

# An Introduction to Radio Interferometry

5-5 Error recognition



You can find relevant material  
on my personal webpage

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# Types of errors (I) origin

- Absolute flux errors
- Passband errors (rarely need to face this)
- Complex gain or delay errors
- Other additive errors (RFI, other instrumental emission or standing wave, etc)

# Types of errors (II) form

## Additive errors

$$\widetilde{V(u, v)} = V(u, v) + \epsilon(u, v)$$

## Multiplicative errors

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For the complex visibility  $\widetilde{V_{ij}(t)}$  taken with antennae  $i$  and  $j$

- Baseline based:  $\epsilon_{ij}(t) = g_{ij}(t)$ , complex variable
- Antenna based:  $\epsilon_{ij}(t) = A_i(t)A_j(t)e^{\phi_i(t)}e^{-\phi_j(t)}$ ,  $\longrightarrow$   $A_i(t)$ : real variable, antenna-based amplitude error  
 $\phi_i(t)$ : real variable, antenna-based phase error



# Recognizing errors

- There is no rule of thumb. Rely on the first principles.
- Easier when the source structure is simpler. You can observe a bright, compact source (called check source) near your target source for the purpose of diagnosing errors.
- The errors that have large effects on one or a few visibility points are more easily seen in the visibility domain. Those that have effects on a large number of visibilities (e.g., all the visibilities that are associated with a specific antenna) may be seen in the visibility or image domain.

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For examples of the effects of antenna position errors or delay errors, see Figure 7-10 in [this link](#)

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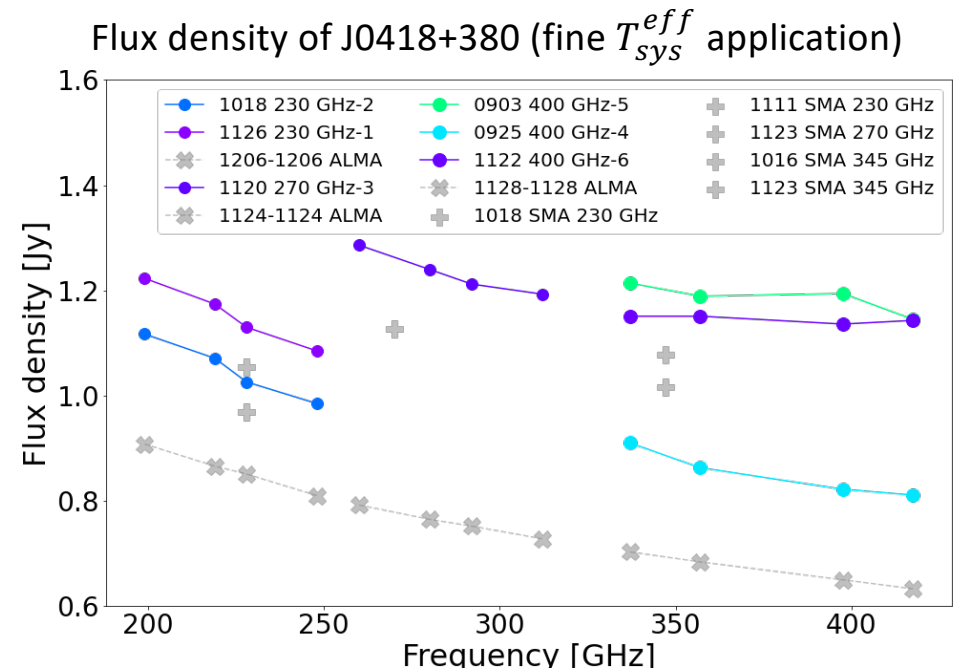
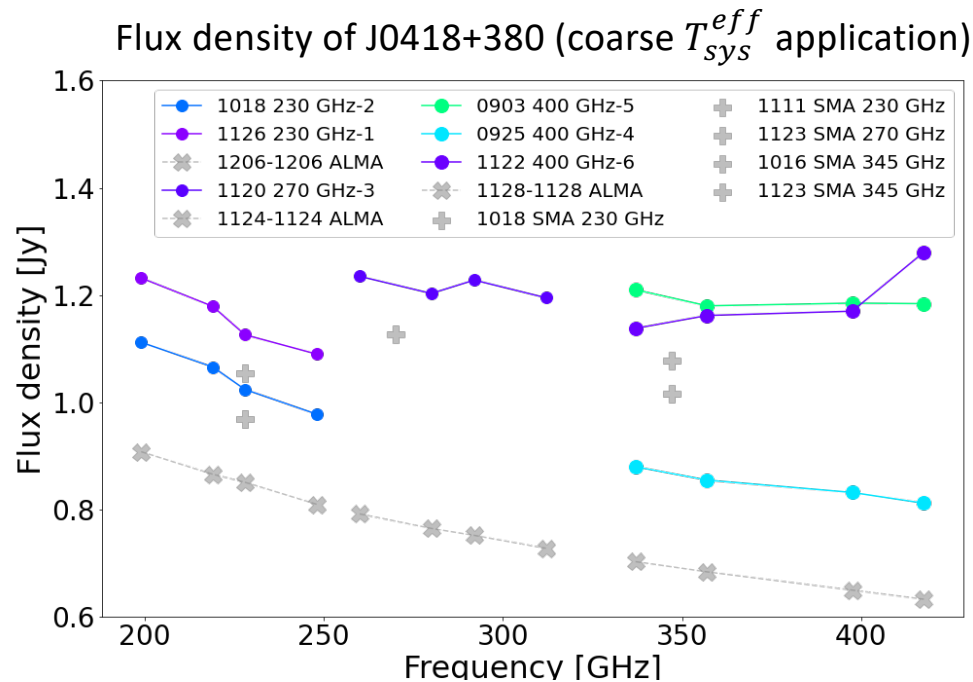
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# Absolute flux issues

