First Observation of Multiple Transverse Wobbling Bands of Different Kinds in ¹⁸³Au

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(Received 7 June 2020; revised 17 July 2020; accepted 27 August 2020; published 25 September 2020)

We report the first observation of two wobbling bands in 183 Au, both of which were interpreted as the transverse wobbling (TW) band but with different behavior of their wobbling energies as a function of spin. It increases (decreases) with spin for the positive (negative) parity configuration. The crucial evidence for the wobbling nature of the bands, dominance of the E2 component in the $\Delta I=1$ transitions between the partner bands, is provided by the simultaneous measurements of directional correlation from the oriented states ratio and the linear polarization of the γ rays. Particle rotor model calculations with triaxial deformation reproduce the experimental data well. A value of spin, I_m , has been determined for the observed TW bands below which the wobbling energy increases and above which it decreases with spin. The nucleus 183 Au is, so far, the only nucleus in which both the increasing and the decreasing parts are observed and thus gives the experimental evidence of the complete transverse wobbling phenomenon.

DOI: 10.1103/PhysRevLett.125.132501

Nuclear wobbling excitation is a manifestation of nonaxial nuclear shape, which was first discussed by Bohr and Mottelson [1]. The nonaxial (triaxial) nuclear shape appears due to the unequal nuclear mass distribution along the three principal axes and implies three unequal moments of inertia about the three principal axes. A triaxially deformed nucleus always tries to rotate around the medium (m) axis having the largest moment of inertia but the presence of the rotations around the other two axes, i.e., short (s) and long (l), generates a precession of the medium axis rotation about the space-fixed angular momentum axis, similar to the classical wobbling motion of an asymmetric top [2]. The energy spectrum of this excitation is given by [1]:

$$E = E_{\rm rot} + (n_w + 1/2)\hbar\omega_{\rm wob}$$

where, the term $E_{\rm rot}$, corresponds to the rotation about the medium axis while n_w is the wobbling quanta and $\omega_{\rm wob}$ is the wobbling frequency with wobbling energy $E_{\rm wob}=\hbar\omega_{\rm wob}$. This generates a series of rotational bands with different n_w .

This exotic excitation has been observed only in a few odd-A nuclei [3–13]. In case of the odd-A nuclei, the odd particle in high-j orbital couples with a triaxial core and modifies the wobbling motion. Depending on the coupling of the odd particle, two types of wobbling bands can be observed: longitudinal wobbling (LW) and transverse wobbling (TW) [14]. In LW, the angular momentum of the odd particle aligns along the medium axis while in TW, it aligns along one of the perpendicular axes (short or long).

An extensive theoretical description of the wobbling motion has been given by Frauendorf and Dönau [14] in terms of a quasiparticle triaxial rotor model. Analytical expression for $\hbar\omega_{\rm wob}$ has been derived with the assumption of "frozen alignment" and harmonic oscillation (HFA). It was shown that $E_{\rm wob}$ increases as a function of angular momentum (I) in case of LW which has been recently observed experimentally in ¹³³La [10] and ¹⁸⁷Au [12]. However, in case of TW, the variation of $E_{\rm wob}$ is highly dependent on the values of the moments of inertia, \mathcal{J}_m , \mathcal{J}_s , and \mathcal{J}_l along the medium, short, and long axes, respectively, of the triaxial core. In general, $E_{\rm wob}$ decreases with I. But in a situation where \mathcal{J}_m is slightly larger than \mathcal{J}_s and

both are much larger than \mathcal{J}_l , the value of E_{wob} increases with the increase of I for the lower values of I and then decreases with I [14]. So far, in all the cases, in which the experimental evidence of TW bands have been reported, only the decreasing part of the E_{wob} has been observed. Therefore, it was generally believed that the LW and TW can be distinguished from the variation of E_{wob} with I; increasing (decreasing) of which corresponds to LW (TW).

