

# PHY 321 HW 6 p1

1) Review typed in Jupyter notebook.

2) a)  $\vec{F} = k(x\vec{e}_1 + 2y\vec{e}_2 + 3z\vec{e}_3)$   $k$  is a constant so I'll factor that in at the end.

$$\vec{\nabla} \times \vec{F} = \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ x & 2y & 3z \end{vmatrix} = \frac{\partial}{\partial y} 3z \hat{i} + \frac{\partial}{\partial z} x \hat{j} + \frac{\partial}{\partial x} 2y \hat{k} - \frac{\partial}{\partial y} x \hat{k} - \frac{\partial}{\partial z} 2y \hat{i} - \frac{\partial}{\partial x} 3z \hat{j}$$

technically  $k$  this  $\rightarrow$

$$= 0 + 0 + 0 - 0 - 0 - 0 = 0$$

$$\vec{\nabla} \times \vec{F} = 0 \Rightarrow \text{Conservative force}$$

b)  $\vec{F} = y\vec{e}_1 + x\vec{e}_2 + 0\vec{e}_3$

$$\vec{\nabla} \times \vec{F} = \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ y & x & 0 \end{vmatrix} = 0 + \frac{\partial}{\partial z} y \hat{j} + \frac{\partial}{\partial x} x \hat{k} - \frac{\partial}{\partial y} y \hat{k} - \frac{\partial}{\partial z} x \hat{i} - 0$$

$$= 0 + 0 + 1\hat{k} - 1\hat{k} - 0 - 0 = 0$$

$$\vec{\nabla} \times \vec{F} = 0 \Rightarrow \text{Conservative force}$$

c)  $\vec{F} = k(-y\vec{e}_1 + x\vec{e}_2 + 0\vec{e}_3)$  again, I'll factor in the  $k$  at the end.

$$\vec{\nabla} \times \vec{F} = k \cdot \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ -y & x & 0 \end{vmatrix} = 0 + \frac{\partial}{\partial z} x \hat{j} + \frac{\partial}{\partial x} 0 \hat{k} - \frac{\partial}{\partial y} 0 \hat{k} - \frac{\partial}{\partial z} (-y) \hat{i} - 0$$

$$= 0 + 0 + 0 - 0 - 0 = 2k\hat{k}$$

factoring back in the  $k$ , we get  $\vec{\nabla} \times \vec{F} = 0\hat{i} + 0\hat{j} + 2k\hat{k} \neq 0 \Rightarrow \text{Not conservative force}$

2d) We need to find  $V$  for parts a and b.

\* part a:  $\vec{F} = k(x\vec{e}_1 + 2y\vec{e}_2 + 3z\vec{e}_3)$ . ~~we need  $\vec{F}$~~

We need  $F = -\nabla V$  i.e.  $-\frac{\partial}{\partial x} = kx$ ,  $-\frac{\partial}{\partial y} = k2y$ ,  $-\frac{\partial}{\partial z} = k3z$ .

In this we can just integrate.

$$\Rightarrow \boxed{V = -k\left(\frac{1}{2}x^2 + y^2 + \frac{3}{2}z^2\right)}$$

$$\begin{aligned}\text{To confirm, } -\nabla V &= -\left(\frac{\partial}{\partial x}V\vec{e}_1 + \frac{\partial}{\partial y}V\vec{e}_2 + \frac{\partial}{\partial z}V\vec{e}_3\right) = -(-k(x\vec{e}_1 + 2y\vec{e}_2 + 3z\vec{e}_3)) \\ &= k(x\vec{e}_1 + 2y\vec{e}_2 + 3z\vec{e}_3) = \vec{F} \quad \checkmark\end{aligned}$$

\* part b:  $\vec{F} = y\vec{e}_1 + x\vec{e}_2 + 0\vec{e}_3 = y\vec{e}_1 + x\vec{e}_2$ .

