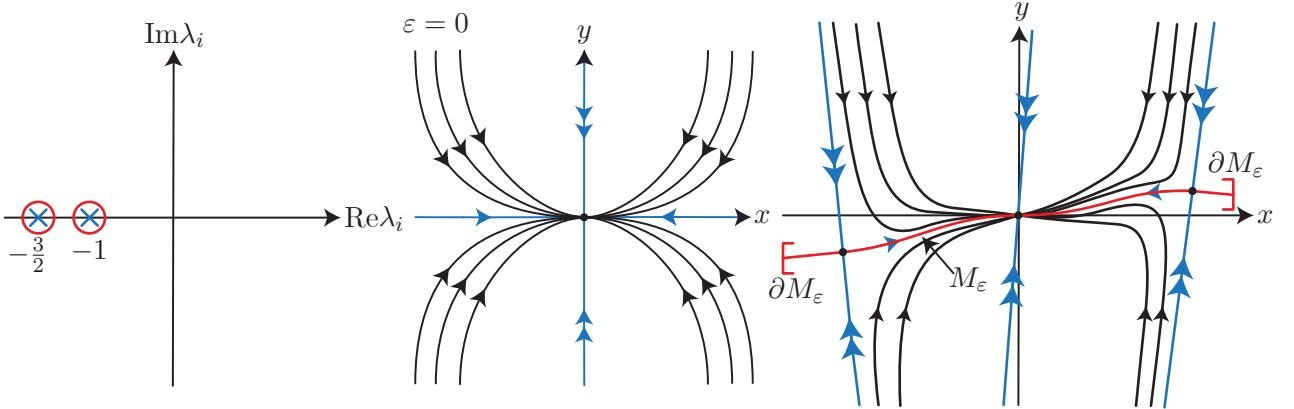


Figure 9.15: Left: The eigenvalue constellation for the linearization around B_0 for $\varepsilon = 0$. Middle: The linearized phase portrait around B_0 . Right: The nonlinear phase portrait with the nearby overflowing invariant manifold M_ε .

One can check here that $\dot{y} = (r+\mu)y + \varepsilon x$ would give C^r differentiability for M_ε for $0 \leq \mu \ll 1$. Formally, for $\mu = 0$ we would have C^∞ differentiability, but only for the linear part, and this would not be robust.

The general challenge is to quantify the growth/contraction rates in the normal and tangential directions, i.e. how do we determine the rates away from fixed points along trajectories.

Recall our setup with the compact, overflowing invariant C^r manifold $M_0 \subset \mathbb{R}^n$ of dimension k , for $r \geq 1$ and the dynamical system

$$\dot{x} = f(x); \quad x \in \mathbb{R}^n; \quad f \in \mathcal{C}^r. \quad (9.17)$$

Now consider the *inverse linearized flow map* on the tangent bundle of \mathbb{R}^n for some $t \geq 0$

$$(F^{-t}, DF^{-t}) : T\mathbb{R}^n \rightarrow T\mathbb{R}^n; (p, v) \mapsto (F^{-t}(p), DF^{-t}(p)v). \quad (9.18)$$

This map is depicted in Fig. 9.16.

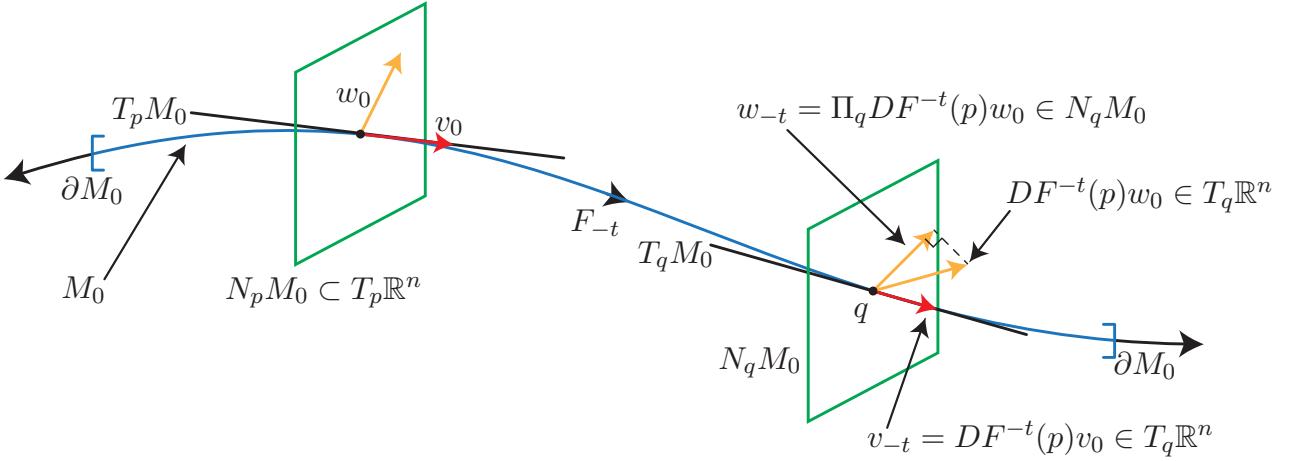


Figure 9.16: The inverse linearized flow map. We have $q = F^{-t}(p) \in M_0$ and the red arrow on the right is $v_{-t} = DF^{-t}(p)v_0 \in T_q M_0$ for $v_0 \in T_p M_0$. The map Π_q is the orthogonal projection onto $N_q M_0$.

To quantify the rate of normal attraction we define for $w_0 \in N_p M_0$, as in Fig. 9.16, $w_{-t} \in N_q M_0$ to be $\Pi_q DF^{-t}(p)w_0$. If we have repulsion in backward time, $\|w_{-t}\| \gg \|w_0\|$, with an exponential rate, i.e.

$$\frac{\|w_0\|}{\|w_{-t}\|} \sim e^{-\alpha t} = (e^{-\alpha})^t; \quad \alpha > 0; \quad t \rightarrow \infty, \quad (9.19)$$

then the best estimate for such an $e^{-\alpha}$ is

$$\nu(p) = \inf \left\{ a \in \mathbb{R}^+ : \lim_{t \rightarrow \infty} \frac{\|w_0\|}{\|w_{-t}\|}/a^t = 0, \forall w_0 \in N_p M_0 \right\}. \quad (9.20)$$

The inf is defined as the largest lower bound. The idea behind this, is that for large a , the quotient will go to zero, if we start to lower a until this trend changes. We then pick the largest lower bound for this change.

If $e^{-\alpha} = a \sim \nu(p) < 1$ then M_0 is exponentially attracting in backward time. Note that the orbit from p will generally not approach a fixed point. Furthermore, $\log(\nu(p))$ will generally not be the eigenvalue of a Jacobian (the linearized flow map at a fixed point).

Next, we would like to quantify the relationship between the normal and tangential rates. If we have

$$\frac{\|w_0\|}{\|w_{-t}\|} / \frac{\|v_0\|}{\|v_{-t}\|} \sim \frac{e^{-\alpha t}}{e^{-\alpha t \cdot \frac{1}{\rho}}} \quad (9.21)$$

for some $\rho \in \mathbb{R}$ and $\alpha > 0$ as $t \rightarrow \infty$ then define for the best asymptotic estimate for $1/\rho$, define $\sigma(p)$ to be

$$\boxed{\sigma(p) = \inf \left\{ b \in \mathbb{R} : \lim_{t \rightarrow \infty} \left(\frac{\|w_0\|}{\|w_{-t}\|} \right)^b / \frac{\|v_0\|}{\|v_{-t}\|} = 0, \forall w_0 \in N_p M_0, \forall v_0 \in T_p M_0 \right\}.} \quad (9.22)$$

The numbers $\nu(p)$ and $\sigma(p)$ are called *Fenichel-type* numbers. They are sometimes also referred to as Lyapunov-type numbers or sometimes simply Type-numbers. If $\sigma(p) \sim \frac{1}{\rho} < \frac{1}{s}$ for some $s \in \mathbb{Z}^+$, then the normal decay to M_0 is s -times stronger than the tangential compression along M_0 , at least along the trajectory through p .

Remark 9.2 (Properties of Fenichel-type numbers). First, each are constant along trajectories, and dependent only on the α -limit set of the trajectory (the set of limit points as $t \rightarrow -\infty$). Also, both $\sigma(p)$ and $\nu(p)$ are upper-semicontinuous in p ($\limsup_{p \rightarrow p_0} \sigma(p) \leq \sigma(p_0)$).

to discuss the computation of the Fenichel-type numbers. To this end, define two functions

$$A_t(p) = DF^{-t}(p)|_{T_p M_0} : T_p M_0 \rightarrow T_q M_0; \quad q = F^{-t}(p); \quad (9.23)$$

$$B_t(p) = \Pi_p DF^t(F^{-t}(p))|_{N_{F^{-t}(p)} M_0} : N_{F^{-t}(p)} M_0 \rightarrow N_p M_0. \quad (9.24)$$

The reason for the $F^{-t}(p)$ being the inner argument for B_t is that we need to step back in order to safely evaluate the linearized map in forward time. With these functions in hand we can calculate the Fenichel-type numbers equipped with the next theorem.

Theorem 9.3 (Fenichel (1971)). *We have that the following holds*

$$\boxed{\nu(p) = \limsup_{t \rightarrow \infty} \|B_t(p)\|^{\frac{1}{t}};} \quad (9.25)$$

and if $\nu(p) < 1$, then

$$\boxed{\sigma(p) = \limsup_{t \rightarrow \infty} \frac{\log(\|A_t(p)\|)}{-\log(\|B_t(p)\|)}}. \quad (9.26)$$

Example 9.6. Consider the dynamical system

$$\begin{cases} \dot{x} = -x(1-x)(1+x) \\ \dot{y} = -by \end{cases}; \quad b > 0. \quad (9.27)$$

Further, define $M_0 = \{(x, y) : x \in [-\frac{3}{2}, \frac{3}{2}], y = 0\}$. Using the previous theorem, it is possible to show

$$\nu(A_0) = \nu(C_0) = e^{-b}; \quad \sigma(A_0) = \sigma(C_0) = -\frac{2}{b}; \quad \nu(B_0) = e^{-b}; \quad \sigma(B_0) = \frac{1}{b}. \quad (9.28)$$

The points A_0 , B_0 , and C_0 along with the manifold M_0 are illustrated in the phase space in Fig. 9.17.

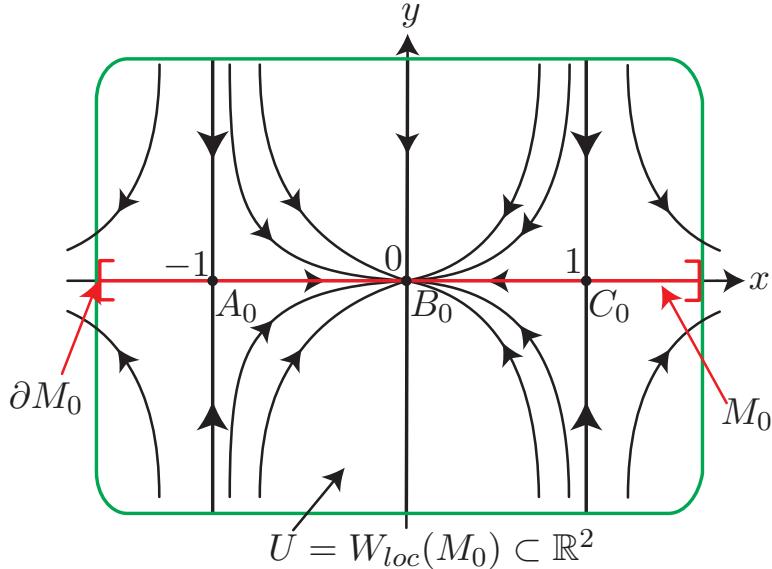


Figure 9.17: A_0 , B_0 , and C_0 along with the manifold M_0 are illustrated in the phase space.

Definition 9.11. The manifold M_0 is an r -normally hyperbolic invariant manifold (NHIM) if

- (i) $\nu(p) < 1$ for all $p \in M_0$;
- (ii) $\sigma(p) < \frac{1}{r}$ for all $p \in M_0$.

Remark 9.4. If there is no tangential compression, then $\sigma(p) < 0$ and (ii) holds automatically for all r .

Building on the previous results, we have the next the next result.

Theorem 9.5 (Fenichel (1971)). *Assume we have*

$$\dot{x} = f_{pert}(x) = f_0(x) + \Delta f(x); \quad f_0, \Delta f \in \mathcal{C}^r, \quad r \geq 1, \quad (9.29)$$

for a compact, overflowing invariant (under $f_0(x)$) manifold M_0 which is \mathcal{C}^r and r -normally hyperbolic. Then for $\|\Delta f\|_{\mathcal{C}^1}$ small enough, there exists an overflowing invariant manifold \tilde{M}_0 for $\dot{x} = f_{pert}(x)$, such that

- (i) \tilde{M}_0 is \mathcal{C}^1 -close to M_0 and \mathcal{C}^r -diffeomorphic to M_0 ;
- (ii) All close enough trajectories converge to \tilde{M}_0 as $t \rightarrow \infty$, as long as they stay in a neighborhood of \tilde{M}_0 .

This result is depicted in Fig. 9.18.

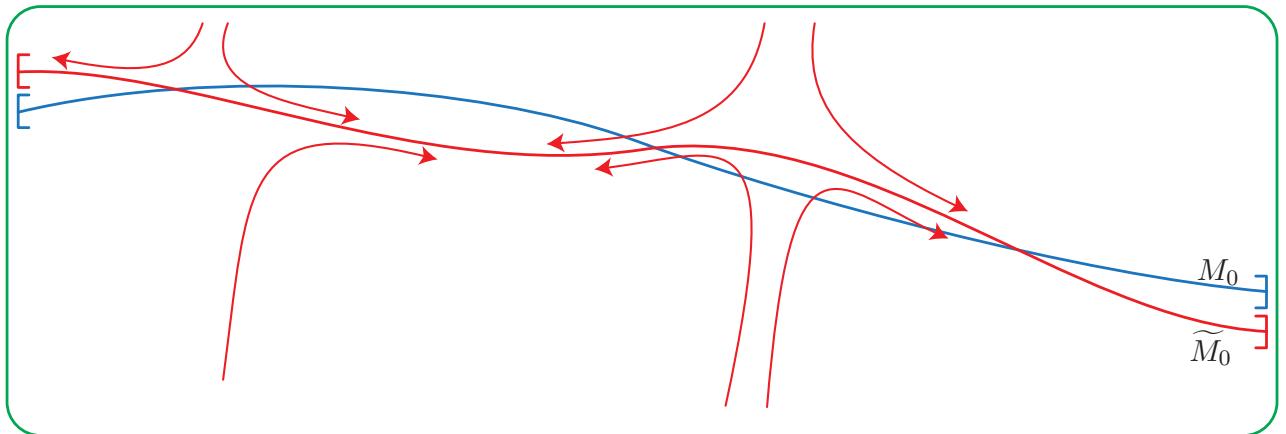


Figure 9.18: The manifolds M_0 and \tilde{M}_0 illustrated within the locally invariant manifold $W_{loc}^S(\tilde{M}_0)$ (green).

Remark 9.6. The following hold:

- (i) If M_0 is compact, normally hyperbolic invariant (without boundary), then it attracts all nearby trajectories (M_0 is an attractor that smoothly persists);
- (ii) The integer r featured in the theorem is the minimum of the integers $r_i \in \mathbb{Z}^+$, where $f \in \mathcal{C}^{r_1}$, $M_0 \in \mathcal{C}^{r_2}$, and $\sigma(p) < \frac{1}{r_3}$;
- (iii) Similar results hold for inflowing invariant manifolds that repel nearby initial conditions (M_0 and \tilde{M}_0 are persistent repellers);

- (iv) The theory also extends to saddle-type invariant manifolds too, e.g. the unstable manifold of an overflowing invariant manifold. In this case assume that the tangent bundle $T\mathbb{R}^n|_{M_0}$ has an invariant splitting, i.e. $T_p\mathbb{R}^n = T_pM_0 \oplus N_p^U M_0 \oplus N_p^S M_0$, then there exists an additional Fenichel-type number for $N^U M_0$ and $W_{\text{loc}(M_0)}^U$ is robust. This is shown in Fig. 9.19.

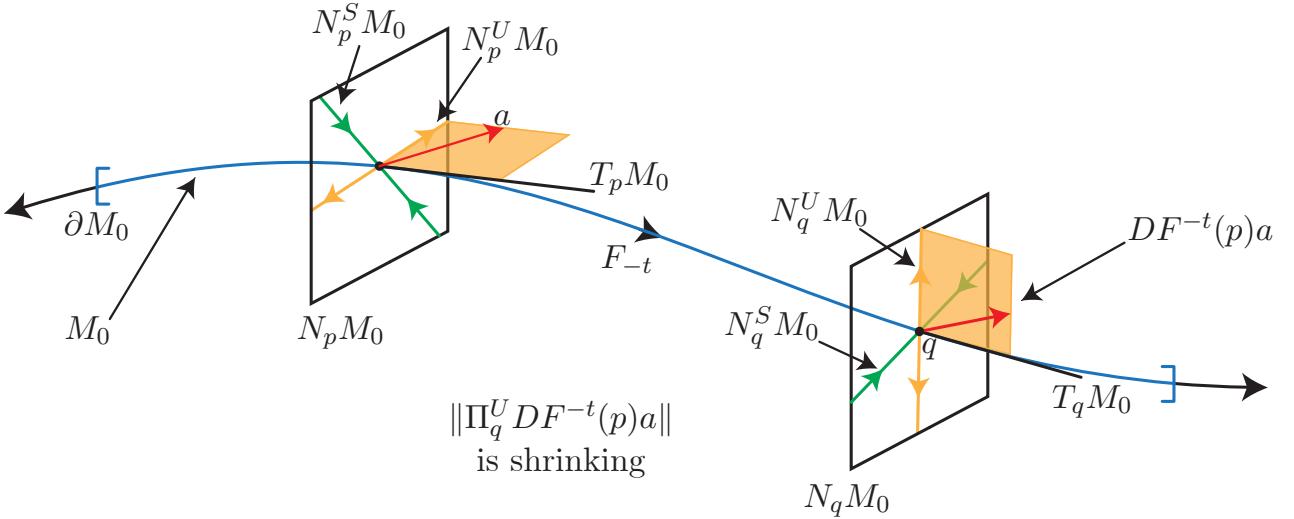


Figure 9.19: The effect of the tangent bundle having an invariant splitting with the addition of a Fenichel-type number for $N^U M_0$ and robust $W_{\text{loc}}^U(M_0)$.

For more see [Wiggins, 1994].

Exercise 9.4. Compute the Lyapunov-type numbers $\nu(p)$ and $\sigma(p)$ in the example

$$\begin{cases} \dot{x} &= -x(1-x^2) \\ \dot{y} &= -by, \end{cases} \quad (9.30)$$

for all points $p \in M_0$, with the parameter $b \in \mathbb{R}^+$ and with overflowing-invariant manifold M_0 defined as

$$M_0 = \{(x, y) \in \mathbb{R}^2 : y = 0, x \in [-3/2, 3/2]\}. \quad (9.31)$$

Hint: Use the operators $A_t(p)$ and $B_t(p)$ defined in class.

9.3 Geometric singular perturbation theory

Singular perturbation theory is the study of systems with two different time scales, i.e. of the form

$$\begin{cases} \dot{x} = f(x, y, \varepsilon) \\ \varepsilon \dot{y} = g(x, y, \varepsilon) \end{cases}; \quad 0 \leq \varepsilon \ll 1; \quad f, g \in \mathcal{C}^1. \quad (9.32)$$

In such systems $x \in \mathbb{R}^n$ is called the *slow variable* and $y \in \mathbb{R}^m$ is the *fast variable*.

Example 9.7 (One degree of freedom mechanical oscillator with a small mass). An example of a system with two different time scales is a single degree of freedom mechanical oscillator with a small mass, the dynamical system is given by

$$\varepsilon \ddot{x} + F(x, \dot{x}, \varepsilon) = 0 \implies \begin{cases} \dot{x} = y \\ \varepsilon \dot{y} = -F(x, y, \varepsilon) \end{cases}. \quad (9.33)$$

Note that $\dot{y} = \frac{1}{\varepsilon} F(x, y, \varepsilon)$ implies that y changes very quickly. The equations for \dot{y} become singular as $\varepsilon \rightarrow 0$, hence we have singular perturbation (the smooth dependence of solutions on ε does not apply).

The classical approach to such systems is to set $\varepsilon = 0$ in (9.32), obtaining

$$\begin{cases} \dot{x} = f(x, y, 0) \\ 0 = g(x, y, 0) \end{cases}. \quad (9.34)$$

This is a mixed differential-algebraic system of equations. We solve $g(x, y, 0) = 0$ for $y = G_0(x)$, expand in ε , and plug this into $\dot{x} = f(x, y, \varepsilon)$. Sometimes this works, however this is not guaranteed. Instead we will follow [Fenichel, 1979].

First, turn this into a regular perturbation problem by introducing the fast time $t = \varepsilon\tau$, i.e. $\tau = \frac{t}{\varepsilon}$ for $0 < \varepsilon \ll 1$. This change of coordinates induces via the chain rule the Jacobian

$$\frac{d(\cdot)}{dt} = \frac{d(\cdot)}{d\tau} \frac{d\tau}{dt}. \quad (9.35)$$

The differential of a function f with respect to t will be denoted by \dot{f} , and is equal to $f'\frac{1}{\varepsilon}$, with f' denoting the differential of f with respect to τ . Hence we obtain the rescaled system

$$\begin{cases} x' = \varepsilon f(x, y, \varepsilon) \\ y' = g(x, y, \varepsilon) \end{cases}. \quad (9.36)$$

Note that (9.36) is a regular perturbation problem from a fictitious limit. The systems (9.32) and (9.36) are not equivalent for $\varepsilon = 0$, but are equivalent for any $\varepsilon > 0$.

We analyze the $\varepsilon = 0$ limit next, i.e.

$$\begin{cases} x' = 0 \\ y' = g(x, y, 0) \end{cases}. \quad (9.37)$$

Assume that there exists a pair (x_0, y_0) such that $g(x_0, y_0, 0) = 0$, i.e. a fixed point of (9.37). Then we have m equations for m unknowns y_0 , by the Implicit Function Theorem a unique nearby zero $y = \varphi_0(x)$ with $\varphi \in \mathcal{C}^1$ exists whenever we have that the following hold

- (i) The matrix $D_y g(x_0, y_0, 0)$ is nonsingular;
- (ii) The derivative $D_x g(x_0, y_0, 0)$ exists near (x_0, y_0) .

In which case we can define the n -dimensional invariant manifold filled with fixed points

$$M_0 = \{(x, y) : y = \varphi_0(x), x \in U, U \text{ compact}, \det [D_y g(x, y, 0)] \neq 0\}. \quad (9.38)$$

This manifold is called the *critical manifold* and an example is shown in Fig. 9.20.

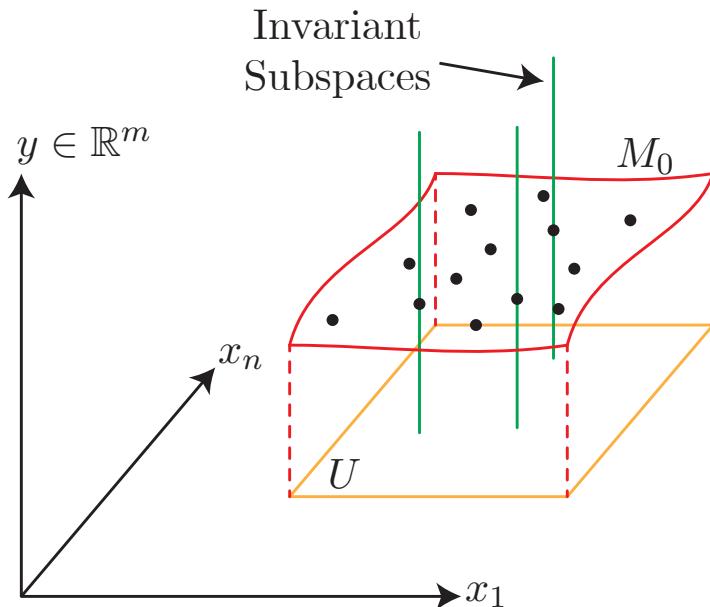


Figure 9.20: An example of the critical manifold M_0 (red) filled with fixed points, an n -dimensional invariant subspace (green line) passes through each fixed point.

Next, we linearize. Begin by defining $\eta = y - \varphi_0(x)$, then calculate the derivative of η with respect to τ

$$\eta' = y' - \underbrace{\varphi'_0 x'}_{=0} = g(x, \varphi_0(x) + \eta, 0) = \underbrace{g(x, \varphi_0(x), 0)}_{=0} + D_y g(x, \varphi_0(x), 0)\eta + \mathcal{O}(\eta^2). \quad (9.39)$$

Thus our linearization is $\eta' = D_y g(x, \varphi_0(x), 0)\eta$, due to the nonsingularity condition before we know that $\lambda_i(D_y g(x, y, 0)) \neq 0$ for $i = 1, \dots, m$. If, in addition, we have that $\text{Re}[\lambda_i(D_y g(x, y, 0))] \neq 0$ for $i = 1, \dots, m$, then M_0 is normally hyperbolic.

The most important case is when M_0 is normally attracting, i.e. when $\text{Re}[\lambda_i(D_y g(x, y, 0))] < 0$ for each $i = 1, \dots, m$. In which case M_0 is r -normally hyperbolic, where if $\varphi_0(x)$ is C^r for any $r \in \mathbb{Z}^+$ then M_0 persists for ε small enough.

The classical formal approach often fails when there exist eigenvalues with positive real parts, i.e. when M_0 is unstable.

For $\varepsilon > 0$ and with eigenvalues of only strictly negative real parts (as before), we get the perturbation M_ε of M_0 . This perturbed manifold is the *slow manifold*, motion on it is of $\mathcal{O}(\varepsilon)$ speed, and in this case it is attracting. We have that for the C^1 function φ_ε (also in ε as long as $g(x, y, \varepsilon)$ is C^1 in ε)

$$y = \varphi_\varepsilon = \varphi_0(x) + \varepsilon \varphi_1(x) + \mathcal{O}(\varepsilon^2). \quad (9.40)$$

Thus y is a smooth graph over the domain U . The perturbed manifold and the representation of y as a smooth graph over U is shown in Fig. 9.21.

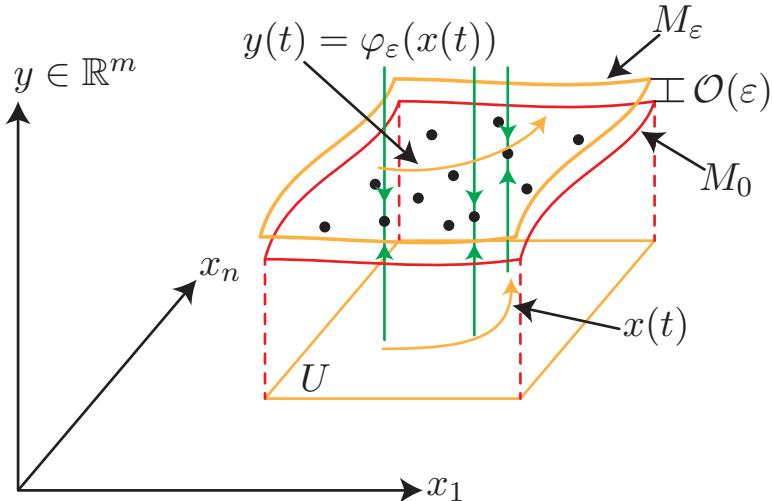


Figure 9.21: The perturbed manifold M_ε near the critical manifold M_0 as well as the representation of y as a smooth graph over the domain U .

We can reduce to the slow manifold and find the reduced flow

$$x'|_{M_\varepsilon} = \varepsilon f(x, y, \varepsilon)|_{M_\varepsilon} = \varepsilon f(x, \varphi_\varepsilon(x), \varepsilon) \quad (9.41)$$

$$= \varepsilon [f(x, \varphi_0(x), 0) + \varepsilon (D_\varepsilon f(x, \varphi_0(x), 0) + D_y f(x, \varphi_0(x), 0)\varphi_1(x))] + \mathcal{O}(\varepsilon^3). \quad (9.42)$$

In the original time t we find the *reduced flow*

$$\dot{x} = f(x, \varphi_0(x), 0) + \varepsilon [D_\varepsilon f(x, \varphi_0(x), 0) + D_y f(x, \varphi_0(x), 0) \varphi_1(x)] + \mathcal{O}(\varepsilon^2). \quad (9.43)$$

The leading order restricted flow on the attractor M_ε is given by $\dot{x} = f(x, \varphi_0(x), 0)$. Any structurally stable feature of this leading-order reduced flow will persist on the attractor. For other features, we need higher-order terms. The dynamics near the slow manifold is attractive towards M_ε if $\text{Re}[\lambda_i(D_y g(x, y, 0))] < 0$ for $i = 1, \dots, m$ for all $(x, y) \in M_0$, i.e. $W_{\text{loc}}^U(M_\varepsilon)$ is empty. Furthermore, we get the stable fibers $f_\varepsilon^S(p)$ which are \mathcal{C}^1 -close to their respective $f_0^S(p)$, and are of the same dimension (m). These facts are illustrated in Fig. 9.22.

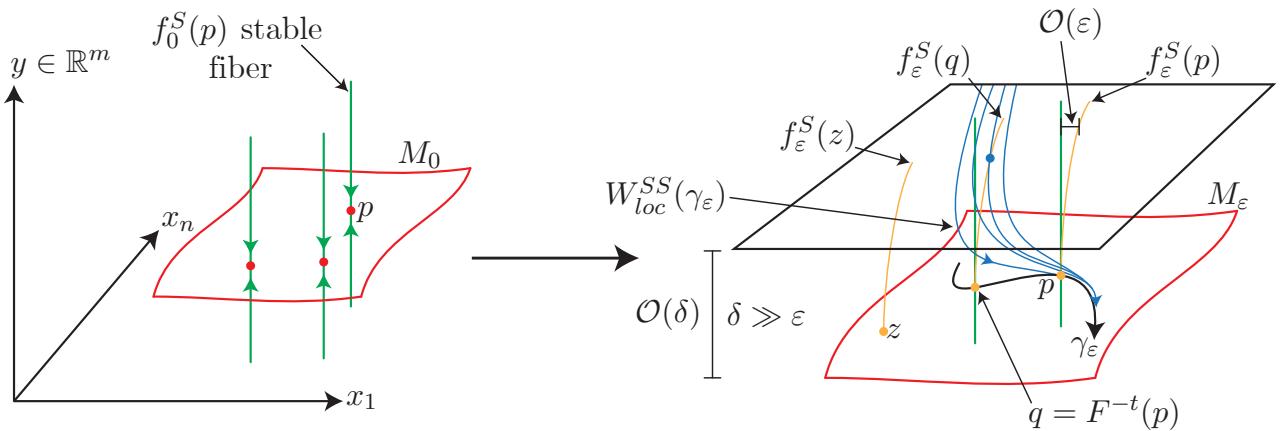


Figure 9.22: The dynamics near the slow manifold and the stable fibers with their relationship to M_0 and the stable fibers of the linearization.

Define the $m+1$ -dimensional *strong stable manifold* of γ_ε to be

$$W_{\text{loc}}^{SS}(\gamma_\varepsilon) = \bigcup_{p \in \gamma_\varepsilon} f_\varepsilon^S(p), \quad (9.44)$$

i.e. the collection of fastest-converging trajectories to γ_ε . By construction, the stable fibers form an invariant family along γ_ε , i.e. $F^t(f_\varepsilon^S(q)) \subset f_\varepsilon^S(p)$.

Example 9.8 (Clustering of inertial particles in fluids). When finite-size inertial particles are suspended in fluids they tend to form clusters, even in incompressible fluids (as opposed to fluid particles). For a specific case, consider the steady flow in the wake of a bluff body [Burns et al., 1999]. This can be seen in Fig. 9.23.

The notation here will be $x \in \mathcal{S}^1 \times \mathbb{R}$ for the particle position in a moving frame and periodic in the first variable. The steady fluid velocity in the moving frame is given by $u(x) \in \mathbb{R}^2$, the nondimensionalized particle mass is given by $\varepsilon \ll 1$, and the velocity of the inertial particle

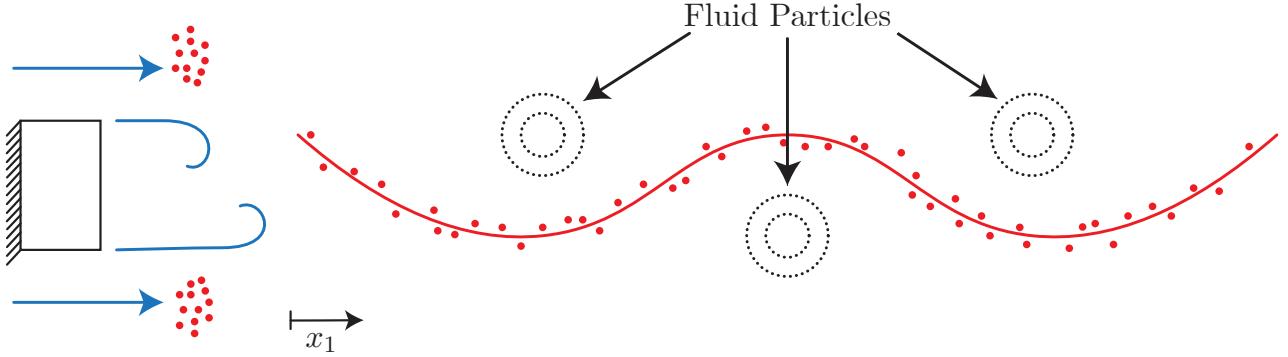


Figure 9.23: The clustering of inertial particles observed experimentally, referred to as the *von Kármán vortex street*.

at time t given by $y = \dot{x} \in \mathbb{R}^2$. The equation of motion in its simplest form is given by [Maxey and Riley, 1983]

$$\begin{cases} \dot{x} &= y \\ \varepsilon \dot{y} &= u(x) - y \end{cases}. \quad (9.45)$$

Thus we have the slow variable x and the fast variable y , performing the time transformation as before ($t = \varepsilon\tau$) we obtain

$$\begin{cases} x' = \varepsilon y \\ y' = u(x) - y \end{cases}. \quad (9.46)$$

the $\varepsilon = 0$ limit yields

$$\begin{cases} x' = 0 \\ y' = u(x) - y = g(x, y, 0) \end{cases}. \quad (9.47)$$

Hence, we get the critical manifold

$$M_0 = \{(x, y) \in U \subset \mathcal{S}^1 \times \mathbb{R}^3 : y = u(x)\}. \quad (9.48)$$

To examine the stability of M_0 calculate $D_y g(x, u(x), 0) = -I_2$, hence $\text{Re}[\lambda_i] = \lambda_i = -1 < 0$ for $i = 1, 2$. Hence, M_0 is normally hyperbolic and attracting, the unperturbed fibers $f_0^S(p)$ for $p \in M_0$ are 2-dimensional planes. The unperturbed geometry is depicted in Fig. 9.24.

For $\varepsilon > 0$ we get the perturbed manifold M_ε which is $\mathcal{C}^1 \mathcal{O}(\varepsilon)$ -close to M_0 . Trajectories in $W^{SS}(\gamma_\varepsilon)$ quickly synchronize with γ_ε . The perturbed geometry is shown in Fig. 9.25.

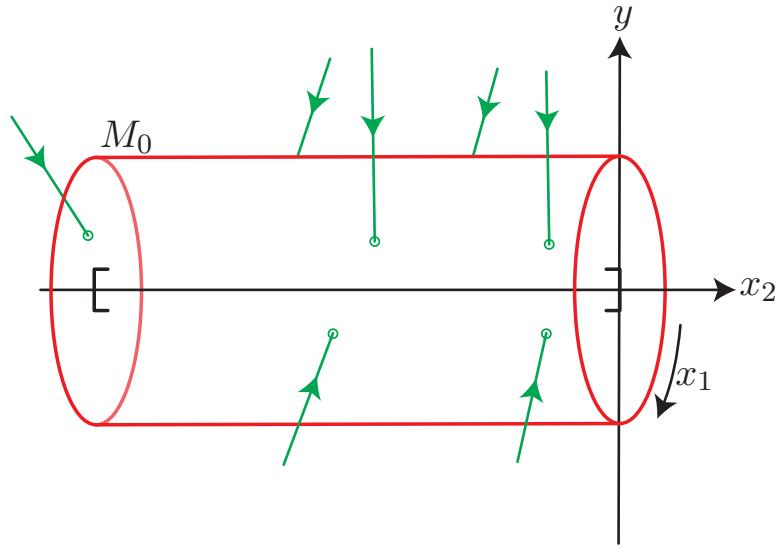


Figure 9.24: The unperturbed geometry of the dynamics given in the example of inertial particles suspended in the fluid.