In this Letter we report the observation of two TW bands in 183 Au. $E_{\rm wob}$ has been found to increase with I in one of these bands, the first example of such a behavior of a TW band, while in the other band $E_{\rm wob}$ decreases with I. Thereby, it constitutes the only example of experimental observation of a complete scenario of the transverse wobbling phenomenon.

The excited $n_w = 1$ wobbling bands decay to the $n_w = 0$ yrast bands via $\Delta I = 1$, collectively enhanced E2 transitions [12,14]. On the other hand, in case of the signature partner (SP) bands (which occurs when the odd-particle angular momentum is not fully aligned with the rotation axis), the connecting transitions are predominantly $\Delta I = 1$, M1 type [12]. Therefore, experimentally the wobbling partner band and the SP band can be distinguished from

the nature of the connecting transitions between the partner band and the yrast band. For clear identification of wobbling, it is important to experimentally observe both the signature partner band and the wobbling partner band decaying to the main band.

The excited states of ¹⁸³Au in this work were populated using the $^{169}\text{Tm}(^{20}\text{Ne}, 6n)^{183}\text{Au}$ fusion evaporation reaction with 146-MeV ²⁰Ne beam delivered from the K-130 cyclotron facility at the Variable Energy Cyclotron Centre, Kolkata. A thick (23 mg/cm²) ¹⁶⁹Tm foil was used as the target for this experiment. The γ rays emitted from the residual nuclei were detected by the Indian National Gamma Array facility which was composed of eight Compton-suppressed (bismuth germanate anti-Compton shields) clover HPGe detectors and two HPGe planar LEPS (low energy photon spectrometer). A total of $1.5 \times$ 10⁹ two- and higher-fold coincidence data with time stamp were recorded in a fast (250 MHz) digital data acquisition system based on Pixie-16 modules of XIA [15]. The data were sorted using the IUCPIX package [15], and were analyzed using the Radware package [16]. Several E_{γ} - E_{γ} coincidence matrices and an E_{γ} - E_{γ} - E_{γ} cube were

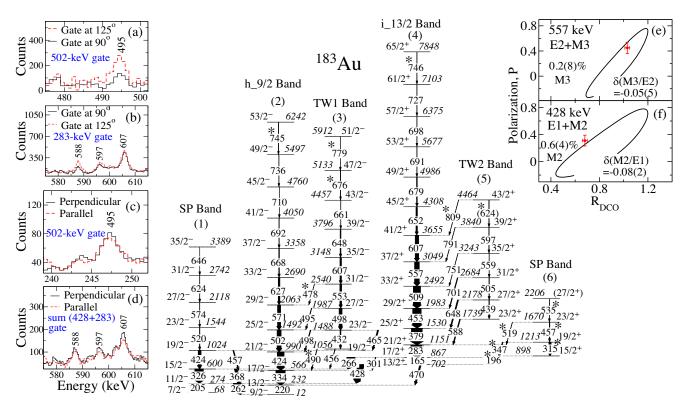


FIG. 1. Level scheme of ¹⁸³Au, established from the present work. Line widths are proportional to their intensities. The level energies are obtained by fitting the γ -ray energies using the code GTOL [19]. Representative gated γ -ray spectra for a few transitions, projected from the two-dimensional DCO and polarization matrices, are shown in the insets (a), (b), (c), and (d). The spectra in (a) and (b) are used for DCO ratio (R_{DCO}) whereas (c) and (d) are used for linear polarization (P) measurements. The insets (e) and (f) show the experimental (symbol) and calculated (solid line) values of R_{DCO} and P of two of the transitions in ¹⁸³Au. The 557 keV is a known E2 transition decaying from $37/2^+$ to $33/2^+$ in band (4) and 428 keV is a known E1 transition from $13/2^+$ to $11/2^-$ between band (4) and band (1). These transitions show very small mixing ratios (δ) as they should be. The new transitions in the level scheme are marked by asterisks.

constructed from which single- and double-gated spectra were projected for the analysis.

