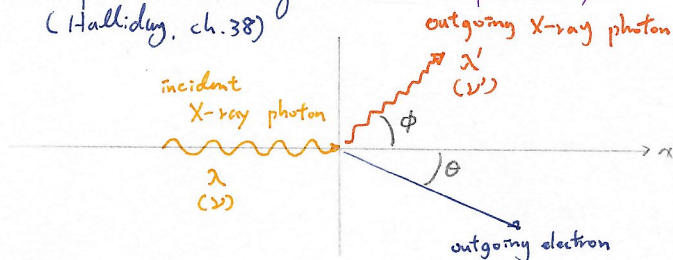


General Physics (II) Quantum Physics

Compton scattering (Arthur Compton 1922)
(Halliday, ch.38)



He also verified Maxwell's prediction that light is EM wave.

discovered by Heinrich Hertz (1887)

photoelectric effect — experimentally implies that the energy of a photon is $h\nu$

Planck constant

Assumption made by Max Planck in 1890s in the derivation of the formula to describe black body radiation.

energy conservation = (rest energy of the electron has been subtracted from the l.h.s. & r.h.s. of the equation)

$$\underbrace{h\nu}_{\text{energy of incoming photon}} = \underbrace{h\nu'}_{\text{energy of outgoing photon}} + \underbrace{mc^2(\gamma - 1)}_{\text{energy of outgoing electron}}$$

$$\gamma = \frac{1}{\sqrt{1 - v^2/c^2}}$$

speed of outgoing electron.

momentum conservation =

$$\text{horizontal: } \frac{h\nu}{c} = \frac{h\nu'}{c} \cos\phi + \gamma m v \cos\theta$$

$$\text{vertical: } 0 = \frac{h\nu'}{c} \sin\phi - \gamma m v \sin\theta$$

There are 3 equations for us to eliminate 3 of the 5 independent variables ($\nu, \nu', v, \phi, \theta$)
Defining $\Delta\lambda \equiv \lambda' - \lambda$, with a little (unimportant) algebra, we got:

$$\underbrace{\Delta\lambda}_{\text{Compton shift}} = \frac{h}{mc} (1 - \cos\phi)$$

if we place detectors at various outgoing angle ϕ , this theory predict that we should detect the scattered light at different wavelength.

Note that this theory is only true when the hypothesis of photons is true, which is distinct from the prediction of the classical EM wave propagation.

photons behave like particles. This is a problem!

Can particles behave like wave?

de Broglie (1923) : $\lambda = \frac{h}{|\vec{p}|}$
(wavelength)

verified by

L.H. Germer & C.J. Davisson (1927)
G.P. Thomson.

利用 X-ray Bragg diffraction 測量 crystal 晶面间距. 所用電子束驗證電子的 diffraction pattern

粒子: 侷限於小的空間中, 具有一定的動量與能量.
其位置與速度隨時間之變化由 equation of motion 決定.
波: 彌漫於空間中之物理量, 由波函數 $\psi(r, t)$ 描述.
波函數隨時間的演化由 wave equation 決定.

Experiments that show photons behave like particles.

1. Compton scattering (Arthur Compton 1922)

2. Blackbody radiation (the earliest development about QM, since 1859)
↳ even before we know the light is EM wave.

(i) Before 1900, Gustav Kirchhoff

Argued based on thermal physics, that the intensity of blackbody radiation only depends on the temperature T and frequency ν .

(ii) mid-1880s, { Ludwig Boltzmann
Josef Stefan

Argued that the energy density of blackbody emission u [Joule cm^{-3}] is proportional to intensity. In addition, its value is

$$u = \sigma T^4 \quad \text{Stefan-Boltzmann law.}$$

↳ Stefan-Boltzmann constant

(iii) Wilhelm Wien

Attempted to derive the blackbody intensity based on statistical physics, assuming that the photons are similar to classical gas particles.

He obtained the Wien's law which is consistent with observations at short wavelength

$$I_{\text{Wien}}(\lambda, T) = \frac{b}{\lambda^5} \exp(-a/\lambda T)$$

↳ constants to be determined experimentally.

(iv) Early 1900, Max Planck.

