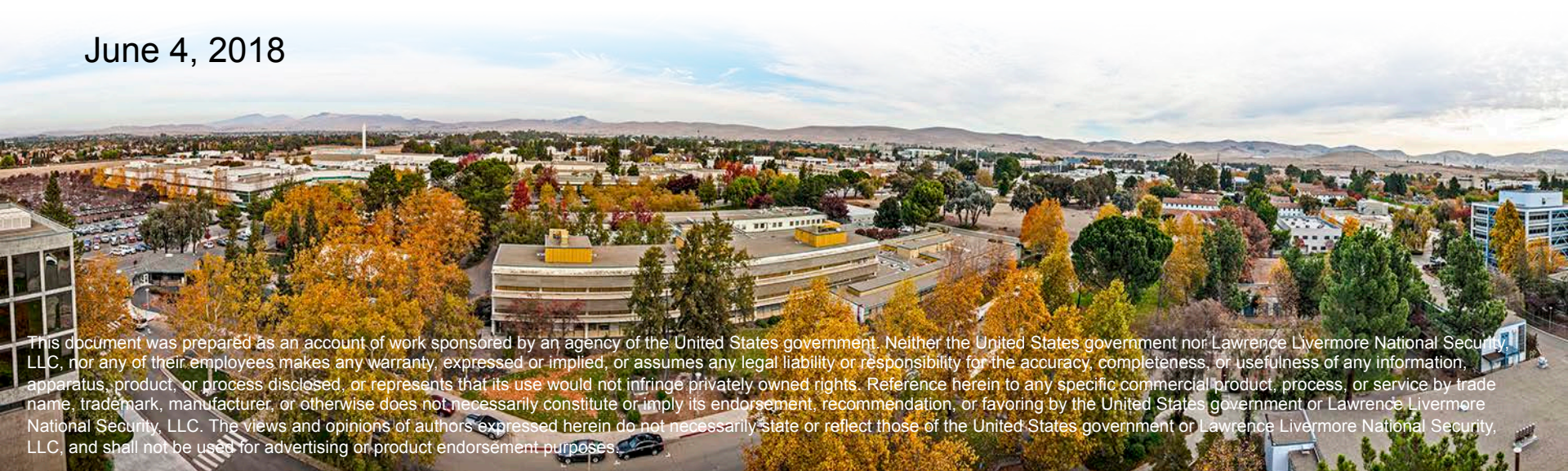


# CRETIN

## Session 9

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June 18, 2018

An aerial photograph of a city, likely Livermore, California, showing a mix of residential and commercial buildings, trees with autumn foliage, and a clear sky with some clouds. The city is nestled in a valley with mountains visible in the distance.

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# Session topics

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1. Stark lineshapes
2. Escape factors

# Line profiles

Line profiles are determined by multiple effects:

- Natural broadening ( $A_{I2}$ ) - Lorentzian
- Collisional broadening ( $n_e, T_e$ ) - Lorentzian
- Doppler broadening ( $T_i$ ) - Gaussian
- **Stark effect** (plasma microfields) - complex

Stark effect

- Splits and shifts lines  $\rightarrow$  additional components + broadening
- Increases strongly with quantum number  $n$  ( $n^{6-7}$ )
- Increases with  $n_e$  ( $n_e^{2/3}$  for  $e^-$  broadening)
- Decreases with  $Z$  ( $Z^{-1}$  for hydrogenic  $\rightarrow \Delta E / E \sim Z^{-3}$ )

**Total** is a lineshape code which runs standalone or as part of Cretin

- Stark effect from electron collisions + ion microfields + ion dynamics
- Zeeman splitting

**SLS** is a rewritten version of Total not completely integrated into Cretin

**Total (and SLS) are located in /usr/apps/cretin/bin**

# Specifying transitions for Stark lineshapes

For line transport:

**line** *iline iz iso i1 iso i2*  
**linetype** *total (totalb)*

For spectral calculations:

**stark transition** *iz iso i1 i2* - or -  
**stark manifold** *iz iso n1 n2*

Total requires ( $n, l, j$ ) detailed levels + transition matrix elements

- available in dca\_xxk models
- additional switches control Total options (88, 114-117)

