

# E/gamma energy calibration for Higgs mass measurement

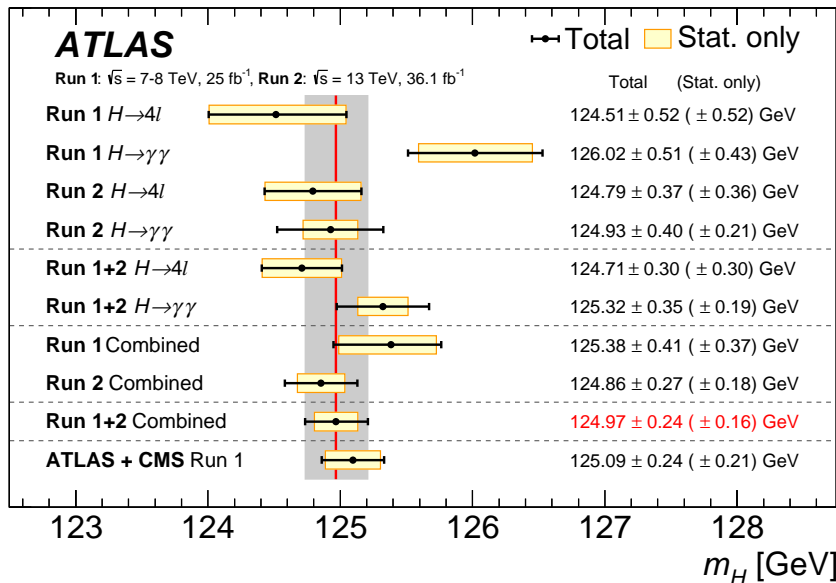
ATLAS week - Berlin 2019

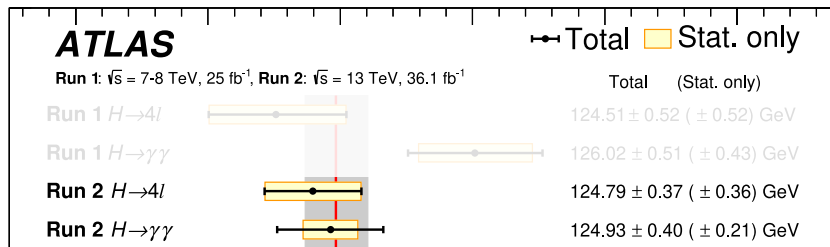
Ruggero Turra

on behalf of the e/gamma calibration and the Higgs groups

INFN Milano

8 Oct 2019



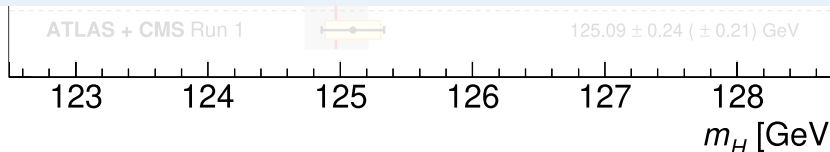


$\gamma\gamma$  starts to be dominated by systematic uncertainty (0.34 GeV)

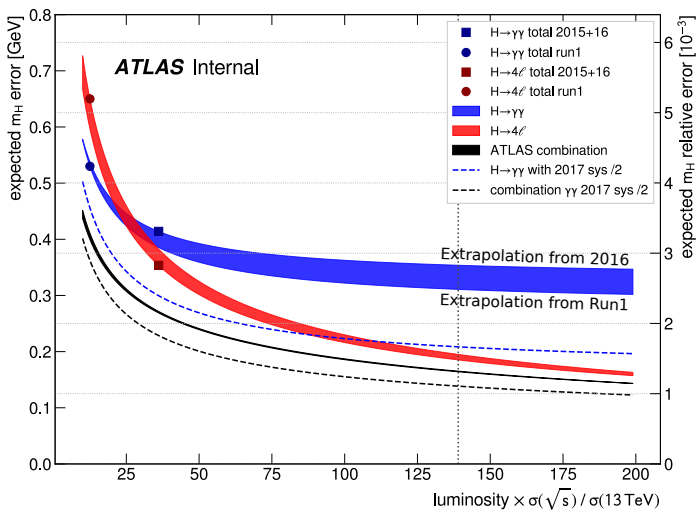
- completely from photon energy scale

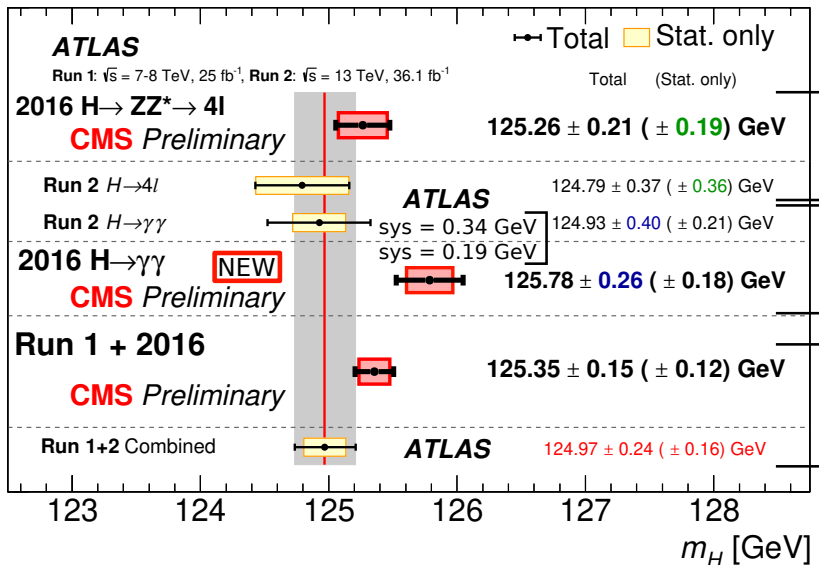
Huge effort needed on systematic reduction if we want  $\gamma\gamma$  to be competitive at 140 fb<sup>-1</sup>

Full Run2 measurement will be dominated by  $H \rightarrow 4\ell$ ...



... unless we reduce energy scale systematic by a factor  $\simeq 2$



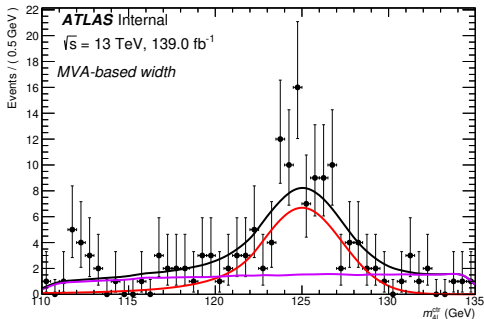
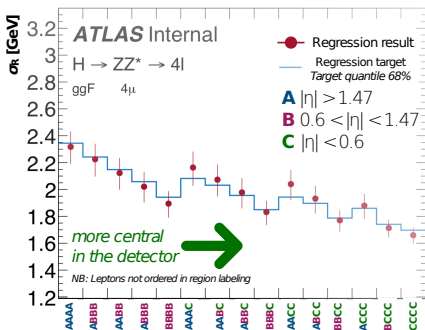


# $H \rightarrow 4\ell$ ongoing analysis Full Run2

- Dominated by statistical error
- Category by lepton flavour
- Improve resolution modeling and purity:  $Z$ -mass constraint, FSR recovery, bkg BDT, per-event resolution
- 2D fit  $m_{4\ell} \times BDT_{ZZ}$  (signal vs  $ZZ$ ):

$$P[m_{4\ell}, BDT_{ZZ} | m_H, \sigma_{4\ell}] = P[BDT_{ZZ} | m_{4\ell}] P[m_{4\ell} | m_H, \sigma_{4\ell}]$$

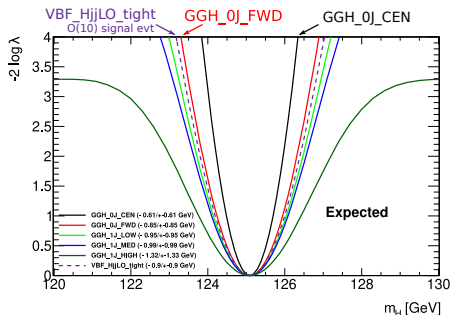
- NEW: resolution value from a NN regression:  $\sigma_{4\ell} = NN[\text{kinematic, flavours}]$



Same selection as other  $H(125) \rightarrow \gamma\gamma$  analyses:  $E_T^\gamma/m_{\gamma\gamma} > 0.25/0.35$ , tight, isolated.

