

Statistical Estimation for Physics

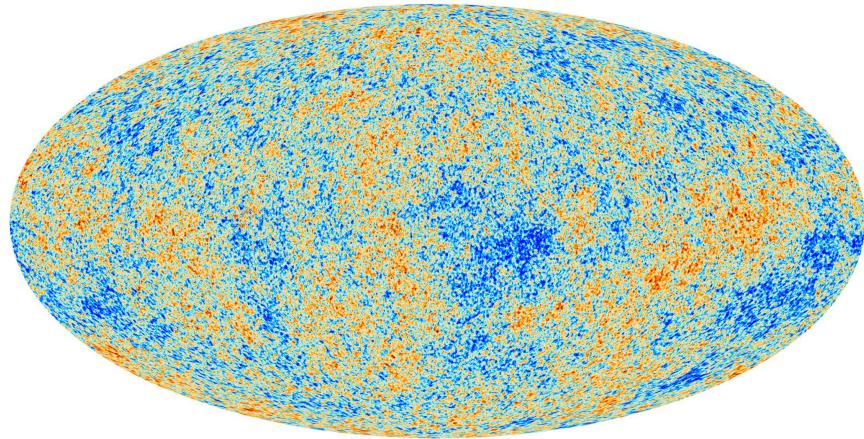
Bayesian Statistical Inference*

Boris Bolliet
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(DAMTP)

*[Link](#) to GitHub repository

The Cosmic Microwave Background (CMB) radiation

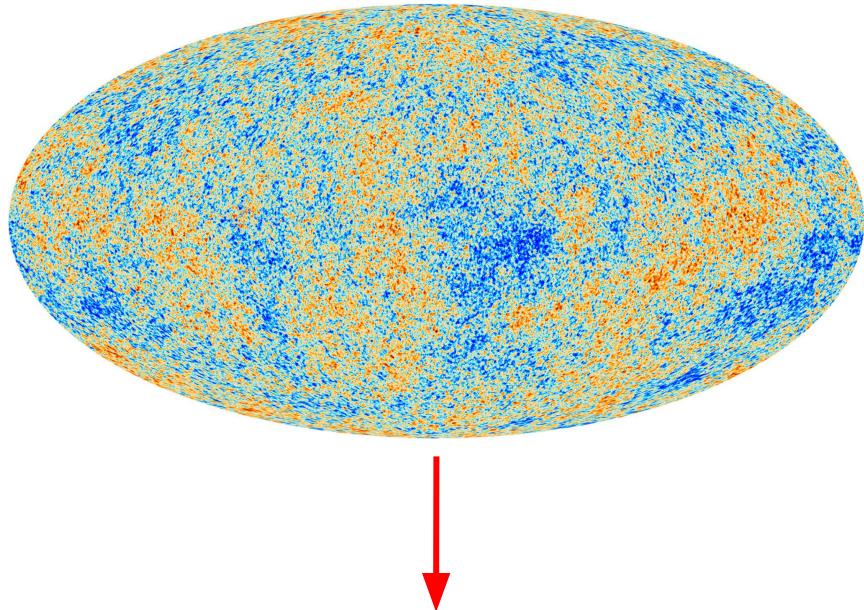
Full-sky map of the CMB temperature fluctuations



- Relic radiation from the Big Bang
- Average temperature $\bar{T} = 2.7\text{K}$
- Fluctuations $\frac{\Delta T}{\bar{T}} \approx 10^{-5}$

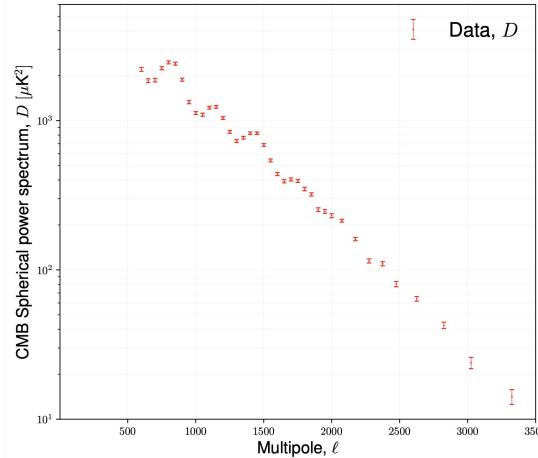
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“Data compression”

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large scales

small scales

A two-parameter model $D^{\text{th}} = f(\text{age}, M_\nu)$

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 - Parameter of interest

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→ $\theta = \{\text{age}, M_\nu\}$