Example: dca\_18k H- $\alpha$  transitions

data	phxs								
d	1	1	1	2	1.37957E-01	3.73646E+00	0.00000E+00	5.82580E-02	
d	1	1	1	3	1.90219E-08	3.73629E+00	0.00000E+00	2.16322E-05	
d	1	1	1	4	2.73463E-01	3.73105E+00	0.00000E+00	8.19629E-02	

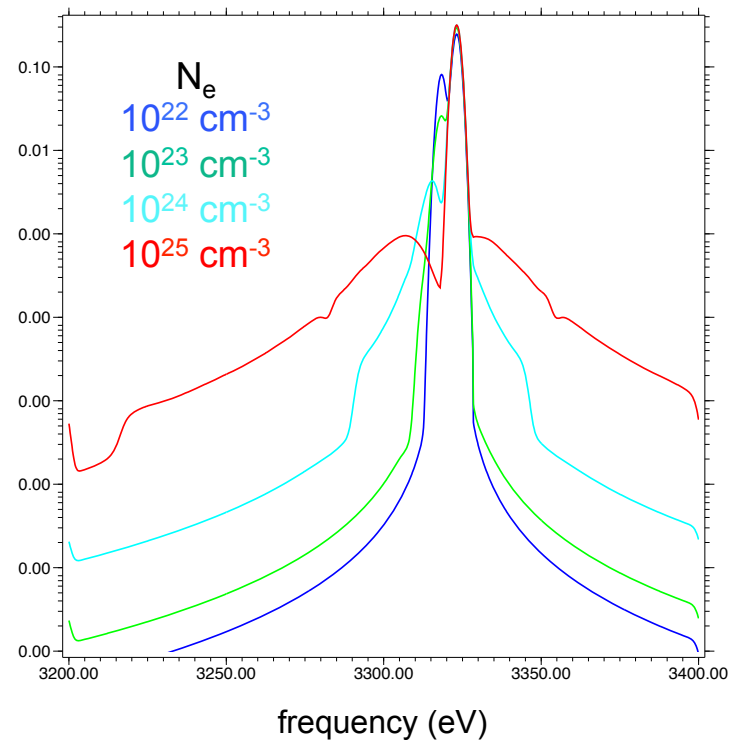
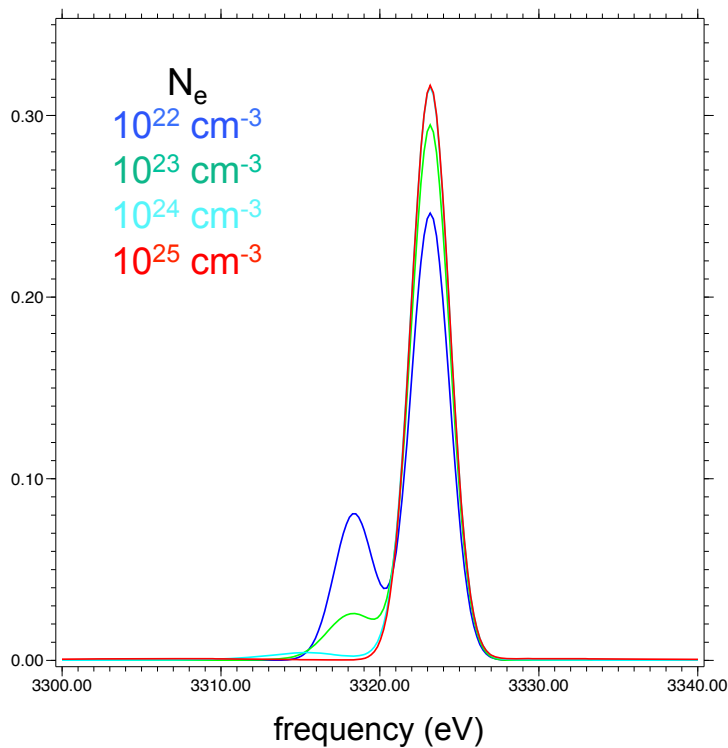
For transitions not treated with Total:

- broadening can be added if  $\text{switch}(52) < 0$
- formulas provided by Hans Griem (aimed at dense plasmas)
- both Doppler core and Lorentzian wing are affected if  $\text{switch}(52) = -1$