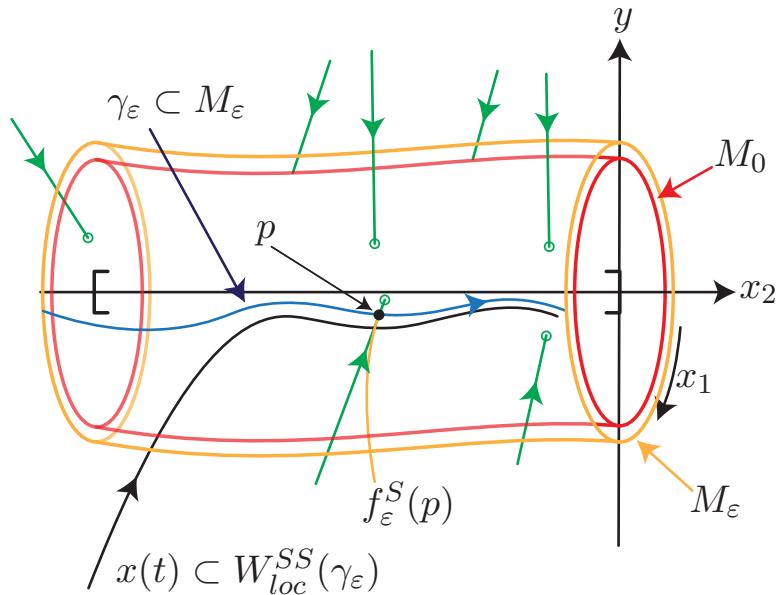


Figure 9.25: The perturbed geometry of the dynamics given in the example of inertial particles suspended in the fluid.

To study the dynamics on the slow manifold, write the perturbed manifold M_ε as the graph

of the function φ_ε

$$M_\varepsilon = \left\{ (x, y) : y = \varphi_\varepsilon(x) = \underbrace{\varphi_0(x)}_{u(x)} + \varepsilon h_1(x) + \mathcal{O}(\varepsilon^2) \right\}. \quad (9.49)$$

By the invariance of M_ε the formula $y = \varphi_\varepsilon(x)$ holds along full trajectories. Plugging this in and differentiating with respect to τ we find

$$y' = (Du)x' + \varepsilon(Dh_1)x' + \dots = (Du)\varepsilon u(x) + \mathcal{O}(\varepsilon^2). \quad (9.50)$$

In the second equation the fact $x' = \varepsilon\varphi_\varepsilon(x)$ was used. Also from the equation of motion we find

$$y'|_{M_\varepsilon} = (u(x) - y)|_{M_\varepsilon} = u(x) - (u(x) + \varepsilon h_1(x) + \dots) = -\varepsilon h_1(x) + \mathcal{O}(\varepsilon^2). \quad (9.51)$$

By comparing coefficients we then find

$$h_1 = -(Du)u. \quad (9.52)$$

Hence, the reduced flow on the slow manifold is

$$x' = \varepsilon\varphi_\varepsilon(x) = \varepsilon(u(x) - \varepsilon(Du(x))u(x)) + \mathcal{O}(\varepsilon^2). \quad (9.53)$$

Transforming this back into the original time we find

$$\dot{x} = u(x) - \varepsilon(Du(x))u(x) + \mathcal{O}(\varepsilon^2). \quad (9.54)$$

The $\mathcal{O}(1)$ term, $u(x)$, denotes the incompressible terms, meanwhile the $\mathcal{O}(\varepsilon)$ term is the dissipative terms (i.e. has nonzero divergence). This nonzero divergence can be calculated

$$\nabla \cdot [(Du)u] = \text{Tr}[DuDu] = -2Q; \quad Q = \frac{1}{2} |\|\Omega\|^2 - \|S\|^2|, \quad (9.55)$$

where $\|\cdot\|$ denotes the Euclidean matrix norm. The unperturbed system and $Q > 0$ case are depicted in Fig. 9.26.

Exercise 9.5. Consider now the three-dimensional nonlinear dynamical system

$$\begin{cases} \dot{x} = -\varepsilon(x + y^2) \\ \dot{y} = -y \\ \dot{z} = z, \end{cases} \quad (9.56)$$

with the small parameter $\varepsilon \geq 0$.

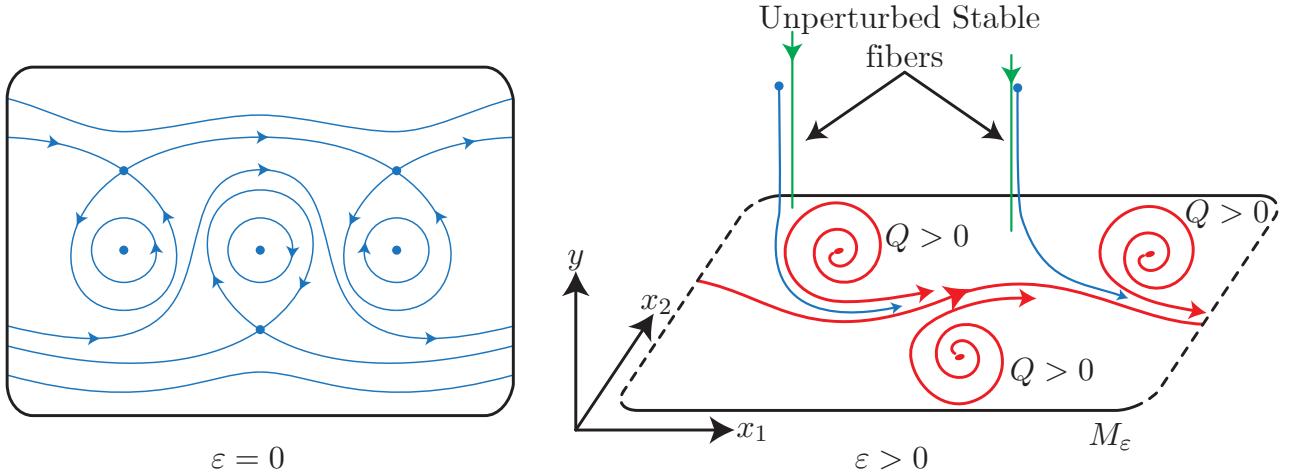


Figure 9.26: Left: The unperturbed phase space. Right: The perturbed phase space for $Q > 0$, the green arrows denotes unperturbed stable fibers, and the red line designates the limit cycle. By identifying the left and right sides, we get a periodic variable.

- (i) Show that the set $M_0 = \{y = z = 0, x \in [-1, 1]\}$ is a normally hyperbolic invariant manifold for $\varepsilon = 0$.
- (ii) Find the manifold M_ε into which M_0 perturbs for small $\varepsilon > 0$.
- (iii) Using the property (v), show that for any base point $p \in W_{\text{loc}}^s(M)$, the corresponding stable fiber is the nonlinear surface.

$$f^s(p) = \left\{ (x, y, z) \mid x = x_p + \frac{\varepsilon}{2 - \varepsilon} y^2, z = 0 \right\}. \quad (9.57)$$

- (iv) Find a similar expression for the unstable fibers $f^u(p)$.

A more general approach to particle motion is given by [Haller and Sapsis, 2008] using the Maxey-Riley equations. For neutrally buoyant particles (ones which are suspended in the fluid) we have

$$\varepsilon \dot{v} - \varepsilon \frac{Du}{Dt} = -(v - u); \quad \varepsilon = \frac{3St}{2} \ll 1, \quad (9.58)$$

where St denotes the Stokes number. Formulating this in terms of geometric singular pertur-

bation theory, with $x \in \mathbb{R}^d$ ($d = 2, 3$) we get

$$\begin{cases} \dot{x} = y \\ \varepsilon \dot{y} = u(x, t) - y + \varepsilon \frac{Du(x, t)}{Dt} \end{cases} \implies \begin{cases} t' = \varepsilon \\ x' = \varepsilon y \\ y' = u(x, t) - y + \varepsilon \frac{Du(x, t)}{Dt} \end{cases}. \quad (9.59)$$

The time transformation used was $t = t_0 + \varepsilon\tau$. From here we have the slow variables x and t , and the fast variable y . There then exists a $(d+1)$ -dimensional slow manifold M_ε where we have $y = u(x, t) + \varepsilon u_1(x, t) + \mathcal{O}(\varepsilon)$. The reduced particle dynamics here is given by

$$y' = \nabla u x' + u_t \varepsilon + \varepsilon \nabla u_1 x' + u_{1t} \varepsilon^2 + \mathcal{O}(\varepsilon^2) = \varepsilon (\nabla u y + u_t) + \mathcal{O}(\varepsilon^2) \quad (9.60)$$

$$= \varepsilon (\nabla u u + u_t) + \mathcal{O}(\varepsilon^2) = \varepsilon \frac{Du}{Dt} + \mathcal{O}(\varepsilon^2). \quad (9.61)$$

Also, we have

$$y' = u - (u + \varepsilon u_1 + \mathcal{O}(\varepsilon^2)) + \varepsilon (u_t + u_x u) \quad (9.62)$$

$$= -\varepsilon u_1 + \varepsilon \frac{Du}{Dt} + \mathcal{O}(\varepsilon^2). \quad (9.63)$$

By comparing terms we see that $u_1 = 0$, in fact this holds up to any order.

The reduced dynamics on the slow manifold follow $\dot{x} = v = u(x, t)$ up to any order independent of ε . This suggests a fast synchronization with the fluid, yet scattering of the particles is still observed along fluid motions. The reason for this can be seen by rewriting (9.58)

$$\dot{v} - u_t - \nabla u u = \mu(u - v); \quad \dot{v} - u_t - \nabla u v = \mu(u - v) + \nabla u(u - v). \quad (9.64)$$

This implies that

$$\begin{cases} \frac{d}{dt}(v - u(x, t)) = -(\nabla u + \mu I)(v - u) \\ \frac{d}{dt}x = v \end{cases}. \quad (9.65)$$

These in turn imply that $v = u(x, y)$ is an invariant manifold for any $\varepsilon = \frac{1}{\mu}$, but we only know its stability for $\varepsilon \ll 1$. This means that different parts of the manifold may have different stability types.

To analyze the stability in more detail set $z = v - u(x, y)$, i.e. the distance from M_ε in the y -direction. The dynamics with this variable now is

$$\begin{cases} \dot{z} = -[\nabla u(x, t) + \mu I] z \\ \dot{x} = z + u(x, t) \end{cases} \implies \frac{1}{2} \frac{d}{dt} \|z\|^2 = -\langle z, (\nabla u + \mu I)z \rangle = -\langle z, (S + \mu I)z \rangle, \quad (9.66)$$

Where $S = \frac{1}{2} (\nabla u + \nabla u^T) = S^T$ is the rate of strain tensor.

We can derive conditions now for guaranteed attraction or repulsion:

(i) **Exponential attraction** is guaranteed when we have

$$\frac{1}{2} \frac{d}{dt} \|z\|^2 \leq \lambda_{\max} [-(S + \mu I)] \|z\|^2. \quad (9.67)$$

By the Gronwall-inequality, this implies

$$\|z(t)\| \leq \|z(t_0)\| \exp \left(\int_{t_0}^t \lambda_{\max} [-(S + \mu I)] ds \right). \quad (9.68)$$

Thus we have that $\|z(t)\|$ decays as long as we have

$$\lambda_{\max} [-S(x(t), y(t), t) - \mu I] < 0 \Leftrightarrow \lambda_{\min} [S(x, t) + \mu I] > 0 \Leftrightarrow \mu > \sqrt{|\det(S)|}. \quad (9.69)$$

(ii) **Exponential repulsion** is guaranteed when we have

$$\frac{1}{2} \frac{d}{dt} \|z\|^2 \geq \lambda_{\min} [-(S + \mu I)] \|z\|^2. \quad (9.70)$$

Thus we have that $\|z(t)\|$ grows instantaneously when

$$\lambda_{\max} [-(S + \mu I)] > 0 \implies \lambda_{\max} [S + \mu I] < 0. \quad (9.71)$$

More generally, an attracting/repelling invariant manifold may have localized repulsion/attraction along it, even if all trajectories approach it asymptotically [Sapsis and Haller, 2010].

To study the local repulsion/attraction, consider the following dynamical system

$$\dot{x} = f(x, t); \quad x \in \mathbb{R}^n, \quad (9.72)$$

with the flow map $F_{t_0}^t : x_0 \mapsto x(t; t_0, x_0)$ and a locally invariant k -dimensional manifold $\mathcal{M}(t) \subset \mathbb{R}^n$, i.e. $F_{t_0}^t(\mathcal{M}(t_0)) \subset \mathcal{M}(t)$. Now define the *normal infinitesimal Lyapunov exponent* $\sigma(p, t)$ (abbreviated NILE) by taking the $t \rightarrow t_0$ limit and looking at the largest growth for a tangent vector w_0 , i.e.

$$\sigma(p, t) = \lim_{s \rightarrow 0} \frac{1}{s} \log \left\| \Pi_{F_t^{t+s}(p)}^{t+s} DF_t^{t+s}(p) \Big|_{N_p \mathcal{M}(t)} \right\|. \quad (9.73)$$

The graphical depiction of the terms used for the calculation of the normal infinitesimal Lyapunov exponent are depicted in Fig. 9.27.

To compute the NILE we can use the following formula

$$\sigma(p, t) = \max_{n \in N_p \mathcal{M}(t), |n|=1} \langle n, S(p, t)n \rangle; \quad S = \frac{1}{2} (Df + Df^T). \quad (9.74)$$

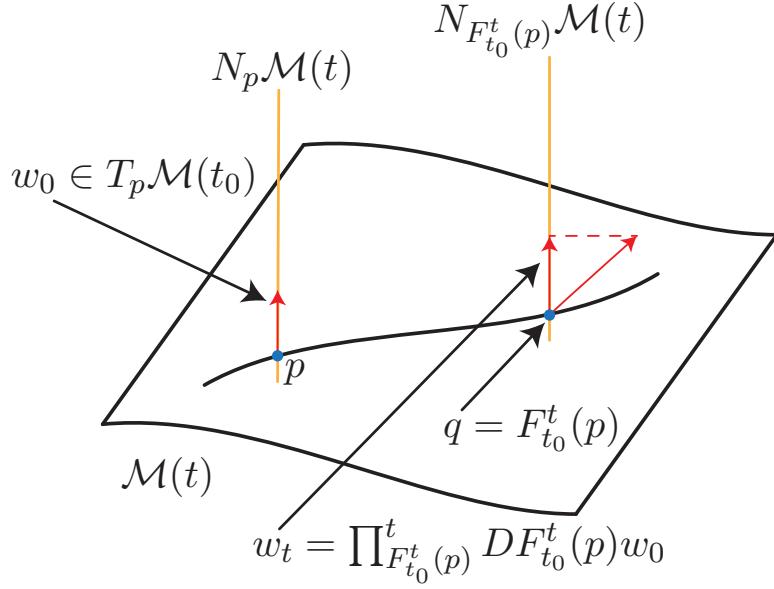


Figure 9.27: The geometry of tangent vectors along the invariant manifold M . The quantities used to define the NILE (9.73) are indicated in the figure.

Let us express the invariant manifold as a graph over k variables, i.e. write x as $x = (a, b)$, where $a \in \mathbb{R}^k$ and $b \in \mathbb{R}^{k-n}$, $f = (A, B)$ and on the invariant manifold M we have $b = h(a)$. We can then write

$$\sigma(a, t) = \frac{1}{2} \lambda_{\max} [\Gamma(a, t) + \Gamma(a, t)^T]; \quad (9.75)$$

where we have defined the tensors

$$\Gamma(a, t) = B_b(a, h(a), t) - Dh(a, t)A_b(a, h(a), t). \quad (9.76)$$

Example 9.9 (Neutrally buoyant particles). Consider the dynamical system with $x, z \in \mathbb{R}^2$

$$\begin{cases} \dot{x} = z + u(x, t) \\ \dot{z} = -(\nabla u(x, t) + \mu I)z \end{cases}; \implies \mathcal{M}_0(t) \{(x, z) \in \mathbb{R}^2 \times \mathbb{R}^2 : z = 0\}. \quad (9.77)$$

Therefore we have

$$\Gamma(a, t) = B_b(x, 0, t) = -(\nabla u(x, t) + \mu I). \quad (9.78)$$

With this we can calculate the NILE

$$\sigma(p, t) = \lambda_{\max} [-(S(x, t) + \mu I)]; \quad S = \frac{\nabla u + \nabla u^t}{2}. \quad (9.79)$$

As concluded before, where $\sigma(p, t) > 0$ we have a region of local repulsion, analogously where $\sigma(p, t) < 0$ we have a region of local attraction.

Exercise 9.6. Consider the slowly forced pendulum with viscous damping

$$\ddot{\varphi} + k\dot{\varphi} + \sin \varphi = F_0 \sin \varepsilon t,$$

where $k > 0$ is the damping coefficient, $0 < F_0 < 1$ is the forcing amplitude, and $0 < \varepsilon \ll 1$ is the forcing frequency. Give a complete qualitative description of the geometry of this system in the extended phase space for small enough ε . (Hint: Make the system autonomous by introducing the phase variable $\psi = \varepsilon t$.)

Exercise 9.7. Thermohaline circulation (THC) is a part of the large-scale ocean circulation that is driven by global density gradients created by surface heat and freshwater fluxes. Stommel's box model (1961) is a qualitative description of the trends and equilibria in THC. This model couples the two fundamental drivers of THC, temperature (thermo) and salt concentration (-haline), in a nonlinear fashion (see Fig. 9.7).

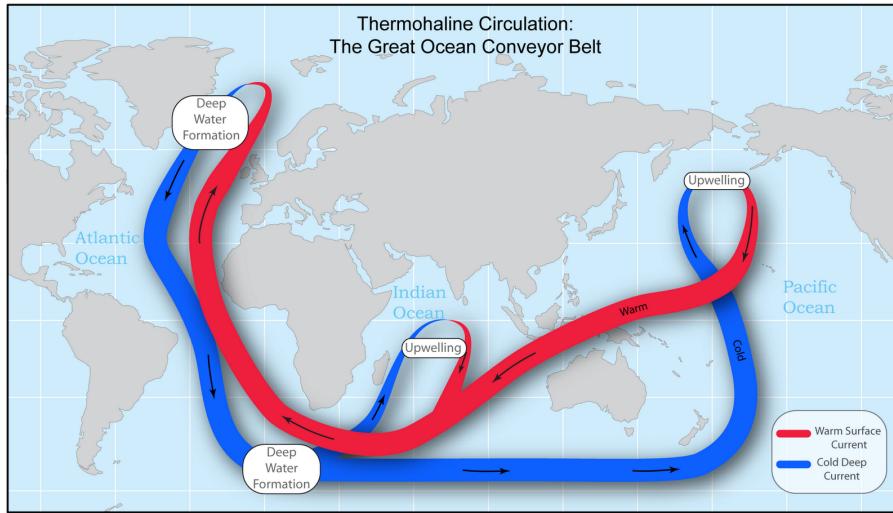


Figure 9.28: Depiction of thermohaline circulation.

The non-dimensional variables of Stommel's model are:

$x(t)$: temperature difference between the tropics (lower latitudes) and the North-Atlantic (higher latitudes);

$y(t)$: salinity (i.e., salt concentration) difference between the above two regions of the ocean.

The non-dimensional parameters of the model are:

τ_x : relaxation time to a constant temperature difference between northern and southern latitudes in the absence of coupling

τ_y : relaxation time to zero salinity difference between higher and lower latitudes in the absence of coupling. In practice, $\tau_x/\tau_y = \epsilon \ll 1$.

μ : measure of freshwater flux through clouds moving from lower to higher latitudes

η : nonlinear coupling parameter between temperature and salinity evolution

With this notation, Stommel's model can be written as

$$\begin{cases} \dot{x} &= -\frac{1}{\tau_x}(x-1) + \frac{1}{\tau_y}x[1+\eta^2(x-y)^2] \\ \dot{y} &= \frac{\mu}{\tau_y} - \frac{1}{\tau_y}y[1+\eta^2(x-y)^2]. \end{cases} \quad (9.80)$$

- (i) Show that Stommel's model has a globally attracting slow manifold that governs the asymptotic behavior of THC. Find a leading order approximation to this manifold. (Hint: rescale time by letting $s = t/\tau_y$.)
- (ii) Compute the leading-order reduced flow on the slow manifold. Determine qualitatively the possible dynamical behaviors on the slow manifold as the parameters μ and η are varied.

9.4 Existence of invariant manifolds

We now turn to the various methods to prove the existence of an invariant manifold:

- (i) **Lyapunov-Perron Method** The manifold is the solution of a global integral equation, this approach requires global coordinates;
- (ii) **Hadamard's Method** The manifold is sought as an invariant graph, this was the approach Fenichel used and local coordinates are sufficient for this approach;
- (iii) **Topological Methods** There are multiple topological approaches, for instance the Wasowsky principle.

Next, the Lyapunov-Perron method will be sketched for stable manifolds on the plane. Consider the dynamical system

$$\dot{x} = f(x); \quad x \in \mathbb{R}^2; \quad f(0) = 0; \quad \text{spect}[Df(0)] = \{-\lambda_S, \lambda_U\}; \quad \lambda_S, \lambda_U > 0. \quad (9.81)$$

The objective is to construct a stable manifold tangent to E_S . This will be achieved in multiple steps.

- (i) First, choose a good basis

$$x = Ty; \quad T = (e_S \ e_U); \quad y = \begin{pmatrix} y_S \\ y_U \end{pmatrix}. \quad (9.82)$$

And implement this transformation

$$\dot{x} = Df(0)x + \mathcal{O}(\|x\|^2) \implies T\dot{y} = Df(0)Ty + \mathcal{O}(\|y\|^2). \quad (9.83)$$

By multiplying from the left by T^{-1} we find

$$\dot{y} = \underbrace{T^{-1}DF(0)T}_{\text{diag}[-\lambda_S, \lambda_U]} y + \mathcal{O}(\|y\|^2); \quad \lim_{y \rightarrow 0} \frac{\|g(y)\|}{\|y\|^2} < \infty. \quad (9.84)$$

And now we find

$$\begin{cases} \dot{y}_S = -\lambda_S y_S + g_S(y) \\ \dot{y}_U = \lambda_U y_U + g_U(y) \end{cases}. \quad (9.85)$$

- (ii) Second, define the local stable manifold. The classic definition of $W^S(0)$ is that it is the collection of initial coordinates y_0 which under the limit of the flow map have $\lim_{t \rightarrow \infty} F_t(y_0) = 0$. This requires detailed knowledge of asymptotics, and thus is hard to work with. For linear systems an alternative definition is, that $W^S(0)$ is the collection of forward-bounded trajectories, but this is generally nothing special after the addition of nonlinear terms. However, we may change the nonlinear system away from the origin to ensure that boundedness remains a distinguishing property.

To go about this we separate the phase space into the rectangular area (of half-width δ) around the origin and the rectangular annulus of outer half-width 2δ . In the inner rectangle, the original nonlinear dynamics hold, outside of the annulus the dynamics are 0 (i.e. $f(x) = 0$). Interpolating in between these regimes takes place on the annulus using a so-called *bump function* Ψ_δ , a smooth function which is equal to 1 on the interval $[-\delta, \delta]$ and equal to 0 on the intervals $(-\infty, -2\delta)$ and $(2\delta, \infty)$, in between these regions

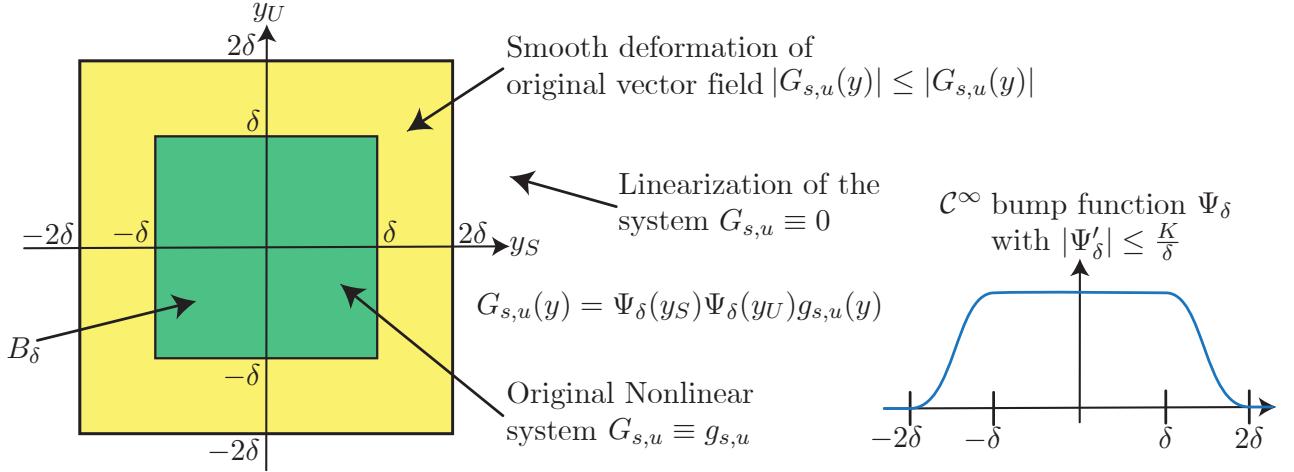


Figure 9.29: The interpolated dynamical system, within the green area the original nonlinear dynamics hold, on the yellow area the smoothly deformed (interpolated) system holds, and outside of these two regions, the null dynamics hold. The bump function is also depicted.

is the interpolation. The bump function and the interpolated new dynamical system are depicted in Fig. 9.29.

Thus we have the modified system

$$\begin{cases} \dot{y}_S = -\lambda_S y_S + G_S(y) \\ \dot{y}_U = \lambda_U y_U + G_U(y) \end{cases}. \quad (9.86)$$

Furthermore, we can bound this system

$$|G'_{S,U}(y)| = \Psi'_\delta g_{S,U} + \Psi_\delta g'_{S,U} \leq |\Psi_\delta| |g_{S,U}| + |\Psi_\delta| |g'_{S,U}| \quad (9.87a)$$

$$\leq \frac{K}{\delta} C \delta^2 + 1 \cdot 2\delta \leq C_0 \delta. \quad (9.87b)$$

Then our goal becomes to find globally bounded solutions in the modified system, i.e.

$$W_{\text{loc}}^S(0) = \left\{ y_0 \in B_\delta : \sup_{t \geq 0} |\mathcal{F}^t(y_0)| < \infty \right\}. \quad (9.88)$$

- (iii) Now pretend that $y(t)$ is known to find the *equivalent integral equations*. The variation of constants formula yields

$$y_S(t) = y_S(t_S) e^{-\lambda_S(t-t_S)} + \int_{t_S}^t e^{-\lambda_S(t-s)} G_S(y(s)) ds \quad (9.89a)$$

$$y_U(t) = y_U(t_U) e^{\lambda_U(t-t_U)} + \int_{t_U}^t e^{\lambda_U(t-s)} G_U(y(s)) ds. \quad (9.89b)$$

By taking $t_U \rightarrow \infty$ and $t_S = 0$ we restrict to $W_{\text{loc}}^S(0)$, thus we find the integral equations for $W_{\text{loc}}^S(0)$

$$\begin{cases} y_S(t) = e^{-\lambda_S t} y_S(0) + \int_0^t e^{-\lambda_S(t-s)} G_S(y(s)) ds \\ y_U(t) = \int_\infty^t e^{\lambda_U(t-s)} G_U(y(s)) ds \end{cases}. \quad (9.90)$$

- (iv) Finally we have a fixed point problem, the equation (9.90) is fulfilled when $y(t) = \mathcal{F}(y(t); y_S)$, thus we have a fixed point problem with the initial condition as a parameter. There exists a unique solution if \mathcal{F} is a contraction mapping from a complete metric space into itself, i.e. when we can apply the Banach fixed point theorem. For \mathcal{F} to be contracting, we require

$$d(\mathcal{F}(x), \mathcal{F}(y)) \leq q d(x, y); \quad 0 < q < 1, \quad (9.91)$$

for the metric d defined over the underlying metric space.

Define the set

$$\mathcal{B}_\delta = \{y(t) : [0, \infty) \rightarrow \mathbb{R}^2 : y_S(0) = \delta, \|y\| \leq K, y \in \mathcal{C}^0[0, \infty)\}, \quad (9.92)$$

where the norm is $\|y\| = \sup_{t \geq 0} \|y(t)\|$. The space can be equipped with the metric

$$d(x, y) = \sup_{t \geq 0} \|x(t) - y(t)\| = \|y - x\|, \quad (9.93)$$

and is complete with respect to this metric by the uniform convergence of Cauchy sequences of \mathcal{C}^0 functions to \mathcal{C}^0 functions, with limits obeying the same bound. We also have to show that \mathcal{F} maps into itself, i.e. that $\|\mathcal{F}(y)\| \leq K$ when $\|y\| \leq K$. To this end, calculate

$$\|y(t)\| \leq |y_S(t)| + |y_U(t)| \quad (9.94a)$$

$$\leq e^{-\lambda_S t} |y_S(0)| + \int_0^t e^{-\lambda_S(t-s)} |G_S(y(s))| ds + \int_t^\infty e^{\lambda_U(t-s)} |G_U(y(s))| ds \quad (9.94b)$$

$$\leq \delta + e^{-\lambda_S t} \underbrace{\int_0^t e^{\lambda_S s} |G_S(y(s))| ds}_{\leq C\delta^2} + e^{\lambda_U t} \underbrace{\int_t^\infty e^{-\lambda_U s} |G_U(y(s))| ds}_{\leq C\delta^2} \quad (9.94c)$$

$$\leq \delta + C\delta^2 \frac{1}{\lambda_S} \left[e^{-\lambda_S(t-s)} \right]_0^t - \frac{C\delta^2}{\lambda_U} \left[e^{\lambda_U(t-s)} \right]_t^\infty \quad (9.94d)$$

$$= \delta + C\delta^2 \left(\frac{1}{\lambda_S} (1 - e^{-\lambda_S t}) + \frac{1}{\lambda_U} (1 - 0) \right) \quad (9.94e)$$

$$\leq \delta + C\delta^2 \left(\frac{1}{\lambda_S} + \frac{1}{\lambda_U} \right) \leq K. \quad (9.94f)$$

This shows that $y(t) \in \mathcal{B}_\delta$ for δ small enough. Furthermore, we have that \mathcal{F} is a contraction mapping for δ small enough.

$$\|\mathcal{F}(y) - \mathcal{F}(x)\| \leq \int_0^t e^{-\lambda_S(t-s)} |G_S(x(s)) - G_S(y(s))| ds \quad (9.95)$$

$$+ \int_t^\infty e^{\lambda_U(t-s)} |G_U(x(s)) - G_U(y(s))| ds \quad (9.96)$$

$$\leq \int_0^t e^{-\lambda_S(t-s)} C_0 \delta \|x(s) - y(s)\| ds \quad (9.97)$$

$$+ \int_t^\infty e^{\lambda_U(t-s)} C_0 \delta \|x(s) - y(s)\| ds \quad (9.98)$$

$$\leq C_0 \delta \left(\frac{1}{\lambda_S} + \frac{1}{\lambda_U} \right) \|x - y\|, \quad (9.99)$$

where in (9.98) the mean value theorem was used. This inequality implies that for δ small enough,

$$\|\mathcal{F}(y) - \mathcal{F}(x)\| \leq \underbrace{C_0 \delta \left(\frac{1}{\lambda_S} + \frac{1}{\lambda_U} \right)}_{\leq q < 1} \|x - y\|. \quad (9.100)$$

For the general, multi-dimensional case see [Chicone, 1999]. A similar argument also works for the strong stable manifold.

Exercise 9.8. Use the Lyapunov-Perron integral-equation approach discussed in class to prove the existence of a local strong stable manifold W for two-dimensional, autonomous dynamical systems. In particular, consider a system of the form

$$\dot{x} = f(x); \quad x \in \mathbb{R}^2, \quad f(0) = 0, \quad \text{spect}[Df(0)] = \{-\lambda_{ss}, -\lambda_s\}, \quad (9.101)$$

where $\lambda_s < \gamma < \lambda_{ss}$.

Chapter 10

Spectral submanifolds and model reduction

We now outline a dynamical systems approach to model reduction. Our goal is to identify globally attracting invariant sets S and understand the reduced dynamics on said sets, i.e. to find the reduced order model. However, we must address the issue of the existence, uniqueness, and smoothness of these sets. Smoothness is not a given, as attractors are very often not smooth manifolds, as seen in Fig. 10.1.

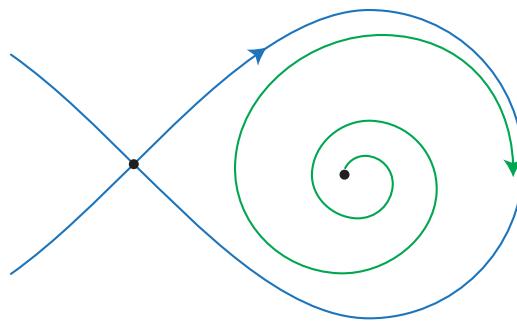


Figure 10.1: An example of a globally attracting set which is not a smooth manifold.

In order to study these objects, we must first set out a few definitions.

Definition 10.1. The *global attractor* is the largest bounded, invariant, set (in forward and backward time) that attracts trajectories.

Definition 10.2. An *inertial manifold* is a manifold of lower dimension than the full ambient space, that is attracting, forward-invariant manifold that is at least Lipschitz and contains the global attractor.

The existence of inertial manifolds is not guaranteed in general, however the following proposition provides a sufficient condition [Foias et al., 1988].

Proposition 10.1. *If we have the dynamical system defined on a Hilbert space H*

$$\dot{u} = Au + R(u, u); \quad u \in H. \quad (10.1)$$

Then given the following conditions:

- (i) *A is self-adjoint, negative definite, and densely defined;*
- (ii) *A^{-1} is compact;*
- (iii) *$R(u, u)$ is bilinear and is suitably dominated in norm by Au ;*

then we have that the following hold:

- (i) *All trajectories enter a ball B at some point;*
- (ii) *The limit set of the whole ball is the global attractor (inside the ball);*
- (iii) *An inertial manifold M exists.*

Such a global attractor within a ball B and the inertial manifold M is depicted in Fig. 10.2.

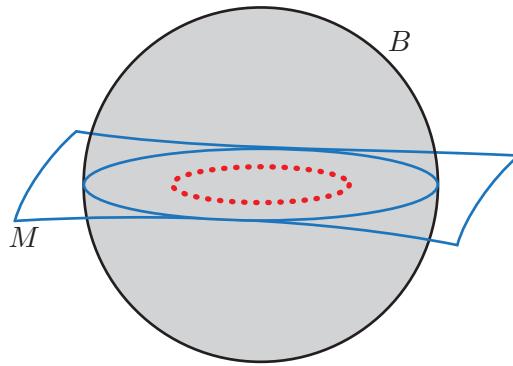


Figure 10.2: The red global attractor contained within the ball B and the inertial manifold M in the case of proposition 10.1.

An example in which such an M is proven to exist can be found in the reaction-diffusion equation

$$\partial_t u = D\Delta u + R(u), \quad (10.2)$$

for $u(x, t)$, $x \in \mathbb{R}^3$ and D being a positive definite matrix. In contrast, the proposition is not applicable to Navier-Stokes equations,

$$\partial_t u = -\nabla p + \frac{1}{\text{Re}} \Delta u - (u \nabla) u, \quad (10.3)$$

even in 2 dimensions. Due to this restrictiveness we may instead approximate the inertial manifold, following [Foias et al., 1989].

Assume we have the decomposition $u = p + q$ where $p = Pu$ is the projection of u , and $q = (I - P)$ is the residual after projection. We envision p to be the slow variable and q to be the fast variable. Using the self-adjointness of A , we obtain the projected equations

$$\begin{cases} \dot{p} = Ap + PR(p + q, p + q) \\ \dot{q} = Aq + (I - P)R(p + q, p + q). \end{cases} \quad (10.4)$$

Next, we mimick geometric singular perturbation theory and set $\dot{p} = 0$ which implies that p is constant $p = p_0$. Therefore we get that the solutions to

$$Aq + (I - P)R(p_0 + q, p_0 + q) = 0 \quad (10.5)$$

yield the critical manifold. Next if we assume that $R(p_0 + q, p_0 + q) \approx R(p_0, p_0)$ then we can write the fast variable q as a function of the slow variable p , as

$$q = \phi_0(p) = -A^{-1}(I - P)R(p, p). \quad (10.6)$$

The goal here is similar to that of a critical manifold. It should be kept in mind that we have relied on strong assumptions.

10.1 The simplest setting for nonlinear model reduction

Let us discuss nonlinear model reduction of a finite dimensional system near a fixed point. Consider the dynamical system

$$\dot{x} = Ax + f(x); \quad x \in \mathbb{R}^n, \quad A \in \mathbb{R}^{n \times n}, \quad f \in \mathcal{C}^r, \quad f = \mathcal{O}(\|x\|^2). \quad (10.7)$$

Assume that $x = \begin{pmatrix} x_s \\ x_f \end{pmatrix}$ where x_s is aligned with the s slower decaying modes and x_f is aligned with the f faster decaying modes. This enables us to write

$$\begin{cases} \dot{x}_s = A_s x_s + f_s(x_s, x_f) \\ \dot{x}_f = A_f x_f + f_f(x_s, x_f). \end{cases} \quad (10.8)$$

In the simplest form, this entails a 2-dimensional dynamical system with $\begin{pmatrix} x_s \\ x_f \end{pmatrix} = \begin{pmatrix} x \\ y \end{pmatrix}$ and the operators $f = 0$ and $A_s, A_f = -a, -b$ where $0 < a < b$. This case is illustrated in Fig. 10.3. The solutions of this system are

$$\begin{cases} x(t) = x_0 e^{-at} \\ y(t) = y_0 e^{-bt}. \end{cases} \quad (10.9)$$

A family of attracting, invariant, manifolds is formed by the trajectories described by (10.9). Equivalently, this family is given by the solutions to

$$\frac{1}{a} \log \left(\frac{x}{x_0} \right) = \frac{1}{b} \log \left(\frac{y}{y_0} \right) \quad (10.10)$$

The solution is given in this case by $y = Cx^{\frac{b}{a}}$, which is generally only $\lfloor \frac{b}{a} \rfloor$ -times differentiable except when $C = 0$ in which case the manifold is \mathcal{C}^∞ .

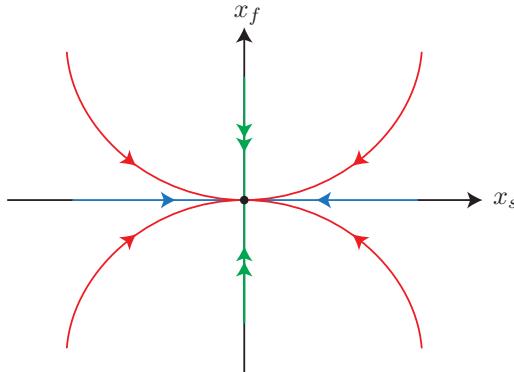


Figure 10.3: An example of the simplest setting for a nonlinear model reduction.

The reduced dynamics is the same on all of these invariant manifolds, namely

$$\dot{x} = -ax. \quad (10.11)$$

The conclusions are similar for general $n \geq 1$ in (10.8) as long as $f(x) = 0$. Now we may ask ourselves if it is possible to find similar low-dimensional attracting, invariant, manifolds for $f(x) \neq 0$? To this end the simplest approach is to use linearization theorems to relate the invariant manifolds of the $f(x) = 0$ case to those of the $f(x) \neq 0$ case. These come in various forms

- (i) Hartman-Grobman theorem: If $\text{Re}(\lambda_j) \neq 0$ for $j = 1, \dots, n$ then there exists a \mathcal{C}^0 linearization. However, this is insufficient for constructing smooth manifolds.

- (ii) Hartman theorem: If $\operatorname{Re}(\lambda_j) < 0$ for $j = 1, \dots, n$ and $f \in \mathcal{C}^2$ then there exists a \mathcal{C}^1 linearization. This is insufficient for any Taylor expansion and does not carry over the smoothness of any manifold smoother than \mathcal{C}^1 in the linearized system.
- (iii) Poincaré-Dulac theorem [Poincaré, 1952]: If we have $\operatorname{Re}(\lambda_j) < 0$ for $j = 1, \dots, n$, $f \in \mathcal{C}^a$ and the following non-resonance conditions hold

$$\sum_{i=1}^n m_i \lambda_i \neq \lambda_j; \quad \forall m \in \mathbb{N}^n : \sum_{i=1}^n m_i \geq 2 \quad (10.12)$$

then there exists a unique \mathcal{C}^a linearization. In this case, all invariant manifolds of the linear part of (10.8) continue to exist for $f(x) \neq 0$ without any change in their smoothness. It should be noted that the nonresonance condition still allows for 1:1 resonances, i.e. when $\lambda_i = \lambda_j$ for some i, j , but no others.

- (iv) Strong Sternberg linearization theorem: Under all of the conditions of the Poincaré-Dulac theorem, but for $f \in \mathcal{C}^r$, then there exists a \mathcal{C}^r linearization including the case of $r = \infty$. Under nonresonance, all \mathcal{C}^r manifolds of the linearization continue to exists for $r \in \mathbb{N}^+ \cup \{\infty\}$.
- (v) Sternberg linearization theorem [Sternberg, 1957]: With the same non-resonance conditions as above along with $f \in \mathcal{C}^r$ for $r \in \mathbb{N}^+ \cup \infty$ then there exists a \mathcal{C}^ρ linearization where $\rho \leq r$ and $\lim_{r \rightarrow \infty} \rho = \infty$. Note that this requires no assumptions on the real parts of the eigenvalues as we had before, which means Sternberg's theorem also applies to unstable hyperbolic fixed points. In the general non-resonant case, only $f \in \mathcal{C}^\infty$ gives specific enough linearization that preserves the smoothness of all invariant manifolds of the linearization.

Remark 10.2. Note that the nonresonance condition implies hyperbolicity of the fixed point. If there is a pair of purely imaginary eigenvalues, $\lambda = i\omega$ and $\bar{\lambda} = -i\omega$, we have, for example,

$$\lambda + 2\bar{\lambda} = \lambda \quad (10.13)$$

and thus the nonresonance condition is violated.

In summary the results listed above guarantee that under sufficient non-resonance conditions the existence of various invariant manifolds tangent to eigenspaces even for the nonlinear system. These manifolds are called *spectral submanifolds*.

We are now left with a few questions

- (i) How do we compute spectral submanifolds in the nonlinear system?

- (ii) Which of the infinitely many submanifolds do we use for model reduction?
- (iii) As linearization of the full dynamics is more than what we need for model reduction, do such manifolds exist under less stringent conditions as well?

To answer this we introduce the theory of (primary) spectral submanifolds and follow [Haller and Ponsioen, 2016]. We begin by introducing the dynamical system

$$\dot{x} = Ax + f_0(x), \quad x \in \mathbb{R}^n, \quad f_0(x) = \mathcal{O}(\|x\|^2), \quad f_0 \in \mathcal{C}^r, \quad (10.14)$$

for $r \in \mathbb{N}^+ \cup \{\infty, a\}$ where a signifies f_0 being analytic. Note here that the first condition on f_0 implies that $x = 0$ is a fixed point. The eigenvalues of A given by $\lambda_1, \dots, \lambda_n$ fulfill

$$\operatorname{Re}(\lambda_n) \leq \operatorname{Re}(\lambda_{n-1}) \leq \dots \leq \operatorname{Re}\lambda_1 \quad (10.15)$$

This implies that $x = 0$ is asymptotically stable. Assume that the eigenspaces are such that every eigenspace E_j is the span of the real and imaginary parts of all eigenvectors and generalized eigenvectors corresponding to λ_j for $j = 1, \dots, n$, thus we have

$$\mathbb{R}^n = \bigoplus_{j=1}^N E_j. \quad (10.16)$$

Such a set of eigenspaces is shown in Fig. 10.4.

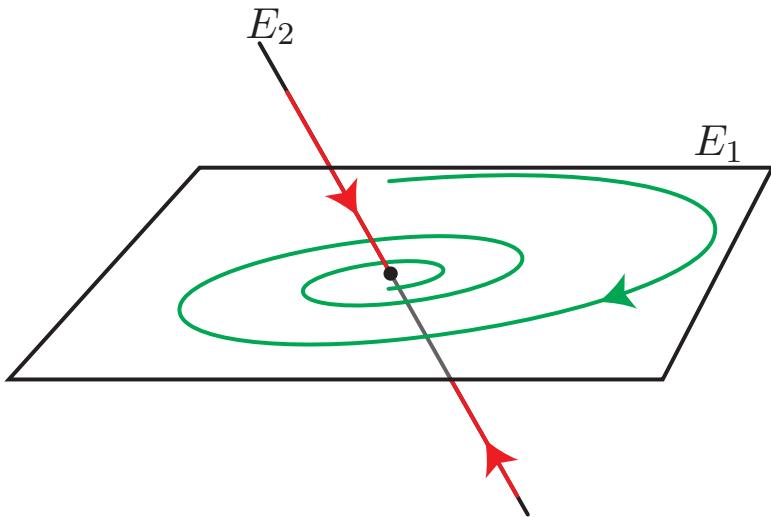


Figure 10.4: A set of eigenspaces fulfilling the assumption that they are generated by the eigenvectors and generalized eigenvectors.

Next we lay out a few definitions which will be useful for studying such dynamical systems.

Definition 10.3. We call the direct sum of an arbitrary number of distinct eigenspaces the *spectral subspace*

$$E = \bigoplus_{j=1}^k E_j; \quad 1 \leq k \leq N. \quad (10.17)$$

This always forms an invariant subspace of $\dot{x} = Ax$.

Definition 10.4. The *spectral quotient* of A is given by

$$\sigma(E) = \left[\frac{\text{smallest } \operatorname{Re}[\lambda_k] \text{ outside of } E}{\text{largest } \operatorname{Re}[\lambda_i] \text{ inside of } E} \right]. \quad (10.18)$$

An illustration of these two values are given in Fig. 10.5.

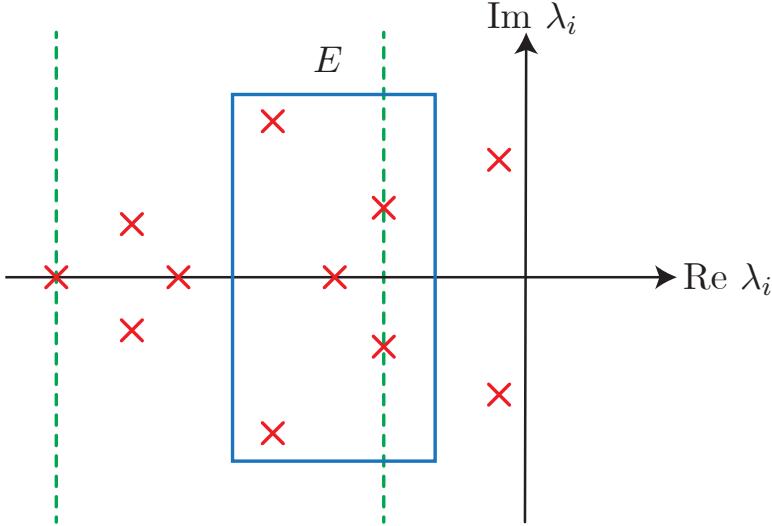


Figure 10.5: The eigenvalue constellation, the left dotted green line designates the smallest real part outside of E and the right dotted green line designates the largest real part inside of E .

Definition 10.5. The *external non-resonance condition* for E with $\dim[E] = q$ is

$$\langle m, \lambda \rangle_E = m_1 \lambda_{j1} + \dots + m_q \lambda_{jq} \neq \lambda_l, \quad m_i \in \mathbb{N}, \quad (10.19)$$

where $\lambda_{j1}, \dots, \lambda_{jq} \in \operatorname{spect}[A|_E]$ and $\lambda_l \notin \operatorname{spect}[A|_E]$.

We proceed by discussing the existence of a spectral submanifold (SSM) in the full nonlinear system, i.e. a nonlinear continuation of E .

Theorem 10.3. *Assume the low-order external non-resonance condition*

- (i) $\langle m, \lambda \rangle_E \neq \lambda_l$ for $\lambda_l \neq \text{spect}[A|_E]$;
- (ii) $2 \leq |m| \leq \sigma(E)$ with the norm $|m| = m_1 + \dots + m_k$;

then we have that the following hold:

- (i) There exists a \mathcal{C}^r spectral submanifold $W(0)$ tangent to E at $x = 0$ with the same dimension as E ;
- (ii) $W(0)$ is unique among all $\mathcal{C}^{\sigma(E)+1}$ invariant manifolds satisfying (i);
- (iii) If f_0 is jointly \mathcal{C}^r in (x, μ) for some parameter $\mu \in \mathbb{R}^p$, then so is $W(0)$, this also holds for $r = \infty$ and $r = a$.

Furthermore if f_0 is smooth or analytic in x and μ , then so is $W(0)$.

Proof. See [Haller and Ponsioen, 2016] □

With this theorem, we can see that high-enough-order Taylor-series expansion puts us on $W(0)$ if we look for it as an invariant graph over E .

Example 10.1. Consider the dynamical system

$$\begin{cases} \dot{x} = -\frac{1}{4}x + xy + x^3 \\ \dot{y} = -y - 2x^2, \end{cases} \quad (10.20)$$

whose dynamics is depicted in Fig. 10.6. The slow spectral subspace is given by $E = E_1$ where E_1 is defined as the x -axis and can be seen in Fig. 10.6. The spectral quotient of this manifold can be calculated as

$$\sigma(E) = \left[\frac{-1}{-\frac{1}{4}} \right] = 4. \quad (10.21)$$

It can also be seen that m_1 exists such that $m_1 \lambda_1 = -m_1 \frac{1}{4} = -1 = \lambda_2$ with $2 \leq |m| = m_1 \leq 4$, namely $m_1 = 4$, therefore the non-resonance condition for E is not fulfilled and no unique spectral submanifold as a nonlinear continuation of E exists. Hence, we modify the system to

$$\begin{cases} \dot{x} = -\frac{1}{\alpha}x + xy + x^3 \\ \dot{y} = -y - 2x^2 \end{cases}; \quad \alpha \notin \mathbb{N}^+ - \{1\}. \quad (10.22)$$

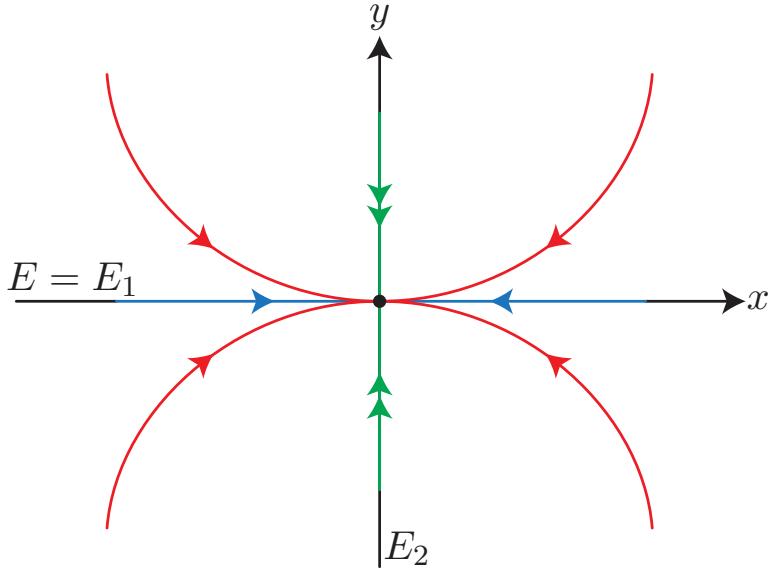


Figure 10.6: The dynamics of the dynamical system given in the example.

With this modification we find that $\sigma(E) = \lfloor \alpha \rfloor$, the non-resonance condition becomes $-m_1 \frac{1}{\alpha} \neq 1$ for $2 \leq m_1 \leq \lfloor \alpha \rfloor$, thus we can see why we need the restriction on α . Assuming such an α we find that there exists a unique spectral submanifold in the class of $\mathcal{C}^{\lfloor \alpha \rfloor + 1}$.

Next, we look for a spectral submanifold as a Taylor expansion over E . To do this write the manifold as

$$y = ax^2 + bx^3 + cx^4 + \mathcal{O}(x^5) = h(x). \quad (10.23)$$

We can differentiate this to find

$$\dot{y} = h'(x)\dot{x} = (2ax + 3bx^2 + 4cx^3 + \mathcal{O}(x^4)) \left(-\frac{1}{\alpha}x + xh(x) + x^3 \right) \quad (10.24)$$

$$= -\frac{2a}{\alpha}x^2 - \frac{3b}{\alpha}x^3 - \frac{4c}{\alpha}x^4 + 2a(a+1)x^4 + \mathcal{O}(x^5). \quad (10.25)$$

On the other hand, from equation (10.22) we find

$$\dot{y} = -y|_{W_E(0)} - 2x^2 = - (ax^2 + bx^3 + cx^4 + \mathcal{O}(x^5)) - 2x^2 \quad (10.26)$$

$$= -(a+2)x^2 - bx^3 - cx^4 + \mathcal{O}(x^5). \quad (10.27)$$

By comparing (10.25) and (10.27) we can see

$$\mathcal{O}(x^2) : \frac{2a}{\alpha} = a + 2 \implies a(2 - \alpha) = 2\alpha; \quad (10.28)$$

$$\mathcal{O}(x^3) : \frac{3b}{\alpha} = b \implies b(3 - \alpha) = 0; \quad (10.29)$$

$$\mathcal{O}(x^4) : 2a(a + 1) - \frac{4c}{\alpha} = -c \implies c(4 - \alpha) = 2a(a + 1)\alpha. \quad (10.30)$$

These equations cannot be solved when a resonance is present. Furthermore, below order $\sigma(E)$ all manifolds have the same expansion.

We may select $\alpha = \frac{3}{2}$ which implies $\lambda_1 = -\frac{2}{3}$ and $\lambda_2 = -1$, thus we have that $a = 6$, $b = 0$, and $c = \frac{2a(a+1)\alpha}{\alpha-1} > 0$. Hence the spectral quotient can be calculate as $\sigma(E) = \left\lfloor \frac{-1}{-\frac{2}{3}} \right\rfloor = 1$, and thus $W_E(0)$ is already unique among the class of \mathcal{C}^2 invariant manifolds. Based on this we can see that the second order Taylor expansion only exists for $W_E(0)$

$$y = 6x^2 + cx^4 + \mathcal{O}(x^5). \quad (10.31)$$

We may now write the reduced dynamics

$$\dot{x} = -\frac{2}{3}x + x(6x^2 + cx^4 + \mathcal{O}(x^5)) + x^3 = -\frac{2}{3}x + 7x^3 + \mathcal{O}(x^4). \quad (10.32)$$

From here we derive the leading order model

$$\dot{x} = x \left(-\frac{2}{3} + 7x^2 \right) \quad (10.33)$$

which has two nontrivial fixed points at $x = \pm\sqrt{\frac{2}{21}}$. Now we may ask if the full system has fixed points, namely by attempting to solve

$$\begin{cases} -\frac{2}{3}x + xy + x^3 = 0 \\ -y - 2x^2 = 0. \end{cases} \quad (10.34)$$

By setting the condition implied from the second equation into the first we find

$$-\frac{2}{3}x - x(2x^2) + x^3 = 0 \implies x \left(-\frac{2}{3} - x^2 \right) = 0. \quad (10.35)$$

This implies that no fixed point exist outside of the origin! The first order approximation of the spectral submanifold may not suffice, but we have the opportunity to go to *any* order, without increasing the dimension of the reduced order model.

Following the results of the previous example, we now look to extending the spectral submanifold to non-autonomous (forced) dynamical systems. For this we will assume the setup with $0 \leq \varepsilon \ll 1$

$$\dot{x} = Ax + f_0(x) + \varepsilon f_1(x, \Omega t, \varepsilon); \quad x \in \mathbb{R}^n, \quad A \in \mathbb{R}^{n \times n}, \quad f_0, f_1 \in \mathcal{C}^r, \quad f_0 = \mathcal{O}(\|x\|^2). \quad (10.36)$$

Further we assume that $\Omega = (\Omega_1, \dots, \Omega_k) \in \mathbb{R}^k$ has rationally independent entries and that f_1 is 2π -periodic in the Ωt arguments. Furthermore we assume that the real parts of the eigenvalues are ascending and are all strictly less than 0, i.e. for $\varepsilon = 0$ the origin is asymptotically stable (and hyperbolic). We are interested in the spectral subspace $E = \bigoplus_i E_i$, the direct sum of eigenspaces forming a q dimensional space with $q \leq n$. Expanding on the definition of the spectral quotient, let us define the absolute spectral quotient as

$$\Sigma(E) = \left[\frac{\text{smallest } \operatorname{Re}(\lambda_i) \text{ in } \operatorname{spect}(A)}{\text{largest } \operatorname{Re}(\lambda_2) \text{ in } \operatorname{spect}(A|_E)} \right] \quad (10.37)$$

We may introduce the k -dimensional angle ϕ , such that the extended phase space is given by $\mathbb{R}^n \times \mathbb{T}^k$ where the dynamics is

$$\begin{cases} \dot{x} = Ax + f_0(x) + \varepsilon f_1(x, \phi, \varepsilon) \\ \dot{\phi} = \Omega \end{cases}; \quad \phi = \phi_0 + \Omega(t - t_0) \in \mathbb{T}^k. \quad (10.38)$$

The geometry of this system for $\varepsilon = 0$ is shown in Fig. 10.7.

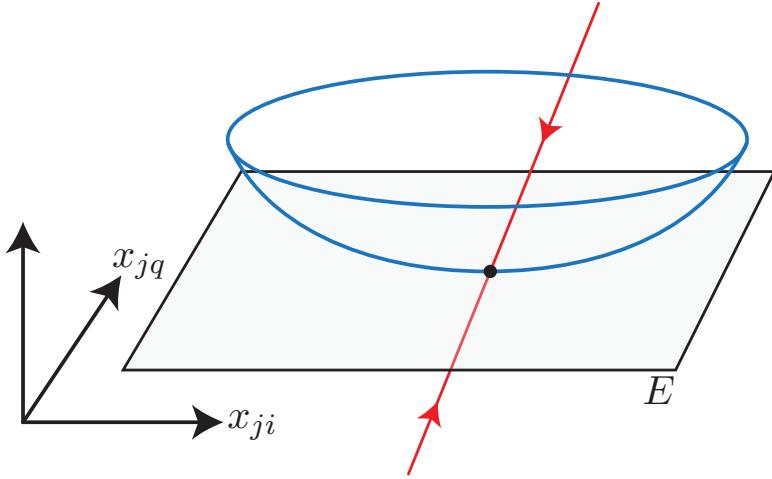


Figure 10.7: The geometry of the \mathbb{R}^k portion of the dynamical system with $\dim(E) = q$.

Theorem 10.4. *If we have that $\langle m, \lambda \rangle_E \neq \lambda_l$ for $2 \leq |m| \leq \sigma(E)$, $\lambda_i \in \operatorname{spect}(A|_E)$, and $\lambda_l \in \operatorname{spect}(A) - \operatorname{spect}(A|_E)$, then the following hold for $\varepsilon > 0$ small enough*

- (i) There exists a unique \mathcal{C}^r invariant torus T_ε ;
- (ii) There exists a unique spectral submanifold $W_E(T_\varepsilon)$ of dimension $q + k$ which is $\mathcal{O}(\varepsilon)$ \mathcal{C}^r -close to $W_E(0) \times T_0$;
- (iii) $W_E(T_\varepsilon)$ is already unique among all $\mathcal{C}^{\Sigma(E)+1}$ invariant manifolds satisfying (ii);
- (iv) If f_0 and f_1 are smooth or analytic, then so is $W_E(T_\varepsilon)$ in ε too.

Remark 10.5. Note that the non-resonance condition here is stronger than that of the Theorem, and is already violated when the real part of the spectrum has an outer resonance.

Example 10.2. Consider the dynamical system

$$\begin{cases} \dot{x} = -x \\ \dot{y} = -\sqrt{24}y + \underbrace{x^2 + x^3 + x^4 + x^5}_{f_0} + \underbrace{\varepsilon(\sin(t) + \sin(\sqrt{2}t))}_{\varepsilon f_1}. \end{cases} \quad (10.39)$$

Therefore we have $\Omega = (1, \sqrt{2})$, i.e. $k = 2$. We have $E = E_1$ and $\Sigma(E) = \left[\frac{-\sqrt{24}}{-1} \right] = 4$, this already tells us that if it exists then $W_E(T_\varepsilon)$ is unique in the smoothness class \mathcal{C}^5 . Next we check that the non-resonance condition is satisfied

$$\forall m_1 \in \mathbb{N}^+ : \quad m_1(-1) \neq -\sqrt{24}. \quad (10.40)$$

Thus we have that there exists a unique, invariant, 2 dimensional torus T_ε which is $\mathcal{O}(\varepsilon)$ \mathcal{C}^r -close to $T_0 = 0 \times \Pi^2$ and that there exists a unique spectral submanifold $W_E(T_\varepsilon)$ which is 3 dimensional. Both T_ε and $W_E(T_\varepsilon)$ are explicitly computable and are indeed unique in the appropriate sense. These are shown in Fig. 10.8.

10.2 Spectral Submanifolds in Mechanical Systems

We may now reduce the setting to mechanical systems, i.e. with n degrees-of-freedom, damped-forced, weakly nonlinear mechanical systems. Consider the dynamical system

$$M\ddot{q} + C\dot{q} + Kq + N(q, \dot{q}) = \varepsilon F(q, \dot{q}, \Omega t, \varepsilon); \quad q = \begin{pmatrix} q_1 \\ \vdots \\ q_n \end{pmatrix} \in \mathbb{R}^n. \quad (10.41)$$

Again we require that F is quasi-periodic in the argument Ωt with a rationally independent frequency vector $\Omega = (\Omega_1, \dots, \Omega_k)$ with $\Omega_i \in \mathbb{R}^k$. The matrices M , K , and C are all symmetric

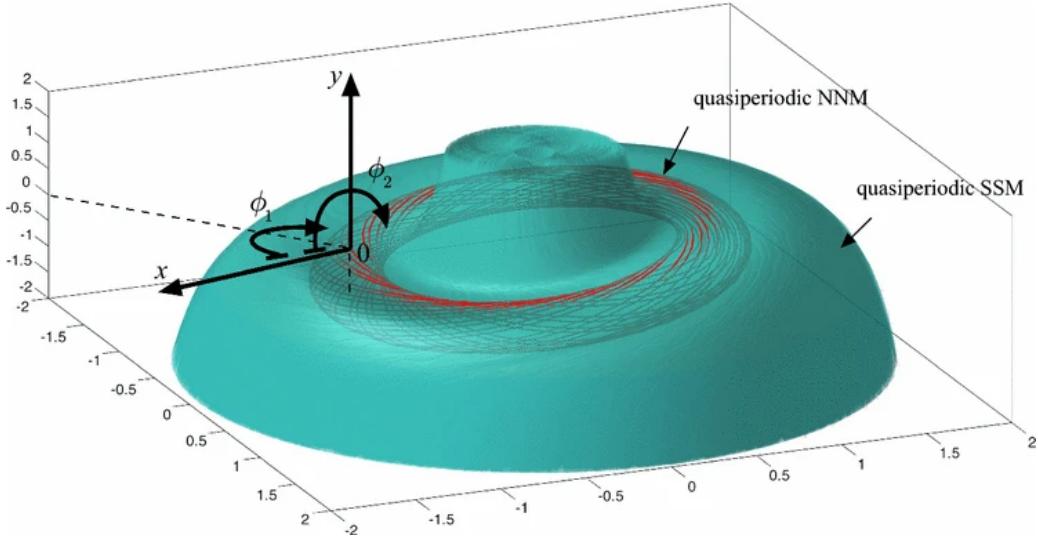


Figure 10.8: A projection of the extended phase space, the teal surface is the unique, analytic, quasiperiodic spectral submanifold, emanating from the unique quasiperiodic nonlinear normal mode in red [Haller and Ponsioen, 2016].

$\in \mathbb{R}^{n \times n}$ and are called the mass, stiffness, and damping matrices respectively, and all fulfill $M, K, C \succ 0$.

The structural damping hypothesis states that $C = \alpha K + \beta M$, is convenient but is not required. Note that for $\varepsilon = 0$ the eigenvalues of the linearization satisfy the characteristic equation $\det(\lambda^2 M + \lambda C + K) = 0$. This implies that $\text{Re}(\lambda_{2n}) \leq \dots \leq \text{Re}(\lambda_1) < 0$. For the linear system we have the ansatz $q = e^{\lambda t} A$ for $A \neq 0$. The corresponding eigensolution are the linear normal modes, written as

$$q_j(t) = \exp(\text{Re}(\lambda_j)t) \sin(\text{Im}(\lambda_j)t + \varphi_j) q_0^j. \quad (10.42)$$

The j -th mode shape is the constant vector $q_0^j \in \mathbb{R}^n$. If $\text{Im}(\lambda_j) = 0$ we have an overdamped mode, otherwise if $\text{Im}(\lambda_j) \neq 0$ we have an underdamped mode. The modal coordinates are

$$q = Tx, \quad T = (q_0^1, \dots, q_0^n); \quad x \in \mathbb{R}^n. \quad (10.43)$$

In these coordinates (10.41) becomes

$$\ddot{x}_i = 2\xi_i \omega_i \dot{x}_i + \omega_i^2 x_i + s_i(x, \dot{x}) = \varepsilon f_i(x, \dot{x}, \Omega t, \varepsilon). \quad (10.44)$$

The natural frequency ω_i is equal to $\text{Im}(\lambda_i)$ and the modal damping factor ξ_i is considered overdamping if $\xi_i > 1$ and underdamped if $\xi_i < 1$. The modal forcing terms are contained in f_i . Possible nonlinear responses are shown in Fig. 10.9.

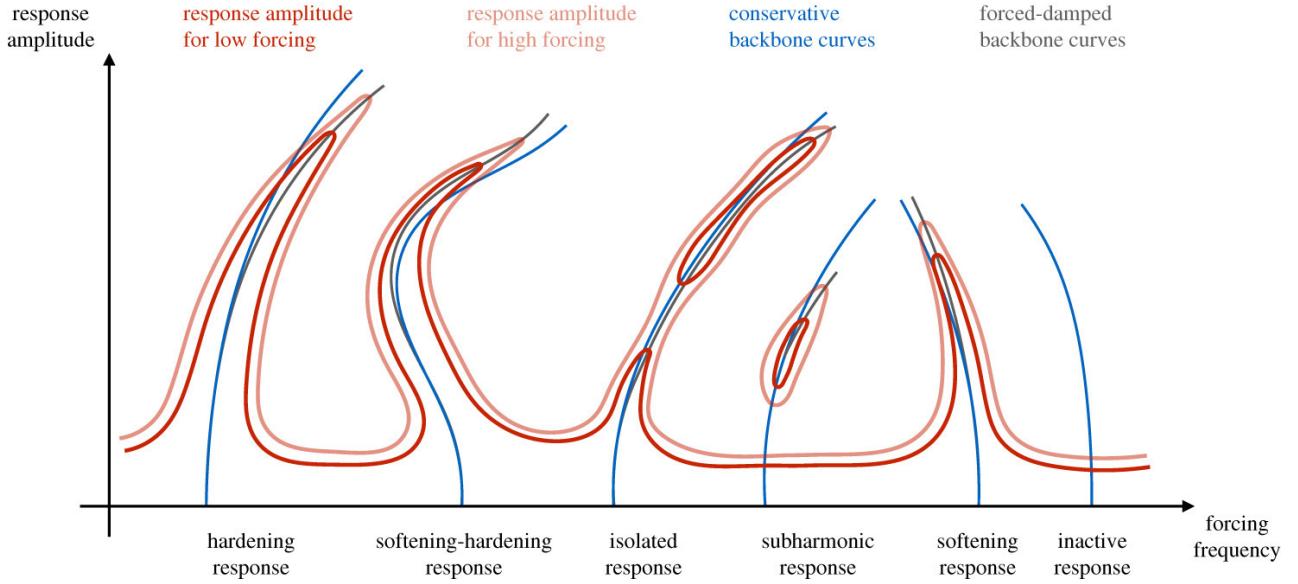


Figure 10.9: Illustration of the frequency response phenomena in mechanical systems. The dark and light red curves identify the frequency response for low and high forcing amplitudes, respectively, while blue curves depict conservative backbone curves and grey curves represent force-damped backbone curves [Cenedese and Haller, 2020].

We can formulate this autonomously in the extended phase space $\mathbb{R}^{2n} \times \mathbb{T}^k$

$$\begin{cases} \dot{x}_i = y_i \\ \dot{y}_i = 2\xi_i\omega_i y_i - \omega_i^2 x_i + s_i(x, y) + \varepsilon f(x, y, \phi, \varepsilon) \\ \dot{\phi} = \Omega. \end{cases} \quad (10.45)$$

The generalized quasiperiodic spectral submanifold results apply here. Note that for a typical forced response curve (e.g Fig. 10.9) the forcing is periodic, $k = 1$. The modal subspace is the spectral subspace of the form $E = E_j$ with E_j the eigenspace corresponding to the 2 eigenvalues of the j -th mode. Implied by the mode being overdamped, i.e. $\xi_j > 1$, is that we only have negative real eigenvalues, thus the linearized dynamics on E_j is described by a node. Meanwhile in the underdamped case, we have a complex conjugate pair of eigenvalues, thus the dynamics on E_j is described by a stable focus.

The modal submanifolds are the nonlinear continuations of $E_j \times \Pi^k$ under forcing, i.e. the spectral subbundle of perturbed invariant tori $W_{E_j}(T_\varepsilon)$.

Since we have the freedom to choose the dimension of the reduced order model (of the relevant spectral submanifold), this should be selected depending on how many of the transient time scales we would like to model (often stemming from the resonances with external forcing).

Another consideration in this selection is that the lower-order internal resonances must be included in the spectral submanifold.

10.3 System Identification with Periodic External Forcing

The most important case in nonlinear system identification and testing is with periodic external forcing. Let us assume periodic external forcing ($k = 1$, i.e. $T_0 \sim T_\varepsilon \sim S^1$). Our goal is to predict the forced response curve. Ideally we would find a nontrivial response near resonance which can be analyzed by reduction to $W_{E_j}(T_\varepsilon)$. In this case E_j is the modal subspace corresponding to the resonance.

The normal form for the reduced dynamics a 2 dimensional modal submanifold

$$\begin{cases} \dot{\rho} = a(\rho) + \varepsilon(b_1(\rho, \Omega) \cos(\Psi) + b_2(\rho, \Omega) \sin(\Psi)) + \mathcal{O}(\varepsilon^2) \\ \dot{\Psi} = b(\rho) - \Omega + \frac{\varepsilon}{\rho}(c_1(\rho, \Omega) \cos(\Psi) + c_2(\rho, \Omega) \sin(\Psi)) + \mathcal{O}(\varepsilon^2), \end{cases} \quad (10.46)$$

where we have

$$a(\rho) = \operatorname{Re}(\lambda_j) + \sum_i a_i \rho^{2i+1} \quad (10.47)$$

$$b(\rho) = \operatorname{Im}(\lambda_j) + \sum_i b_i \rho^{2i}, \quad (10.48)$$

further $b_{1,2}$ and $c_{1,2}$ in (10.46) are Taylor series in ρ , which are convergent if $W_{E_j}(T_\varepsilon)$ is analytic. We call ρ the polar radius, Ψ the phase lag between the polar angle φ and phase of the forcing, i.e. $\Psi = \varphi - \Omega t$.

Theorem 10.6. *Nontrivial non-spurious transverse zeros of $a(\rho)$ (at ρ_0) give rise to isola, created around the point $(\Omega, \rho) = (b(\rho_0), \rho_0)$.*

An illustration of the time periodic spectral submanifold and isola as in the theorem are given in Fig. 10.10.

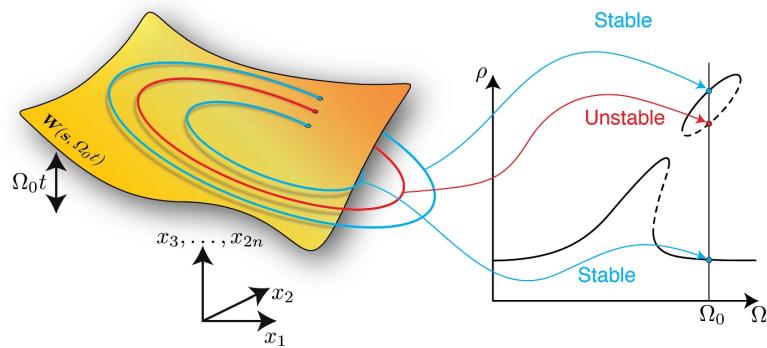
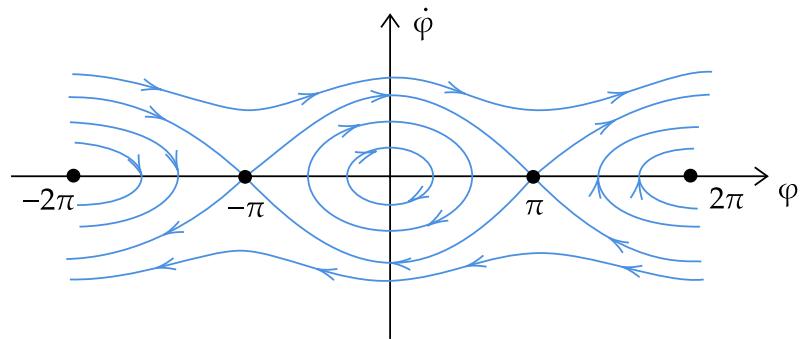


Figure 10.10: Illustration of a time periodic spectral submanifold. For a given forcing frequency Ω , here it is illustrated how the spectral submanifold may contain three limit cycles, of which two fall in an isola [Ponsioen et al., 2018].