31 categories as in the STXS analysis: (ggF\_0J\_CEN, ggF\_0J\_FWD, ..., ttH):

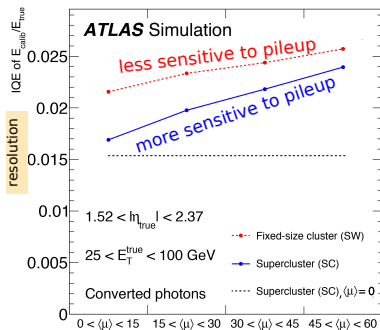
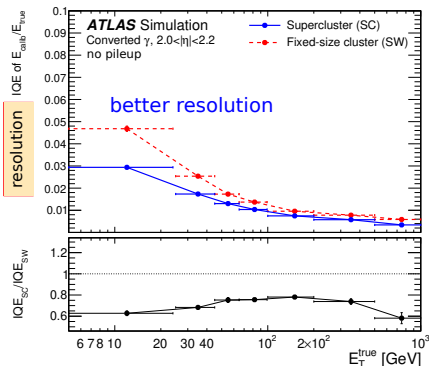
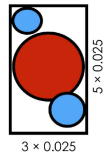
- Separate events with **different s/b** (VBF\_HjjLO\_tight has s/b=1)
- Quite separate events with **different resolution** (CEN/FWD, high-pT (VBF) / low-pT (ggF))



| method                              | ADD  |     | SUB  |     |
|-------------------------------------|------|-----|------|-----|
|                                     | down | up  | down | up  |
| sys [MeV]                           |      |     |      |     |
| Background model                    | 70   | 67  | 53   | 55  |
| Conversion reconstruction           | 49   | 50  | 52   | 47  |
| ID material                         | 107  | 109 | 108  | 109 |
| <u>Lar cell non-linearity</u>       | 193  | 190 | 182  | 178 |
| <u>Lateral shower shape</u>         | 110  | 109 | 111  | 110 |
| <u>Layer calibration</u>            | 178  | 177 | 170  | 169 |
| Luminosity                          | 1    | 0   | 18   | 0   |
| <u>Other material (not ID)</u>      | 124  | 125 | 120  | 121 |
| Resolution                          | 20   | 0   | 20   | 22  |
| Signal model                        | 17   | 18  | 15   | 16  |
| Primary vertex effect on mass scale | 37   | 38  | 37   | 38  |
| $Z \rightarrow ee$ calibration      | 76   | 78  | 76   | 76  |

Change in reconstruction: use **super-cluster** instead of **sliding window**

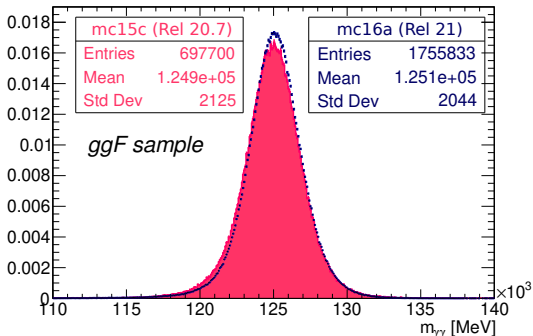
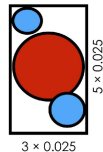
- Recover energy from satellite clusters (e.g. soft conversions)
- Better resolution, in particular for converted photons in the endcap
- More sensitive to pileup





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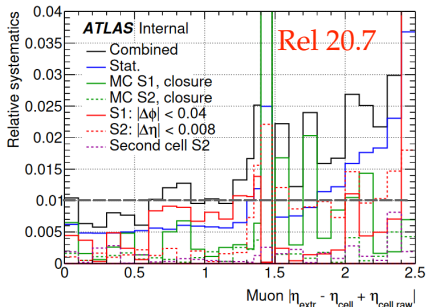
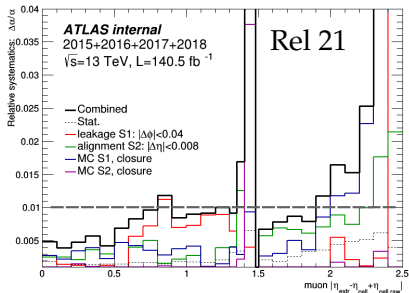
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4% improvement in mc16a (2015+16 data) ggF resolution. Smaller at higher pileup

