

Uncertainty quantification for activation cross sections

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Cross section evaluations

Region	Knowledge	Approach
Stable and very long-lived isotopes	Direct measurements of cross-sections, all structure details	Direct fit to measurements, R-matrix fits, fine-tuned models
Intermediate region, $T_{1/2} \sim$ seconds, minutes	Masses, lifetimes, some excited states, average level spacings	"Regional evaluations"
Very exotic, very short-live nuclei	Z,A, theoretical binding energies	Global systematics, global structure models (RPA, DFT)

ENDF, JEFF, JENDL ...
s-process, reactors

Well-defined uncertainties, detailed covariances

Supernovae,
neutron-bursts,
Material activation

?

Astrophysics libraries
R-process
nucleosynthesis

Strong model dependence, many parameters

Cross section evaluations

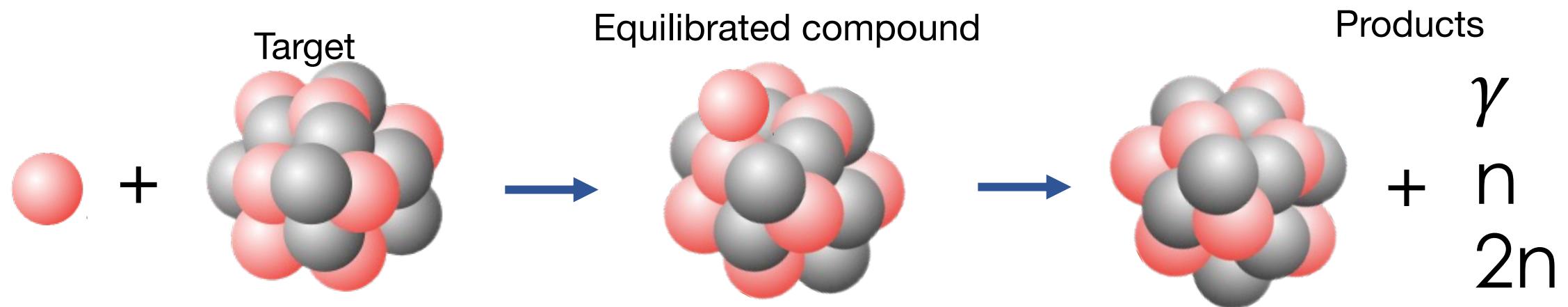
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Stable and very long-lived isotopes	Direct measurements of cross-sections, all structure details	Direct fit to measurements, R-matrix fits, fine tuned models	<i>ENDF, JEFF, JENDL ...</i> s-process, reactors	Well-defined uncertainties, detailed covariances
Intermediate region, $T_{1/2} > \text{hours}$	Masses, lifetimes, some excited states, average level spacings	"Regional evaluations"	Supernovae, neutron-bursts, Material activation	?
Very exotic, very short-live nuclei	Z,A, theoretical binding energies	Global systematics, global structure models (RPA, DFT)	<i>Astrophysics libraries</i> R-process nucleosynthesis	Strong model dependence, many parameters

Model parameters are experimentally determined for stable targets

- Focus on transmutation (activation) cross sections for fast neutrons
- Using Hauser-Feshbach statistical model (YAHFC)
- Experimental data (RIPL-3) that constrains model input:
 - Average level spacings D_0
 - Average radiative widths $\langle \Gamma_\gamma \rangle$
- Develop systematic trends of reaction model parameters
- measured cross sections for validation

Nucleus 89Y STABLE 100%	90Mo 5.56 h $\epsilon+\beta+=100\%$	91Mo 15.49 min $\epsilon+\beta+=100\%$	92Mo STABLE 14.649%	93Mo 4000 y $\epsilon=100\%$	94Mo STABLE 9.187%	95Mo STABLE 15.873%	96Mo STABLE 16.673%	97Mo STABLE 9.582%	98Mo STABLE 24.292%
148Pr / Z#	89Nb 2.03 h $\epsilon+\beta+=100\%$	90Nb 14.6 h $\epsilon+\beta+=100\%$	91Nb 680 y $\epsilon+\beta+=100\%$	92Nb 3.47e+7 y $\epsilon+\beta+=100\%$	93Nb STABLE 100%	94Nb 2.04e+4 y $\beta-=100\%$	95Nb 34.991 d $\beta-=100\%$	96Nb 23.35 h $\beta-=100\%$	97Nb 72.1 min $\beta-=100\%$
86Zr 16.5 h $\epsilon+\beta+=100\%$	87Zr 1.68 h $\epsilon+\beta+=100\%$	88Zr 83.4 d $\epsilon=100\%$	89Zr 78.364 h $\epsilon+\beta+=100\%$	90Zr STABLE 51.45%	91Zr STABLE 11.22%	92Zr STABLE 17.15%	93Zr 1.61e+6 y $\beta-=100\%$	94Zr STABLE 17.38%	95Zr 64.032 d $\beta-=100\%$
85Y 2.67 h $\epsilon+\beta+=100\%$	86Y 14.74 h $\epsilon+\beta+=100\%$	87Y 79.89 h $\epsilon+\beta+=100\%$	88Y 106.626 d $\epsilon+\beta+=100\%$	89Y STABLE 100%	90Y 64.046 h $\beta-=100\%$	91Y 58.56 d $\beta-=100\%$	92Y 3.54 h $\beta-=100\%$	93Y 10.17 h $\beta-=100\%$	94Y 18.7 min $\beta-=100\%$
84Sr STABLE 0.56%	85Sr 64.849 d $\epsilon=100\%$	86Sr STABLE 9.86%	87Sr STABLE 7%	88Sr STABLE 82.58%	89Sr 50.56 d $\beta-=100\%$	90Sr 28.91 y $\beta-=100\%$	91Sr 9.68 h $\beta-=100\%$	92Sr 2.61 h $\beta-=100\%$	93Sr 7.43 min $\beta-=100\%$
83Rb 86.2 d $\epsilon=100\%$	84Rb 32.82 d $\epsilon+\beta+=96.1\%$ $\beta-=3.9\%$	85Rb STABLE 72.17%	86Rb 18.671 d $\beta-=99.994\%$ $\epsilon=5.2e-3\%$	87Rb 4.97e+10 y 27.83% $\beta-=100\%$	88Rb 17.775 min $\beta-=100\%$	89Rb 15.39 min $\beta-=100\%$	90Rb 158 s $\beta-=100\%$	91Rb 58 s $\beta=n?$	92Rb 4.48 s $\beta=n?$
82Kr STABLE 11.593%	83Kr STABLE 11.5%	84Kr STABLE 56.987%	85Kr 10.73 y $\beta-=100\%$	86Kr STABLE 17.279%	87Kr 76.09 min $\beta-=100\%$	88Kr 2.803 h $\beta-=100\%$	89Kr 3.147 min $\beta-=100\%$	90Kr 32.32 s $\beta=n?$	91Kr 8.57 s $\beta=n?$
									92Kr 1.841 s $\beta=n=3.32e-2\%$

The Hauser-Feshbach statistical model (HF)



Model inputs can combine

- Experimental data
- Microscopic models
- Parameterizations

$$\sigma_{jk} \sim \sum \frac{T_j T_k}{T_{tot}}$$

- Experimental data (RIPL-3) that constrains model input:
 - Average level spacings D_0
 - Average radiative widths $\langle \Gamma_\gamma \rangle$
- Optical model with UQ: KDUQ

We use [YAHFC](#) [E. Ormand et al.]

