

# CRETIN

## Session 6

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# Session topics

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## Radiation transport

- "Flavors": continuum, lines, spectral
- Goals: energetics, kinetics / populations / spectra
- Main commonalities
  - considerations for meshes in space, angle and frequency
  - edit specifications
- Major differences
  - setting up frequency meshes
  - options for including physical effects
  - methods for achieving convergence

# Radiation transport “flavors”

Continuum, lines and spectra are treated separately for efficiency

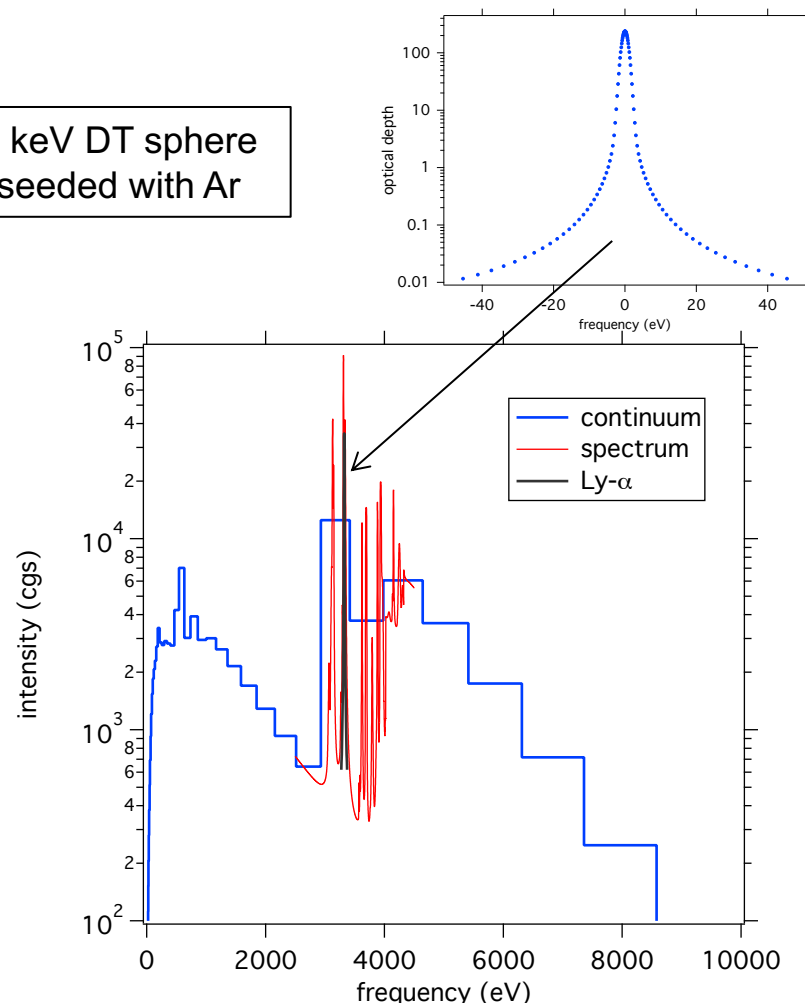
Iterated to consistency with atomic kinetics (and other processes):

- continuum radiation coarsely-binned over full energy range for evaluating photo rates
- line radiation finely-binned for resolving individual line profiles

Evaluated after convergence:

- spectral radiation at arbitrary energies to resolve features of interest

1 keV DT sphere  
seeded with Ar



# Spatial meshes

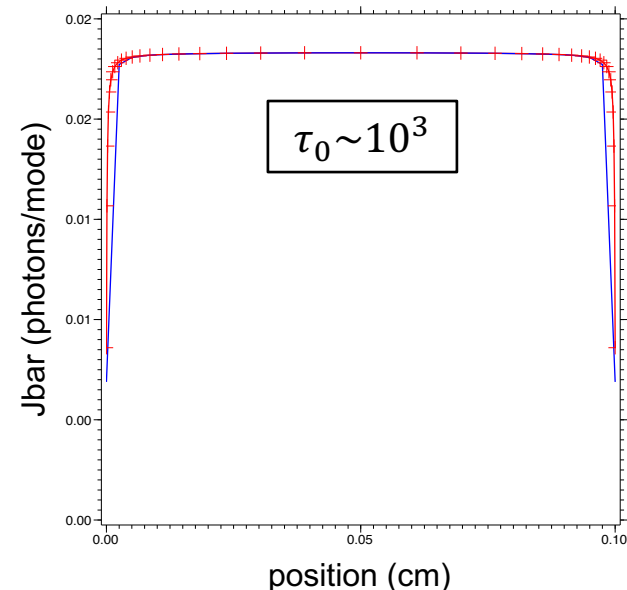
- All types of radiation transport in a Cretin run use the same spatial mesh
  - mixing centerings (for rad-hydro runs) is possible but discouraged
- The main consideration is resolving boundary layers
  - a) Boundaries produce non-uniformities in the radiation field within  $\tau_v \sim 1$
  - b) The radiation field produces non-uniformities in the material properties
  - c) Escaping radiation may originate primarily in the boundary layer
- This can be an issue for quantities sensitive to the boundary layer

## Example:

Optically thick (multiple) line transfer

Two mesh specifications:

- 1) `rlin 1 41 0.0 1.0`
- 2) `rgeom 1 21 0.0 0.5 1.e-4 +1`  
`rgeom 21 41 0.5 1.0 1.e-4 -1`



# Symmetries (reflective boundary conditions)

- Symmetries are specified with **switch(34)** (or by boundary conditions)
  - 1D:  $\text{switch}(34) > 0 \Rightarrow$  symmetric about  $ir=1$
  - 2D/ 3D: each digit specifies symmetries on a logical boundary
    - digit - 2D k: 10's, l: 1's      3D k: 100's, l:10's, m: 1's
    - value - 0: none, 1 or 3: lower boundary, 2 or 3: upper boundary
- e.g.  $\text{switch}(34) = 13$  in 2D or 133 in 3D gives one open boundary at  $k=k_{\max}$

Example: sphere in RZ geometry

**geometry rz**

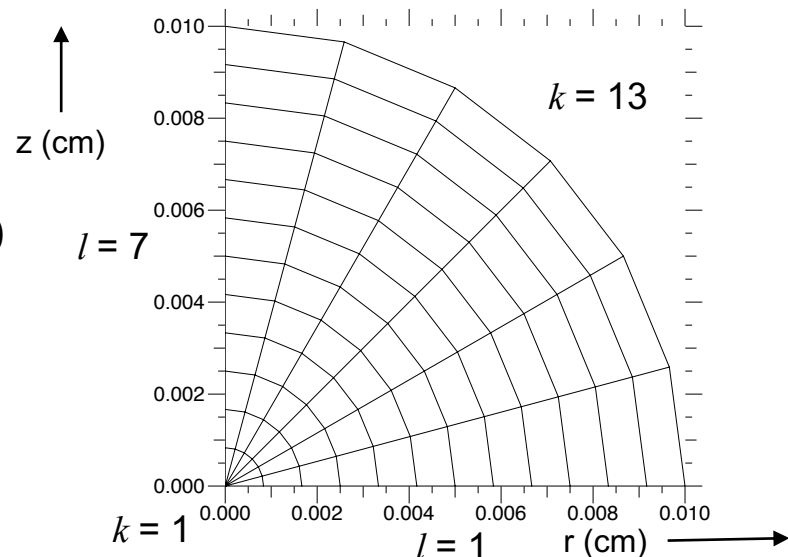
**wedge 1 13 1 7 0.001 90.0 0.0 1.0 1.0**