(after applying  $T_{sys}$ , calibrating gain amplitude, and referencing the calibrated amplitudes to an absolute flux standard) Multiplicative

Unexpected time variation (e.g., compare with multi-epoch observations)

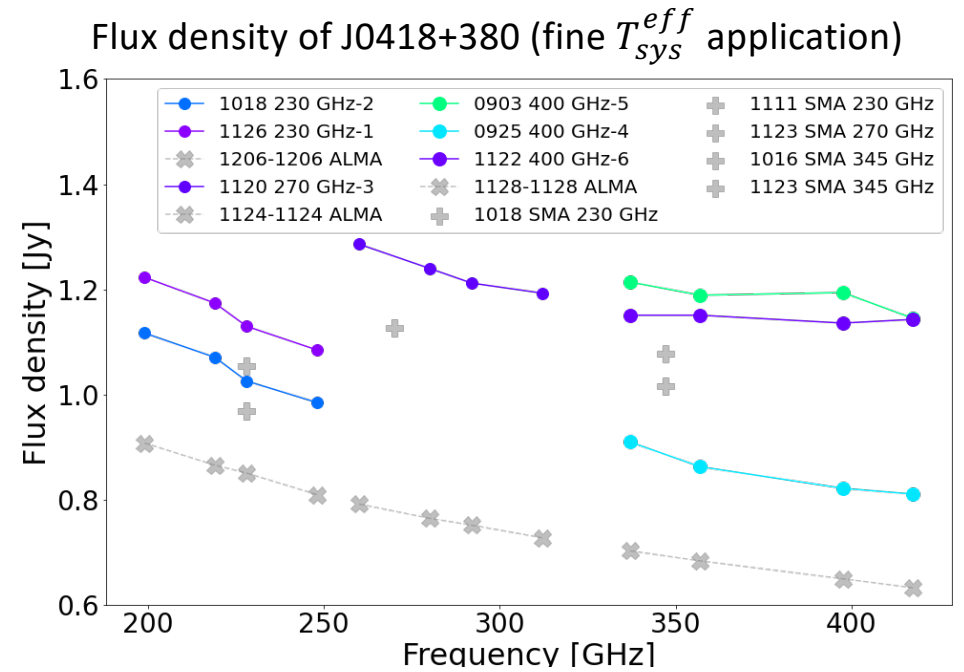
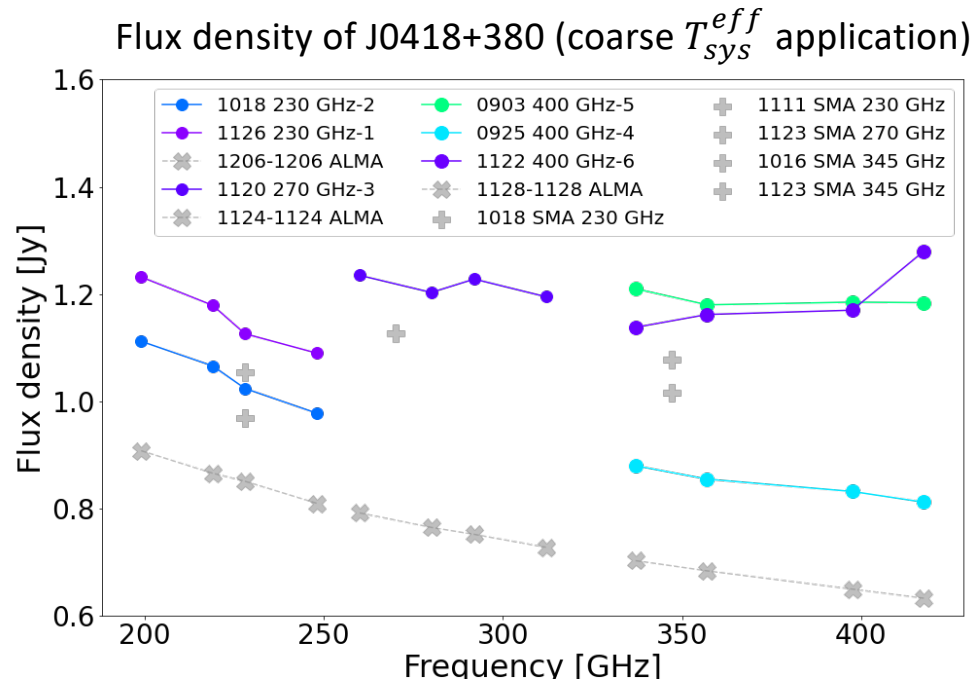


(Chung, Chia-Ying, Master's thesis, 2023, NTU)

# Absolute flux issues

Unexpected time variation (e.g., compare with multi-epoch observations)

Unexpected frequency variation (e.g., unphysical spectral indices )