<https://github.com/LLNL/Yet-Another-Hauser-Feshbach-Code>

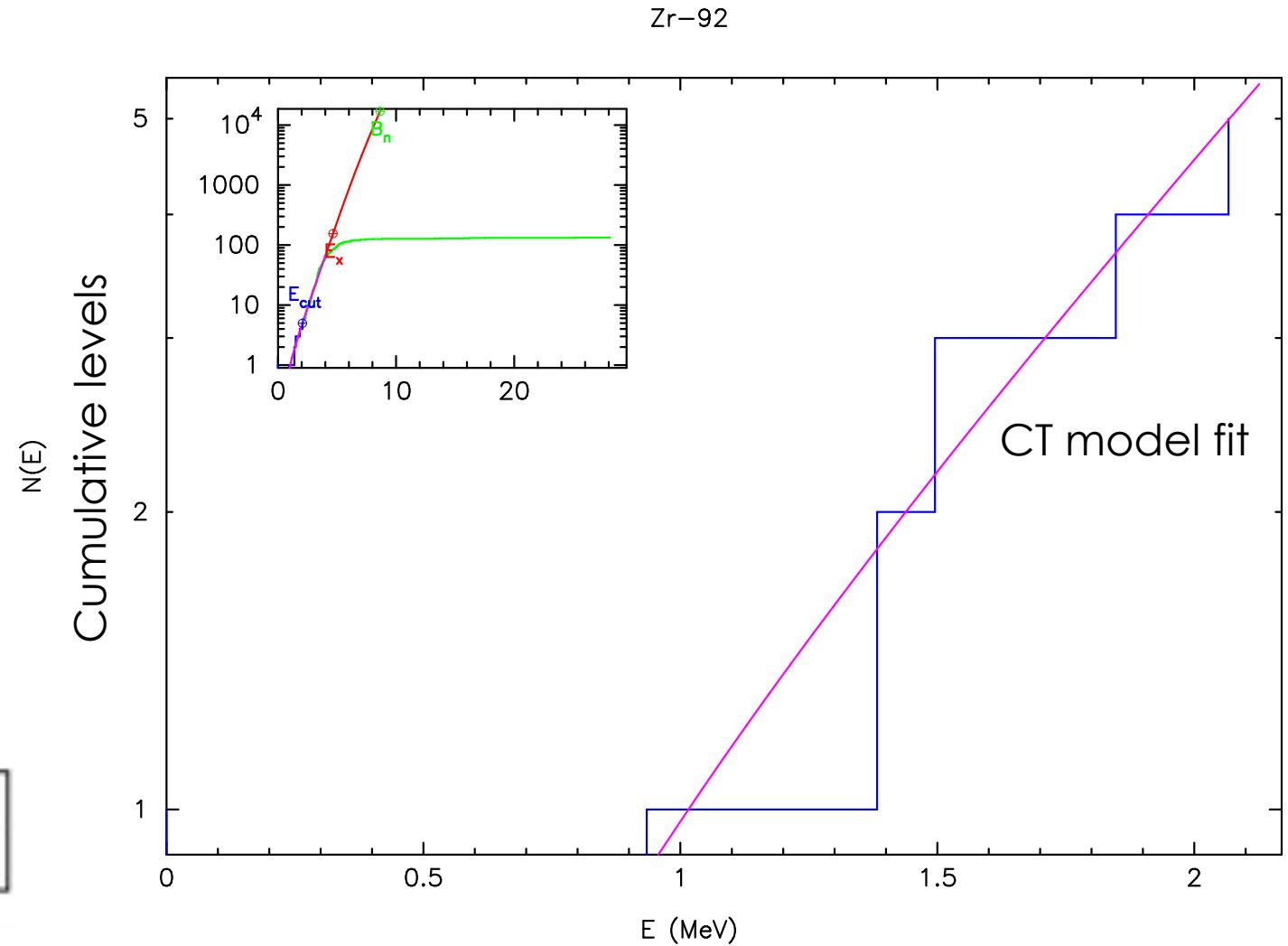
Nuclear Level Densities – low energy

- At low incident particle energy the LD is calculated from the known spectroscopic levels, which increase exponentially.
- A constant temperature model is fit to these levels

$$N(E) = \exp \left[\frac{E - E_0}{T} \right]$$

- Observable level density is

$$\rho_2(E) = \frac{dN(E)}{dE} = \frac{1}{T} \exp \left[\frac{E - E_0}{T} \right]$$



Level densities: Back-shifted Fermi gas model

$$\rho_{FG}(E_x, J^\pi) = \frac{1}{12\sqrt{2}\sigma} \frac{\exp[2\sqrt{a}U]}{a^{1/4}U^{5/4}} P_J(E_x) P_\pi(E_x), \quad (2)$$

where $U = E_x - \Delta$ and $a = a(E_x)$ are functions of the excitation energy with the backshift Δ and *sigma* is the spin cutoff parameter. The latter also enters in the angular momentum distribution

$$P_J(E_x) = \frac{2J+1}{2\sigma^2} \exp\left[-\frac{(J+1/2)^2}{2\sigma^2}\right]. \quad (3)$$

The spin cutoff σ is in principle a free parameter. However, in YAHFC, we assume the following relationship

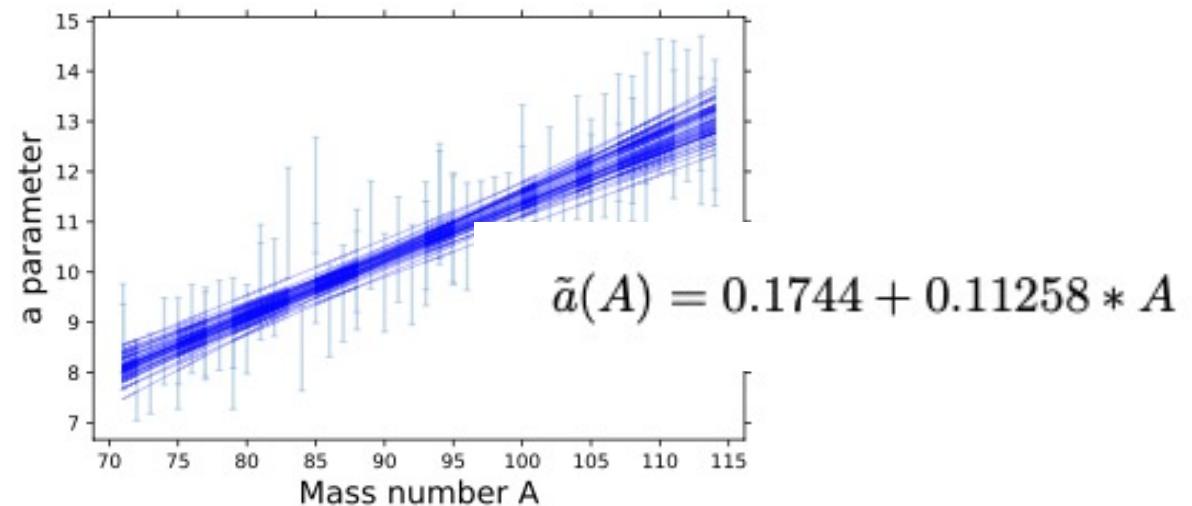
$$\sigma(E_x) = X A^{5/3} \sqrt{\frac{U}{a(U)}} \frac{a(U)}{\tilde{a}} \quad (4)$$

$$a(U, Z, N) = \tilde{a}(A) \left[1 + \delta W(Z, N) \frac{f(U)}{U} \right]$$

with

$$f(U) = 1 - \exp(-\gamma U)$$

- Back-shifted Fermi gas description:
 - Multiple parameters ($\Delta, \delta W, \tilde{a}, \sigma, \gamma$) are constrained by one experimental quantity: D_0
 - Keep fixed: σ, γ
 - Δ from comparison to mass formula
 - Parametrized \tilde{a}
 - Fit δW to known D_0



Level Densities – shell corrections

- The fitted systematic is similar to hoffman:

$$N \leq 50$$

$$c_0 = -364 \pm 70$$

$$c_1 = 17.2 \pm 3.0$$

$$c_2 = -0.20 \pm 0.03$$

$$N > 50$$

$$c_0 = -202 \pm 12$$

$$c_1 = 6.67 \pm 0.43$$

$$c_2 = -0.053 \pm 0.004$$

- Uncertainties represent the 1σ quantile of the distributions that results from the MC fitting procedure.
Close to normal distributions.