## Layer calibration (ongoing)

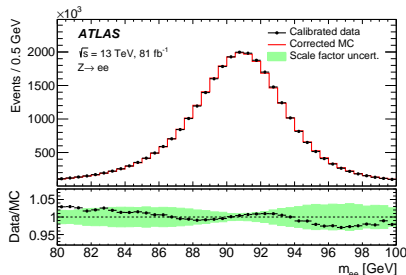
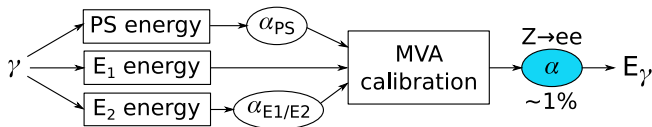
- $E_1/E_2$  scale measured with muons (insensitive to material in front)
- Full Run2 data used: will enter in fall recommendation
- Overall uncertainty decreased (vs Rel 20.7 2015+16 data)



- PS scale study just started

Still quite large extrapolation uncertainty on muon  $\rightarrow$  photons

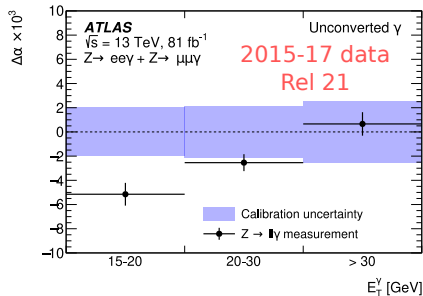
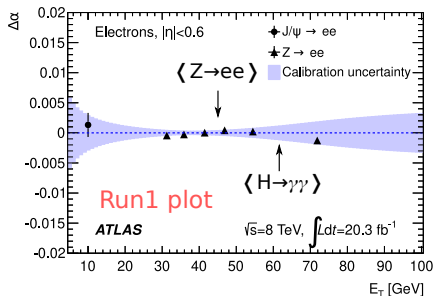
The alternative is to study  $E_1/E_2$  directly with electrons



- Energy systematic dominated by extrapolation from  $\simeq 45 \text{ GeV}$  electrons  $\rightarrow \simeq 60 \text{ GeV}$  photons
- Systematics split in  $\sim 70$  components with non-trivial correlation in  $\eta$
- Most relevant for  $m_H^{\gamma\gamma}$ : high/medium gain,  $E_1/E_2$  calibration, non-ID material

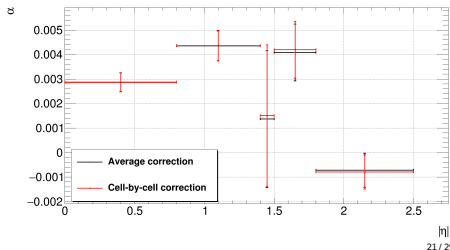
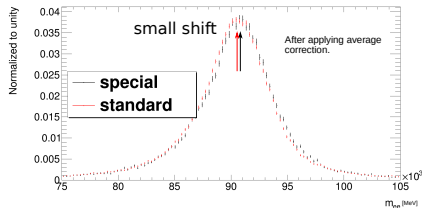
# Scales as a function of $p_T$ and from $Z \rightarrow \ell\ell\gamma$

- Presently used as cross checks: compute residual scale factors on top of the ones from  $Z \rightarrow ee$
- Some tension between  $\gamma/e$  scales



Can be used to reduce extrapolation uncertainties 45 GeV electrons  $\rightarrow$  60 GeV photons

- Calibrate medium / high-gain using **special runs**: lower DAC threshold to record most of Z decay events in MG (instead of HG)
- Derive scale factor between MG/HG from  $Z \rightarrow ee$
- Presently: assume no correction, use the deviation as systematic (100%)
- May change in future: correct and use error from the measurement  $\rightarrow$  strong reduction in systematic



## Projection to $140 \text{ fb}^{-1}$ $H \rightarrow \gamma\gamma$ (expected)

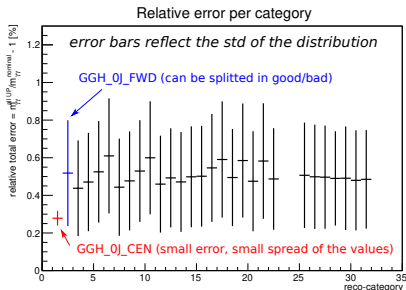
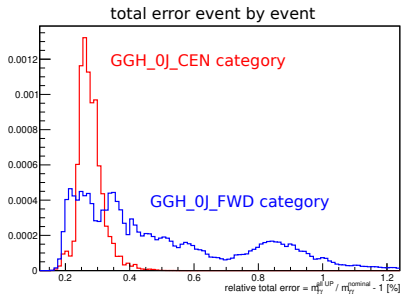
Use workspace from coupling 80/fb, try scenarios provided by egamma with different values of systematic