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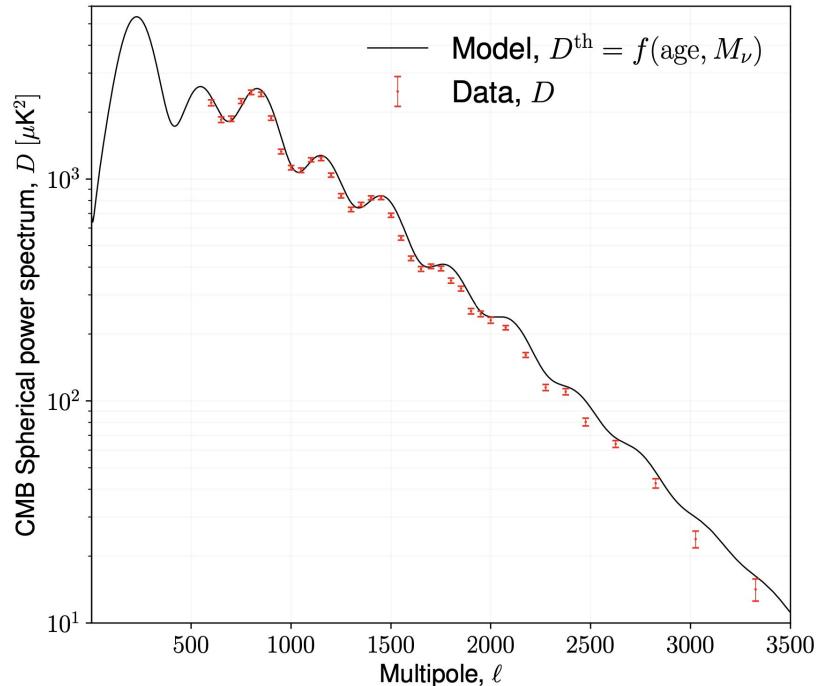
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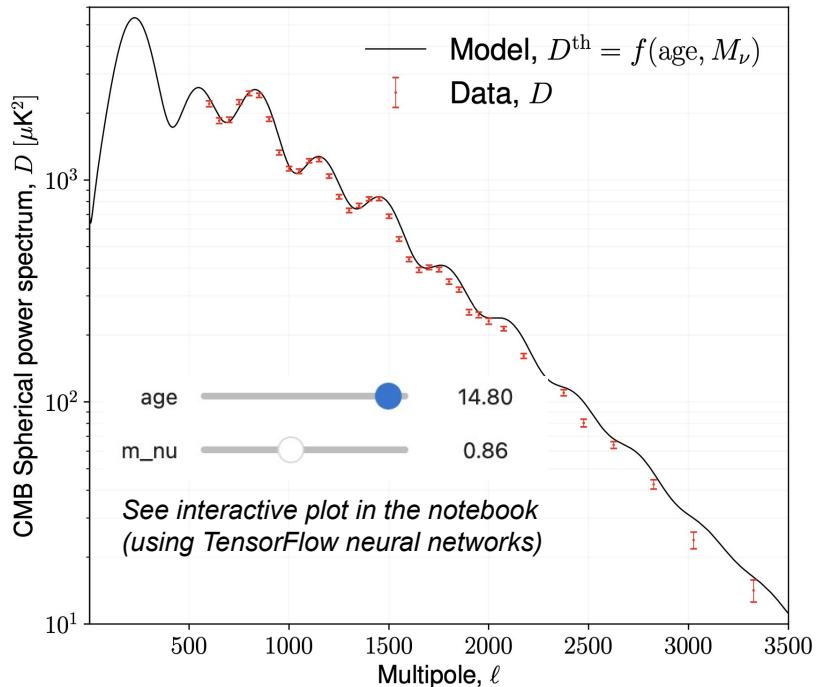
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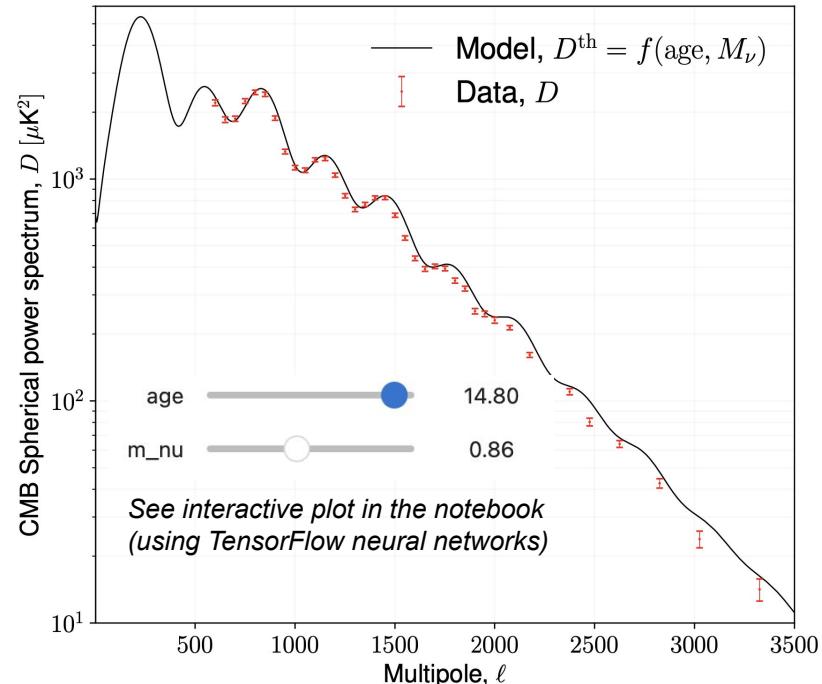
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How old is the universe?



Bayes' theorem for Bayesian inference

- Bayes' theorem: $P(A|B) = \frac{P(B|A)P(A)}{P(B)}$
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$D^{\text{th}} = f(\theta)$: model

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 - Evidence: normalizing constant $\longrightarrow P(\theta|D) \propto P(D|\theta)P(\theta)$
- D : data
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 θ : parameters

The likelihood function

$$P(\theta|D) \propto P(D|\theta)P(\theta)$$

- *Definition:* probability of data for a given θ , a function of θ , $\mathcal{L}(\theta) = P(D|\theta)$
- Summarises experimental data
- Main focus of classical/frequentist analysis

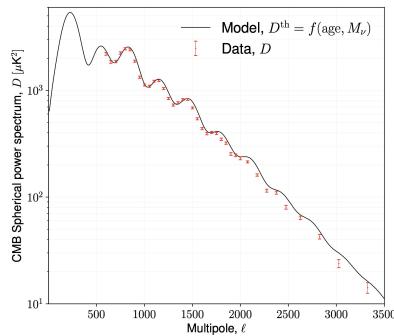
The CMB likelihood

- Gaussian approximation $P(D|\theta) = \frac{1}{\sqrt{2\pi|C|}} \exp\left\{-\frac{1}{2}(D^{\text{th}}(\theta) - D)^T C^{-1} (D^{\text{th}}(\theta) - D)\right\}$

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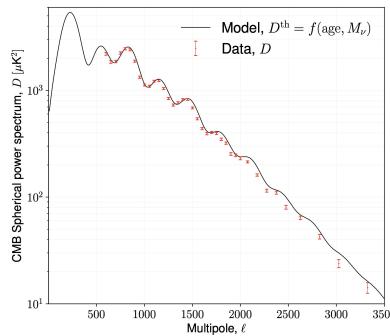
Model and data



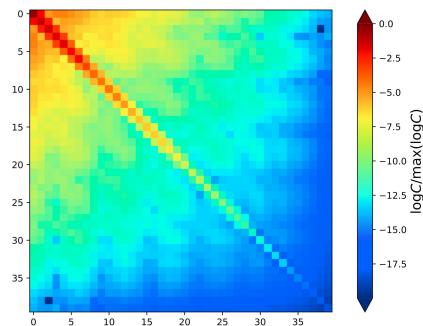
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Model and data



*Covariance matrix
(part of experimental data)*

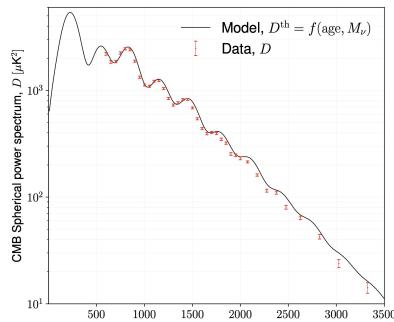


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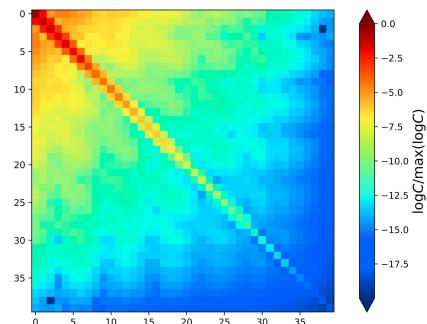
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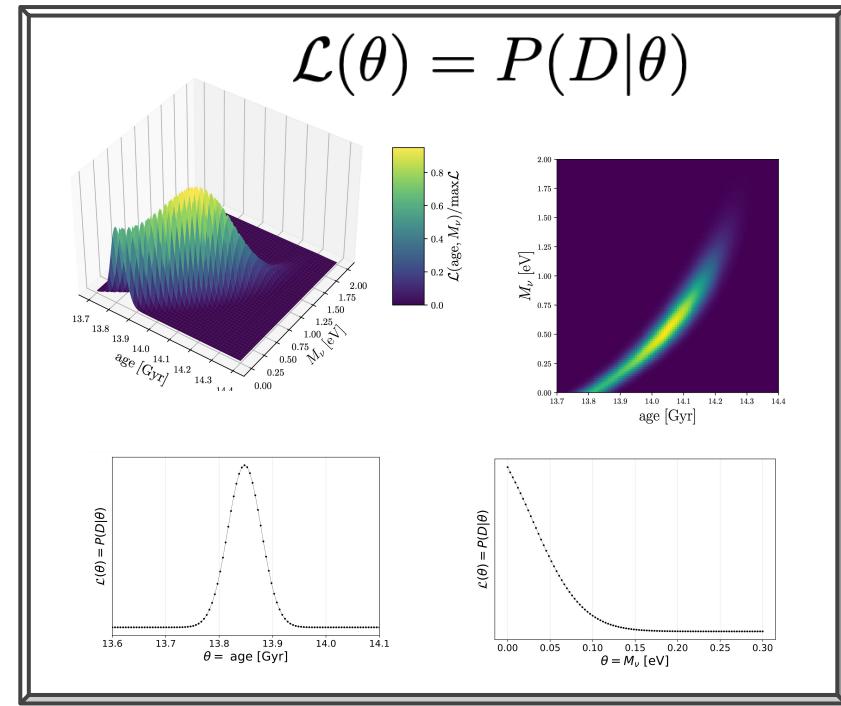
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$$\mathcal{L}(\theta) = P(D|\theta)$$



The prior

$$P(\theta|D) \propto P(D|\theta)P(\theta)$$

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- *Definition:* probability distribution for the range of possible parameter values
- Key advantage of Bayesian approach
- **Summarises prior knowledge** on the model
- Prior choice
 - Guiding principles
 - Educated intelligence

The posterior

$$P(\theta|D) \propto P(D|\theta)P(\theta)$$

- *Definition:* multivariate/joint probability distribution for the model parameters (conditioned on D and model, given prior)
- Goal of Bayesian inference: solution to the problem
- **Degree of belief**

The posterior

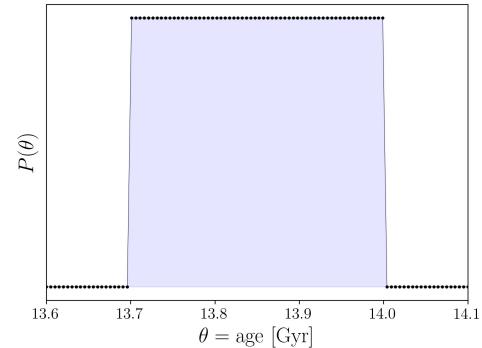
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- Does not exist in the classical/frequentist approach

Evaluating the posterior, a 1-parameter example

- Pick a **uniform prior**
- Idea: explore parameter space (sampling), compute prior and multiply by likelihood

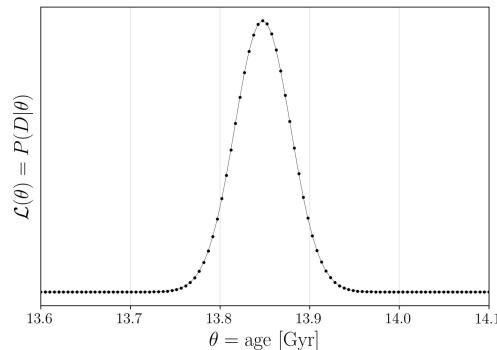
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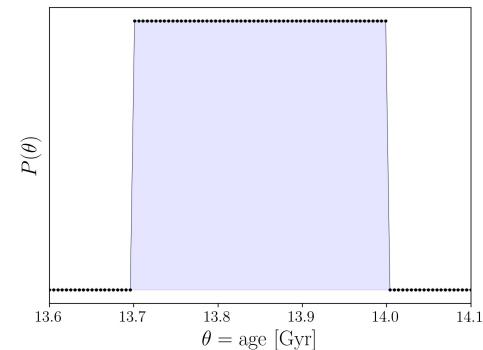
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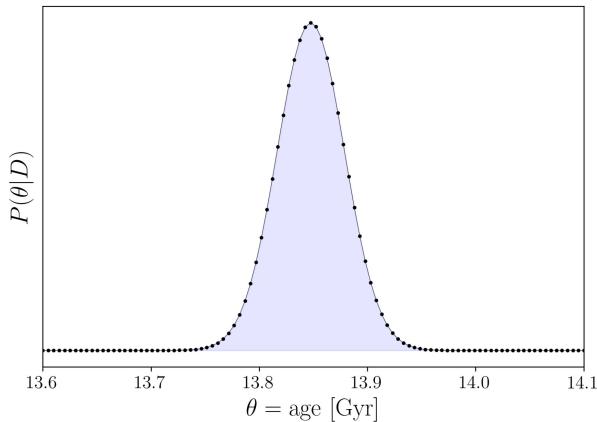
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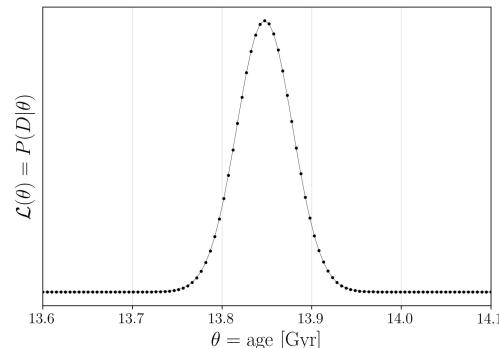
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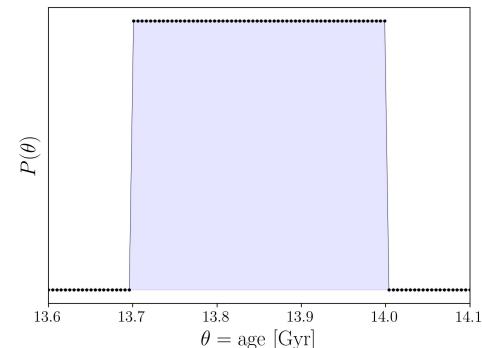
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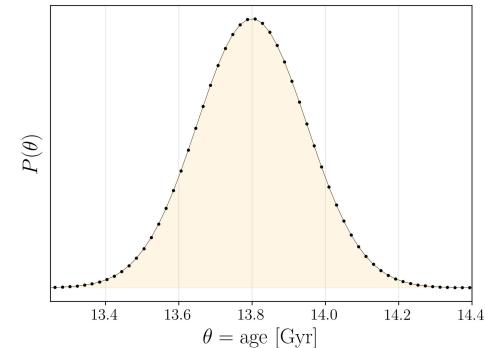
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Evaluating the posterior, a 1-parameter example

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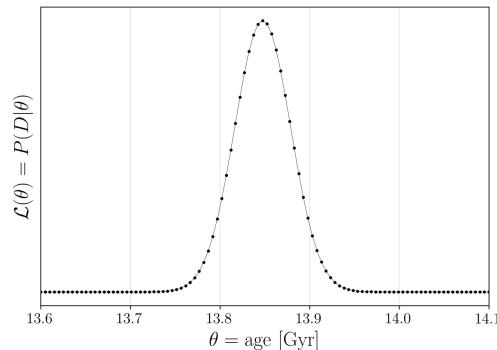
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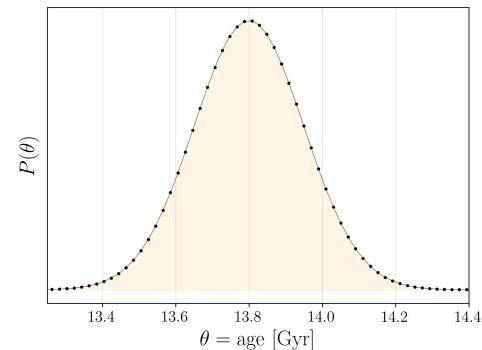
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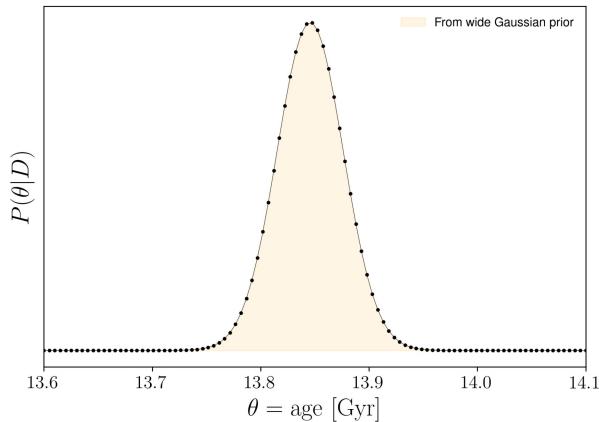
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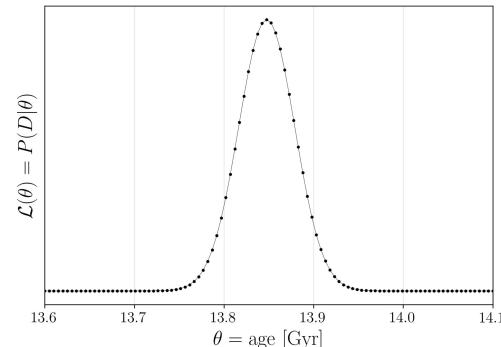
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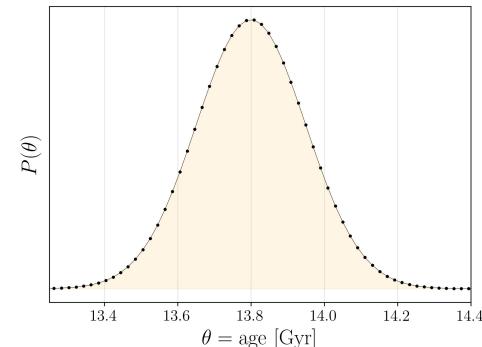
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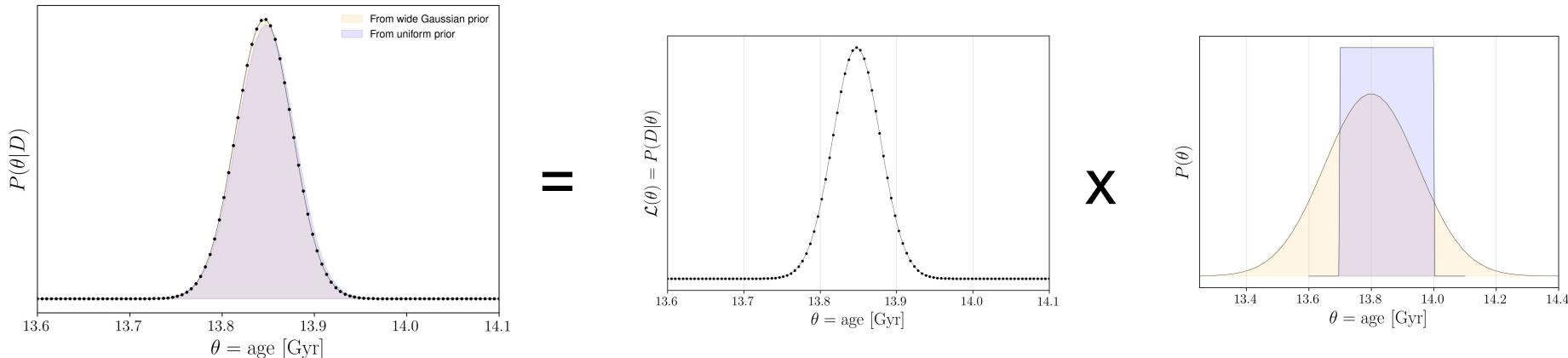
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Two very different priors lead to the same posterior!

