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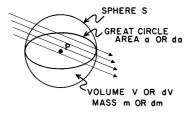
Quantities and concepts in interaction theory and dosimetry

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Ionizing radiation field

- Field of ionizing particles, where the particles may have a directional- and energy distribution
- Radiation field striking a small sphere:



 Number of particles N striking the sphere is proportional to dose



Fluence

• Fluence *F*: number of particles *dN* striking the sphere per unit area *da*:

$$\Phi = \frac{dN}{da}$$
 (da is the great circle area)

- The small sphere defines a point in space
- Fluence is as an expectation value; *N* is in reality a stochastic quantity
- For a radiation field through a medium, the fluence varies due to absorption, scattering and creation of new particles $\rightarrow \Phi = \overrightarrow{\Phi}(r)$

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Fluence 2

• The fluence may vary in time – the fluence rate is defined as:

$$\Phi_{t} = \frac{d\Phi}{dt} = \frac{d^{2}N}{dtda}$$

Thus

$$\Phi = \int_{0}^{t_0} \Phi_t dt$$

• For a time-independent field:

$$\Phi = \Phi_t \Delta t$$

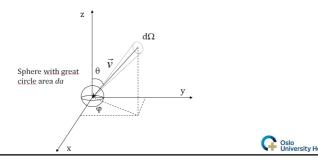


Fluence 3

 The radiation field may have a an energy and directional dependence. The differential fluence is:

$$\Phi_{\rm T} = \frac{{\rm d}\Phi}{{\rm d}T}$$
, $\Phi_{\Omega} = \frac{{\rm d}\Phi}{{\rm d}\Omega}$ (${\rm d}\Omega = \sin\theta {\rm d}\theta {\rm d}\phi$)

• Φ_T is the number of particles per energy and area in the energy interval [T, T+dT] striking the sphere



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Energy fluence

- How much energy 'strikes' the sphere?
- The energy fluence is defined as:

$$\Psi = \int_{0}^{T_{\text{max}}} T\Phi_{T} dT$$

• For a monoenergetic field:

$$\Psi = T\Phi = T\frac{dN}{da}$$

• Differentiated:

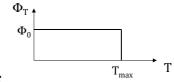
$$\Psi_{\scriptscriptstyle T} = \frac{d\Psi}{dT} = T\Phi_{\scriptscriptstyle T} \ , \ \Psi_{\scriptscriptstyle \Omega} = \frac{d\Psi}{d\Omega} = \int\limits_{\scriptscriptstyle 0}^{\scriptscriptstyle T_{max}} T \frac{d\Phi_{\scriptscriptstyle T}}{d\Omega} dT$$

Fluence vs energy fluence

Differential fluence with respect to energy is constant up to T_{max}:

$$\Phi_{T} = \Phi_{0} \implies \Phi = \int_{0}^{T_{max}} \Phi_{T} dT$$

$$\Rightarrow \underline{\Phi = T_{max} \Phi_{0}}$$

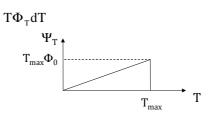


Differential energy fluence is:

Differential energy fluence is:
$$\Psi_{T} = T\Phi_{T} \implies \Psi = \int_{0}^{T_{max}} \Psi_{T} dT = \int_{0}^{T_{max}} T\Phi_{T} dT$$

$$\Rightarrow \underbrace{\Psi = \frac{1}{2} T_{max}^{2} \Phi_{0}}_{T_{max}}$$

$$T_{max} \Phi_{0}$$



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Average particle energy in field

- Differential fluence and energy fluence are distribution functions
- Average energy defined as:

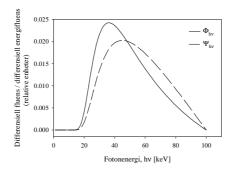
$$< T>_{\Phi} = \frac{\int\limits_{0}^{T_{max}} T\Phi_{T} dT}{\int\limits_{0}^{T_{max}} \Phi_{T} dT} = \frac{\Psi}{\Phi}$$

$$< T>_{\Psi} = \frac{\int\limits_{0}^{T_{max}} T\Psi_{T} dT}{\int\limits_{0}^{T_{max}} \Psi_{T} dT} = \frac{\int\limits_{0}^{T_{max}} T^{2}\Phi_{T} dT}{\int\limits_{0}^{T_{max}} \Psi_{T} dT} \neq < T>_{\Phi}$$



Fluence vs energy fluence 2

- X-ray spectrum is either differential fluence or differential energy fluence
- Problem: is often given as "intensity"



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Fluence vs energy fluence 3

• In our example:

$$<$$
 T $>_{\Phi} \approx 48 \text{ keV}$
 $<$ T $>_{\Psi} \approx 54 \text{ keV}$

Always ask what the unit of the ordinate is in X-ray (or e.g. e⁻) spectra!