|  | total [GeV] | stat [GeV] | sys [GeV] |
|--|-------------|------------|-----------|
| Latest result with $36 \text{ fb}^{-1}$ Rel 20.7               | 0.40        | 0.25       | 0.31      |
| ... extrapolated at $140 \text{ fb}^{-1}$                      | 0.35        | 0.13       | 0.31      |
| Refit latest analysis Rel 21 with $140 \text{ fb}^{-1}$        | 0.26        | 0.10       | 0.24      |
| ... with reduced systematics <sup>1</sup>                      | 0.20        | 0.10       | 0.18      |
| Latest $4\ell$ extrapolated at $140 \text{ fb}^{-1}$           | 0.18        | 0.18       | $\sim 0$  |
| ATLAS Combination $\gamma\gamma + 4\ell$ $140 \text{ fb}^{-1}$ | 0.13        |            |           |

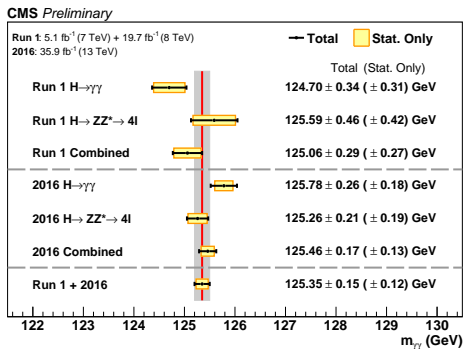
With reduced systematic scenario  $\gamma\gamma$  is competitive with  $4\ell$  (CMS extrapolated  $\sim 0.09 \text{ GeV}$ )

<sup>1</sup>gain and leakage (now 100% uncertainty) by 3, layer by 2

- Past strategy: separate events with good/bad resolution and high/low purity
- Need to focus also on the value of the systematics (e.g. separate by conv/unconv)
- Ideally one wants to use for each event its value of the systematic and not the value computed as average of the category  $\rightarrow$  per-event systematic error
- Feedback from  $4\ell$  (which did similarly for the resolution): computationally too expensive
- Alternatively: split present categories in bad/good systematic



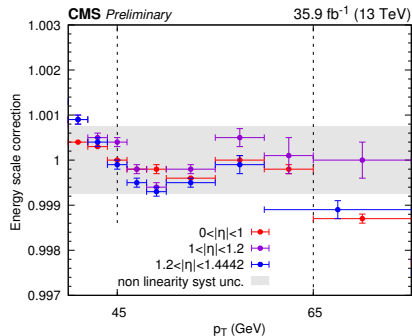
- 3 VBF categories + 4 untagged (BDT) (STXS in ATLAS)
- MVA regression with shower shapes ( $R_9$ ), pileup as inputs (no conv/unconv separated regression); regression of the expected resolution (no shower shapes in ATLAS, large mismodeling)
- Best resolution  $\sigma_{\text{eff}} = 1.35 \text{ GeV}$  (ATLAS  $1.59 \text{ GeV}$  in GGH\_0J\_CEN)
- BDT photon identification (ATLAS: cut-based). Diphoton event BDT



Energy scale assumed to be uncorr. between measurements  $\Rightarrow$  inflate uncertainty by 5%



- Scales from  $Z \rightarrow ee$  as a function of  $|\eta| \otimes p_T \otimes R_9$ .
- $R_9$  different from e/conv/unconv
- Systematic error from the non-linearity: non-closure of the  $p_T$ -dependent scale factors (0.075-0.15%)



|  |                      | ATLAS |
|--|----------------------|-------|
| Source   | Contribution [ GeV ] |       |
| Electron energy scale and resolution corrections     | 0.10                 | 0.08  |
| Residual $p_T$ dependence of the photon energy scale | 0.11                 |       |
| Modelling of the material budget                     | 0.03                 | 0.16  |
| Non-uniformity of the light collection               | 0.11                 |       |
| Statistical uncertainty                              | 0.18                 | 0.21  |
| Total uncertainty                                    | 0.26                 | 0.40  |
| total systematic                                     | 0.18                 | 0.34  |

- $H \rightarrow 4\ell$  almost systematic free
- $H \rightarrow \gamma\gamma$  will be dominated by systematics, not competitive  $140 \text{ fb}^{-1}$
- unless energy scale systematics reduced by a factor 2
- this is challenging, but within reach assuming we have the people to work on it (for the gain and for the leakage, we currently quote 100% uncertainty on the source of uncertainty)
- From the analysis point of view need to separate events with high/low-systematic value

## Person power in e/gamma calibration

Timeline of e/gamma limited by person-power. Some tasks just started (PS scale,  $Z \rightarrow \ell\ell\gamma$ , lateral leakage, several are uncovered (see [here](#)):

- non-ID material
- $Z \rightarrow ee$  scales

Also  $W$  mass is affected.