**switch 34 13**

Symmetry point @ origin:  $k=l=1$

Symmetry line @ z-axis:  $l=7$

Symmetry plane @  $z=0$ :  $l=1$



**Symmetries are only valid along coordinate axes / planes**

# Boundary conditions and sources

Boundary conditions require position, direction, magnitude and frequency

- Position and direction come from **xfilebc** or **boundary** commands
- Intensity distribution can come from a simple formula or tabulated data
- **xfilebc** *ix mu phi multiplier isotropy k1 kmin kmax lmin lmax mmin mmax*  
**xfilebc** *ix mu phi multiplier isotropy xy xmin xmax ymin ymax zmin zmax*  
*ix*: index of xfile containing boundary data  
*isotropy* ≠ 0: boundary condition is treated as isotropic external field
- **boundary** *package type k1 [k2 l1 l2 m1 m2] {history | xfile | source} id [multiplier]*  
*package*: **radiation** or **all** (other options apply to other physics)  
*type*: **reflecting, isotropic, streaming** “
- **source jbdry** *ix black-body trad [multiplier]*  
**source jbdry** *ix emin emax {black-body | power-law a b} {history | xfile | profile | constant}*  
*{value | rate | integral | initial} {id | value} [multiplier] [ireg | k1 k2 l1 l2 m1 m2]*
- **source jnu ...** ← provides internal radiation sources

Source specifications sum to give the final spectrum  
Sources apply to continuum and spectral radiation

# Angular meshes

All types of radiation transport in Cretin use discrete ordinates ( $S_N$ ) methods with the same angular mesh

- exception: spectral transport may not need an angular mesh

The number of angles in the angular mesh depends on geometry

- 1D planar / spherical : 1 angle  $\mu = \cos(\theta)$
- 1D cylindrical, 2D, 3D : 2 angles  $(\mu, \phi)$

The angular mesh

1. Provides a basis for doing solid-angle integrals:

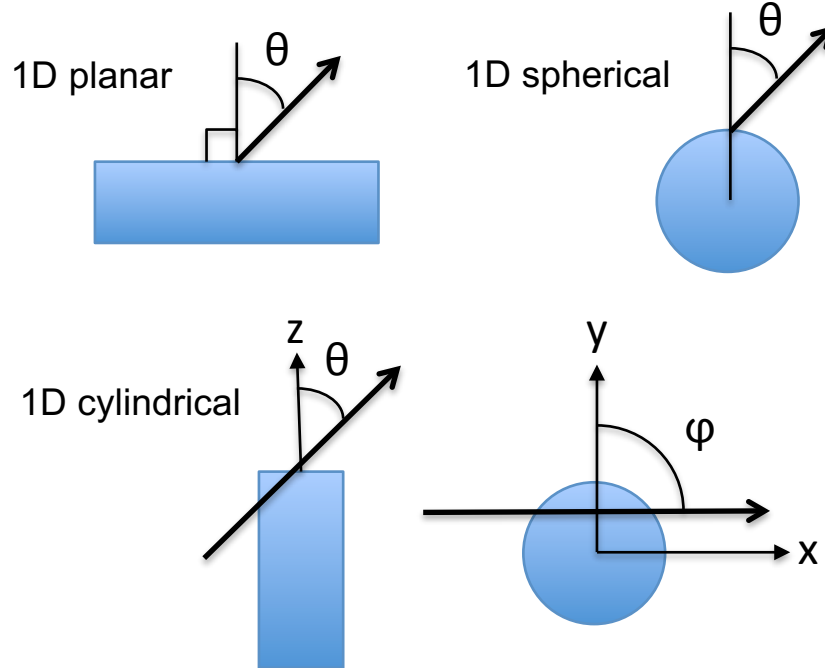
$$\int I_v d\Omega = \int_0^{2\pi} d\phi \int_{-1}^1 d\mu I_v(\mu, \phi) \rightarrow \sum_j \Delta\phi_j \sum_i \Delta\mu_i I_v(\mu_i, \phi_j)$$

2. Resolves angular variations in the radiation field

**Warning: resolving angular scales in 2D / 3D may be very difficult**  
— see John Castor's book "Radiation Hydrodynamics"