Indirectly ionizing radiation

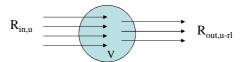
- Indirectly ionizing radiation experience few interactions, but releases relatively large amounts of energy in each interaction
- Example: photons, neutrons
- Secondary charged particles (electrons most relevant) will deposit the transferred energy over a short distance
- How large are the energy transfers from e.g. photons to matter for a given volume element?
- The energy-mass budget is important!

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Energy transferred, ε_{tr}

• A photon field with total energy $R_{in,u}$ enters a volume, while $R_{out,u-rl}$ is the energy leaving the volume:



• Energy transferred:

$$\varepsilon_{tr} = R_{in.u} - R_{out.u-rl} + \Sigma Q$$

• ϵ_{tr} is the total energy transferred from photons to electrons, and is the sum of all kinetic energy released



Energy transferred, ε_{tr} 2

- u-rl: uncharged minus radiative losses; radiative losses by secondary electrons should not be included
- ϵ_{tr} is a stochastic quantity
- ΣQ: energy from conversion of rest mass or *vice versa*
- Example, pair production $\Sigma Q = -2m_e c^2$

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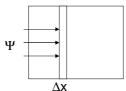


KERMA

• Kinetic Energy Release per MAss:

$$K = \frac{d\epsilon_{\rm tr}}{dm} \quad \text{ unit: [J/kg]}$$

- K is the expectation value of the energy transferred per unit mass in a point of interest
- Consider monoenergetic photons (quantum energy hv) passing a thin layer:



S: cross section of photon field



KERMA 2

- Probability per unit lenght for photon interaction multiplied with fraction of energy transferred: μ_{tr}
- Total energy transferred to electrons: $\varepsilon_{tr} = N(h\nu)\mu_{tr}\Delta x$
- Energy fluence for monoenergetic photons:

$$\Psi = (h\nu)\Phi = \frac{N(h\nu)}{S}$$

• KERMA becomes: $K = \frac{\varepsilon_{tr}}{m} = \frac{N(hv)\mu_{tr}\Delta x}{\rho V} = \frac{N(hv)\mu_{tr}\Delta x}{\rho S\Delta x}$ $= \frac{\Psi \frac{\mu_{tr}}{\rho}}{\rho}$

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KERMA 3

- KERMA is determined by the energy fluence and the mass energy transfer coeffecient
- For a distribution of photons:

$$K = \int_{0}^{h\nu_{max}} \Psi_{h\nu} \frac{\mu_{tr}}{\rho} d(h\nu)$$

• Remember that μ_{tr}/ρ is dependent on the photon energy and atomic number of the absorber



Components of KERMA

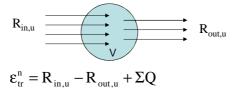
- Kerma includes all kinetic energy given to secondary electrons, and this energy may be lost by:
 - Collisions
 - Radiative losses
- Kerma may be divided into two components: $K{=}K_c{+}K_r$
- K_c: collision Kerma; provides a measure of the energy loss per unit mass from photons resulting in collisional losses for secondary electrons!

