Thermal Physics: entropy, temperature, free energy

*Entropy熵: logarithm of number of (stationary quantum) states of a system of particles. 量子定态

For a system in a stationary quantum state, all observable physical properties (energy) and number of particles are <u>independent of time</u>.

Each quantum state has a definite energy. States with identical energies belong to the same energy level.

*Multiplicity/Degeneracy of an energy level: number of quantum states with very nearly the same energy.

The energy of a system is the total energy of all particles(K+P+interactions between particles)

A quantum state of system is a state of all particles.

State of Particle: QM state of one element of system of particles (SHO n, spin 1/1)

Microstate of System: a single possible configuration of particle system.

Macrostate of System: gross parameters of, and constraints on system. ???

*Orbital: quantum states of a one-particle system.

energy of quantum state s of the N particle system $\varepsilon_s(N)$: $\varepsilon_s(N) = \frac{\hbar^2}{2M} \left(\frac{\pi}{L}\right)^2 (n_x^2 + n_y^2 + n_z^2)$

Binary Model System: magnetic moment +m/-m >>magnet points up/down.

- **Total model system magnetic moment/net system magnetization M**: (N magnets each of magnetic moment m) M=Nm-2km =2sm ∈[-Nm,Nm] (possible values of M: N+1)
- ◆Total number states of system: 2^N.

Enumeration of States:

N: even, number of states

 N_{\uparrow} :# of states with magnets up $N_{\uparrow}=\frac{1}{2}N+s$

 N_{\downarrow} :# of states with magnets down $N_{\downarrow}=\frac{1}{2}N-s$

Spin Excess 2s= N_{\uparrow} - N_{\downarrow} s∈[-½ N_{\uparrow} 2 N_{\uparrow} 2 N_{\uparrow} 3

net magnetic moment M=2sm

All possible choice of system state:

$$(\uparrow + \downarrow)^{N} = \sum_{t=0}^{N} \frac{N!}{(N-t)!t!} \uparrow^{N-t} \downarrow^{t} = \sum_{s=-\frac{1}{2}N}^{\frac{1}{2}N} \frac{N!}{(\frac{1}{2}N+s)!(\frac{1}{2}N-s)!} \uparrow^{(\frac{1}{2}N+s)} \downarrow^{(\frac{1}{2}N-s)}$$
(where t=½N-s)

Multiplicity Function g(N,s): number of states having the same value of s/fixed spin excess 2s.

$$g(N,s) = \frac{N!}{(\frac{1}{2}N+s)!(\frac{1}{2}N-s)!} = \frac{N!}{N_{\uparrow}!N_{\downarrow}!}$$

Sharpness: g(N,s) is peaked sharply about s=0

$$\ln g(N,s) = \ln N! - \ln(\frac{1}{2}N + s)! - \ln(\frac{1}{2}N - s)!$$

[Stirling Approximation]
$$\ln N! = N \ln N - N$$
 $N! = \sqrt{2\pi N} N^N e^{[-N]}$ for N>>1

[Approximation]
$$\ln(1+x) = x - \frac{1}{2}x^2 + \dots$$
 for x<<1

$$g(N,s) \cong g(N,0)e^{(-2s^2/N)} = \sqrt{\frac{2}{\pi N}}2^N e^{(-2s^2/N)}$$

for s/N<<1 with N>>1 Gaussian Approximation

g(N,s) center/maximum at s=0 gmax=g(N,0)

s²=½N value of g=(1/e)g_{max} fractional width of distribution

N1 distribution sharped peak

Average Value of s2

 $P(N,s)=g(N,s)/2^N$ $< f>=\Sigma f(s)P(N,s)$ $f(s)=s^2$ change Σ to $\int (-\infty \text{ to } \infty \infty)$

$$\langle s^2 \rangle = \frac{\int_{-\infty}^{\infty} ds s^2 g(N, s)}{\int_{-\infty}^{\infty} ds g(N, s)} = \frac{1}{4} N$$

$$\langle s^2 \rangle = \frac{1}{4}N$$

mean square spin excess $<(2s)^2>=N$

root mean square spin excess $\sqrt{\langle (2s)^2 \rangle} = \sqrt{N}$

$$J = \frac{\sqrt{\langle (2s)^2 \rangle}}{N} = \frac{1}{\sqrt{N}}$$

Fractional Fluctuation of s2 in 2s

Energy of Binary Magnetic System:

Potential Energy $U = -\vec{m} \cdot \vec{B}$

$$U(s) = \sum_{i=1}^{N} U_{i} = -\vec{B} \cdot \sum_{i=1}^{N} \vec{m}_{i} = -2smB = -MB$$

$$\Delta \epsilon = U(s) - U(s-1) = 2mB$$

Multiplicity function for Harmonic Oscillators

Fundamental Assumption (of thermal phy): a closed system is equally likely/probable to be in any of the quantum states accessible to it.

