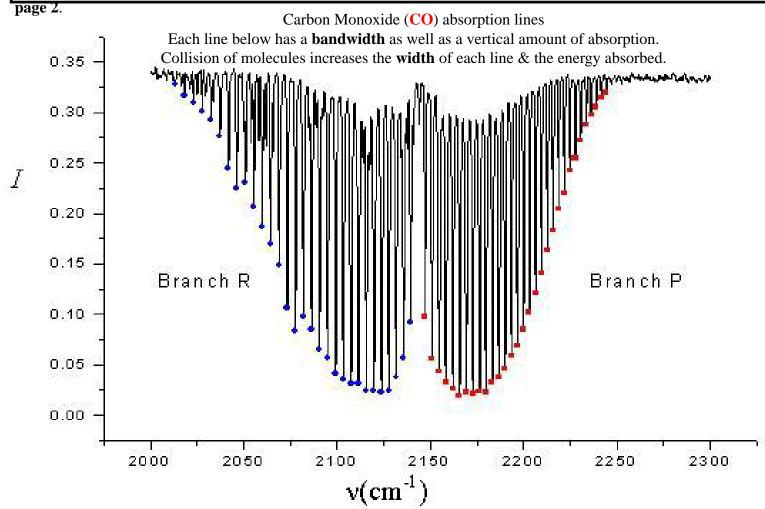


The Effect of Gas Mixtures on Absorption spectra of gases App. Note A82

Mixtures of gas compounds may result in the broadening of the absorption lines for each compound. This is similar to increased atmospheric pressure effects or Gas Law effects (App note A12).

In the field of Infrared Spectroscopy, the phenomenon of pressure broadening is well know and documented (see attached references on page 4). The mechanism of pressure broadening includes **natural broadening**, **collisional broadening**, and **doppler broadening**. Of these three, **collisional broadening** has the greatest effect on the width and shape of the spectroscopic transitions. See **definitions** of Pressure or Collisional Broadening & Gas Law on



Although **IR** absorption spectral responses are often shown as smooth curve responses (as is shown on page 2 with **CO**₂) they are more accurately shown as individual narrow frequency lines of absorption as shown in the plot of carbon monoxide above. The relatively narrow absorption lines created by the **CO** molecules are **broadened** by the **collisions** with **CO**₂ molecules. This may increase the apparent interference of **CO**₂ in the **CO** channel. The vertical length or intensity of the absorption depends upon the number of gas molecules in the optical path between the **IR** source and the **IR** detector/band pass filter. The Gas Laws affect the density and hence the number of molecules (see page 2). The impact is greater with heavy molecules. It increases the bandwidth of each line and hence the amount of energy absorbed by each line. See test data on page 3 & references on page 4.

Gas calibration with a mixture of compounds that will be measured will help to minimize this effect on accuracy. When we attempt to measure a gas compound using **NDIR** there is a **backgrond compound** like air with 71% nitrogen or pure nitrogen (N_2) with a **molecular weight** of 28. If you are trying to measure N_2 0 and you have a background of nitrogen the collision effect of the N_2 1 molecules gives a baseline effect on the absorption lines of N_2 2. If the background gas is N_2 3 with a molecular weight of 44 the molecular collisions will have more impact & a more significant effect on the absorption line broadening.

Gas Law definitions:

Boyle's Law: The proportionality between the pressure and volume of a gas may be expressed as follows: V = 1/P, or PV = 1/P.

V is proportional to 1/P or using a proportionality constant V = k/P or PV = kFor a given sample of gas at constant temperature the product of the pressure of the gas and its volume will always be the same value

Charles's Law: At constant pressure, the volume of a gas is directly proportional to its absolute, or Kevin temperature V - mT

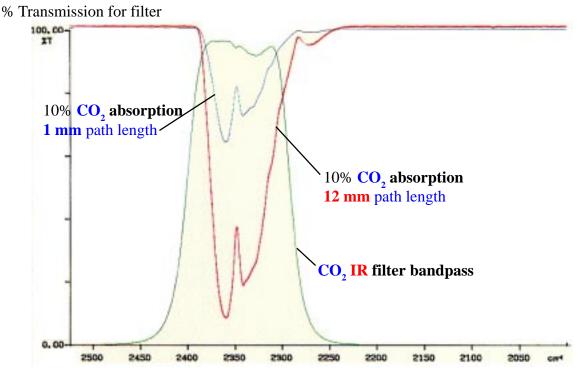
Gay-Lussac's Law:At constant volume, the pressure of a gas is directly proportional to its absolute, or Kevin temperature. $P_1/T_1 = P_2/T_2 = P_3/T_3 = P_n/T_n = k$

Combined Law Equation: $P_1V_1/T_1 = P_2V_2/T_2$

This equation shows there is a linear relationship when changes in absolute pressure, volume, or absolute temperature occur. Any change in either one of the three will affect either or both of the remaining ones.