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Net energy transferred ϵ_{tr}^n

• ε_{tr}^{n} is defined as:



- R_{out,u} is all photon energy leaving the volume element (including brehmsstrahlung)
- ϵ_{tr}^n is thus the total kinetic energy of secondary electrons which is not lost as brehmsstrahlung



Collision Kerma

Is defined by:

$$K_{c} = \frac{d\varepsilon_{tr}^{n}}{dm}$$

May take radiative losses into account by defining the quntity g; the fraction of kinetic energi lost as brehmsstrahlung

$$K_c = K(1-g) = \Psi \frac{\mu_{tr}}{\rho} (1-g)$$

- $K_{c} = K(1-g) = \Psi \frac{\mu_{tr}}{\rho} (1-g)$ Definition: $\frac{\mu_{en}}{\rho} = \frac{\mu_{tr}}{\rho} (1-g)$
- $\mu_{en}/\rho\colon mass\ energy\ absorption\ coeffecient$

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Collision Kerma 2

K_c is thus:

$$K_c = \Psi \frac{\mu_{en}}{\rho}$$

- Generally: K_c<K
- Special case: Low energy photons releases low energy electrons in an absorber of low atomic number Z. Radiative losses are insignificant, and g≈0 and $K_c \approx K$



Energy imparted and absorbed dose

• Look at all energy transport (both charged and uncharged particles) through the volume of interest:

$$\begin{aligned} &R_{\text{in},u} + R_{\text{in},c} & & \\ & & & \\ & & \\ & & \\ & & \\ & \epsilon = R_{\text{in},u} + R_{\text{in},c} - R_{\text{out},u} - R_{\text{out},c} + \Sigma Q \end{aligned}$$

• Absorbed dose is (at last) defined as:

$$D = \frac{d\epsilon}{dm}$$
 unit: [Gy] = [J/kg]

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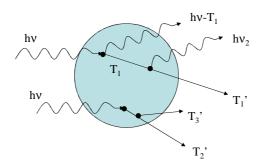
Absorbed dose

- The absorbed dose is all energy imparted to the volume per mass
- May not be directly related to photon interaction coefficients
- However, in some cases the dose may be approximated by K_c



ε_{tr} , ε_{tr}^{n} , ε : example

• Two photons interacts in a volume of interest ($\Sigma Q=0$):



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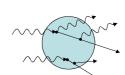


$\boldsymbol{\epsilon}_{tr}$, $\boldsymbol{\epsilon}_{tr}^{n}$, $\boldsymbol{\epsilon}$: example 2

• <u>Photon 1:</u>

$$\begin{split} & \epsilon_{tr} = R_{in,u} - R_{out,u-rl} = h\nu - (h\nu - T_1) = T_1 \\ & \epsilon_{tr}^{\ n} = R_{in,u} - R_{out,u} = h\nu - (h\nu - T_1) - h\nu_2 = T_1 - h\nu_2 \\ & \epsilon = & R_{in,u} + R_{in,c} - R_{out,u} - R_{out,c} \\ & = h\nu + 0 - (h\nu - T_1) - h\nu_2 - T_1' = T_1 - h\nu_2 - T_1' \end{split}$$





• <u>Photon 2:</u>

$$\begin{split} & \epsilon_{tr} = \ h\nu - 0 = h\nu \\ & \epsilon_{tr}^{\ n} = h\nu - 0 = h\nu \\ & \epsilon_{tr} = h\nu + 0 - \ T_2 - T_3 = h\nu - \ T_2' - T_3' \end{split}$$



Charged particle equilibrium (CPE)

 Photons enter a volume V, which includes a smaller volume v:

Photons V

- CPE: Number of charged particles of a given type and energy entering v is equal to the number of particles of the same type and energy leaving
- Certain conditions must be fullfilled:
 - V must be homogeneous
 - Photon attenuation must be negligible

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CPE 2

CPE:

- If CPE is present, $R_{in,c} = R_{out,c}$
- Energy imparted:

$$\epsilon = R_{_{\mathrm{in},u}} + R_{_{\mathrm{in},c}} - R_{_{out,u}} - R_{_{out,c}} = R_{_{\mathrm{in},u}} - R_{_{out,u}} = \epsilon_{\mathrm{tr}}^{^{n}}$$

• In this case, absorbed dose equals collision Kerma:

$$D = \frac{\epsilon}{m} = \frac{\epsilon^{\text{CPE}}}{m} = K_c = \Psi \frac{\mu_{en}}{\rho}$$