# Angle specifications – 1D

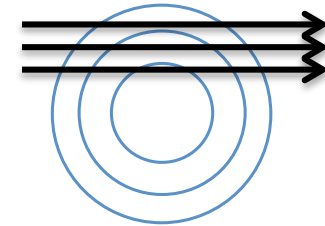
- The number(s) of angles are specified by **angles** *nmu* (*nphi*)
- The angular mesh depends on dimensionality and geometry:



Planar / cylindrical:  
*nmu* = # angles in  $90^\circ$

Spherical (*nmu*) / cylindrical (*nphi*):  
minimum # used for integrations

“long characteristics”



**In non-planar geometry, the total # of “rays” increases with the # of zones**



# Angle specifications – 1D

- Gaussian integration is used to integrate over  $\mu$  in planar / cylindrical geometry  $\Rightarrow$  values of  $\mu$  correspond to zeroes of Legendre polynomials

**switch 8** = 0 : Gaussian integration over  $[-1,1]$   
 $\neq 0$  : double-Gaussian integration over  $[0,1]$

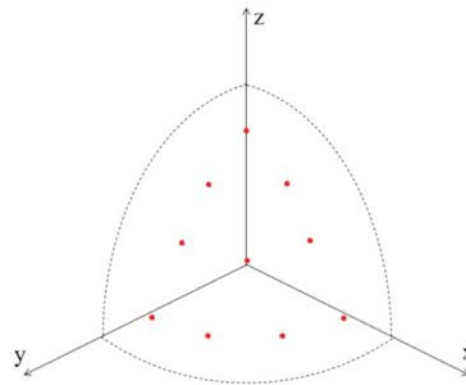
$$\int_{-1}^1 I_v(\mu) d\mu \quad \text{vs} \quad 2 \int_0^1 [I_v(\mu) + I_v(-\mu)] d\mu$$

A few general comments:

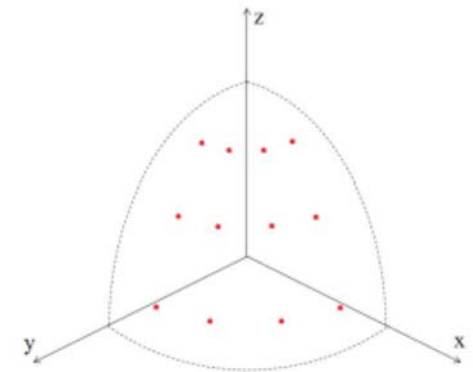
- All transport is 3-dimensional, with symmetries used to decrease the work
- Correspondence with  $S_N$  designation:  $N = 2 \times nmu$

# Angle specifications – 2D, 3D

- The number(s) of directions in a quadrant / octant are specified by angles *nmu* (*nphi*)
- Angle sets are determined by **switch 8** :
  - 0: Carlson set A
  - 1: Carlson set B (equivalent to double-Gaussian)
  - 2: product set with Gaussian *mu*
  - 3: product set with double-Gaussian *mu*
  - 4: level symmetric set (valid for *nmu* up to 10  $\Leftrightarrow S_{20}$ )
- Total number of directions per quadrant / octant:
  - 0, 1, 4:  $\frac{1}{2} \times nmu \times (nmu \pm 1)$
  - 2, 3:  $nmu \times nphi$



level symmetric – *nmu* = 4



product set – *nmu, nphi* = 3, 4

Angle sets A, B are described in: B. Carlson, “The Numerical Theory of Neutron Transport”, Methods in Comp. Physics, Vol. 1, 1963

## Edit specifications

- Edits for transport types have very similar names  
 e.g. *inu+*, *isp+*, *iline+*      for specific intensity  
       *cflux+*, *spflux+*, *iflux+*    for flux  
       *jnu*, *jsp*, *jline*             for angle-averaged intensity
  - Continuum / spectral transport are done in cgs units (*jnu*: erg/cm<sup>2</sup>/s/Hz/ster)  
 Line transport uses the photon distribution function (*jline*: photons/mode)  
 conversion factor  $2h\nu^3/c^2$  is available as: jmode iline
  - Directions are specified by the 4<sup>th</sup> edit index and must be defined:  
**editray idir mu phi type**                  *type\_ = 0*: use the closest existing direction  
    *≠ 0*: add this exact direction
- edit**
- |                                     |  |                         |  |                                 |
|-------------------------------------|--|-------------------------|--|---------------------------------|
| <b>xvar</b> energy                  |  | <i>sp_energy</i>        |  | <i>evline iline</i>             |
| <b>yvar</b> <i>inu+ 0 ir 0 idir</i> |  | <i>isp+ 0 ir 0 idir</i> |  | <i>iline+ iline 0 ir 0 idir</i> |
- In 1D geometries, +(-) specifies direction of increasing (decreasing) node index  
 In 2D/3D, +(-) designations are identical and (*mu,phi*) give direction of photons

# Other useful transport edits

Spectral frequency sub-mesh specification:

**spectrum** nspec E0 E1 ratio ["wavelength"]

**spectral-group** is ES0 ES1 ["wavelength"]

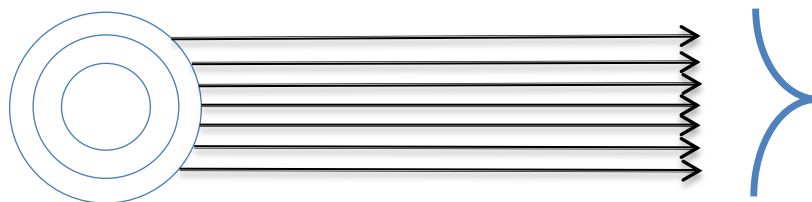
Area-integrated edits – *jnuarea*, *jsparea*, *lnuarea* – 1D spherical / cylindrical

**xvar** *sp\_energy* *is*

← *is* = 0 means full spectral range

**yvar** *jsparea* *is* *ir*

← area-integrated out to radius *ir*



2D/3D equivalent: detector edits – *cdetect*, *spdetect*, *idetect*

**detector** *idtct* x y [z] mu phi dΩ

← specifies position and orientation of detector

**detector** *idtct* x y [z] mu phi dΩ **extent** dr

+ spatial extent

Directions for detectors are specified from photons leaving the detector

# Frequency meshes

- Continuum radiation uses the same frequency mesh as the atomic kinetics

**ebins** *nbins emin emax [ratio]*

← or **fbins**, **wbins**

**energy** *e0 e1 e2 e3 ...*

← or **frequency**, **wavelength**

– or –

**pbins** | **rbins** *i0 i1 e0 e1*

← in units of keV

**pbinlin** | **rbin\_lin** *i0 i1 e0 e1*

- Spectral radiation

**spectrum** *nbins emin emax [ratio]*

Default behavior is to evaluate spectral absorption / emission coefficients at the exact spectral bin energies, which may miss narrow lines or peaks