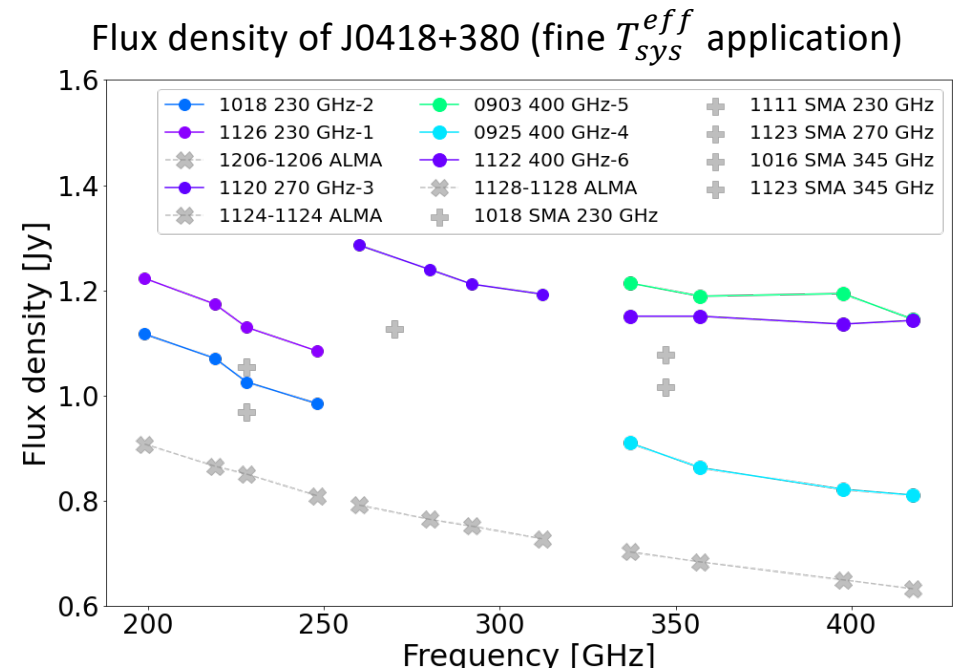
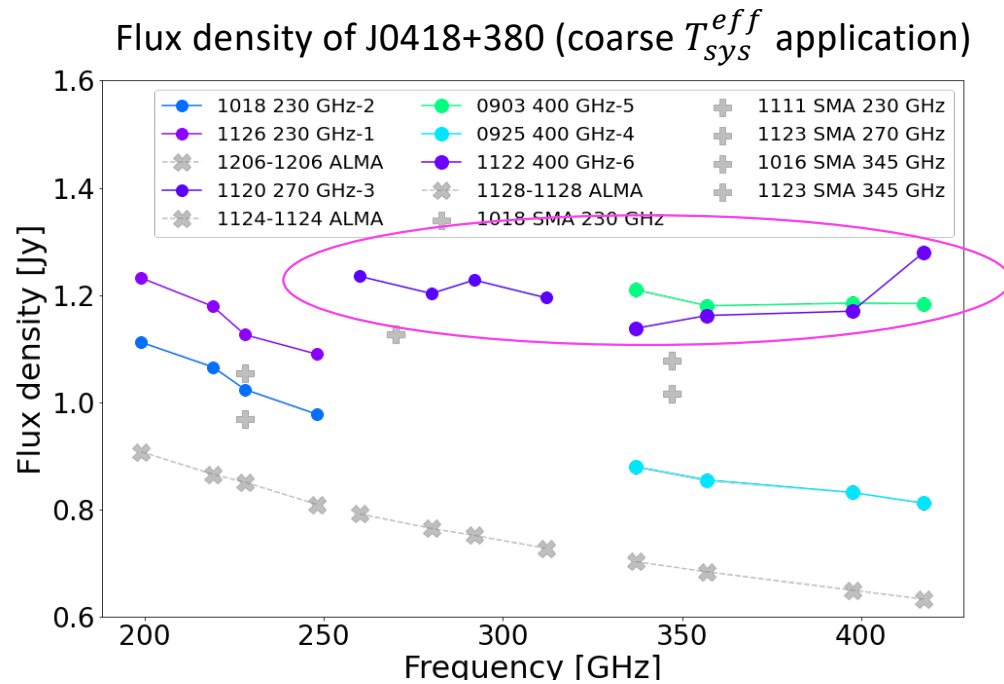


