

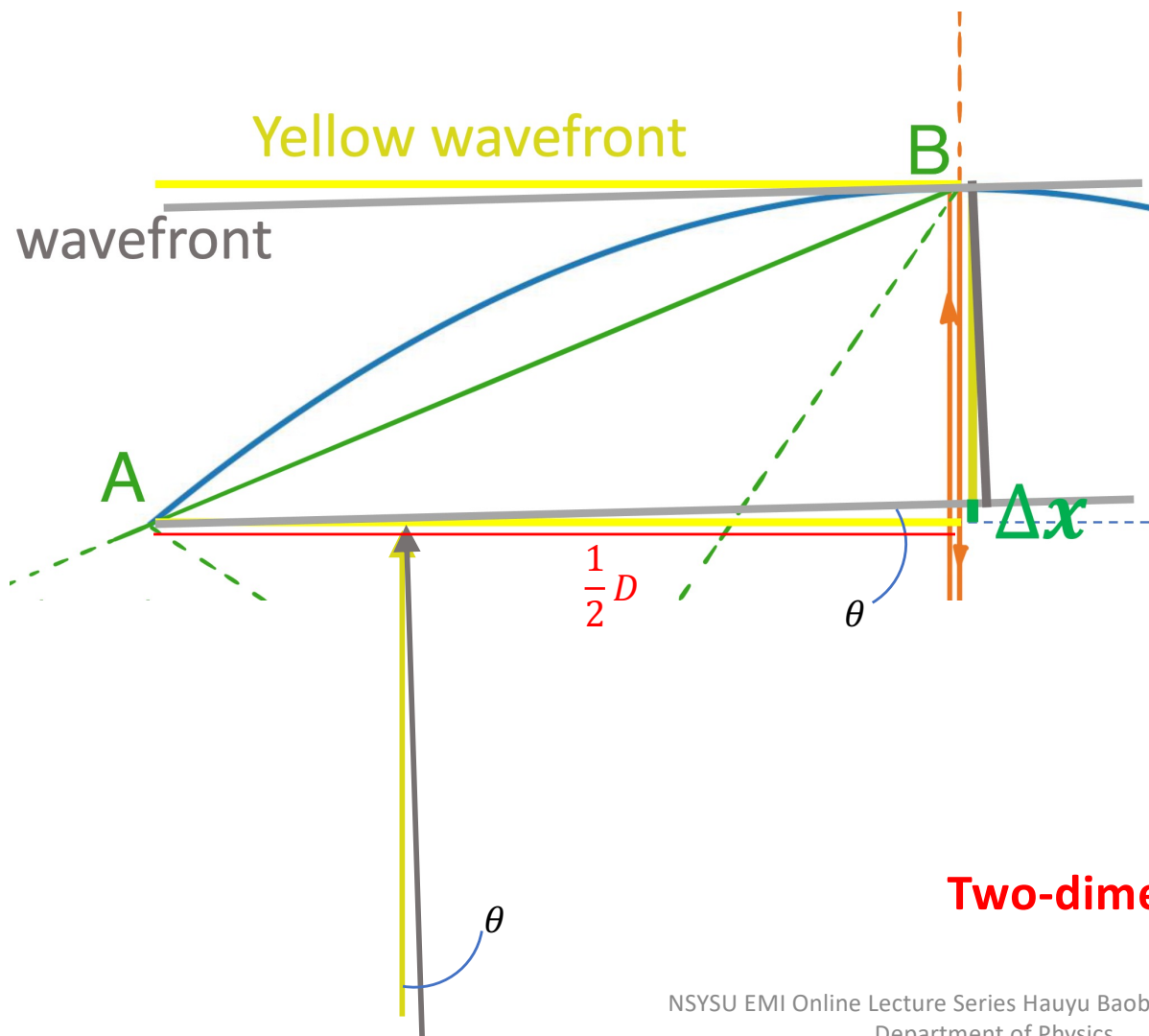
# An Introduction to Radio Interferometry

2-4 Observations with single-dish telescopes



You can find relevant material  
on my personal webpage

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How do we find the first minimum?

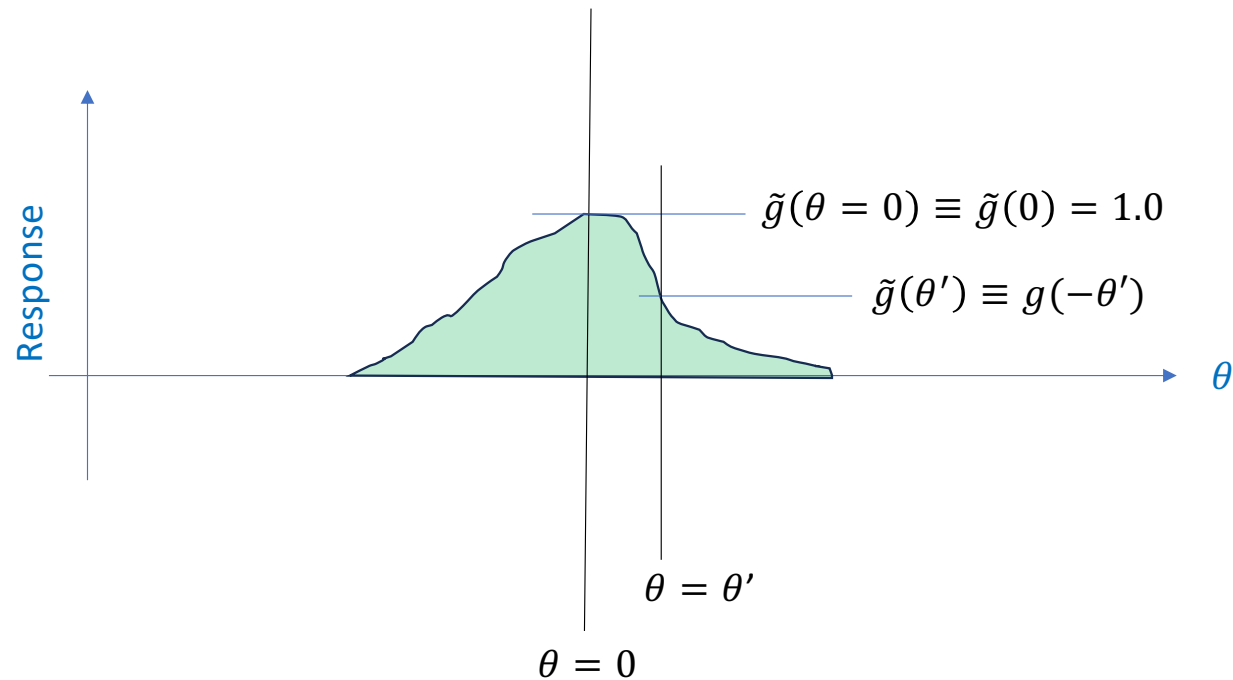
The trick we knew when working on single-slit.