Bayesian credible intervals

- Set a X% Confidence Level (CL), e.g., X=68 or 95
- Find narrowest interval $t_1 < \theta < t_2$ such that:

$$X = 100 \int_{t_1}^{t_2} P(\theta|D) d\theta$$

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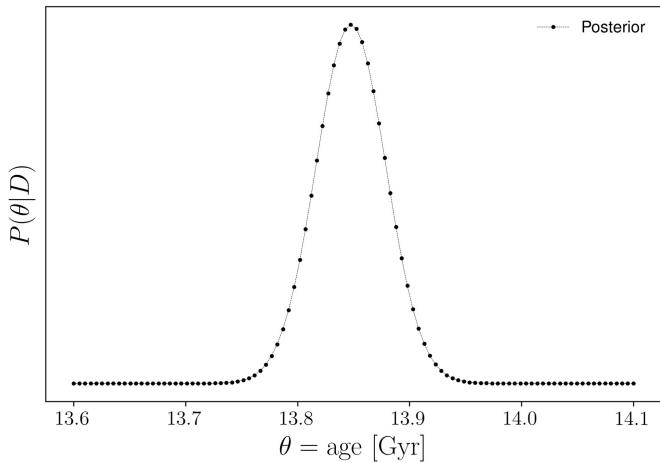
- Bayesian interpretation: **degree of belief** that the true value of theta lies within this interval (given D and model)
- Different from frequentist view!

Obtaining the 68% and 95%CL credible intervals

- General method: Kernel Density Estimation (KDE)

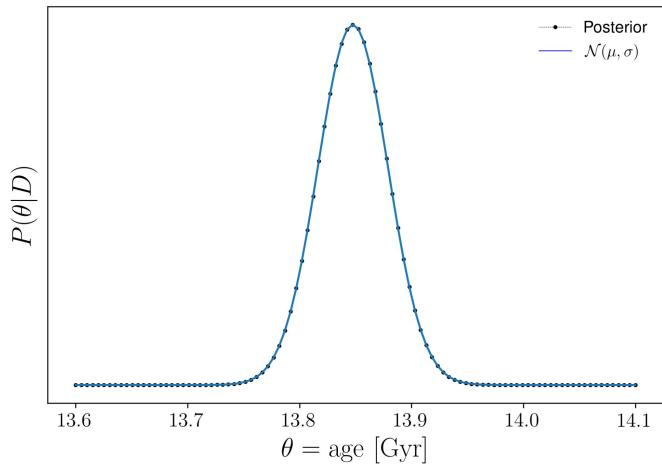
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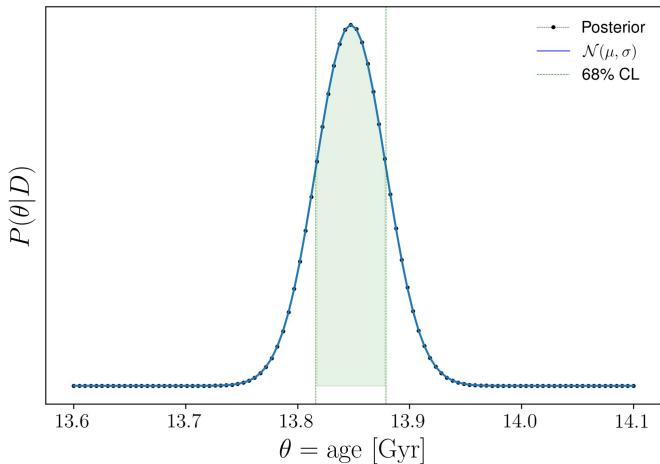
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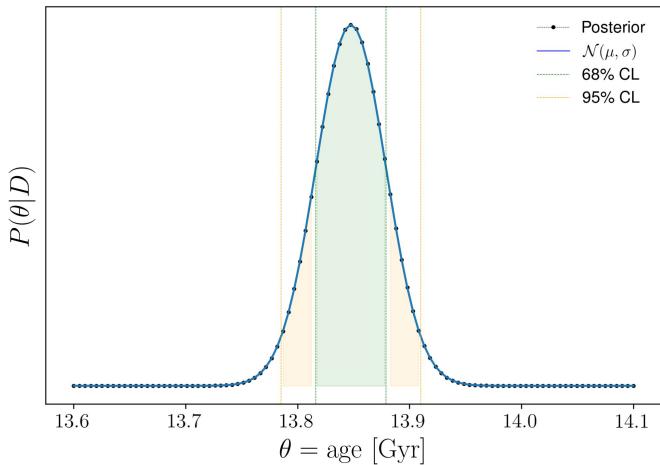
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age = 13.847 ± 0.031 Gyr (68%CL)