Closed System: constant energy, number of particles, volume, relevant external parameters(which influence system, gEB)

Accessible (quantum state): properties compatible with system physical specification.(in range) state non-accessible: 0 probability P(s)=0

Accessible State of a System: a system microstate compatible with the state specification of energy, number of particles, etc.-- a microstate compatible with the system macrostate.

 $\delta U/U << 1 \delta N/N << 1$

Suppose a closed system with g equally likely accessible quantum states, s be a general state label P(s) = 1/g

 $< X > = \sum_{s} X(s)P(s) = \sum_{s} X(s)1/g$

Ensemble Average: average value of X

(s:accessible state)

(property average over ensemble system)

Ensemble of Systems: group of systems constructed alike.

Thermal Contact: freely energy transfer between 2 system

Most Probable Configuration: system S₁,S₂ combine S=S₁+S₂ energy U=U₁+U₂ number of particle $N=N_1+N_2$ spin excess $s=s_1+s_2$ P34????? energy of combined system U(s)=U(s₁)+U(s₂)=-2mBs constant so $s=s_1+s_2$ constant multiplicity function of combines system: (number of accessible states of combine system) $g(N,s) = \sum_{i=1}^{n} g_1(N_1, s_1)g_2(N_2, s_2) = \sum_{i=1}^{n} g_1(N_1, s_1)g_2(N_2, s - s_1)$ $S_1 \in [-\frac{1}{2}N_1, \frac{1}{2}N_1]$ $N_1 < N_2$

Configuration: set of all states with specified values of s₁ and s₂.

Different configuration of combines system are characterized by different values of s₁ **Most Probable Configuration**: configuration for which g_1g_2 maximum with $s_1=s_1$ hat/caret. Note: For system with large N, product q_1q_2 is a very sharply peaked function of s_1 .

Thermal Equilibrium Values:

Average of a physical quantity over all accessible configurations >> average over MPC

1. Two spin systems in the thermal contact:

MPC: $\frac{s_1}{N_1} = \frac{s_2}{N_2}$ $\frac{\hat{s}_1}{N_1} = \frac{\hat{s}_2}{N_2}$

2. Thermal Equilibrium:

The most probable division of total energy is the one for which the combined system has the maximum number of accessible states.

$$g(N,U) = \sum_{U_1} g_1(N_1, U_1) g_2(N_2, U - U_1)$$

Entropy: $\sigma(N,U) = \ln g(N,U)$ [dimensionless] g:# of accessible states of system

$$\left[\frac{\partial \sigma_1(N_1, U_1)}{\partial U_1}\right]_{N_1} = \left[\frac{\partial \sigma_2(N_2, U_2)}{\partial U_2}\right]_{N_2} \text{ in equilibrium}$$

 $\frac{1}{T} = k_B \left(\frac{\partial \sigma}{\partial U} \right)_N$ Boltzmann Constant k_B=1.381×10⁻²³ J/K (Absolute) **Temperature** (in Kelvin):

 $\frac{1}{\tau} = \left(\frac{\partial \sigma}{\partial U}\right)_{N}$ [J] $\tau = k_{B}T$ Conventional **Entropy**: S classic Fundamental Temperature:

 $\frac{1}{T} = \left(\frac{\partial S}{\partial U}\right)_{N}$ definition

 $\Delta \sigma = \left(\frac{\partial \sigma_1}{\partial U_1}\right)_{N_1} (-\Delta U) + \left(\frac{\partial \sigma_2}{\partial U_2}\right)_{N_2} (\Delta U) = \left(-\frac{1}{\tau_1} + \frac{1}{\tau_2}\right) \Delta U$ \tau1>\tau2 so \Delta \sigma 0 Total entropy change

Law of Increase of Entropy: σ_{final}

$$\sigma_{final} \simeq \ln(g_1 g_2)_{\max} \ge \sigma_{initial} = \ln(g_1 g_2)_{initial}$$

total entropy always increases when 2 systems are brought into thermal contact.

entropy of a closed system tends to remain constant or increase when a constraint internal to system is removed.

Increase entropy: add particles, add energy, increase volume, decompose molecules, let a linear polymer curl up

The thermal equilibrium values of physical properties of a system are defined as averages over all states accessible when the system is in contact with a large system/reservoir.

If first system also large, thermal equilibrium properties 可以只考虑 MPC

Laws of Thermodynamics:

0Zeroth: τ1=τ2 τ2=τ3 then τ1=τ3 [$\partial ln(g_1)/\partial U_1]_{N_1}$ =[$\partial ln(g_2)/\partial U_2]_{N_2}$ If 2 systems are each in thermodynamic equilibrium with a third system, they must be in equilibrium with each other.

- 1) First: heat is a form of energy. conservation of energy
- 2Second: increase of entropy

removing internal constraints in a closed system tends to cause the system entropy to remain constant or increase.