Path length difference:  $\Delta x = x_1 - x_2 \sim \frac{1}{2} D \theta$

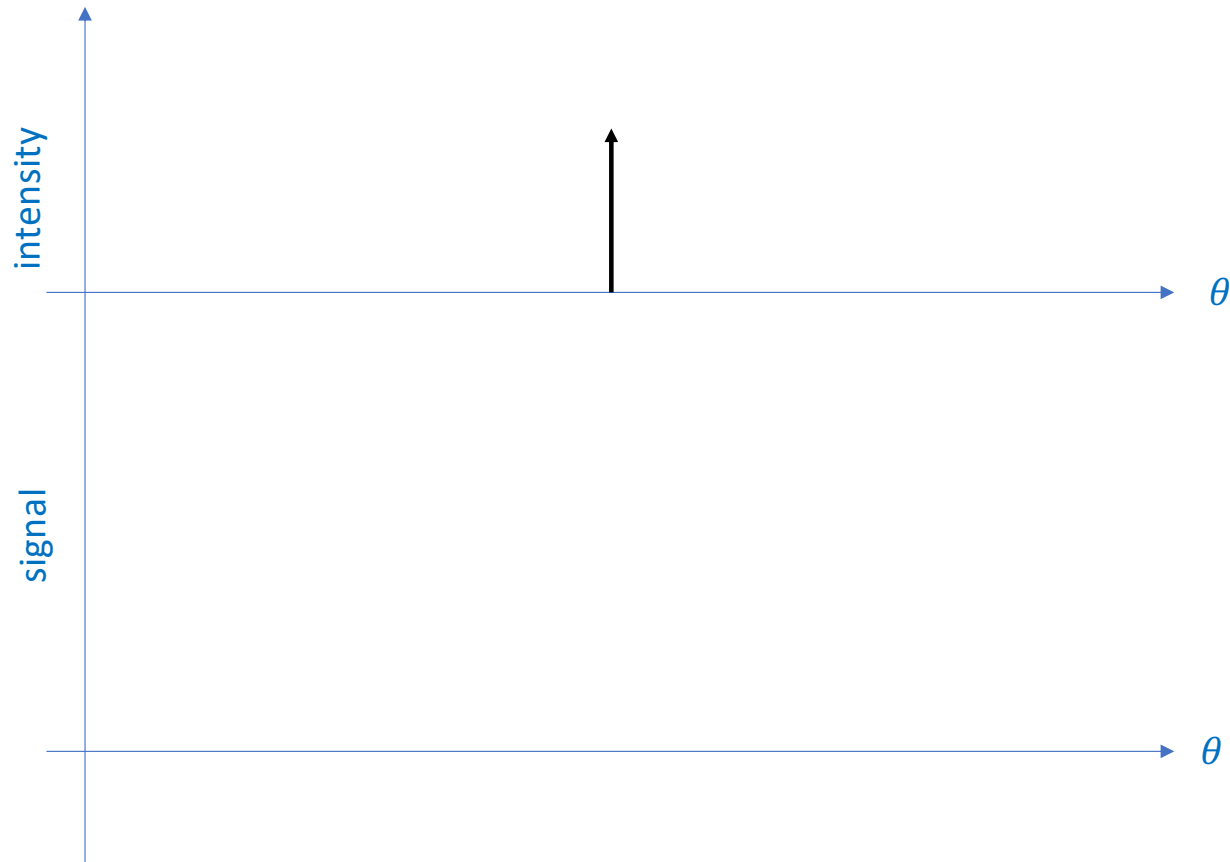
$$\begin{aligned} \text{First zero: } k\Delta x &= \frac{1}{2} k D \theta = \pi \\ \Rightarrow \frac{1}{2} \frac{2\pi}{\lambda} D \theta &= \pi \Rightarrow \theta = \frac{\lambda}{D} \end{aligned}$$

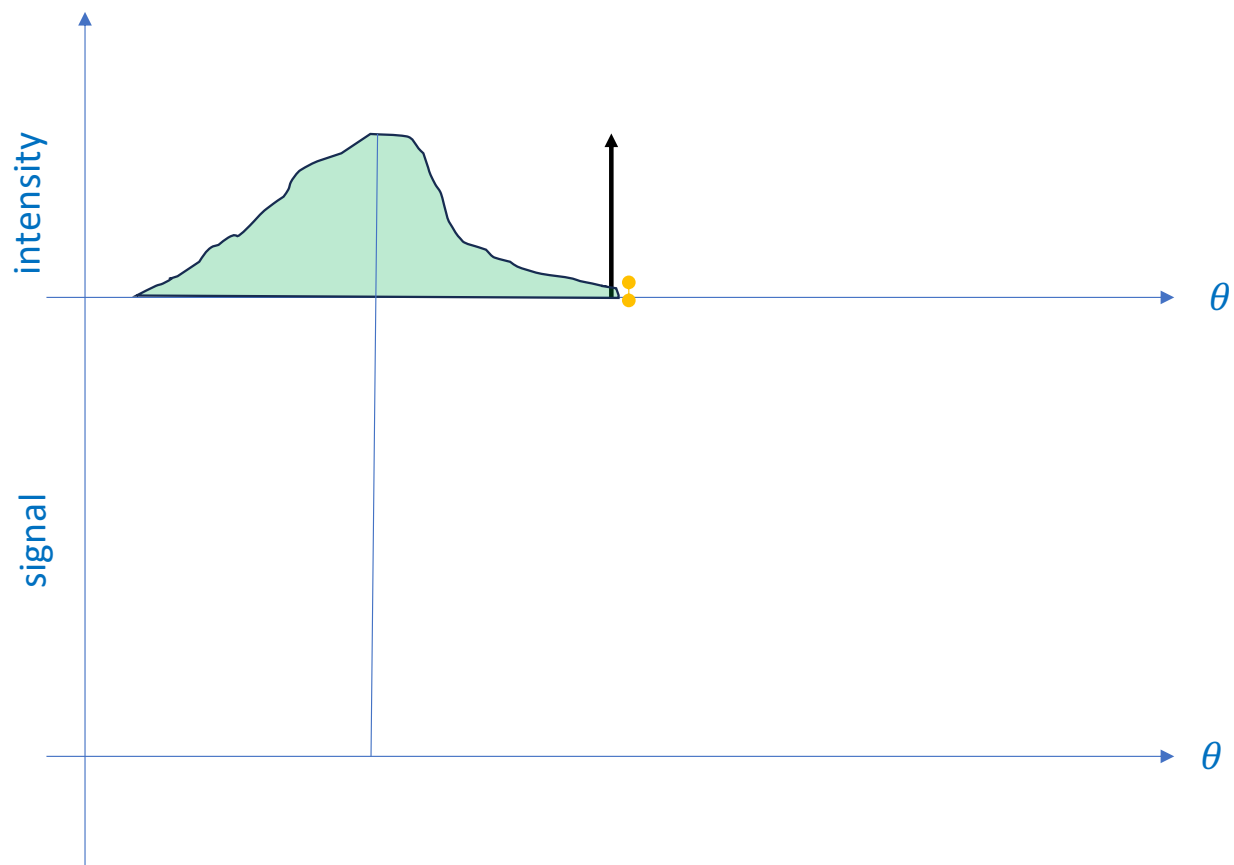
$$\text{Two-dimensional dish: FWHM} = 1.22 \frac{\lambda}{D}$$

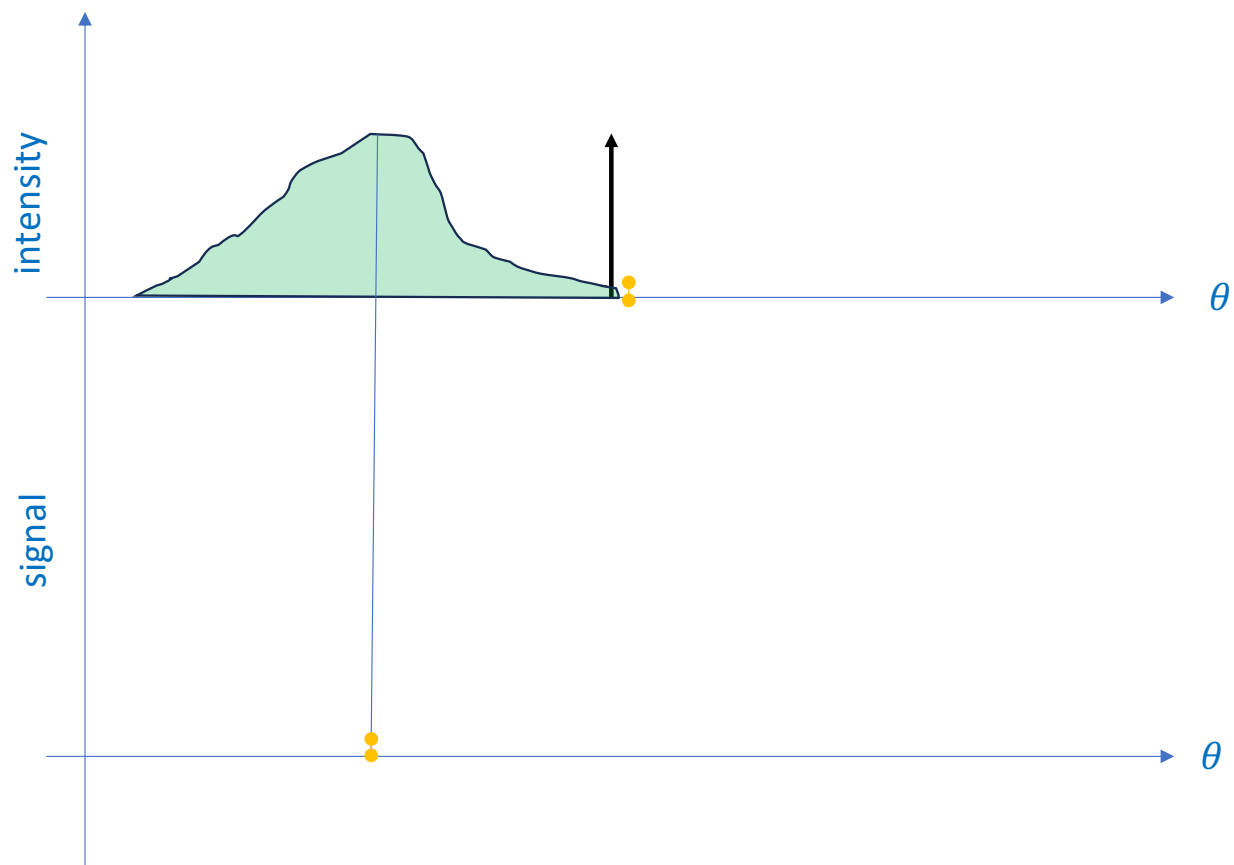
## Response of a (one-dimensional) single dish telescope