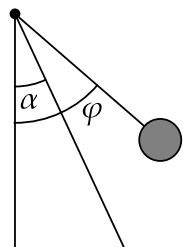
Chapter 11

Exercise Solutions

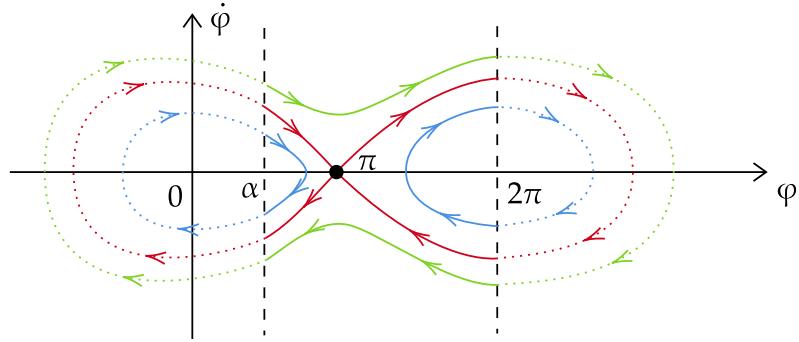
Solution (1.1). *Without any walls:*



(i) *Case $\alpha > 0$:*



(a) Post impact velocity $v(t_+) = -v(t_-)$ pre-impact velocity

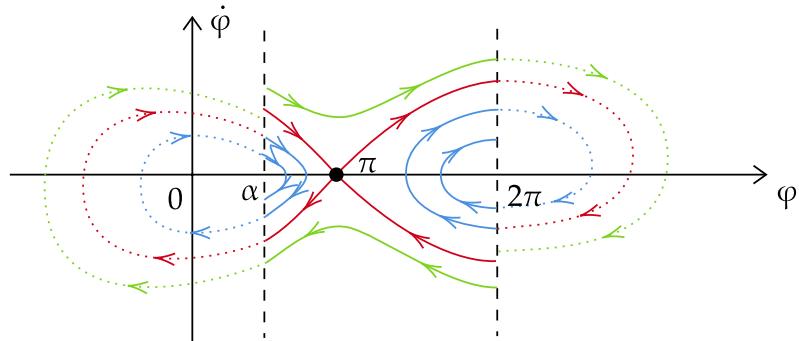


The dashed lines represent the instance of impact.

Possible asymptotic behavior: depending on initial conditions $\varphi(0), \dot{\varphi}(0)$

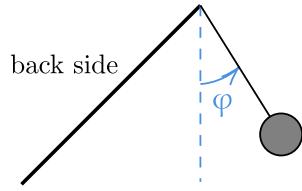
- Bouncing against the inclined wall or the vertical wall;
- Bouncing back and forth between the two walls;
- Convergence to the upright position $\varphi = \pi, \dot{\varphi} = 0$.

(b) Post impact velocity $v(t_+) = -\frac{1}{2}v(t_-)$ Possible asymptotics:

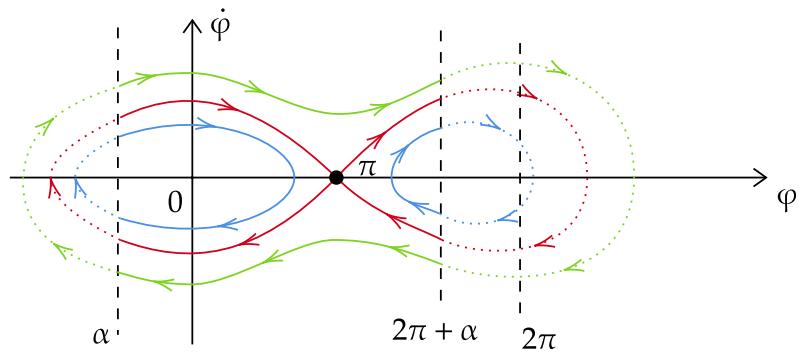


- Convergence to the upright position
- Lying against the inclined wall
- Lying against the vertical wall

(ii) Case $\alpha < 0$: Solution 1



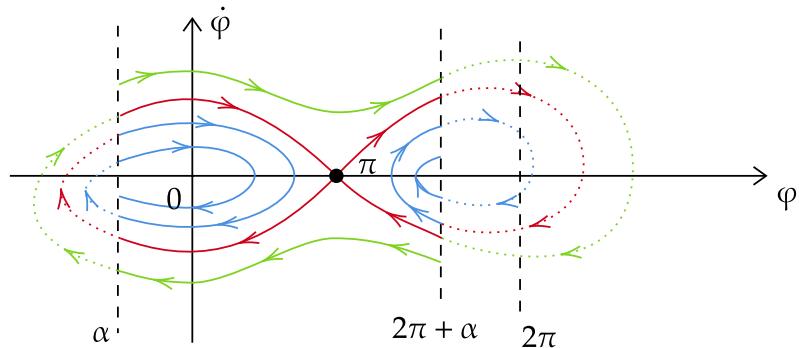
(a) $v(t_+) = -v(t_-)$:



Possible asymptotics:

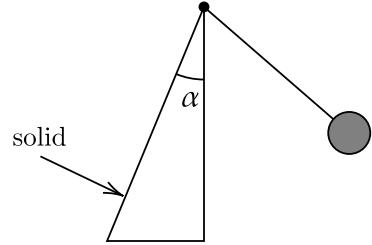
- Convergence to upright position
- Bouncing against either side of the wall
- Bouncing back and forth between the two sides of the wall
- Oscillating around the vertical position $\varphi = 0$

(b) $v(t_+) = -\frac{1}{2}v(t_-)$: Possible asymptotics:

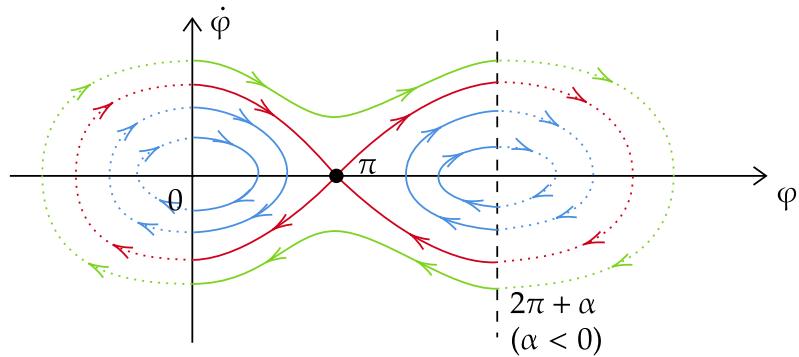


- *Convergence to the upright position*
- *Oscillating around the vertical position $\varphi = 0$*
- *Lying on the back of the wall*

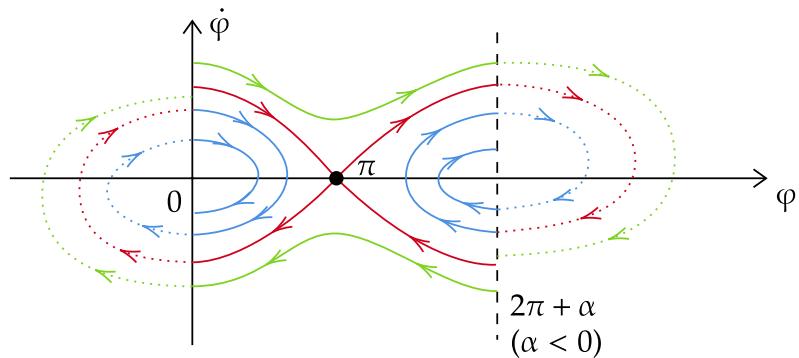
(iii) Case $\alpha < 0$: Solution 2



(a) $v(t_+) = -v(t_-)$:



(b) Post impact velocity $v(t_+) = -\frac{1}{2}v(t_-)$ pre-impact velocity Possible asymptotics:



- Convergence to the upright position
- Lying against the inclined wall
- Lying against the vertical wall

Solution (2.1).

(i) Define

$$h(t) = C + \int_{t_0}^t u(\tau)v(\tau) d\tau \quad (11.1)$$

$$\implies \dot{h}(t) = u(t)v(t) \leq^{(*)} v(t)h(t) \quad (11.2)$$

From the definition of $h(t)$ and because $u, v > 0$ we have $u(t) \leq h(t)$

$$\implies \frac{\dot{h}(t)}{h(t)} \leq v(t) \implies \frac{d}{dt} \log [h(t)] \leq v(t) \quad (11.3)$$

Integrate both sides of (11.3) to get:

$$\log \left[\frac{h(t)}{h(t_0)} \right] \leq \int_{t_0}^t v(\tau) d\tau \quad (11.4)$$

$$\implies h(t) \leq h(t_0) \exp \left(\int_{t_0}^t v(\tau) d\tau \right) \quad (11.5)$$

From (11.1) we have $h(t_0) = C$; hence:

$$u(t) \leq h(t) \leq C \exp \left(\int_{t_0}^t v(\tau) d\tau \right) \quad (11.6)$$

(ii) The solutions $x(t, x_0)$ and $x(t, \tilde{x}_0)$ of the IVP satisfy the integral aligns

$$x(t, x_0) = x_0 + \int_{t_0}^t f(x(s, x_0), s) ds \quad (11.7)$$

$$x(t, \tilde{x}_0) = x_0 + \int_{t_0}^t f(x(s, \tilde{x}_0), s) ds \quad (11.8)$$

Subtracting these two from each other and taking the absolute values we get

$$|x(t, x_0) - x(t, \tilde{x}_0)| = \left| x_0 - \tilde{x}_0 + \int_{t_0}^t [f(x(s, x_0), s) - f(x(s, \tilde{x}_0), s)] ds \right| \quad (11.9)$$

Using triangular and Jensen's inequalities we get

$$|x(t, x_0) - x(t, \tilde{x}_0)| \leq |x_0 - \tilde{x}_0| + \int_{t_0}^t |f(x(s, x_0), s) - f(x(s, \tilde{x}_0), s)| ds \quad (11.10)$$

By Lipschitz continuity of f we have

$$|f(x(s, x_0), s) - f(x(s, \tilde{x}_0), s)| \leq L |x(t, x_0) - x(t, \tilde{x}_0)| \quad (11.11)$$

This together with (11.10) gives

$$|x(t, x_0) - x(t, \tilde{x}_0)| \leq |x_0 - \tilde{x}_0| + \int_{t_0}^t L |x(s, x_0) - x(s, \tilde{x}_0)| ds \quad (11.12)$$

Using Gronwall's inequality from Exercise 1.1 with $u(t) = |x(t, x_0) - x(t, \tilde{x}_0)|$, $C = |x_0 - \tilde{x}_0|$ and $v(t) = L$ (constant), we get

$$|x(t, x_0) - x(t, \tilde{x}_0)| \leq |x_0 - \tilde{x}_0| \exp\left(\int_{t_0}^t L ds\right) = |x_0 - \tilde{x}_0| e^{L(t-t_0)} \quad (11.13)$$

Solution (2.3).

$$\begin{cases} \ddot{x} + \omega_0^2 x = \varepsilon M x^2, & 0 < \varepsilon \ll 1, \quad M > 0, \quad \omega_0 \neq 0 \\ x(0) = a_0 \\ \dot{x}(0) = 0. \end{cases} \quad (11.14)$$

- Seek solutions of the form:

$$x_\varepsilon(t) = \varphi_0(t; \varepsilon) + \varepsilon \varphi_1(t; \varepsilon) + \mathcal{O}(\varepsilon^2) \quad (11.15)$$

$$\varphi_i(t, \varepsilon) = \varphi_i(t + T_\varepsilon; \varepsilon) \quad (11.16)$$

- rewrite period as

$$T_\varepsilon = \frac{2\pi}{\omega(\varepsilon)}, \quad \omega(\varepsilon) = \omega_0 + \varepsilon \omega_1 + \mathcal{O}(\varepsilon^2) \quad (11.17)$$

- Rescale time:

$$\tau = \omega(\varepsilon)t \implies \boxed{\frac{d}{d\tau} = \frac{1}{\omega(\varepsilon)} \frac{d}{dt}} \implies \boxed{(\omega(\varepsilon))^2 x'' + \omega_0^2 x = \varepsilon M x^2} \quad (11.18)$$

- Plug in the new Ansatz into the rescaled equation

$$[\omega_0^2 + 2\varepsilon\omega_0\omega_1 + \mathcal{O}(\varepsilon^2)][\varphi_0'' + \varepsilon\varphi_1'' + \mathcal{O}(\varepsilon^2)] + \omega_0^2[\varphi_0 + \varepsilon\varphi_1 + \mathcal{O}(\varepsilon^2)] = \varepsilon M[\varphi_0^2 + \mathcal{O}(\varepsilon)] \quad (11.19)$$

- Collect terms of equal power of ε

$\mathcal{O}(1)$:

$$\omega_0^2 \varphi_0'' + \omega_0^2 \varphi_0 = 0, \quad \varphi_0(0) = a_0, \quad \dot{\varphi}_0(0) = 0 \quad (11.20)$$

$$\implies \varphi_0 = a_0 \cos(\tau) \quad (11.21)$$

$\mathcal{O}(2)$:

$$\omega_0^2 \varphi_1'' + \omega_0^2 \varphi_1 = M \varphi_0^2 - 2\omega_0 \omega_1 \varphi_0'' = Ma_0^2 \cos^2(\tau) + 2a_0 \omega_0 \omega_1 \cos(\tau) \quad (11.22)$$

$$= M \frac{a_0^2}{2} [1 + \cos(2\tau)] + \underbrace{2a_0 \omega_0 \omega_1 \cos(\tau)}_{\text{resonance}} \quad (11.23)$$

Select $\omega_1 = 0$ to eliminate resonance terms and obtain periodic solution.

Solve for φ_1 :

$$\varphi_1'' + \varphi_1 = \frac{Ma_0^2}{2\omega_0^2} [1 + \cos(2\tau)], \quad \varphi_1(0) = 0, \quad \dot{\varphi}_1(0) = 0 \quad (11.24)$$

- Pick solution Ansatz:

$$\varphi_1(\tau) = A \cos(\tau) + B \sin(\tau) + C \cos(2\tau) + D \sin(2\tau) + E \quad (11.25)$$

- Substituting in (11.24):

$$-A \cos(\tau) - B \sin(\tau) - 4C \cos(2\tau) - 4D \sin(2\tau) + A \cos(\tau) + B \sin(\tau) \quad (11.26)$$

$$+ C \cos(2\tau) + D \sin(2\tau) + E = \frac{Ma_0^2}{2\omega_0^2} \cos(2\tau) + \frac{Ma_0^2}{2\omega_0^2} \quad (11.27)$$

- Comparing coefficients:

$$\implies E = \frac{Ma_0^2}{2\omega_0^2}, \quad C = -\frac{Ma_0^2}{6\omega_0^2}, \quad D = 0 \quad (11.28)$$

$$\varphi_1(0) = 0 \implies A + C + E = 0 \implies A = -\frac{Ma_0^2}{3\omega_0^2} \quad (11.29)$$

$$\varphi_1'(0) = 0 \implies B + 2D = 0 \implies B = 0 \quad (11.30)$$

$$\implies \boxed{\varphi_1(\tau) = -\frac{Ma_0^2}{3\omega_0^2} \cos(\tau) - \frac{Ma_0^2}{6\omega_0^2} \cos(2\tau) + \frac{Ma_0^2}{2\omega_0^2}} \quad (11.31)$$

- In original time:

$$x_\varepsilon(t) = a_0 \cos(\omega t) + \varepsilon \frac{Ma_0^2}{\omega_0^2} \left[-\frac{1}{3} \cos(\omega t) - \frac{1}{6} \cos(2\omega t) + \frac{1}{2} \right] + \mathcal{O}(\varepsilon^2) \quad (11.32)$$

where

$$\omega = \omega_0 + \mathcal{O}(\varepsilon^2) \quad (11.33)$$

Solution (2.4).

- (i) Seek solutions of the form:

$$x_\varepsilon(t) = \varphi_0(t) + \varepsilon \varphi_1(t) + \mathcal{O}(\varepsilon^2) \quad (11.34)$$

Substituting this solution in the ODE $\ddot{x} + \varepsilon(x^2 - 1)\dot{x} + x = F \cos(\omega t)$ we get:

$$\ddot{\varphi}_0 + \varphi_0 + \varepsilon(\ddot{\varphi}_1 + \varphi_1 + \varphi_0^2 \dot{\varphi}_0 - \dot{\varphi}_0) + \mathcal{O}(\varepsilon^2) = F \cos(\omega t) \quad (11.35)$$

$$\implies \mathcal{O}(1) : \ddot{\varphi}_0 + \varphi_0 = F \cos(\omega t) \quad (11.36)$$

$$\implies \mathcal{O}(2) : \ddot{\varphi}_1 + \varphi_1 = \dot{\varphi}_0(1 - \varphi_0^2) \quad (11.37)$$

Since we seek solutions with period $\frac{2\pi}{\omega}$ for any $0 \leq \varepsilon \ll 1$, the period of each φ_i must be $\frac{2\pi}{\omega}$. Because the initial condition for the system was not specified we will use those to enforce the periodicity condition. That is, we assume that the initial conditions also depend on ε .

$$x_\varepsilon(0) = a + b\varepsilon + \mathcal{O}(\varepsilon^2) \quad (11.38)$$

$$\dot{x}_\varepsilon(0) = c + d\varepsilon + \mathcal{O}(\varepsilon^2) \quad (11.39)$$

$$(11.40)$$

The solution that is $\mathcal{O}(\varepsilon)$ accurate, can be obtained as a solution to equation (11.36). We take the following trial function

$$\varphi_0(t) = A \sin(t) + B \cos(t) + C \cos(\omega t). \quad (11.41)$$

The constant C can be determined by substituting into (11.36) as

$$-C\omega^2 \cos(\omega t) + C \cos(\omega t) = F \cos(\omega t).$$

Since this has to hold for all t , the coefficients of $\cos(\omega t)$ must balance. This leads to $C = F/(1 - \omega^2)$. Therefore, the solution $\varphi_0(t)$ reads as

$$\varphi_0(t) = A \sin(t) + B \cos(t) + \frac{F \cos(\omega t)}{1 - \omega^2}. \quad (11.42)$$

To have a period $\frac{2\pi}{\omega}$, we must select $A = B = 0$. This condition can be enforced by choosing appropriate initial conditions for the ODE (11.36):

$$A = B = 0 \implies a = C = \frac{F}{1 - \omega^2} \quad d = 0. \quad (11.43)$$

$$\boxed{\varphi_0(t) = \frac{F \cos(\omega t)}{1 - \omega^2}, \quad \varphi_0(0) = \frac{F}{1 - \omega^2}, \quad \dot{\varphi}_0(0) = 0} \quad (11.44)$$

$$x_\varepsilon(t) = \varphi_0(t) + \underbrace{\mathcal{O}(\varepsilon)}_{\text{error term}} \quad (11.45)$$

Bonus: $O(\varepsilon^2)$ accurate calculation

We can proceed by solving the $O(\varepsilon)$ equation (11.37). With the solution (11.44) the forcing term on the right-hand side can be written as

$$\dot{\varphi}_0(1 - \varphi_0^2) = -C\omega \sin(\omega t) (1 - C^2 \cos^2(\omega t)) \quad (11.46)$$

$$= -C\omega \sin(\omega t) (1 - C^2 + C^2 \sin^2(\omega t)) \quad (11.47)$$

$$= -C\omega \sin(\omega t)(1 - C^2) - C^3 \omega \sin^3(\omega t) \quad (11.48)$$

$$= -C\omega(1 - C^2) \sin(\omega t) - C^3 \omega \frac{3 \sin(\omega t) - \sin(3\omega t)}{4} \quad (11.49)$$

$$= C\omega \left(\frac{C^2}{4} - 1 \right) \sin(\omega t) + \frac{C^3 \omega}{4} \sin(3\omega t). \quad (11.50)$$

Equation (11.37) now takes the form

$$\ddot{\varphi}_1 + \varphi_1 = C\omega \left(\frac{C^2}{4} - 1 \right) \sin(\omega t) + \frac{C^3 \omega}{4} \sin(3\omega t); \quad \varphi_1(0) = b \quad \dot{\varphi}_1(0) = d. \quad (11.51)$$

The appropriate trial functions is

$$\varphi_1(t) = D \cos t + E \sin t + G \sin(\omega t) + H \sin(3\omega t). \quad (11.52)$$

Upon substitution we find the following equation

$$-G\omega^2 \sin(\omega t) - 9H\omega^2 \sin(3\omega t) + G \sin(\omega t) + H \sin(3\omega t) = C\omega \left(\frac{C^2}{4} - 1 \right) \sin(\omega t) + \frac{C^3\omega}{4} \sin(3\omega t).$$

Balancing the coefficients of $\sin(\omega t)$ and $\sin(3\omega t)$ we end up with

$$G = \frac{C\omega}{(1 - \omega^2)} \left(\frac{C^2}{4} - 1 \right) \quad (11.53)$$

$$H = \frac{C^3\omega}{4(1 - 9\omega^2)}. \quad (11.54)$$

To enforce periodicity with period $T = \frac{2\pi}{\omega}$ to order ε we must have $D = E = 0$. This can be achieved by choosing the initial conditions of the φ_1 equation appropriately. This means we must have

$$\varphi_1(0) = \boxed{b = 0} \quad (11.55)$$

$$\dot{\varphi}_1(0) = \boxed{G\omega + 3H\omega = d}. \quad (11.56)$$

Combining the solution $\varphi_1(t)$ with (11.44) we finally obtain, that the T -periodic solution is given by

$$x_\varepsilon(t) = \frac{F}{1 - \omega^2} \cos(\omega t) + \varepsilon \frac{F\omega}{(1 - \omega^2)^2} \left(\frac{F^2}{4(1 - \omega^2)^2} - 1 \right) \sin(\omega t) \quad (11.57)$$

$$+ \varepsilon \frac{F^3\omega}{4(1 - \omega^2)^3(1 - 9\omega^2)} \sin(3\omega t) \quad (11.58)$$

$$x_\varepsilon(0) = \frac{F}{1 - \omega^2} + O(\varepsilon^2) \quad (11.59)$$

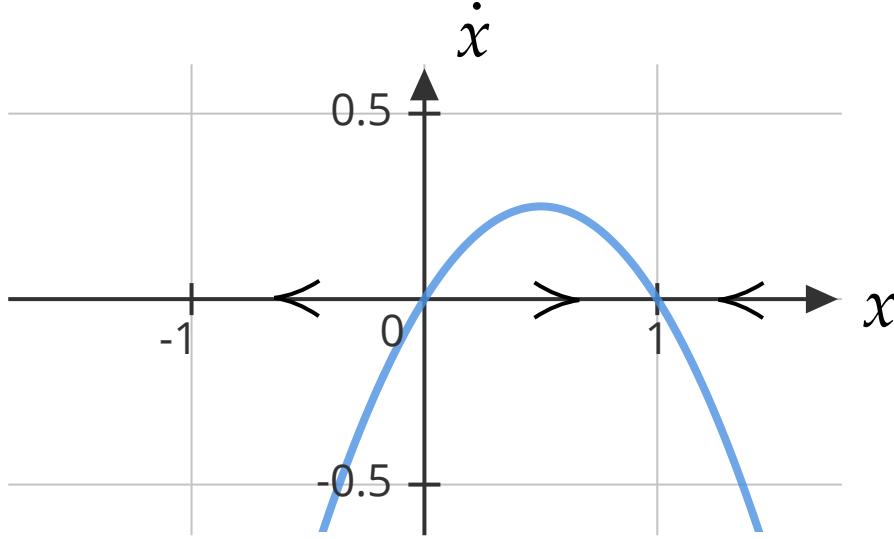
$$\dot{x}_\varepsilon(0) = \varepsilon \left[\frac{F\omega^2}{(1 - \omega^2)^2} \left(\frac{F^2}{4(1 - \omega^2)^2} - 1 \right) + 3 \frac{F^3\omega^2}{4(1 - \omega^2)^3(1 - 9\omega^2)} \right] + O(\varepsilon^2) \quad (11.60)$$

- (ii) We solve the ODE numerically to obtain a solution $x(t)$ and compare this solution to the perturbed approximation $x_\varepsilon(t)$ given by (11.44) up to $O(\varepsilon)$.

The initial conditions for the ODE are chosen such that $x(0) = x_\varepsilon(0)$ and $\dot{x}(0) = \dot{x}_\varepsilon(0)$. Therefore, $x(0) = \varphi_0(0)$, $\dot{x}(0) = \dot{\varphi}_0(0)$ where $\varphi_0(0)$ and $\dot{\varphi}_0(0)$ are given in (11.43).

Equivalent first order system of differential equations:

$$z_1 = x, \quad z_2 = \dot{x}, \quad \begin{bmatrix} \dot{z}_1 \\ \dot{z}_2 \end{bmatrix} = \begin{bmatrix} z_2 \\ F \cos(\omega t) - \varepsilon(z_1^2 - 1)z_2 - z_1 \end{bmatrix} \quad (11.61)$$



- (iii) We now solve the same ODE (11.61) numerically to obtain a solution $x(t)$ started from the $O(\varepsilon^2)$ accurate initial conditions given by (11.55) and compare this solution to refined perturbed approximation $x_\varepsilon(t)$ given by (11.57) up to $O(\varepsilon^2)$.

Solution (3.1).

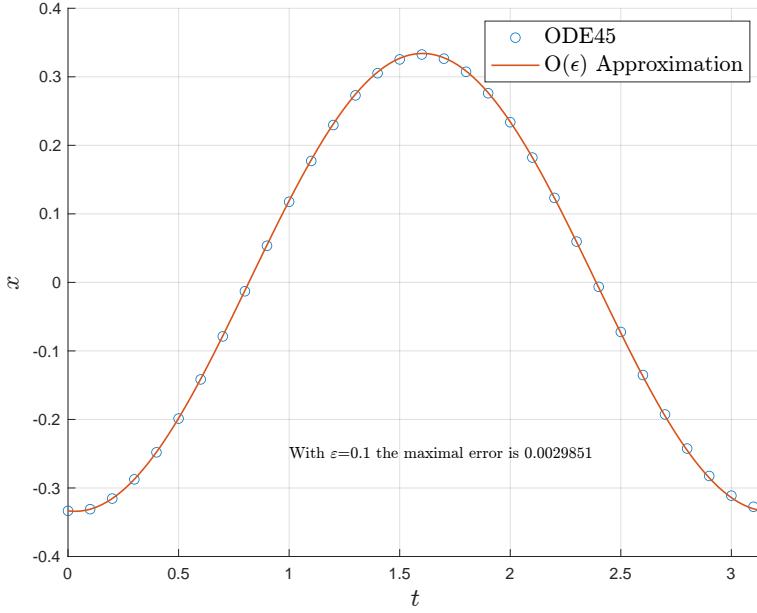
- (i) Let x_k be near the fixed point P and define $y_k = x_k - P$. Then

$$x_{k+1} = f(x_k) = f(P + y_k) = f(P) + Df(P)y_k + \mathcal{O}(\|y_k\|^2) \quad (11.62)$$

$$= P + Df(P)y_k + \mathcal{O}(\|y_k\|^2) \quad (11.63)$$

$$\implies y_{k+1} = x_{k+1} - P = Df(P)y_k + \mathcal{O}(\|y_k\|^2) \quad (11.64)$$

Now for $\|y_k\|$ small enough the linear approximation of the map $x_{k+1} = f(x_k)$ is $y_{k+1} = Ay_k$ with $A = Df(P)$.



(ii) Take any $y_0 \in \mathbb{R}^n$. Since $s_1, \dots, s_n \in \mathbb{C}^n$ are linearly independent there are constants $c_1, \dots, c_n \in \mathbb{C}$ such that $y_0 = c_1 s_1 + \dots + c_n s_n$.

Now define

$$y_1 = Ay_0 = c_1 As_1 + \dots + c_n As_n \quad (11.65)$$

$$= c_1 \lambda s_1 + \dots + c_n \lambda s_n \quad (11.66)$$

$$= c_1 \varphi_1(1) + \dots + c_n \varphi_n(1) \quad (11.67)$$

Similarly, for any $k \geq 1$,

$$y_k = Ay_{k-1} = c_1 \lambda_1^{k-1} As_1 + \dots + c_n \lambda_n^{k-1} As_n \quad (11.68)$$

$$= c_1 \varphi_1(k) + \dots + c_n \varphi_n(k) \quad (11.69)$$

It's easy to check that $y_{k+1} = Ay_k$ for any $k \geq 0$. Since $y_0 \in \mathbb{R}^n$ was arbitrary, $c_1 \varphi_1(k) + c_2 \varphi_2(k) + \dots + c_n \varphi_n(k)$ is a general solution of $y_{k+1} = Ay_k$.

(iii) We claim that the necessary and sufficient condition for asymptotic stability of the origin is $|\lambda_i| < 1$ for $i = 1, 2, \dots, n$

Sufficient: From (c) any solution of $y_{k+1} = Ay_k$ can be written as:

$$y_{k+1} = \sum_{i=1}^n c_i \lambda_i^k s_i \quad (11.70)$$

Without loss of generality, we assume that the eigenvectors s_i are normalized, i.e., $\|s_i\| = 1 \forall i \in \{1, 2, \dots, n\}$. Then

$$\|y_{k+1}\| \leq \sum_{i=1}^n |c_i| |\lambda_i|^k \|s_i\| = \sum_{i=1}^n |c_i| |\lambda_i|^k \quad (11.71)$$

But since $|\lambda_i| < 1$, we have $\lim_{k \rightarrow \infty} |\lambda_i|^k = 0$. Which implies

$$\lim_{k \rightarrow \infty} \sum_{i=1}^n |c_i| |\lambda_i|^k = 0 \quad (11.72)$$

Hence,

$$\lim_{k \rightarrow \infty} \|y_{k+1}\| = 0 \quad (11.73)$$

Also note that since $|\lambda_i| < 1 \forall i \in \{1, \dots, n\}$, the matrix norm $\|A\| < 1$.

Hence $\|y_{k+1}\| = \|Ay_k\| < \|y_k\| \implies y = 0$ is also stable. This together with (11.73) implies asymptotic stability of the fixed point $y = 0$.

Necessity: Assume there is $i_0 \in \{1, 2, \dots, n\}$ such that $|\lambda_{i_0}| \geq 1$.

It is enough to show that there exists $y_0 \in \mathbb{R}^n$ such that $\lim_{k \rightarrow \infty} \|A^k y_0\| \neq 0$.

This is due to the fact that $y_k = A^k y_0$ and that one can rescale y_0 as ry_0 for $0 < r \ll 1$ small enough such that $\|ry_0\| < \delta, \forall \delta > 0$.

To show that such $y_0 \in \mathbb{R}^n$ exists, note that $\|A^k s_{i_0}\| = \|\lambda_{i_0}^k s_{i_0}\| = |\lambda_{i_0}|^k \geq 1 \forall k \geq 0$.

Remark 11.1. This is, however, not enough since $s_{i_0} \in \mathbb{C}^n$ while we need a vector in \mathbb{R}^n .

To complete the proof, note that $s_{i_0} = \xi + i\eta$ with $\xi, \eta \in \mathbb{R}^n$.

$$\implies 1 \leq \|A^k s_{i_0}\|^2 = \|A^k \xi + iA^k \eta\|^2 = \|A^k \xi\|^2 + \|A^k \eta\|^2. \quad (11.74)$$

Therefore, either $\|A^k \xi\| \geq \frac{1}{\sqrt{2}}$ or $\|A^k \eta\| \geq \frac{1}{\sqrt{2}}$.

Without loss of generality assume $\|A^k \xi\| \geq \frac{1}{\sqrt{2}}$. Now let $y_0 = \xi$ to get

$$\underbrace{\|y_k\| = \|A^k y_0\|}_{\text{true for every } k \geq 0} \geq \frac{1}{\sqrt{2}} \implies \lim_{k \rightarrow \infty} \|y_k\| \neq 0. \quad (11.75)$$

Solution (3.2).

(i) The equation of motion for the sliding ball is given by

$$mR^2\ddot{\alpha} + bR^2\dot{\alpha} + mR^2(g/R - \Omega^2 \cos(\alpha)) \sin(\alpha) = 0 \quad (11.76)$$

Where we'll define $\nu = R\Omega^2/g$

We can define $x_1 = \alpha$, $x_2 = \dot{\alpha}$ in order to write the ODE as $\dot{x} = f(x)$ where

$$x = \begin{pmatrix} x_1 \\ x_2 \end{pmatrix} \quad \text{and} \quad f(x) = \begin{pmatrix} x_2 \\ -\frac{g}{R}[1 - \nu \cos(x_1)] \sin(x_1) - \frac{b}{m}x_2 \end{pmatrix} \quad (11.77)$$

In other words

$$\begin{bmatrix} \dot{x}_1 \\ \dot{x}_2 \end{bmatrix} = \begin{bmatrix} x_2 \\ -\frac{g}{R}[1 - \nu \cos(x_1)] \sin(x_1) - \frac{b}{m}x_2 \end{bmatrix} \quad (11.78)$$

Fixed points are found when $f(x) = 0$. This implies that $x_2 = 0$ and $[1 - \nu \cos(x_1)] \sin(x_1) = 0$

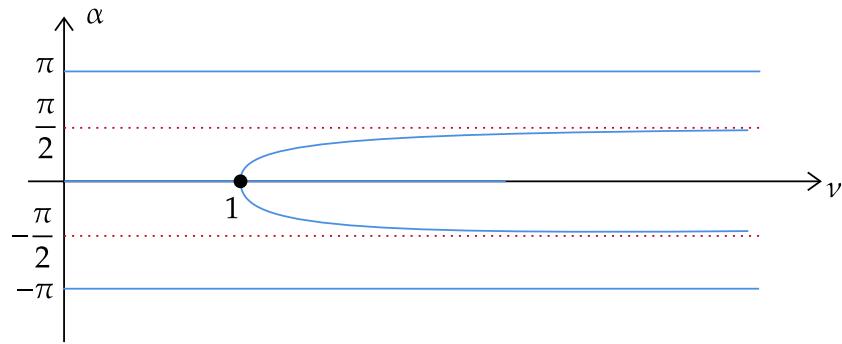
Case 1: $\nu < 1$: Only two fixed points exist: $(0, 0)$ and $(\pi, 0)$ [Note: the fixed point $(-\pi, 0)$ is physically identical to the fixed point $(\pi, 0)$. Therefore, we only discuss $(\pi, 0)$]

Case 2: $\nu > 1$:

Two additional fixed points emerge that correspond to the solutions of $\cos(x_1) = \frac{1}{\nu}$.

Let $\alpha_0 \in (0, \pi)$ be the positive solution: $\cos(\alpha_0) = \frac{1}{\nu}$. Then the fixed points in this case are:

$(0, 0), (\pi, 0), (\alpha_0, 0)$ and $(-\alpha_0, 0)$



The blue curves mark the location of the fixed points.

(ii) First we compute $\nabla f(x_1, x_2)$:

$$\nabla f(x_1, x_2) = \begin{pmatrix} 0 & 1 \\ \frac{g}{R}[2\nu \cos^2(x_1) - \cos(x_1) - \nu] & -\frac{b}{m} \end{pmatrix} \quad (11.79)$$

who's eigenvalues are given by

$$\lambda_{\pm} = -\frac{b}{2m} \pm \sqrt{\left(\frac{b}{2m}\right)^2 + \frac{g}{R}[2\nu \cos^2(x_1) - \cos(x_1) - \nu]} \quad (11.80)$$

Now we investigate the linear stability of each fixed point:

Fixed point $(0, 0)$:

$$\lambda_{\pm} = -\frac{b}{2m} \pm \sqrt{\left(\frac{b}{2m}\right)^2 + \frac{g}{R}(\nu - 1)} \quad (11.81)$$

- $\nu < 1 \implies \operatorname{Re}(\lambda_+) < 0$ and $\operatorname{Re}(\lambda_-) < 0$. Therefore $(0, 0)$ is asymptotically stable.
- $\nu > 1 \implies \operatorname{Re}(\lambda_+) > 0$ and $\operatorname{Re}(\lambda_-) < 0$. Therefore $(0, 0)$ is unstable.

Fixed points $(\pm\pi, 0)$:

$$\lambda_{\pm} = -\frac{b}{2m} \pm \sqrt{\left(\frac{b}{2m}\right)^2 + \frac{g}{R}(\nu + 1)} \quad (11.82)$$

For any $\nu \geq 0$, $\operatorname{Re}(\lambda_+) > 0 \implies (\pm\pi, 0)$ is unstable for any $\nu \geq 0$.

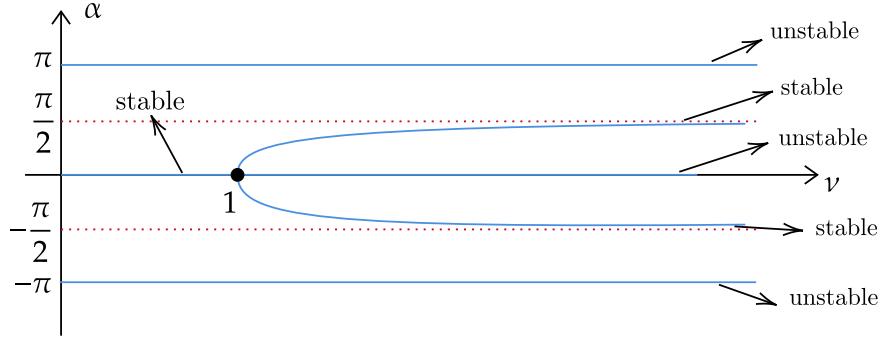
Fixed points $(\pm\alpha_0, 0)$ Remember that these fixed points only exist when $\nu > 1$. Also $\cos(\pm\alpha_0) = \frac{1}{\nu}$

$$\lambda_{\pm} = -\frac{b}{2m} \pm \sqrt{\left(\frac{b}{2m}\right)^2 + \frac{g}{R} \left(\frac{1 - \nu^2}{\nu}\right)} \quad (11.83)$$

For any $\nu > 1$, $\operatorname{Re}(\lambda_+) < 0$ and $\operatorname{Re}(\lambda_-) < 0$. Therefore the fixed points $(\pm\alpha_0, 0)$ are asymptotically stable.

The bifurcation of equilibria occurs at $\nu = 1 \implies \Omega^2 = \frac{g}{R} \implies \Omega = \pm\sqrt{\frac{g}{R}}$

Solution (3.3).