③Third: entropy of a system approaches a constant value as the temperature approaches (absolute) zero. (system goes to ground state)

Boltzmann Distribution:

one system S & reservoir R, $U_0=U_R+U_S$ $U_R=U_0-\epsilon_s$

 $g_{R\times S}=g_{R\times 1}=g_F$

$$\frac{P(\varepsilon_1)}{P(\varepsilon_2)} = \frac{g_R(U_0 - \varepsilon_1)}{g_R(U_0 - \varepsilon_2)} = \frac{e^{\sigma_R(U_0 - \varepsilon_1)}}{e^{\sigma_R(U_0 - \varepsilon_2)}} = e^{\Delta\sigma_R} \\ \text{where Taylor Expansion} \\ \sigma_R(U_0 - \varepsilon) = \sigma_R(U_0) - \varepsilon \left(\frac{\partial\sigma_R}{\partial U}\right)_{V,N} = \sigma_R(U_0) - \frac{\varepsilon}{\tau}$$

$$\Delta\sigma_{R} = -\frac{\varepsilon_{1} - \varepsilon_{2}}{\tau} \qquad \qquad \frac{P(\varepsilon_{1})}{P(\varepsilon_{2})} = \frac{e^{-\frac{\varepsilon_{1}}{\tau}}}{e^{-\frac{\varepsilon_{2}}{\tau}}} = e^{\left(-\frac{\Delta\varepsilon}{\tau}\right)}$$

Boltzmann Factor e^(-Δε/τ)

Ratio of probability of finding system in a single quantum state 1 to state 2.

$$Z(\tau) = \sum e^{(-\frac{\mathcal{E}_s}{\tau})}$$

Partition Function:

sum over Boltzmann factor for all state s of system

$$P(\varepsilon_s) = \frac{e^{\frac{(-\varepsilon_s)}{\tau}}}{Z} \qquad (\Sigma P(\varepsilon_s) = 1)$$

$$U = <\varepsilon> = \sum \varepsilon_{S} P(\varepsilon_{S}) = \frac{\sum \varepsilon_{S} e^{-\varepsilon_{S}/\tau}}{Z} = \tau^{2} \frac{\partial \ln Z}{\partial t}$$

(Thermal/Ensemble) Average energy of system

Reversible Process: system remains infinitesimally close to the equilibrium state at all times during the process.

Pressure:
$$p = -\left(\frac{\partial U}{\partial \mathcal{V}}\right)_{\sigma} = \tau \left(\frac{\partial \sigma}{\partial \mathcal{V}}\right)_{\upsilon}$$

number of states σ in the ensemble is unchanged in reversible compression

Thermodynamic Identity: $\tau d\sigma = dU + pdV$ $dU = \tau d\sigma - pdV = TdS - pdV$

TdS heat added to system,-pdV work done on system

Helmholtz Free Energy: $F = U - \tau \sigma = -\tau \ln Z$

$$\sigma = -\left(\frac{\partial T}{\partial \tau}\right)$$

$$\frac{\partial F}{\partial \tau} \bigg|_{\mathcal{V}} \qquad p = -\left(\frac{\partial F}{\partial \mathcal{V}}\right)_{\tau} = -\left(\frac{\partial U}{\partial \mathcal{V}}\right)_{\tau} + \tau \left(\frac{\partial \sigma}{\partial \mathcal{V}}\right)_{\tau}$$

differential of F

For an ideal monatomic gas of N atoms of spin zero,

$$Z_N = (n_0 V)^N / N! ,$$

if $n = N/V \ll n_0$. The quantum concentration $n_0 \equiv (M\tau/2\pi\hbar^2)^{3/2}$. Further,

Ideal Monatomic Gas:

$$pV = N\tau;$$
 $\sigma = N\left[\log(n_Q/n) + \frac{5}{2}\right];$ $C_V = \frac{3}{2}N.$

Quantum-Mechanical Simple Harmonic Oscillators(QM-SHOs) frequency ω, state of oscillator specified by quantum number n:

$$Z = \sum_{s=0}^{\infty} e^{-s\frac{\hbar\omega}{\tau}} = \frac{1}{1 - e^{-\frac{\hbar\omega}{\tau}}}$$

Partition Function

$$\langle s \rangle = \frac{1}{\exp(\hbar\omega/\tau) - 1}.$$

Planck Distribution Function:

Thermal average energy

$$\langle \varepsilon \rangle = \langle s \rangle \hbar \omega = \frac{\hbar \omega}{\exp(\hbar \omega / \tau) - 1}$$
 $U = \sum_{n} \langle \varepsilon_{n} \rangle = \sum_{n} \frac{\hbar \omega_{n}}{\exp(\hbar \omega_{n} / \tau) - 1}$

$$U = \sum_{n} \langle \varepsilon_{n} \rangle = \sum_{n} \frac{\hbar \omega_{n}}{\exp(\hbar \omega_{n}/\tau) - 1}$$

on:
$$\frac{U}{V} = \frac{\pi^2}{15\hbar^3 c^3} \tau^4$$

Stefan-Boltzmann law of radiation:
$$\frac{U}{V} = \frac{\pi^2}{15\hbar^3c^3}\tau^4 \; ,$$

$$u_\omega = \frac{\hbar}{\pi^2c^3}\frac{\omega^3}{\exp(\hbar\omega/\tau)-1}.$$

Planck Radiation Law: