

University of Illinois at Chicago
Department of Physics

Quantum Mechanics
Qualifying Examination

January 4, 2010.
9:00 am – 12:00 pm

Question 1

In heavy (large Z) hydrogen-like atoms, where to a very good approximation the reduced mass μ is equal to the electron mass m_e , relativistic corrections to the electron's kinetic energy need to be taken into account due to its large orbital velocity.

- a) Show that the first-order relativistic correction to the electron's kinetic energy is

given by
$$\hat{H}_1 = -\frac{1}{2m_e c^2} \left(\frac{\mathbf{p}^2}{2m_e} \right)^2.$$

- b) Verify that $\left\langle \frac{1}{r} \right\rangle = \frac{Z}{a_0}$ for the 1s state of the hydrogen-like atom.

- c) Similarly, show that $\left\langle \frac{1}{r^2} \right\rangle = \frac{2Z^2}{a_0^2}$ for the 1s state of the hydrogen-like atom.

- d) Using the fact that the unperturbed Hamiltonian, $\hat{H}_0 = \frac{\mathbf{p}^2}{2m_e} - \frac{Ze^2}{4\pi\epsilon_0 r}$, and the results from parts (a), (b) and (c), evaluate the first-order perturbation due to the relativistic correction to the electron's kinetic energy.

① a) Since $E = \sqrt{p^2 c^2 + m_e^2 c^4} = K + m_e c^2$, where K is the kinetic energy

$$\Rightarrow K = \sqrt{p^2 c^2 + m_e^2 c^4} - m_e c^2 = m_e c^2 \sqrt{1 + \frac{p^2 c^2}{m_e^2 c^4}} - m_e c^2$$

$$\approx m_e c^2 \left(1 + \frac{p^2}{2 m_e^2 c^2} - \frac{1}{8} \left(\frac{p^2}{m_e^2 c^2} \right)^2 + \dots \right) - m_e c^2$$

$$\approx \frac{p^2}{2 m_e} - \frac{1}{8} \frac{p^4}{m_e^3 c^2} \approx \frac{p^2}{2 m_e} - \frac{1}{2 m_e c^2} \left(\frac{p^2}{2 m_e} \right)^2$$

∴ First-order correction $\hat{H}_1 = -\frac{1}{2 m_e c^2} \left(\frac{p^2}{2 m_e} \right)^2$

b) Since $\langle r | \phi_{100} \rangle = 2 \left(\frac{z}{a_0} \right)^{3/2} \exp \left[-\frac{zr}{a_0} \right]$, we have

$$\left\langle \frac{1}{r} \right\rangle = \int_0^\infty r^2 dr \cdot \frac{1}{r} \cdot 4 \left(\frac{z}{a_0} \right)^3 e^{-2zr/a_0}$$

$$= 4 \left(\frac{z}{a_0} \right)^3 \int_0^\infty dr r e^{-2zr/a_0} = 4 \left(\frac{z}{a_0} \right)^3 \frac{a_0^2}{4 z^2}$$

∴ $\left\langle \frac{1}{r} \right\rangle = \frac{z}{a_0}$

① c) Similarly $\langle \frac{1}{r^2} \rangle$ for $\langle r | \phi_{100} \rangle$ is given by

$$\begin{aligned} \langle \frac{1}{r^2} \rangle &= \int_0^\infty r^2 dr \cdot \frac{1}{r^2} \cdot 4 \left(\frac{z}{a_0} \right)^3 e^{-2zr/a_0} \\ &= 4 \left(\frac{z}{a_0} \right)^3 \left[-\frac{a_0}{2z} e^{-2zr/a_0} \right]_0^\infty \\ \therefore \langle \frac{1}{r^2} \rangle &= \underline{\underline{2 \left(\frac{z}{a_0} \right)^2}} \end{aligned}$$

d) As $\hat{H}_0 = \frac{p^2}{2m_e} - \frac{ze^2}{4\pi\epsilon_0 r}$, we have

$$\hat{H}_1 = -\frac{1}{2m_e c^2} \left(\frac{p^2}{2m_e} \right)^2 = -\frac{1}{2m_e c^2} \left(\hat{H}_0 + \frac{ze^2}{4\pi\epsilon_0 r} \right)^2$$

$$\Rightarrow \langle \phi_{100} | \hat{H}_1 | \phi_{100} \rangle$$

$$= -\frac{1}{2m_e c^2} \langle \phi_{100} | \hat{H}_0^2 + 2\hat{H}_0 \frac{ze^2}{4\pi\epsilon_0 r} + \left(\frac{ze^2}{4\pi\epsilon_0 r} \right)^2 | \phi_{100} \rangle$$

$$= -\frac{1}{2m_e c^2} \left[E_{100}^2 + \frac{ze^2 E_{100}}{2\pi\epsilon_0} \langle \frac{1}{r} \rangle + \left(\frac{ze^2}{4\pi\epsilon_0} \right)^2 \langle \frac{1}{r^2} \rangle \right]$$

$$\therefore E_{100} = -13.6 z^2 \text{ eV}$$

① d) contd.

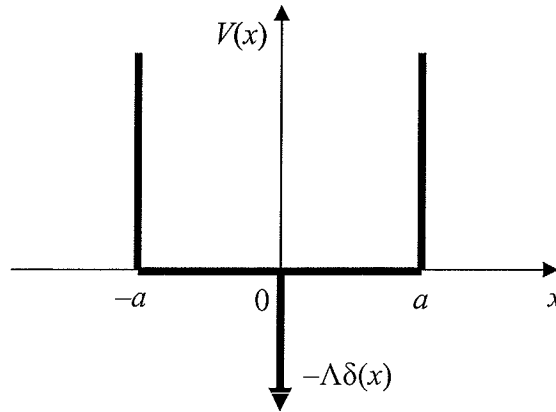
Since $E_{100} = -\frac{m_e}{2\hbar^2} \left(\frac{ze^2}{4\pi\epsilon_0} \right)^2$, $\langle \frac{1}{r} \rangle = \frac{z}{a_0}$ and $\langle \frac{1}{r^2} \rangle = 2 \left(\frac{z}{a_0} \right)^2$ with $a_0 = \frac{4\pi\epsilon_0\hbar^2}{m_e e^2}$, we have

$$\begin{aligned} \langle \phi_{100} | \hat{H} | \phi_{100} \rangle &= -\frac{1}{2m_e c^2} \left(\frac{m_e^2 z^4 e^8}{\hbar^4 \epsilon_0^4} \right) \left[\frac{1}{64} - \frac{1}{16} + \frac{1}{8} \right] \\ &= -\frac{5m_e z^4 e^8}{128 c^2 \hbar^4 \epsilon_0} \\ &= -\frac{5}{2m_e c^2} (E_{100})^2 \end{aligned}$$

Question 2

Consider an *attractive* delta-function potential $-\Lambda\delta(x)$, where the parameter Λ characterizes the strength of the delta-function, positioned at the center of an infinite square well of width $2a$; that is, a potential given by

$$\begin{aligned} V(x < -a) &= \infty \\ V(-a < x < 0) &= 0 \\ V(x = 0) &= -\Lambda\delta(x) \\ V(0 < x < a) &= 0 \\ V(x > a) &= \infty \end{aligned}$$



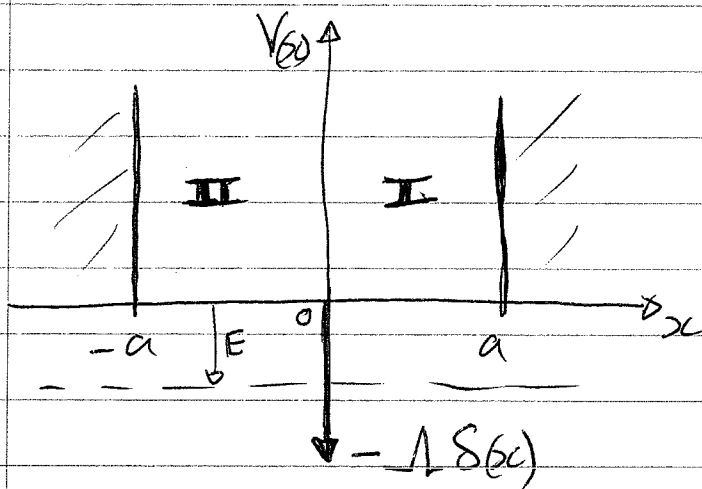
- a) Show that the equation determining the energy E of the eigenstate bound to the delta-function is given by

$$\tanh(\kappa a) = \frac{\hbar^2 \kappa}{m\Lambda},$$

where $\hbar\kappa = \sqrt{2m|E|}$ and m is the mass of the particle.

- b) What is the minimum strength of the delta-function for which a state with $E < 0$ exists?

(2) a)



$$\psi_I = Ae^{Kx} + Be^{-Kx}$$

$$\psi_{II} = Ce^{Kx} + De^{-Kx}$$

$$K = \sqrt{\frac{2m|E|}{\hbar^2}}$$

Boundary conditions at $x = \pm a$ give

$$\psi_I(a) = 0 = Ae^{Ka} + Be^{-Ka} \Rightarrow A = -Be^{-2Ka} \quad (1)$$

$$\psi_{II}(-a) = 0 = Ce^{-Ka} + De^{Ka} \Rightarrow C = -De^{+2Ka} \quad (2)$$

Boundary conditions at the δ -function give

$$\psi_I(0) = \psi_{II}(0) \Rightarrow A + B = C + D \quad (3)$$

$$\left. \frac{\partial \psi_I}{\partial x} \right|_{x=0} - \left. \frac{\partial \psi_{II}}{\partial x} \right|_{x=0} = -\frac{2m\Lambda}{\hbar^2} \psi(0)$$

$$\Rightarrow K(A - B) - K(C - D) = -\frac{2m\Lambda}{\hbar^2} (A + B) \quad (4)$$

Inserting (1) and (2) into (3) and (4) to eliminate A and C gives

② a) contd.