A new and improved level scheme of 183 Au has been established (Fig. 1) with 14 new γ rays including a new band, the band (6), compared to the previous studies [17,18].

For unambiguous assignment of multipolarity (λ) and type (E/M) of a γ -ray transition as well as its mixing ratio δ , in case of a mixed transition (e.g., M1+E2), the combined measurements of linear polarization (P) and DCO (directional correlation from the oriented states) ratio $(R_{\rm DCO})$, as described in Refs. [20–24], have been performed. Representative gated spectra for P and $R_{\rm DCO}$ measurements are shown in Figs. 1(a)–1(d). The measured values of P and $R_{\rm DCO}$ of a γ ray have been compared with the values calculated for several δ . For a pure transition, the measured values would lie close to the ones calculated with $\delta \sim 0$. This is demonstrated for two transitions in 183 Au, known to be E2 and E1 [see Figs. 1(e) and 1(f)]. In the case of $\Delta I=1$ transition with enhanced E2 component, measured values would lie close to the calculated ones with large δ .

Both negative- and positive-parity yrast bands, bands (2) and (4), in 183 Au have two side bands each and are connected by $\Delta I=1$ transitions. The measured and the calculated values of P and $R_{\rm DCO}$ for these connecting transitions are plotted in Fig. 2. The low values of δ , obtained for the connecting transitions from bands (1) and (6) to the respective negative and positive parity yrast bands, suggest that they are mostly of pure M1 type (about 1% E2). On the other hand, quite large values of δ are

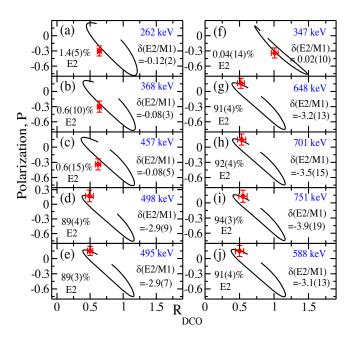


FIG. 2. Experimental (symbol) and calculated (solid line) values (for different mixing ratios δ) of DCO ratios ($R_{\rm DCO}$) and linear polarization (P) of the connecting transitions that decay to the negative parity band (2) (a)–(e) and to the positive parity band (4) (f)–(j).

resulted for the transitions connecting the side bands (3) and (5) to the respective yrast bands. It clearly indicates the predominantly E2 nature (about 90% E2) of these transitions. Therefore, the bands (1) and (6) have been identified as the SP bands, while the bands (3) and (5) are identified as the one-phonon wobbling ($n_w = 1$) bands. It is the first instance that a pair of wobbling bands with positive and negative parity has been identified in a nucleus.

The positive and the negative parity bands in ¹⁸³Au were assigned the configurations of $\pi i_{13/2}$ and $\pi h_{9/2}$, respectively [18]. Both of these orbitals are situated above the Z = 82 shell closure and, hence, they are of particle character with large contribution from the low- Ω $(\Omega = \text{projection of the particle angular momentum on to})$ the symmetry axis) Nilsson configuration. The quasiparticle aligned angular momentum $i_x \approx 6.5\hbar$, estimated for the positive parity yrast band from its level energies, indicates fully aligned $\Omega = 1/2$, $i_{13/2}$ configuration. In case of the negative parity band, the estimated value of i_x $(\approx 3.5\hbar)$ is somewhat less than the fully aligned value $(=4.5\hbar)$, suggesting a small mixing with larger Ω orbitals. However, both of these situations favor the alignment of the odd particle perpendicular to the medium axis and, hence, the occurrence of transverse wobbling.