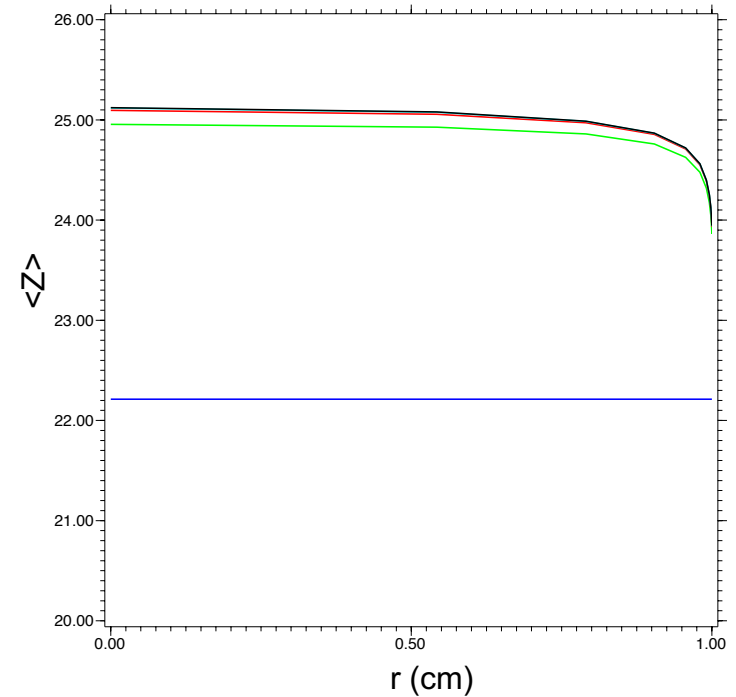
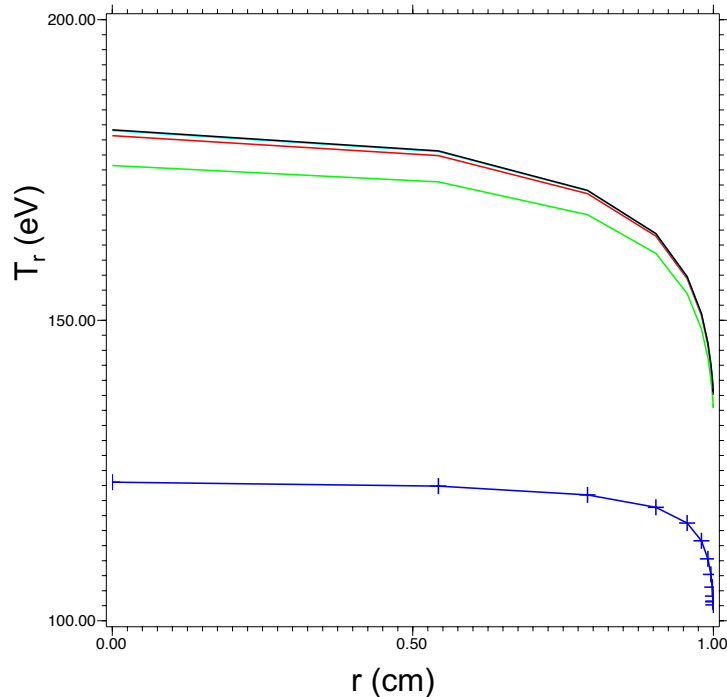
**switch 125**  $\neq 0$  : integrate over (constructed) spectral bin intervals

When starting from a restart or spectral dump, specifying a new frequency mesh in the (optional) generator replaces the original frequency mesh

**Multiple ebins, etc. or spectrum commands may be used**  
**Frequencies do not need to be specified in consecutive order**