$$B(1 - e^{-2Ka}) = D(1 - e^{2Ka})$$

$$-KB(1 + e^{-2Ka}) + KD(1 + e^{2Ka}) = -\frac{2mL}{\hbar^2} B(1 - e^{-2Ka})$$

$$\Rightarrow -K(1 + e^{-2Ka}) + K \frac{(1 + e^{2Ka})(1 - e^{-2Ka})}{(1 - e^{2Ka})} = -\frac{2mL}{\hbar^2} (1 - e^{-2Ka})$$

$$\Rightarrow 2K \cosh(Ka) = \frac{2mL}{\hbar^2} \sinh(Ka)$$

$$\circ \circ \tanh(Ka) = \frac{\hbar^2 K}{mL}$$

b) For a state with $E < 0$ to exist, the above equation must have a solution, other than at $K=0$. This means that

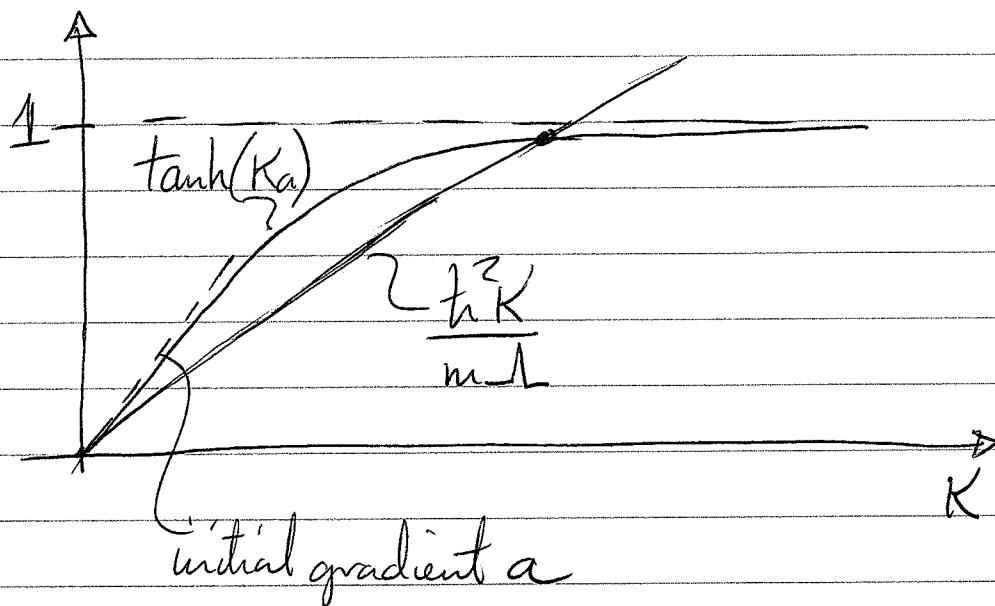
$$\frac{\partial}{\partial K} (\tanh(Ka)) > \frac{\hbar^2}{mL} \quad @ \quad K=0$$

$$\Rightarrow a > \frac{\hbar^2}{mL}$$

(2) b) contd.

So, $\lambda > \frac{\hbar^2}{ma}$ for state with $E < 0$.

Graphically, this is self-evident:



Question 3

Quantum dots are important material systems in nanoscience.

Consider an electron of charge e and mass m_e confined in an idealized quantum dot with $V(r) = 0$ for $r < a$ and $V(r) = \infty$ for $r > a$; that is, an infinite spherical well of radius a . The eigenstates of this potential are given by

$$\langle \mathbf{r} | \psi_{nlm} \rangle = C_{nl} j_l \left(\frac{s_{nl} r}{a} \right) Y_{lm}(\theta, \phi),$$

where $Y_{lm}(\theta, \phi)$ is the spherical harmonic associated with the usual angular momentum quantum numbers l and m , and C_{nl} is the normalization constant for the l^{th} spherical Bessel function j_l with roots s_{nl} . The first three spherical Bessel functions with $\rho = s_{nl}r/a$ are

$$j_0(\rho) = \frac{\sin \rho}{\rho} \quad j_1(\rho) = \frac{\sin \rho}{\rho^2} - \frac{\cos \rho}{\rho} \quad j_2(\rho) = \left(\frac{3}{\rho^3} - \frac{1}{\rho} \right) \sin \rho - \frac{3}{\rho^2} \cos \rho$$

and the first few roots are given below:

s_{nl}	$l = 0$	$l = 1$	$l = 2$	$l = 3$	$l = 4$
$n = 1$	π	4.49	5.76	6.99	9.36
$n = 2$	2π	7.73	9.10	10.42	
$n = 3$	3π				

The spherical Bessel functions also satisfy the relation

$$\int_0^a r^2 dr j_l \left(\frac{s_{nl} r}{a} \right) j_l \left(\frac{s_{n'l'} r}{a} \right) = \frac{a^3}{2} [j_{l-1}(s_{nl})]^2 \delta_{nn'}.$$

- What are the energies of the lowest four eigenstates?
- Evaluate the normalization constants C_{10} and C_{11} .

- Consider the state $\langle \mathbf{r} | \psi \rangle = A \left[\left(\frac{\sin \rho}{\rho^2} - \frac{\cos \rho}{\rho} \right) \sin \theta \cos \phi + \frac{\sin \rho}{\rho} \right]$, where A is a constant.

With what probability is the ground state energy measured? And what is the probability of measuring a z -component of the angular momentum with a value $+\hbar$?

- In the presence of a strong homogeneous magnetic field \mathbf{B} , the energy states split – the Zeeman effect. Neglecting spin (the normal Zeeman effect), the additional contribution to the Hamiltonian can be written as $H_1 = \frac{e}{2m_e} \mathbf{L} \cdot \mathbf{B}$. Amongst the four lowest states

identified in part (a), at what magnetic field strength will states originating from one energy level overlap energetically with those from another?

③ a) Since the states are given as $j_l(k_{nl}r)$ with $k_{nl} = s_{nl}/a$, the wave vector, the state energies will be

$$E_{nl} = \frac{k_{nl}^2 \hbar^2}{2m_e} = \frac{\hbar^2 s_{nl}^2}{2m_e a^2}$$

Hence, the first 4 are:

$$n=1, l=0 : E_{10} = \frac{\pi^2 \hbar^2}{2m_e a^2} = 4.93 \left(\frac{\hbar^2}{m_e a^2} \right) \dots \text{ground state}$$

$$n=1, l=1 : E_{11} = 10.08 \left(\frac{\hbar^2}{m_e a^2} \right)$$

$$n=1, l=2 : E_{12} = 16.95 \left(\frac{\hbar^2}{m_e a^2} \right)$$

$$n=2, l=0 : E_{20} = \frac{4\pi^2 \hbar^2}{2m_e a^2} = 19.74 \left(\frac{\hbar^2}{m_e a^2} \right)$$

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b) Since the Y_{lm} are already normalized, we need only evaluate the integral

$$1 = |C_{nl}|^2 \int_0^a r^2 dr \left| j_l\left(\frac{s_{nl}r}{a}\right) \right|^2$$

(3b) contd. For C_{10} : $1 = |C_{10}|^2 \int_0^a r^2 dr \frac{\sin^2(\frac{\pi r}{a})}{(\frac{\pi r}{a})^2} = |C_{10}|^2 \frac{a^3}{2\pi^2}$

$$\Rightarrow C_{10} = \frac{\pi}{a} \sqrt{\frac{2}{a}} = \frac{4.443}{a\sqrt{a}}$$

For C_{11} : $1 = |C_{11}|^2 \int_0^a r^2 dr j_1^2\left(\frac{s_{11}r}{a}\right) = \frac{a^3}{2} [j_0(s_{11})]^2$

... using given relation.

Now; $j_0(s_{11}) = \frac{\sin(4.49)}{4.49} = -0.217$

$$\Rightarrow C_{11} = \frac{6.51}{a\sqrt{a}}$$

c) $\psi = A \left[\underbrace{\left(\frac{\sin \theta}{r^2} - \frac{\cos \theta}{r} \right)}_{j_1\left(\frac{s_{11}r}{a}\right)} \sin \theta \cos \phi + \underbrace{\frac{\sin \theta}{r}}_{j_0\left(\frac{s_{10}r}{a}\right)} \right]$

$$\frac{1}{2} \sin \theta \cdot (e^{i\phi} + e^{-i\phi})$$

$$= \frac{1}{2} \sqrt{\frac{8\pi}{3}} (\gamma_{1-1} - \gamma_{11})$$

③ c) contd

So, we clearly have

$$\psi = \alpha C_{11} j_1\left(\frac{s_{11} r}{a}\right) (Y_{1-1} - Y_{11}) + \beta C_{10} j_0\left(\frac{s_{10} r}{a}\right) Y_{00}$$

with

$$\alpha = \sqrt{\frac{8\pi}{3}} \frac{A}{C}$$

$$\beta = \sqrt{4\pi} \frac{A}{C_{10}}$$

$$\text{as } 1 = \sqrt{4\pi} Y_{00}.$$

Now, we also know that $2|\alpha|^2 + |\beta|^2 = 1$, thus

$$1 = \frac{8\pi}{6} \frac{A^2}{C_{11}^2} + 4\pi \frac{A^2}{C_{10}^2} = 4\pi A^2 \left(\frac{1}{3C_{11}^2} + \frac{1}{C_{10}^2} \right)$$

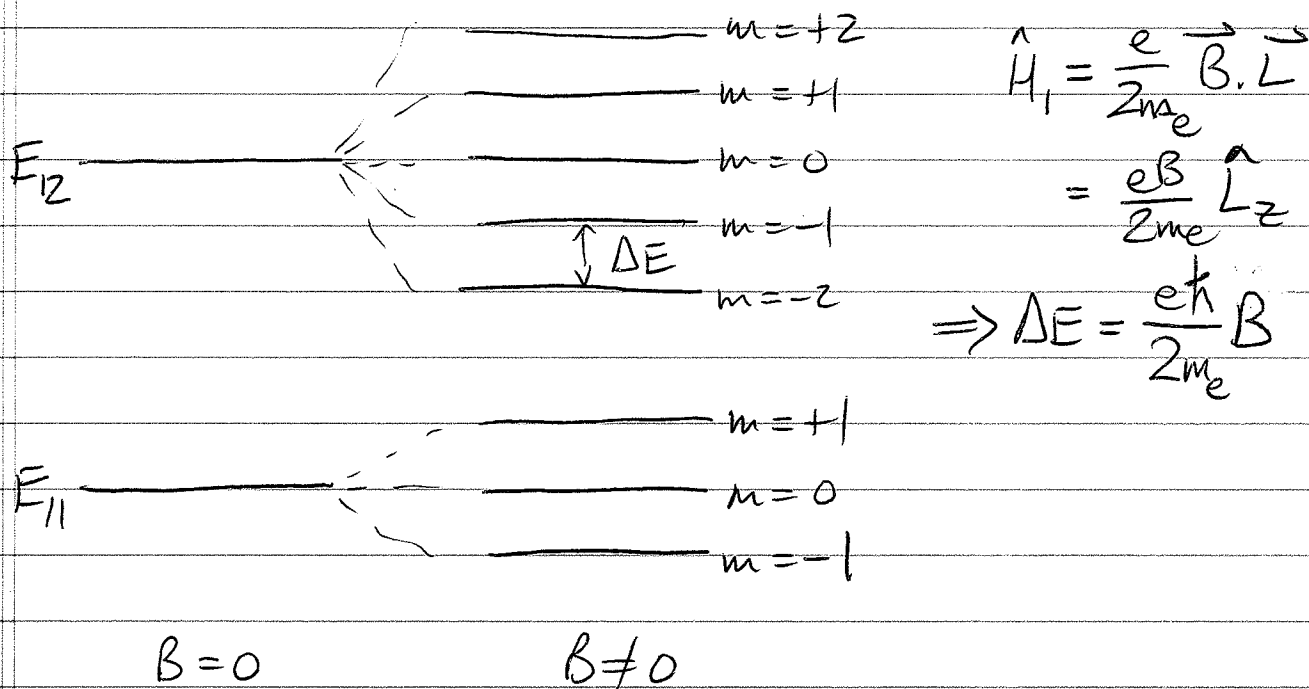
$$\Rightarrow A = \frac{1.165}{a\sqrt{a}} \text{ from part (b)}$$

$$\therefore \text{Probability of measuring } E_{10}, |\beta|^2 = \frac{4\pi A^2}{C_{10}^2} = \underline{\underline{0.866}}$$

$$\therefore \text{Probability of measuring } L_z = +\hbar, |\alpha|^2 = \frac{2\pi A^2}{3C_{11}^2} = \underline{\underline{0.067}}$$

③ d) contd.

Both $l=0$ states do not split, but the $l=1$ and $l=2$ states do. Since the energy gap between the ground state E_{10} and E_{11} is roughly equal to that between E_{11} and E_{12} , we expect the $l=1$ and $l=2$ states to be involved;



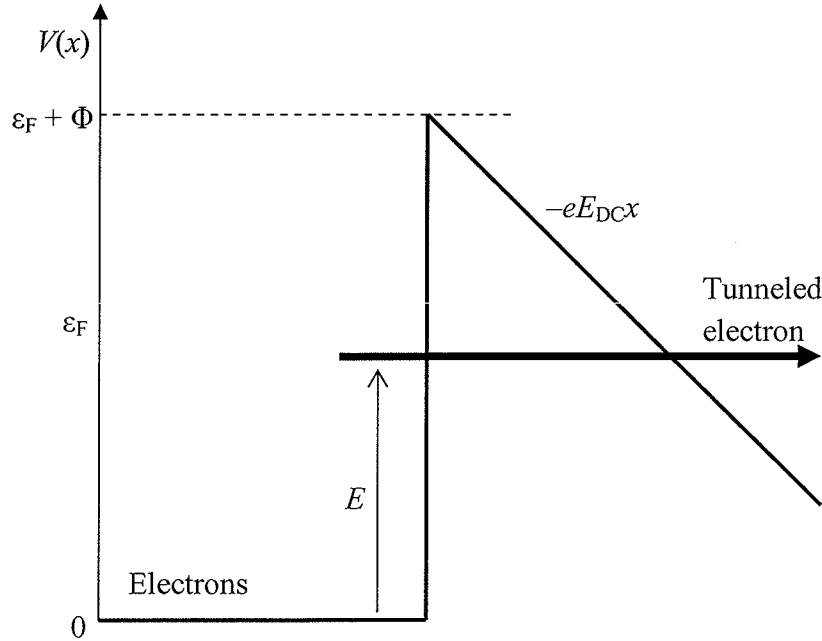
So, states overlap when

$$E_{21} - E_{11} = 3\Delta E = \frac{3ehB}{2m_e} = 6.51 \frac{\hbar^2}{mea^2}$$

$$\Rightarrow B = \underline{\underline{4.34 \frac{\hbar}{ea^2}}}$$

Question 4

Field emission, a quantum tunneling phenomenon, occurs when a DC electric field E_{DC} applied perpendicular to a material surface becomes sufficiently strong to allow electrons in the material to tunnel out into the vacuum.



In materials like metals, the zero-point energy E of the electrons in the metal is usually defined in terms of the Fermi energy ε_F (the energy of the last filled electronic state at zero Kelvin), which is located at an energy Φ (the material work function) below the vacuum energy.