It was shown in Ref. [14] that the variation of the wobbling energy with spin (I) depends on the type (LW or TW) of wobbling motion. Experimentally, the wobbling energy $E_{\rm wob} = \hbar \omega_{\rm wob}$ can be obtained from the energy differences between the $n_{\rm w} = 1$ wobbling partner band and $n_{\rm w} = 0$ yrast band using the relation [9–12]

$$\begin{split} E_{\text{wob}} &= E(I, n_w = 1) \\ &- [E(I-1, n_w = 0) + E(I+1, n_w = 0)]/2, \end{split}$$

where E(I) is the excitation energy of the state with angular momentum I.

The experimental values of $E_{\rm wob}$ for both the wobbling bands, identified in $^{183}{\rm Au}$ in the present work, are plotted in Fig. 3. It can be seen that the wobbling energy decreases (increases) with I for the negative (positive) parity band and hence, it may be, in general, considered as TW (LW). However, the low- Ω , $i_{13/2}$ configuration for the positive parity band is in contradiction to the LW geometry of the coupling of the odd quasiparticle. Therefore, this band is, most likely, the initial part of a TW band, predicted in Ref. [14] for the same $\pi i_{13/2}$ configuration in $^{163}{\rm Lu}$ but has not been observed in any nucleus prior to this Letter.

In order to understand the observation of wobbling bands in $^{183}\mathrm{Au}$, theoretical calculations in the framework of particle rotor model (PRM) [14,25–28] have been performed. As the input to the calculations, the deformation parameters of $\beta=0.30$ and $\gamma=20.0^{\circ}$ and the moments of inertia $\mathcal{J}_{m,s,l}=36.85,25.70,5.45~\hbar^2/\mathrm{MeV}$ for the

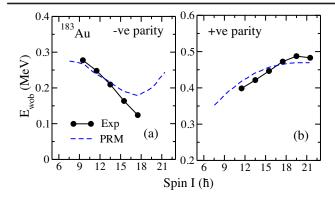


FIG. 3. Experimental wobbling energy $E_{\rm wob}$ as a function of angular momentum I for the negative parity (a) and the positive parity (b) bands in 183 Au. The theoretically calculated values are also shown. The error bars on experimental values are within the size of the data points.

negative parity band, while $\beta = 0.29$, $\gamma = 21.4^{\circ}$, and $\mathcal{J}_{m,s,l} = 50.00, 37.52, 2.38 \ \hbar^2/\text{MeV}$ for the positive parity band have been used. The deformation parameters are obtained by the covariant density functional theory (CDFT) [29,30] with PC-PK1 effective interaction [31], and the moments of inertia are fitted to the energy spectra. In both calculations, the pairing gap $\Delta = 12/\sqrt{A} = 0.89$ MeV is adopted. It should be noted that the calculations can reproduce the experimental B(E2) values [32], which justifies the correct prediction of the deformation parameters by CDFT. We have also carried out calculations by including cranking terms. The equilibrium deformations for both negative and positive parity bands change only within 0.01 for β and 1° for γ from the bandhead to the highest spin, which justifies the assumption of a constant deformation for the PRM calculations.

The results of the calculations are shown in Figs. 3 and 4. The wobbling energies $(E_{\rm wob})$ for both the bands are well reproduced in the calculations. Moreover, the experimental $B(E2)_{\rm out}/B(E2)_{\rm in}$ values, determined from the measured γ -ray intensities, agree well with the calculated ones (Fig. 4) which justifies the correct inputs of triaxiality by CDFT. The overestimation of the $B(M1)_{\rm out}/B(E2)_{\rm in}$ values is attributed to the absence of scissors mode in the calculations [33]. The large $B(E2)_{\rm out}/B(E2)_{\rm in}$ and small $B(M1)_{\rm out}/B(E2)_{\rm in}$ values further support the wobbling interpretation for both the bands.