# Ar H- $\alpha$ Stark profile

- Doppler widths dominate at low density
- Larger ( $J=3/2$ ) component has no linear Stark broadening
- Wings fall off as  $(\Delta\nu)^{-5/2}$  for frequencies  $\Delta\nu >$  collision frequency

$$T_e = T_i = 5 \text{ keV}$$



# Escape factors

An approximate treatment of trapped radiation which (usually) assumes:  
uniform spatial conditions (+ line profile remains the same)  
2-level atom treatment is sufficient (with negligible upper-level population)  
isotropic emission with complete redistribution

Definitions:

monochromatic escape probability

$$p_e(\nu) = e^{-\tau_\nu}, \quad \tau_\nu = \int \alpha_\nu^0(r) \phi_\nu dr$$

escape probability

$$\langle p_e \rangle = \int \phi_\nu p_e(\nu) d\nu$$

escape factor

$$P_e = \frac{1}{4\pi} \oint \langle p_e \rangle d\Omega$$

$$\bar{P}_e = \frac{\int P_e dV}{\int dV}$$

averaged escape factor

- Escape probabilities require knowledge of the line profile
- Escape factors also require knowledge of the complete problem geometry, but are applied locally

# Applying escape factors

Two-level atom (steady-state) rate equation:

$$n_1(B_{12}\bar{J}_{12} + C_{12}) = n_2(A_{21} + B_{21}\bar{J}_{12} + C_{21}) \quad \bar{J}_{12} = \int_0^\infty J_\nu \phi(\nu) d\nu$$

The escape factor replaces radiative terms with decreased spontaneous emission:

$$n_2(A_{21} + B_{21}\bar{J}_{12}) - n_1B_{12}\bar{J}_{12} \rightarrow n_2A_{21}P_e \quad \Rightarrow \quad n_1C_{12} = n_2(P_eA_{21} + C_{21})$$

Comments:

- Iron's theorem says this is correct on average (emissivity-weighted spatial)
- This substitution happens for all transitions treated with escape factors
$$A_{ij} \rightarrow P_e^{ij} A_{ij}$$
- Since  $P_e$  depends on populations, this adds non-linearity to the rate equations
- $P_e$  incorporates global information through the optical depth
- Escape factors are used to calculate populations (absorption, emission)  
Radiation transport is used to calculate intensities along a given path

$$I_\nu = \int_0^s e^{-\alpha_\nu s} \eta_\nu ds$$

# Choosing the appropriate escape factor

- Escape factors depend on line shape and geometry
- Analytic approximations to  $P_e$  are available for
  - Doppler, Lorentzian, Voigt profiles + Stark (in wings)
  - planar, cylindrical, spherical geometry (central location)

What about Doppler shifts?

- Sobolev approximation accounts for a constant velocity gradient  $dV/ds$ 
  - photons shift frequency w.r.t. local frame by  $\frac{\Delta v}{v} = \frac{\Delta s}{c} \frac{dV}{ds}$
  - in the Sobolev limit of large  $dV/ds$

$$\langle p_e \rangle \rightarrow \frac{1 - e^{-\tau_s}}{\tau_s} \quad \tau_s = \frac{\Delta v_D}{dV/ds} \quad \leftarrow \text{escape probability}$$

Other issues –

- Overlapping lines / continuum opacity
- Continuum opacity
- Non-central locations in non-planar geometry
- Doppler shifts not in Sobolev limit
- Other radiation fields



# Escape factors in 0D simulations

In a 0D simulation, the goal is (usually) to calculate the average emissivity

- Calculate averaged escape factors  $\bar{P}_e$  by evaluating  $P_e$  for optical depth

$$\tau_v = \alpha_v \langle R \rangle, \quad \langle R \rangle = 4V/S \quad \leftarrow \text{mean chord}$$

- Evaluate escaping intensity for desired line-of-sight

Ref: G.J. Phillips, J.S. Wark, F.M. Kerr, S.J. Rose, R.W. Lee, HEDP **4**, 18-25 (2008)

Important points:

- This can give the correct average emission for a single transition
- Averaging transitions does not produce average plasma properties  
→ may not give the correct emission for multiple transitions
- Averaging spreads out the radiative boundary layer over the whole plasma  
→ does not produce boundary layer effects

# Escape factors in Cretin

## 0-dimensional

- uses  $\bar{P}_e$  for each transition
- special spectral edits for integrated quantities (*isp\_0d*, *esp\_0d*, *tausp\_0d*, *trsp\_0d*)

