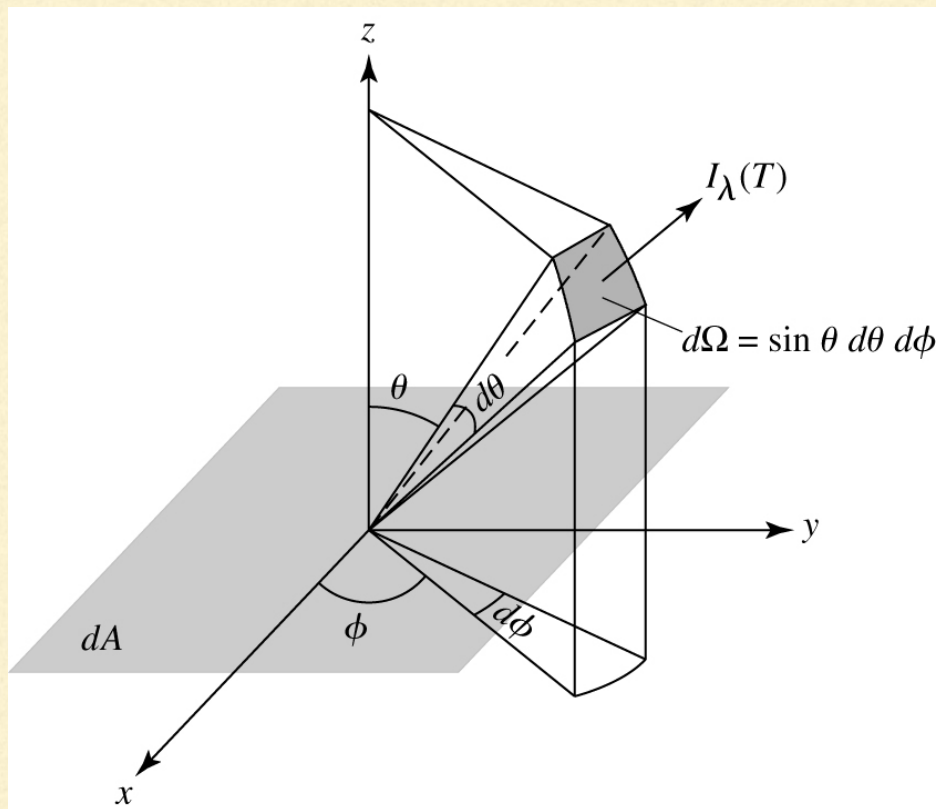


# SPECIFIC INTENSITY



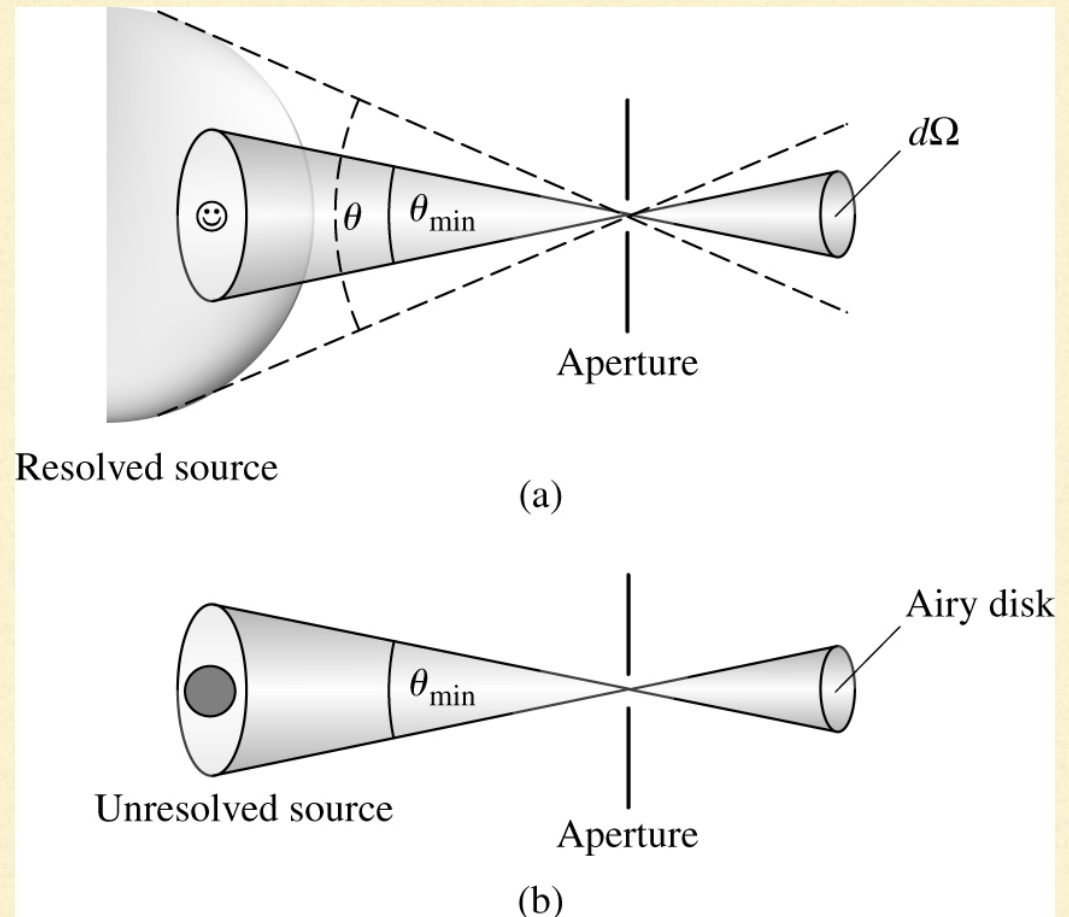
- $I_\lambda$  = energy / time / area / solid angle / wavelength
- $I_\nu$  = energy / time / area / solid angle / frequency
- Function of position, direction, wavelength, and polarization