Absorbed doses under CPE

- K_c , and thus dose, is given by $\Psi \mu_{en}/\rho$, and is thus proportional to the interaction probability in a given absorber
- Two different absorbers A og B placed in the same point in a radiation field:

$$\frac{D_{_{A}}}{D_{_{B}}} = \frac{\Psi\left(\frac{\mu_{en}}{\rho}\right)_{_{A}}}{\Psi\left(\frac{\mu_{en}}{\rho}\right)_{_{B}}} = \frac{\left(\frac{\mu_{en}}{\rho}\right)_{_{A}}}{\left(\frac{\mu_{en}}{\rho}\right)_{_{B}}}$$

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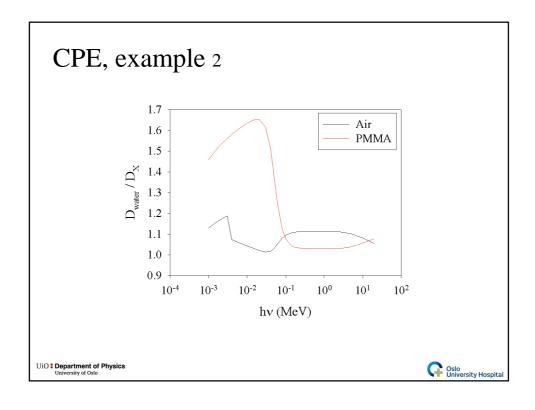


CPE, example

- Two small volumes of water and air is placed in same point in a radiation field (1 MeV photons) where CPE exists. What is the dose ratio?
- Use tabulated values for μ_{en}/ρ (Attix):

$$\begin{split} &\mu_{en}/\rho(water) = 0.0309 \\ &\mu_{en}/\rho(air) = 0.0278 \\ &\rightarrow D(air) \ / \ D(water) = 0.90 \end{split}$$





CPE, example

- A (real) patient was treated with 60 kV X-rays with the dose reported as 400 R x 3. Estimate the dose to tissue (water).
- ✓ R is the Roentgen unit, giving a measure of exposure (liberated charges in air; see later lectures on ionometry).
- ✓ 1 R = 0.00877 Gy dose to *air*
- ✓ 60 kV X-rays have mean X-ray energy of roughly 30 keV (strongly depends on filtration; see lectures on X-ray production)



CPE, example

- ✓ Assuming CPE, Dair/Dw is proportional to μ_{en}/ρ ratio
- ✓ Use NIST tables -> $\mu_{en}/\rho(air) = 0.1537$ $\mu_{en}/\rho(water) = 0.1557$
- ✓ Dose:

$$\frac{D_{w}}{D_{air}} = \left(\frac{\mu_{en}}{\rho}\right)_{air}^{w}$$

$$\Rightarrow D_{w} = D_{air} \left(\frac{\mu_{en}}{\rho}\right)_{air}^{w} = 3 \times 400R \times 0.00877Gy / R \times \frac{0.1557}{0.1537}$$

$$D_{w} = 10.7Gy$$

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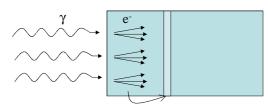
CPE, problems

 When the photon energy increases, the range of the secondary electrons increases more than the photon pathlenght

Photon energy(MeV)	Photon attenuation (%) in water within the range of a secondary electron
0.1	0
1	1
10	7
30	15



CPE, problems 2



e⁻ with long range contributes to the dose at the layer. Photon beam significantly attenuated between the interaction point and the layer – fewer electrons are generated in the layer than what was generated upstream.

- Thus: $R_{in,c} > R_{out,c}$ and: $\Rightarrow \varepsilon = R_{in,u} + R_{in,c} - R_{out,u} - R_{out,c} > \varepsilon_{tr}^{n}$ $\Rightarrow D > K_{c}$
- Most relevant for high photon energies

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TCPE

- Transient Charged Particle Equilibrium: electrons originating from upstream contributes to the dose, while the photon contribution ($R_{in,u}$ - $R_{out,u}$) is given by the collision Kerma
- Assumption: absorbed dose propotional to K_c

$$\begin{array}{c} \text{TCPE} \\ D = K_c (1 + f_{\text{TCPE}}) \\ \text{INDIRECTLY} \\ \text{INDIRECTLY} \\ \text{IONIZING} \\ \text{RADIATION} \\ \text{RADIATION} \\ \text{O DEPTH IN MEDIUM} \\ \end{array}$$