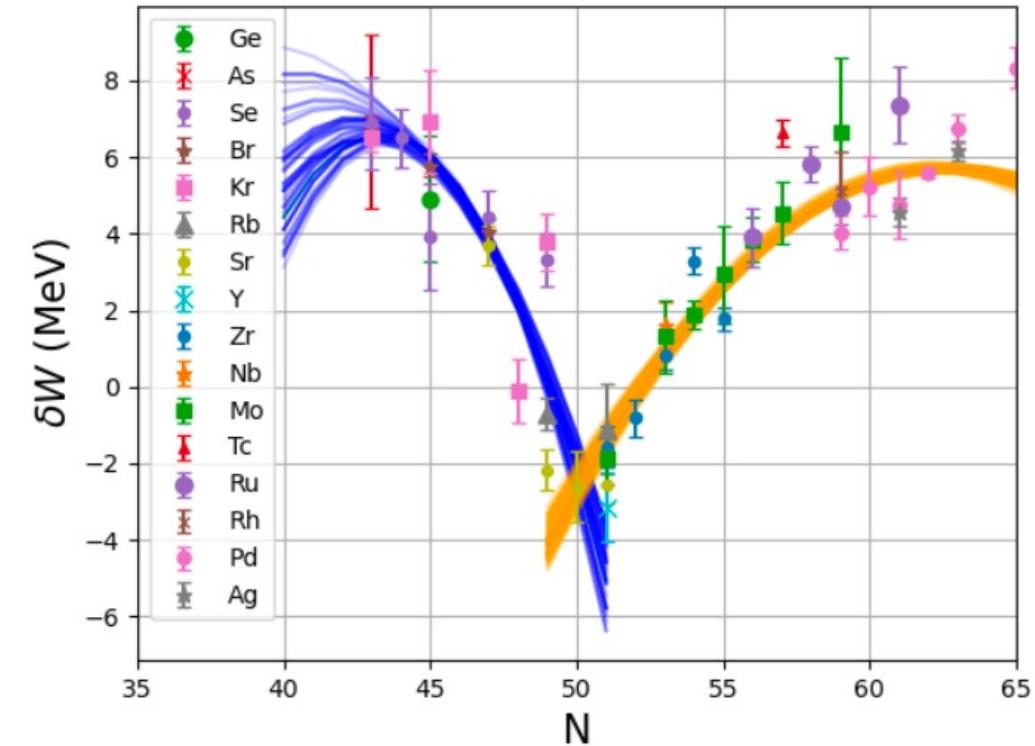
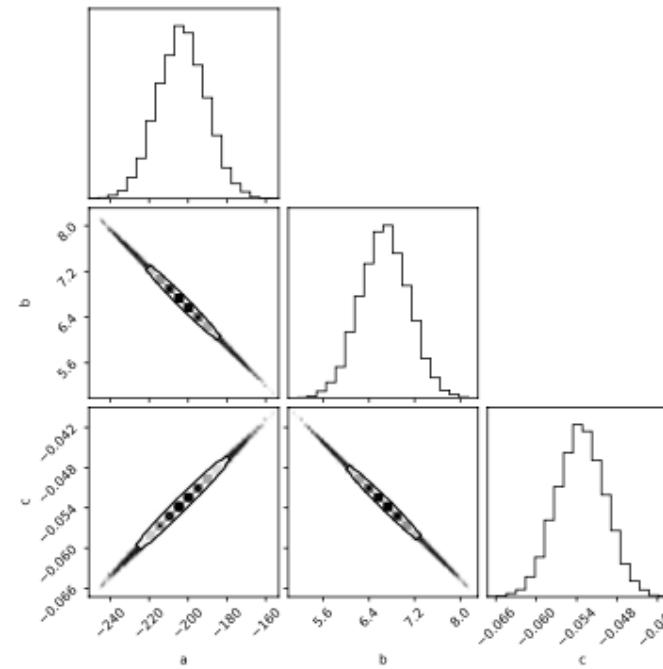


Figure 2: Results of the Monte-Carlo fitting procedure for the shell correction parameter δW .

Gamma-ray Transmission Coefficients

- γ -transmission coefficients calculated with a model that depends on multi-pole type (XL) and transition energy (ϵ) related to the γ -ray strength function

$$T_{XL}^{\gamma}(\epsilon) = 2\pi\epsilon^{2L+1}f_{XL}^{\gamma}(\epsilon)$$

- GDR model with enhanced generalized Lorentzian line shape

$$f_{E1}^{\gamma}(\epsilon) = \mathcal{N} \frac{4}{3\pi} \frac{e^2}{\hbar c} \frac{1}{M_p c^2} \times \left[\frac{\epsilon \Gamma_{GDR}(\epsilon, T_f)}{(\epsilon^2 - E_{GDR}^2)^2 + (\Gamma_{GDR}(\epsilon, T_f)\epsilon)^2} + 0.7 \frac{\Gamma_{GDR}(0, T_f)}{\epsilon^3} \right]$$

$$\langle \Gamma_{\gamma} \rangle_0 = \frac{J+1}{2J+1} \left\langle \Gamma_{\gamma} \left(B_n, J + \frac{1}{2} \right) \right\rangle + \frac{J}{2J+1} \left\langle \Gamma_{\gamma} \left(B_n, J - \frac{1}{2} \right) \right\rangle$$

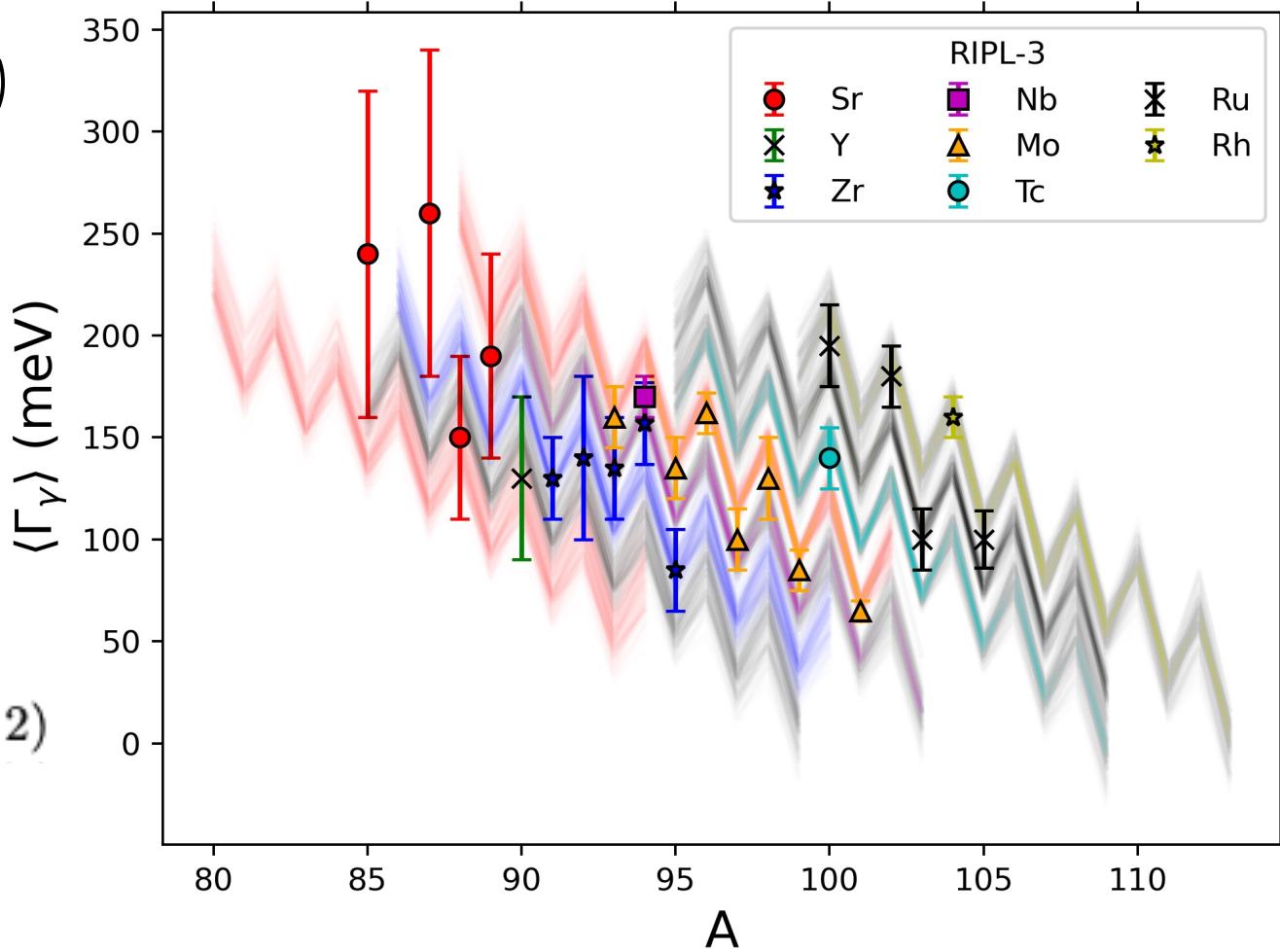
$$\Gamma_{\gamma}(E, J) = \frac{T_{\gamma}(E, J)}{2\pi\rho(E, J)} \text{ (meV)}$$