# SPECIFIC FLUX

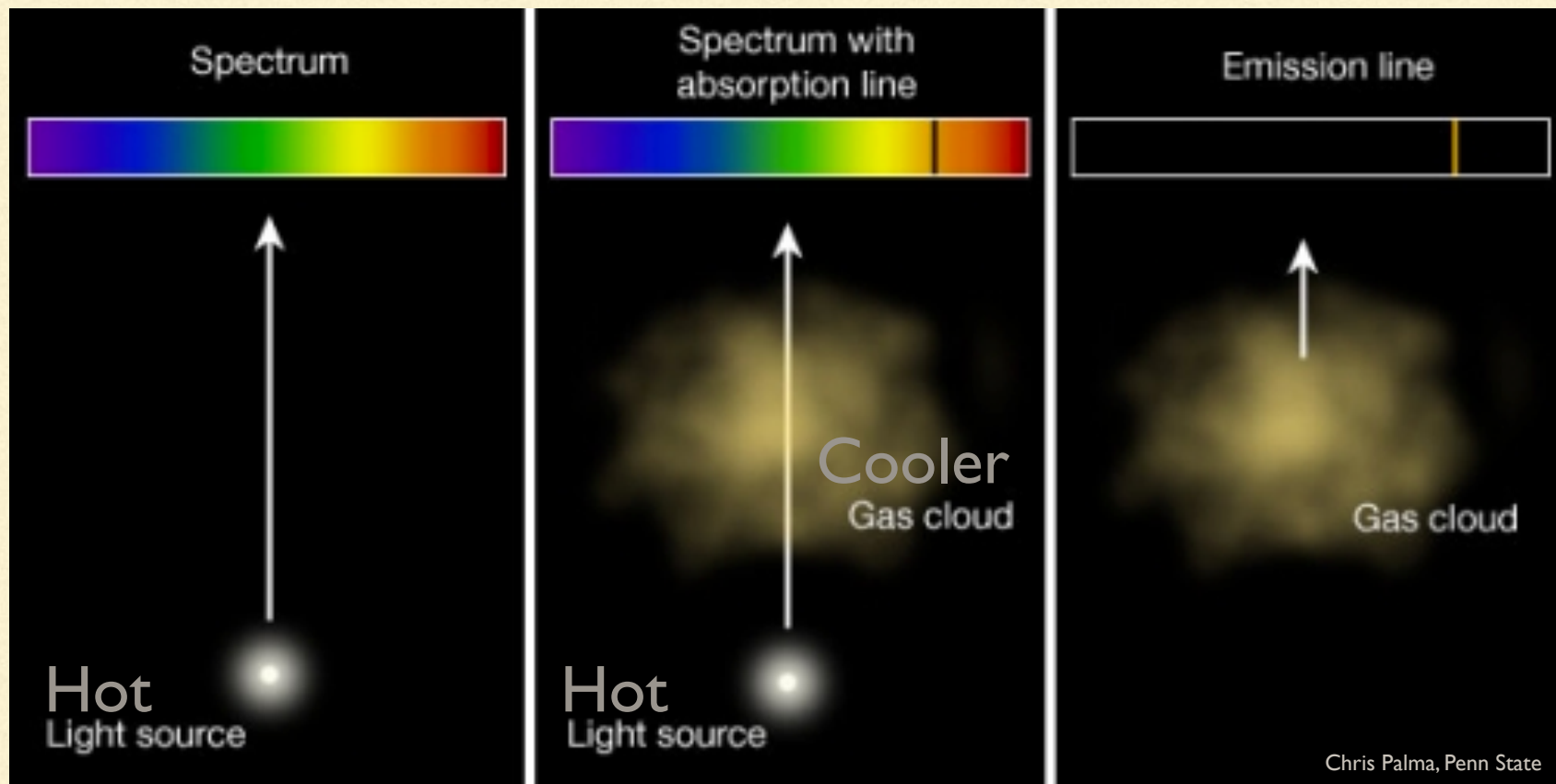
- If a source is resolved by our telescope, we can measure  $I_\lambda$
- If the source is unresolved, we integrate over solid angle and measure the *specific flux* or *flux density*:

$$f_\lambda d\lambda = \oint I_\lambda d\lambda \cos\theta d\Omega$$

(energy / time / area / wavelength)



# KIRCHHOFF'S LAWS





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# MOMENTS OF SPECIFIC INTENSITY

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- Mean intensity  $J_\lambda = \langle I_\lambda \rangle = (1/4\pi) \oint I_\lambda d\Omega$
  - Specific flux,  $f_\lambda$   $F_\lambda d\lambda = \oint I_\lambda d\lambda \cos\theta d\Omega$
  - Specific energy density  $u_\lambda = (4\pi/c) \langle I_\lambda \rangle$
  - Pressure  $P = (1/c) \int \oint I_\lambda \cos^2\theta d\Omega d\lambda$
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