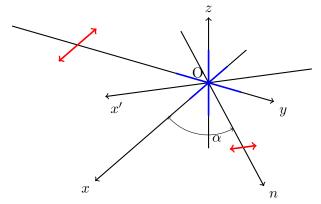
- Broken spherical symmetry (privileged direction) because of
 - anisotropic radiation
 - magnetic field
- Polarization amplitudes of most spectral lines in the sun are of order 10^{-3} : good accuracy needed
- achieved 10^{-5} accuracy with ZIMPOL

Radiation of a non relativistic charge accelelerated free or bound (electron in the classical description with oscillators)

$$\vec{E_{rad}} \propto \vec{n} \times (\vec{n} \times \vec{a}) \tag{1}$$

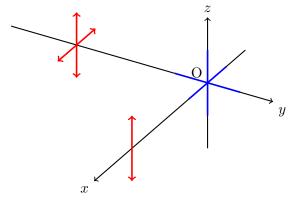
Electron accelerated by a force due to the electric field of polarized radiation (Thomson scattering)



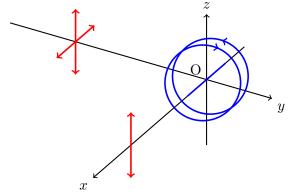
 $\vec{x}, \vec{y}, \vec{n}, \vec{x'}$ same plane

- Anisotropic radiation creates a net dipole moment in scattering atoms or molecules (Rayleigh scattering)
- shorter wavelengths more scattered (blue sky)
- in the Sun: Thomson scattering by electrons and Rayleigh scattering by neutral H

Incident unpolarized light, cartesian oscillators



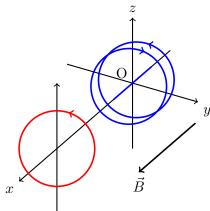
Incident unpolarized light, circular oscillators



coherence between circular

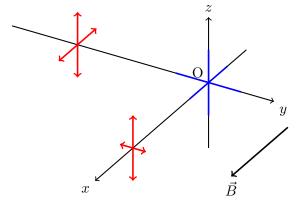
oscillators

Magnetic field along l.o.s , circular oscillators $\,$



ciclotron radiation

Unpolarized incident light and magnetic field along l.o.s



'Rosette' effect (circular oscillators)

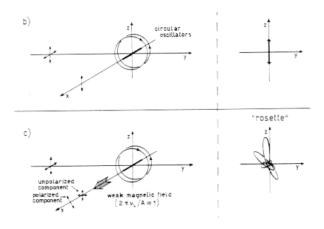
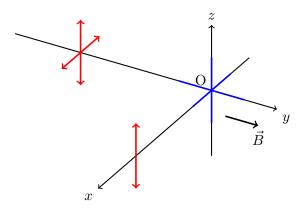


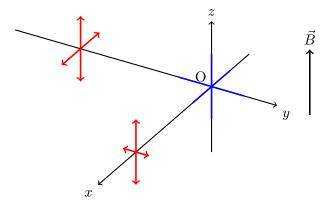
Figure 1: Incoherence in the circular oscillators - Hanle effect

Vertical magnetic field



effect adding to that of the radiation: linear polarization (no V signal, no Hanle effect)

Horrizontal magnetic field, perpendicular to l.o.s.



horrizontal component(Hanle effect), no V signal

Quantic representation (bound electron)

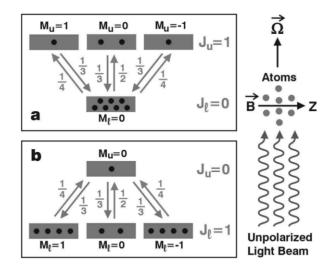


Figure 2: Emission/Absorption

Quantic representation (bound electron)

- Correspondence between atomic transitions and radiation components
 - $\Delta M = 1$, $\sigma_+ (\nu + \nu_L)$, clockwise oscillator
 - $\Delta M = 0, \pi (\nu), \text{ x oscillator}$
 - $\Delta M = -1$, $\sigma_{-}(\nu \nu_{L})$, anticlockwise oscillator
- similar mechanism(Fig 2 b) for absorption lines (experimetally done by observing same line in filaments: absorption and in prominences: emission)
- Anisotropic radiation populates/depopulates twice M=0 sublevel
- Magnetic field:
 - splitting of energy level J into 2J+1 sublevels (Zeeman effect)
 - \bullet creates in coherence between population of levels corresponding to σ components which leads to incoherence in the 2 circular oscillators (Hanle effect)

Second specter of the sun

With the same geometry: Q represents linear polarization parallel to the closest limb

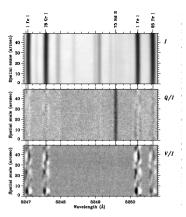


Fig. 1. Examples of the different characters of the ordinary intensity spectrum (Stokes I, no diagram), the linearly polarized spectrum (Stokes QI), here called the "second solar spectrum"), and the circularly polarized spectrum (Stokes V/I, bottom diagram). While the circular polarization shows Zeeman-effect signatures due to magnetic canopies at the supergranulation size scale, the linear polarization is dominated by the signature of coherent scattering, which here unexpected is due to a transition in insinzed nepdomium. The recogning

Second specter of the sun

assume U = V = 0 \Longrightarrow no U, V->Q cross talk all signals divided by I (no zero I), no I->Q cross talk Q->I, U, V would reduce Q ampl by 0.5% Q signal generated by other mechanisms (is the assumption U, V->Q cross talk = 0 correct?)

- instrumental
 - atmospheric seeing and CCD noise: resolved by modulating the signal at high frequency: 2 modulators one at 42 kHz and 84 kHz
 - polarization by reflection (in the plane perpendicular to the plane of reflection for angles bigger than a certain angle which depends on refraction indices) - not important
- real effects of magnetic field
 - position near the (geographic) north/south pole where influence of the magnetic field is smaller
 - cross talk Stokes V generated by the magnetic field along l.o.s recognized (sign reversed periodically in spatial repr.)
 - vertical magnetic field no effect, horrizontal magnetic field perpendicular to line of sight - not considered

Observation details

- instrument: ZIMPOL, the Zurich Imaging Stokes Polarimeter
- telescope: McMath-Pierce(National Solar Observatory (Kitt Peak, USA).)
- accuracy: 10^{-5}
- April 1995, slit positioned 5 seconds of arc inside the north polar limb (where the cosine μ of the heliocentric angle is 0.1)
- 10 minutes exposure time