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age = 13.847 ± 0.031 Gyr (68%CL)

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General case

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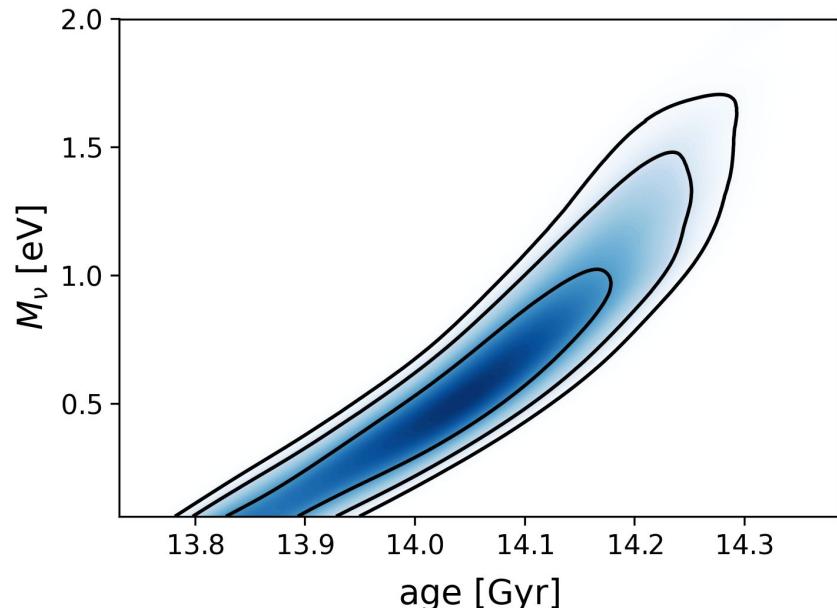
$$P(\text{age}, M_\nu | D) \propto P(D | \text{age}, M_\nu) P(\text{age}, M_\nu)$$

- Uniform priors
 - $0.06 \text{ eV} < M_\nu < \text{few eV}$
 - $t_{\text{sun}} < \text{age} < \text{few 10Gyr}$

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 - $0.06 \text{ eV} < M_\nu < \text{few eV}$
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- 2D posterior probability distribution
- Obtained with MCMC* and KDE*
(see *notebook*)

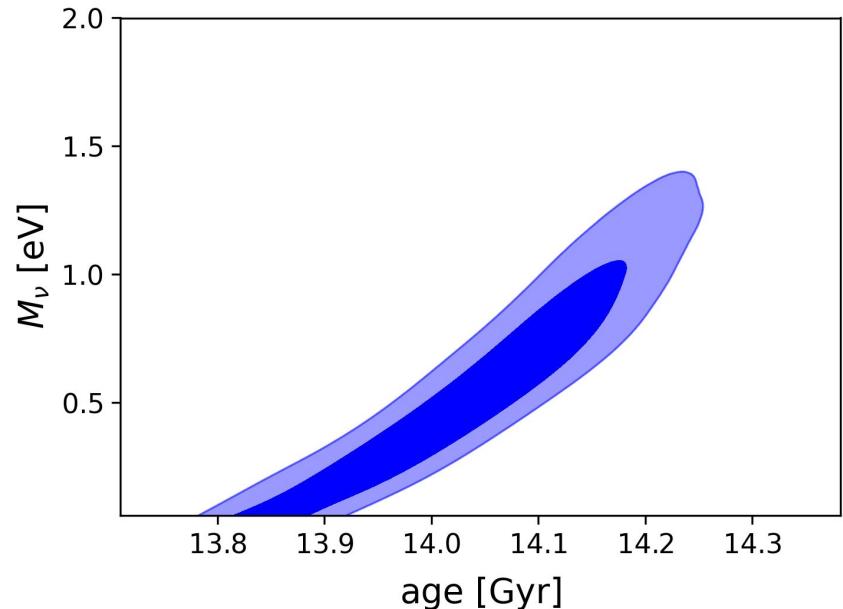


*MCMC: Monte Carlo Markov Chains, KDE: Kernel Density Estimation

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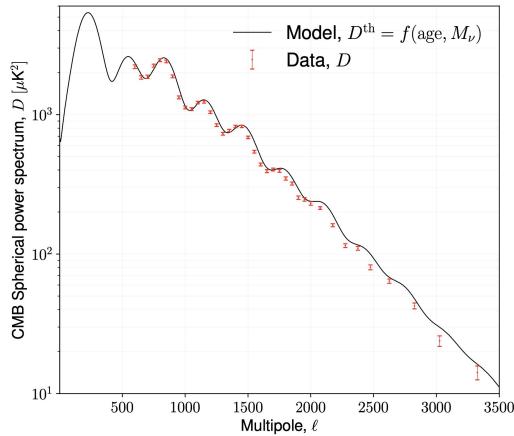
- **Bayesian credible regions** (68% and 95%CL)



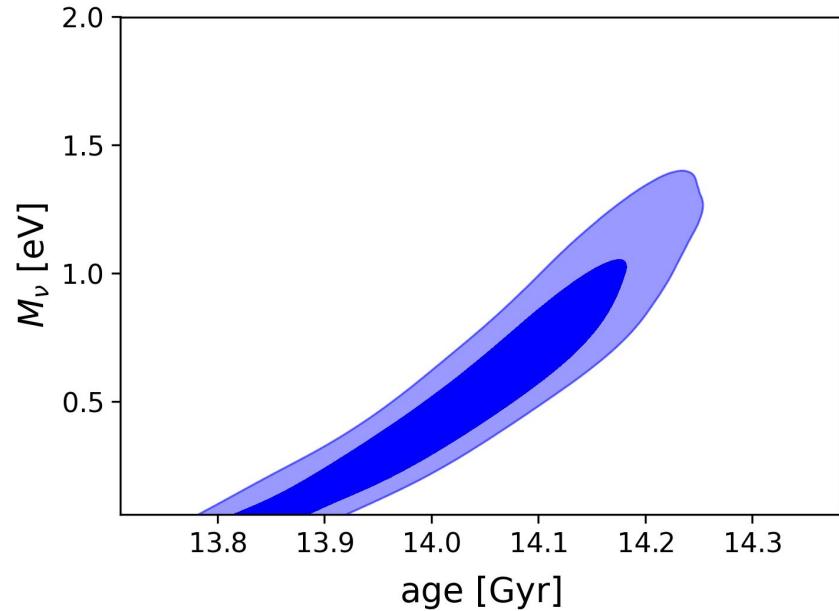
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- **Bayesian credible regions** (68% and 95%CL)
- Parameter degeneracies