(i) By setting $f(x) = 0$, we obtain three fixed points for $\dot{x} = f(x)$. This can be seen by noting that the first equation $a(y_0 - x_0) = 0$ implies $x_0 = y_0$. From the second equation, we get

$$bx_0 - y_0 - x_0 z_0 = 0 \quad (11.84)$$

$$z_0 = b - 1. \quad (11.85)$$

Then the third equation reduces to

$$x_0 y_0 - c z_0 = 0 \quad (11.86)$$

$$x_0 = y_0 = \pm \sqrt{c(b-1)}. \quad (11.87)$$

The resulting three fixed points are

$$P_1 : x_0 = y_0 = z_0 = 0 \quad (11.88)$$

$$P_2 : x_0 = y_0 = \sqrt{c(b-1)}, z_0 = b-1 \quad (11.89)$$

$$P_3 : x_0 = y_0 = -\sqrt{c(b-1)}, z_0 = b-1. \quad (11.90)$$

For the system to have these three fixed points we must have $b > 1$. If $0 < b \leq 1$, the only fixed point is P_1 .

Let A denote $Df(x_0, y_0, z_0)$. Then:

$$A = \begin{pmatrix} -a & a & 0 \\ b - z_0 & -1 & -x_0 \\ y_0 & x_0 & -c \end{pmatrix} \quad (11.91)$$

The eigenvalues λ of A are defined as the roots of the characteristic polynomial

$$\det |A - \lambda I| = 0.$$

For the matrix A this takes the form

$$A = \begin{vmatrix} -a - \lambda & a & 0 \\ b - z_0 & -1 - \lambda & -x_0 \\ y_0 & x_0 & -c - \lambda \end{vmatrix} = 0. \quad (11.92)$$

We may expand this determinant according to the first row as

$$(-a - \lambda) \begin{vmatrix} -1 - \lambda & -x_0 \\ x_0 & -c - \lambda \end{vmatrix} - a \begin{vmatrix} b - z_0 & -x_0 \\ y_0 & -c - \lambda \end{vmatrix} = 0 \quad (11.93)$$

$$- (a + \lambda) [(1 + \lambda)(c + \lambda) + x_0^2] - a(b - z_0)(-c - \lambda) - ax_0y_0 = 0 \quad (11.94)$$

After collecting the coefficients of the different powers of λ , the characteristic equation of A is:

$$\lambda^3 + (a + c + 1)\lambda^2 + [ac + a + c + x_0^2 + a(z_0 - b)]\lambda + ac(z_0 - b + 1) + x_0^2a + ax_0y_0 = 0 \quad (11.95)$$

Stability of P_1 :

$$\lambda_3 + (a + c + 1)\lambda^2 + (ac + a + c - ab)\lambda - ac(b - 1) = 0 \quad (11.96)$$

A necessary condition for all roots of the above polynomial to be negative is that all its coefficients have the same sign. But here $-ac(b - 1) < 0$ while λ^3 has a positive coefficient (i.e. +1). $\Rightarrow A$ has a positive eigenvalue.

$\Rightarrow P_1$ is unstable.

Stability of P_2, P_3 : Note that in the characteristic equation, we only have products of x_0 and y_0 , i.e. x_0^2 and x_0y_0 . This means that the equation is invariant to changing the sign of x_0 and y_0 and we get the same eigenvalues at P_2 and at P_3 .

The Routh-Hurwitz determinants are:

$$d_1 = 2ac(b - 1) > 0 \quad (11.97)$$

$$d_2 = (a + b)c > 0 \quad (11.98)$$

$$d_3 = \begin{vmatrix} (a + b)c & 2ac(b - 1) \\ 1 & a + c + 1 \end{vmatrix} = (a + b)(a + c + 1)c - 2ac(b - 1) \quad (11.99)$$

For P_2 and P_3 to be unstable, we must have $d_3 < 0$

$$d_3 < 0 \iff b > \frac{a(3 + a + c)}{a - (c + 1)} \underbrace{\text{follows from } a > c + 1}_{>} 1 \quad (11.100)$$

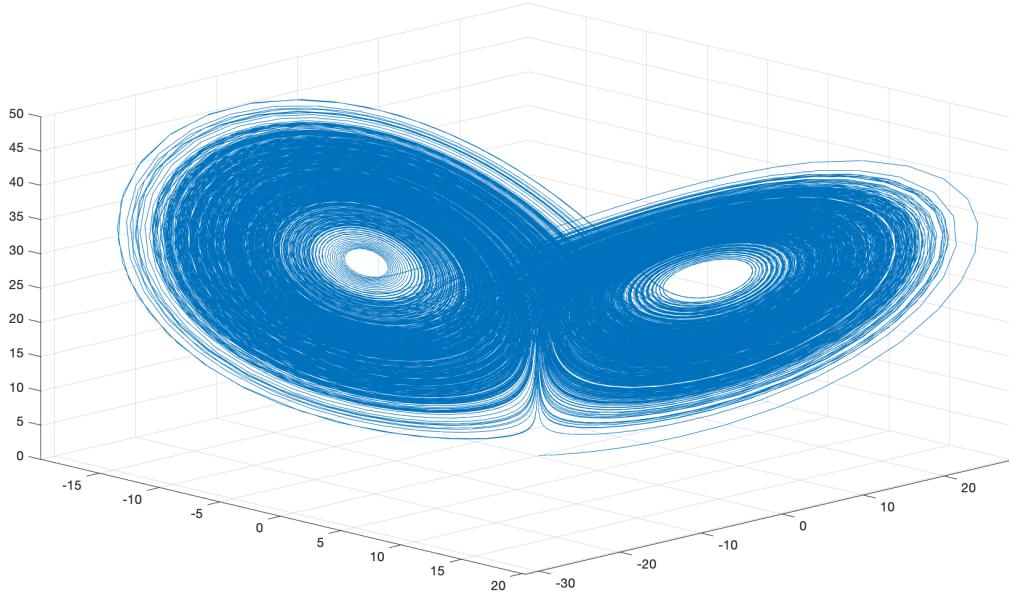


Figure 11.1: Simulation of the Lorenz equations.

(ii)

Solution (3.4).

$$E(x, \dot{x}) = \frac{1}{2}\dot{x}^2 + (1 - \cos(x)) \quad (11.101)$$

$$y = (y_1, y_2) := (x, \dot{x}) \quad (11.102)$$

$$y = f(y) = \begin{bmatrix} y_2 \\ -cy_2 - \sin(y_1) \end{bmatrix} \implies E(y) = \frac{1}{2}y_2^2 + (1 - \cos(y_1)) \quad (11.103)$$

(i)

$$E(0) = 0, DE(0) = 0, D^2E(0) = \begin{bmatrix} 1 & 0 \\ 0 & 1 \end{bmatrix} \quad (11.104)$$

\implies Hessian is positive definite.

\implies E is positive-definite near the origin

(ii)

$$\dot{E}(y) = \langle DE(y), f(y) \rangle = (\sin(y_1), y_2) \cdot (y_2, -cy_2 - \sin(y_1)) \quad (11.105)$$

$$= \sin(y_1)y_2 - cy_2^2 - \sin(y_1)y_2 \quad (11.106)$$

$$= -cy_2^2 \leq 0 \quad (11.107)$$

E is positive definite around the origin and \dot{E} is negative semi-definite.

Indeed, we cannot find an open set U around the origin where

$$\dot{E}(y) < 0 \quad \forall y \in U \setminus \{0\} \quad [\dot{E}(y) = 0 \text{ for any } y = (y_1, 0) \text{ with } y_1 \neq 0] \quad (11.108)$$

Thus, theorem 2 is not applicable to conclude nonlinear asymptotic stability of the origin.

Solution (3.5).

(i) By setting $f(x) = 0$, we obtain three fixed points for $\dot{x} = f(x)$. This can be seen by noting that the first equation $a(y_0 - x_0) = 0$ implies $x_0 = y_0$. From the second equation, we get

$$bx_0 - y_0 - x_0 z_0 = 0 \quad (11.109)$$

$$z_0 = b - 1. \quad (11.110)$$

Then the third equation reduces to

$$x_0 y_0 - c z_0 = 0 \quad (11.111)$$

$$x_0 = y_0 = \pm \sqrt{c(b-1)}. \quad (11.112)$$

The resulting three fixed points are

$$P_1 : x_0 = y_0 = z_0 = 0 \quad (11.113)$$

$$P_2 : x_0 = y_0 = \sqrt{c(b-1)}, z_0 = b-1 \quad (11.114)$$

$$P_3 : x_0 = y_0 = -\sqrt{c(b-1)}, z_0 = b-1. \quad (11.115)$$

For the system to have these three fixed points we must have $b > 1$. If $0 < b \leq 1$, the only fixed point is P_1 .

Let A denote $Df(x_0, y_0, z_0)$. Then:

$$A = \begin{pmatrix} -a & a & 0 \\ b - z_0 & -1 & -x_0 \\ y_0 & x_0 & -c \end{pmatrix} \quad (11.116)$$

The eigenvalues λ of A are defined as the roots of the characteristic polynomial $\det|A - \lambda I| = 0$. For the matrix A this takes the form

$$A = \begin{vmatrix} -a - \lambda & a & 0 \\ b - z_0 & -1 - \lambda & -x_0 \\ y_0 & x_0 & -c - \lambda \end{vmatrix} = 0. \quad (11.117)$$

We may expand this determinant according to the first row as

$$(-a - \lambda) \begin{vmatrix} -1 - \lambda & -x_0 \\ x_0 & -c - \lambda \end{vmatrix} - a \begin{vmatrix} b - z_0 & -x_0 \\ y_0 & -c - \lambda \end{vmatrix} = 0 \quad (11.118)$$

$$- (a + \lambda) [(1 + \lambda)(c + \lambda) + x_0^2] - a(b - z_0)(-c - \lambda) - ax_0y_0 = 0 \quad (11.119)$$

After collecting the coefficients of the different powers of λ , the characteristic equation of A is:

$$\lambda^3 + (a + c + 1)\lambda^2 + [ac + a + c + x_0^2 + a(z_0 - b)]\lambda + ac(z_0 - b + 1) + x_0^2a + ax_0y_0 = 0 \quad (11.120)$$

Stability of P_1 :

$$\lambda_3 + (a + c + 1)\lambda^2 + (ac + a + c - ab)\lambda - ac(b - 1) = 0 \quad (11.121)$$

A necessary condition for all roots of the above polynomial to be negative is that all its coefficients have the same sign. But here $-ac(b - 1) < 0$ while λ^3 has a positive coefficient (i.e. +1). $\implies A$ has a positive eigenvalue.

$\implies P_1$ is unstable.

Stability of P_2, P_3 : Note that in the characteristic equation, we only have products of x_0 and y_0 , i.e. x_0^2 and x_0y_0 . This means that the equation is invariant to changing the sign of x_0 and y_0 and we get the same eigenvalues at P_2 and at P_3 .

The Routh-Hurwitz determinants are:

$$d_1 = 2ac(b - 1) > 0 \quad (11.122)$$

$$d_2 = (a + b)c > 0 \quad (11.123)$$

$$d_3 = \begin{vmatrix} (a + b)c & 2ac(b - 1) \\ 1 & a + c + 1 \end{vmatrix} = (a + b)(a + c + 1)c - 2ac(b - 1) \quad (11.124)$$

For P_2 and P_3 to be unstable, we must have $d_3 < 0$

$$d_3 < 0 \iff b > \frac{a(3 + a + c)}{a - (c + 1)} \underbrace{\text{follows from } a > c + 1}_{>} 1 \quad (11.125)$$

Solution (3.6). We use Krasovski's theorem with $V = E, U \subset (-\pi, \pi) \times \mathbb{R}$ open set around the origin in $S' \times \mathbb{R}$ such that the statements (i) & (ii) in the hypothesis of Krasovski are satisfied as shown above in part a).

$$S = \{y \in U \mid \dot{E}(y) = 0\} \subset \underbrace{\{(y_1, 0) \mid y_1 \in (-\pi, \pi)\}}_{\tilde{S}} \quad (11.126)$$

Indeed, the only trajectory of the system completely contained in the set \tilde{S} on the y_1 -axis is the origin (cf. phase portrait). $\Rightarrow S$ contains only the fixed point as a trajectory of the system.

Hence, the hypothesis of Krasovski's theorem is satisfied and the origin is asymptotically stable for the nonlinear damped pendulum.

Solution (3.7).

$$\dot{E}(y) = \langle DE(y), f(y) \rangle = (\sin(y_1), y_2) \cdot (y_2, -cy_2 - \sin(y_1)) \quad (11.127)$$

$$= \sin(y_1)y_2 - cy_2^2 - \sin(y_1)y_2 \quad (11.128)$$

$$= -cy_2^2 \leq 0 \quad (11.129)$$

E is positive definite around the origin and \dot{E} is negative semi-definite.

Indeed, we cannot find an open set U around the origin where

$$\dot{E}(y) < 0 \quad \forall y \in U \setminus \{0\} \quad [\dot{E}(y) = 0 \text{ for any } y = (y_1, 0) \text{ with } y_1 \neq 0] \quad (11.130)$$

Thus, theorem 2 is not applicable to conclude nonlinear asymptotic stability of the origin.

Solution. First construct the function:

$$\bar{E}(q, \dot{q}) = E(q, \dot{q}) - V(q_0) \quad (11.131)$$

$$= \frac{1}{2}\dot{q}^T M \dot{q} + V(q) - V(q_0) \quad (11.132)$$

Now, at $(q, \dot{q}) = (q_0, 0)$ we have $\bar{E}(q_0, 0) = 0$

Note that $M(q)$ is positive definite for all q and $V(q) - V(q_0)$ is positive around $q = q_0$. (Since V has a local minimum at q_0)

$\Rightarrow \bar{E}(q, \dot{q})$ is positive definite around $(q_0, 0)$.

But $\frac{d\bar{E}}{dt} = \frac{dE}{dt}$ since $V(q_0)$ is a constant.

We show that $\frac{dE}{dt} = 0$.

First note that, in general, the Lagrangian equation of motion is a system of n coupled equations with each equation given by:

$$\frac{d}{dt} \left[\frac{\partial L}{\partial \dot{q}_k} \right] - \frac{\partial L}{\partial q_k} = 0 \quad k = 1, 2, \dots, n \quad (11.133)$$

Multiply each equation by \dot{q}_k and sum over k to get:

We'll use Einstein's notation: sum over repeated indices.

$$\dot{q}_k \left[\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_k} \right) - \frac{\partial L}{\partial q_k} \right] = 0 \quad (11.134)$$

Since $L = \frac{1}{2} M_{ij} \dot{q}_i \dot{q}_j - V$ we have:

$$\frac{\partial L}{\partial \dot{q}_k} = M_{ik} \dot{q}_i \quad , \quad \frac{\partial L}{\partial q_k} = \frac{1}{2} \frac{\partial M_{ij}}{\partial q_k} \dot{q}_i \dot{q}_j - \frac{\partial V}{\partial q_k} \quad (11.135)$$

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{q}_k} = M_{ik} \ddot{q}_i + \frac{\partial M_{ik}}{\partial q_j} \dot{q}_j \dot{q}_i \quad (11.136)$$

Substituting (11.135) into (11.134), we get

$$M_{ik} \ddot{q}_i \dot{q}_k + \frac{\partial M_{ik}}{\partial q_j} \dot{q}_j \dot{q}_i \dot{q}_k - \frac{1}{2} \frac{\partial M_{ij}}{\partial q_k} \dot{q}_i \dot{q}_j \dot{q}_k + \frac{\partial V}{\partial q_k} \dot{q}_k = 0 \quad (11.137)$$

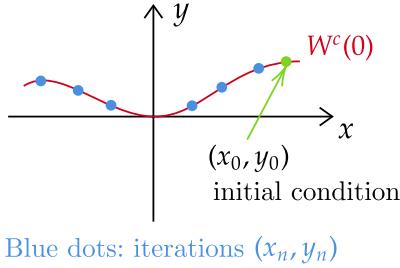
Since there is a sum over repeated indices we have:

$$M_{ik} \ddot{q}_i \dot{q}_k \equiv M_{ij} \ddot{q}_i \dot{q}_j \quad \frac{\partial M_{ik}}{\partial q_j} \dot{q}_j \dot{q}_i \dot{q}_k \equiv \frac{\partial M_{ij}}{\partial q_k} \dot{q}_i \dot{q}_j \dot{q}_k \quad (11.138)$$

$$\begin{aligned} & \Rightarrow M_{ij} \ddot{q}_i \dot{q}_j + \underbrace{\frac{1}{2} \frac{\partial M_{ij}}{\partial q_k} \dot{q}_i \dot{q}_j \dot{q}_k}_{= \frac{d}{dt} \left[\frac{1}{2} M_{ij} \dot{q}_i \dot{q}_j + V(q) \right]} + \frac{\partial V}{\partial q_k} \dot{q}_k = 0 \\ & \quad (11.139) \end{aligned}$$

$$\Rightarrow \left[\frac{1}{2} \dot{q}^T M(q) \dot{q} + V(q) \right] = 0 \Rightarrow \frac{dE}{dt} = 0 \Rightarrow \frac{d\bar{E}}{dt} = 0 \quad (11.140)$$

Using \bar{E} as the Lyapunov function, we conclude that $(q_0, 0)$ is a stable equilibrium point.



Solution (4.1). (i) Let the graph of the center manifold near the origin be given by $y = h(x)$, $h : \mathbb{R}^c \rightarrow \mathbb{R}^d$. By the invariance of the center manifold we have $y_n = h(x_n)$ for all n .

$$y_{n+1} = h(x_{n+1}) \quad (11.141)$$

But

$$y_{n+1} = By_n + g(x_n, y_n) \quad (11.142)$$

$$= Bh(x_n) + g(x_n, h(x_n)) \quad (11.143)$$

and

$$x_{n+1} = Ax_n + f(x_n, y_n) \quad (11.144)$$

$$= Ax_n + f(x_n, h(x_n)) \quad (11.145)$$

Hence

$$Bh(x_n) + g(x_n, h(x_n)) = h[Ax_n + f(x_n, h(x_n))] \quad (11.146)$$

Therefore, the function $h : \mathbb{R}^c \rightarrow \mathbb{R}^d$ satisfies

$$\boxed{Bh(x) + g(x, h(x)) = h[Ax + f(x, h(x))]} \quad (11.147)$$

(ii) Here, $[A] = [1]$, $[B] = [\lambda]$, $f(x_n, y_n) = x_n y_n$ and $g(x_n, y_n) = -x_n^2$.

Since h passes through the origin and it is tangent to the x -axis, we have $h(0) = 0$ and $h'(0) = 0$. (Note that here $c = d = 1 \Rightarrow h : \mathbb{R} \rightarrow \mathbb{R}$).

Therefore, the Taylor expansion of h around $x = 0$ has the form

$$h(x) = ax^2 + bx^3 + \mathcal{O}(x^4) \quad (11.148)$$

Equation (11.147) for the current system is: $\lambda h(x) - x^2 = h(x + xh(x))$.

Substituting (11.148) in this equation we get

$$\lambda(ax^2 + bx^3 + \mathcal{O}(x^4)) - x^2 = h[x + ax^3 + \mathcal{O}(x^4)] \quad (11.149)$$

$$= a(x + ax^3 + \mathcal{O}(x^4))^2 + b(x + ax^3 + \mathcal{O}(x^4))^3 + \dots \quad (11.150)$$

$$\implies (\lambda a - 1)x^2 + \lambda b x^3 + \mathcal{O}(x^4) = ax^2 + bx^3 + \mathcal{O}(x^4) \quad (11.151)$$

Matching the exponents from both sides we obtain:

$$\lambda a - 1 = a \implies a = \frac{1}{\lambda - 1} \quad (11.152)$$

$$\lambda b = b \implies b = 0 \quad (11.153)$$

and finally

$$\boxed{h(x) = \frac{1}{\lambda - 1}x^2 + \mathcal{O}(x^4)} \quad (11.154)$$

- (iii) The center manifold near the origin satisfies $h(x) \approx \frac{1}{\lambda-1}x^2$. Hence, the dynamics on the center manifold satisfy

$$\begin{aligned} x_{n+1} &= x_n + \frac{1}{\lambda - 1}x_n^3 \\ \implies x_{n+1} &= x_n \left(1 + \frac{1}{\lambda - 1}x_n^2 \right) \end{aligned} \quad (11.155)$$

In the following, we show that the fixed point $x = 0$ of (11.155) is asymptotically stable.

First let $x_0 \in \mathbb{R}$ with $|x_0| > 0$ small enough be an initial condition. Then if

$$\left| 1 + \frac{1}{\lambda - 1}x_0^2 \right| < 1 \quad (11.156)$$

we have

$$|x_1| \leq \left| x_0 \left(1 + \frac{1}{\lambda - 1}x_0^2 \right) \right| < |x_0| \quad (11.157)$$

Inequality (11.156) holds if and only if $|x_1| < |x_0| < \sqrt{2(1-\lambda)}$. That is for any x_0 with $|x_0| < \sqrt{2(1-\lambda)}$ we have $|x_1| < |x_0| < \sqrt{2(1-\lambda)}$.

For such initial conditions, we have (by induction):

$$\cdots < |x_{n+1}| < |x_n| < \cdots < |x_1| < |x_0| < \sqrt{2(1-\lambda)} \quad (11.158)$$

This proves that the stability of the fixed point $x = 0$. To prove asymptotic stability, we need

$$\lim_{n \rightarrow \infty} x_n = 0 \quad (11.159)$$

Note that the sequence $\{|x_n|\}$ is a decreasing (due to (11.158)) sequence that is bounded from below ($|x_n| \geq 0$). Therefore, it must have a limit:

$$\lim_{n \rightarrow \infty} |x_n| = \alpha \quad (11.160)$$

This limit, in general, doesn't have to be zero. But taking the limit $n \rightarrow \infty$ in equation (11.155) we get

$$\lim_{n \rightarrow \infty} |x_{n+1}| = \lim_{n \rightarrow \infty} |x_n| \left(1 + \frac{1}{\lambda - 1} \lim_{n \rightarrow \infty} |x_n|^2 \right) \quad (11.161)$$

$$\alpha = \alpha \left(1 + \frac{1}{\lambda - 1} \alpha^2 \right) \implies \alpha = 0 \implies \lim_{n \rightarrow \infty} x_n = 0 \quad (11.162)$$

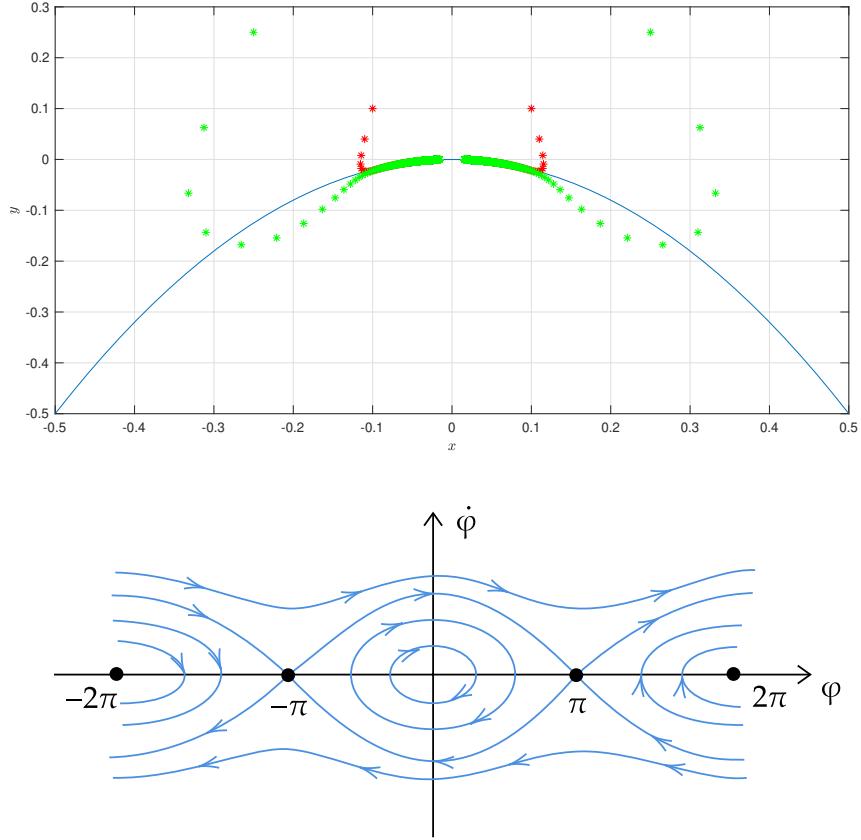
Therefore the fixed point is asymptotically stable. The following figure shows the iterations of the map for four initial conditions marked by square symbols. The higher iterations are marked by dots. The black curve marks the approximate center manifold:

$$y = \frac{1}{\lambda - 1} x^2 \quad (11.163)$$

Solution (4.2). Let $x_1 = x$ and $x_2 = \dot{x}$. Then:

$$\begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -\sin(x_1) \end{cases} \quad (11.164)$$

By linearization, one can show that the fixed point $(\pi, 0)$ is an unstable hyperbolic fixed point with stable and unstable linear subspaces spanned by $\begin{pmatrix} 1 \\ 1 \end{pmatrix}$ and $\begin{pmatrix} 1 \\ -1 \end{pmatrix}$, respectively: For convenience,



we shift the origin by the transformation $\xi_1 = x_1 - \pi$ and $\xi_2 = x_2$ such that in (ξ_1, ξ_2) the origin is the hyperbolic fixed point.

In this coordinate system, the dynamical system becomes:

$$\begin{cases} \dot{\xi}_1 = \xi_2 \\ \dot{\xi}_2 = \sin(\xi_1) \end{cases} \quad (\text{since } -\sin(x_1) = -\sin(\xi_1 + \pi) = \sin(\xi_1)) \quad (11.165)$$

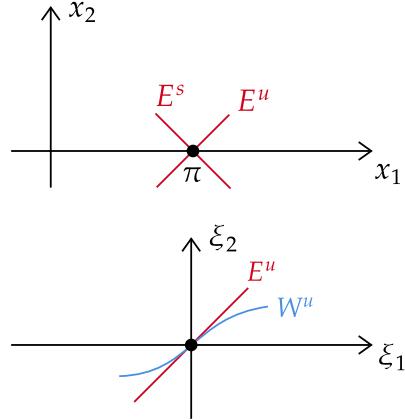
The unstable manifold passing through the origin is a graph over ξ_1 and tangent to E^U . If this graph is given by $\xi_2 = h(\xi_1)$, the Taylor expansion of h looks like:

$$h(\xi_1) = \underbrace{0}_{h(0)=0} + \underbrace{\xi_1}_{h'(0)=1} + a\xi_1^2 + b\xi_1^3 + \mathcal{O}(\xi_1^4) \quad (11.166)$$

By invariance of the unstable manifold we have $\dot{\xi}_2 = h'(\xi_1)\dot{\xi}_1$

Therefore,

$$\sin(\xi_1) = (1 + 2a\xi_1 + 3b\xi_1^2 + \mathcal{O}(3))(\xi_1 + a\xi_1^2 + b\xi_1^3 + \mathcal{O}(4)) \quad (11.167)$$



The Taylor expansion of $\sin(\xi_1)$ around $\xi_1 = 0$ reads

$$\sin(\xi_1) = \xi_1 - \frac{1}{6}\xi_1^3 + \mathcal{O}(\xi_1^5) \quad (11.168)$$

Matching exponents we get $a = 0$ and $b = -\frac{1}{24}$.

Therefore, the graph of unstable manifold satisfies:

$$\xi_2 = \xi_1 - \frac{1}{24}\xi_1^3 + \mathcal{O}(\xi_1^4) \quad (11.169)$$

or

$$x_2 = x_1 - \pi - \frac{1}{24}(x_1 - \pi)^3 + \mathcal{O}(|x_1 - \pi|^4) \quad (11.170)$$

Solution (4.3).

(i) Linearized dynamics around fixed point $(0, 0)$

$$\dot{\eta} = A\eta, \quad A = \begin{bmatrix} 0 & 1 \\ \beta & -\delta \end{bmatrix}, \quad \text{eig}(A) = \lambda_{1,2} = -\delta \pm \sqrt{\delta^2 - \beta} \quad (11.171)$$

Note that $\lambda_1 = 0$, $\lambda_2 = -\delta$ for $\beta = 0$. Thus, by the center manifold theorem, we have a 1-dimensional center manifold passing through the origin and a unique 1-dimensional stable manifold.

Consider the extended system

$$\dot{\beta} = 0 \quad (11.172)$$

$$\begin{bmatrix} \dot{u} \\ \dot{v} \end{bmatrix} = \underbrace{\begin{bmatrix} 0 & 1 \\ 0 & -\delta \end{bmatrix}}_B \begin{bmatrix} u \\ v \end{bmatrix} + \begin{bmatrix} 0 \\ \beta u - u^2 \end{bmatrix} \quad (11.173)$$

$$\text{Eigenvalues of } B : \lambda_1 = 0, \lambda_2 = -\delta \quad (11.174)$$

$$\text{Eigenvectors of } B : e_1 = \begin{bmatrix} 1 \\ 0 \end{bmatrix}, e_2 = \begin{bmatrix} \frac{1}{\delta} \\ -1 \end{bmatrix} \quad (11.175)$$

From the eigenvalues and eigenvectors, we can perform a change of coordinates

$$\begin{bmatrix} u \\ v \end{bmatrix} = T \begin{bmatrix} x \\ y \end{bmatrix}, T = [e_1 | e_2] = \begin{bmatrix} 1 & \frac{1}{\delta} \\ 0 & -1 \end{bmatrix}, T^{-1} = \begin{bmatrix} 1 & \frac{1}{\delta} \\ 0 & -1 \end{bmatrix} = T \quad (11.176)$$

$$\implies u = x + \frac{y}{\delta}, v = -y \quad (11.177)$$

$$\begin{aligned} \implies \begin{bmatrix} \dot{x} \\ \dot{y} \end{bmatrix} &= T^{-1} B T \begin{bmatrix} x \\ y \end{bmatrix} + T^{-1} \begin{bmatrix} 0 \\ \beta u - u^2 \end{bmatrix} \\ &= \begin{bmatrix} 0 & 0 \\ 0 & -\delta \end{bmatrix} \begin{bmatrix} x \\ y \end{bmatrix} + \begin{bmatrix} \frac{1}{\delta} (\beta (x + \frac{y}{\delta}) - (x + \frac{y}{\delta})^2) \\ -\beta (x + \frac{y}{\delta}) + (x + \frac{y}{\delta})^2 \end{bmatrix} \end{aligned} \quad (11.178)$$

Seek center manifold as a graph over center subspace locally as

$$\begin{aligned} y &= h(x, \beta) = a_1 x^2 + a_2 x \beta + \underbrace{a_3 \beta^2}_{=0} + \mathcal{O}(3) \\ \dot{y} &= \frac{\partial h}{\partial x} \dot{x} + \frac{\partial h}{\partial \beta} \underbrace{\dot{\beta}}_0 \end{aligned} \quad (11.179)$$

Note: We cancel the term $a_3 \beta^2$ to respect the existence of the fixed point.

Use invariance in ((11.179)):

$$\implies \dot{y} = (2a_1 x + a_2 \beta) \left[\frac{1}{\delta} \left(\beta \left(x + \frac{h(x, \beta)}{\delta} \right) - \left(x + \frac{h(x, \beta)}{\delta} \right)^2 \right) \right] \quad (11.180)$$

$$\text{But also } \dot{y} = -\delta h(x, \beta) - \beta \left(x + \frac{h(x, \beta)}{\delta} \right) + \left(x + \frac{h(x, \beta)}{\delta} \right)^2 \quad (11.181)$$

Comparing $\mathcal{O}(2)$ terms in (11.180) & (11.181) we get:

$$x^2 : \quad -\delta a_1 + 1 = 0 \implies a_1 = \frac{1}{\delta} \quad (11.182)$$

$$x\beta : \quad -\delta a_2 - 1 = 0 \implies -a_2 = \frac{1}{\delta} \quad (11.183)$$

Thus, the β -dependent center manifold is given by

$$h(x, \beta) = \frac{x^2}{\delta} - \frac{x\beta}{\delta} + \mathcal{O}(3) \quad (11.184)$$

Substitute (11.184) into first equation in (11.178) to obtain reduced dynamics on the center manifold: $W_\beta^C(0)$ up to quadratic order.

$$\dot{x} = \frac{1}{\delta} \left[\beta \left(x + \frac{h(x, \beta)}{\delta} \right) - \left(x + \frac{h(x, \beta)}{\delta} \right)^2 \right] \quad (11.185)$$

$$= \frac{1}{\delta} [\beta x - x^2] + \mathcal{O}(3) \quad (11.186)$$

(ii)

$$\dot{x} = \frac{1}{\delta} [\beta x - x^2] \quad (11.187)$$

Fixed points:

$$x = 0, \quad (11.188)$$

$$\beta = x \quad (11.189)$$

Solution (4.4). (i) Let $x_1 = x$ and $x_2 = \dot{x}$. Then the system can be written as

$$\begin{cases} \dot{x}_1 = x_2 \\ \dot{x}_2 = -\frac{k}{m}x_1 - \frac{c}{m}x_2 + \frac{1}{m}F_f(x_2). \end{cases} \quad (11.190)$$

The fixed point is at

$$x_1^0 = \frac{1}{k}F_f(0) = \frac{mg\mu_0}{k} (1 + e^{-\beta|v_0|}) \operatorname{sign}(v_0), \quad x_2^0 = 0. \quad (11.191)$$

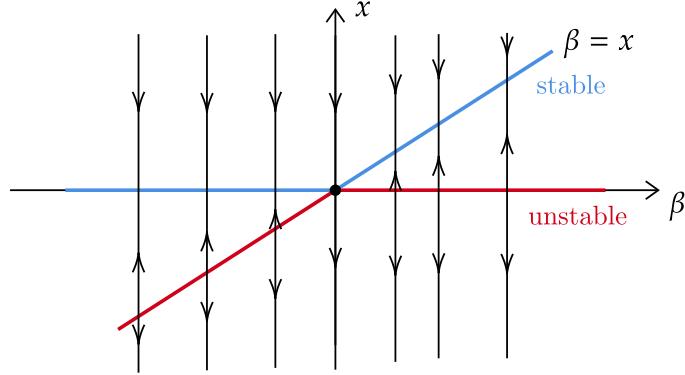


Figure 11.2: Transcritical bifurcation

Let us now shift the origin to the fixed point by introducing new coordinates as $z_1 = x_1 - x_1^0$ and $z_2 = x_2$. Then

$$\dot{z}_1 = z_2 \quad (11.192)$$

$$\dot{z}_2 = -\frac{k}{m}z_1 - \frac{c}{m}z_2 + \frac{1}{m}F_f(z_2) - \frac{1}{m}F_f(0), \quad (11.193)$$

with the fixed point at $z_1 = z_2 = 0$. The linearized system is given by

$$\dot{\xi} = \begin{pmatrix} 0 & 1 \\ -\frac{k}{m} & \mu \end{pmatrix} \xi, \quad (11.194)$$

where we have introduced the parameter

$$\boxed{\mu = \frac{1}{m} (F'_f(0) - c) = g\beta\mu_0 e^{-\beta|v_0|} - \frac{c}{m}}. \quad (11.195)$$

The eigenvalues of the coefficient matrix are

$$\lambda_{1,2} = \frac{\mu}{2} \pm \frac{1}{2} \sqrt{\mu^2 - \frac{4k}{m}}. \quad (11.196)$$

For the remainder of the discussion, let us assume that $|\mu|$ is not too big; specifically $\mu^2 < \frac{4k}{m}$. This is in line with our previous assumption to treat μ as a bifurcation parameter. In this case, the real part of $\lambda_{1,2}$ can simply be read off from (11.196) as

$$\operatorname{Re} (\lambda_{1,2}) = \frac{\mu}{2}. \quad (11.197)$$

As a result, if $\mu < 0$ then $\operatorname{Re}(\lambda_1) < 0$ and $\operatorname{Re}(\lambda_2) < 0$, hence the fixed point is asymptotically stable by the Hartman-Grobman theorem. This condition translates into

$$\mu < 0 \Leftrightarrow g\beta\mu_0 e^{-\beta|v_0|} < \frac{c}{m} \Leftrightarrow |v_0| > \frac{1}{\beta} \log \left(\frac{mg\beta\mu_0}{c} \right). \quad (11.198)$$

Hence, if

$$|v_0| > \frac{1}{\beta} \log \left(\frac{mg\beta\mu_0}{c} \right), \quad (11.199)$$

then $(z_1 = z_2 = 0)$ is an asymptotically stable fixed point.

Remark 11.2. Note that if the viscous damping c is large enough such that $c > mg\beta\mu_0$ then $\log \left(\frac{mg\beta\mu_0}{c} \right) < 0$. Then the condition (11.199) is satisfied for any $v_0 \neq 0$ and $(z_1 = z_2 = 0)$ is asymptotically stable for any v_0 .

- (ii) For $0 < \mu \ll 1$ we have $\operatorname{Re}(\lambda_{1,2}) > 0$. At $\mu = 0$ the eigenvalues cross the imaginary axis at

$$\lambda_{1,2}(\mu = 0) = \pm i \sqrt{\frac{4k}{m}}. \quad (11.200)$$

Now let $-1 \ll \mu < 0$. Then the eigenvalues are

$$\lambda_{1,2} = \frac{\mu}{2} \pm i \frac{1}{2} \sqrt{\frac{4k}{m} - \mu^2}. \quad (11.201)$$

Define

$$\delta(\mu) = \frac{\mu}{2}, \quad \omega(\mu) = \frac{1}{2} \sqrt{\frac{4k}{m} - \mu^2}.$$

(11.202)

Then, we can separate the linear and nonlinear terms from the system (11.192) and write it as

$$\begin{pmatrix} \dot{z}_1 \\ \dot{z}_2 \end{pmatrix} = A \begin{pmatrix} z_1 \\ z_2 \end{pmatrix} + \begin{pmatrix} 0 \\ \frac{1}{m}F_f(z_2) - \frac{1}{m}F'_f(0)z_2 - \frac{1}{m}F_f(0) \end{pmatrix}, \quad (11.203)$$

where we have denoted the linear part as

$$A = \begin{pmatrix} 0 & 1 \\ -\frac{k}{m} & \mu \end{pmatrix}.$$

The eigenvalues of A are $\delta(\mu) \pm i\omega(\mu)$. To simplify the calculation of the eigenvectors, we note that (as a consequence of (11.202))

$$\mu = 2\delta \quad \frac{k}{m} = \delta^2 + \mu^2 \quad (11.204)$$

and hence

$$A = \begin{pmatrix} 0 & 1 \\ -\delta^2 - \omega^2 & 2\delta \end{pmatrix}. \quad (11.205)$$

We then search for a vector s such that

$$As - (\delta + i\omega)s = \begin{pmatrix} -\delta - i\omega & 1 \\ -\delta^2 - \omega^2 & \delta - i\omega \end{pmatrix} \begin{pmatrix} s_1 \\ s_2 \end{pmatrix}. \quad (11.206)$$

For example, the non-normalized vector

$$s = \begin{pmatrix} 1 \\ \delta + i\omega \end{pmatrix} \quad (11.207)$$

is a good choice. To separate the real and imaginary parts of the eigenvector we write it as

$$s = \begin{pmatrix} 1 \\ \delta(\mu) \end{pmatrix} \pm i \begin{pmatrix} 0 \\ \omega(\mu) \end{pmatrix}. \quad (11.208)$$

Selecting the real and imaginary parts of s as basis vectors then puts A in the desired block-diagonal form, that is

$$A = VDV^{-1}, \quad (11.209)$$

where

$$D = \begin{pmatrix} \delta & \omega \\ -\omega & \delta \end{pmatrix} \text{ and } V = \begin{pmatrix} 1 & 0 \\ \delta & \omega \end{pmatrix}. \quad (11.210)$$

This means, that under the change of variables

$$\begin{pmatrix} z_1 \\ z_2 \end{pmatrix} = V \begin{pmatrix} u \\ v \end{pmatrix} \quad (11.211)$$

we get the transformed dynamical system

$$\begin{pmatrix} \dot{u} \\ \dot{v} \end{pmatrix} = \begin{pmatrix} \delta & \omega \\ -\omega & \delta \end{pmatrix} \begin{pmatrix} u \\ v \end{pmatrix} + \frac{1}{m\omega} \begin{pmatrix} 0 \\ F_f(\delta u + \omega v) - F'_f(0)(\delta u + \omega v) - F_f(0) \end{pmatrix}. \quad (11.212)$$

This is the desired form for the dynamical system defined in the problem description. Note that we may put

$$f(x, y, \mu) = 0 \quad (11.213)$$

$$g(x, y, \mu) = \frac{1}{m\omega} F_f(\delta u + \omega v) - F'_f(0)(\delta u + \omega v) - F_f(0) \quad (11.214)$$

to compute the desired parameters

$$d_0 = \delta'(0) = \frac{1}{2}, \quad \omega_0 = \omega(0) = \sqrt{\frac{k}{m}}, \quad a_0 = \frac{kg\mu_0\beta^3 e^{-\beta|v_0|}}{16m}, \quad (11.215)$$

where we have used that

$$g_{vvv}(0, 0, 0) = \frac{k}{m^2} F'''(0) = \frac{kg\mu_0\beta^3 e^{-\beta|v_0|}}{m}. \quad (11.216)$$

According to the Hopf-Bogdanov theorem, the radial component of the normal form of the dynamics can be written as

$$\boxed{\dot{r} = r \left(\frac{\mu}{2} + \frac{kg\mu_0\beta^3 e^{-\beta|v_0|}}{16m} r^2 \right)}, \quad (11.217)$$

which has fixed points

$$r = 0 \text{ which corresponds to the stable fixed point} \quad (11.218)$$

$$r = \pm \sqrt{\frac{-8\mu m e^{\beta|v_0|}}{kg\mu_0\beta^3}}, \text{ which corresponds to the unstable limit cycle.} \quad (11.219)$$

Expressed as a function of v_0 , the bifurcation occurs at

$$v_C = \frac{1}{\beta} \log \left(\frac{mg\beta\mu_0}{c} \right). \quad (11.220)$$

(iii) For $\mu < 0$ the amplitude of the unstable limit cycle is

$$r_0 = \sqrt{\frac{-8\mu m e^{\beta|v_0|}}{kg\mu_0\beta^3}} \quad (11.221)$$

Solution (5.1). Consider a planar dynamical system with the following phase portrait:

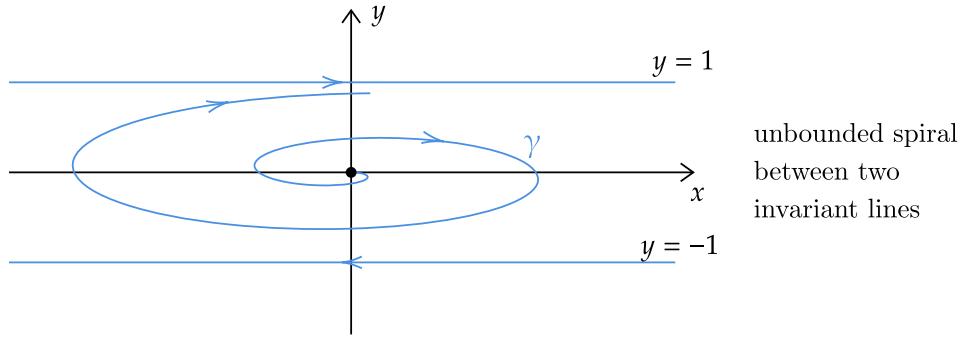


Figure 11.3: Phase portrait of the planar dynamical system

Which of the following statement is true?

- (i) The ω -limit set of γ is empty.
- (ii) By the Poincaré-Bendixson theorem, the ω -limit set of γ is composed of the lines $y = 1$ and $y = -1$.
- (iii) The Poincaré-Bendixson theorem does not apply to γ .
- (iv) None of the above

Solution (5.2).

$$\begin{bmatrix} \dot{x} \\ \dot{y} \end{bmatrix} = \underline{F}(x, y) := \begin{bmatrix} \frac{\partial H}{\partial y}(x, y) + f_1(x, y) \\ -\frac{\partial H}{\partial x}(x, y) + f_2(x, y) \end{bmatrix} \quad (11.222)$$

$$\text{div}(\underline{F}) = \frac{\partial^2 H}{\partial x \partial y}(x, y) + \frac{\partial f_1}{\partial x}(x, y) - \frac{\partial^2 H}{\partial y \partial x}(x, y) + \frac{\partial f_2}{\partial y} \quad (11.223)$$

$$= \text{div}(\underline{f}) \quad (H \in C^2) \quad (11.224)$$

$$\neq 0 \quad \forall (x, y) \in \mathbb{R}^2 \quad (11.225)$$

Thus, by the Bendixson's criterion, (11.222) does not have a periodic solution in \mathbb{R}^2 .

Solution (6.1). Remember from the lecture on averaging that $\dot{x} = \varepsilon f(x, t, \varepsilon)$ can be transformed to the differential equation

$$\dot{\tilde{x}} = \varepsilon \bar{f}_0(\tilde{x}) + \varepsilon^2 f_1(\tilde{x}, t, \varepsilon) \quad (11.226)$$

Through the near-identity transformation $x = \tilde{x} + \varepsilon w(\tilde{x}, t)$.

Moreover, f_1 is globally bounded, i.e., there exists $L_1 > 0$ such that

$$|f_1(\tilde{x}, t, \varepsilon)| < L_1 \quad \forall t > 0 \text{ and } \forall \tilde{x} \in \mathbb{R}^n \quad (11.227)$$

Now by construction $|x(t) - \tilde{x}(t)| = \varepsilon |w(\tilde{x}(t), t)| = \mathcal{O}(\varepsilon)$. Therefore, it suffices to show that solutions of the averaged equation

$$\dot{y} = \varepsilon \bar{f}_0(y) \quad (11.228)$$

remain $\mathcal{O}(\varepsilon)$ close to the solutions of (11.226).

Subtracting (11.228) from (11.226), integrating and dropping the tilde ($\tilde{\cdot}$) sign, we get

$$x(t) - y(t) = x_0 - y_0 + \varepsilon \int_0^t (\bar{f}_0(x(s)) - \bar{f}_0(y(s))) ds + \varepsilon^2 \int_0^t f_1(x(s), s, \varepsilon) ds \quad (11.229)$$

$$\implies |x(t) - y(t)| \leq |x_0 - y_0| + \varepsilon \int_0^t L_2 |x(s) - y(s)| ds + \varepsilon^2 \int_0^t L_1 ds \quad (11.230)$$

where we used boundedness of f_1 and Lipschitz continuity of \bar{f}_0 :

$$|f_0(x) - f_0(y)| \leq L_2 |x - y| \quad (11.231)$$

Therefore,

$$|x(t) - y(t)| \leq |x_0 - y_0| + \varepsilon^2 L_1 t + \int_0^t \varepsilon L_2 |x(s) - y(s)| ds \quad (11.232)$$

Now apply Gronwall's inequality with $v(t) = |x(t) - y(t)|$, $u(t) = \varepsilon L_2$ and $c(t) = |x_0 - y_0| + \varepsilon^2 L_1 t$ to get

$$|x(t) - y(t)| \leq |x_0 - y_0| e^{\varepsilon L_2 t} + \int_0^t \varepsilon L_1 e^{\varepsilon L_2 (t-s)} ds \quad (11.233)$$

$$= |x_0 - y_0| e^{\varepsilon L_2 t} + \varepsilon \frac{L_1}{L_2} (e^{\varepsilon L_2 t} - 1) \leq \left[|x_0 - y_0| + \varepsilon \frac{L_1}{L_2} \right] e^{\varepsilon L_2 t} \quad (11.234)$$

Since $|x_0 - y_0| = \mathcal{O}(\varepsilon)$, we conclude that $|x(t) - y(t)| = \mathcal{O}(\varepsilon)$ as long as $t \in \left[0, \frac{1}{\varepsilon L_2}\right)$, i.e., time scales of $\mathcal{O}(1/\varepsilon)$.

Solution (6.2). We start with the Taylor expansions of $u(x, y, t)$ and $v(x, y, t)$ in y near $y = 0$:

$$\begin{cases} u(x, y, t) = u(x, 0, t) + \frac{\partial u}{\partial y}(x, 0, t)y + \mathcal{O}(|y|^2) \\ v(x, y, t) = v(x, 0, t) + \frac{\partial v}{\partial y}(x, 0, t)y + \frac{1}{2} \frac{\partial^2 v}{\partial y^2}(x, 0, t)y^2 + \mathcal{O}(|y|^3) \end{cases} \quad (11.235)$$

But $u(x, 0, t) = v(x, 0, t) = 0$ for any x .

Differentiating $u(x, 0, t) = 0$ with respect to x we get $\frac{\partial u}{\partial x}(x, 0, t) = 0$. By incompressibility we have

$$\frac{\partial u}{\partial x} = -\frac{\partial v}{\partial y} \quad (11.236)$$

$$\Rightarrow \frac{\partial v}{\partial y}(x, 0, t) = -\frac{\partial u}{\partial x}(x, 0, t) = 0, \quad \forall x \quad (11.237)$$

Hence (11.235) simplifies to

$$\begin{cases} u(x, y, t) = \frac{\partial u}{\partial y}(x, 0, t)y + \mathcal{O}(|y|^2) = yU(x, y, t) \\ v(x, y, t) = \frac{1}{2} \frac{\partial^2 v}{\partial y^2}(x, 0, t)y^2 + \mathcal{O}(|y|^3) = y^2V(x, y, t) \end{cases} \quad (11.238)$$

Also note that

$$\begin{cases} U(x, y, t) = \frac{\partial u}{\partial y}(x, 0, t) \\ V(x, y, t) = \frac{1}{2} \frac{\partial^2 v}{\partial y^2}(x, 0, t) \end{cases} \quad (11.239)$$

Higher-order terms are identically zero. Therefore,

$$\begin{cases} \dot{x} = yU(x, y, t) \\ \dot{y} = y^2V(x, y, t) \end{cases} \quad (11.240)$$

Scaling y as $y = \varepsilon\eta$, we get:

$$\begin{cases} \dot{x} = \varepsilon\eta U(x, \varepsilon\eta, t) \\ \dot{\eta} = \varepsilon\eta^2 V(x, \varepsilon\eta, t) \end{cases} \quad (11.241)$$

Since U and V are also T -periodic, averaging theory applies to (11.241) with the averaged equations

$$\begin{cases} \dot{x} = \varepsilon\eta \bar{U}(x) \\ \dot{\eta} = \varepsilon\eta^2 \bar{V}(x) \end{cases} \quad (11.242)$$

where

$$\begin{cases} \bar{U}(x) = \frac{1}{T} \int_0^T U(x, 0, s) ds = \frac{1}{T} \int_0^T \frac{\partial u}{\partial y}(x, 0, s) ds \\ \bar{V}(x) = \frac{1}{T} \int_0^T V(x, 0, s) ds = \frac{1}{2T} \int_0^T \frac{\partial^2 v}{\partial y^2}(x, 0, s) ds \end{cases} \quad (11.243)$$

Rescaling time as $\frac{d\tau}{dt} = \eta(t)$ and denoting the derivative with respect to τ by prime sign ('), we get

$$\dot{x} = \frac{dx}{dt} = \frac{dx}{d\tau} \frac{d\tau}{dt} = \eta x' \quad , \quad \dot{\eta} = \eta \eta' \quad (11.244)$$

Substituting these expressions in (11.243), we get

$$\begin{cases} x' = \varepsilon \bar{U}(x) \\ \eta' = \varepsilon \eta \bar{V}(x) \end{cases} \quad (11.245)$$

Equation (11.245) has a fixed point $(x_0, \eta = 0)$ on the wall if and only if $\bar{U}(x_0, 0) = 0$. Using (11.243), we have

$$\bar{U}(x_0) = 0 \iff \boxed{\int_0^T \frac{\partial u}{\partial y}(x_0, 0, s) ds = 0} \quad (11.246)$$

Now we turn to the stability of the fixed point $(x_0, 0)$ on the wall by linearising (11.245) around this fixed point:

$$\underline{\xi}' = \varepsilon \underbrace{\begin{pmatrix} \frac{\partial \bar{U}}{\partial x}(x_0) & 0 \\ 0 & \bar{V}(x_0) \end{pmatrix}}_{:=A} \underline{\xi} \quad (11.247)$$

The matrix A has eigenvalues $\varepsilon \frac{\partial \bar{U}}{\partial x}(x_0)$ and $\varepsilon \bar{V}(x_0)$ corresponding to eigenvectors $\begin{pmatrix} 1 \\ 0 \end{pmatrix}$ and $\begin{pmatrix} 0 \\ 1 \end{pmatrix}$ respectively.

For the unstable manifold to be off the wall, we need $\bar{V}(x_0) > 0$. Using (11.243), we have

$$\bar{V}(x_0) > 0 \iff \boxed{\int_0^T \frac{\partial^2 v}{\partial y^2}(x_0, 0, s) ds > 0} \quad (11.248)$$

Solution (7.1). A periodic orbit is characterized by a string of k symbols. There is a total of 2^k possible permutations of the symbols 1 and 2 over k positions. However if this string consists of k/i repetitions of a string of length i , that periodic orbit is already an i -orbit, and so will have to be excluded from the set of k -orbits. The number of i -orbits is $N(i)$, and there will be i occurrences of k -strings consisting of repetitions of some shift of each i -orbit. Let $\langle i, k \rangle$ denote the set of $i \in \mathbb{N}$ such that k/i is an integer. Then

$$2^k - \sum_{i \in \langle i, k \rangle} iN(i) \quad (11.249)$$

is the number of permutations of two symbols that are not a repetition of some string. Furthermore, since shifting a string does not change its represented periodic orbit, there will be k shifted versions of each string, which we handle by dividing with k :

$$N(k) = \frac{1}{k} \left(2^k - \sum_{i \in \langle i, k \rangle} iN(i) \right) \quad (11.250)$$

Solution (7.2). Given any two periodic orbits $\bar{s}.\bar{s}$ and $\bar{z}.\bar{z}$, any sequence s^* defines a heteroclinic orbit between them by $\bar{s}s^*\bar{z}$. To show this is a heteroclinic orbit, observe that

$$\lim_{N \rightarrow \infty} d(\sigma^N(\bar{s}s^*\bar{z}), \bar{z}.\bar{z}) = 0 \quad (11.251)$$

$$\lim_{N \rightarrow -\infty} d(\sigma^N(\bar{s}s^*\bar{z}), \bar{s}.\bar{s}) = 0. \quad (11.252)$$

That is, any such trajectory approaches $\bar{z}.\bar{z}$ in forward iterations and $\bar{s}.\bar{s}$ in backward iterations of the shift map. Since there are infinitely many sequences s^* , and the periodic orbits were arbitrary, it follows that there are infinitely many heteroclinic orbits connecting any two periodic orbits.

Solution (7.3). Fix "points" $a \in A$ and $b \in B$. Since A and B are open sets, we can choose a neighborhood U around a and V around b such that all points closer to a than δ_U lie in U , and all points closer to b than δ_V lie in V . Pick $\delta = \min(\delta_U, \delta_V)$. The existence of a dense orbit for σ in Σ implies that there are integers $N(a)$, $N(b)$, and a sequence $s^* \in \Sigma$, such that $d(\sigma^{N(a)}(s^*), a) < \delta$ and $d(\sigma^{N(b)}(s^*), b) < \delta$. Since this sequence comes closer than δ to a and b , it passes through both U and V . Set $N = N(b) - N(a)$ and $z = \sigma^{N(a)}(s^*)$. Then $z \in A$ and $\sigma^N(z) = \sigma^{N(b)}(s^*) \in B$. The claim follows.

Solution (7.4). Given δ and $s \neq z$, by the definition of distance in Σ there must be some finite integer $N \in \mathbb{Z}$ such that the N th elements $s_N \neq z_N$. Now iterate both sequences N times and compute the distance:

$$\begin{aligned} d(\sigma^N(s), \sigma^N(z)) &= \sum_i \frac{|\sigma^N(s)_i - \sigma^N(z)_i|}{2^{|i|}} = \\ &= \sum_i \frac{|s_{i+N} - z_{i+N}|}{2^{|i|}} \geq \frac{|s_N - z_N|}{2^{|0|}} = 1. \end{aligned} \tag{11.253}$$

Thus the claim follows from setting e.g. $\Delta = \frac{1}{2}$.

Solution (7.5). (i) The space of admissible sequences is

$$\Sigma_A^N = \{s \in \Sigma^N : A_{s_i, s_{i+1}} \neq 0 \quad \forall i \in \mathbb{Z}\}. \tag{11.254}$$

For a fixed point $\bar{a}.\bar{a}$, $a \in \{1, \dots, N\}$, we must therefore have $A_{a,a} \neq 0$. So the a th element on the diagonal of A must be non-zero for $\bar{a}.\bar{a}$ to be in the space of admissible sequences. Furthermore since Σ^N contains all permutations and therefore all possible fixed points, the condition $A_{a,a} \neq 0$ is both sufficient and necessary for $\bar{a}.\bar{a}$ to be an admissible fixed point. Therefore the number of fixed points is equal to the number of non-zero diagonal elements of A , which, since $A_{i,j} \in \{0, 1\}$, is equal to $\text{tr}(A)$.

(ii) Apply the reasoning from a) to the map σ^k , i.e. the shift map repeated k times. The matrix of admissible sequences is then A repeated k times, i.e. A^k . Again if $A_{a,a}^k = 0$, there is no admissible sequence going from a back to a in k steps. Note that e.g. $A_{i,j}^3 = \sum_{k,l} A_{i,k} A_{k,l} A_{l,j}$. Each of the terms represents an admissible sequence in 3 steps from i to j . The sum is the total number of such admissible sequences. As in a), all possible sequences are in Σ^N , so $A_{a,a}^k$ is the number of ways to map a into itself in k steps. Again the total number of such admissible k -periodic points is the sum of $A_{i,i}^k$, i.e. $\text{tr}(A^k)$.

Solution (8.1). Fix the origin at the suspension point. Then, the Cartesian coordinates of the two point masses are

$$x_1 = L \sin \alpha; \quad x_2 = L \sin \beta + x_1 = L(\sin \alpha + \sin \beta) \tag{11.255}$$

$$y_1 = -L \cos \alpha; \quad y_2 = -L(\cos \alpha + \cos \beta). \tag{11.256}$$

The Lagrangian for the double pendulum in Cartesian coordinates is

$$L(x, y, \dot{x}, \dot{y}) = \frac{1}{2}m(\dot{x}_1^2 + \dot{x}_2^2 + \dot{y}_1^2 + \dot{y}_2^2) - mg(y_1 + y_2). \tag{11.257}$$

Calculating the velocities, in terms of the generalized coordinates α, β , we get

$$\dot{x}_1^2 + \dot{y}_1^2 = L^2 \dot{\alpha}^2 \tag{11.258}$$

$$\dot{x}_2 = L \cos \alpha \dot{\alpha} + L \cos \beta \dot{\beta} \quad (11.259)$$

$$\dot{y}_2 = L \sin \alpha \dot{\alpha} + L \sin \beta \dot{\beta} \quad (11.260)$$

$$\dot{x}_2^2 + \dot{y}_2^2 = L^2[\dot{\alpha}^2 + \dot{\beta}^2 + 2\dot{\alpha}\dot{\beta}(\cos \alpha \cos \beta + \sin \alpha \sin \beta)] = L^2[\dot{\alpha}^2 + \dot{\beta}^2 + 2\dot{\alpha}\dot{\beta} \cos(\alpha - \beta)]. \quad (11.261)$$

The Lagrangian is

$$L(\alpha, \beta, \dot{\alpha}, \dot{\beta}) = \frac{mL^2}{2}(2\dot{\alpha}^2 + \dot{\beta}^2 + 2\dot{\alpha}\dot{\beta} \cos(\alpha - \beta)) + mgL(2 \cos \alpha + \cos \beta). \quad (11.262)$$

The generalized momenta are

$$p_\alpha = \frac{\partial L}{\partial \dot{\alpha}} = mL^2(2\dot{\alpha} + \dot{\beta} \cos(\alpha - \beta)) \quad (11.263)$$

$$p_\beta = \frac{\partial L}{\partial \dot{\beta}} = mL^2(\dot{\beta} + \dot{\alpha} \cos(\alpha - \beta)). \quad (11.264)$$

To invert this relation, express $\dot{\alpha}$ and $\dot{\beta}$ with p_α and p_β . From the second equation, we get

$$\dot{\beta} = \frac{p_\beta}{mL^2} - \dot{\alpha} \cos(\alpha - \beta). \quad (11.265)$$

Substituting this into the first,

$$\dot{\alpha} = \frac{p_\alpha}{2mL^2} - \frac{\dot{\beta} \cos(\alpha - \beta)}{2} = \frac{p_\alpha}{2mL^2} - \frac{1}{2} \cos(\alpha - \beta) \left(\frac{p_\beta}{mL^2} - \dot{\alpha} \cos(\alpha - \beta) \right) \quad (11.266)$$

$$\dot{\alpha} \left(1 - \frac{1}{2} \cos^2(\alpha - \beta) \right) = \frac{p_\alpha - p_\beta \cos(\alpha - \beta)}{2mL^2} \quad (11.267)$$

$$\dot{\alpha} = \frac{p_\alpha - p_\beta \cos(\alpha - \beta)}{mL^2(2 - \cos^2(\alpha - \beta))}. \quad (11.268)$$

Then, using $\dot{\alpha}$ in the first equation,

$$\dot{\beta} = \frac{p_\beta}{mL^2} - \frac{p_\alpha - p_\beta \cos(\alpha - \beta)}{mL^2(2 - \cos^2(\alpha - \beta))} \cos(\alpha - \beta) = \frac{2p_\beta - p_\alpha \cos(\alpha - \beta)}{mL^2(1 + \sin^2(\alpha - \beta))}. \quad (11.269)$$

The Hamiltonian is obtained by a Legendre transform

$$H(\alpha, p_\alpha, \beta, p_\beta) = \dot{\alpha}p_\alpha + \dot{\beta}p_\beta - L = \frac{p_\alpha^2 + 2p_\beta^2 - 2p_\beta p_\alpha \cos(\alpha - \beta)}{mL^2(1 + \sin^2(\alpha - \beta))} - \quad (11.270)$$

$$-\frac{mL^2}{2}(2\dot{\alpha}^2 + \dot{\beta}^2 + 2\dot{\alpha}\dot{\beta} \cos(\alpha - \beta)) - mgL(2 \cos \alpha + \cos \beta) \quad (11.271)$$

The second term is

$$\frac{mL^2}{2}(2\dot{\alpha}^2 + \dot{\beta}^2 + 2\dot{\alpha}\dot{\beta} \cos(\alpha - \beta)) = \frac{mL^2}{2}[2\left(\frac{p_\alpha - p_\beta \cos(\alpha - \beta)}{mL^2(1 + \sin^2(\alpha - \beta))}\right)^2 + \quad (11.272)$$

$$+ \left(\frac{2p_\beta - p_\alpha \cos(\alpha - \beta)}{mL^2(1 + \sin^2(\alpha - \beta))}\right)^2 + 2\frac{2p_\beta - p_\alpha \cos(\alpha - \beta)}{mL^2(1 + \sin^2(\alpha - \beta))}\frac{p_\alpha - p_\beta \cos(\alpha - \beta)}{mL^2(1 + \sin^2(\alpha - \beta))} \cos(\alpha - \beta)] = \quad (11.273)$$

$$= \frac{2p_\alpha^2 + 2p_\beta^2 \cos^2(\alpha - \beta) - 4p_\alpha p_\beta \cos(\alpha - \beta) + 4p_\beta^2 + p_\alpha^2 \cos^2(\alpha - \beta) - 4p_\alpha p_\beta \cos(\alpha - \beta)}{2mL^2(1 + \sin^2(\alpha - \beta))^2} + \quad (11.274)$$

$$+ \frac{4p_\beta p_\alpha \cos(\alpha - \beta) - 4p_\beta^2 \cos^2(\alpha - \beta) - 2p_\alpha^2 \cos^2(\alpha - \beta) + 2p_\beta p_\alpha \cos^3(\alpha - \beta)}{2mL^2(1 + \sin^2(\alpha - \beta))^2} = \quad (11.275)$$

$$= \frac{p_\alpha^2(2 - \cos^2(\alpha - \beta)) + 2p_\beta^2(2 - \cos^2(\alpha - \beta)) - 2p_\beta p_\alpha \cos(\alpha - \beta)(2 - \cos^2(\alpha - \beta))}{2mL^2(1 + \sin^2(\alpha - \beta))^2} = \quad (11.276)$$

$$= \frac{p_\alpha^2 + 2p_\beta^2 - 2p_\beta p_\alpha \cos(\alpha - \beta)}{2mL^2(1 + \sin^2(\alpha - \beta))}. \quad (11.277)$$

Substituting it into the Hamiltonian,

$$H(\alpha, p_\alpha, \beta, p_\beta) = \frac{p_\alpha^2 + 2p_\beta^2 - 2p_\beta p_\alpha \cos(\alpha - \beta)}{2mL^2(1 + \sin^2(\alpha - \beta))} - mgL(2 \cos \alpha + \cos \beta). \quad (11.278)$$

Hamilton's Equations are

$$\dot{\alpha} = \frac{p_\alpha - p_\beta \cos(\alpha - \beta)}{mL^2(2 - \cos^2(\alpha - \beta))} \quad (11.279)$$

$$\dot{\beta} = \frac{2p_\beta - p_\alpha \cos(\alpha - \beta)}{mL^2(1 + \sin^2(\alpha - \beta))} \quad (11.280)$$

$$\dot{p}_\alpha = -\frac{\partial H}{\partial \alpha} \quad (11.281)$$

$$\dot{p}_\beta = -\frac{\partial H}{\partial \beta} \quad (11.282)$$

For the momentum equations, let $\lambda(\alpha, \beta) = 2mL^2(1 + \sin^2(\alpha - \beta))$.

$$\dot{p}_\alpha = -\frac{\partial H}{\partial \alpha} = -\frac{1}{\lambda^2} \left(2\lambda p_\alpha p_\beta \sin(\alpha - \beta) - \frac{\partial \lambda}{\partial \alpha} (p_\alpha^2 + 2p_\beta^2 - 2p_\alpha p_\beta \cos(\alpha - \beta)) \right) - 2mgL \sin \alpha \quad (11.283)$$

$$= \frac{-2p_\alpha p_\beta \sin(\alpha - \beta)}{\lambda} + \frac{2mL^2(2 \sin(\alpha - \beta) \cos(\alpha - \beta))(p_\alpha^2 + 2p_\beta^2 - 2p_\alpha p_\beta \cos(\alpha - \beta))}{\lambda^2} - 2mgL \sin \alpha \quad (11.284)$$

$$\dot{p}_\alpha = -2mgL \sin \alpha - \frac{p_\alpha p_\beta \sin(\alpha - \beta)}{mL^2(1 + \sin^2(\alpha - \beta))} + \frac{\sin(2(\alpha - \beta))(p_\alpha^2 + 2p_\beta^2 - 2p_\alpha p_\beta \cos(\alpha - \beta))}{2mL^2(1 + \sin^2(\alpha - \beta))} \quad (11.285)$$

$$\dot{p}_\beta = -mgL \sin \beta + \frac{p_\alpha p_\beta \sin(\alpha - \beta)}{mL^2(1 + \sin^2(\alpha - \beta))} - \frac{\sin(2(\alpha - \beta))(p_\alpha^2 + 2p_\beta^2 - 2p_\alpha p_\beta \cos(\alpha - \beta))}{2mL^2(1 + \sin^2(\alpha - \beta))} \quad (11.286)$$

Solution (8.2). For a steady flow, $\frac{\partial \rho}{\partial t} = 0$, which means $\rho \mathbf{v}$ is divergence free, by the continuity equation. This condition means

$$\frac{\partial(\rho u)}{\partial x} = -\frac{\partial(\rho v)}{\partial y}. \quad (11.287)$$

If we extend $\rho \mathbf{v}$ to be a 3 dimensional vector, the above relation means that there is a vector-potential $\mathbf{A} : \mathbb{R}^3 \rightarrow \mathbb{R}^3$. In particular, with a scalar function $\Psi : \mathbb{R}^2 \rightarrow \mathbb{R}$,

$$\begin{bmatrix} \rho u \\ \rho v \\ 0 \end{bmatrix} = \text{curl } \mathbf{A} = \text{curl} \begin{bmatrix} 0 \\ 0 \\ \Psi \end{bmatrix} = \begin{bmatrix} \partial_y \Psi \\ -\partial_x \Psi \\ 0 \end{bmatrix}. \quad (11.288)$$

The (massless) fluid particles' motion obeys the differential equation

$$\dot{x} = u \quad (11.289)$$

$$\dot{y} = v. \quad (11.290)$$

Multiplying the equations by the density, and substituting Ψ , gives the desired form

$$\rho \dot{x} = \rho u \quad (11.291)$$

$$\rho \dot{y} = \rho v \quad (11.292)$$

$$\dot{x} = \frac{1}{\rho} \frac{\partial \Psi}{\partial y} \quad (11.293)$$

$$\dot{y} = -\frac{1}{\rho} \frac{\partial \Psi}{\partial x}. \quad (11.294)$$

We can bring it to the canonical form, by introducing a rescaling of time, $\tau = \int_0^t \frac{1}{\rho(x(s), y(s))} ds$.

$$\frac{dx}{d\tau} = \frac{\partial \Psi}{\partial y} \quad (11.295)$$

$$\frac{dy}{d\tau} = -\frac{\partial \Psi}{\partial x} \quad (11.296)$$

Solution (8.3). (i) The orbits of the linear equation are described in the space space, spanned by the variables $x, \dot{x} = v_x, y, \dot{y} = v_y$.

The differential equations are satisfied by

$$x(t) = r_x \cos(t + \delta_x) \quad y(t) = r_y \cos(t + \delta_y). \quad (11.297)$$

This can be verified by direct substitution,

$$\ddot{x} = -r_x \cos(t + \delta_x) = -x(t) \quad \ddot{y} = -r_y \cos(t + \delta_y) = -y(t). \quad (11.298)$$

The velocity variables are

$$v_x = \dot{x} = -r_x \sin(t + \delta_x) \quad v_y = \dot{y} = -r_y \sin(t + \delta_y). \quad (11.299)$$

The trajectory is given by the parametrization, using the notation $\phi_1 = t + \delta_x$, $\phi_2 = t + \delta_y$

$$\begin{bmatrix} x \\ v_x \\ y \\ v_y \end{bmatrix} = \begin{bmatrix} r_x \cos(\phi_1) \\ -r_x \sin(\phi_1) \\ r_y \cos(\phi_2) \\ -r_y \sin(\phi_2) \end{bmatrix} \quad (11.300)$$

Which describes a 2-torus, embedded in \mathbb{R}^4 .

(ii) Assume the converse, that the system is Hamiltonian, in the generalized sense. That is, there is a smooth function $H : \mathbb{T}^2 \rightarrow \mathbb{R}$ and $F : \mathbb{T}^2 \rightarrow \mathbb{R}$, $F(\phi_1, \phi_2) \neq 0$

$$\dot{\phi}_1 = F(\phi_1, \phi_2) \frac{\partial H}{\partial \phi_2} \quad (11.301)$$

$$\dot{\phi}_2 = -F(\phi_1, \phi_2) \frac{\partial H}{\partial \phi_1}. \quad (11.302)$$

Because F cannot be zero, all possible fixed points must correspond to critical points of H . We also know that H is a continuous function, defined on a compact domain (\mathbb{T}^2 is

compact). Then, by a theorem from analysis, H must have a minimum and a maximum value on \mathbb{T}^2 .

Since the 2-torus is a manifold without a boundary, these extremum points must correspond to critical points of H , where $DH = 0$.

We conclude that if the original system, $[\dot{\phi}_1, \dot{\phi}_2]$, is Hamiltonian, then it must have a fixed point.

If we assume that $[\dot{\phi}_1, \dot{\phi}_2]$ does not have a fixed point, then it cannot be Hamiltonian.

Solution (8.4). Consider the function $H : \mathbb{R}^n \times \mathbb{R}^n \times \mathbb{R} \rightarrow \mathbb{R}$, defined by

$$H(x, p, t) = \mathbf{f}(x, t) \cdot \mathbf{p}. \quad (11.303)$$

This function defines a Hamiltonian dynamical system on $(x, p) \in \mathbb{R}^n \times \mathbb{R}^n$. The evolution equations are

$$\dot{x} = f(x, t) \quad (11.304)$$

$$\dot{p} = -\nabla_x H(x, p, t) = -\nabla_x \mathbf{f}(x, t) \cdot \mathbf{p} \quad (11.305)$$

Projecting this system to any $\mathbf{p} = \text{constant}$ subspace gives the original dynamics, defined by the ODEs

$$\dot{q} = f(q, t). \quad (11.306)$$

Solution (8.5). Given a Hamiltonian system $\dot{x} = JDH(x)$, for $x \in \mathbb{R}^{2n}$ its linearization is

$$\dot{y} = (JD^2H)y, \quad (11.307)$$

with the matrix

$$J = \begin{bmatrix} 0 & \mathbb{I}_{n \times n} \\ -\mathbb{I}_{n \times n} & 0 \end{bmatrix}. \quad (11.308)$$

From the hint, BA and $B^{-1}(BA)B$ have the same eigenvalues. This is a similarity transformation. I will call $BA = G$, and the similar matrix $B^{-1}GB = F$. The inverse transform is $G = BFB^{-1}$.

Assume λ is an eigenvalue of G with eigenvector v . Then, for F we have that

$$BFB^{-1}v = \lambda v \quad (11.309)$$

and

$$F(B^{-1}v) = \lambda(B^{-1}v), \quad (11.310)$$

Showing that λ is an eigenvalue of F with eigenvector $B^{-1}v$.

Also note, that for the linearized Hamiltonian, the matrix D^2HJ has the property (written in index notation)

$$(D^2H)_{ij}J_{jk} = -J_{kj}(D^2H)_{ji}, \quad (11.311)$$

because of $J_{ij} = -J_{ji}$ and $D^2H_{ij} = D^2H_{ji}$.

$$D^2HJ = -(JD^2H)^T \quad (11.312)$$

Substituting $B = J$ and $A = D^2H$ into (11.310) gives

$$J^{-1}J(D^2H)J(J^{-1}v) = \lambda J^{-1}v \quad (11.313)$$

$$(D^2H)J(J^{-1}v) = \lambda J^{-1}v. \quad (11.314)$$

Using (11.312) we get

$$(JD^2H)^T(J^{-1}v) = -\lambda J^{-1}v. \quad (11.315)$$

This means that $-\lambda$ is an eigenvalue of $(JD^2H)^T$ (with eigenvector $J^{-1}v$), and since the eigenvalues of a matrix and its transpose coincide, it proves the statement.

Solution (8.6). (i) The linearized system for the gradient system is

$$\dot{y} = -D^2V y. \quad (11.316)$$

V is assumed to be smooth, so the Hessian D^2V is a symmetric matrix, because of the symmetry of second partial derivatives.

Since symmetric matrices have real eigenvalues, the linearized gradient system too, has real eigenvalues.

(ii) The fixed points satisfy $DV(x_0) = 0$. That is, x_0 is a fixed point if and only if it is a critical point of V .

Claim: If x_0 is a local minimum of V , then x_0 is asymptotically stable.

To show this, we can use V as a Lyapunov function:

If $V(x_0)$ is a minimum, that means there is an open subset U , such that $\forall x \in U, V(x) > V(x_0)$.

Then, define $L(x) := V(x) - V(x_0)$. This function has the property that $L(x) > 0 \forall x \in U \setminus \{x_0\}$. Its time derivative is

$$\frac{dL}{dt} = DV(x) \cdot \dot{x} = -DV(x) \cdot DV(x) = -\|DV(x)\|^2 < 0. \quad (11.317)$$

The norm of the gradient is always positive if $x \neq x_0$. This shows that $L(x)$ is positive definite and \dot{L} is negative definite on $U \setminus \{x_0\}$, which, by Lyapunov's theorem, guarantees asymptotic stability for x_0 .

(iii) The statement is, if x_0 is not a fixed point, there is no $T > 0$ such that $x(T) = x_0$. Assume the converse:

$\exists T > 0$ such that $x(T) = x(0) = x_0$. Consider the net change in V along this trajectory, using Newton-Leibniz:

$$V(x(T)) - V(x(0)) = \int_0^T \frac{dV(t)}{dt} dt. \quad (11.318)$$

From the previous exercise,

$$\int_0^T \frac{dV(t)}{dt} dt = - \int_0^T \|DV(x(t))\|^2 dt < 0. \quad (11.319)$$

This is strictly negative, if x_0 is not a fixed point. By assumption, the left hand side of (11.318) is 0, which is in contradiction with (11.319). This shows that there is no periodic orbit in a gradient system.

(iv) Based on this result, we can numerically find the minima of a smooth function $f(x)$ by considering the gradient system

$$\dot{x} = -Df(x). \quad (11.320)$$

If \hat{x} is a local minimum of $f(x)$, then there is an open subset U , such that $\forall x \in U, f(x) > f(\hat{x})$. By the result in (b), \hat{x} is an asymptotically stable fixed point of (11.320).

To numerically find \hat{x} , take an initial guess x_0 in U . Then, we solve the initial value problem $x(0) = x_0$, with $x(t)$ satisfying (11.320). The solution will converge to x_0 . To be specific, simply using the Euler-scheme

$$x_{n+1} = x_n - Df(x_n), \quad (11.321)$$

is the gradient descent method. For finding maxima, put $F(x) := -f(x)$, and repeat the process for the gradient system $\dot{x} = -DF(x)$.

Solution (8.7). (i) The system can be written as

$$\dot{h} = hp \left(\frac{a_1}{p} - a_1 b \right) \quad (11.322)$$

$$\dot{p} = hp \left(a_2 c - \frac{a_2}{h} \right). \quad (11.323)$$

We can rescale time by $A(h, p) = hp$, which is a positive function for $h, p > 0$, introducing a new time variable

$$\tau = \int_0^t h(s)p(s)ds. \quad (11.324)$$

$$\begin{bmatrix} \frac{dh}{d\tau} \\ \frac{dp}{d\tau} \end{bmatrix} = \begin{bmatrix} \frac{a_1}{p} - a_1 b \\ a_2 c - \frac{a_2}{h} \end{bmatrix} := \begin{bmatrix} C(p) \\ D(h) \end{bmatrix} \quad (11.325)$$

For this system to be Hamiltonian, C and D must be the appropriate partial derivatives of a function $H : \mathbb{R}^+ \times \mathbb{R}^+ \rightarrow \mathbb{R}$. That is,

$$\frac{\partial H}{\partial p} = C(p) \quad (11.326)$$

and

$$\frac{\partial H}{\partial h} = -D(h). \quad (11.327)$$

We can integrate the equations to get

$$\frac{a_1}{p} - a_1 b = \frac{\partial H}{\partial p} \quad (11.328)$$

$$H(h, p) = a_1 \log p - a_1 bp + F(h) \quad (11.329)$$

$$F'(h) = \frac{\partial H}{\partial h} = -a_2 c + \frac{a_2}{h} \quad (11.330)$$

$$F(h) = -a_2 ch + a_2 \log h + K \quad (11.331)$$

$$H(h, p) = a_1 \log p - a_1 bp + a_2 \log h - a_2 ch + K, \quad (11.332)$$

where K is a constant. Taking H as the Hamiltonian, the system can be written as (in the rescaled time)

$$\frac{dh}{d\tau} = \frac{\partial H}{\partial p} \quad (11.333)$$

$$\frac{dp}{d\tau} = -\frac{\partial H}{\partial h}. \quad (11.334)$$

- (ii) In the region $h, p > 0$, (where the rescaling is valid), the system has a single fixed point, defined by

$$C(h) = D(h) = 0. \quad (11.335)$$

This is satisfied by $\frac{a_1}{p_0} = a_1 b$ and $\frac{a_2}{h_0} = a_2 c$, or

$$(h_0, p_0) = \left(\frac{1}{b}, \frac{1}{c} \right). \quad (11.336)$$

To establish stability of (h_0, p_0) , take H as a Lyapunov function. (h_0, p_0) is a critical point of H , which is conserved along trajectories, $\dot{H} = 0$.

The Hessian matrix of H is

$$D^2 H = \begin{bmatrix} \frac{\partial^2 H}{\partial h^2} & \frac{\partial^2 H}{\partial h \partial p} \\ \frac{\partial^2 H}{\partial h \partial p} & \frac{\partial^2 H}{\partial p^2} \end{bmatrix} = \begin{bmatrix} -\frac{a_2}{h^2} & 0 \\ 0 & -\frac{a_1}{p^2} \end{bmatrix} \quad (11.337)$$

$$D^2 H_{(h_0, p_0)} = \begin{bmatrix} -a_2 b^2 & 0 \\ 0 & -a_1 c^2 \end{bmatrix}. \quad (11.338)$$

This is negative definite at (h_0, p_0) , so by taking $V = -H$ as a Lyapunov function, we can conclude nonlinear stability.

Solution (9.1). $X \subset \mathbb{R}^n$ is a manifold, which means $\forall x \in X$ there is an open subset $V \subset \mathbb{R}^k$ around $\xi \in \mathbb{R}^k$ and a local parametrization $\varphi : V \rightarrow \mathbb{R}^n$, $\varphi(\xi) = x$. I also assume that $Y \subset \mathbb{R}^l$ for some l . The product $X \times Y$ is a subset of the ambient space $\mathbb{R}^n \times \mathbb{R}^l$.

Let us call the graph of f , M . The goal is to show that for each $(x, y) \in M$, there is an open subset U of \mathbb{R}^k , with $\xi' \in U$ and a local parametrization $F : U \rightarrow \mathbb{R}^n \times \mathbb{R}^l$ with $F(\xi') = (x, y) \in M$, which is a diffeomorphism from U to $M \cap W$, for an open set $W \subset \mathbb{R}^n \times \mathbb{R}^l$ [$W \cap M$ is relatively open].

Consider the map

$$F : V \subset \mathbb{R}^k \rightarrow \mathbb{R}^n \times \mathbb{R}^l \quad (11.339)$$

given by

$$F(\xi) = (\varphi(\xi), f \circ \varphi(\xi)). \quad (11.340)$$

This map is smooth, because its components are smooth. It is also a bijection on $M \cap W$:

- It is injective, since if $F(\xi_1) = F(\xi_2)$, we must have $\varphi(\xi_1) = \varphi(\xi_2)$ because of the first component. φ is a bijection, so $\xi_1 = \xi_2$.
- It is surjective, because if $(x, y) \in M \cap W$, then we have $f(x) = y$ and $x = \varphi(\xi)$ for some $\xi \in V$, since φ is a local parametrization for X . This means we can write $(x, y) = F(\xi) = (\varphi(\xi), f \circ \varphi(\xi))$.

Its inverse is given by

$$F^{-1} : M \cap W \rightarrow V; \quad \varphi^{-1} \circ \pi, \quad (11.341)$$

where $\pi : X \times Y \rightarrow X$, $\pi(x, y) = x$ is the projection onto the first factor, so $F^{-1}(x, y) = \varphi^{-1}(x)$. This function is smooth, because it has a smooth extension to $W \subset \mathbb{R}^n \times \mathbb{R}^l$: π is already smooth on W , and because φ was a local parametrization, φ^{-1} has to have a smooth extension.

This shows that around a point $(x, y) \in M$, the map F is a local parametrization from an open subset $U = V \subset \mathbb{R}^k$, with $x \in \varphi(V)$, for some local parametrization of X around x .

Solution (9.2). Given an m -dimensional manifold $M \subset \mathbb{R}^n$, the tangent space at $p \in M$ is defined by

$$df_x(\mathbb{R}^m), \quad (11.342)$$

for some local parametrization around p , $f : U \subset \mathbb{R}^m \rightarrow \mathbb{R}^n$ and $f(x) = p$. Assume that there is a different local parametrization, $g : V \subset \mathbb{R}^m \rightarrow \mathbb{R}^n$, $p \in f(U) \cap g(V)$, with $f(x) = p = g(y)$.

The differentials are the linear maps

$$df_x : \mathbb{R}^m \rightarrow \mathbb{R}^n \quad dg_y : \mathbb{R}^m \rightarrow \mathbb{R}^n. \quad (11.343)$$

Since both f and g are local parametrizations, df_x and dg_y are injective and they (as linear maps) have $\text{Rank } f = \text{Rank } g = m$.

$\text{Rank } f = \dim(\text{Image } f) = \dim(\text{Image } g) = m$. This shows that the tangent spaces we obtain using f and g are both m dimensional subspaces of \mathbb{R}^n . That is, the vector spaces $df_x(\mathbb{R}^m)$ and $dg_y(\mathbb{R}^m)$ are isomorphic, because they are finite dimensional, with the same dimension.

Solution (9.3). Let $M \subset \mathbb{R}^n$ be an m dimensional manifold. Its tangent bundle is

$$TM = \bigcup_{p \in M} \{p\} \times TM_p. \quad (11.344)$$

We show, that TM is a $2m$ dimensional manifold. Consider a point $(p, v) \in TM$, such that $v \in TM_p$. Since M is a manifold, there is a local parametrization $\varphi : \tilde{U} \subset \mathbb{R}^m \rightarrow \mathbb{R}^n$, with $x \in \tilde{U}, \varphi(x) = p$.

Let $U = \tilde{U} \times \mathbb{R}^m$ be an open set, and define a function $\Phi : U \rightarrow \mathbb{R}^n \times \mathbb{R}^n$, by

$$\Phi(x, \eta) = (\varphi(x), d\varphi_x(\eta)). \quad (11.345)$$

Since both φ and $d\varphi_x$ are smooth functions, Φ is also smooth. We still need to show that Φ is a diffeomorphism onto $TM \cap W$ for an open set $W \subset \mathbb{R}^n \times \mathbb{R}^n$.

Φ is a bijection, because:

If $\Phi(x, \eta) = \Phi(x', \eta')$, for $(x, \eta), (x', \eta') \in TM \cap W$ then $\varphi(x) = \varphi(x')$ from the first component, which means $x = x'$. This is because φ is a local parametrization. From the second component: $d\varphi_x(\eta) = d\varphi_{x'}(\eta')$ means $\eta = \eta'$, since the linear map $d\varphi_x$ is injective. This shows Φ is injective on the relatively open set $TM \cap W$.

For surjectivity, for a point $(p, v) \in TM \cap W$ we can find $x = \varphi^{-1}(p)$, since the local parametrization is already bijective. We also know that its inverse is smooth, so with $\eta := d(\varphi^{-1})_p(v)$, we see that

$$d\varphi_p(\eta) = d\varphi_p[d(\varphi^{-1})_p(v)] = d\varphi_{\varphi^{-1}(p)} \circ d(\varphi^{-1})_p(v) = d(\varphi^{-1} \circ \varphi)_p(v) = v, \quad (11.346)$$

by the chain rule. So, we have found a point (x, η) for each $(p, v) \in TM \cap W$, with $\Phi(x, \eta) = (p, v)$.

This also gives the inverse function on this relatively open set:

$$\Phi^{-1} : TM \cap W \rightarrow U \quad (11.347)$$

$$\Phi^{-1}(p, v) = [\varphi^{-1}(p), d(\varphi^{-1})_p(v)]. \quad (11.348)$$

We know that φ^{-1} can be extended to a smooth function, and the same is true for its differential. Then, we have that $\Phi : U \rightarrow \mathbb{R}^n \times \mathbb{R}^n$ is a diffeomorphism, and is a suitable local parametrization for TM , around a point $(p, v) = \Phi(x, \eta)$,

Solution (9.4). To define the Lyapunov-type numbers at a point p on an overflowing invariant manifold M , let $w_0 \in NM_p$ be a vector in the normal space of M at p and $v_0 \in TM_p$ be a vector in the tangent space at p . Their images under the linearized backward flow are

$$v_{-t} = DF_p^{-t}(v_0) \text{ and} \quad (11.349)$$

$$w_{-t} = \Pi_{F^{-t}(p)} \circ DF_p^{-t}(w_0). \quad (11.350)$$

Here, DF_p^{-t} is the differential of the (backward) flow-map at the point p (which is a linear map on the tangent space of the full phase-space at p) and Π_q is the orthogonal projection onto the normal space of M at q .

With this notation, the Lyapunov-type numbers, for a point p on an overflowing invariant manifold M , are defined to be

$$\nu(p) = \inf \left\{ a > 0 : \frac{\|w_0\|}{\|w_{-t}\| a^t} \rightarrow 0 \text{ as } t \rightarrow \infty, \forall w_0 \in NM_p \right\} \quad (11.351)$$

$$\sigma(p) = \inf \left\{ b \in \mathbb{R} : \frac{\|w_0\| / \|w_{-t}\|^b}{\|v_0\| / \|v_{-t}\|} \rightarrow 0 \text{ as } t \rightarrow \infty, \forall w_0 \in NM_p, v_0 \in TM_p \right\}. \quad (11.352)$$

To calculate $\nu(p)$ and $\sigma(p)$, we first define the two operators

$$A_t(p) := DF_p^{-t}|_{TM_p} \quad (11.353)$$

$$B_t(p) := \Pi_p \circ DF_{F^{-t}(p)}^t|_{NM_{F^{-t}(p)}}. \quad (11.354)$$

Then, a theorem by Fenichel states that

$$\nu(p) = \limsup_{t \rightarrow \infty} \|B_t\|^{1/t}. \quad (11.355)$$

If in addition, $\nu(p) < 1$ then

$$\sigma(p) = \limsup_{t \rightarrow \infty} \frac{\log \|A_t(p)\|}{-\log \|B_t(p)\|}. \quad (11.356)$$

To compute the operators $A_t(p)$ and $B_t(p)$, we calculate the flow-map of system $\dot{x} - \dot{y}$ explicitly. Equations for \dot{x} and \dot{y} are decoupled, so we can find the flow map by integrating the equations separately. The equation for y is a homogeneous, linear, scalar equation, which has the solution, assuming the initial condition is given as $y(t_0) := y_0$,

$$y(t; y_0, t_0) = y_0 e^{-b(t-t_0)}. \quad (11.357)$$

The system is autonomous, which means we can also set $t_0 = 0$ without loss of generality:

$$y(t; y_0) = y_0 e^{-bt}. \quad (11.358)$$

We can solve the equation in \dot{x} by separation of variables:

$$\dot{x} = -x(1 - x^2) \quad (11.359)$$

$$t = \int_{x_0}^x -\frac{dx'}{x'(1 - x'^2)} dx'. \quad (11.360)$$

The integral on the right side can be evaluated using the expression (partial fraction decomposition)

$$-\frac{1}{x(1 - x^2)} = -\frac{1}{x} - \frac{1}{2(1 - x)} + \frac{1}{2(1 + x)}. \quad (11.361)$$

Then, using Newton-Leibniz, we get

$$t = -(\log x - \log x_0) + \frac{1}{2} [\log(1 - x) - \log(1 - x_0)] + \frac{1}{2} [\log(1 + x) - \log(1 + x_0)] \quad (11.362)$$

$$e^t = \frac{x_0}{x} \sqrt{\frac{1 - x^2}{1 - x_0^2}} \quad (11.363)$$

This expression only makes sense if x and x_0 have the same sign. In this case, we can square it to get

$$\frac{x^2 e^{2t}}{x_0^2} = \frac{1 - x^2}{1 - x_0^2} \quad (11.364)$$

Rearranging to solve for x^2 , we have

$$x^2 = \frac{x_0^2}{x_0^2 + (1 - x_0^2)e^{2t}}. \quad (11.365)$$

In the relevant case, when x and x_0 have the same sign, we can take the (positive) square root

$$x(t; x_0) = \frac{x_0}{\sqrt{x_0^2 + (1 - x_0^2)e^{2t}}}. \quad (11.366)$$

The components of the flow-map are then given by the functions (11.358) and (11.366).

$$F^t(x_0, y_0) := \begin{bmatrix} x(t; x_0) \\ y(t; y_0) \end{bmatrix} = \begin{bmatrix} \frac{x_0}{\sqrt{x_0^2 + (1 - x_0^2)e^{2t}}} \\ y_0 e^{-bt} \end{bmatrix} \quad (11.367)$$

The linearized backward flow-map is the matrix

$$DF_{(x_0, y_0)}^{-t} := \begin{bmatrix} \frac{\partial x(-t; x_0)}{\partial x_0} & 0 \\ 0 & \frac{\partial y(-t; y_0)}{\partial y_0} \end{bmatrix}. \quad (11.368)$$

To compute the first entry, I use the short-hand $a := x_0^2 + (1 - x_0^2)e^{-2t}$.

$$\frac{\partial x(t; x_0)}{\partial x_0} = \frac{1}{a} \left[\sqrt{a} - \frac{x_0^2(1 - e^{-2t})}{\sqrt{a}} \right] = \frac{a - x_0^2(1 - e^{-2t})}{a^{3/2}} = \frac{e^{-2t}}{a^{3/2}} \quad (11.369)$$

$$DF_{x_0}^{-t} := \begin{bmatrix} \frac{e^{-2t}}{a^{3/2}} & 0 \\ 0 & e^{bt} \end{bmatrix} \quad (11.370)$$

Restricting this matrix onto the tangent space of M at $p = x_0$, means that it acts on vectors of the ambient space, that have the form

$$\mathbf{v}_0 = \begin{bmatrix} v_0 \\ 0 \end{bmatrix}, \quad (11.371)$$

as

$$DF_{x_0}^{-t}|_{TM_{x_0}} \mathbf{v}_0 = \begin{bmatrix} \frac{e^{-2t}}{a^{3/2}} v_0 \\ 0 \end{bmatrix} \in TM_{F^{-t}(x_0)} \quad (11.372)$$

Which shows that the operator $A_t(x_0)$ is simply the scalar multiplication by

$$A_t(x_0) = \frac{e^{-2t}}{(x_0^2 + (1 - x_0^2)e^{-2t})^{3/2}}. \quad (11.373)$$

Similarly, with the notation $x_{-t} = F^{-t}(x_0)$,

$$DF_{F^{-t}(x_0)}^t = \begin{bmatrix} \frac{e^{2t}}{(x_{-t}^2 + (1 - x_{-t}^2)e^{2t})^{3/2}} & 0 \\ 0 & e^{-bt} \end{bmatrix} \quad (11.374)$$

Restricting it to the normal space at x_{-t} (vectors parallel to the y axis), and then projecting to the normal space at x_0 (again, vectors parallel to the y axis) we have

$$B_t(x_0) = e^{-bt}. \quad (11.375)$$

Substituting the expressions (11.373) and (11.375) into the theorem (11.355) and (11.356) gives

$$\nu(x_0) = \lim \sup_{t \rightarrow \infty} \|B_t\|^{1/t} = \lim \sup_{t \rightarrow \infty} e^{-b} = e^{-b}. \quad (11.376)$$

$$\sigma(x_0) = \lim \sup_{t \rightarrow \infty} \frac{\log \|A_t(p)\|}{-\log \|B_t(p)\|} = \lim \sup_{t \rightarrow \infty} \frac{-2t - \frac{3}{2} \log(x_0^2 + (1 - x_0^2)e^{-2t})}{bt} = \quad (11.377)$$

$$= -\frac{2}{b} - \lim \sup_{t \rightarrow \infty} \frac{3}{2b} \frac{\log(x_0^2 + (1 - x_0^2)e^{-2t})}{t} \quad (11.378)$$

Since $\nu(x_0) = e^{-b} < 1$, we could use the second part of the theorem as well. In the last line, we have the $\lim \sup$. of product of two functions, one of which is $1/t$, which goes to zero as $t \rightarrow \infty$. The limit of the numerator is

$$\lim_{t \rightarrow \infty} \frac{3}{2b} \log(x_0^2 + (1 - x_0^2)e^{-2t}) = \frac{3}{b} \log x_0. \quad (11.379)$$

If this limit exists, then the whole limit is zero, since it is a product of two convergent functions.

However, for $x_0 = 0$, $\log x_0$ is divergent. Then, we use the original expression and set $x_0 = 0$:

$$\lim \sup_{t \rightarrow \infty} \frac{3}{2b} \frac{\log(x_0^2 + (1 - x_0^2)e^{-2t})}{t} = \lim \sup_{t \rightarrow \infty} \frac{3}{2b} \frac{\log e^{-2t}}{t} = -\frac{3}{b}. \quad (11.380)$$

Then, the Lyapunov-type numbers on the manifold M are

$$\nu(x_0) = e^{-b} \quad (11.381)$$

$$\sigma(x_0) = \begin{cases} \frac{1}{b}, & \text{for } x_0 = 0 \\ -\frac{2}{b} & \text{otherwise.} \end{cases} \quad (11.382)$$

Solution (9.5). (i) The set M_0 is an invariant manifold, since setting $\varepsilon = 0$ and restricting to $y = z = 0$, the dynamics becomes $\dot{x} = \dot{y} = \dot{z} = 0$, there is no dynamics on this set. It is also a 1-manifold with boundary, $\partial M_0 = \{-1, 1\}$.

The system can be seen as the fast-time formulation of a singularly perturbed problem. In the limit $\varepsilon = 0$, the fast subsystem is

$$\dot{y} = -y \quad (11.383)$$

$$\dot{z} = z, \quad (11.384)$$

and the critical manifold is M_0 . To establish normal hyperbolicity, we restrict the dynamics to invariant subspaces $x = \text{const.}$, where it is given by the fast subsystem. In this

case, the fast subsystem is linear for all x , and has nonzero eigenvalues, showing that the dynamics in the direction normal to M_0 is hyperbolic.

A direct application of Fenichel's definition of normal hyperbolicity is not applicable, since the manifold is not overflowing invariant. This can be fixed by a local modification of the vector field near the boundary points [Wiggins, 1994].

- (ii) Restricting the perturbed system onto the set $M_\varepsilon := \{y = z = 0, x \in [-1, 1]\}$, the dynamics is $\dot{x} = -\varepsilon x$. This means that M_ε is an invariant manifold of the perturbed system, which coincides with the unperturbed manifold M_0 .
- (iii) The local stable manifold is contained in the $x - y$ plane. Because of the normal hyperbolicity of M_ε , it admits an invariant foliation:

$$W_{loc}^s(M_\varepsilon) = \bigcup_{p \in M_\varepsilon} f^s(p). \quad (11.385)$$

In particular, this foliation has the property, that given a base point $p \in M_\varepsilon$, the trajectories starting in $f^s(p) \subset W_{loc}^s(M_\varepsilon)$ have the fastest rate of convergence towards the trajectory $\gamma(p) \subset M_\varepsilon$ through p . That is, given $q \in f^s(p)$ and $q' \in f^s(p')$

$$\frac{\|F^t(q) - F^t(p)\|}{\|F^t(q') - F^t(p)\|} \rightarrow 0, \text{ as } t \rightarrow \infty. \quad (11.386)$$

To use this property to find the nonlinear surface $f^s(p)$, we first calculate the flow-map of system (??). The y and z coordinates decouple from x , and they can be integrated first

$$y(t; y_0) = y_0 e^{-t} \quad z(t; z_0) = z_0 e^t. \quad (11.387)$$

Substituting this into the x equation, we get the inhomogeneous, linear scalar equation

$$\dot{x} = -\varepsilon x - \varepsilon y_0^2 e^{-2t}, \quad (11.388)$$

which we can solve by variation of constants. The solution of the homogeneous part is $x_h = ce^{-\varepsilon t}$. To obtain a particular solution for the inhomogeneous equation, look for a solution of the form $x(t) = c(t)e^{-\varepsilon t}$. This leads to

$$\dot{c} = -y_0^2 \varepsilon e^{(\varepsilon-2)t} \quad (11.389)$$

$$c = \frac{y_0^2 \varepsilon}{2 - \varepsilon} e^{(\varepsilon-2)t}. \quad (11.390)$$

The full solution is the sum of the general solution to the homogeneous part and the particular solution

$$x(t) = c_0 e^{-\varepsilon t} + \frac{y_0^2 \varepsilon}{2 - \varepsilon} e^{-2t}. \quad (11.391)$$

Eliminating the constant of integration c_0 , using the initial condition $x(0) = x_0$, we get

$$x(t; x_0) = \left(x_0 - \frac{y_0^2 \varepsilon}{2 - \varepsilon} \right) e^{-\varepsilon t} + \frac{y_0^2 \varepsilon}{2 - \varepsilon} e^{-2t}. \quad (11.392)$$

The flow-map of the system is

$$F^t(x_0, y_0, z_0) = \begin{bmatrix} \left(x_0 - \frac{y_0^2 \varepsilon}{2 - \varepsilon} \right) e^{-\varepsilon t} + \frac{y_0^2 \varepsilon}{2 - \varepsilon} e^{-2t} \\ y_0 e^{-t} \\ z_0 e^t \end{bmatrix}. \quad (11.393)$$

Consider the points $p = (x_p, 0, 0), p' = (x'_p, 0, 0) \in M_\varepsilon$ and $q = (x_q, y_q, 0) \in f^s(p) \subset W_{loc}^u(M_\varepsilon)$, $q' = (x'_q, y'_q, 0) \in f^s(p') \subset W_{loc}^u(M_\varepsilon)$. Then,

$$F^t(q) - F^t(p) = \begin{bmatrix} \left(x_q - \frac{y_q^2 \varepsilon}{2 - \varepsilon} \right) e^{-\varepsilon t} + \frac{y_q^2 \varepsilon}{2 - \varepsilon} e^{-2t} - x_p e^{-\varepsilon t} \\ y_q e^{-t} \\ 0 \end{bmatrix} = \begin{bmatrix} e^{-\varepsilon t} \left(x_q - \frac{y_q^2 \varepsilon}{2 - \varepsilon} - x_p \right) + \frac{y_q^2 \varepsilon}{2 - \varepsilon} e^{-2t} \\ y_q e^{-t} \\ 0 \end{bmatrix}. \quad (11.394)$$

Similarly, for the denominator of (11.386)

$$F^t(q') - F^t(p) = \begin{bmatrix} e^{-\varepsilon t} \left(x'_q - \frac{y'^2 q \varepsilon}{2 - \varepsilon} - x_p \right) + \frac{y'^2 q \varepsilon}{2 - \varepsilon} e^{-2t} \\ y'_q e^{-t} \\ 0 \end{bmatrix} \quad (11.395)$$

In general, both the denominator and the numerator of (11.386) go to 0 as $t \rightarrow \infty$ (because they are the sums of exponentially decaying terms), the limit of the ratio is undetermined. However, since $\varepsilon \ll 1$, the slowest decaying term in the first component of the flow-map is the one which is multiplied by $e^{-\varepsilon t}$. We can set its coefficient to 0, which guarantees that the numerator goes to 0 at a faster rate than the denominator. This way, the ratio will go to 0.

This establishes a relation between x_p , the base point and the points $(x_q, y_q, 0)$ on the stable fiber connected to it. This relation is

$$x_q = x_p + \frac{y_q^2 \varepsilon}{2 - \varepsilon}, \quad (11.396)$$

as claimed.

- (iv) The local unstable manifold of M_ε , is the local stable manifold under the backward flow. Reversing time in the original system, we obtain

$$\dot{x} = \varepsilon(x + y^2) \quad (11.397)$$

$$\dot{y} = y \quad (11.398)$$

$$\dot{z} = -z, \quad (11.399)$$

for which the local stable manifold is contained in the $x - z$ plane. To obtain the foliation, we use the property (11.386) for the backward flow map.

$$F^{-t}(x_0, y_0, z_0) = \begin{bmatrix} \left(x_0 - \frac{y_0^2 \varepsilon}{2 - \varepsilon}\right) e^{\varepsilon t} + \frac{y_0^2 \varepsilon}{2 - \varepsilon} e^{2t} \\ y_0 e^t \\ z_0 e^{-t} \end{bmatrix} \quad (11.400)$$

Taking points $p = (x_p, 0, 0)$, $p' = (x'_p, 0, 0) \in M_\varepsilon$ and $q = (x_q, 0, z_q) \in f^u(p) \subset W_{loc}^u(M_\varepsilon)$, $q' = (x'_q, 0, z'_q) \in f^u(p') \subset W_{loc}^u(M_\varepsilon)$, and evaluating the expressions $F^{-t}(p) - F^{-t}(q)$ and $F^{-t}(q') - F^{-t}(p)$ gives

$$F^{-t}(q) - F^{-t}(p) = \begin{bmatrix} e^{\varepsilon t} (x_q - x_p) \\ 0 \\ z_q e^{-t} \end{bmatrix}; \quad F^{-t}(q') - F^{-t}(p) = \begin{bmatrix} e^{\varepsilon t} (x'_q - x_p) \\ 0 \\ z'_q e^{-t} \end{bmatrix}. \quad (11.401)$$

Following a similar argument as we made with the stable fibers, the ratio (11.386) converges to zero, if the numerator goes to zero much faster than the denominator. In this case, the numerator and the denominator do not go to zero term-by-term, since the exponent of the first term is positive. The ratio is

$$\frac{\|F^{-t}(q) - F^{-t}(p)\|}{\|F^{-t}(q') - F^{-t}(p)\|} = \frac{e^{2\varepsilon t} (x_q - x_p)^2 + z_q^2 e^{-2t}}{e^{2\varepsilon t} (x'_q - x_p)^2 + z'_q^2 e^{-2t}}. \quad (11.402)$$

To ensure a faster convergence in the denominator, we can set $x_q = x_p$ and see that the ratio goes to zero

$$\frac{\|F^{-t}(q) - F^{-t}(p)\|}{\|F^{-t}(q') - F^{-t}(p)\|} = \frac{z_q^2 e^{-2t}}{e^{2\varepsilon t} (x'_q - x_p)^2 + z_q'^2 e^{-2t}} = \frac{z_q^2}{(x'_q - x_p)^2 e^{2t(\varepsilon+1)} + z_q'^2} \rightarrow 0. \quad (11.403)$$

The unstable fibers are the straight lines parallel to the z axis, given by

$$f^u(p) = \{(x, y, z) : x = x_p, y = 0\}. \quad (11.404)$$

(e) Verify explicitly the properties of the stable fibers in this example.

- $f^s(p)$ is a C^r smooth (given by a smooth graph), 1 dimensional submanifold of the 2 dimensional manifold $W_{loc}^s(M_\varepsilon)$. Its intersection with the manifold is $f^s(p) \cap M_\varepsilon = \{(x, y, z) : x = x_p, y = 0, z = 0\} = p$, the base point.
- The local section of the stable subbundle $N_p^s M_\varepsilon$ is the stable subspace of the normal space at p : $N_p^s M_\varepsilon$. In this example, the normal space is the plane $N_p M_\varepsilon = \{(x, y, z) : x = x_p\}$, the stable subspace is the line $N_p^s M_\varepsilon = \{(x, y, z) : x = x_p, z = 0\}$. The local tangent vector to $f^s(p)$ is $(0, 1, 0)$, which is in this subspace.
- The stable fibers form a positively invariant family. The stable fibers of the images of the basepoints are

$$f^s(F^t(p)) = f^s(F^t(x_p, 0, 0)) = \left\{ (x, y, z) : x = x_p e^{-\varepsilon t} + \frac{\varepsilon}{2-\varepsilon} y^2, z = 0 \right\}. \quad (11.405)$$

While the images of the stable fibers are:

$$F^t(f^s(p)) = F^t \left(\left\{ (x, y, z) : x = x_p + \frac{\varepsilon}{2-\varepsilon} y^2, z = 0 \right\} \right) = \quad (11.406)$$

$$= \left\{ F^t(\bar{x}, \bar{y}, \bar{z}) : \bar{x} = x_p + \frac{\varepsilon}{2-\varepsilon} \bar{y}^2, \bar{z} = 0 \right\} = \quad (11.407)$$

$$= \{(x, y, z) : x = x(t; \bar{x}), y = y(t; \bar{y}); z = z(t; \bar{z})\} = \quad (11.408)$$

$$= \left\{ (x, y, z) : x = \left(x_p + \frac{\varepsilon}{2-\varepsilon} \bar{y}^2 - \frac{\varepsilon}{2-\varepsilon} \bar{y}^2 \right) e^{-\varepsilon t} + \frac{\varepsilon}{2-\varepsilon} \bar{y}^2 e^{-2t}, y = \bar{y} e^{-t}; z = 0 \right\} = \quad (11.409)$$

$$= \left\{ (x, y, z) : x = x_p e^{-\varepsilon t} + \frac{\varepsilon}{2-\varepsilon} \bar{y}^2 e^{-2t}, y = \bar{y} e^{-t}; z = 0 \right\} \quad (11.410)$$

Since we can eliminate \bar{y} using the second coordinate $\bar{y} = y e^t$, this is the same set as

$$\left\{ (x, y, z) : x = x_p e^{-\varepsilon t} + \frac{\varepsilon}{2-\varepsilon} y^2; z = 0 \right\}, \quad (11.411)$$

which coincides with $f^s(F^t(p))$.

- There exist positive constants C_s and λ_s , such that for any $q \in f^s(p)$ we have $|F^t(q) - F^t(p)| < C_s e^{-\lambda_s t}$.

For a point $q = (x_q, y_q, 0) \in f^s(p)$, we can use the expression describing the fiber to get

$$F^t(q) - F^t(p) = \begin{bmatrix} \left(x_q - \frac{y_q^2 \varepsilon}{2-\varepsilon}\right) e^{-\varepsilon t} + \frac{y_q^2 \varepsilon}{2-\varepsilon} e^{-2t} - x_p e^{-\varepsilon t} \\ y_q e^{-t} \\ 0 \end{bmatrix} = \begin{bmatrix} \frac{y_q^2 \varepsilon}{2-\varepsilon} e^{-2t} \\ y_q e^{-t} \\ 0 \end{bmatrix}. \quad (11.412)$$

$$|F^t(q) - F^t(p)|^2 = \frac{y_q^2 \varepsilon^2}{(2-\varepsilon)^2} e^{-4t} + y_q^2 e^{-2t} = y_q^2 e^{-2t} \left(\frac{\varepsilon^2}{(2-\varepsilon)^2} e^{-2t} + 1 \right) \quad (11.413)$$

$$|F^t(q) - F^t(p)| = |y_q| e^{-t} \sqrt{\frac{\varepsilon^2}{(2-\varepsilon)^2} e^{-2t} + 1} \quad (11.414)$$

The square root is bounded by the constant $L = \sqrt{\frac{\varepsilon^2}{(2-\varepsilon)^2} + 1}$, so we can bound the distance between the trajectory through p and the trajectory through q as

$$|F^t(q) - F^t(p)| \leq |y_q| L e^{-t}. \quad (11.415)$$

The local stable manifold is compact set, that contains M_ε intersected with the $x-y$ plane, so we can choose a $C > 0$, such that $C > |y_q|$ for any $(x_q, y_q, z_q) \in W_{loc}^s(M_\varepsilon)$. Then, we can set $C_s = CL$ and $\lambda_s = 1$.

- Given a base point $p \in M_\varepsilon$, the trajectories starting in $f^s(p) \subset W_{loc}^s(M_\varepsilon)$ have the fastest rate of convergence towards the trajectory $\gamma(p) \subset M_\varepsilon$ through p . That is, given $q \in f^s(p)$ and $q' \in f^s(p')$

$$\frac{\|F^t(q) - F^t(p)\|}{\|F^t(q') - F^t(p)\|} \rightarrow 0, \text{ as } t \rightarrow \infty. \quad (11.416)$$

We already used this property, to arrive at the expression for the stable fibers.

- $f^s(p) \cap f^s(p') = \{\}$ if $p \neq p'$.

$$f^s(p) \cap f^s(p') = \left\{ (x, y, z) : x = x_p + \frac{\varepsilon}{2-\varepsilon} y^2, z = 0 \right\} \cap \left\{ (x', y', z') : x' = x'_p + \frac{\varepsilon}{2-\varepsilon} y'^2, z' = 0 \right\} \quad (11.417)$$

For this to be nonempty, we need $x = x'$, $y = y'$, $z = z'$, but this means

$$x_p + \frac{\varepsilon}{2-\varepsilon} y^2 = x'_p + \frac{\varepsilon}{2-\varepsilon} y'^2, \quad (11.418)$$

which cannot be equal if $x_p \neq x'_p$.

- The stable fiber $f^s(p)$ is a C^{r-1} smooth function of its base point p . The set $f^s(p) = f^s(x_p)$, which is given by a graph of a C^∞ function of x_p .
- The stable fibers C^r -smoothly persist under small C^1 perturbations of the dynamical system.

First, consider that the perturbation is only in ε : $\varepsilon := \varepsilon_0 + \Delta\varepsilon$. The stable fiber is well defined for any $\varepsilon \neq 0$ and it is also differentiable in ε .

For a general C^1 -small perturbation to the system, (say, order δ) the flow-map is also perturbed by an order δ term, since we can Taylor-expand it to first order

$$F^t(p; \delta) = F^t(p; 0) + \delta \frac{\partial F^t}{\partial \delta} + O(\delta^2). \quad (11.419)$$

Then, it can probably (?) be argued that this is only an order δ perturbation to the equation that prescribes the maximal convergence rate for points in the fiber (exercise (c)).

(f) For any trajectory $\gamma \subset M_\varepsilon$, find explicit expressions for $W_{loc}^{ss}(\gamma)$, $W_{loc}^{uu}(\gamma)$, $W_{loc}^s(\gamma)$, $W_{loc}^u(\gamma)$.

From the definition,

$$W_{loc}^{ss}(\gamma) = \bigcup_{\bar{p} \in \gamma(p)} f^s(\bar{p}). \quad (11.420)$$

On the manifold, the trajectory is given by $\gamma(p) = \gamma(x_p) = x_p e^{-\varepsilon t}$. After substituting, we obtain

$$W_{loc}^{ss}(\gamma) = \bigcup_{x \in \gamma(x_p)} \left\{ (\bar{x}, \bar{y}, \bar{z}) : \bar{x} = x + \frac{\varepsilon}{2-\varepsilon} \bar{y}^2, \bar{z} = 0 \right\} = \bigcup_{t \geq 0} \left\{ (\bar{x}, \bar{y}, \bar{z}) : \bar{x} = x_p e^{-\varepsilon t} + \frac{\varepsilon}{2-\varepsilon} \bar{y}^2, \bar{z} = 0 \right\}, \quad (11.421)$$

which is the same set as (if $x_p > 0$) $W_{loc}^{ss}(\gamma) = \{(x, y, z) : \frac{\varepsilon}{2-\varepsilon} y^2 \leq x \leq x_p + \frac{\varepsilon}{2-\varepsilon} y^2, z = 0\}$. And in case $x_p < 0$ $W_{loc}^{ss}(\gamma) = \{(x, y, z) : \frac{\varepsilon}{2-\varepsilon} y^2 + x_p \leq x \leq \frac{\varepsilon}{2-\varepsilon} y^2, z = 0\}$.

For the unstable fibers, we have (if $x_p > 0$)

$$W_{loc}^{uu}(\gamma) = \bigcup_{\bar{p} \in \gamma(p)} f^u(\bar{p}) = \bigcup_{t \geq 0} \{x_p e^{-\varepsilon t}, y = 0\} = \{(x, y, z) : x \leq x_p, y = 0\} \quad (11.422)$$

and if $x_p < 0$

$$W_{loc}^{uu}(\gamma) = \{(x, y, z) : x \geq x_p, y = 0\} \quad (11.423)$$

The local stable manifold of trajectory $\gamma(p)$ is the set of initial conditions, through which trajectories converge to $\gamma(p)$.

$$W_{loc}^s(\gamma) = \{(x, y, z) : \|F^t(x, y, z) - \gamma(p)\| \rightarrow 0 \text{ as } t \rightarrow \infty\}. \quad (11.424)$$

Using (11.394), we can write the difference between the trajectories as

$$F^t(x, y, z) - \gamma(x_p) = F^t(x, y, z) - F^t(x_p, 0, 0) = \begin{bmatrix} e^{-\varepsilon t} \left(x - \frac{y^2\varepsilon}{2-\varepsilon} - x_p \right) + \frac{y^2\varepsilon}{2-\varepsilon} e^{-2t} \\ ye^{-t} \\ ze^t \end{bmatrix}. \quad (11.425)$$

For convergence, we need to set $z = 0$ and then we have

$$\|F^t(x, y, 0) - F^t(x_p, 0, 0)\| = \left[e^{-\varepsilon t} \left(x - \frac{y^2\varepsilon}{2-\varepsilon} - x_p \right) + \frac{y^2\varepsilon}{2-\varepsilon} e^{-2t} \right]^2 + y^2 e^{-2t}, \quad (11.426)$$

which converges for all x, y . $W_{loc}^s(\gamma) = \{(x, y, z) : z = 0\}$.

The local unstable manifold of trajectory $\gamma(p)$ is the set of initial conditions, through which trajectories converge to $\gamma(p)$ in backward time.

Using the expression (11.401), we have

$$F^{-t}(x, y, z) - F^{-t}(x_p, 0, 0) = \begin{bmatrix} e^{\varepsilon t} \left(x - \frac{y^2\varepsilon}{2-\varepsilon} - x_p \right) + \frac{y^2\varepsilon}{2-\varepsilon} e^{2t} \\ ye^t \\ ze^{-t} \end{bmatrix}. \quad (11.427)$$

Setting $y = 0$, we get

$$\|F^{-t}(x, 0, z) - F^{-t}(x_p, 0, 0)\| = e^{2\varepsilon t} (x - x_p)^2 + z^2 e^{-2t}, \quad (11.428)$$

which converges if $x = x_p$. $W_{loc}^u(\gamma) = \{(x, y, z) : x = x_p, y = 0\}$.

Solution (9.6). Consider the slowly forced pendulum with viscous damping

$$\ddot{\varphi} + k\dot{\varphi} + \sin \varphi = F_0 \sin(\varepsilon t)$$

$k > 0$ is the damping coefficient, $F_0 < 1$ and $0 < \varepsilon \ll 1$. To write the system in first order, autonomous form, let $v = \dot{\varphi}$ and $\psi = \varepsilon t$.

$$\begin{aligned}\dot{\varphi} &= v \\ \dot{v} &= -kv - \sin \varphi + F_0 \sin \psi \\ \dot{\psi} &= \varepsilon.\end{aligned}$$

This is a problem that has separated timescales, and can be viewed as a singularly perturbed problem, on the fast timescale. Here, φ and v are the fast variables, while ψ is the slow variable. Setting $\varepsilon = 0$, we get the fast subsystem:

$$\begin{aligned}\dot{\varphi} &= v \\ \dot{v} &= -kv - \sin \varphi + F_0 \sin \psi \\ \dot{\psi} &= 0.\end{aligned}$$

To obtain the critical manifold C_0 , we look for fixed points of this system. Because of the first equation, we must have $v = 0$. Substituting it into the second equation, we get

$$0 = -\sin \varphi + F_0 \sin \psi.$$

Since $F_0 < 1$, this has two solutions for all ψ .

$$\varphi_1(\psi) = \arcsin(F_0 \sin \psi) \tag{11.429}$$

or

$$\varphi_2(\psi) = -\arcsin(F_0 \sin \psi) + \pi. \tag{11.430}$$

The extended phase space is spanned by the variables (φ, v, ψ) . But since φ and ψ are cyclic, the phase space can be viewed as the manifold $S^1 \times S^1 \times \mathbb{R} \sim T^2 \times \mathbb{R}$. Because we were able to write the fixed points as a graph over ψ , the critical manifold is

$$C_0 = \{(\varphi, \psi, v) \in S^1 \times S^1 \times \mathbb{R} : v = 0, \varphi = \arcsin(F_0 \sin \psi), \varphi = -\arcsin(F_0 \sin \psi) + \pi\}.$$

This is a union of two disjoint curves. The normal hyperbolicity of the manifold depends on the hyperbolicity of the fixed points, in the fast subsystem.

$$\begin{aligned}\frac{d}{dt} \begin{bmatrix} \varphi \\ v \end{bmatrix} &= \begin{bmatrix} v \\ -kv - \sin \varphi + F_0 \sin \psi \end{bmatrix} := f(\varphi, v) \\ \dot{\psi} &= 0\end{aligned}$$

The Jacobian matrix $D_{\varphi,v}f(\varphi, v)$ is

$$D_{\varphi,v}f(\varphi, v) = \begin{bmatrix} 0 & 1 \\ -\cos \varphi & -k \end{bmatrix},$$

which we have to evaluate at the points (11.429) and (11.430) and look for its eigenvalues. The characteristic polynomial is

$$\lambda^2 + k\lambda + \cos \varphi_{1,2}(\psi) = 0,$$

with roots

$$\lambda_{\pm} = -\frac{k}{2} \pm \sqrt{\frac{k^2}{4} - \cos \varphi_{1,2}(\psi)}. \quad (11.431)$$

Here, by substituting (11.429) and (11.430), we get

$$\cos \varphi_1 = \cos(\arcsin(F_0 \sin \psi)) = \sqrt{1 - F_0^2 \sin^2 \psi} > 0$$

and

$$\cos \varphi_2 = \cos(-\arcsin(F_0 \sin \psi) + \pi) = -\cos(\arcsin(F_0 \sin \psi)) = -\sqrt{1 - F_0^2 \sin^2 \psi} < 0,$$

by the properties $\cos(\pi - \theta) = -\cos \theta$ and $\cos(\arcsin(x)) = \sqrt{1 - x^2}$. Evaluating (11.431) at these points, we see that in case of the first root, φ_1 , the square root is strictly smaller than $k/2$, both eigenvalues are negative. At the root φ_2 , the square root is strictly larger than $k/2$, which means λ_+ becomes positive.

Both fixed-curves $\varphi_{1,2}(\psi)$ of the fast subsystem are hyperbolic, φ_1 is asymptotically stable and φ_2 is unstable. This means that the critical manifold C_0 is a normally hyperbolic invariant manifold.

The full system is an ε -size perturbation to the fast subsystem, and we can conclude that there exists a normally hyperbolic invariant manifold C_ε , ε close to C_0 , which is the slow manifold.

The full extended phase space $S^1 \times S^1 \times \mathbb{R}$, can be viewed as the product of a 2-torus and the real line. The critical manifold (and ε close to it, the slow manifold) then, is an union of two curves, that run on the 2-torus, corresponding to the section $v = 0$.

We can represent this torus as an embedded submanifold of \mathbb{R}^3 , parametrized by (φ, ψ) .

The slow manifold is shown in Fig. 11.4. The red curve is the unstable part of C_ε and the blue one is the stable part. They trace out closed curves, φ_1 oscillating between $-\arcsin(F_0)$ and $\arcsin(F_0)$ and φ_2 between $\pi + \arcsin(F_0)$ and $\pi - \arcsin(F_0)$. In the extended phase space, $S^1 \times S^1 \times \mathbb{R}$ (the torus shown is just a submanifold of this), the curve φ_2 has a local stable manifold and an unstable manifold. Trajectories in the stable manifold would converge to φ_2 , while all other nearby trajectories would converge to φ_1 .

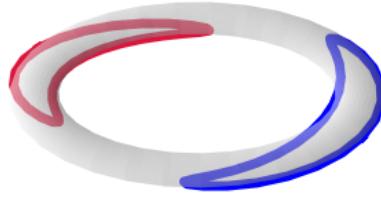


Figure 11.4: Slow manifold of the slowly forced pendulum. The torus shown is the $v = 0$ section of phase space. The colored curves make up the slow manifold, the blue one is $\varphi_1(\psi)$ and the red one is $\varphi_2(\psi)$.

Solution (9.7). Stommel's model for the Thermohaline circulation is

$$\dot{x} = -\frac{1}{\tau_x}(x - 1) + \frac{1}{\tau_y}x[1 + \eta^2(x - y)^2] \quad (11.432)$$

$$\dot{y} = \frac{\mu}{\tau_y} - \frac{1}{\tau_y}y[1 + \eta^2(x - y)^2]. \quad (11.433)$$

The variables are x, y , where $x(t)$ is the temperature-gradient and $y(t)$ is the salinity gradient. The parameter μ is the freshwater flux, η is the nonlinear coupling constant. The relaxation times for the two processes are τ_x and τ_y , which satisfy $\tau_x/\tau_y = \varepsilon \ll 1$.

(a) Show that Stommel's model has a globally attracting slow manifold that governs the asymptotic behavior of THC. Find a leading order approximation to this manifold.

Solution

Introducing the new timescale $s = t\tau_y$, $d/dt = 1/\tau_y d/ds$:

$$\begin{aligned} \frac{1}{\tau_y} \frac{dx}{ds} &= -\frac{1}{\tau_x}(x - 1) + \frac{1}{\tau_y}x[1 + \eta^2(x - y)^2] \\ \frac{1}{\tau_y} \frac{dy}{ds} &= \frac{\mu}{\tau_y} - \frac{1}{\tau_y}y[1 + \eta^2(x - y)^2]. \end{aligned}$$

Using the definition of ε , we have the singular perturbation problem, with x the fast variable

and y the slow variable.

$$\varepsilon \frac{dx}{ds} = -(x - 1) + \varepsilon x[1 + \eta^2(x - y)^2] := f(x, y, \varepsilon) \quad (11.434)$$

$$\frac{dy}{ds} = \mu - y[1 + \eta^2(x - y)^2] := g(x, y). \quad (11.435)$$

Switching timescales again, by $s = \varepsilon\tau$ and denoting the differentiation with respect to τ , by $(\cdot)'$

$$\begin{aligned} x' &= -(x - 1) + \varepsilon x[1 + \eta^2(x - y)^2] \\ y' &= \varepsilon(\mu - y[1 + \eta^2(x - y)^2]). \end{aligned}$$

Setting $\varepsilon = 0$, we get the fast subsystem:

$$x' = -(x - 1)$$

and $y' = 0$.

We look for the fixed point of this subsystem, which is $x = 1$, for any y , which means we can write the critical manifold as

$$C_0 = \{(x, y) : x = 1\}. \quad (11.436)$$

The stability of the critical manifold depends on the stability of the fixed point $x = 1$ in the fast subsystem. Since

$$\frac{d}{dx}f(x, y, 0) = -1$$

this fixed point is hyperbolic and also attracting with a rate $e^{-\tau}$. Because of this, the critical manifold C_0 is a normally attracting invariant manifold for the system, for $\varepsilon = 0$.

As a result of Fenichel's theorem, we conclude that for small enough ε , there is a normally attracting invariant manifold C_ε for the system (11.432), which is $O(\varepsilon)$ close to C_0 . This is the slow manifold, which we can represent as a graph

$$C_\varepsilon = \{(x, y) : x = \varphi(y, \varepsilon)\}.$$

Because of the smoothness properties, we can expand it in terms of ε : $x = \varphi(y, \varepsilon) = \varphi_0(y) + \varepsilon\varphi_1(y) + O(\varepsilon^2) = 1 + \varphi_1(y) + O(y^2)$.

On one hand, on the manifold we have $x = \varphi(y, \varepsilon)$, which we can differentiate with respect to τ .

$$x' = \frac{d\varphi}{dy}y' = \varepsilon^2 \frac{d\varphi_1(y)}{dy} (\mu - y - y\eta^2(1 + \varepsilon\varphi_1(y) - y)^2) = O(\varepsilon^2). \quad (11.437)$$

On the other hand, we can use the x' equation restricted to the manifold.

$$\begin{aligned} x' &= f(\varphi(y, \varepsilon), y, \varepsilon) = -\varepsilon\varphi_1(y) + \varepsilon x[1 + \eta^2(x - y)^2] = \\ &= -(x - 1) + \varepsilon x[1 + \eta^2(x^2 - 2xy + y^2)] = \\ &= -\varepsilon\varphi_1(y) + \varepsilon + \varepsilon\eta^2(1 - 2y + y^2) + O(\varepsilon^2). \end{aligned}$$

For the manifold to be invariant, these two expressions must match in all orders of ε . Since in (11.437), there were no order ε terms, we must get 0 for the coefficient of the $O(\varepsilon)$ term in the latter expression. This means

$$\varphi_1(y) = 1 + \eta^2(1 - y)^2.$$

To leading-order, the slow manifold is described by the graph $x = 1 + \varepsilon(1 + \eta^2(1 - y)^2)$.

(b) Compute the leading-order reduced flow on the slow manifold. Determine qualitatively the possible dynamical behaviors on the slow manifold as the parameters μ and η are varied.

Solution

To derive the reduced order flow, we Taylor-expand the equation governing the slow variable (on the original timescale, s):

$$g(x, y) = g(\varphi_0(y), y) + \varepsilon \left(\frac{dg}{dx} \Big|_{(\varphi_0(y), y)} \right) \varphi_1(y) + O(\varepsilon^2)$$

From (11.435), $g(x, y) = \mu - y[1 + \eta^2(x - y)^2]$

$$g(x, y) = \mu - y[1 + \eta^2(1 - y)^2] - 2\varepsilon\varphi_1(y)\eta^2y(1 - y) + O(\varepsilon^2)$$

Plugging in the expression for φ_1 , we have

$$g(x, y) = \mu - y[1 + \eta^2(1 - y)^2] - 2\varepsilon(1 + \eta^4y(1 - y)^3) + O(\varepsilon^2)$$

Keeping only the leading order term, we get the reduced dynamics on the slow manifold

$$\frac{dy}{ds} = \mu - y[1 + \eta^2(1 - y)^2].$$

On the slow manifold, the long time behavior of solutions is dictated by the fixed points. They are the roots of the right hand side:

$$\mu = y[1 + \eta^2(1 - y)^2]$$

The graph of these functions are shown in Fig. 11.5. The fixed points of the system are given by the intersection points. To see the effect of the parameters, note that the shape of the

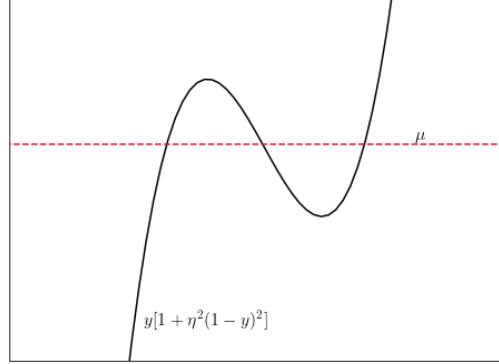


Figure 11.5: Graphs of the functions $f(y) = \mu$ and $f(y) = y[1 + \eta^2(1 - y)^2]$.

cubic function is only governed by η . It can have at most two critical points, one minimum and one maximum. These are:

$$\frac{d}{dy}y[1 + \eta^2(1 - y)^2] = 0$$

$$1 + \eta^2(1 - y)^2 - 2y\eta^2(1 - y) = 0$$

$$3\eta^2y^2 - 4y\eta^2 + 1 + \eta^2 = 0.$$

This equation has two roots (the minimum and maximum of the cubic function) if

$$16\eta^4 - 12\eta^2(1 + \eta^2) > 0,$$

or $\eta^2 > 3$. In this case, the intersection of this graph with $y = \mu$ can give 1, 2 or 3 values. If y is smaller or larger than any of the critical values of the graph, there is only one solution. In these cases the right hand side of the system $\mu - y[1 + \eta^2(1 - y)^2]$ changes sign at the fixed points, and they are stable.

When μ reaches (one of) the critical values, a new fixed point appears in a bifurcation.

When μ is between the critical values, we have 3 fixed points, two stable and one unstable.

When the above discriminant is 0, at $\eta^2 = 3$, the cubic graph has only one critical point, and for all values of μ we only have one fixed point. This fixed point is stable.

Solution (9.8). Assume $\lambda_s > 0$, then $Df(0)$ is diagonalizable with the matrix T , formed from the eigenvectors of $Df(0)$, $T = [e_{ss}, e_s]$

$$T^{-1}Df(0)T = \begin{bmatrix} -\lambda_{ss} & 0 \\ 0 & -\lambda_s \end{bmatrix}. \quad (11.438)$$

Introducing a new variable y , defined by $x = Ty$, the differential equation reads,

$$\dot{y} = T^{-1}\dot{x} = \begin{bmatrix} \dot{y}_{ss} \\ \dot{y}_s \end{bmatrix} = \begin{bmatrix} -\lambda_{ss}y_{ss} + g_{ss}(y) \\ -\lambda_s y_s + g_s(y) \end{bmatrix} \quad (11.439)$$

where g_s and g_{ss} are of order $O(|y|^2)$:

$$\limsup_{|y| \rightarrow 0} \frac{|g_{s,ss}(y)|}{|y|^2} \leq K, \quad (11.440)$$

for some K . To characterize the strong stable manifold W , consider a modified form of (11.439). Outside a small neighborhood $B_\delta := [-\delta, \delta] \times [-\delta, \delta] \subset \mathbb{R}^2$, the vector field is smoothly changed to a linear one.

This construction is done by using C^∞ bump-functions $\psi_\delta : \mathbb{R} \rightarrow \mathbb{R}$, with the properties

$$\psi_\delta(r) = 1 \text{ for } |r| \leq \delta \quad (11.441)$$

$$\psi_\delta(r) \leq 1 \text{ for } \delta \leq |r| \leq 2\delta \quad (11.442)$$

$$\psi_\delta(r) = 0 \text{ for } 2\delta \leq |r| \quad (11.443)$$

$$|\psi_\delta(r)'| \leq \frac{K_0}{\delta}. \quad (11.444)$$

Then, define the new functions, the components of the modified nonlinear part of the vector field (11.439) as

$$G_{s,ss} : \mathbb{R}^2 \rightarrow \mathbb{R} \quad (11.445)$$

$$G_{s,ss}(y) := \psi_\delta(y_s)\psi_\delta(y_{ss})g_{s,ss}(y). \quad (11.446)$$

The gradient of this function is (with the notation $\psi = \psi_\delta(y_{ss})\psi_\delta(y_s)$)

$$\nabla G_{s,ss}(y) = g_{s,ss}(y) \begin{bmatrix} \psi_\delta(y_s)' \psi_\delta(y_{ss}) \\ \psi_\delta(y_{ss})' \psi_\delta(y_s) \end{bmatrix} + \psi(y) \nabla g_{s,ss}(y). \quad (11.447)$$

The norm can be bounded by

$$|\nabla G_{s,ss}(y)| \leq \sqrt{2}|g_{s,ss}| \frac{K_0}{\delta} + |\nabla g_{s,ss}(y)| \leq K_0 C \delta + C_1 \delta = K_2 \delta, \quad (11.448)$$

because of the properties of the bump functions and we use the fact that $|g_{s,ss}|$ can be bounded by a quadratic function on the region B_δ .

The modified vector takes the form

$$\begin{bmatrix} \dot{y}_{ss} \\ \dot{y}_s \end{bmatrix} = \begin{bmatrix} -\lambda_{ss}y_{ss} + G_{ss}(y) \\ -\lambda_s y_s + G_s(y) \end{bmatrix}. \quad (11.449)$$

The strong stable manifold of system (11.449) is \overline{W} , which coincides with the strong stable manifold of the original system on B_δ . It is distinguished by the property, that trajectories on the manifold decay at least as fast as $e^{-\gamma t}$ for positive times:

$$W = \overline{W} \cap B_\delta = \left\{ y_0 \in B_\delta : \sup_{t \geq 0} e^{\gamma t} |F^t(y_0)| < \infty \right\}. \quad (11.450)$$

To prove the existence of the strong stable manifold, we put system (11.449) in a weak form, to make use of Banach's fixed point theorem.

As seen in the lecture notes, the solutions of a system of the form

$$\dot{x} = Ax + f(x, t) \quad (11.451)$$

are the solutions to the integral equation

$$x(t) = e^{A(t-t_0)}x_0 + \int_{t_0}^t e^{A(t-\tau)}f(x(\tau), \tau)d\tau. \quad (11.452)$$

Applying this general form to system (11.449), we have

$$y_{ss}(t) = e^{-\lambda_{ss}(t-t_{ss})}y_{ss}(t_{ss}) + \int_{t_0}^t e^{-\lambda_{ss}(t-\tau)}G_{ss}(y(\tau))d\tau \quad (11.453)$$

$$y_s(t) = e^{-\lambda_s(t-t_s)}y_s(t_s) + \int_{t_0}^t e^{-\lambda_s(t-\tau)}G_s(y(\tau))d\tau. \quad (11.454)$$

Assuming W is nonempty, pick an initial point $(y_{ss}(t_{ss}), y_s(t_s)) \in W$. It has the property

$$e^{\gamma\tau}|y_s(\tau)| \leq e^{\gamma\tau}|y(\tau)| < \infty, \text{ for } \tau \geq 0. \quad (11.455)$$

Moreover, the norm of the first term in (11.454) can be written as

$$e^{-\lambda_s(t-t_s)}|y_s(t_s)| = |y_s(t_s)|e^{\gamma t_s}e^{-\lambda_s t}e^{-\lambda_s t_s}e^{-\gamma t_s} = |y_s(t_s)|e^{\gamma t_s}e^{-\lambda_s t}e^{(\lambda_s - \gamma)t_s}. \quad (11.456)$$

Because $|y_s(t_s)|e^{\gamma t_s}e^{-\lambda_s t}$ stays bounded for given t , (using the previous line, with $\tau = t_s$) and $\lambda_s < \gamma$, we have

$$e^{-\lambda_s(t-t_s)}|y_s(t_s)| \rightarrow 0 \text{ as } t_s \rightarrow \infty. \quad (11.457)$$

Then, we can choose the initial times to be $t_s = \infty$ and $t_{ss} = 0$. Also, $y_{ss}(0)$ is the first coordinate of the initial point of the trajectory, when it enters B_δ , that is, we can write $y_{ss}(0) = \delta$ to obtain the integral equations

$$y_{ss}(t) = e^{-\lambda_{ss}t}\delta + \int_0^t e^{-\lambda_{ss}(t-\tau)}G_{ss}(y(\tau))d\tau \quad (11.458)$$

$$y_s(t) = - \int_t^\infty e^{-\lambda_s(t-\tau)}G_s(y(\tau))d\tau. \quad (11.459)$$

Claim: The set

$$B := \left\{ y(t) : y \in C^0([0, \infty), B_\delta), \quad \sup_{t \geq 0} e^{\gamma t} |y(t)| \leq K < \infty \right\}, \quad (11.460)$$

equipped with the metric

$$d(y_1, y_2) := \sup_{t \geq 0} |y_1 - y_2| e^{\gamma t}. \quad (11.461)$$

becomes a complete metric space, on which the map

$$\mathcal{F}_\delta : B \rightarrow B, \quad (11.462)$$

defined by the integral equations (11.458), (11.459)

$$\mathcal{F}_\delta(y) = \begin{bmatrix} e^{-\lambda_{ss}t}\delta + \int_0^t e^{-\lambda_{ss}(t-\tau)}G_{ss}(y(\tau))d\tau \\ - \int_t^\infty e^{-\lambda_s(t-\tau)}G_s(y(\tau))d\tau \end{bmatrix} \quad (11.463)$$

is a contraction mapping, that is, for $q < 1$,

$$d(\mathcal{F}_\delta(y_1), \mathcal{F}_\delta(y_2)) \leq q d(y_1, y_2). \quad (11.464)$$

First, we show that the map $d(y_1, y_2)$ is a metric, by showing that it satisfies the defining properties of a metric

$$d(y_1, y_2) = 0 \text{ if and only if } |y_1 - y_2| = 0 \iff y_1 = y_2 \quad (11.465)$$

$$d(y_1, y_2) = \sup_{t \geq 0} |y_1 - y_2| e^{\gamma t} = \sup_{t \geq 0} |y_2 - y_1| e^{\gamma t} = d(y_2, y_1) \quad (11.466)$$

$$d(y_1, y_3) = \sup_{t \geq 0} |y_1 - y_3| e^{\gamma t} \leq \sup_{t \geq 0} (|y_1 - y_2| + |y_2 - y_3|) e^{\gamma t} \quad (11.467)$$

$$= \sup_{t \geq 0} |y_1 - y_2| e^{\gamma t} + \sup_{t \geq 0} |y_2 - y_3| e^{\gamma t} = d(y_1, y_2) + d(y_2, y_3). \quad (11.468)$$

In the usual metric, a sequence of uniformly continuous functions does have a continuous limit. If, in addition, the sequence of functions is also bounded in this weighted sense, the limit will also have this property, which means that B is complete.

Next, we show that the map \mathcal{F}_δ defined above maps to B .

$$|\mathcal{F}_\delta(y(t))| \leq |y_s(t)| + |y_{ss}(t)| \quad (11.469)$$

$$\leq \left| e^{-\lambda_{ss}t}\delta + \int_0^t e^{-\lambda_{ss}(t-\tau)}G_{ss}(y(\tau))d\tau \right| + \left| \int_t^\infty e^{-\lambda_s(t-\tau)}G_s(y(\tau))d\tau \right| \quad (11.470)$$

$$\leq e^{-\lambda_{ss}t}\delta + \int_0^t e^{-\lambda_{ss}(t-\tau)}|G_{ss}(y(\tau))|d\tau + \int_t^\infty e^{-\lambda_s(t-\tau)}|G_s(y(\tau))|d\tau. \quad (11.471)$$

Using the mean value theorem, and the bound developed above,

$$|\nabla G_{s,ss}| \leq g_{s,ss} \leq K_2\delta, \quad |G_{s,ss}(y_{s,ss})| \leq K_2\delta|y_{s,ss}|. \quad (11.472)$$

We can derive the following

$$|\mathcal{F}_\delta(y(t))| = e^{-\lambda_{ss}t}\delta + \int_0^t e^{-\lambda_{ss}(t-\tau)}|G_{ss}(y(\tau))|d\tau + \int_t^\infty e^{-\lambda_s(t-\tau)}|G_s(y(\tau))|d\tau \quad (11.473)$$

$$\leq e^{-\lambda_{ss}t}\delta + \int_0^t e^{-\lambda_{ss}(t-\tau)}K_2\delta|y_{ss}(\tau)|d\tau + \int_t^\infty e^{-\lambda_s(t-\tau)}K_2\delta|y_s(\tau)|d\tau. \quad (11.474)$$

Multiply by $e^{\gamma t}$ and take the supremum

$$\sup_{t \geq 0} |\mathcal{F}_\delta(y(t))|e^{\gamma t} \leq \sup_{t \geq 0} \left\{ \delta e^{(\gamma - \lambda_{ss})t} + \int_0^t e^{-\lambda_{ss}(t-\tau)}K_2\delta|y_{ss}(\tau)|e^{\gamma t}d\tau \right\} \quad (11.475)$$

$$+ \int_t^\infty e^{-\lambda_s(t-\tau)}K_2\delta|y_s(\tau)|e^{\gamma t}d\tau \quad (11.476)$$

$$\leq \sup_{t \geq 0} \left\{ \delta + \int_0^t e^{-\lambda_{ss}(t-\tau)}K_2\delta|y_{ss}(\tau)|e^{\gamma \tau}e^{-\gamma \tau}e^{\gamma t}d\tau \right\} \quad (11.477)$$

$$+ \int_t^\infty e^{-\lambda_s(t-\tau)}K_2\delta|y_s(\tau)|e^{\gamma \tau}e^{-\gamma \tau}e^{\gamma t}d\tau \quad (11.478)$$

$$\leq \sup_{t \geq 0} \left\{ \delta + K_2\delta K \int_0^t e^{(t-\tau)(\gamma - \lambda_{ss})}d\tau + K_2\delta K \int_t^\infty e^{(t-\tau)(\gamma - \lambda_s)}d\tau \right\}. \quad (11.479)$$

After carrying out the integrals, and using $\gamma - \lambda_{ss} < 0$ and $\gamma - \lambda_s > 0$

$$\sup_{t \geq 0} |\mathcal{F}_\delta(y(t))| e^{\gamma t} \leq \sup_{t \geq 0} \left\{ \delta + K_2 \delta K \left[\frac{1}{\lambda_{ss} - \gamma} e^{(t-\tau)(\gamma - \lambda_{ss})} \right]_0^t \right. \quad (11.480)$$

$$\left. + K_2 \delta K \left[\frac{1}{\lambda_s - \gamma} e^{(t-\tau)(\gamma - \lambda_s)} \right]_t^\infty \right\} \quad (11.481)$$

$$= \sup_{t \geq 0} \left\{ \delta + K_2 \delta K \frac{1 - e^{t(\gamma - \lambda_{ss})}}{\lambda_{ss} - \gamma} + K_2 \delta K \frac{1}{\gamma - \lambda_s} \right\} \quad (11.482)$$

$$\leq \delta + K_2 \delta K \left(\frac{1}{\lambda_{ss} - \gamma} + \frac{1}{\gamma - \lambda_s} \right) \quad (11.483)$$

$$= K \left[\frac{\delta}{K} + K_2 \delta \left(\frac{1}{\lambda_{ss} - \gamma} + \frac{1}{\gamma - \lambda_s} \right) \right]. \quad (11.484)$$

Choosing δ , such that

$$\delta \leq \frac{1}{\frac{1}{K} + K_2 \left(\frac{1}{\lambda_{ss} - \gamma} + \frac{1}{\gamma - \lambda_s} \right)}, \quad (11.485)$$

guarantees that

$$\sup_{t \geq 0} |\mathcal{F}_\delta(y(t))| e^{\gamma t} \leq K < \infty, \quad (11.486)$$

which shows that the image of a function $y \in B$ under the functional \mathcal{F} is still in B .

So far we have established that B is a complete metric space and \mathcal{F} is a map from B to B . To show that \mathcal{F} is a contraction mapping, consider the difference between two image-points in B

$$|\mathcal{F}(y_1) - \mathcal{F}(y_2)| = \left| \int_0^t e^{-\lambda_{ss}(t-\tau)} [G_{ss}(y_1(\tau)) - G_{ss}(y_2(\tau))] d\tau \right| \quad (11.487)$$

$$+ \left| \int_t^\infty e^{-\lambda_s(t-\tau)} [G_s(y_1(\tau)) - G_s(y_2(\tau))] d\tau \right| \quad (11.488)$$

$$\leq \int_0^t e^{-\lambda_{ss}(t-\tau)} |[G_{ss}(y_1(\tau)) - G_{ss}(y_2(\tau))]| d\tau \quad (11.489)$$

$$+ \int_t^\infty e^{-\lambda_s(t-\tau)} |[G_s(y_1(\tau)) - G_s(y_2(\tau))]| d\tau \quad (11.490)$$

$$\leq \int_0^t e^{-\lambda_{ss}(t-\tau)} K_2 \delta |y_1(\tau) - y_2(\tau)| d\tau \quad (11.491)$$

$$+ \int_t^\infty e^{-\lambda_s(t-\tau)} K_2 \delta |y_1(\tau) - y_2(\tau)| d\tau. \quad (11.492)$$

Multiplying by $e^{\gamma t}$

$$|\mathcal{F}(y_1) - \mathcal{F}(y_2)|e^{\gamma t} \leq \int_0^t e^{-\lambda_{ss}(t-\tau)} K_2 \delta |y_1(\tau) - y_2(\tau)| e^{\gamma t} d\tau \quad (11.493)$$

$$+ \int_t^\infty e^{-\lambda_s(t-\tau)} K_2 \delta |y_1(\tau) - y_2(\tau)| e^{\gamma t} d\tau \quad (11.494)$$

$$= \int_0^t e^{-\lambda_{ss}(t-\tau)} K_2 \delta |y_1(\tau) - y_2(\tau)| e^{\gamma \tau} e^{\gamma t} e^{-\gamma \tau} d\tau \quad (11.495)$$

$$+ \int_t^\infty e^{-\lambda_s(t-\tau)} K_2 \delta |y_1(\tau) - y_2(\tau)| e^{\gamma \tau} e^{\gamma t} e^{-\gamma \tau} d\tau \quad (11.496)$$

$$\leq \int_0^t e^{(\gamma - \lambda_{ss})(t-\tau)} K_2 \delta d\tau \sup_{t \geq 0} |y_1(t) - y_2(t)| e^{\gamma t} \quad (11.497)$$

$$+ \int_t^\infty e^{(\gamma - \lambda_s)(t-\tau)} K_2 \delta d\tau \sup_{t \geq 0} |y_1(t) - y_2(t)| e^{\gamma t} \quad (11.498)$$

$$= K_2 \delta d(y_1, y_2) \left(\int_0^t e^{(\gamma - \lambda_{ss})(t-\tau)} d\tau + \int_t^\infty e^{(\gamma - \lambda_s)(t-\tau)} d\tau \right) \quad (11.499)$$

$$(11.500)$$

Thus we can surmise

$$\sup_{t \geq 0} |\mathcal{F}(y_1) - \mathcal{F}(y_2)| e^{\gamma t} \leq \delta K_2 d(y_1, y_2) \left(\frac{1}{\lambda_{ss} - \gamma} + \frac{1}{\gamma - \lambda_s} \right). \quad (11.501)$$

If we choose δ such that

$$\delta < \frac{1}{K_2 \left(\frac{1}{\lambda_{ss} - \gamma} + \frac{1}{\gamma - \lambda_s} \right)} \quad (11.502)$$

then we obtain the contraction-mapping criterion for $q < 1$. Also note that if δ satisfies the earlier inequality, it automatically satisfies the contraction mapping criterion too.

Then, we can use Banach's fixed point theorem, and conclude that \mathcal{F} has a unique fixed point in B , which, by construction must coincide with the strong stable manifold of the origin, within the small neighborhood B_δ , which was to be proved.

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