The values of the three moments of inertia obtained for the two wobbling bands in 183 Au indicate that the ratios of both $\mathcal{J}_m/\mathcal{J}_l$ and $\mathcal{J}_s/\mathcal{J}_l$ are much higher in case of the positive parity $\pi i_{13/2}$ band. It makes this band an ideal candidate for the observation of the, hitherto unobserved, initial part of a TW band in which E_{wob} increases with I until a certain value I_m after which, it starts to decrease. The value of I_m must be sufficiently large in order to experimentally observe the increasing part. According to the harmonic wobbling model with frozen orbital

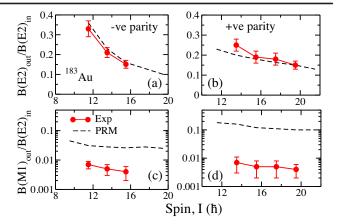


FIG. 4. Measured values of the ratio of transition probabilities, $B(E2)_{\rm out}/B(E2)_{\rm in}$ and $B(M1)_{\rm out}/B(E2)_{\rm in}$, determined from the γ -ray intensities, as a function of angular momentum I for the negative parity (a) and (c) and the positive parity (b) and (d) bands in 183 Au. The theoretical values calculated from PRM are also shown.

approximation of Ref. [14], the larger is the value of the quasiparticle angular momentum j, and closer is the value of $\mathcal{J}_m/\mathcal{J}_s$ to unity, the larger will be the value of I_m .

The I_m values estimated from the HFA model in ¹⁸³Au have been compared, in Table I, with the TW bands reported in the two other normal deformed nuclei, ¹⁰⁵Pd [9] and ¹³⁵Pr [11]. The configurations of the TW bands in ^{105}Pd and ^{135}Pr are $\nu h_{11/2}$ and $\pi h_{11/2}$, respectively. The value of I_m is the largest for the $\pi i_{13/2}$ band in ¹⁸³Au among these nuclei and is much more than the initial spin of its wobbling band. Therefore, E_{wob} for this band would fall in the increasing part of the predicted TW band. This is further established from Fig. 5 in which E_{wob} , calculated from the HFA model using the fitted moments of inertia values $\mathcal{J}_{m,s,l}$ from Table I, are plotted as a function of spin along with the measured E_{woh} . The calculated values are suitably normalised for comparison. Excellent agreement between experiment and theory has been achieved for all the cases. It can be seen that the data points corresponding to the $\pi i_{13/2}$ band in ¹⁸³Au match with the increasing part while for the

TABLE I. The moments of inertia along medium (\mathcal{J}_m) , short (\mathcal{J}_s) and long (\mathcal{J}_l) axes obtained for the wobbling bands in ¹⁸³Au, ¹³⁵Pr, and ¹⁰⁵Pd. The values for the later two nuclei are taken from Refs. [14] and [9], respectively. The estimated values of I_m based on HFA approximation are also given.

	¹⁸³ Au	¹⁸³ Au	¹³⁵ Pr	¹⁰⁵ Pd
	$\pi i_{13/2}$ band	$\pi h_{9/2}$ band	$\pi h_{11/2}$ band	$\nu h_{11/2}$ band
$\overline{{\mathcal I}_m}$	50.00	36.85	21.0	9.24
\mathcal{J}_s	37.52	25.70	13.0	5.87
$\mathcal{J}_{1}^{\tilde{i}}$	2.38	5.45	4.0	1.99
$\mathcal{J}_m/\mathcal{J}_s$	1.33	1.43	1.61	1.57
$I_m(\hbar)$	16.5	7.5	5.5	6.5

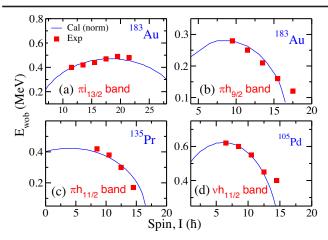


FIG. 5. Experimental and calculated values (see text for details) of wobbling energies as a function of angular momentum (I) for the (a) positive $(i_{13/2})$ and (b) negative $(h_{9/2})$ parity wobbling bands in 183 Au. For comparison, the same for the normal deformed TW bands in 135 Pr (c) and 105 Pd (d) are also shown. Data for the later two nuclei are obtained from Refs. [11] and [9], respectively.

others they fall nicely on the decreasing part of the curves for which the values of I_m are lower than the lowest spin of the observed wobbling bands.