- Overall GSF is fitted to the measured $\langle \Gamma_{\gamma} \rangle_0$ by adding a E1 resonance
- **Sensitivity to GDR parameters is strongly reduced**

Gamma-ray Transmission Coefficients

- Average radiative width trends:
- $\langle \Gamma_\gamma \rangle_0$ increases with charge number (Z)
- $\langle \Gamma_\gamma \rangle_0$ generally decreases with mass number along an isotopic chain (N)
- $\langle \Gamma_\gamma \rangle_0$ shows an odd-even staggering in A , only observable for even- Z nuclei
- We choose to fit with an empirical form:

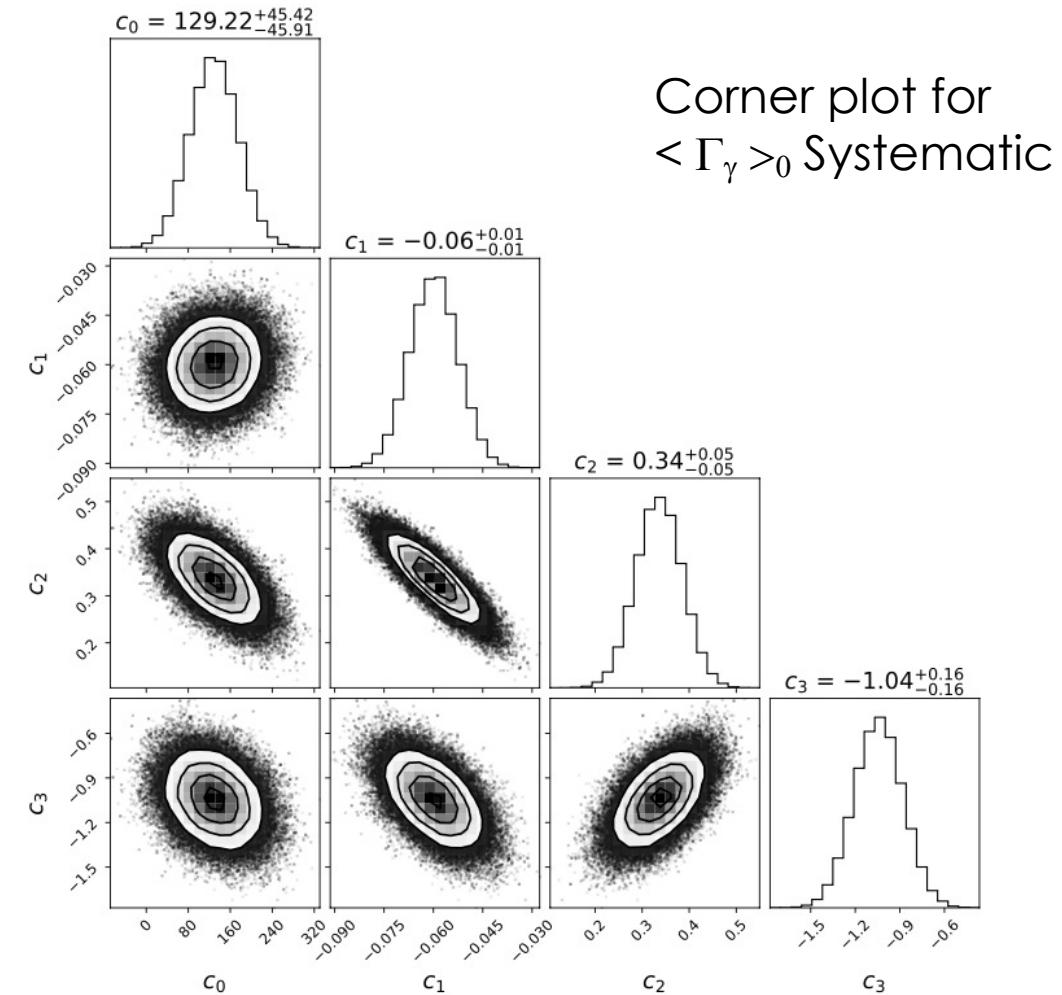
$$\langle \Gamma_\gamma \rangle_0(Z, A) = c_0 + c_1 A^2 + c_2 Z^2 + c_3 Z \bmod(A, 2)$$



Gamma-ray Transmission Coefficients

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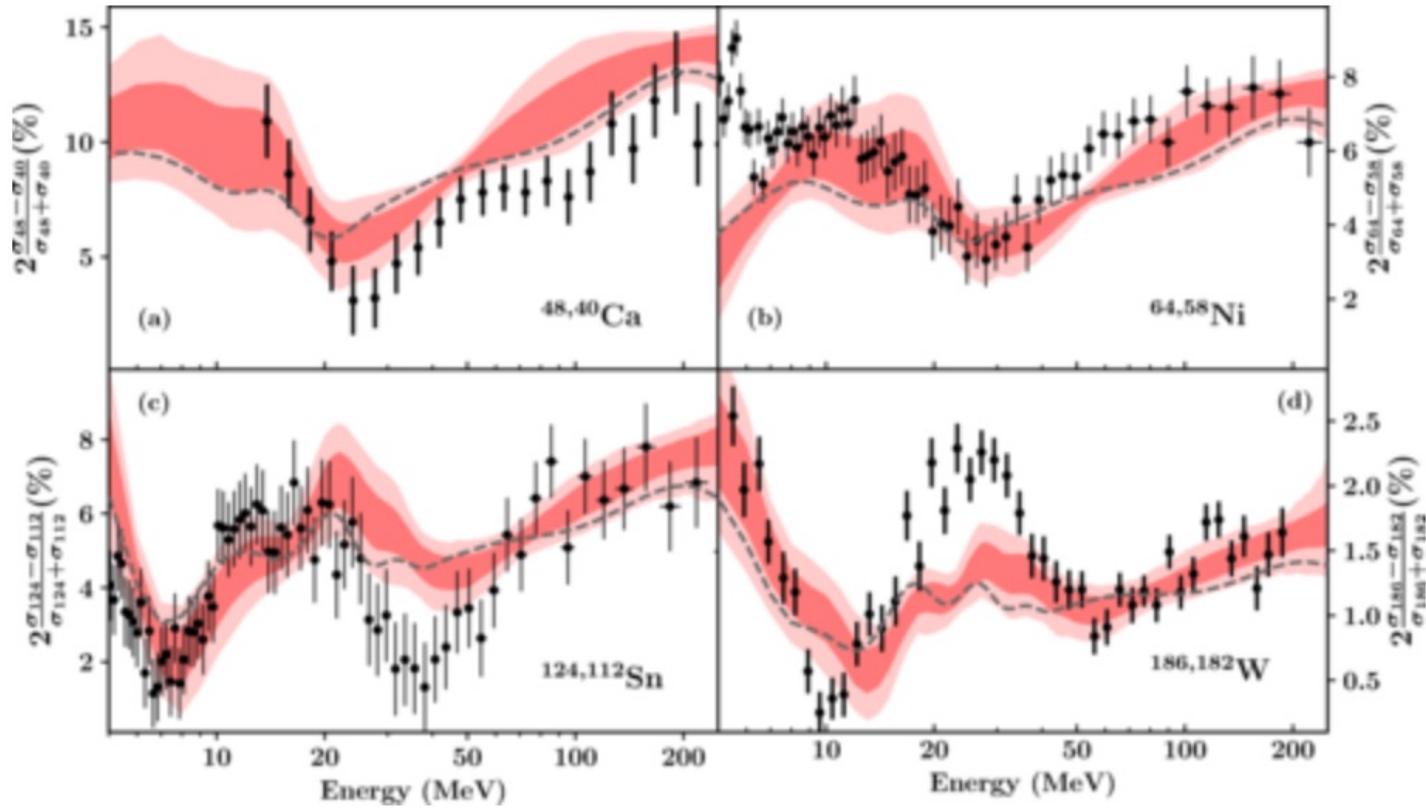
$$\langle \Gamma_\gamma \rangle_0(Z, A) = c_0 + c_1 A^2 + c_2 Z^2 + c_3 Z \bmod(A, 2)$$