## 1-dimensional

- applies  $P_e$  at each position averaged over +/- directions
- $\tau$  in each direction calculated with scaled integrated column density of charge state
- assumes constant line profile, fractional lower state population

## 2- / 3-dimensional

- applies  $P_e$  at each position averaged over +/- directions along 1 or more axes
- option to specify directions for averaging is partially complete

Escape factors are also applied to extant radiation fields

# Escape factors in Cretin

## Main controls

- switch 33:** use escape factors for lines only if  $<0$ , all photoexcitations if  $>0$   
     $\pm 1$ : static     $\pm 2$ : Sobolev     $\pm 3$ : interpolation between static & Sobolev  
     $\pm 4$ : generalized escape factor                      ← **recommended**  
            includes continuum, velocity gradients, arbitrary profile & geometry
- param 54:** minimum optical depth for calculating escape factors

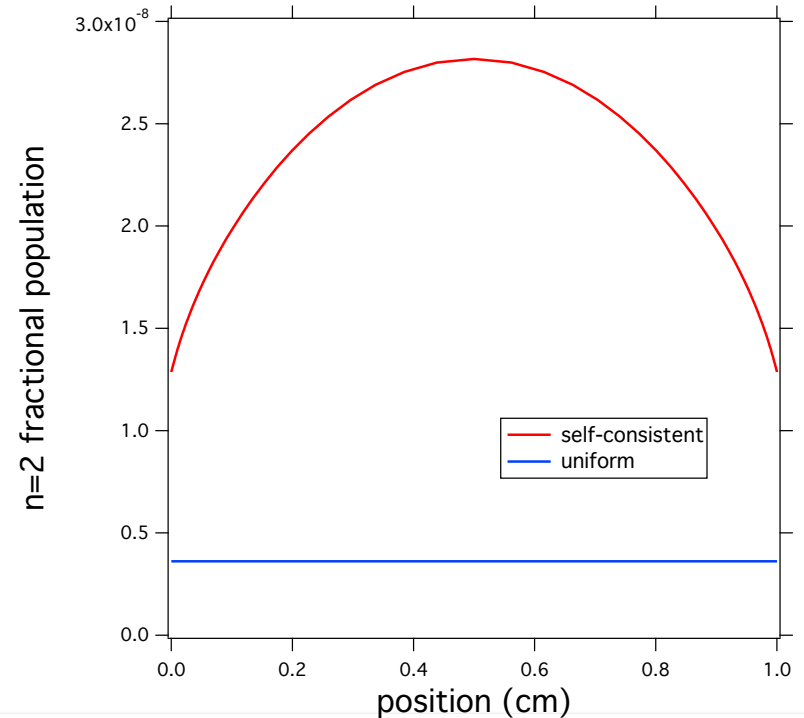
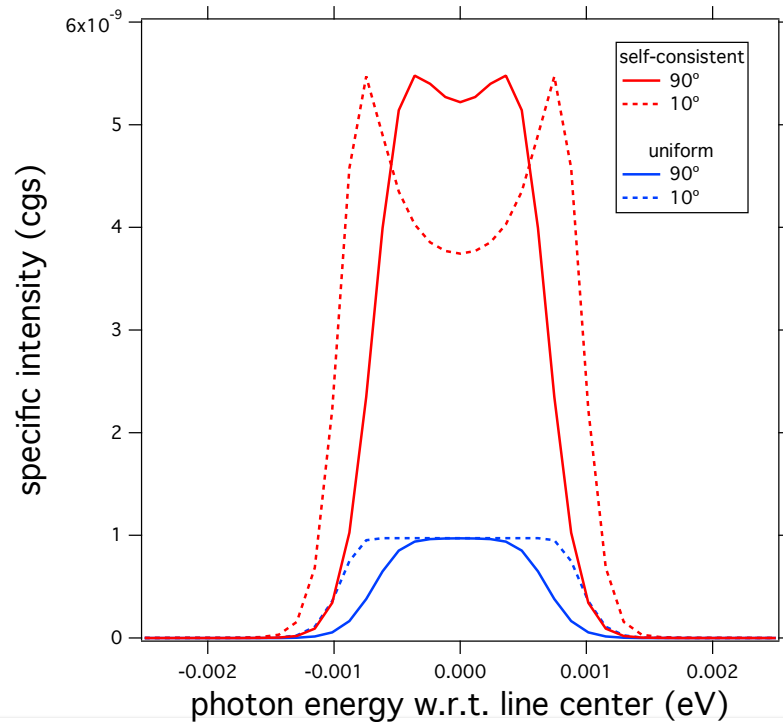
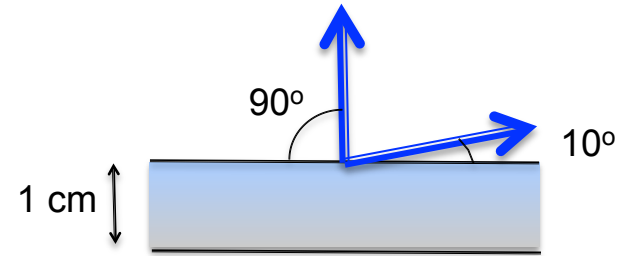
## Additional controls (mostly for OD applications)

