### Question 1

#### Part A)

Derivation of Lame-Emden Equation

$$\frac{dm(r)}{dr} = 4\pi r^2 \rho(r)$$

(1)

$$\frac{dp(r)}{dr} = -G\frac{m(r)\rho(r)}{r^2}$$

(2)

$$\frac{1}{\rho}\frac{dp(r)}{dr} = -G\frac{m(r)}{r^2}$$

(3)

Then

$$\frac{d}{dr}\left(\frac{1}{\rho}\frac{dp(r)}{dr}\right) = G\frac{m}{r^3} - G\frac{1}{r^2}\frac{dm}{dr}$$

(4)

$$\frac{d}{dr}\left(\frac{1}{\rho}\frac{dp(r)}{dr}\right) = -\frac{2}{\rho * r}\frac{dp}{dr} - 4\pi G\rho$$

(5)

Multiply both sides with r<sup>2</sup>, which results.

$$\frac{d}{dr} \left( \frac{r^2}{\rho} \frac{dp(r)}{dr} \right) = -4\pi G \rho$$

(6)

The equation of state (EOS) is the relation between P and  $\rho$ . Take an example from the ideal gas equation

$$PV = \frac{k_B}{\mu m_H} T \rho$$

(7)

Assume polytropic EOS

$$P = K\rho^{\gamma} = K\rho^{1 + \frac{1}{n}}$$

(8)

Where

$$\rho = \rho_c \theta^n$$

(9)

Which makes

$$P = K \rho_c^{1 + \frac{1}{n}} \theta^{n+1}$$

(10)

Insert equaiton 10 into 6.

$$\frac{1}{r^2}\frac{d}{dr}(r^2K\rho_c^{\frac{1}{n}}(n+1)\frac{d\theta}{dx}) = -4\pi G\rho_c\theta^n$$

(11)

Say that  $r = \alpha \, \xi$  , where

$$\alpha^2 = \frac{K\rho_c^{\frac{1}{n}-1}(n+1)}{4\pi G}$$

(12)

Then

$$\left(\frac{1}{\alpha^2}\frac{1}{\xi^2}\right)\frac{1}{\alpha}\frac{d}{d\xi}\left(\alpha^2\xi^2\alpha^2\frac{1}{\alpha}\frac{d\theta}{d\xi}\right)+\theta^n=0$$

(13)

This is nothing but Lame-Emden Equation:

$$\frac{1}{\xi^2}\frac{d}{d\xi}(\xi^2\frac{d\theta}{d\xi})+\theta^n=0$$

(14)

Derivation of Total Mass of the Star

$$M = \int_0^R 4\pi r^2 \rho dr$$

(15)

From  $\rho = \rho_c \theta^n$  and  $r = r_n \xi \rightarrow dr = r_n d\xi$ 

$$M = \int_0^R 4\pi r^2 \rho dr = 4\pi \rho_c \int_0^R \theta^n r^2 dr = 4\pi \rho_c \frac{r^3}{\xi^3} \int_0^{\xi} \theta^n \xi^2 d\xi$$
(16)

(16)

Than  $\frac{r^3}{\xi^3}=r_n^3$  and from Lane-Emden equation  $\frac{1}{\xi^2}\frac{d}{d\xi}(\xi^2\frac{d\theta}{d\xi})=-\theta^n \to \xi^2\theta^n=-\frac{d}{d\xi}(\xi^2\frac{d\theta}{d\xi})$  wiht that

$$M = 4\pi\rho_c \frac{r^3}{\xi^3} \int_0^{\xi} -\frac{d}{d\xi} (\xi^2 \frac{d\theta}{d\xi}) = 4\pi\rho_c r^3 (-\frac{1}{\xi} \frac{d\theta}{d\xi})$$

(17)

Which makes

$$M = 4\pi R^3 \left(-\frac{\theta'(\xi)}{\xi_n}\right) \tag{18}$$

#### Mass and Radius Proportionality.

The proportionality relation can be found by looking at central density( $ho_c$ ) term power such that

$$R = C_1 \rho_c^{\frac{1-n}{2n}} \tag{19}$$

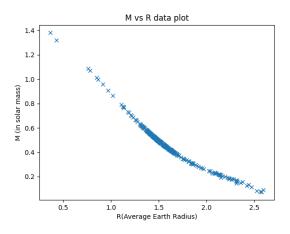
$$M = C_2 \rho_c^{\frac{3-n}{2n}} \tag{20}$$

Which results

$$M \propto R^{\frac{3-n}{1-n}} \end{matrix}$$

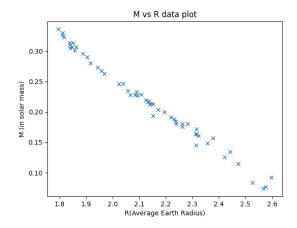
See the Mathematica file for other related derivations.

#### Part B)

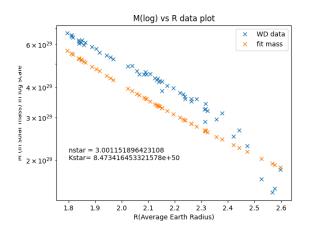


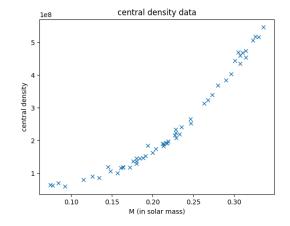
Part C)

I limit the M value to satisfy the low mass condition (M<0.34)

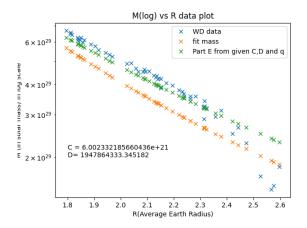


Then I make fit this data to find  $n_* and K_*$ 





## Part E)

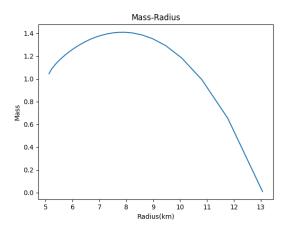


$$M_{ch} = 1.45832M$$

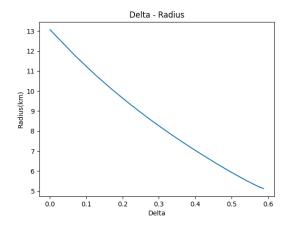
$$M_{ch}^{data} \approx 1.3 \mathrm{M}$$

### Question 2

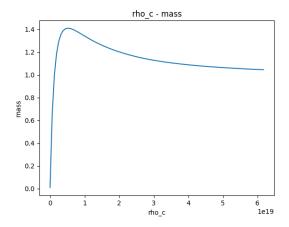
### Part A)



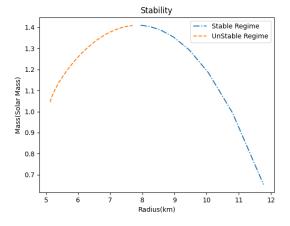
## Part B)



## Part C)



# Part D)



Part E)

See the Mathematica file