We need  $-\frac{\partial}{\partial x} = y$  and  $-\frac{\partial}{\partial y} = x$ .

$$\text{So we get } \boxed{V = -xy}$$

$$\begin{aligned}\text{To confirm, } -\nabla V &= -\left(\frac{\partial}{\partial x}V\vec{e}_1 + \frac{\partial}{\partial y}V\vec{e}_2 + \frac{\partial}{\partial z}V\vec{e}_3\right) = -(-y\vec{e}_1 - x\vec{e}_2 + 0\vec{e}_3) \\ &= y\vec{e}_1 + x\vec{e}_2 + 0\vec{e}_3 = \vec{F} \quad \checkmark\end{aligned}$$

3) a) See jupyter notebook! Looks like



b) Equilibrium points where  $F=0$ . Since  $F=-\nabla V$ , using  $V=V_0\left(\left(\frac{a}{r}\right)^{12}-\left(\frac{b}{r}\right)^6\right)$

Technically ignoring direction here

$$F = -\frac{d}{dr}\left[V_0\left(\left(\frac{a}{r}\right)^{12}-\left(\frac{b}{r}\right)^6\right)\right] = -V_0 \cdot \frac{d}{dr}\left(a^{12}r^{-12} - b^6r^{-6}\right) = 0$$

$$= -V_0 \cdot \left(-\frac{12a^{12}}{r^{13}} + \frac{6b^6}{r^7}\right) = \frac{12V_0a^{12}}{r^{13}} - \frac{6V_0b^6}{r^7} = 0$$

$$\frac{12V_0a^{12}}{r^{13}} = \frac{6V_0b^6}{r^7} \Rightarrow \frac{2a^{12}}{r^{13}} = \frac{b^6}{r^7}$$

$$\Rightarrow 2a^{12} = b^6r^6 \Rightarrow r^6 = \frac{2a^{12}}{b^6} \Rightarrow r = (2)^{1/6} \cdot \frac{a^2}{b} \approx 1.122 \frac{a^2}{b}$$

This is a stable equilibrium point, since the force will push nearby particles towards this point.

c) I already found most of it from  $-\frac{dV}{dr} = V_0\left(\frac{12a^{12}}{r^{13}} - \frac{6b^6}{r^7}\right)$

This is our magnitude. Now we just need to consider direction to make the force vector.

$$\vec{F} = V_0\left(\frac{12a^{12}}{r^{13}} - \frac{6b^6}{r^7}\right) \cdot \frac{\vec{r}}{||\vec{r}||} \quad \leftarrow \text{total vector divided by magnitude}$$

Assuming just 2 particles, this gives:

$$\vec{F} = V_0\left(\frac{12a^{12}}{||\vec{r}_i - \vec{r}_j||^{13}} - \frac{6b^6}{||\vec{r}_i - \vec{r}_j||^7}\right) \cdot \frac{\vec{r}_i - \vec{r}_j}{||\vec{r}_i - \vec{r}_j||}$$

Yes, this is a conservative force. It depends only on position, and can be written in terms of a single dimension  $\vec{r}$ .



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Note: I'm combining 4a and 4b, since its intuitive to scale in the analysis.

4a)  $V(x) = k \frac{x^2}{2}$ . No driving force. Drag force  $-bV$ .  
 and 4b)  $\rightarrow$  This is standard for harmonic oscillator. Since  $F = -\nabla V$ , this gives  $\boxed{F = -kx}$  force from potential

As usual, this gives the differential equation below. Note we have the additional force  $\boxed{F = -bV}$

From  $F = ma$   $\Rightarrow m \frac{d^2x}{dt^2} + b \frac{dx}{dt} + kx = 0$   
 with both forces above  $\rightarrow$  dividing by  $m$  gives  $\frac{d^2x}{dt^2} + \frac{b}{m} \frac{dx}{dt} + \frac{k}{m} x = 0$

4b part  $\rightarrow$  setting  $\omega_0^2 = \frac{k}{m} \Rightarrow \omega_0 = \sqrt{\frac{k}{m}}$ ,  $\gamma = \frac{b}{2m\omega_0}$  and  $\tau = \omega_0 t$ , we can rewrite the above equation as

\* Scaled equation:  $\boxed{\frac{d^2x}{d\tau^2} + 2\gamma \frac{dx}{d\tau} + x = 0}$   
~~note that we've divided by  $\omega_0^2$~~

Note I skipped intermediate  $\omega_0^2 \frac{d^2x}{d\tau^2} + \frac{b\omega_0}{m} \frac{dx}{d\tau} + \omega_0^2 x = 0$ , where you then divide by  $\omega_0^2$ , giving this