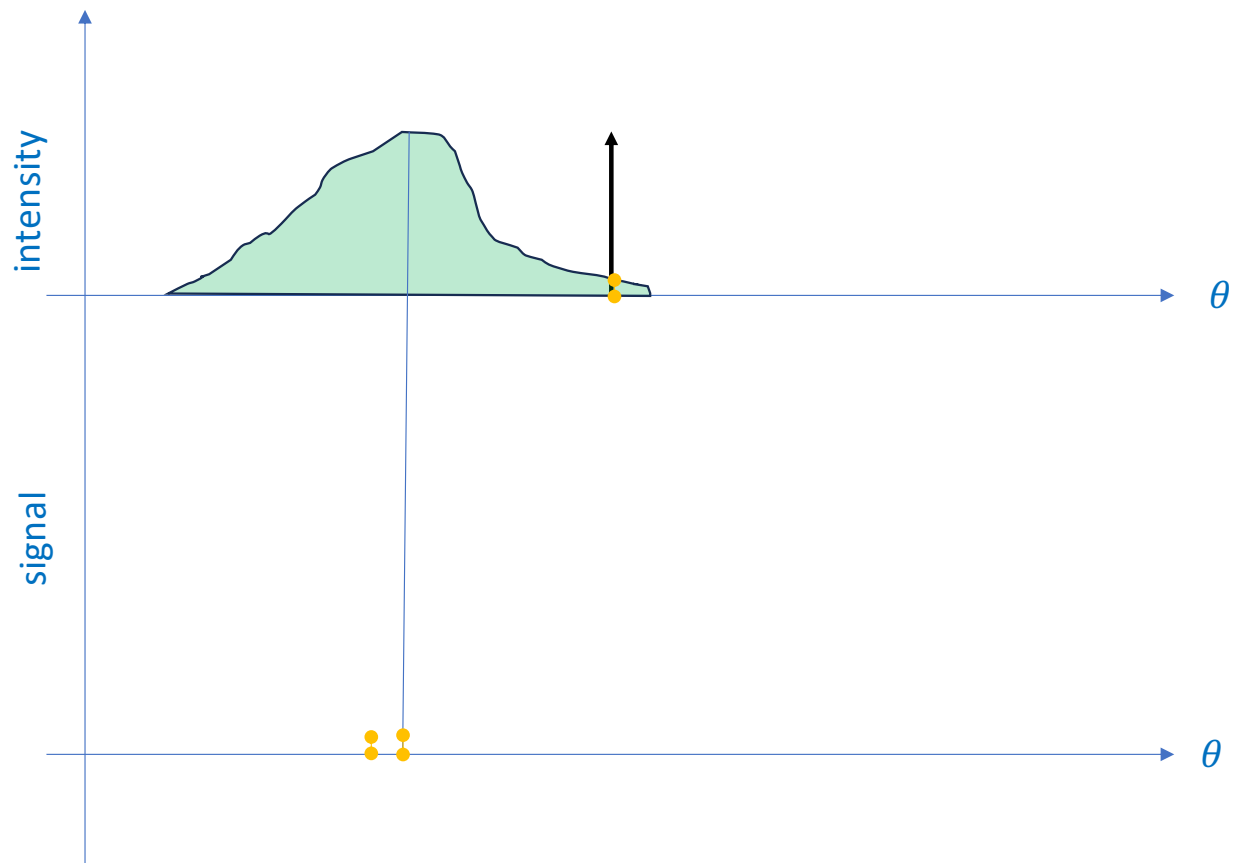


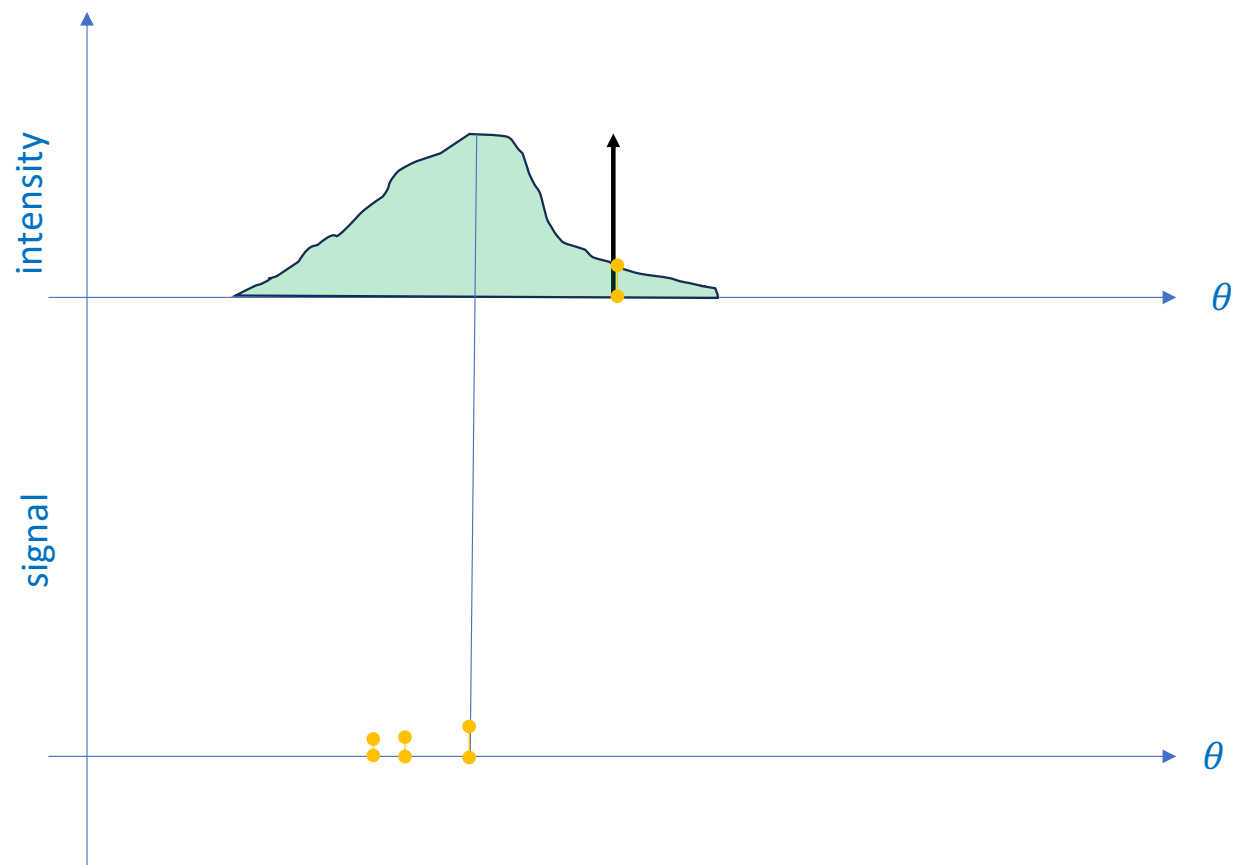
## Flux density distribution of a point-source