- switch 78:** choose line profile (Voigt, Doppler, Lorentzian)  
**switch 79:** include continuum opacity?  
**switch 80:** specify problem geometry  
**switch 81:** single- or double-sided  
**switch 82:** use integrated density, local density, or param(53)  
**switch 154:** choose axes for integrations in 2D/3D  
**param 53:**  $\Delta r$  or  $\Delta(nr)$  to use for escape factors                      (average chord)  
**param 139:**  $\Delta r$  or  $\Delta(nr)$  to use for escape factor edits              (LOS distance)

combinations of switch(79) + switch(80) can also choose specific analytic or tabulated escape factor formulations

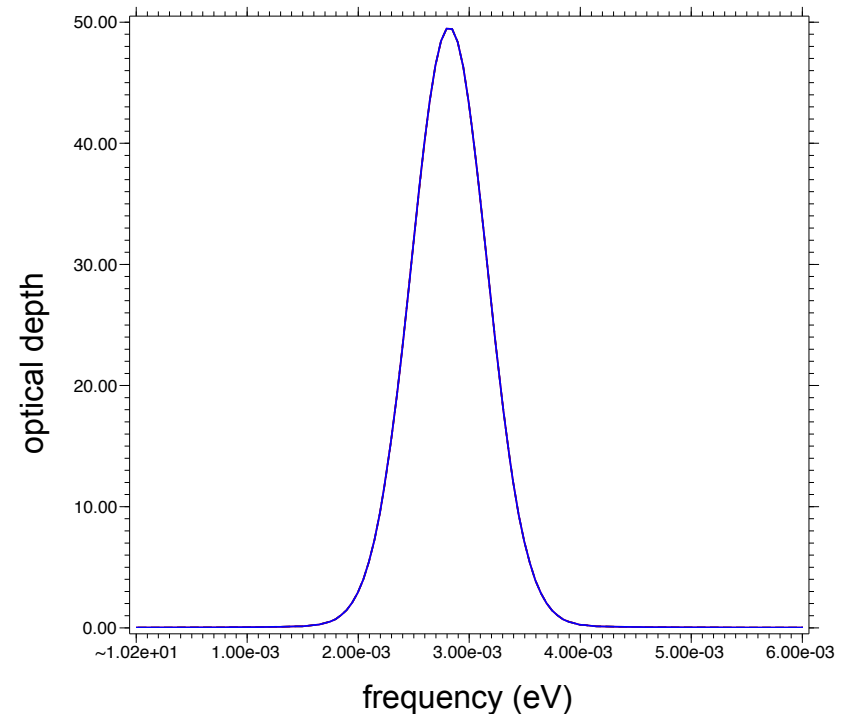
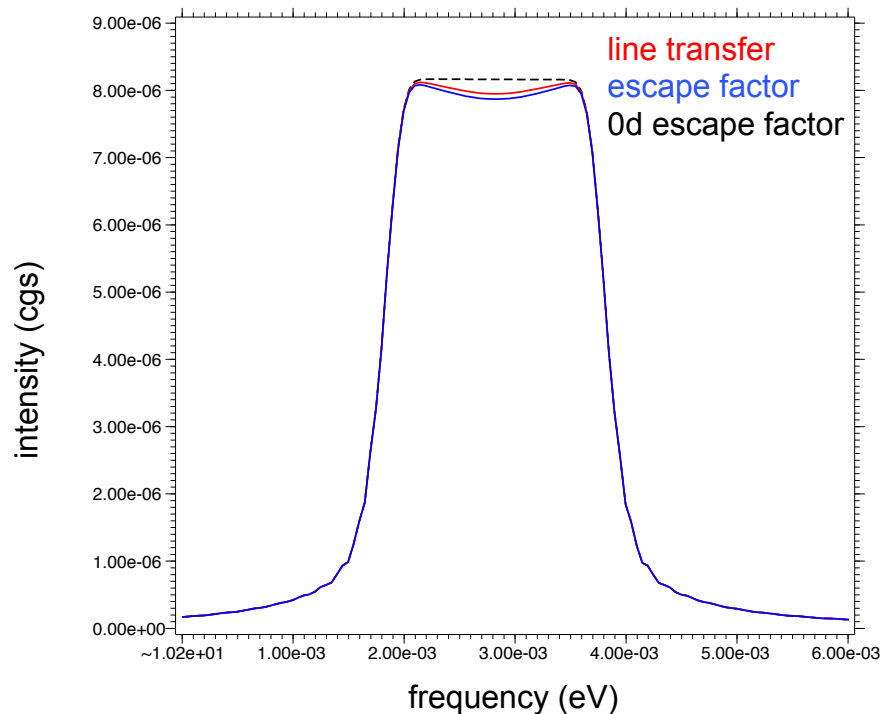
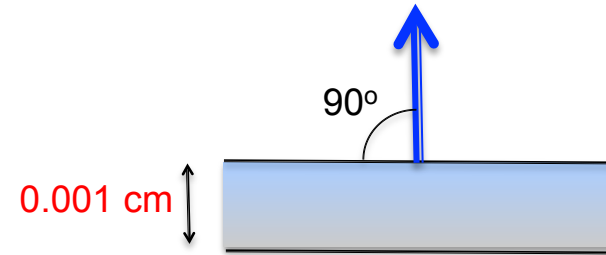
# Example – Hydrogen Ly- $\alpha$