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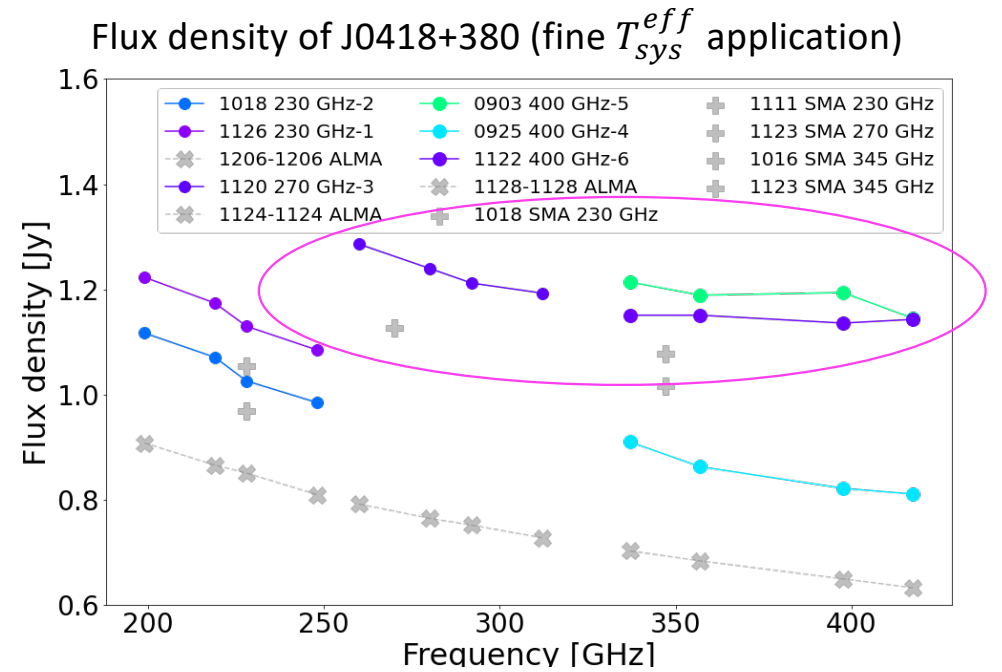
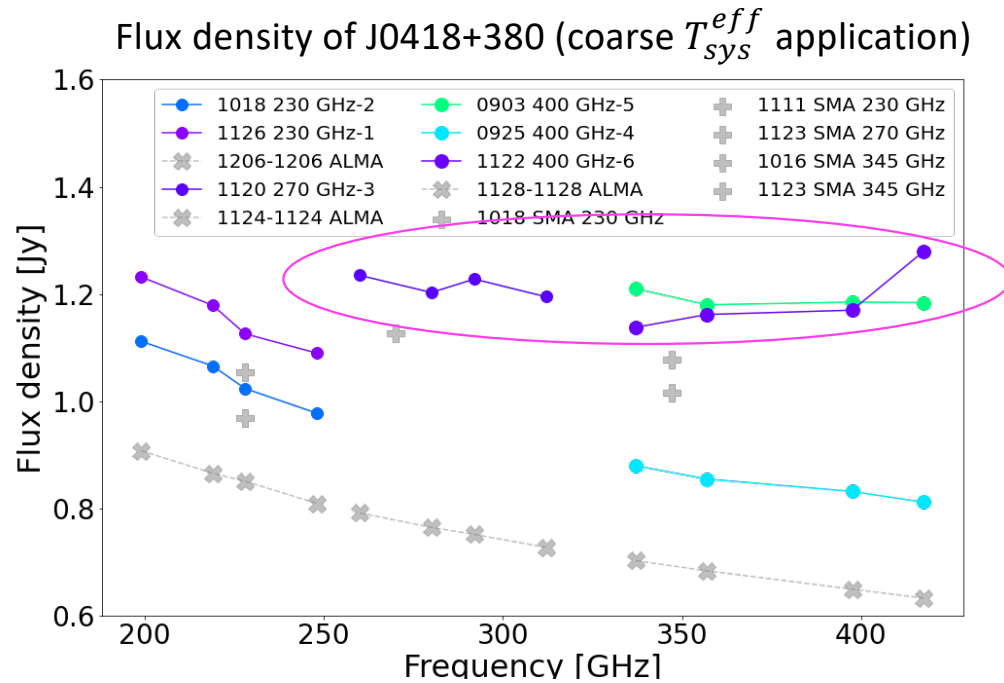