- a) Using the approximation for the one-dimensional tunneling probability through a 'thick' barrier,

$$|T|^2 \approx \exp \left[-2 \int_a^b dx \sqrt{\frac{2m}{\hbar^2} (V(x) - E)} \right],$$

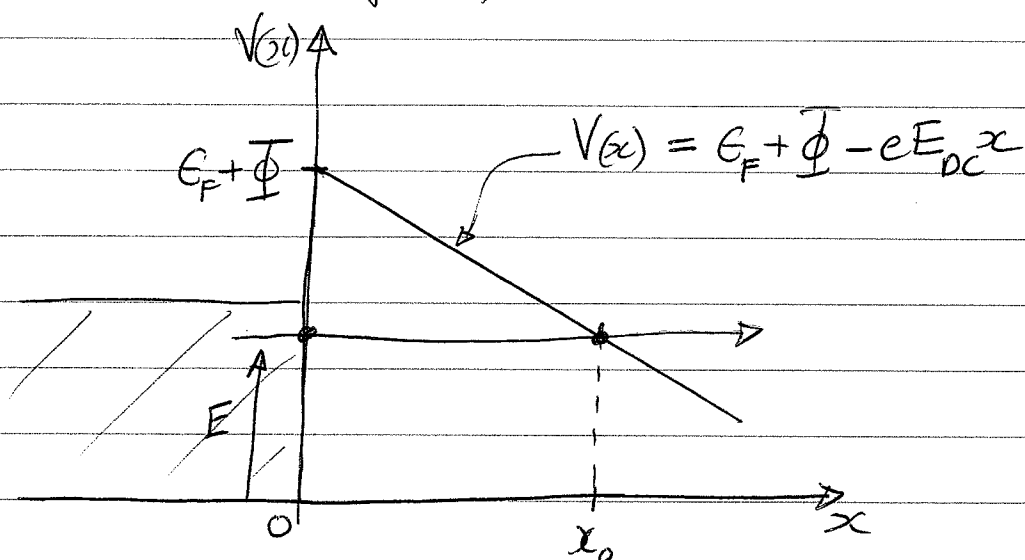
of thickness $b - a$, determine the transmission probability for an electron at an arbitrary energy E ; that is, at an energy $E' = \varepsilon_F + \Phi - E$ below the top of the barrier.

- b) In a real material at non-zero temperature T , the energy distribution of electrons is given by the Fermi function

$$f(E) = \frac{1}{1 + \exp \left(-\frac{\varepsilon_F - E}{k_B T} \right)},$$

where k_B is Boltzmann's constant. For $\exp[(\Phi - E')/k_B T] \gg 1$, determine the energy below the top of the barrier for which electron field emission is most probable.

④ a) To evaluate $|T|^2 \approx \exp\left[-2 \int_a^b dx \sqrt{\frac{2m}{\hbar^2} (V(x) - E)}\right]$, we need to determine $V(x)$ in the $a \rightarrow b$ interval and find a and b (the integral limits). Using $x=0$ at the metal surface, we have



$$V(x_0) = E = E_F + \Phi - eE_{dc}x_0 \Rightarrow x_0 = \frac{E_F + \Phi - E}{eE_{dc}}$$

$$\therefore |T|^2 \approx \exp\left[-2 \int_0^{x_0} dx \sqrt{\frac{2m}{\hbar^2} (E_F + \Phi - eE_{dc}x - E)}\right]$$

$$= \exp\left[-2 \sqrt{\frac{2meE_{dc}}{\hbar^2}} \int_0^{x_0} dx \sqrt{x_0 - x}\right]$$

$$\frac{2}{3} \left(\frac{E_F + \Phi - E}{eE_{dc}} \right)^{3/2}$$

(4) a) contd.

Hence,

$$|T|^2 \propto \exp \left[- \frac{4\sqrt{2m}}{3e\hbar E_{pc}} (\epsilon_F + \Phi - E)^{3/2} \right]$$

b) Including the Fermi distribution, we expect the transmission probability at energy E to be

$$|T(E)|^2 \propto \frac{\exp \left[-\alpha (\epsilon_F + \Phi - E)^{3/2} \right]}{1 + \exp \left[\beta (E - \epsilon_F) \right]}$$

$$\text{with } \alpha = \frac{4\sqrt{2m}}{3e\hbar E_{pc}} \text{ and } \beta = \frac{1}{k_B T}.$$

So, the most probable emission is at

$$\frac{\partial}{\partial E} |T(E)|^2 = 0$$

$$\Rightarrow -\frac{3\alpha}{2} (\epsilon_F + \Phi - E)^{1/2} (1 + e^{\beta(E - \epsilon_F)}) - \beta e^{\beta(E - \epsilon_F)} = 0$$

But, since $E' = \epsilon_F + \Phi - E \Rightarrow E - \epsilon_F = \Phi - E'$
and $\exp[(\Phi - E')/k_B T] \gg 1$, we get

(4) b) contd.

$$-\frac{3\alpha}{2} \sqrt{E'_{\max}} = \beta$$

$$\Rightarrow E'_{\max} = \frac{4\beta^2}{9\alpha^2} = \frac{e^2 \hbar^2 E_{DC}^2}{8m k_B T^2}$$

is the energy below the top of the barrier for most probable field emission.

[NOTE :

$$|T(E'_{\max})|^2 \approx \exp\left[-\frac{e^2 \hbar^2 E_{DC}^2}{12m k_B T^3}\right]$$

as a result of (b).]

Question 5

Consider the following *spatial-coordinate-independent* two-particle Hamiltonian for a spin system

$$\hat{H} = A + BS_1 \cdot S_2 ,$$

where A and B are constants. Calculate the eigenstates and eigenvalues of \hat{H} for two *identical* spin $\frac{1}{2}$ particles

- using the uncoupled basis, matrices, and spinors, and
- using the coupled basis.

Consider the following *spatial-coordinate-independent* two-particle Hamiltonian for a spin system consisting of two *identical* spin $\frac{1}{2}$ particles:

$$\hat{H} = A + BS_1 \cdot S_2$$

where A and B are constants.

- Find the 4×4 matrix representation of \hat{H} in the uncoupled basis set

$$|\uparrow\uparrow\rangle = \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix}, |\uparrow\downarrow\rangle = \begin{pmatrix} 0 \\ 1 \\ 0 \\ 0 \end{pmatrix}, |\downarrow\uparrow\rangle = \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix}, \text{ and } |\downarrow\downarrow\rangle = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 1 \end{pmatrix},$$

where $|\uparrow\rangle$ represents a spin-up particle with $S_z = \frac{1}{2}\hbar$ and a spin-down particle with $S_z = -\frac{1}{2}\hbar$ is $|\downarrow\rangle$.

- Calculate the eigenvalues of \hat{H} .
- What are the eigenstates of the diagonalized Hamiltonian?
- Verify your results by finding the eigenvalues of the four states in the coupled basis $|S, M\rangle$ representing the total spin S and its z -component M .

5) a) Since $\hat{S}_1 \hat{S}_2 = \hat{S}_1^x \hat{S}_2^x + \hat{S}_1^y \hat{S}_2^y + \hat{S}_1^z \hat{S}_2^z$ and $\hat{S}_\pm = \hat{S}_x \pm i \hat{S}_y$, we have

$$\hat{H} = A + B \hat{S}_1 \cdot \hat{S}_2 = A + \frac{1}{2} B \hat{S}_{1+} \hat{S}_{2-} + \frac{1}{2} B \hat{S}_{1-} \hat{S}_{2+} + B \hat{S}_{1z} \hat{S}_{2z}$$

$$\Rightarrow \begin{array}{c|cccc} \hat{H} & |\uparrow\uparrow\rangle & |\uparrow\downarrow\rangle & |\downarrow\uparrow\rangle & |\downarrow\downarrow\rangle \\ \hline |\uparrow\uparrow\rangle & A + \frac{1}{4} B \hbar^2 & 0 & 0 & 0 \\ |\uparrow\downarrow\rangle & 0 & A - \frac{1}{4} B \hbar^2 & \frac{1}{2} B \hbar^2 & 0 \\ |\downarrow\uparrow\rangle & 0 & \frac{1}{2} B \hbar^2 & A - \frac{1}{4} B \hbar^2 & 0 \\ |\downarrow\downarrow\rangle & 0 & 0 & 0 & A + \frac{1}{4} B \hbar^2 \end{array}$$

... since $\hat{S}_z |\uparrow\rangle = \frac{\hbar}{2} |\uparrow\rangle$, $\hat{S}_z |\downarrow\rangle = -\frac{\hbar}{2} |\downarrow\rangle$,
 $\hat{S}_+ |\uparrow\rangle = \hat{S}_- |\downarrow\rangle = 0$, $\hat{S}_+ |\downarrow\rangle = \hbar |\uparrow\rangle$,
 and $\hat{S}_- |\uparrow\rangle = \hbar |\downarrow\rangle$.

b) Clearly $|\uparrow\uparrow\rangle$ and $|\downarrow\downarrow\rangle$ are already eigenstates of \hat{H} with eigenvalues $A + \frac{1}{4} B \hbar^2$. The central 2×2 block can be diagonalized to find the other two eigenvalues;

$$\begin{vmatrix} A - \frac{1}{4} B \hbar^2 - \lambda & \frac{1}{2} B \hbar^2 \\ \frac{1}{2} B \hbar^2 & A - \frac{1}{4} B \hbar^2 - \lambda \end{vmatrix} = 0$$

$$\Rightarrow (A - \frac{1}{4} B \hbar^2 - \lambda)^2 - \frac{1}{4} B^2 \hbar^4 = 0$$

$$\Rightarrow A - \frac{1}{4} B \hbar^2 - \lambda = \pm \frac{1}{2} B \hbar^2$$

(5) b) contd. \therefore Eigenvalues are $\lambda = \underline{A + \frac{1}{4} B \hbar^2}$

$$\text{and } \lambda = \underline{A - \frac{3}{4} B \hbar^2}$$

c) For $\lambda = A + \frac{1}{2} B \hbar^2$, we have for the central 2×2 block

$$\frac{1}{2} B \hbar^2 \begin{pmatrix} -1 & 1 \\ 1 & -1 \end{pmatrix} \begin{pmatrix} a \\ b \end{pmatrix} = 0 \Rightarrow a = b$$