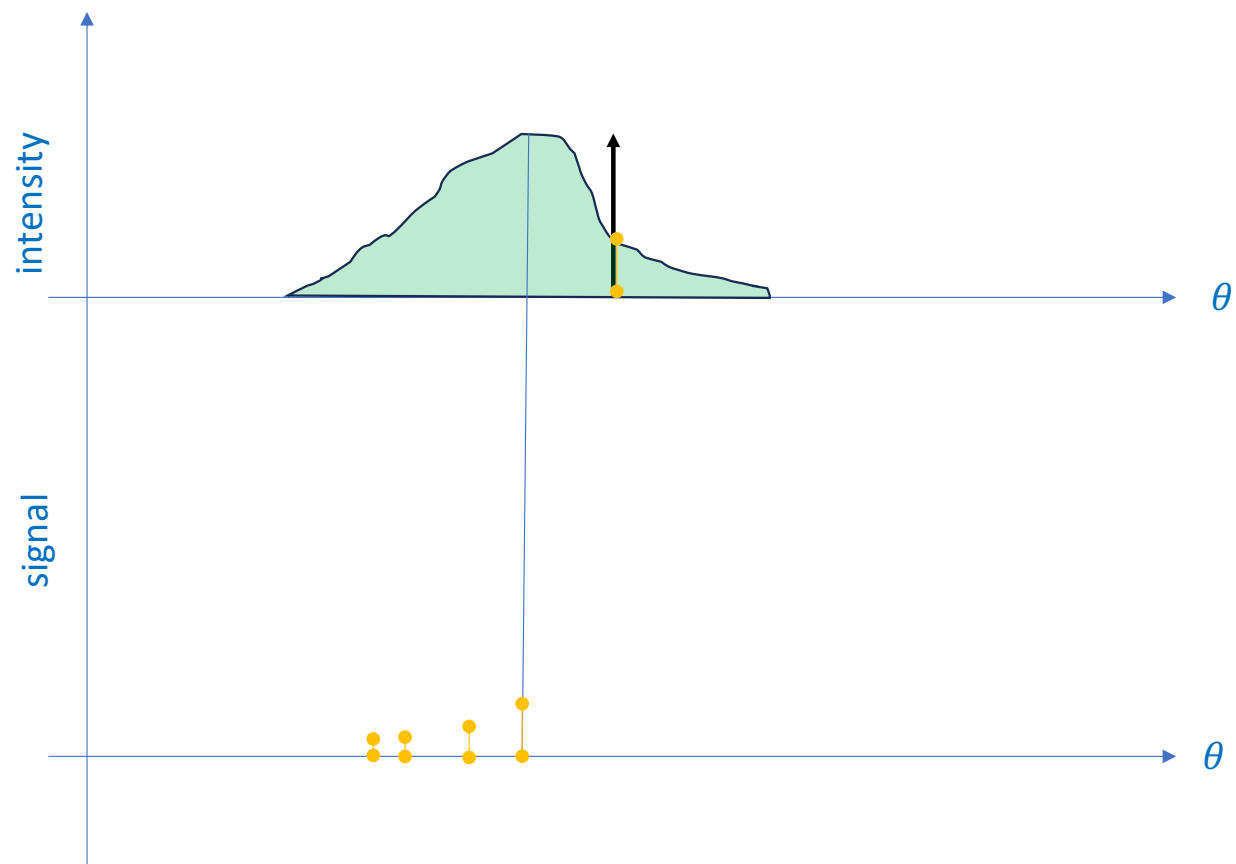


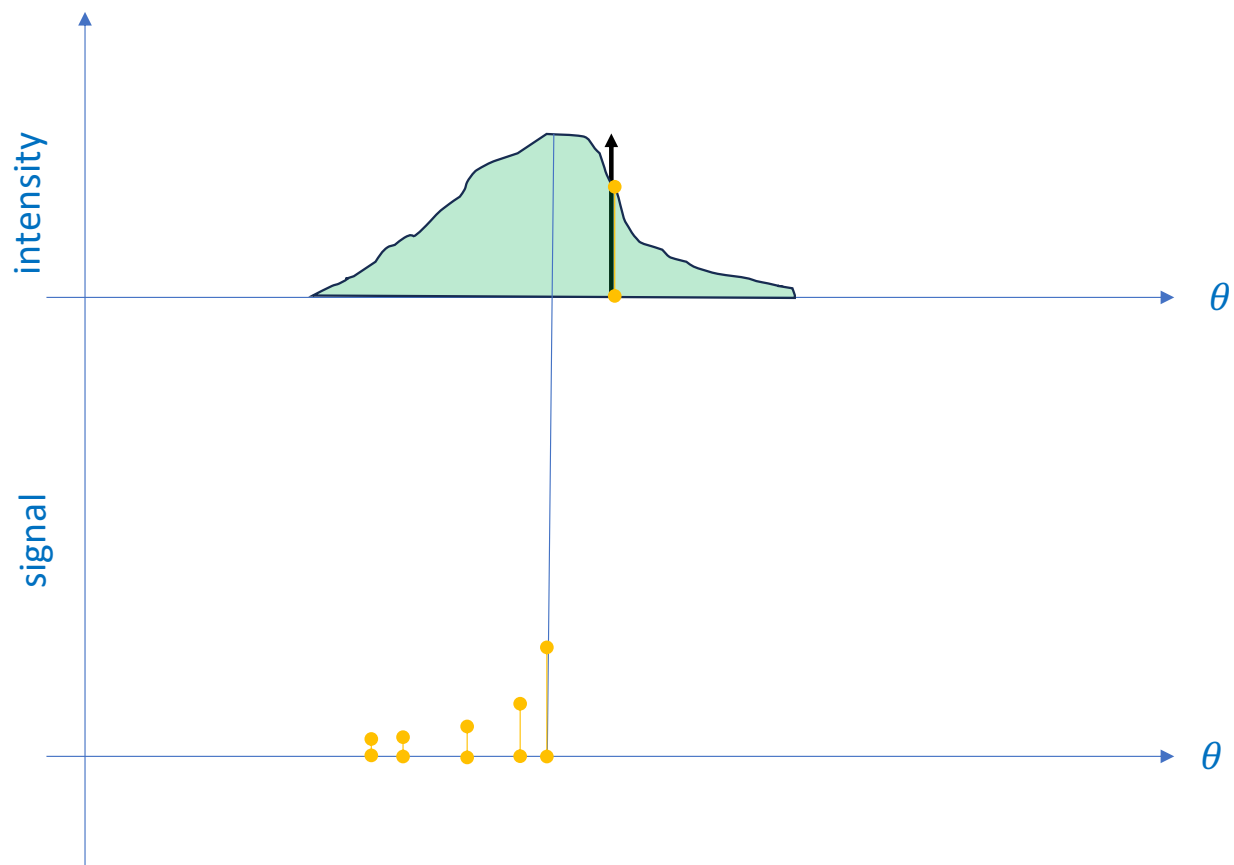


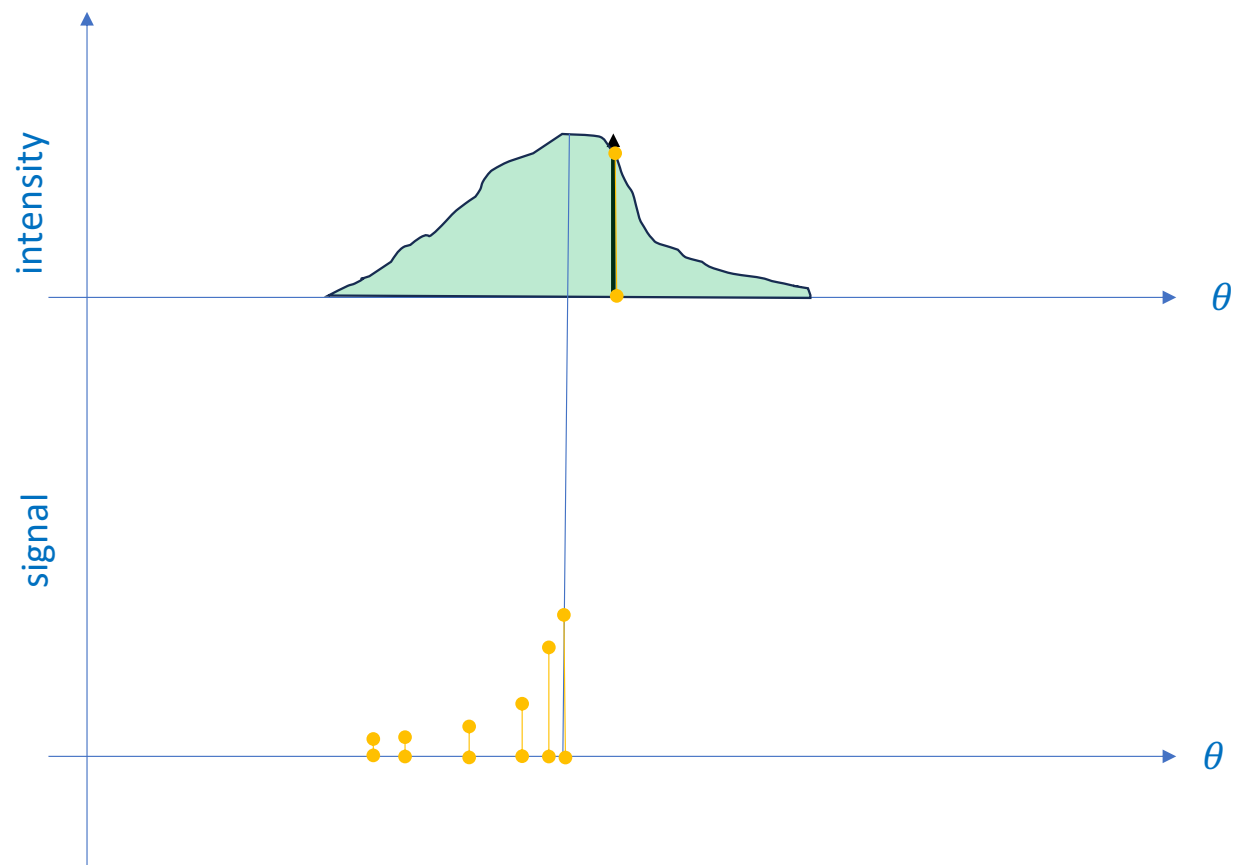


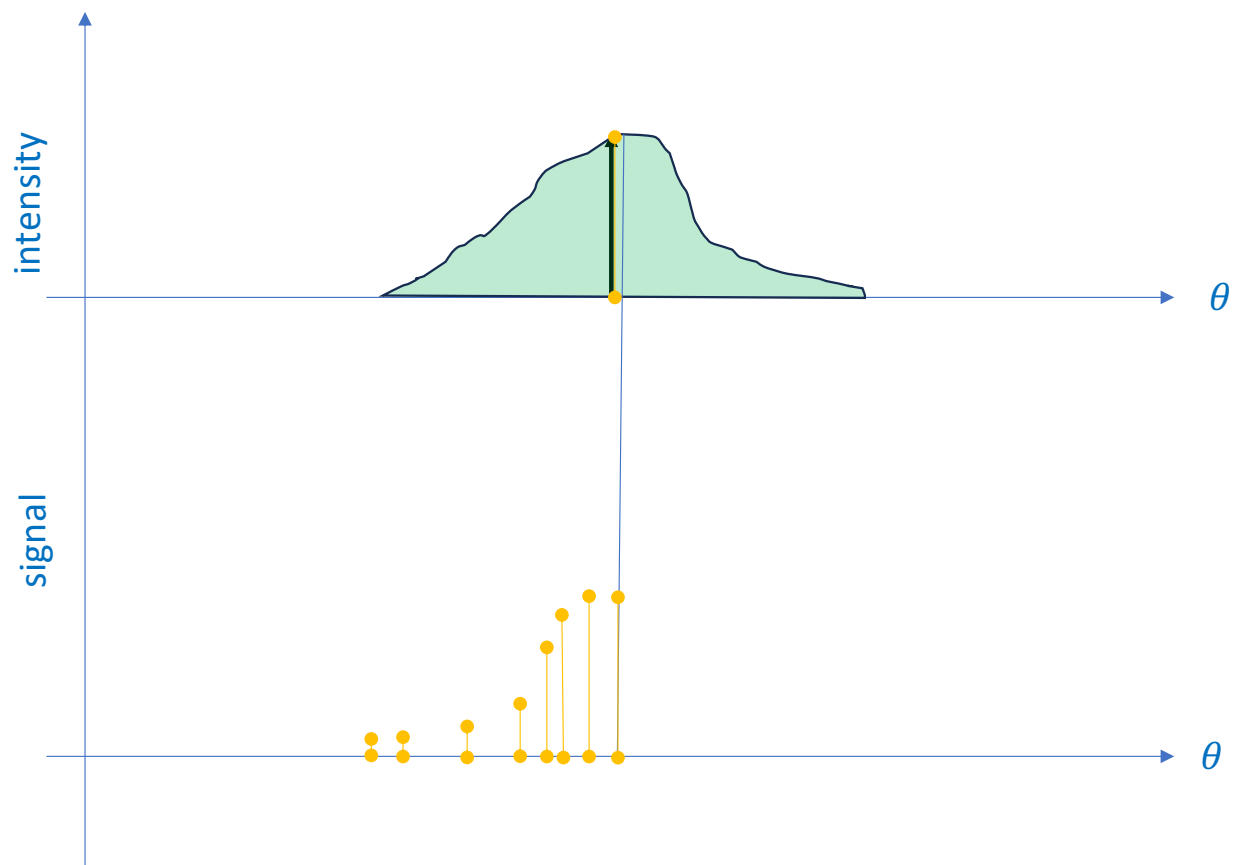


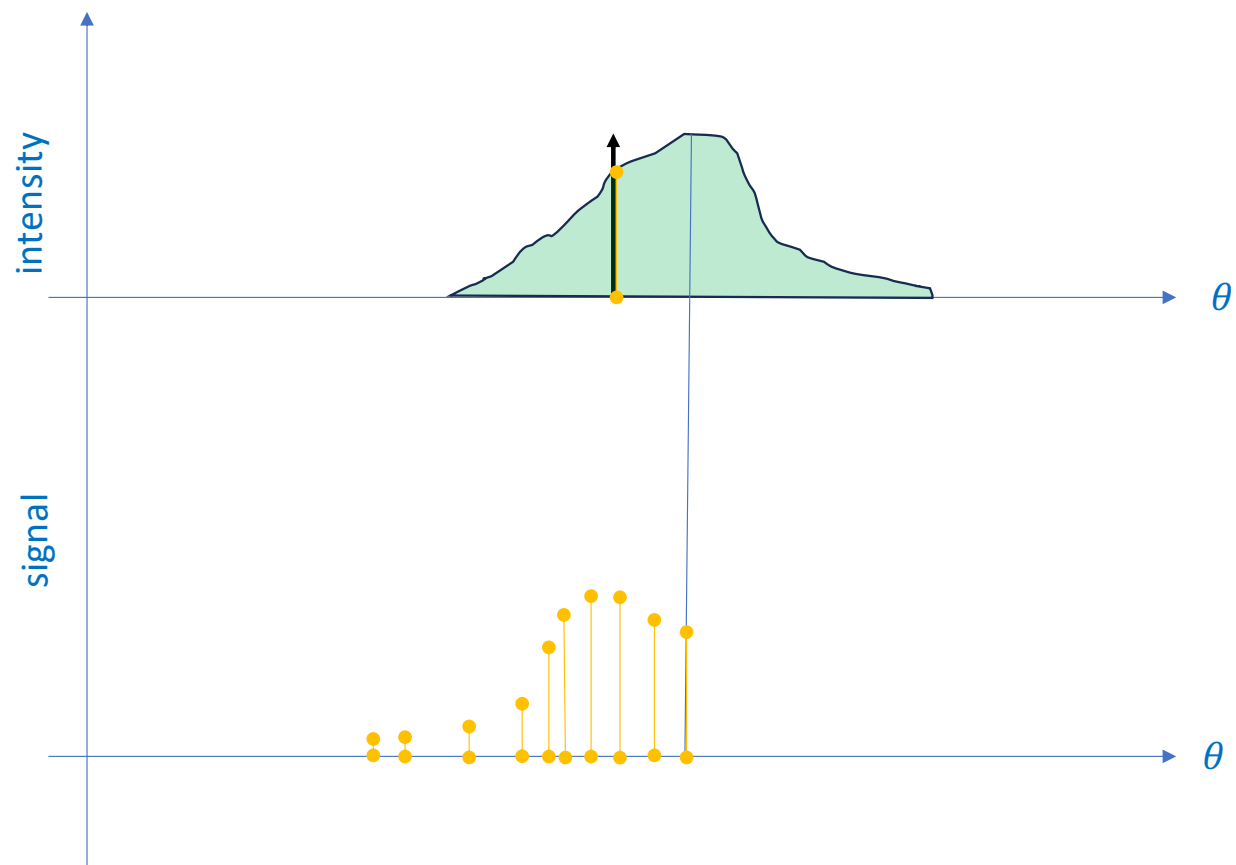


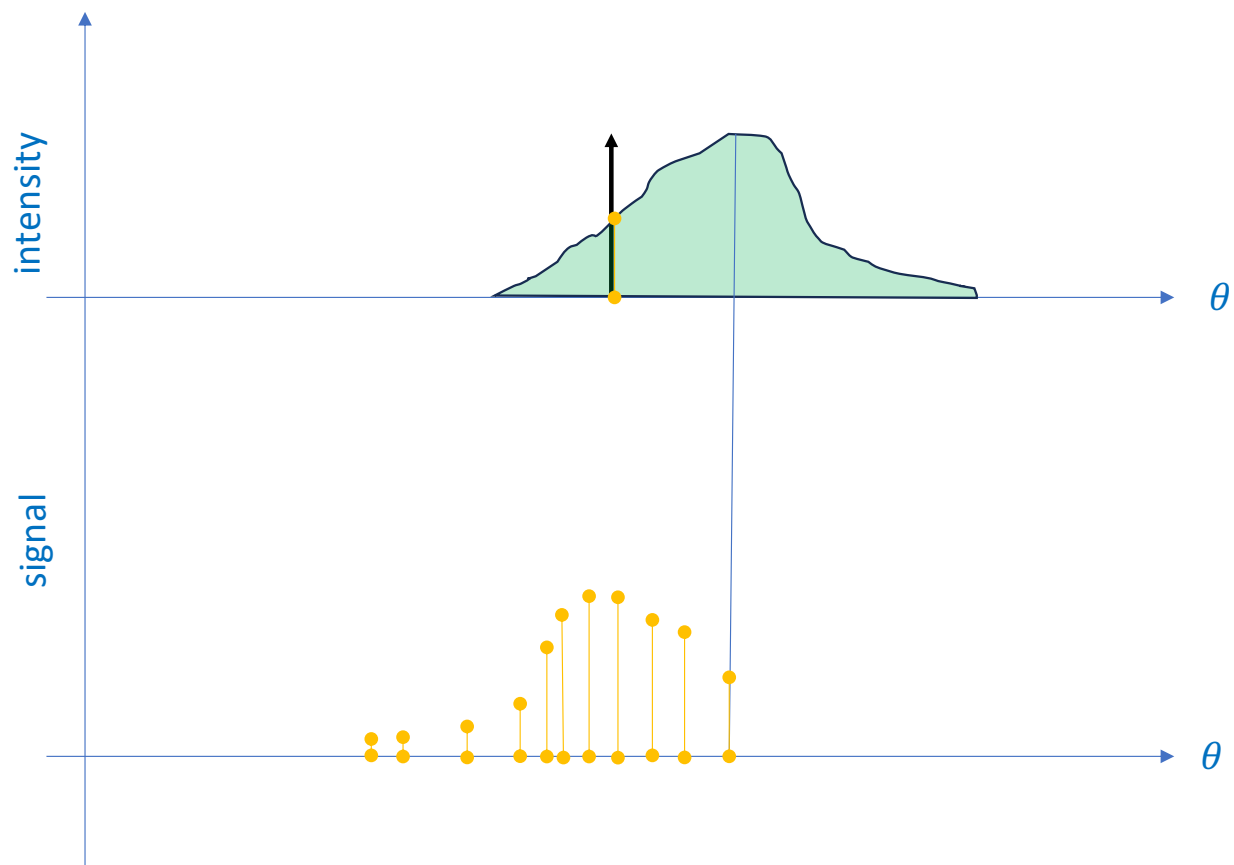


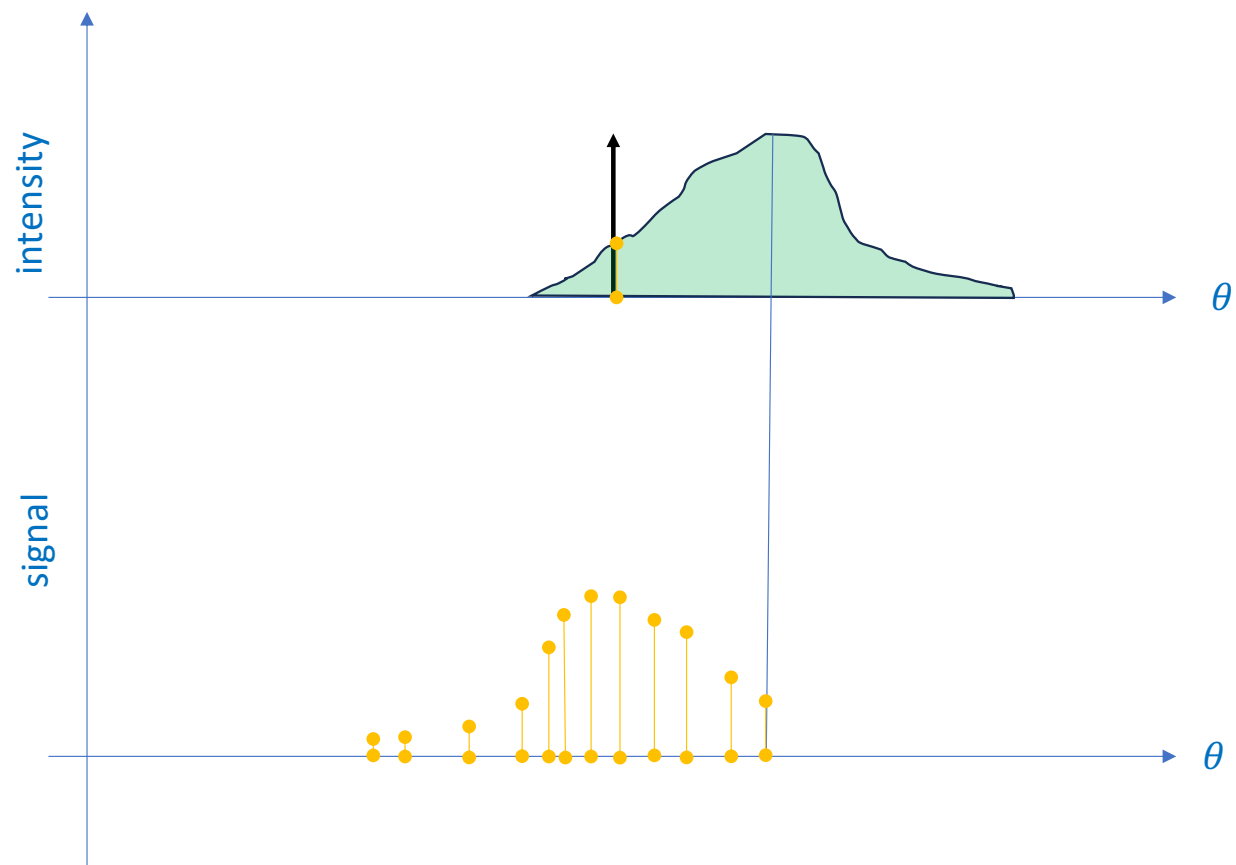


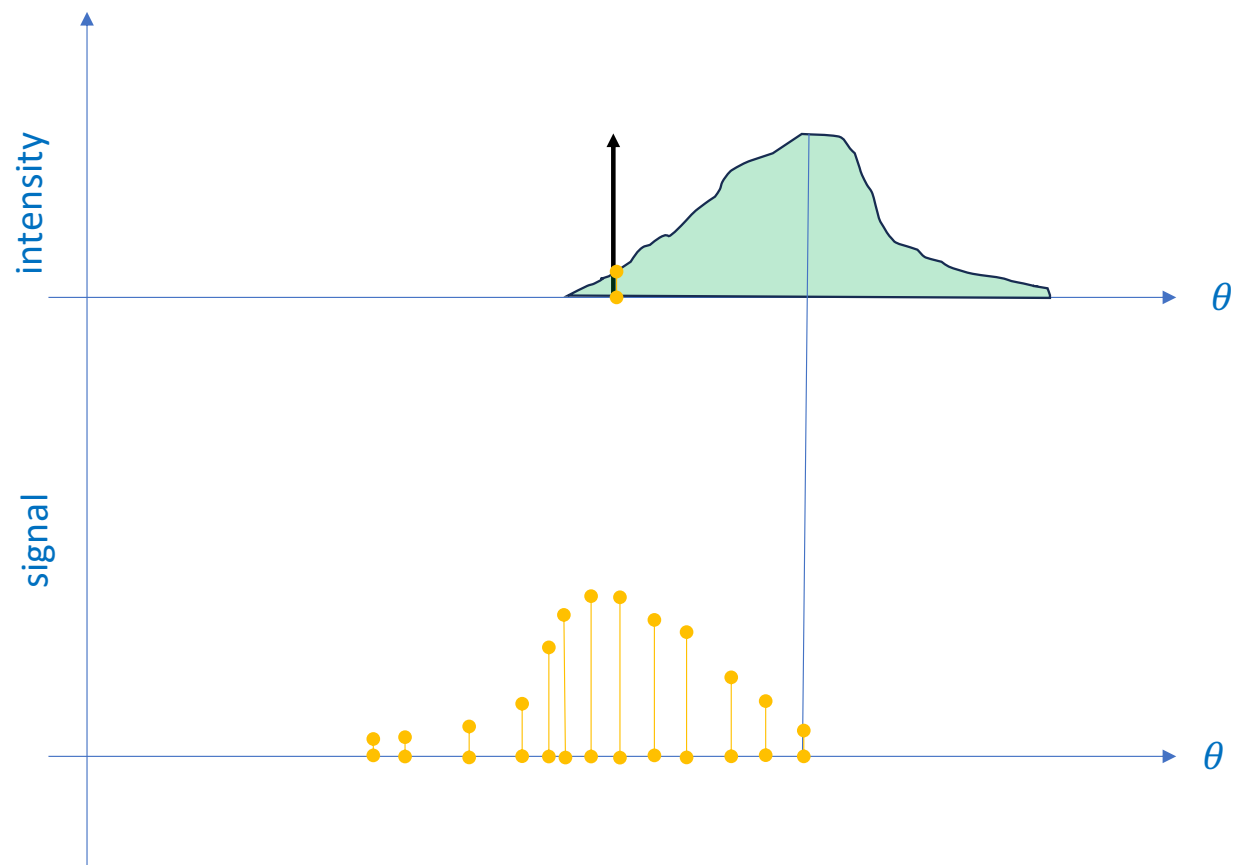




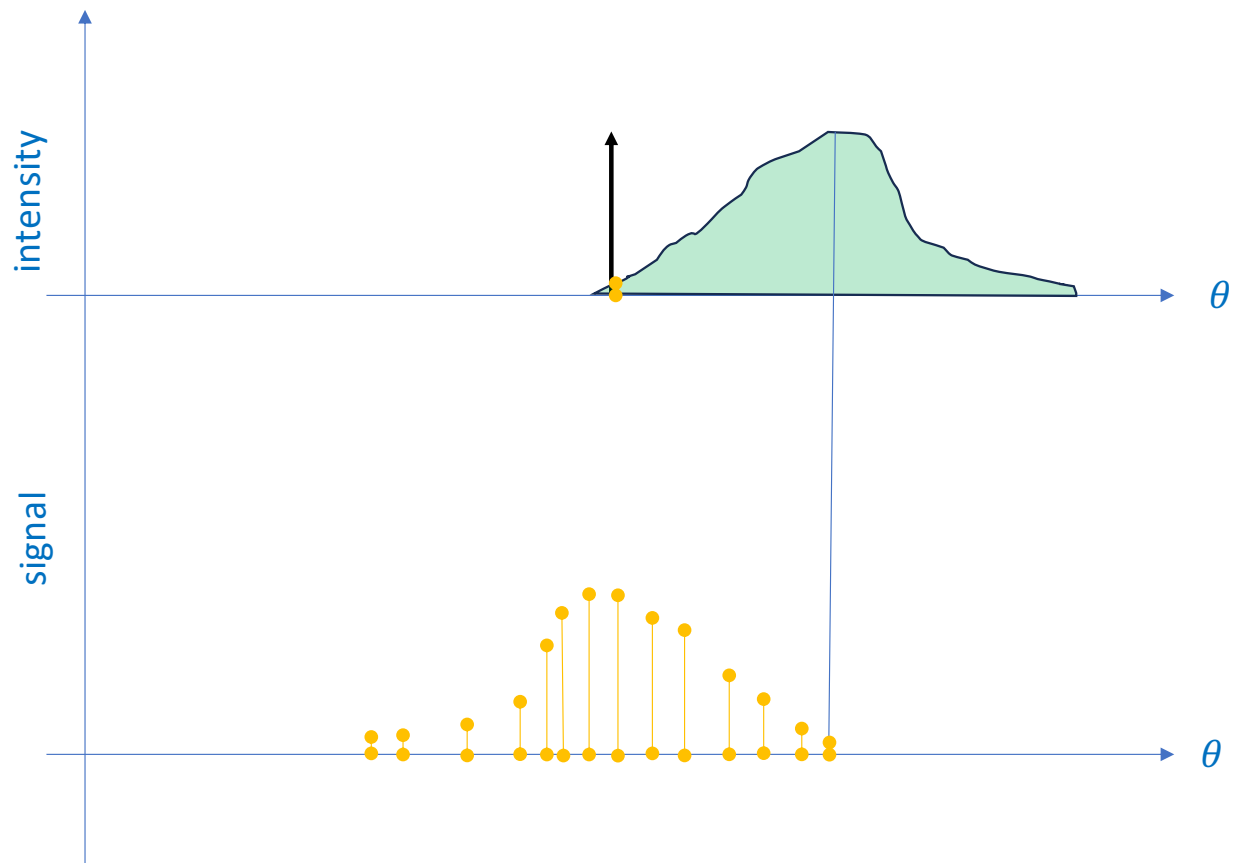


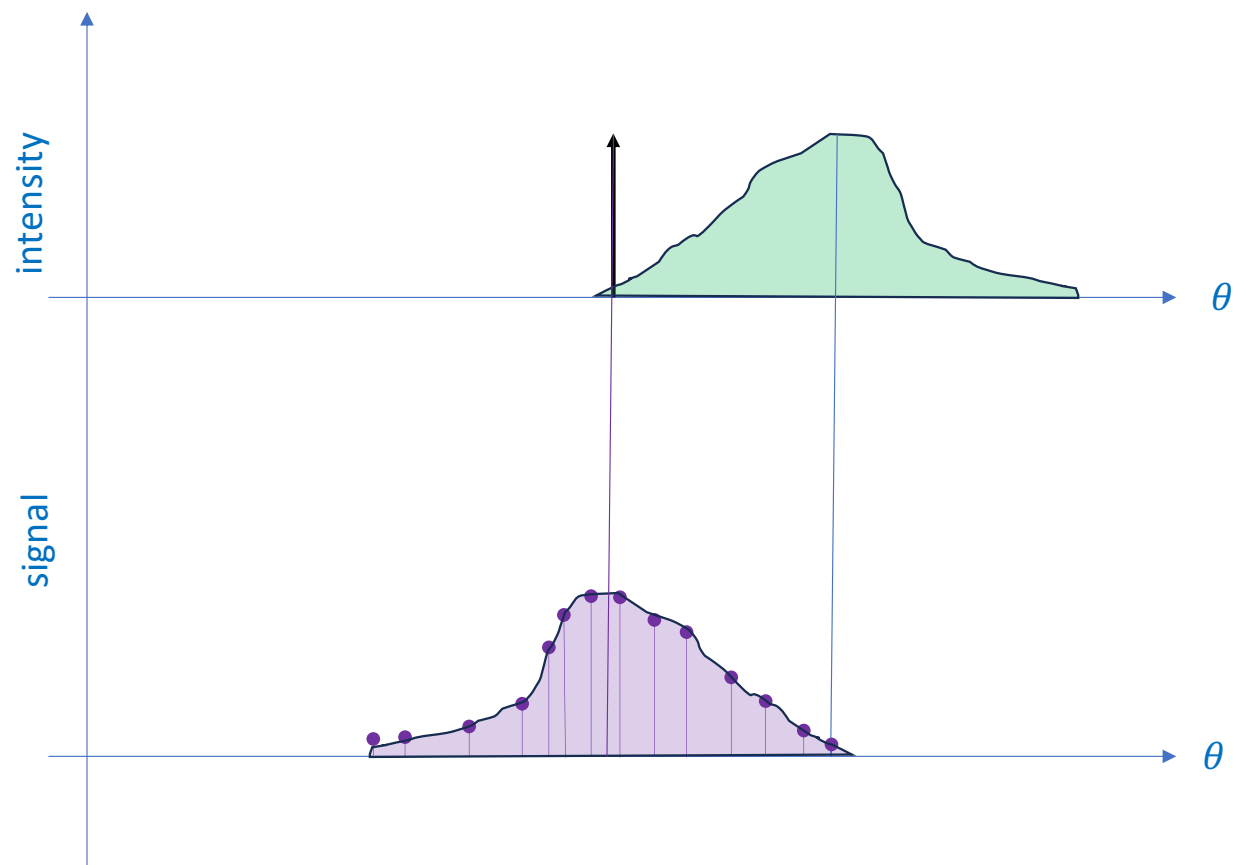


