New results concerning the wobbling properties of $^{183,187}Au$

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1 Introduction

Two wobbling sequences have been identified in ¹⁸³Au by Nandi et. al. [1]. One sequence has two bands with states of negative parity (built on top of the odd $h_{9/2}$ proton) and two bands with states of positive parity (built on top of the odd $i_{13/2}$ proton). Both sequences are considered to have $n_w = 0$ for the *yrast* band and $n_w = 1$ for the one-phonon wobbling band.

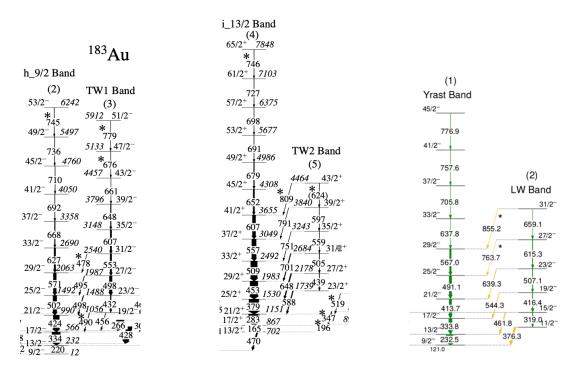


Figure 1: Left: 183 Au: negative parity states based on j = 9/2.Middle: 183 Au: positive parity states based on j = 13/2. Right: The wobbling structure in 187 Au.

On the other hand, Sensharma et. al. [2] has confirmed wobbling motion in ¹⁸⁷Au, with the identification of two such bands, show in figure 1.

2 Numerical application

By using the same formalism as the one applied for ¹⁶³Lu in [3]. Namely, the both positive and negative wobbling sequences from ¹⁸³Au were described with the same analytical expressions for the *excitation energies*.

$$E_{\text{exc}}(I) = \varepsilon_j + \mathcal{H}_{\min}(I) + \Omega_1^I (n_{w_1} + 1) + \Omega_2^I (n_{w_2} + 1) , \qquad (1)$$

such that $E_{\rm exc}(I)=\mathscr{F}(I,j;\mathscr{P})$, where $\mathscr{P}=[\mathscr{I}_1,\mathscr{I}_2,\mathscr{I}_3,V,\gamma]$ is the **free parameter set**. The wobbling frequencies Ω_1 and Ω_2 are the solutions of the algebraic equation:

$$\Omega^4 + B\Omega^2 + C = 0 \tag{2}$$

and

$$\Omega_1 = \sqrt{\frac{1}{2} \left(-B + \sqrt{B^2 - 4C} \right)} \tag{3}$$

$$\Omega_2 = \sqrt{\frac{1}{2} \left(-B - \sqrt{B^2 - 4C} \right)}. \tag{4}$$

3 Coupling schemes

3.1 ¹⁸³Au - positive parity

The spin states with positive parity are created by the coupling of the even-even rotor \vec{R} with the odd-proton $i_{j=13/2}$. As such, the yrast band emerges from a rotor with even angular momentum $\vec{R} = 0, 2, 4, \ldots$, while the first excited wobbling band emerges from the coupling of the same j but with a rotor with odd spin sequence $\vec{R} = 1, 3, 5 \ldots$

References

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