



Lecture 4: Condensation

1. Condensation of corundum

We acknowledge and respect the *ləkʷəŋən* peoples on whose traditional territory the university stands and the Songhees, Esquimalt and *WSÁNEĆ* peoples whose historical relationships with the land continue to this day.



Practice Problem: Condensation of Corundum from the Solar Nebula

Q1: Calculate the temperature that Corundum (Al_2O_3) begins condensing from the solar nebula using the following values (assume no other reactions):

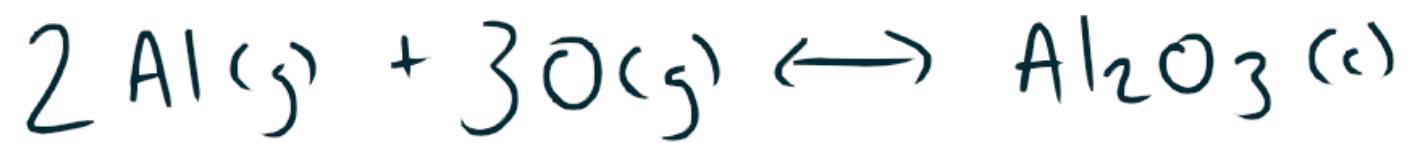


- Solar abundance (molar abundance) of Al: 8.51×10^4
- Solar abundance (molar abundance) of O: 2.36×10^7
- Solar abundance (molar abundance) of H: 2.6×10^{10}
- Pressure in the nebula: 10^{-3} atm
- Gas constant (R): 8.314 J/mol K
- ΔG° (standard free energy of reaction) for condensation of Al_2O_3 : -1.23×10^6 J/mol

Q2: At what temperature will this reaction finish condensing all of the Aluminum in the nebula?



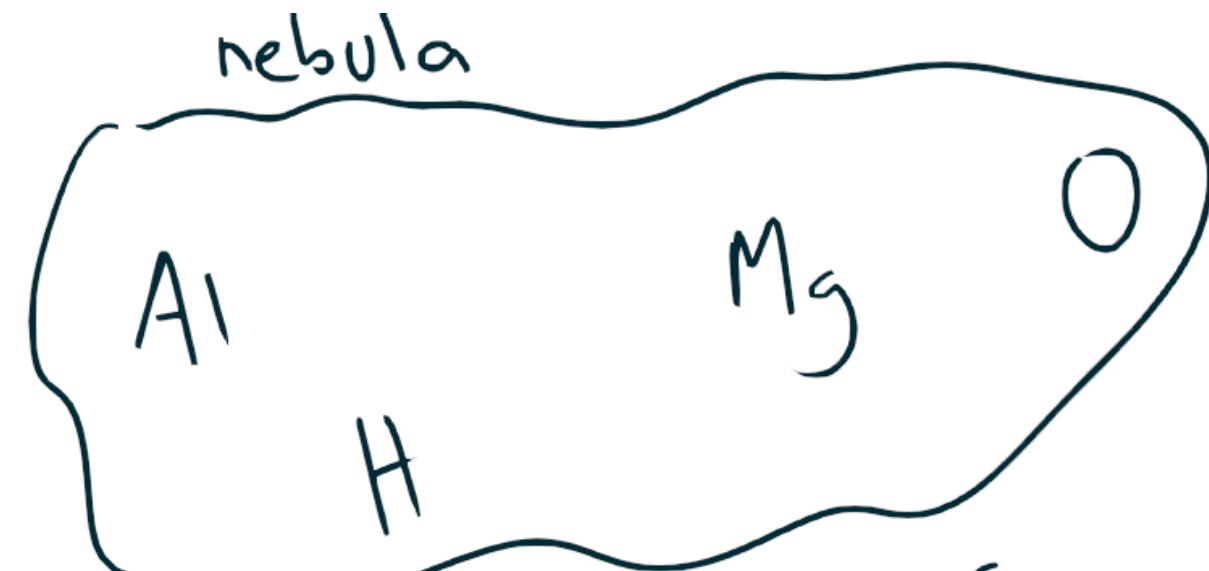
Condensation of corundum (ruby/Sapphire)



$$Q = \frac{\{\text{Al}_2\text{O}_3\}}{\{\text{Al}\}^2 \{\text{O}\}^3} = \frac{1}{P_{\text{Al}}^2 \cdot P_{\text{O}}^3}$$

at high T, low P safe to
assume ideal gas

$$\{x\} = P_x$$



chemistry = Sun

$$P_T = 10^{-3} \text{ atm}$$

assumption: $P_T = P_{\text{H}_2}$

$$\frac{P_x}{P_{\text{H}_2}} = \frac{N_x}{N_{\text{H}_2}}$$

$$N_{\text{H}_2} = \frac{1}{2} N_{\text{H}}$$

$$P_{\text{Al}} = \frac{N_{\text{Al}}}{\frac{1}{2} N_{\text{H}}} \cdot P_{\text{H}_2} = \frac{2 N_{\text{Al}}}{N_{\text{H}}} \cdot P_T$$