- Uniform conditions:  $T_e = 1$  eV,  $n_e = 10^{14}$  cm $^{-3}$
- Self-consistent solution displays effects of
  - Radiation trapping / pumping
  - Non-uniformity due to boundaries



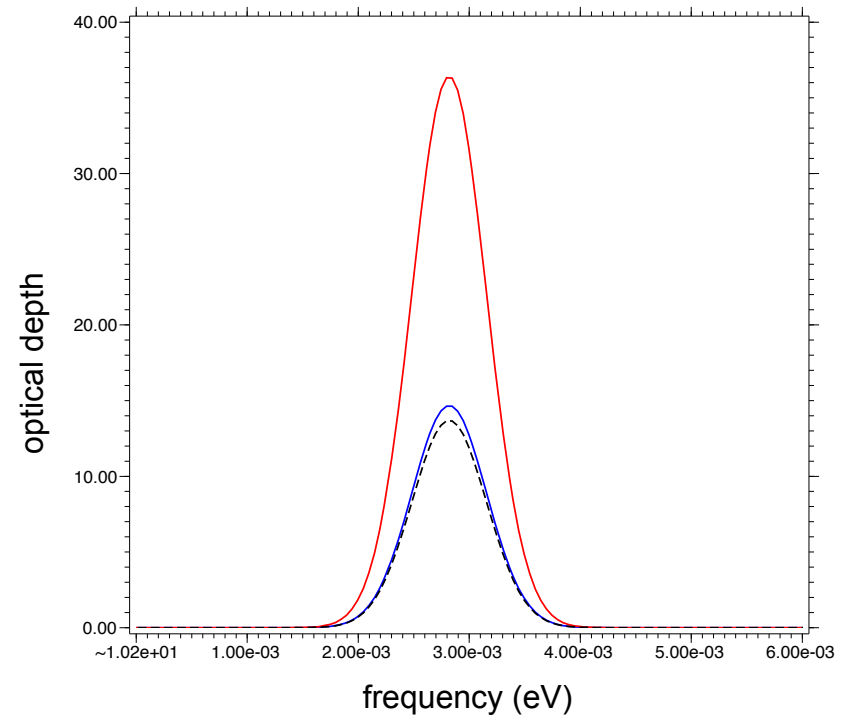
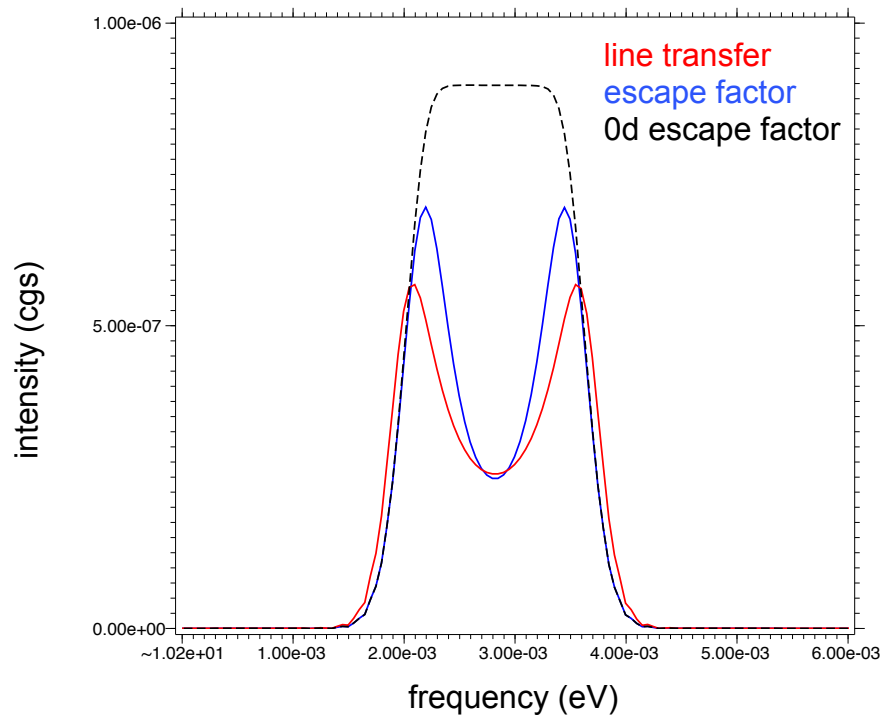
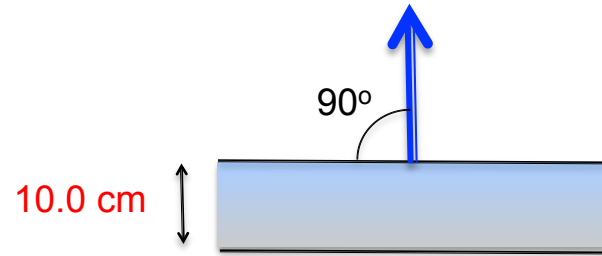
# Hydrogen Ly- $\alpha$ with escape factor – easy case

- Uniform conditions:  $T_e = 1$  eV
- High density:**  $n_i = 10^{18} \text{ cm}^{-3}$