## Section 1

Backup

- Presently we compute the stat-only error fixing the NP-values to the best-fit and look to the width of  $-2 \log \frac{L(\mu, \hat{\hat{\theta}}_\mu)}{L(\hat{\mu}, \hat{\theta})} \sim \chi^2$
- This correspond to assume that there are no systematics error
- The weight the various channels/categories receive is proportional only to the stat-error (and not the total error)
- Discussion about if this is what we want
- Another approach is to propagate the systematic error taking into account the weights from the total error (BLUE, toys, ...)

In this slides: check how much these two approaches differ withing the  $\gamma\gamma$  analysis

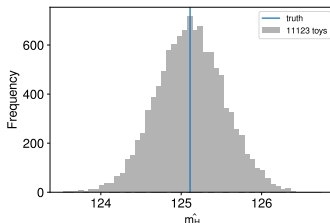
- $\simeq 6000$  toys: much more than the needed ones, precision negligible wrt rounding (e.g. std of  $\hat{m}_H = 0.424 \pm 0.006$ )
- Systematic error always computed as quadrature difference:  $\sqrt{\text{total}^2 - \text{stat-only}^2}$

Showing some results also for couplings: remember differences of workspaces

Conf-note result:  $125.11 \pm 0.42 \text{ GeV} = 125.11 \pm 0.21(\text{stat}) \pm 0.36(\text{sys}) \text{ GeV}$

Toys:  $125.11 \pm 0.42 \text{ GeV} = 125.11 \pm 0.26(\text{stat}) \pm 0.33(\text{sys}) \text{ GeV}$

- No bias to the central value
- Full error perfectly reproduced (it must be)
- Smaller sys-error with toy
  - make sense: if you weight with the full-error, categories with larger sys are weighted less



## Run1 categorization (10 categories)

- 1 Both photon candidates are **unconverted**, and have  $|\eta_{S2}| < 0.75$ ; the diphoton system has  $p_{Tt} < 70$  GeV.
- 2 Both photon candidates are unconverted, and have  $|\eta_{S2}| < 0.75$ ; the diphoton system has  $p_{Tt} > 70$  GeV.
- 3 Both photon candidates are unconverted and at least one candidate has  $|\eta_{S2}| > 0.75$ ; the diphoton system has  $p_{Tt} < 70$  GeV.
- 4 Both photon candidates are unconverted and at least one candidate has  $|\eta_{S2}| > 0.75$ ; the diphoton system has  $p_{Tt} > 70$  GeV.
- 5 Both photon candidates are unconverted and at least one candidate is in the range  $1.3 < |\eta_{S2}| < 1.37$  or  $1.52 < |\eta_{S2}| < 1.75$ .
- 6 At least one photon candidate is converted and both photon candidates have  $|\eta_{S2}| < 0.75$ ; the diphoton system has  $p_{Tt} < 70$  GeV.
- 7 At least one photon candidate is converted and both photon candidates have  $|\eta_{S2}| < 0.75$ ; the diphoton system has  $p_{Tt} > 70$  GeV.
- 8 At least one photon candidate is converted and both photon candidates have  $|\eta_{S2}| < 1.3$  or  $|\eta_{S2}| > 1.75$ , but at least one photon candidate has  $|\eta_{S2}| > 0.75$ . The diphoton system has  $p_{Tt} < 70$  GeV.
- 9 At least one photon candidate is converted and both photon candidates have  $|\eta_{S2}| < 1.3$  or  $|\eta_{S2}| > 1.75$ , but at least one photon candidate has  $|\eta_{S2}| > 0.75$ . The diphoton system has  $p_{Tt} > 70$  GeV.
- 10 At least one photon candidate is converted and at least one photon candidate is in the range  $1.3 < |\eta_{S2}| < 1.37$  or  $1.52 < |\eta_{S2}| < 1.75$ .