This shows, for the first time, an experimental evidence of the general nature of the TW bands in terms of the variation of the wobbling energy in which both increasing and decreasing parts of the $E_{\rm wob}$ are observed.

It may also be pointed out that similar to the multiple chiral doublet $(M\chi D)$ bands [30], the observation of the presence of multiple TW bands is also an evidence of triaxial shape coexistence in ¹⁸³Au. Though the triaxial shape coexistence has been experimentally confirmed in several nuclei through $M\chi D$ bands (see for example [34,35] and references there in), but its realization through the observation of multiple wobbling bands is being reported for the first time.

In summary, clear experimental evidence for the existence of two wobbling bands, based on a positive $(i_{13/2})$ and a negative $(h_{9/2})$ parity configuration, has been observed in ¹⁸³Au. Both the bands are suggested as the transverse wobbling bands based on the calculations by PRM and HFA models. The present work represents the first observation of TW bands in $A \sim 180$ region and a first identification of the initial increasing part of wobbling energy as a function of spin in a TW band in the whole of the nuclear chart. One important conclusion from the present work is that the decrease of E_{wob} with I may not be considered as a necessary condition for the evidence of TW. Moreover, this is, so far, the only example of the observation of wobbling bands based on both positive and negative parity configurations in a single nucleus and represents the first evidence of triaxial shape coexistence in Au isotopes.

The authors gratefully acknowledge the effort of the cyclotron operators at the Variable Energy Cyclotron Centre (VECC), Kolkata for providing a good quality of ²⁰Ne beam. We thank all the members of VECC Indian National Gamma Array (INGA) collaboration for setting up the array. Partial financial support from Department of Science & Technology (DST), Government of India is gratefully acknowledged for some of the clover detectors of INGA (Grant No. IR/S2/PF-03/2003-III). S. N. and R. B. acknowledge with thanks the financial support received as research fellows from the Department of Atomic Energy (DAE), Government of India. Shabir Dar acknowledges with thanks the financial support from the University Grants Commission (UGC), India. H.P. acknowledges the support received from the Ramanujan Fellowship Research Grant under SERB-DST Grant No. SB/S2/RJN-031/2016 of Government of India. The work of O. B. C. was supported by Deutsche Forschungsgemeinschaft (DFG) and National Natural Science Foundation of China (NSFC) through funds provided to the Sino-German CRC 110 "Symmetries and the Emergence of Structure in QCD" (DFG Grant No. TRR110 and NSFC Grant No. 11621131001). The work of S. F. was supported by the U.S. Department of Energy (Grant No. DE-FG02-95ER40934). The work of P. C. is supported by the U.S. Department of Energy, Office of Science, Office of Nuclear Physics, under Grant No. DE-FG02-94ER40848.

- [5] P. Bringe et al., Eur. Phys. J. A 24, 167 (2005).
- [6] G. Schönwaßer et al., Phys. Lett. B 552, 9 (2003).
- [7] H. Amro et al., Phys. Lett. B 553, 197 (2003).
- [8] D. J. Hartley, R. V. F. Janssens, L. L. Riedinger, M. A. Riley, A. Aguilar *et al.*, Phys. Rev. C 80, 041304(R) (2009).
- [9] J. Timár, Q. B. Chen, B. Kruzsicz, D. Sohler, I. Kuti et al., Phys. Rev. Lett. 122, 062501 (2019).
- [10] S. Biswas et al., Eur. Phys. J. A 55, 159 (2019).
- [11] J. T. Matta, U. Garg, W. Li, S. Frauendorf, A. D. Ayangeakaa et al., Phys. Rev. Lett. 114, 082501 (2015).
- [12] N. Sensharma, U. Garg, Q. B. Chen, S. Frauendorf, D. P. Burdette *et al.*, Phys. Rev. Lett. **124**, 052501 (2020).
- [13] N. Sensharma et al., Phys. Lett. B 792, 170 (2019).