We've now isolated an expression for the acceleration

This above circled equation yields solution of the form:

$$x(\tau) = A \cos(\omega_0 t) + B \sin(\omega_0 t) = A \cos(\tau) + B \sin(\tau),$$

didn't necessarily need this

or rewritten in terms of exponentials,  $x(\tau) = Ce^{i\tau} + De^{-i\tau}$

From diff eq standards, this general solution can be ~~of~~ of form

$$x(\tau) = Ae^{r\tau}$$

Plugging this back into scaled differential equation gives

$$r^2 Ae^{r\tau} + 2\gamma r Ae^{r\tau} + Ae^{r\tau} = 0$$

$$Ae^{r\tau} (r^2 + 2\gamma r + 1) = 0$$

$A=0$  is boring, same with  $e^{r\tau} = 0$ .

looking at  $r^2 + 2\gamma r + 1 = 0$  yields

$$r = -\gamma \pm \sqrt{\gamma^2 - 1}$$

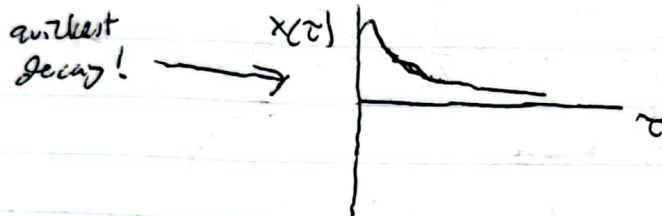
$$r_1 = -\gamma + \sqrt{\gamma^2 - 1}$$

$$r_2 = -\gamma - \sqrt{\gamma^2 - 1}$$

Considering the form earlier,  $x(\tau) = A_1 e^{r_1 \tau} + A_2 e^{r_2 \tau} = A_1 e^{-\gamma \tau + \tau \sqrt{\gamma^2 - 1}} + A_2 e^{-\gamma \tau - \tau \sqrt{\gamma^2 - 1}}$

The  $e^{-\gamma \tau}$  gives exponential decay. Note  $\gamma > 0$ .

critical damping:  $\gamma = 1$ , causes the  $\sqrt{\gamma^2 - 1}$  term to equal 0, leaving us with just  $x(\tau) = A_1 e^{-\gamma \tau} + A_2 \tau e^{-\gamma \tau}$



added since otherwise we can't have 2 initial conditions. Comes back to diff eq requirements. Can show  $\tau e^{-\gamma \tau}$  is also a solution.

over damping:  $\gamma > 1$ , causes the  $\sqrt{\gamma^2 - 1}$  to be  $> 0$ . This value is added/subtracted in the exponents. Exponential decay. More gradual than critical damping.



supposed to look exponential. My drawings skills aren't great

$$A_1 e^{-(\gamma + \sqrt{\gamma^2 - 1})\tau} + A_2 e^{-(\gamma - \sqrt{\gamma^2 - 1})\tau}$$

under damping:  $\gamma < 1$ .  $\sqrt{\gamma^2 - 1} = i\omega$ ,  $\omega = \sqrt{1 - \gamma^2}$ . let  $c = \sqrt{1 - \gamma^2}$

$$r_1 = -\gamma + i\omega, r_2 = -\gamma - i\omega$$

$$x(\tau) = A_1 e^{-\gamma \tau} e^{i\omega \tau} + A_2 e^{-\gamma \tau} e^{-i\omega \tau}$$

Considering rewriting imaginary exponentials with sines and cosines, this comes out to be

$$x(\tau) = B_1 e^{-\gamma \tau} \cos(c \cdot \tau) + B_2 e^{-\gamma \tau} \sin(c \cdot \tau), \text{ giving us oscillatory motion with exponential decay.}$$





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4d) With the driving force, our diff eq now looks like

$$m \frac{d^2 x}{dt^2} + b \frac{dx}{dt} + kx = F_0 \cos(\omega t)$$

We can again scale this using  $\omega_0 = \sqrt{\frac{k}{m}}$  and  $\tau = \omega_0 t$ .

Note that  $F_0 \cos(\omega t)$  can be written as  $F_0 \cos(\frac{\omega}{\omega_0} \tau)$  since  $\tau = \omega_0 t$ .