Neutron Transmission Coefficients



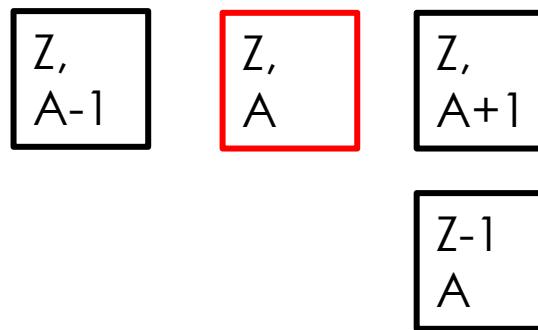
- Pruitt & Escher have re-evaluated the parameters of the Koning-Delaroche optical model
- Including MCMC UQ
- We consider a subset of 50 samples (out 400+) of pre-calculated neutron transmission coefficients





Sampling Procedure

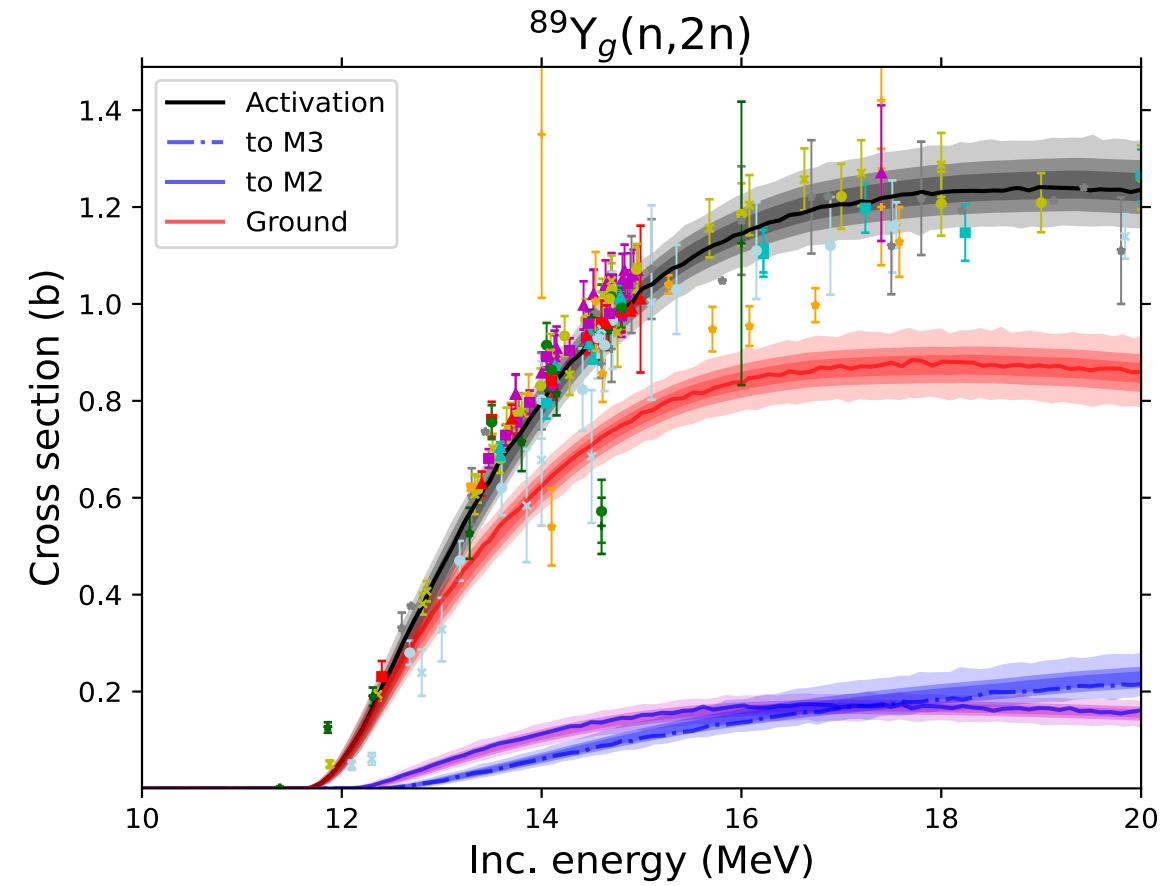
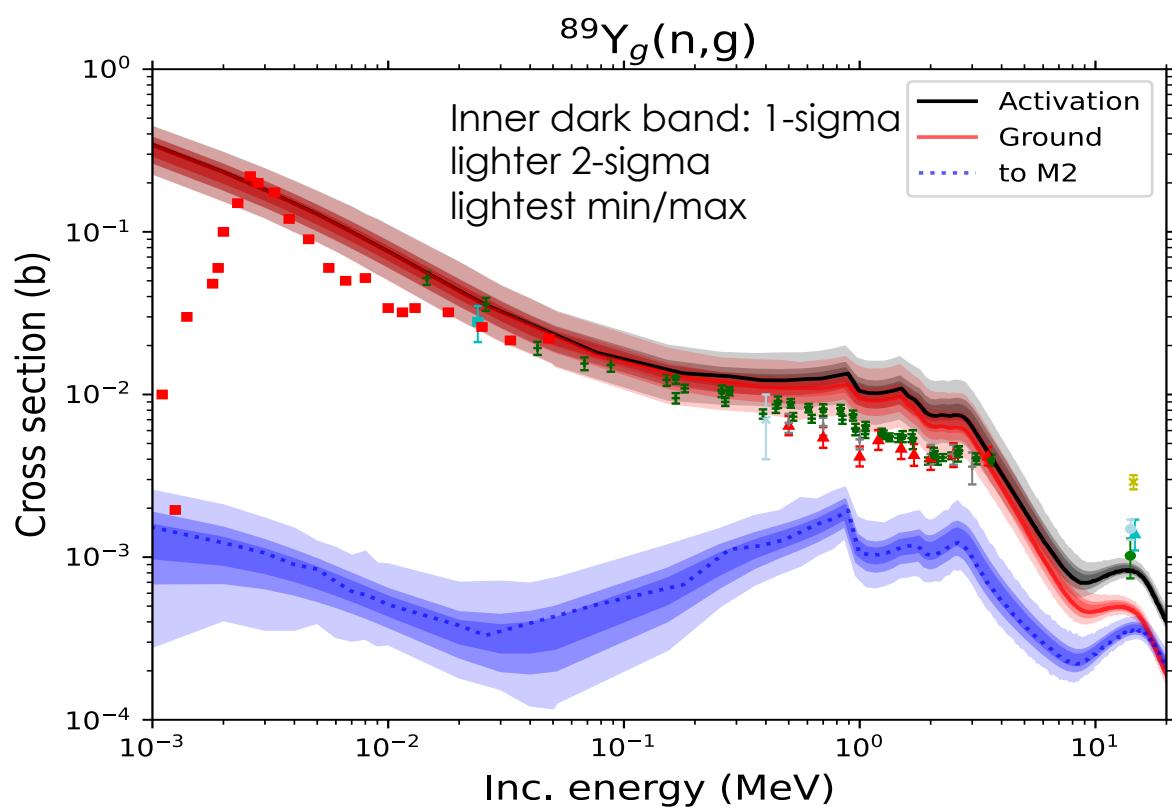
- For neutron-induced reactions on target (Z, A), vary of δW , $\langle \Gamma_\gamma \rangle_0$ parameters **independently** (for now) for three compound nuclei
- Variation of the OMP is **correlated** between all compound nuclei



- Using the LLNL developed COMMCAS framework by Oliver Gorton (using python emcee)
- Draws random samples from specified distributions and sets up input decks, then submits these for parallel on LC machines.
- About 15,000 calculations for each target nucleus.