Gas Density within a gas cell: In an NDIR gas cell the volume is held constant by the volume of the cell. The number of molecules between the IR source of energy and the IR band pass filter / IR detector (path length) is then determined by the absolute temperature and absolute pressure that effects the gas density (how close the molecules are to each other). See Application Note A12 for the quantified effects of tempearure & pressure on gas density.

Effect on Spectral line shape & width by purturbing gases: In addition to the gas density changes that result in more or less gas molecules of interest in the path betwen the IR source & detector, the type of background gas molecules also has an affect on the amount IR energy absorbed. The CO₂ absorption spectra shown below is a smooth approximation of the many individual narrow band absorption lines. The vertical depth of the absorption (inverse of the % transmission shown for the IR band pass filter) depends upon the path length and the gas density. The width of those individual absorption lines is affected by collisions with background gas molecules. Collisions with heavy background molecules cause the lines to get wider and more energy is absorbed.



Wave number is the reciprocal of the wavelength in centimeters (the center wavelength is about 4.24 µM)

See **Application Note A12** for Gas Law effects of pressure and temperature changes from the atmospheric pressure & temperature at gas calibration time. See **Application Note A71** for **NDIR** sensor Theory of Operation.

Table 1 shows the response of an **NDIR CO** sensor for 0 to 20% **CO** by volume with nitrogen as a background gas and then a mixture of nitrogen and **20% carbon dioxide** flow blended in as a background gas. At 0.0% **CO** there would be 20% **CO**₂, at 5% **CO** there would be 15% **CO**₂, at 10% **CO** there would be 10% **CO**₂, at 15% **CO** there would be 5% **CO**₂, and at 20% **CO** there would be 0% **CO**₂ as the chart shows. The 0-1V output response to **CO** of the sensor is increased by the collision of **CO**₂ molecules with the **CO** molecules. **Graph 1** shows the amount of increased % **CO** due to these molecular collisions.

Table 2 shows the 0-1V response of an **NDIR CO** sensor to 0 to 20% **CO** by volume with nitrogen as a background gas. Then the same sensors response to a graduated mixture of 20% **CO** and **100% carbon dioxide** as a background gas. **Graph 2** is a plot of the 0 to 20% **CO** with nitrogen versus 100% CO_2 background. We got nearly an identical response when we used a background of **100% propane** (C_3H_8). Propane has no absorption spectra is this filter's band pass and we were careful to use additional out of band blocking teflon filters and an additional **CO** IR filter.