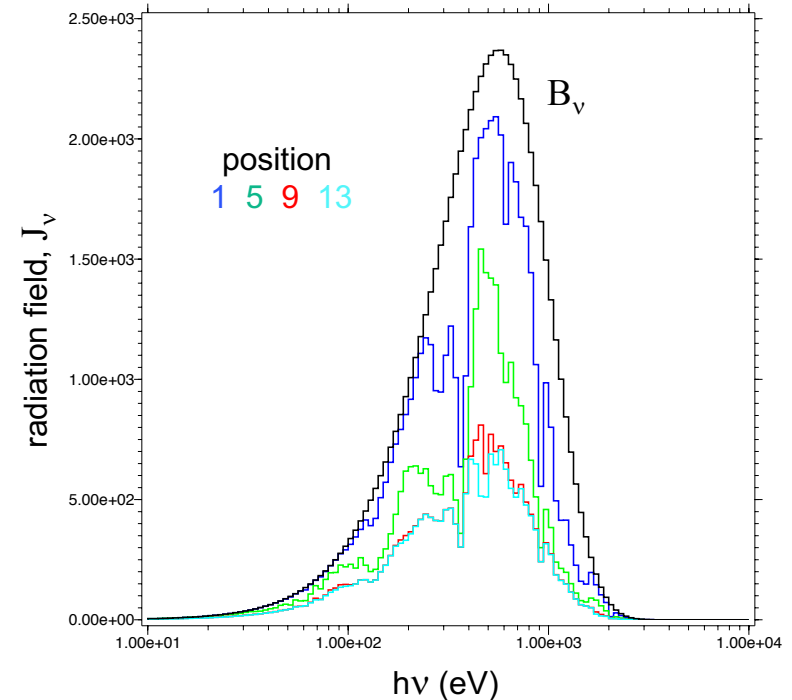
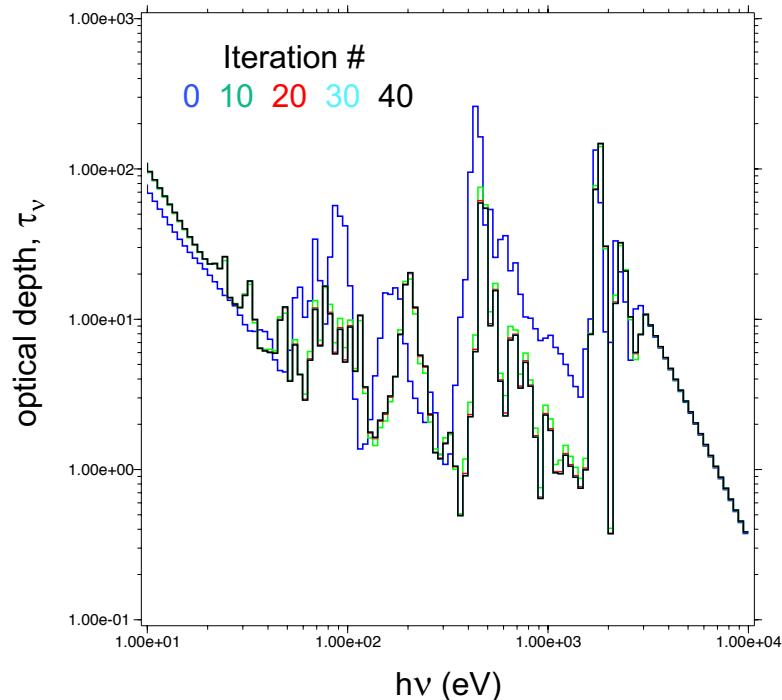
# Example: Kr sphere

- Uniform material conditions:  $T_e = 200$  eV,  $\rho = 0.01$  g/cm<sup>3</sup>
- Continuum transfer w/ vacuum boundary condition
- Initial conditions:  $T_r = 0$
- Spatial mesh: **rgeom 1 13 0.1 1.e-3 -1**
- Frequency mesh: **ebins 1 161 1.e0 1.e4**
- Iterated to  $10^{-3}$  in all charge states  $\rightarrow$  40 iterations