Assumed the walls in a blackbody chamber can only emit and absorb EM wave of frequency ν by energy in unit of $h\nu$

such energy units were named "quanta" by Planck

Based on this assumption, Planck obtained the correct form of blackbody intensity by applying methods of statistical physics.

$$I(\lambda, T) = \frac{b}{\lambda^5} \frac{1}{\exp(a/\lambda T) - 1}$$

(v) Einstein (1905).

At a certain frequency, the energy of EM wave can only be integer times $h\nu$. This predicts the photoelectric effect (光电效应) that can be tested by further experiments.

诠释量子效应与 h 为有限大之结果。若 $h \rightarrow 0$, 即光子能量单位可为无穷小, 则 blackbody radiation formula 趋近于 Rayleigh-Jeans 推导出的古典型式, 在短波长处则发散。

3. Photoelectric effect (Heinrich Hertz, then Robert Millikan in 1904-1913)

电子仅能吸收单一光子

classical expectation { 1. Accumulating the incoming radiative energy with time can eventually excite photoelectron.
2. There can be a time lag between illumination and the ejection of photoelectron.

actual case: { 1. Capability of exciting photoelectron only relates to frequency, not intensity. Kinetic energy of photoelectron depends on frequency and the surface work function.
2. Ejection of photoelectron is instantaneous.

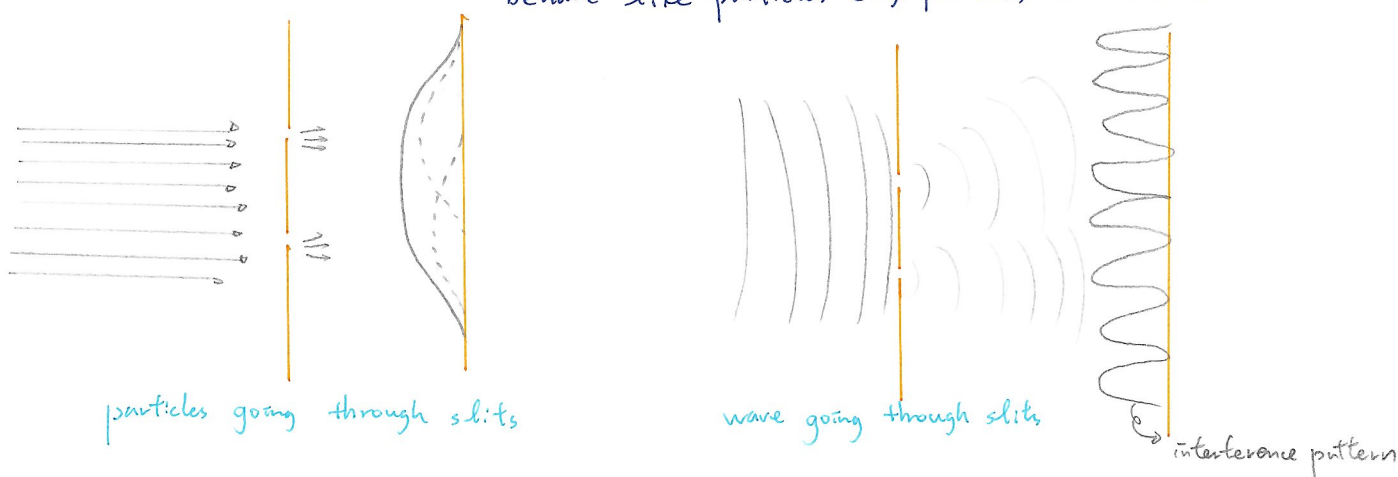
General Physics (II) Quantum Physics.

(此處講一下光的雙狹縫實驗, Halliday ch. 35-36)

Problem: if photon behave like particles, then how we understand the interference of light in the double slit experiments?

- (i) the interference pattern emerge even when there is only one photon at a time
- (ii) when we try to change the apparatus such that we can know which slit the photon is going through, the interference pattern disappear.

The way we measure somehow 'decide' whether light behave like particles (e.g. photons) or waves.



Question: can particle behaves like wave?

Louis de Broglie: matter has wave-like property.
The wavelength is

$$\lambda = \frac{h}{|p|} \quad \text{de Broglie wave-length.}$$

- (i) { C.J. Davisson & L.H. Germer
G.P. Thomson } — observing the interference pattern of electrons scattered by planes in crystals.

- (ii) Tomomura et al. (1989)
slit experiments (see textbook)

Max Born's hypothesis: each photon is associated a wave ψ , called probability amplitude (機率幅)

電場正比於電子出現之機率幅。
機率正比於機率幅平方

In the EM wave, the $|\vec{E}|^2$ determines the probability for a photon to appear.