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^[1] A. Bohr and B. R. Mottelson, *Nuclear Structure* (Benjamin, New York, 1975), Vol. II.

^[2] L. D. Landau and E. M. Lifshitz, *Course of Theoretical Physics*, Mechanics, Vol. 1 (Pergamon Press, London, 1960).

^[3] S. W. Ødegard, G. B. Hagemann, D. R. Jensen, M. Bergstrom, B. Herskind *et al.*, Phys. Rev. Lett. **86**, 5866 (2001).

^[4] D. R. Jensen, G. B. Hagemann, I. Hamamoto, S. W. Odegard, B. Herskind *et al.*, Phys. Rev. Lett. **89**, 142503 (2002).

- [14] S. Frauendorf and F. Dönau, Phys. Rev. C 89, 014322 (2014).
- [15] S. Das et al., Nucl. Instrum. Methods Phys. Res., Sect. A 893, 138 (2018).
- [16] D. C. Radford, Nucl. Instrum. Methods Phys. Res., Sect. A 361, 297 (1995).
- [17] W. F. Mueller et al., Phys. Rev. C 59, 2009 (1999).
- [18] L. T. Song et al., Phys. Rev. C 71, 017302 (2005).
- [19] https://www-nds.iaea.org/public/ensdf_pgm/.
- [20] A. Krämer-Flecken, T. Morek, R. M. Lieder, W. Gast, G. Hebbinghaus, H. M. Jäger, and W. Urban, Nucl. Instrum. Methods Phys. Res., Sect. A 275, 333 (1989).
- [21] K. Starosta *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **423**, 16 (1999).
- [22] C. Droste, S. G. Rohoziński, K. Starosta, T. Morek, J. Srebrny, and P. Magierski, Nucl. Instrum. Methods Phys. Res., Sect. A 378, 518 (1996).
- [23] R. Banik, S. Bhattacharyya, S. Biswas, S. Bhattacharya, G. Mukherjee *et al.*, Phys. Rev. C 101, 044306 (2020).
- [24] S. Nandi, G. Mukherjee, T. Roy, R. Banik, A. Dhal et al., Phys. Rev. C 99, 054312 (2019).

- [25] I. Hamamoto, Phys. Rev. C 65, 044305 (2002).
- [26] W. X. Shi and Q. B. Chen, Chin. Phys. C 39, 054105 (2015).
- [27] E. Streck, Q. B. Chen, N. Kaiser, and U.-G. Meißner, Phys. Rev. C 98, 044314 (2018).
- [28] Q. B. Chen, S. Frauendorf, and C. M. Petrache, Phys. Rev. C 100, 061301(R) (2019).
- [29] Relativistic Density Functional for Nuclear Structure, edited by J. Meng, International Review of Nuclear Physics Vol. 10 (World Scientific, Singapore, 2016).
- [30] J. Meng, J. Peng, S. Q. Zhang, and S.-G. Zhou, Phys. Rev. C 73, 037303 (2006).
- [31] P. W. Zhao, Z. P. Li, J. M. Yao, and J. Meng, Phys. Rev. C 82, 054319 (2010).
- [32] P. Joshi, A. Kumar, G. Mukherjee, R. P. Singh, S. Muralithar, U. Garg, R. K. Bhowmik, and I. M. Govil, Phys. Rev. C 66, 044306 (2002).
- [33] S. Frauendorf and F. Dönau, Phys. Rev. C 92, 064306 (2015).
- [34] A. D. Ayangekaa, U. Garg, M. D. Anthony, S. Frauendorf, J. T. Matta *et al.*, Phys. Rev. Lett. **110**, 172504 (2013).
- [35] T. Roy et al., Phys. Lett. B 782, 768 (2018).