additional out of band blocking teflon filters and an additional CO IR filter.														
Table	1	Balance		Balance		Balance	1	Table 2	Balance		Balance		Balance	
		N2		20% CO ₂		20% CO ₂	Difference	Table 2	N2		100% CO ₂	& N ₂	100% CO ₂	Difference
	%CO	0-1V	%CO ₂	0-1V	0-1V dif	%CO	%CO	%CO	0 - 1 V	%CO ₂	0-1V	0-1V dif	%CO	%CO
	0.0	0.003	20.00	0.003	0.000	0.00	0.00	0.0	0.000	100.0	0.000	0.000	0.00	0.00
	$0.5 \\ 1.0$	$0.091 \\ 0.167$	19.50 19.00	$0.096 \\ 0.176$	0.005 0.010	$\begin{array}{c} 0.52 \\ 1.06 \end{array}$	0.02	0.5	0.103	97.5	0.129	0.026	0.65	0.15
	1.5	0.107	18.50	0.176	0.010	1.65	0.00	1.0	0.183 0.248	95.0 92.5	0.233 0.315	0.050 0.067	1.39 2.15	0.39 0.65
	2.0	0.288	18.00	0.304	0.016	2.19	0.19	2.0	0.301	90.0	0.313	0.081	2.13	0.03
	2.5		17.50	0.355	0.019	2.71	0.21	2.5	0.345	87.5	0.437	0.092	3.83	1.33
	3.0	0.379	17.00	0.400	0.021	3.27	0.27	3.0	0.383	85.0	0.484	0.101	4.67	1.67
	3.5	0.415	16.50	0.438	0.023	3.85	0.35	3.5	0.417	82.5	0.525	0.108	5.48	1.98
	4.0		16.00	0.471	0.024	4.41	0.41	4.0	0.447	80.0	0.560	0.113	6.20	2.20
	4.5 5.0	$0.476 \\ 0.502$	15.50 15.00	0.501 0.528	$0.025 \\ 0.026$	4.98 5.55	0.48 0.55	4.5 5.0	0.475 0.501	77.5 75.0	0.593 0.623	0.118 0.097	6.92 7.63	2.42 2.63
	5.5	0.526	14.50	0.552	0.026	6.09	0.59	5.5	0.526	72.5	0.651	0.125	8.30	2.80
	6.0	0.549	14.00	0.575	0.026	6.65	0.65	6.0	0.550	70.0	0.678	0.128	9.00	3.00
	6.5	0.570	13.50	0.596	0.026	7.15	0.65	6.5	0.574	67.5	0.703	0.129	9.68	3.18
	7.0	0.590	13.00	0.616	0.026	7.65	0.65	7.0	0.596	65.0	0.727	0.131	10.36	3.36
	7.5	0.610	12.50	0.635	0.026	8.15	0.65	7.5	0.618 0.639	$62.5 \\ 60.0$	0.749 0.770	0.131 0.131	11.30 11.69	3.80 3.69
	8.0	0.629	12.00 11.50	0.654 0.673	0.025 0.024	8.65 9.14	0.65 0.64	8.5	0.659	57.5	0.790	0.131	12.33	3.83
	9.0	0.668	11.00	0.691	0.024	9.62	0.62	9.0	0.678	55.0	0.809	0.131	12.93	3.93
	9.5	0.687	10.50	0.710	0.023	10.10	0.60	9.5	0.697	52.5	0.826	0.129	13.50	4.00
	0.0	0.706	10.00	0.728	0.022	10.60	0.60	10.0	0.714	50.0	0.843	0.129	14.33	4.33
	0.5	0.725	9.50	0.746	0.022	11.06	0.56	10.5 11.0	0.732 0.748	47.5 45.0	0.858 0.872	0.126 0.124	14.53 15.33	4.03
	11.0	0.743	9.00	0.764	0.021	11.58	0.58	11.0	0.748	42.5	0.872	0.124	15.33	3.97
	11.5	0.762	8.50	0.782	0.020	12.05	0.55	12.0	0.780	40.0	0.898	0.118	15.93	3.93
	12.0	$0.780 \\ 0.797$	8.00 7.50	$0.799 \\ 0.816$	0.019 0.018	12.55 13.05	0.55 0.55	12.5	0.796	37.5	0.910	0.114	16.37	3.87
	13.0	0.814	7.00	0.832	0.013	13.53	0.53	13.0	0.811	35.0	0.921	0.110	16.77	3.77
	13.5	0.831	6.50	0.847	0.017	14.03	0.53	13.5	0.826	32.5	0.932	0.106	17.20	3.70
1	4.0	0.846	6.00	0.862	0.016	14.54	0.54	14.0 14.5	0.842 0.857	$\frac{30.0}{27.5}$	0.942 0.952	0.100 0.095	17.60 18.33	3.60 3.83
	4.5	0.861	5.50	0.876	0.015	15.03	0.53	15.0	0.871	25.0	0.962	0.091	18.47	3.47
	15.0	0.875	5.00	0.889	0.014	15.54	0.54	15.5	0.886	22.5	0.971	0.085	18.85	3.35
	15.5	$0.888 \\ 0.900$	4.50	0.901 0.913	0.013 0.012	16.03 16.52	0.53	16.0	0.900	20.0	0.979	0.079	19.20	3.20
	6.5	0.912	3.50	0.913	0.012	17.00	0.50	16.5	0.914	17.5	0.986	0.072	19.50	3.00
	7.0	0.923	3.00	0.934	0.010	17.46	0.46	17.0 17.5	0.927 0.939	$15.0 \\ 12.5$	0.991 0.996	0.064 0.057	19.65 19.87	2.65
1	17.5	0.934	2.50	0.943	0.009	17.90	0.40	18.0	0.951	10.0	0.999	0.037	19.97	1.97
	8.0	0.946	2.00	0.953	0.008	18.34	0.34	18.5	0.963	7.5	1.000	0.037	20.00	1.50
	18.5	0.957	1.50	0.964	0.006	18.75	0.25	19.0	0.974	5.0	1.001	0.027	20.00	1.00
	19.0	$0.970 \\ 0.985$	$1.00 \\ 0.50$	$0.975 \\ 0.987$	0.005 0.003	19.17 19.56	$0.17 \\ 0.06$	19.5	0.986	2.5	1.001	0.015	20.00	0.50
	20.0		0.00	1.002	0.003	20.02	0.00	20.0	1.000	0.0	1.000	0.000	20.00	0.00
								['] M	odel 20	15SPI-C	O N ve	ersus CC	backgro	und
	 %	Difference	Э					20			2		1 0	,,,,,,,
0.	7			2015S	PI-CO 14	-Aug-01		-					a partie	
0.	· -						' ' -	-					2	
g	F		90	, , , , , , , , , , , , , , , , , , , 			- 1	-		Gr	aph 2	ø	~ /	_
plend)	.6		-					-		GI	apn 2	P		4
- ဝိ	ţ		1		M		1	15 9				-9°	/	
	ţ		1			8000 N	1	0,			9	0		4
% 0.	.5					P					ø			4
s S	ŀ		Ĭ			þ	+	00			ø			4
ıs.	, F	6	ļ			\		-			90			4
e 0.	.4	/					\	₹10		J. J.	5			
8	ţ	ø	Gr	aph 1			b 1	,		ø				4
2 0.	.3		- 51	apn 1				ģ		ø		Ideal res	oonse	
CO Difference (pure CO versus 20%	1 E	ϕ					\	%CO blended with 100%		8	/ i	f CO ₂ mo	ecules	
9	F	/					٩	9		/		had no		
e 0 .	.2	<i></i> &						9 5	ø		o	n CO abs	sorption	
fer.	ţ	I					φ ‡	* 3	ø					
ΔĬ	ţ	ľ					\		2/					1
O 0.	.1	· 							8/					1
%	- F	∌					g ²		6 /					1
	ړ						7	0	·					1
	0		5	10)	15	20	0		5	10		15	20
	9		•	0/ 0			20	· ·		-	24 0 0			20