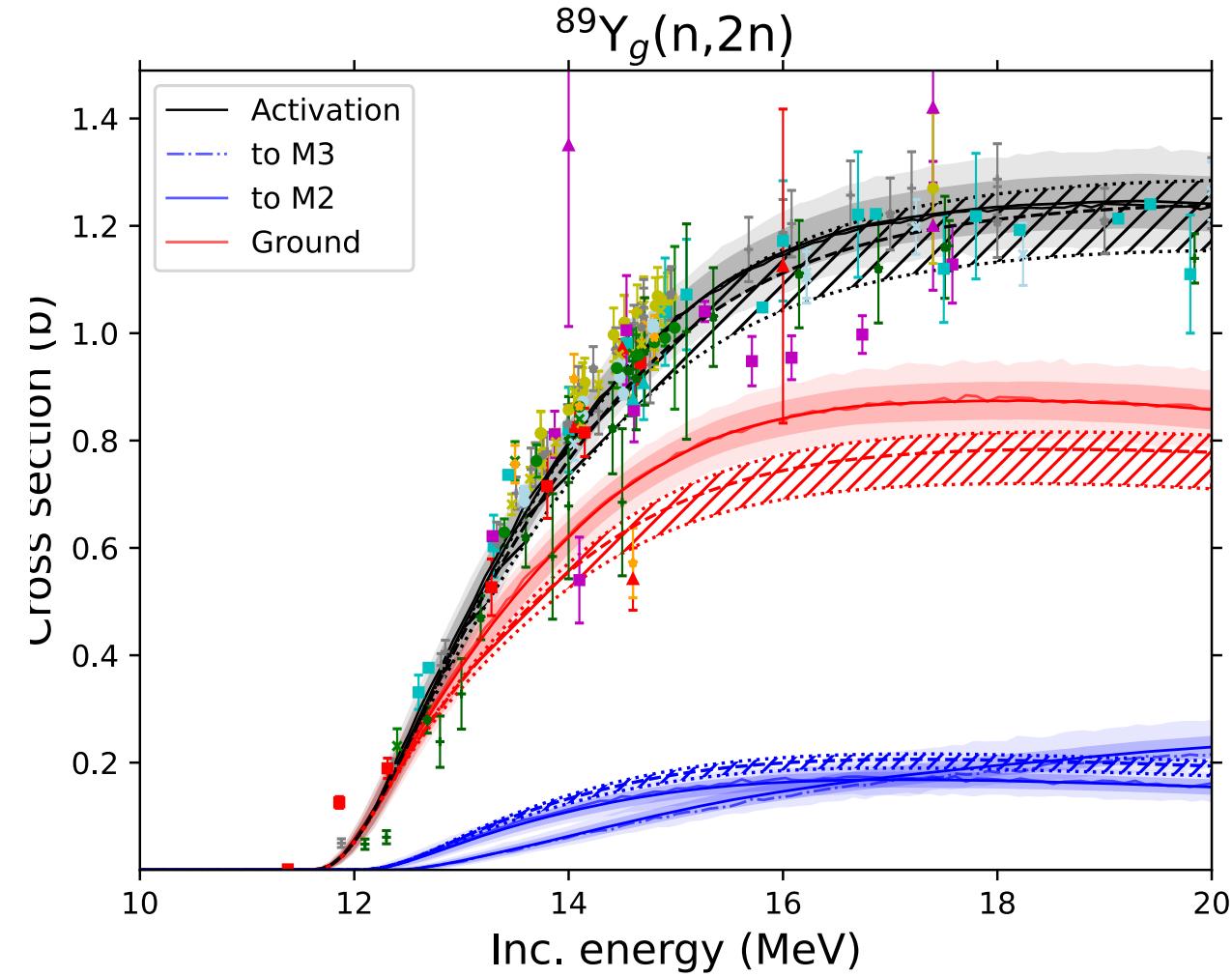
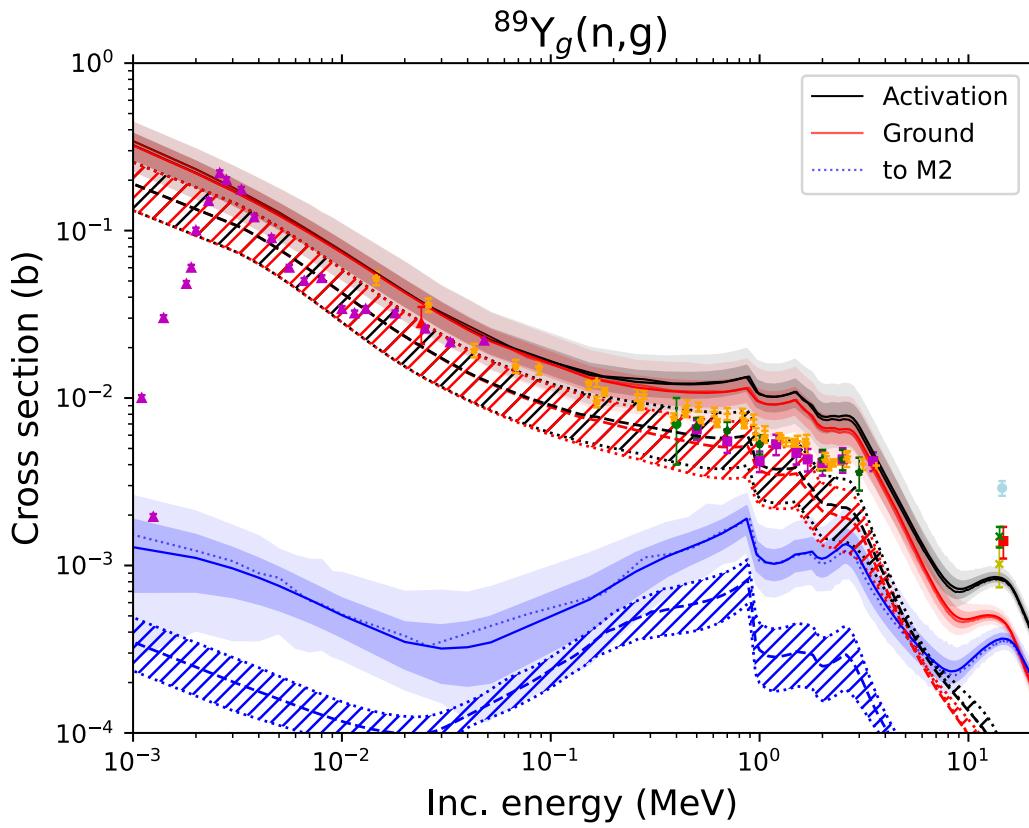
Validation with measured cross sections