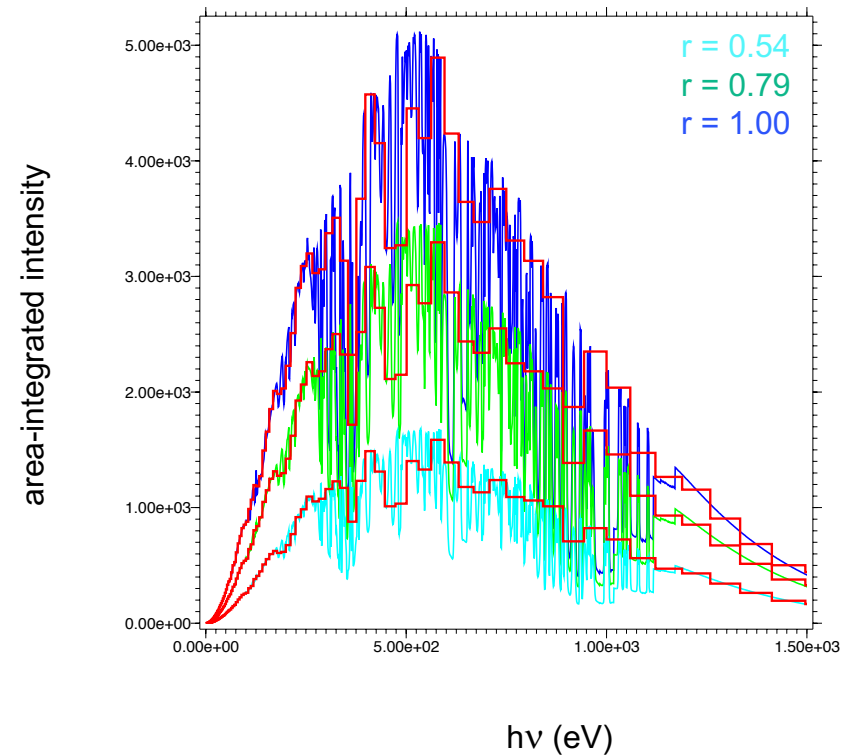
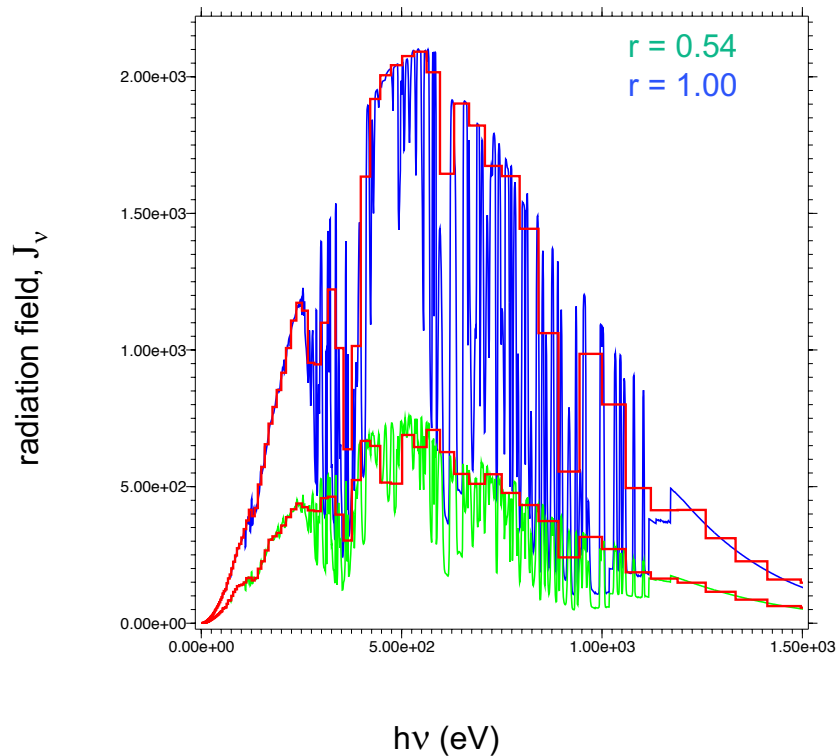
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# Example: Kr sphere

- Frequency mesh: **ebins 1 161 1.e0 1.e4**
- Spectral frequency mesh: **spectrum 2000 1. 2000.**
- Area-integrated edits (jsparea / jnuarea)





# Example: Kr sphere – 2D

- Spatial mesh: **geometry rz**  
**wedge 1 13 1 7 0.001 90.0 0.0 1.0 1.0** [ 1 13 1 91 ]  
**switch 34 13**
- Edit direction: **editray 1 1. 0. 1** (parallel to +z axis)
- Boundary condition: **xfilebc 1 1. 0. 1. 1 kl 13 13 4 5** [ xy .0045 .005 .008 .01]  
**source jbdry 1 2.0e3 3.0e3 black-body constant value 1.e3 1.e-2**