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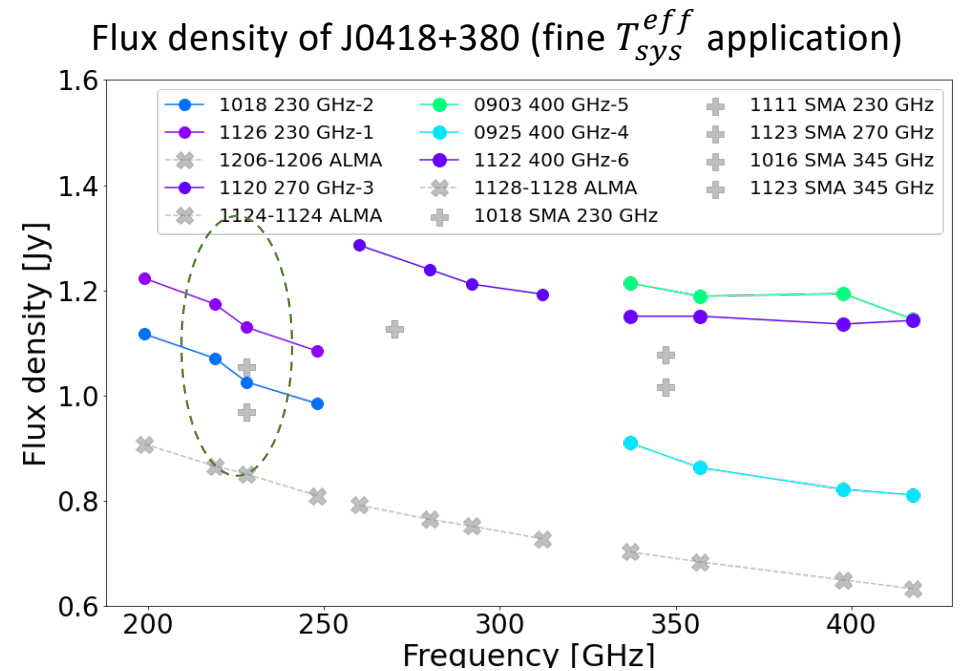
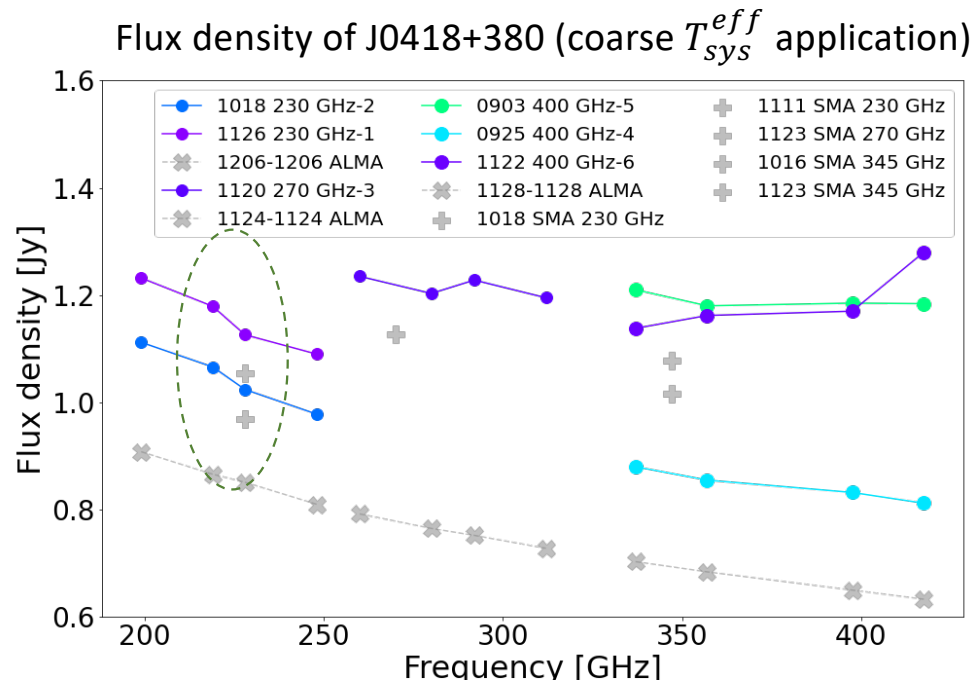
(Chung, Chia-Ying, Master's thesis, 2023, NTU;  
thanks to the help of Dr. Mark Gurwell @ SAO CfA  
to track down the exact problem)



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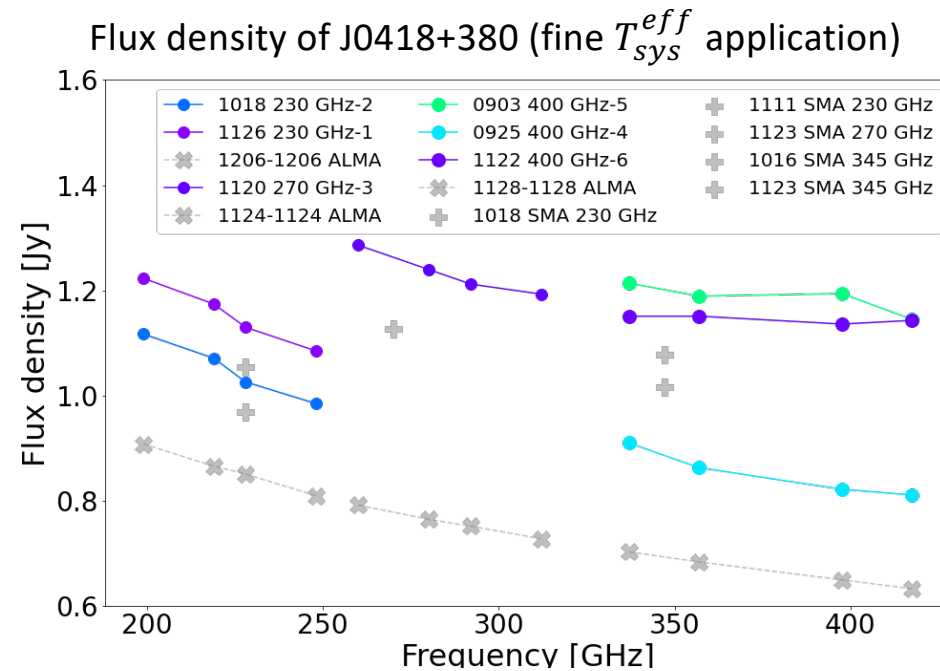