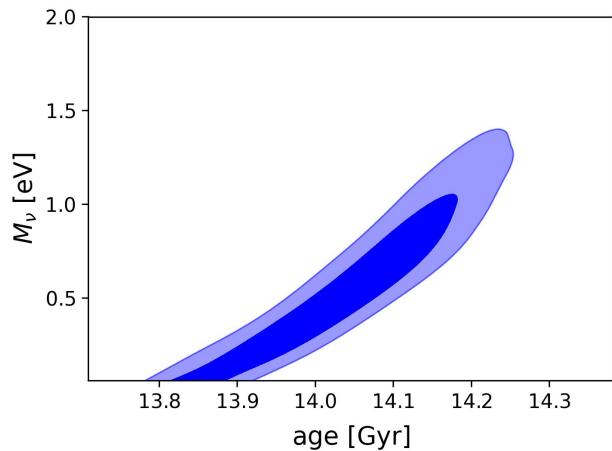


age 14.70
m_nu 0.97



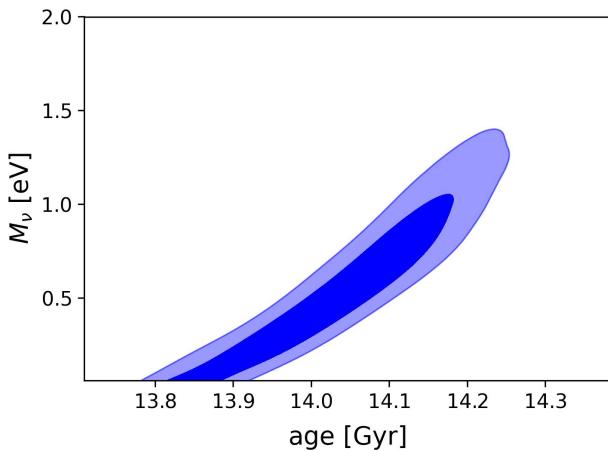
Eliminating a “*nuisance*” parameter

- Give-up hope of knowing neutrino mass precisely
- How old is the universe?



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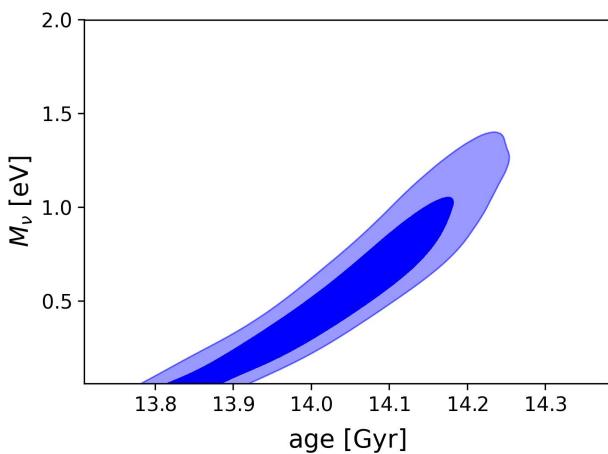
Marginalisation



$$P(\text{age}|D) = \int P(\text{age}, M_\nu|D) dM_\nu$$

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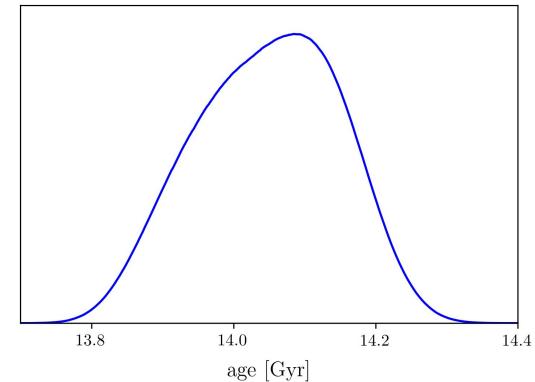


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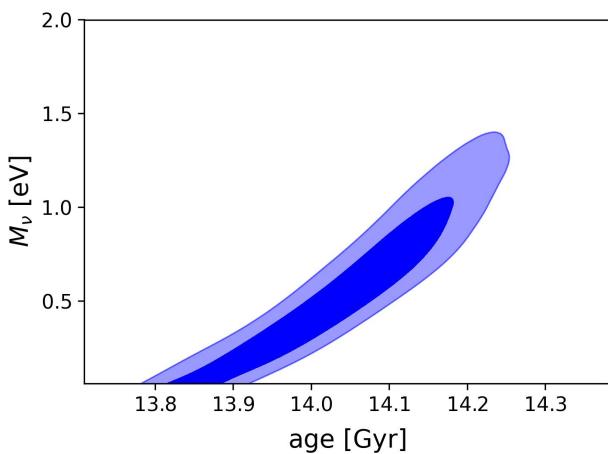
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1D marginalised posterior PDF



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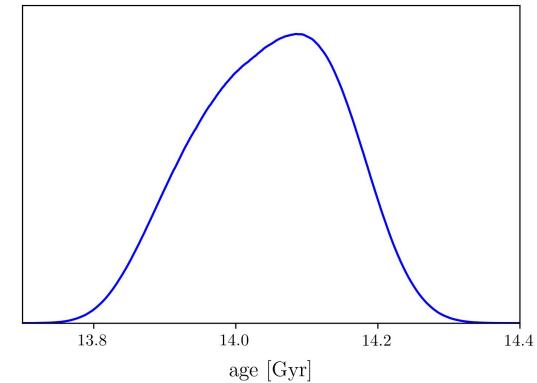