Can electrons be described by another sort of probability wave?

General Physics (II) Quantum Physics

physical Hilbert space = The space of functions that can be either normalized to unity, or to the Dirac delta function.

Postulate = (Max Born
Erwin Schrödinger 1926)

Each electron is associated with a wave function $\psi(\vec{r}, t)$ that obeys some linear equation. (i.e., Schrödinger equation)
 ↳ linearity is a key to allow interference

a function in the physical Hilbert space, s.t. there is a concept of probability conservation

可對比於光波問題中之電場，但 $\psi(\vec{r}, t)$ 無相對應之古典場

ψ can be a complex function, with a complex conjugate ψ^* .

The probability of finding an electron at time t in an infinitesimal space $d^3\vec{r}$ is $|\psi(\vec{r}, t)|^2 d^3\vec{r}$
 ↳ 機率密度 / 機率

The electron must be somewhere in the space. So

$$\int_{\text{all space}} |\psi(\vec{r}, t)|^2 d^3\vec{r} = 1$$

Particle may be described as a ψ wave pulse that has a finite extension in space.

making partial t derivative on both side $\Rightarrow \int_{\text{all space}} \left(\frac{\partial \psi^*(\vec{r}, t)}{\partial t} \psi(\vec{r}, t) + \psi^*(\vec{r}, t) \frac{\partial \psi(\vec{r}, t)}{\partial t} \right) d^3\vec{r} = 0$

a way to satisfy this equation at all time:

requiring $\frac{\partial \psi}{\partial t}$ to be related to ψ in some way. should still be linear.

related by the "Governing equation", which is the Schrödinger Eq. we are constructing.

建構 Ansatz = ψ 必須以類似電磁波的 sin, cos wave 型式干涉，且的時間微分要為 ψ 的 linear operation

最自然的猜法

先猜 ψ 的型式為 $\cos(kx - \omega t)$, $\sin(kx - \omega t)$ 的線性組合

$$\psi(x, t) = A [\cos(kx - \omega t) + i \sin(kx - \omega t)] = A e^{i(kx - \omega t)}$$

$$\frac{\partial \psi}{\partial t} = -i\omega \psi(x, t)$$

(what are k and ω ?)

角頻率，為頻率的2倍

free electron (自由電子)
(non-relativistic)

plane wave: $k = \frac{2\pi}{\lambda}$ $\Rightarrow k = 2\pi \left(\frac{p}{h} \right) = \frac{p}{\hbar}$, $\hbar \equiv h/2\pi$
 wave number \nearrow
 \nwarrow de Broglie wave length $\lambda = \frac{h}{p}$

analogous to light (EM) wave, we expect the energy of the electron to be related to frequency ω by $E = h\omega = \frac{h}{2\pi} (2\pi\omega) = \hbar\omega \Rightarrow \omega = \frac{E}{\hbar}$

\Rightarrow plane-wave form of the electron wave function

$$\psi(x, t) = \underbrace{A}_{\text{normalization constant}} e^{\frac{i(px - Et)}{\hbar}}$$

For a non-relativistic electron, we expect $E = \frac{1}{2} \frac{p^2}{m}$

$$\Rightarrow \frac{\partial \psi}{\partial t} = -i\omega\psi = -i \frac{E}{\hbar} \psi = -i \frac{p^2}{2m\hbar} \psi$$

Observing that $\frac{\partial^2 \psi}{\partial x^2} = -k^2 \psi = -\frac{p^2}{\hbar^2} \psi = -\frac{2m}{\hbar^2} \left(-i \frac{p^2}{2m\hbar} \right) \psi = -i \frac{2m}{\hbar} \frac{\partial \psi}{\partial t}$

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2}$$

time-independent Schrödinger eqn. for free-electron.

electron in potential

modifying from $E = \frac{1}{2} \frac{p^2}{m}$ to $E = \frac{p^2}{2m} + V$

$$\frac{\partial \psi}{\partial t} = i\omega\psi = -i \frac{E}{\hbar} \psi = -i \frac{1}{\hbar} \left(\frac{p^2}{2m} + V \right) \psi$$

\nearrow following the derivation as the free-electron case

$$\Rightarrow i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V\psi$$