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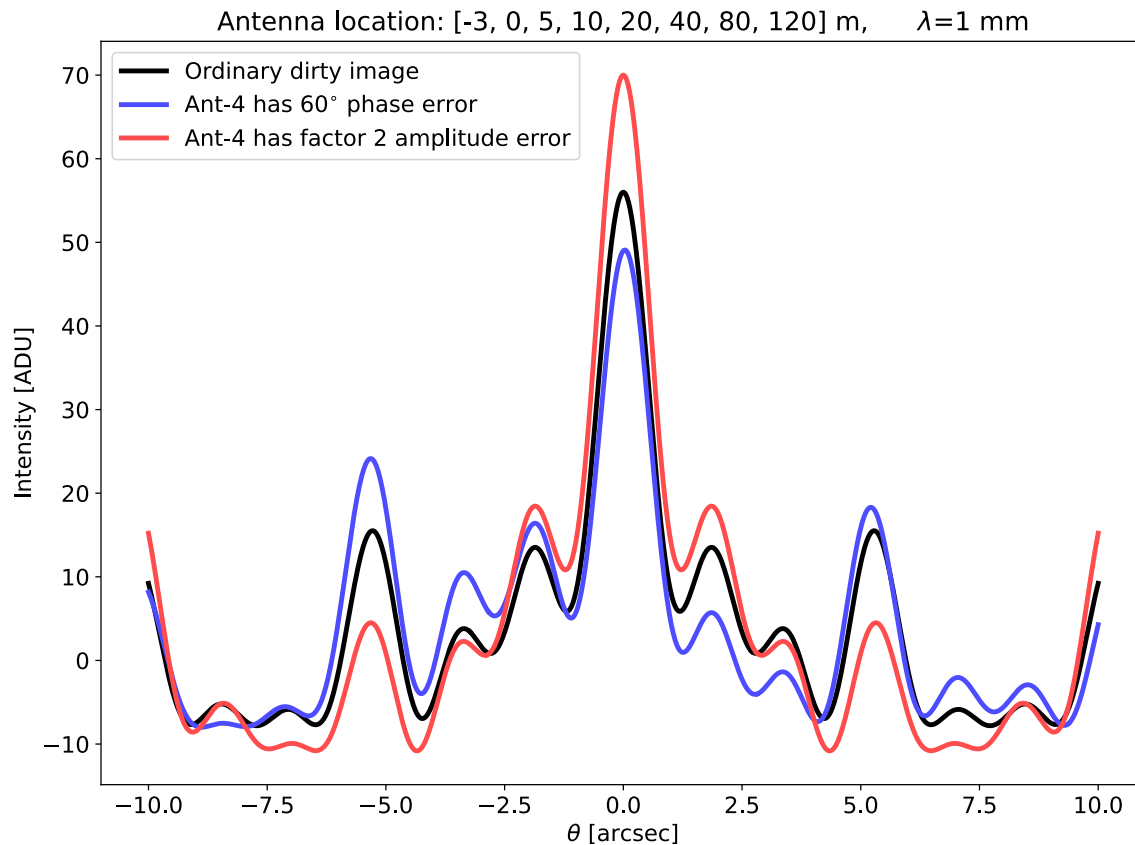
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Is the noise level reasonable? (compare with the theoretical noise level) [Hard. Be careful if you go deeper here]

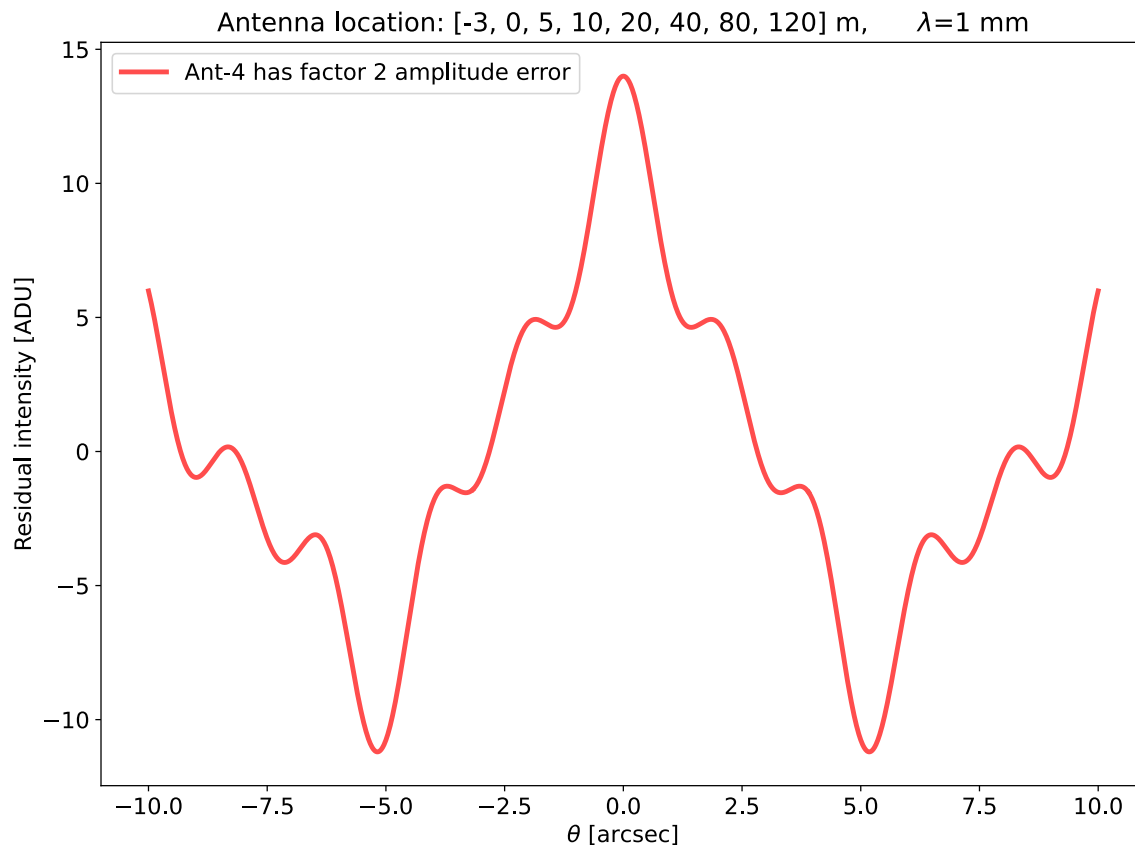
# Effective of multiplicative errors in the image domain