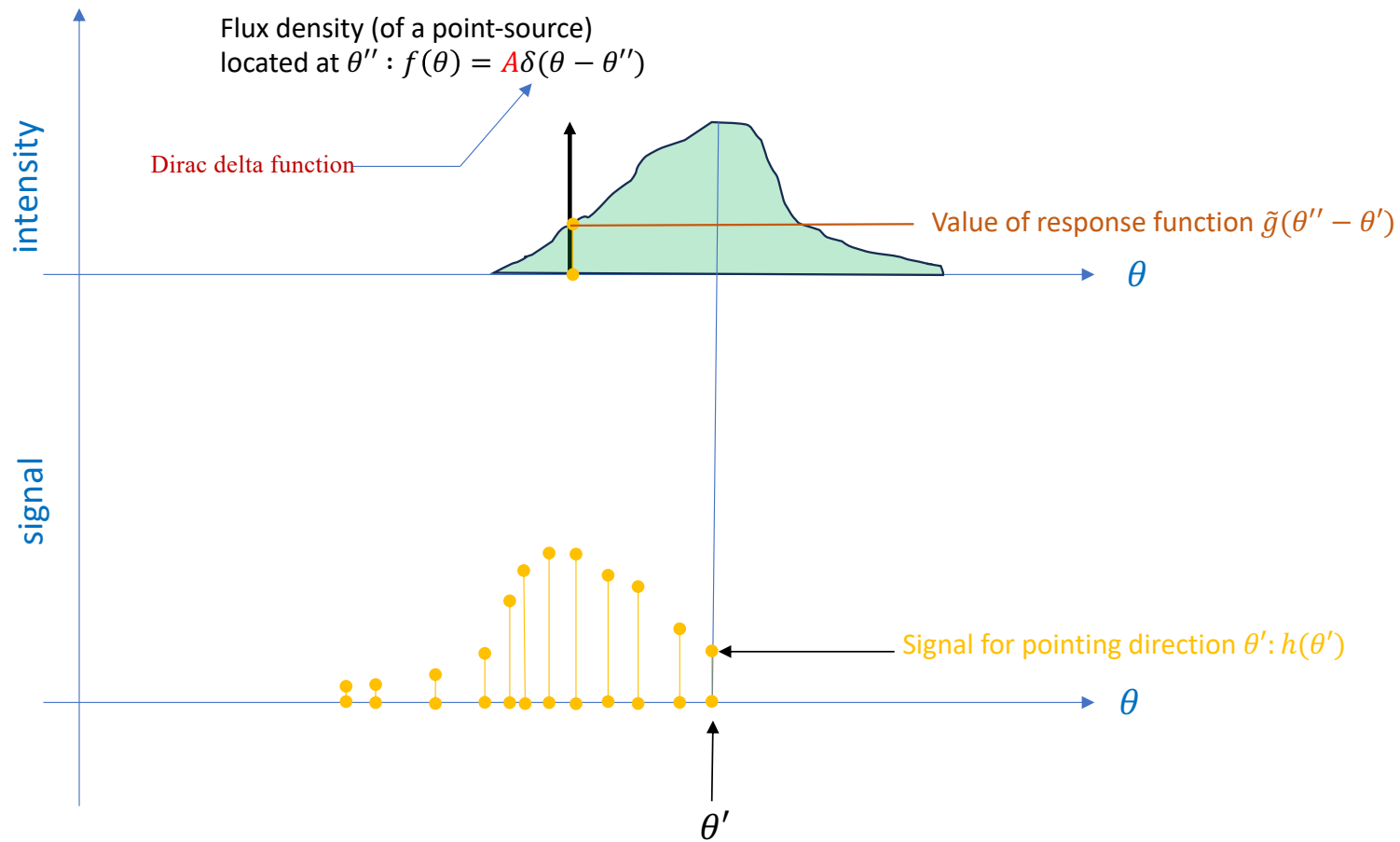


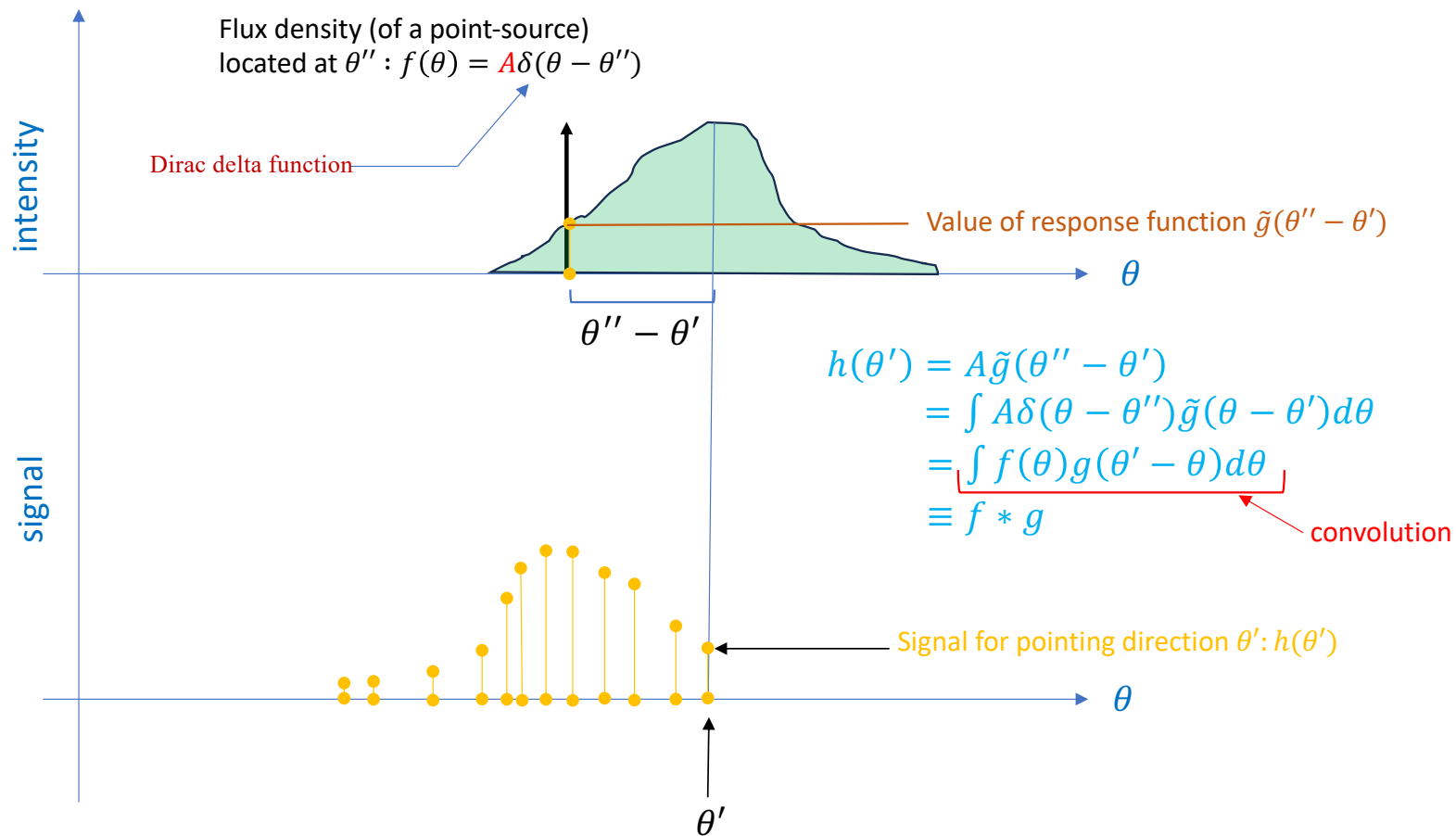












When there are multiple point sources

$$\begin{aligned}h(\theta') &= \sum A_i g(\theta_i'' - \theta') \\&= \int \sum A_i \delta(\theta - \theta_i'') \tilde{g}(\theta - \theta') d\theta \\&= \int f(\theta) g(\theta' - \theta) d\theta \\&\equiv f * g\end{aligned}$$

$$f(\theta) \equiv \sum A_i \delta(\theta - \theta_i'')$$

General expression (source intensity can be continuous)