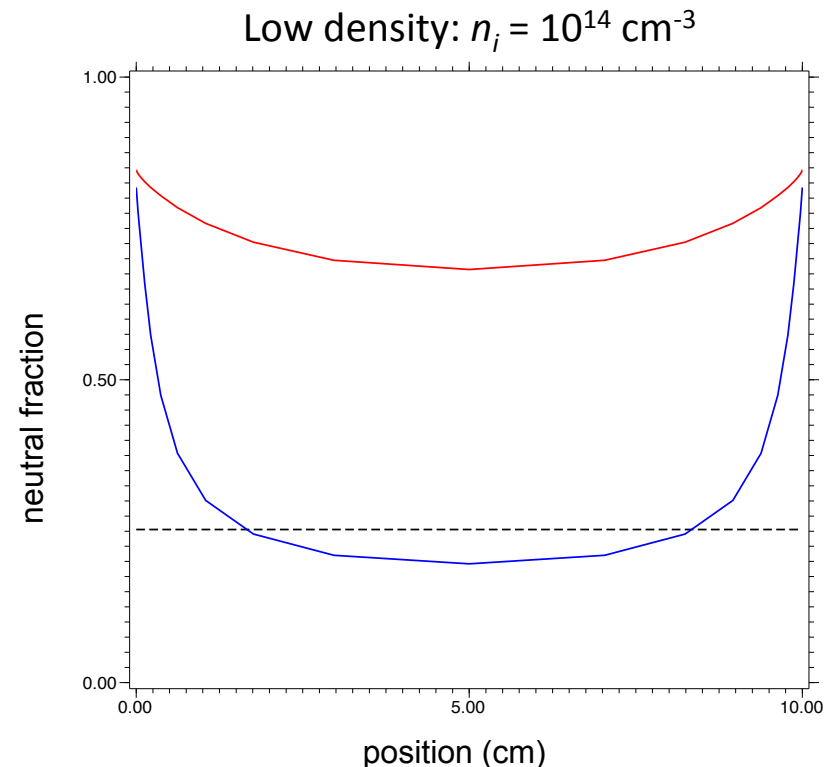
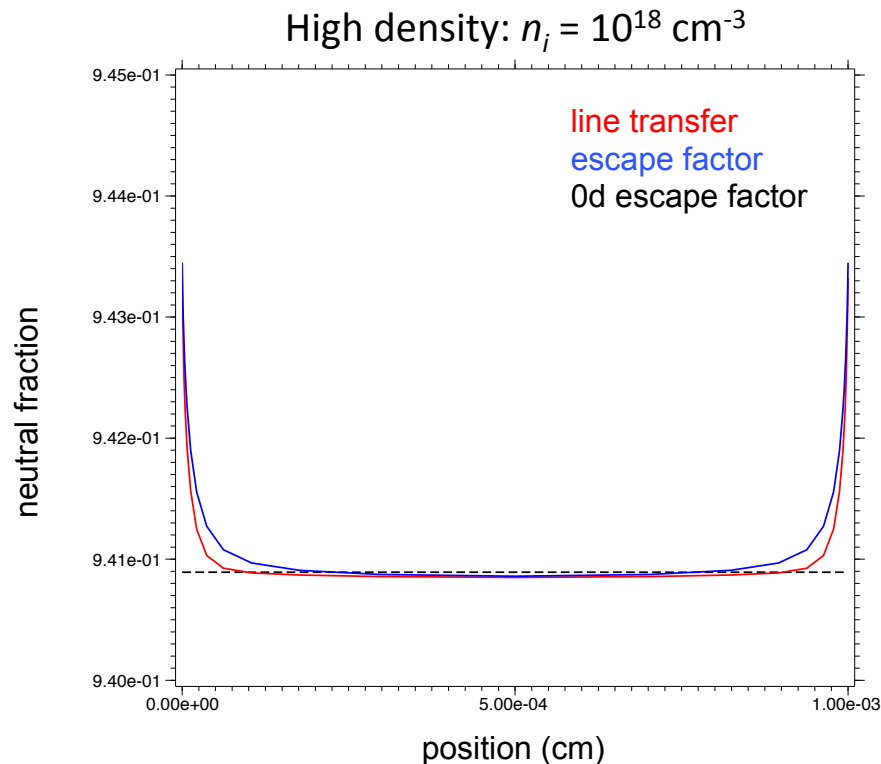
# Hydrogen Ly- $\alpha$ with escape factor – hard case

- Uniform conditions:  $T_e = 1$  eV
- Low** density:  $n_i = 10^{14} \text{ cm}^{-3}$



# Hydrogen Ly- $\alpha$ with escape factor

- Ionization is dominated by collisions at high density, radiation at low density
- Position-dependent escape factors mimic boundary layer, but not ionization
- 0d escape factor assumes average conditions with no boundary layer
- Line transfer allows photons to escape by changing frequency



# Advantages / disadvantages of escape factors

## Advantages:

- Much faster than line transfer, can be applied to all photoexcitations
- Good results for isolated lines (with the right escape factor) under most conditions
- Escape factors give an approximation to radiation transfer

## Disadvantages:

- Choosing or calculating the right escape factor can be difficult
- Applying escape factors in OD for multiple transitions can be problematic
  - average rates for each transition  $\neq$  rates for average conditions
- Calculating escape factors in non-uniform plasmas requires spatial information similar to that of radiation transfer (i.e. ray tracing)
- Escape factors give an approximation to radiation transfer which cannot be iteratively improved

## Other notes:

- Many other variations exist in a large literature
- “Escape factors” for photoionizations appear in the literature (and some codes)
  - may or may not be better than nothing