Calculations use the parameter systematics:
No fit to the measured cross section



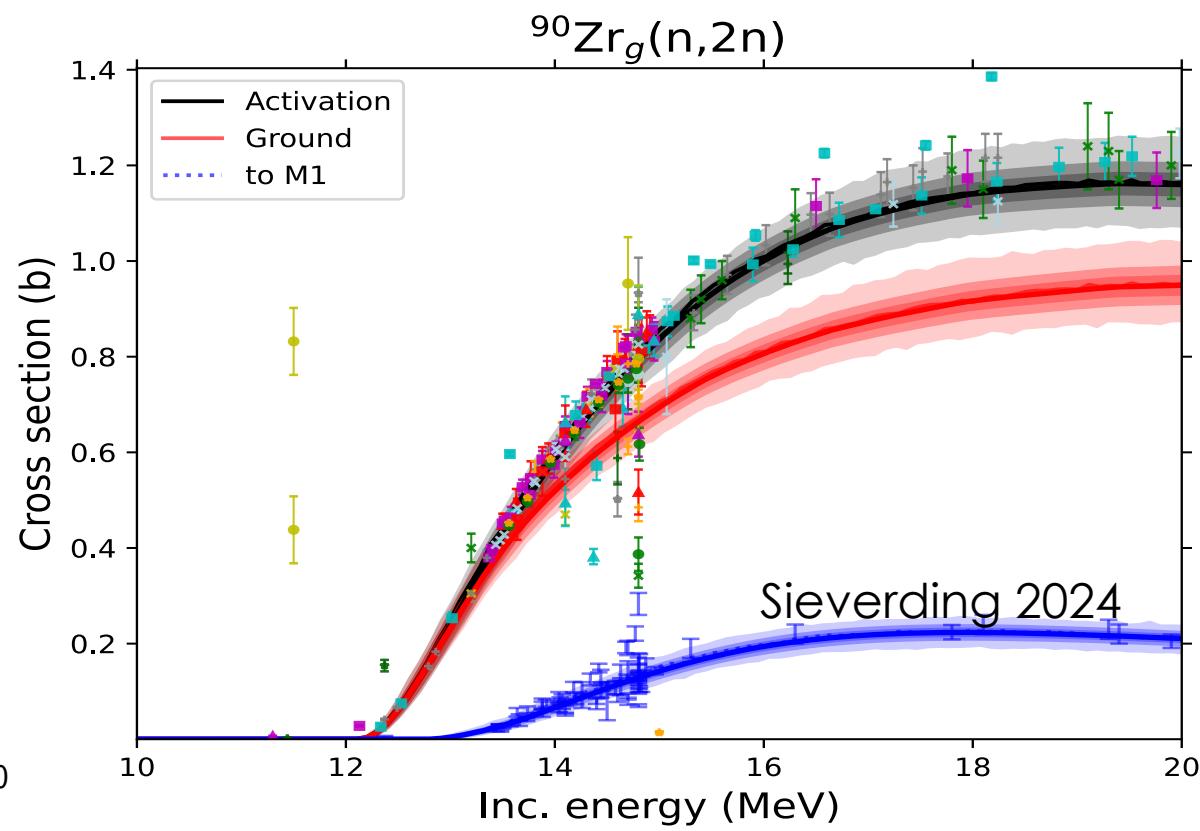
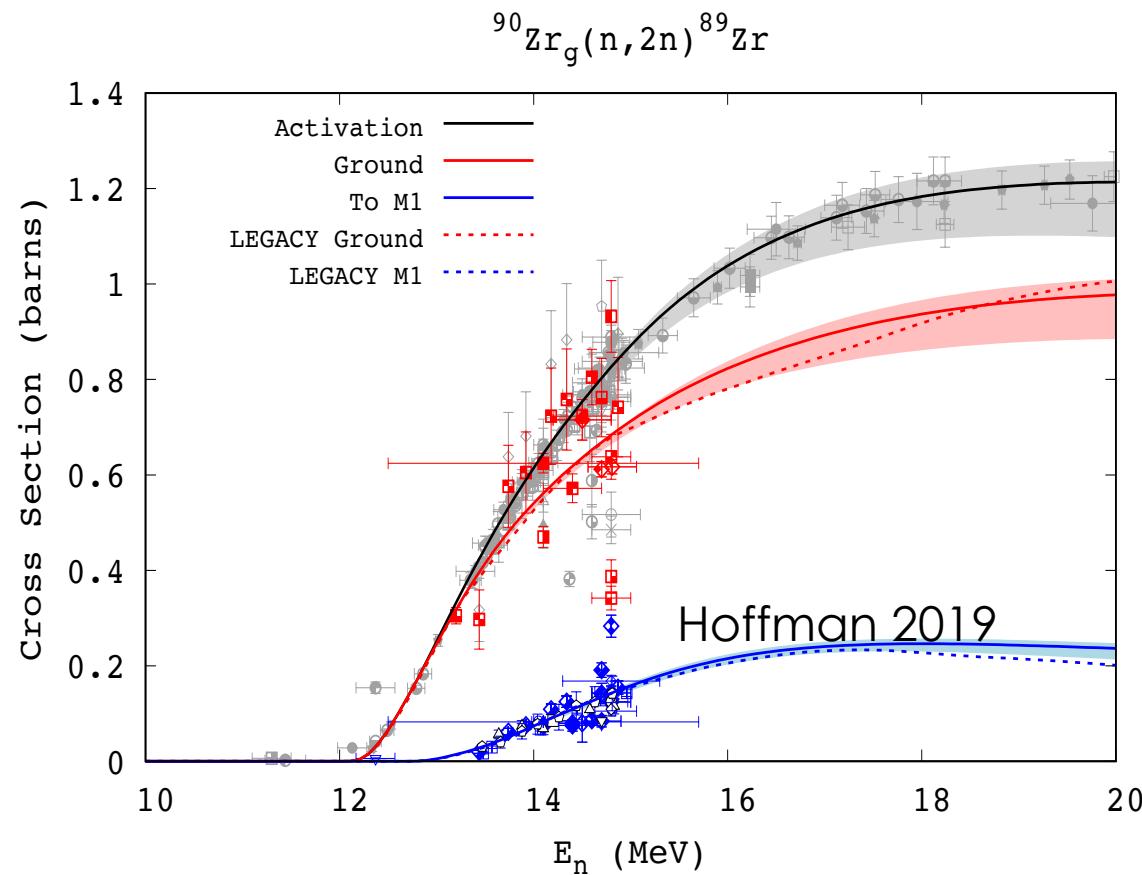
Comparison to previous evaluation