Using scaling the same as in part a with the exception of the driving force written above, we get

$$m \omega_0^2 \frac{d^2 x}{d\tau^2} + b \omega_0 \frac{dx}{d\tau} + kx = F_0 \cos(\frac{\omega}{\omega_0} \tau)$$

divide by  $\omega_0^2 m$

$$\Rightarrow \frac{d^2 x}{d\tau^2} + 2\gamma \frac{dx}{d\tau} + x = \frac{F_0}{m \omega_0^2} \cos(\frac{\omega}{\omega_0} \tau) \quad \text{where we have the same } \gamma = \frac{b}{2m\omega_0}$$

Define  $\tilde{\omega} = \frac{\omega}{\omega_0}$  and  $\tilde{F}_0 = \frac{F_0}{m \omega_0^2}$  to write as

$$\frac{d^2 x}{d\tau^2} + 2\gamma \frac{dx}{d\tau} + x = \tilde{F}_0 \cos(\tilde{\omega} \tau)$$

where  $\gamma_1$  and  $\gamma_2$  are the same as before

At this point we guess solution of the form  $x(\tau) = \underbrace{D \cos(\tilde{\omega} \tau - \delta)}_{x_p(\tau)} + \underbrace{C_1 e^{\gamma_1 \tau} + C_2 e^{\gamma_2 \tau}}_{\text{general solution}}$

making guess on this

Plug  $x_p(\tau)$  back into diff eq giv:

$$D [-\tilde{\omega}^2 \cos(\tilde{\omega} \tau - \delta) - 2\gamma \tilde{\omega} \sin(\tilde{\omega} \tau - \delta) + \cos(\tilde{\omega} \tau - \delta)] = \tilde{F}_0 \cos(\tilde{\omega} \tau)$$

must equal  $\tilde{F}_0$

must equal 0, since no sine

rewrite as

$$D [(-\tilde{\omega}^2 \cos \delta + 2\gamma \tilde{\omega} \sin \delta + \cos \delta) \cos(\tilde{\omega} \tau) + (-\tilde{\omega}^2 \sin \delta - 2\gamma \tilde{\omega} \cos \delta + \sin \delta) \sin(\tilde{\omega} \tau)] = \tilde{F}_0 \cos(\tilde{\omega} \tau)$$

match sine and cosine part. On RHS we have only cosine

Using orthogonality,  $D [-\tilde{\omega}^2 \cos \delta + 2\gamma \tilde{\omega} \sin \delta + \cos \delta] = \tilde{F}_0$

$$-\tilde{\omega}^2 \sin \delta - 2\gamma \tilde{\omega} \cos \delta + \sin \delta = 0$$

$$-\tilde{\omega}^2 \tan \delta - 2\gamma \tilde{\omega} = -\tan \delta \Rightarrow \tan \delta = \frac{2\gamma \tilde{\omega}}{1 - \tilde{\omega}^2}$$

rewrite  $\Rightarrow \sin \delta = \frac{2\gamma \tilde{\omega}}{\sqrt{4\gamma^2 \tilde{\omega}^2 + (1 - \tilde{\omega}^2)^2}}$  using trig equivalencies.

$$\cos \delta = \frac{(1-\tilde{\omega}^2)^2}{\sqrt{4\tilde{\gamma}^2\tilde{\omega}^2 + (1-\tilde{\omega}^2)^2}}$$

Put  $\sin \delta$  and  $\cos \delta$  back into  $\star$  equation.  $\Rightarrow$

$$D = \frac{\tilde{F}_0}{\sqrt{(1-\tilde{\omega}^2)^2 + 4\tilde{\omega}^2\tilde{\gamma}^2}}$$

So finally we have our solution  $x_p(\tau) = D \cos(\tilde{\omega}\tau - \delta)$

$$\text{when } \delta = \tan^{-1}\left(\frac{2\tilde{\gamma}\tilde{\omega}}{1-\tilde{\omega}^2}\right) \text{ and } D = \frac{\tilde{F}_0}{\sqrt{(1-\tilde{\omega}^2)^2 + 4\tilde{\omega}^2\tilde{\gamma}^2}}$$

what we expect to see this this particular solution dominating after sufficient time has passed, since the general solution  $c_1 e^{r_1 \tau} + c_2 e^{r_2 \tau}$  will exhibit exponential decay (recall  $r_1, r_2 = -\gamma \pm \sqrt{\gamma^2 - 1}$ ). since  $\gamma > 0$ , both  $r_1$  &  $r_2 < 0$ , so exponential decay.

we'll check for this in our numerical solution.