\therefore Eigenstate is $\underline{\frac{1}{\sqrt{2}} (| \uparrow \downarrow \rangle + | \downarrow \uparrow \rangle)}$

For $\lambda = A - \frac{3}{4} B \hbar^2$, we get

$$\frac{1}{2} B \hbar^2 \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \begin{pmatrix} a \\ b \end{pmatrix} = 0 \Rightarrow a = -b$$

\therefore Eigenstate is $\underline{\frac{1}{\sqrt{2}} (| \uparrow \downarrow \rangle - | \downarrow \uparrow \rangle)}$

And the other two eigenstates for $\lambda = A + \frac{1}{4} B \hbar^2$ are $| \uparrow \uparrow \rangle$ and $| \downarrow \downarrow \rangle$.

d) In the coupled basis $| S, M \rangle$

$$\hat{H} = A + B \hat{S}_1 \cdot \hat{S}_2 = A + \frac{1}{2} B (\hat{S}^2 - \hat{S}_1^2 - \hat{S}_2^2)$$

⑤ d) contd.

Thus, the energy eigenvalues are

$$E = A + \frac{1}{2}B \left(2\hbar^2 - \frac{3}{4}\hbar^2 - \frac{3}{4}\hbar^2 \right) = A + \frac{1}{4}B\hbar^2 \text{ for } S=1$$

$$E = A + \frac{1}{2}B \left(0 - \frac{3}{4}\hbar^2 - \frac{3}{4}\hbar^2 \right) = A - \frac{3}{4}B\hbar^2 \text{ for } S=0$$

Consequently, since the coupled basis are $|S=1, M\rangle$ and $|S=0, 0\rangle$, we have

$$\text{Eigenstates} \left\{ \begin{array}{l} |S=1, M=1\rangle \\ |S=1, M=0\rangle \\ |S=1, M=-1\rangle \\ |S=0, M=0\rangle \end{array} \right\} \left\{ \begin{array}{l} A + \frac{1}{4}B\hbar^2 \\ A - \frac{3}{4}B\hbar^2 \end{array} \right\} \text{Eigenvalues}$$

in agreement with prior parts of question.

Equation Sheet

$$Y_{00}(\theta, \phi) = \frac{1}{\sqrt{4\pi}} \quad Y_{10}(\theta, \phi) = \sqrt{\frac{3}{4\pi}} \cos \theta \quad Y_{1\pm 1}(\theta, \phi) = \mp \sqrt{\frac{3}{8\pi}} e^{\pm i\phi} \sin \theta$$

$$Y_{20}(\theta, \phi) = \sqrt{\frac{5}{16\pi}} (3 \cos^2 \theta - 1) \quad Y_{2\pm 1}(\theta, \phi) = \mp \sqrt{\frac{15}{8\pi}} e^{\pm i\phi} \sin \theta \cos \theta$$

$$Y_{2\pm 2}(\theta, \phi) = \sqrt{\frac{15}{32\pi}} e^{\pm 2i\phi} \sin^2 \theta$$

$$E_n = - \left[\frac{\mu}{2\hbar^2} \left(\frac{Ze^2}{4\pi\epsilon_0} \right)^2 \right] \frac{1}{n^2} = -13.6 \frac{Z}{n^2} \text{ eV} \quad a_0 = \frac{4\pi\epsilon_0 \hbar^2}{\mu e^2} \quad \frac{1}{\mu} = \frac{1}{m_e} + \frac{1}{M_{\text{nucleus}}}$$

$$R_{10}(r) = 2 \left(\frac{Z}{a_0} \right)^{\frac{3}{2}} \exp \left[-\frac{Zr}{a_0} \right] \quad R_{20}(r) = 2 \left(\frac{Z}{2a_0} \right)^{\frac{3}{2}} \left(1 - \frac{Zr}{2a_0} \right) \exp \left[-\frac{Zr}{2a_0} \right]$$

$$R_{21}(r) = \frac{1}{\sqrt{3}} \left(\frac{Z}{2a_0} \right)^{\frac{3}{2}} \frac{Zr}{a_0} \exp \left[-\frac{Zr}{2a_0} \right]$$

$$\int_0^\infty dx x^m \exp(-ax^2) = \frac{\Gamma\left(\frac{m+1}{2}\right)}{2a^{(m+1)/2}} \quad ; \quad \Gamma(n+1) = n\Gamma(n), \Gamma(n+1) = n!, \Gamma\left(\frac{1}{2}\right) = \sqrt{\pi}$$

$$\text{so that } \int_0^\infty dx x^{2n} \exp(-\lambda^2 x^2) = \frac{1.3.5 \dots (2n+1) \sqrt{\pi}}{2^n \lambda^{2n+1}}$$

$$\int_0^\infty dx x^n e^{-\lambda x} = \frac{n!}{\lambda^{n+1}}$$

$$\int dx \sqrt{A+Bx} = \frac{2}{3B} (A+Bx)^{\frac{3}{2}} \quad \int dx x \sqrt{A+Bx} = -\frac{2(2A-3Bx)(A+Bx)^{\frac{3}{2}}}{15B^2}$$