$$h(\theta') = \int f(\theta)g(\theta' - \theta)d\theta$$
$$\equiv f * g$$

## Convolution theorem

$$h(\theta') = \int f(\theta)g(\theta' - \theta)d\theta$$
$$\equiv f * g$$

$$F.T.(f * g) = F.T.(f) \cdot F.T.(g)$$

$$F.T.(f \cdot g) = F.T.(f) * F.T.(g)$$

### Fourier transform

$$h(\theta') = \int f(\theta)g(\theta' - \theta)d\theta$$
$$\equiv f * g$$

$$F.T.(f) \equiv \tilde{f}(k) = \int_{-\infty}^{\infty} f(x)e^{-2\pi i k x} dx$$

## Convolution theorem

$$F.T.(f * g) = F.T.(f) \cdot F.T.(g) \quad \leftrightarrow \quad F.T.(f \cdot g) = F.T.(f) * F.T.(g)$$

$$\begin{aligned} F.T.(f * g) &= \iint f(\theta)g(\theta' - \theta)d\theta e^{-2\pi i u \theta'} d\theta' \\ &= \iint f(\theta)g(\theta' - \theta)d\theta e^{-2\pi i u(\theta' - \theta + \theta)} d\theta' \\ &= \iint f(\theta) e^{-2\pi i u \theta} d\theta g(\theta' - \theta) e^{-2\pi i u(\theta' - \theta)} d\theta' \\ &= \underbrace{\iint f(\theta) e^{-2\pi i u \theta} d\theta}_{F.T.(f)} \underbrace{g(\theta' - \theta) e^{-2\pi i u(\theta' - \theta)} d(\theta' - \theta)}_{F.T.(g)} \end{aligned}$$



1. What we observe is a **convolution** of the source intensity distribution with the response function of the telescope.

2. Convolution theorem:

$$F.T.(f * g) = F.T.(f) \cdot F.T.(g)$$

$$F.T.(f \cdot g) = F.T.(f) * F.T.(g)$$