Application Note A82 page 3 of 4

%CO

Revised: 10-Dec-01 © 1994-2001 VALTRONICS

Conclusion:

The test data on page 3 clearly shows the upscale effect of pressure or collisional broadening on the ${\bf CO}$ response from the heavier ${\bf CO}_2$ molecule collisions with weight of 44 the instead of a nitrogen (${\bf N}_2$) background with a **molecular** weight of 28. This increased absorption from the **broadening** of the absorption line **bandwidth** has similar effect as increasing the atmospheric pressure (increases gas density) which **lengthens** the absorption lines due to a larger number of molecules in the optical path. See Application Note A12 for more detail about the Gas Law effects.

If you **gas calibrate** with the same **mixture** of **compounds** that you will usually be measuring, yon can minimize any adverse effect on accuracy. If you are measuring two gases like **CO** and **CO**₂ with separate sensors you can cross compensate the **CO** response from the measured **CO**₂ value based on the interference measurement.

In many cases, the concentration of the interfering compound is not large enough to significantly effect your accuracy.

REFERENCES: Note: The URL links to WEB sites may change in the future so you may have to do a web search:

http://www.chem.vt.edu/chem-ed/ac-basic.html - Good reference for basics of IR absorption.

Beer-Lambert Law: The Beer-Lambert law (also called the Beer-Lambert-Bouguer law or simply Beer's law) shows the relationship between absorbance and concentration of an absorber of electromagnetic radiation.

See **page 25** of this **pdf**:

http://www.ime-conrad.com/OEM/CPTCO2/co2guide.pdf

Cardiopulmonary Technologies, Inc.

OEM Compact CO, Waveform Analyzer

MEASUREMENT OF PRESSURE-BROADENING LINE WIDTHS OF NO (nitric oxide) FROM THE FITTING OF LMR SPECTRA WITH CORRECTIONS OF INSTRUMENTAL BROADENING

Modern Physics Letters B, Vol. 14, No. 11 (2000) 401-407 © World Scientific Publishing Company http://www.worldscinet.com/journals/mplb/14/1411/S0217984900000537.html

JIE-LI LIN, YU-YAN LIU, HONG-PING LIU, YUAN-QING QUO, XIAO-YONG LIU, FENG-YAN LI and JIN-RUI LI State Key Laboratory of Magnetic Resonance and Atomic and Molecular Physics, Wuhan Institute of Physics and Mathematics, Chinese Academy of Sciences, Wuhan 430071, People's Republic of China

For absorption lines, the **distortion of spectral lines** by experimental spectrometers must be considered in measuring **linewidth** which has been investigated by many authors. While in the derivative spectrum, we managed to give the corrections of instrumental broadening and distortion to obtain accurate linewidths. We developed a universal fitting program which was explored together with the method of correction in the fundamental band spectra of LMR of NO. The uncorrected halfwidths were compared with the corrected values, which were obtained from the program. The results of comparison are very satisfactory and demonstrate their useful application to analyze experimental spectra.

Applied Physics B Lasers & Optics http://www.neo.no/applphysb98.pdf

Horiba talks about a correction for line stretching due to other gas molecules being present.

http://global.horiba.com/semicon_e/ir-150/ir-150_08.htm

The reason for the IR-150's alternative **gas correction** Infrared property absorption band internal -C-H stretching vibration standard vibrations from 2960~2850 cm-1 (3.38~3.51) appear as if they are near the absorption spectra.