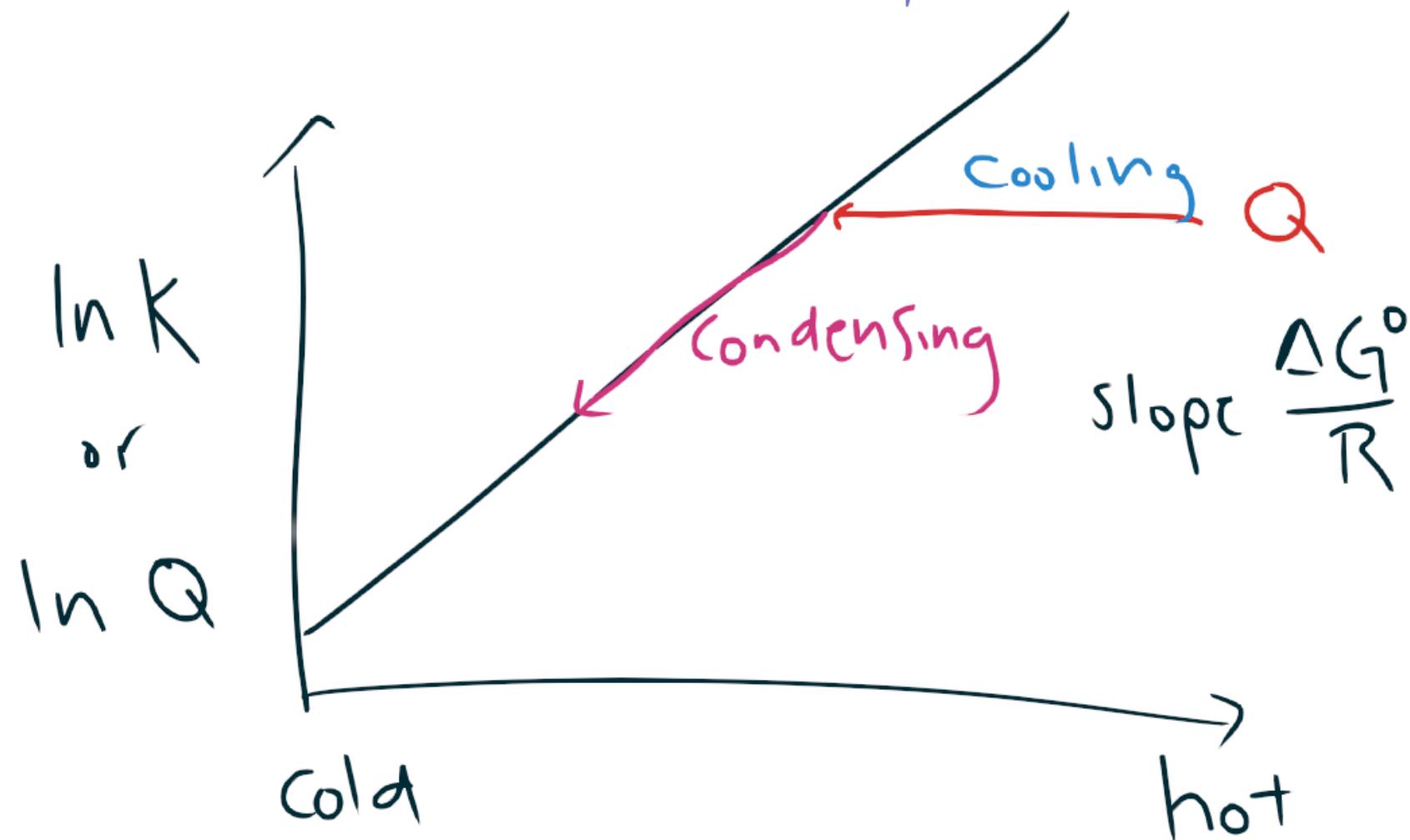


$$\Delta G^\circ = -RT \ln K$$

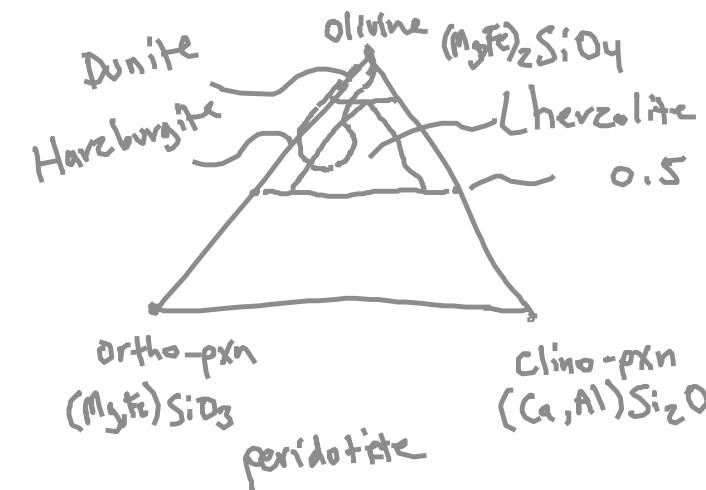
$$-\frac{\Delta G^\circ}{RT} = \ln K \rightarrow \text{ct+} :$$

this sign corresponds to reaction direction

$$x = \frac{-1}{T} \quad y = \ln K$$
$$x - \frac{\Delta G^\circ}{R} = y$$



Olivine Solid Solution Phase Diagram.



Lever Rule:
(mass balance)

unknown fraction

$$q \cdot L + b S = \text{system}$$

$$q + b = 1 \quad \text{mass balance}$$

$$b = 1 - q$$

$$q L + (1-q) S = \text{system}$$

$$q 0.25 + (1-q) 0.8 = 0.7$$