Hoffman et al. 2019, LLNL-TR-737934



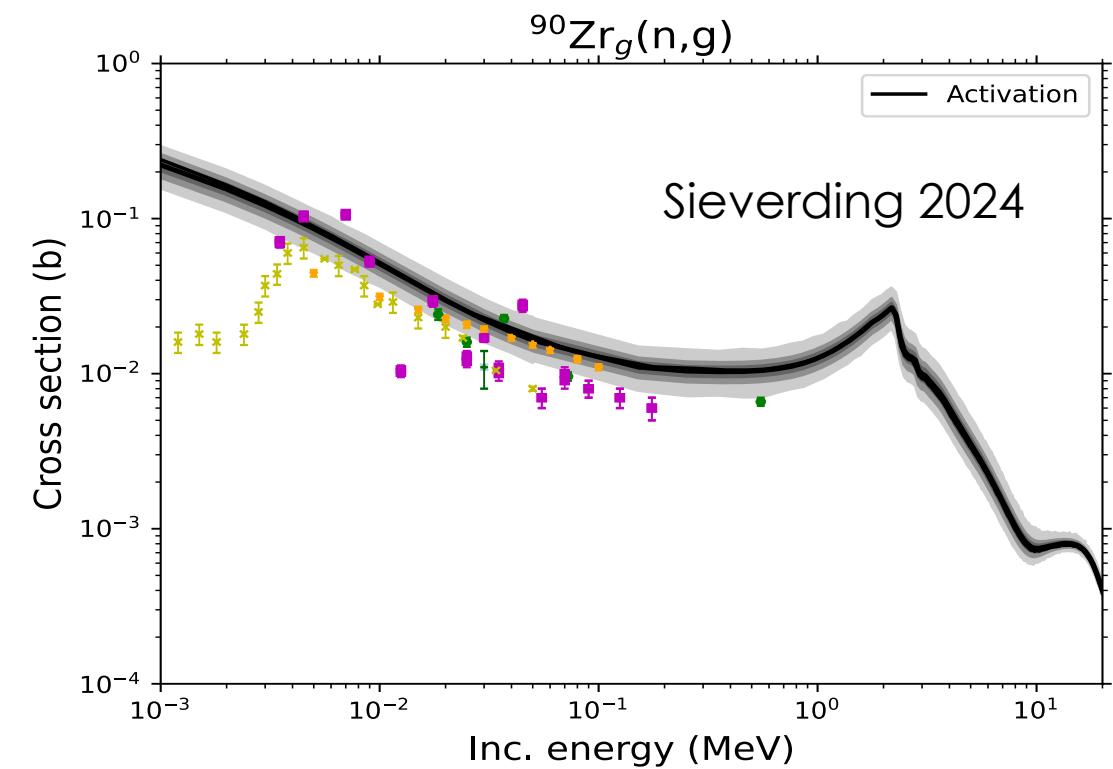
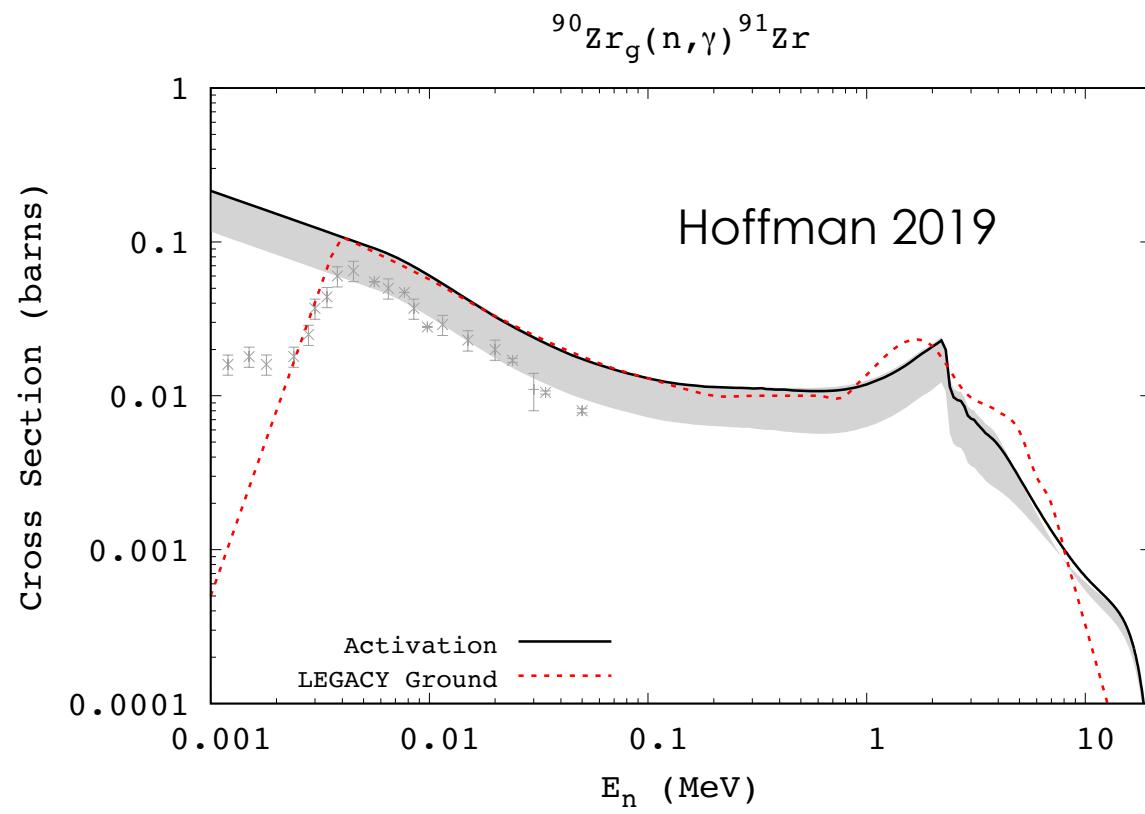
Cross Section comparison: vs UQ

Inner dark band: 1-sigma, lighter 2-sigma, lightest min/max

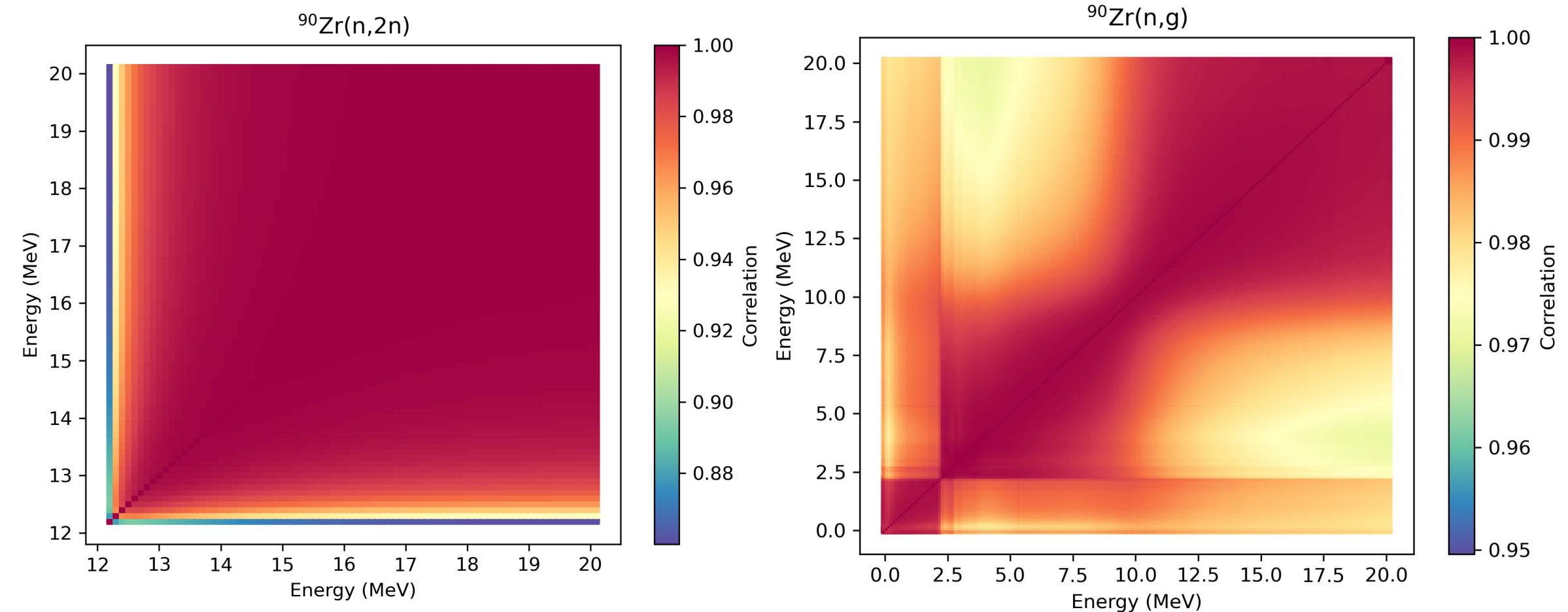


Cross Section comparison: HILO vs UQ

Inner dark band: 1-sigma, lighter 2-sigma, lightest min/max



MCMC approach provides covariance data



- Reaction model induces very strong correlations across all energies

Summary and outlook

- Developed "regional systematics" based on MCMC fits of model parameters to a limited dataset
- The relevant parameters are specific to a region
- Allows systematic UQ including covariance information
- Correlations between cross sections and channels could be extracted
- Reaction model and parameter choice induce very strong correlations across energies
- Possible improvements with advanced data analysis tools:
 - GP or ML for parameters instead of functional form?
 - Emulators for faster sampling
 - Model mixing for some model choices: e.g. GSF model and NLD model
 - Robust UQ for optical model for deformed nuclei



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