Antenna-based amplitude error:  
symmetric residual

Antenna-based phase error:  
asymmetric residual

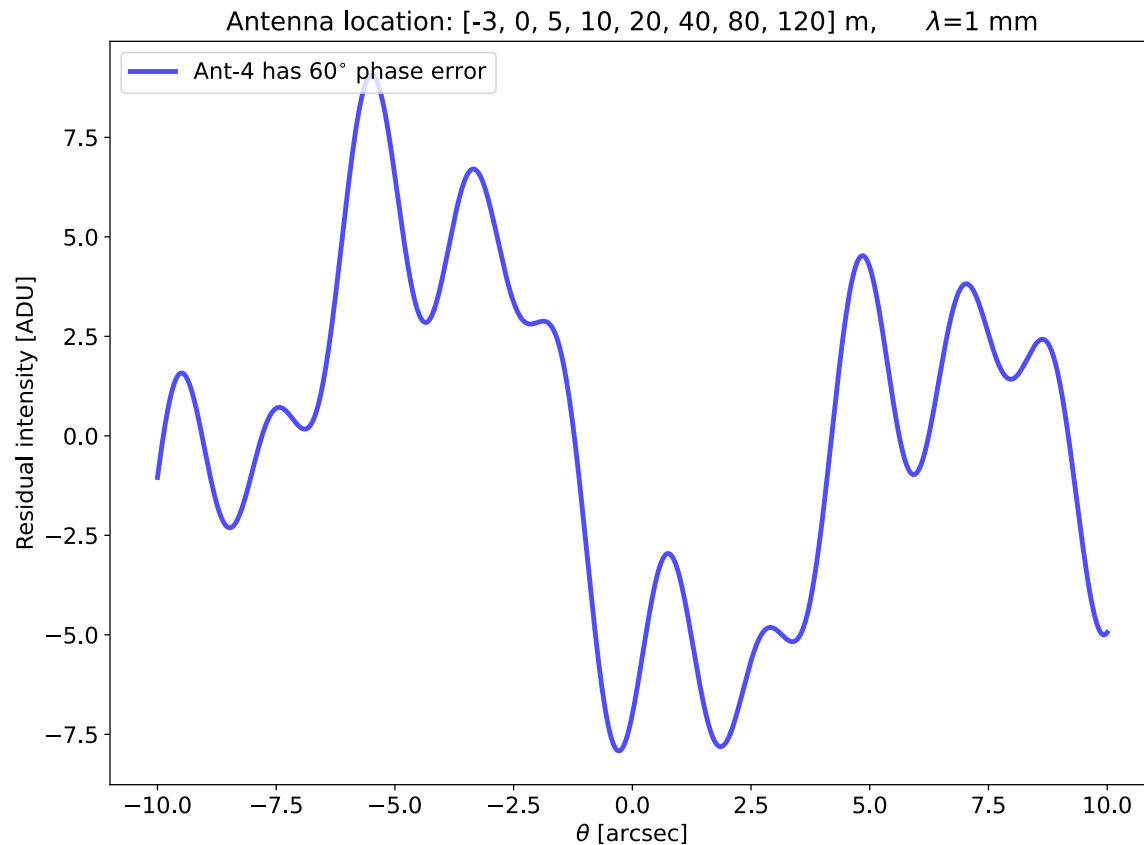
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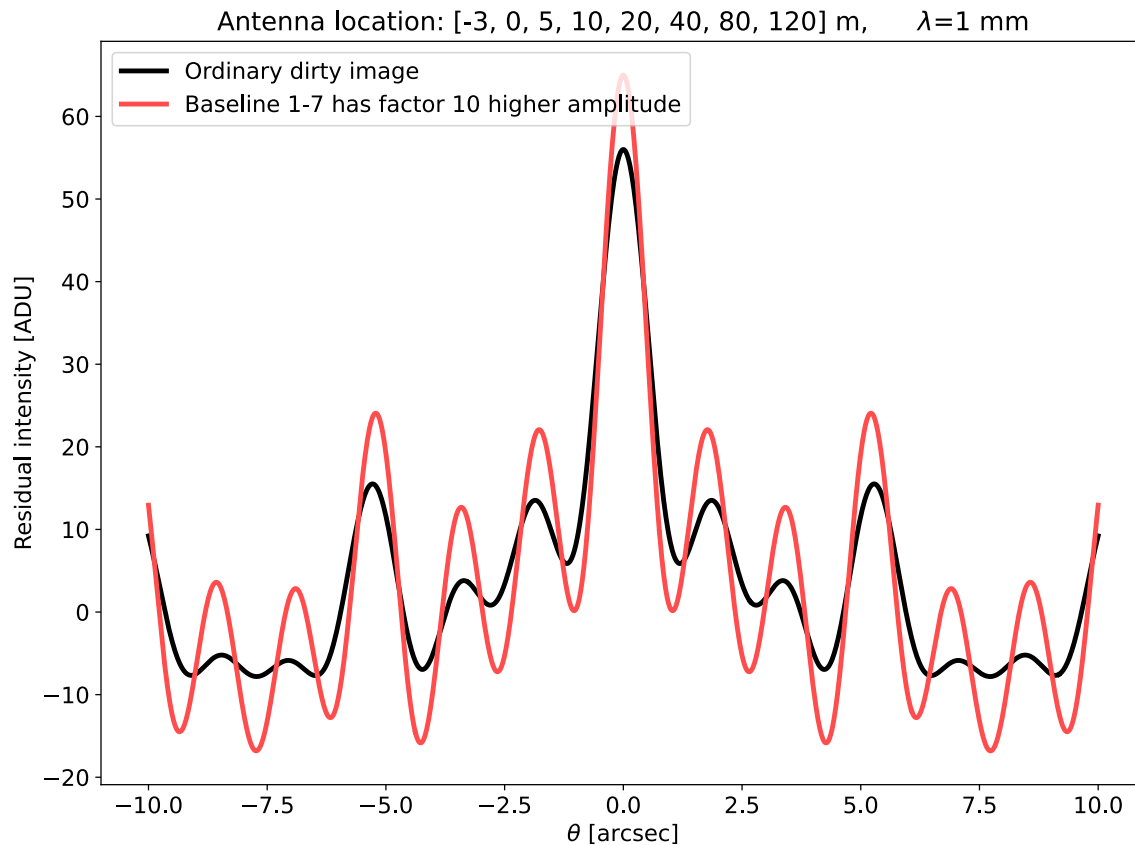
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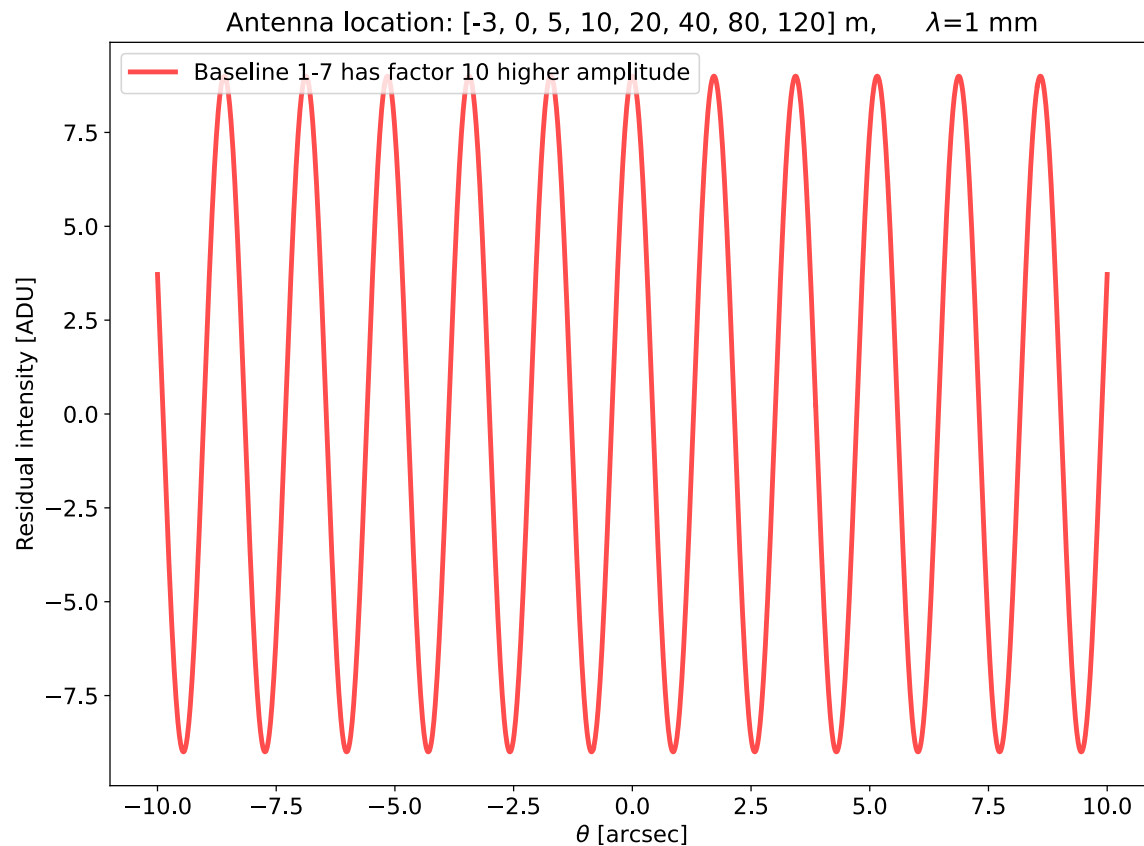
Antenna-based phase error:  
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# Effective of additive errors in the image domain



A high amplitude visibility point:  
stripes

# Effective of additive errors in the image domain



A high amplitude visibility point:  
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1. There are various sources of errors. They can be roughly categorized into additive errors and multiplicative errors.
2. The effects of additive errors are superimposing some artificial intensity distributions onto your image; the effects of multiplicative errors are convolution in the image domain.
3. There is no rule of thumb on which is the best way of treating the errors. You need to know your science case and the effects of the errors very well. Then you can logically decide whether you can calibrate out the errors, trash the affected data, or live with the errors as a compromise.