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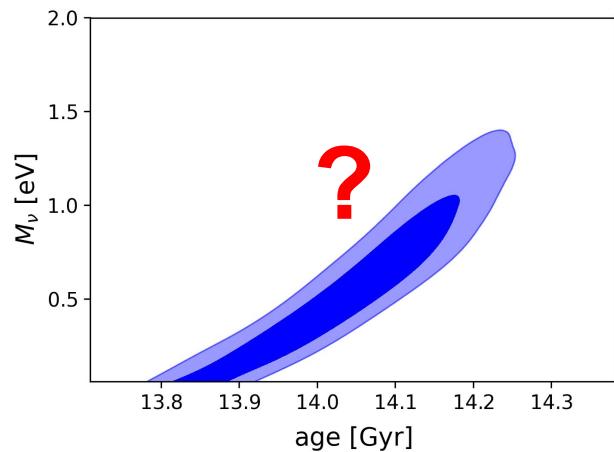
1D marginalised posterior PDF



age = 14.05 ± 0.1 Gyr(68%CL)

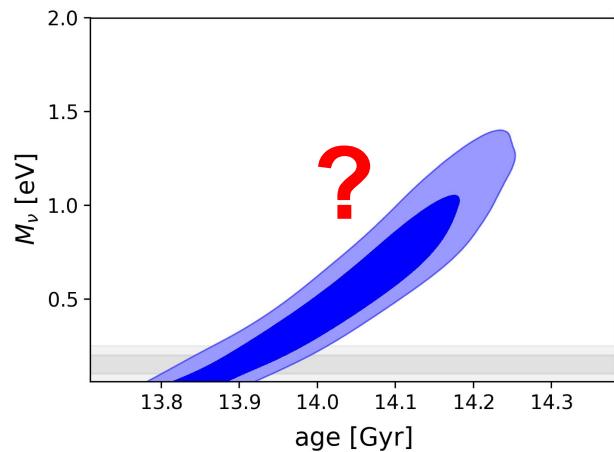
Combining experiments

- A new neutrino mass measurements is available $P(M_\nu) \sim \mathcal{N}(\mu = 0.15, \sigma = 0.05)$



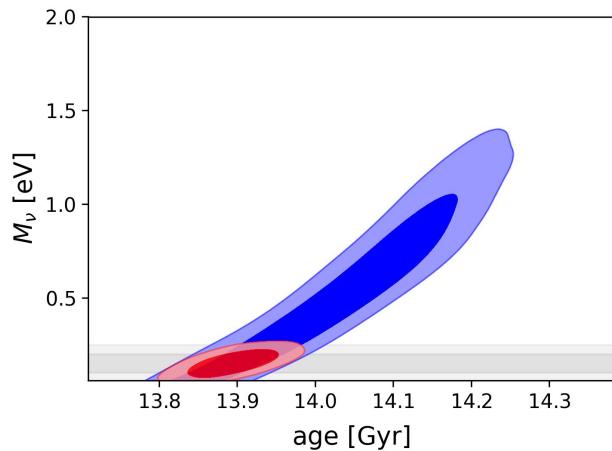
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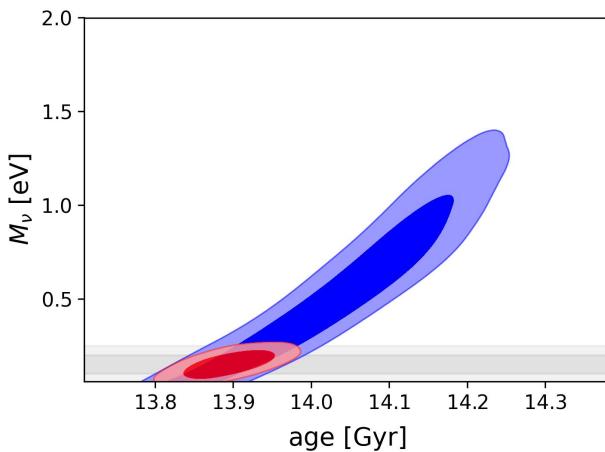
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Combining experiments

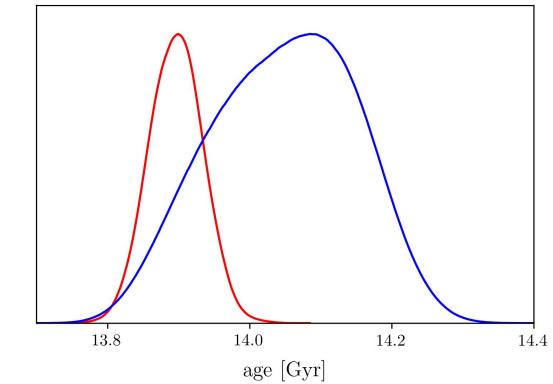
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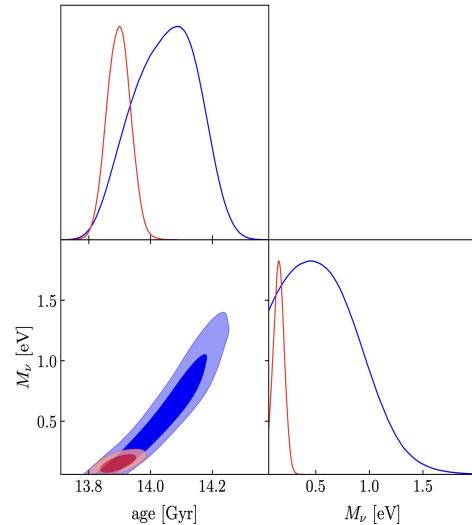


$$\text{age} = 14.90 \pm 0.04 \text{ Gyr}(68\% \text{CL})$$

Bayesian inference summary and key points

"What you know about θ after the data arrive is what you knew before [$P(\theta)$], and what the data told you [$P(D|\theta)$]." (MacKay)

- Interpretation of Bayes theorem $P(\theta|D) = \frac{P(D|\theta)P(\theta)}{P(D)}$
- Effects of prior choices
- Calculating the posterior probability distribution
- Bayesian credible intervals
- Marginalisation over nuisance parameters
- Adding new information/Combining experiments
- Approach directly applicable to large number of free parameters and priors



References

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