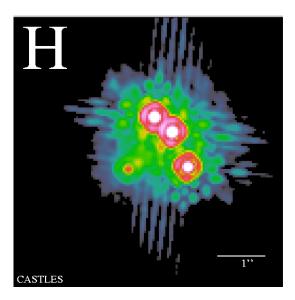
## Representación de B1422+231



 $Figura 1: Imagen \ de \ http://www.cfa.harvard.edu/castles/Postagestamps/Gifs/Fullsize/B1422H.gif$ 

Imagen(nx=1000) obtenida con una fuente de anillos circulares concéntricos (ny = 200) a traves de una SIS (gamma = 0.6, x01 = -0.10, x02 = -0.20)

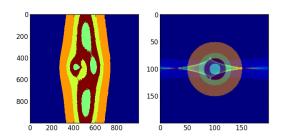


Figura 2: xl=4, yl=4

Imagen de una fuente (hubble deep field) a traves de una lente gravitacional(esfera singular isoterma)

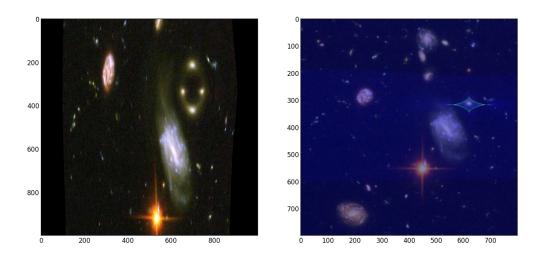


Figura 3: SIS(gamma=0.5,x01=-2.76,x02=3.15),ny=800,nx=1000,xl=8,yl=8

Mapa de magnificación en el caso de un sistema binario

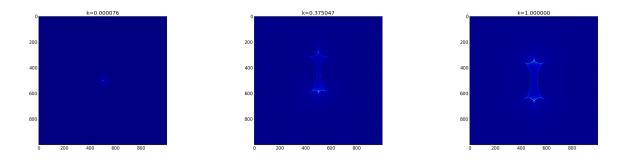
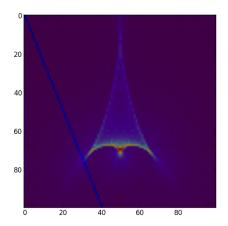


Figura 4: k=M1/M2

Mapa de magnificación en el caso de un sistema binario (estrella-planeta) M1/M2 = 7.6 \* 10 \*\* (-5), a = 1.61, ny = 100, nx = 30000, xl = 1, yl = 0.00025



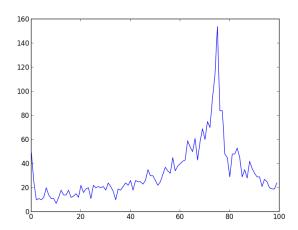
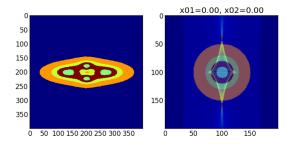
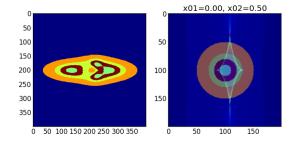
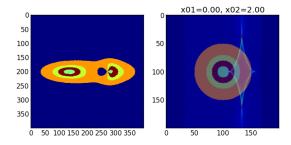


Figura 5: A la derecha el corte del mapa de magnificacion al largo de la curva: y = abs(tan(theta) \* x + u0 \* (cos(theta) + sin(theta) \* tan(theta))), u0 = 0.359, theta = 2.756 rad

Imagenes de una fuente de anillos concéntricos (cruzando la caustica) y mapas de magnificación correspondientes







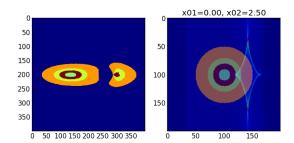


Figura 6:  $SIS \ gamma = -0.5$